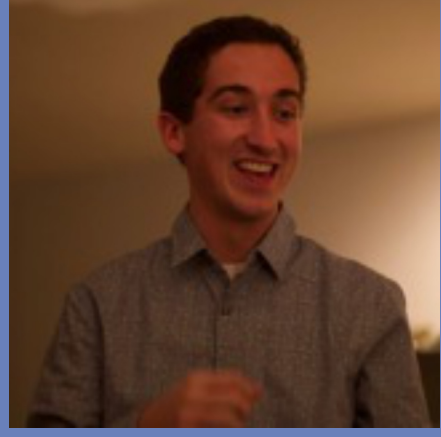


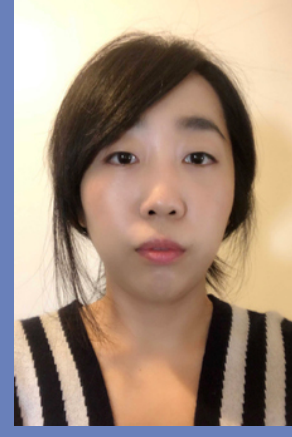
Detecting Dark Matter with a Qubit



Ankur Agrawal



Morgan Lynn

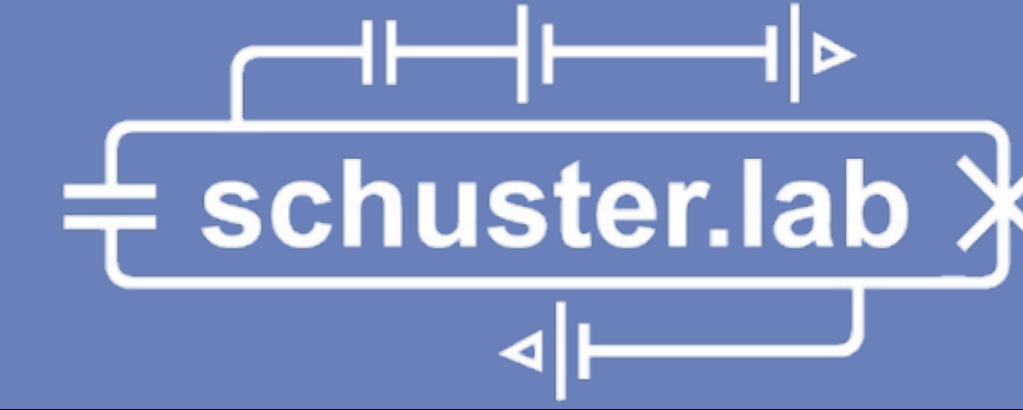
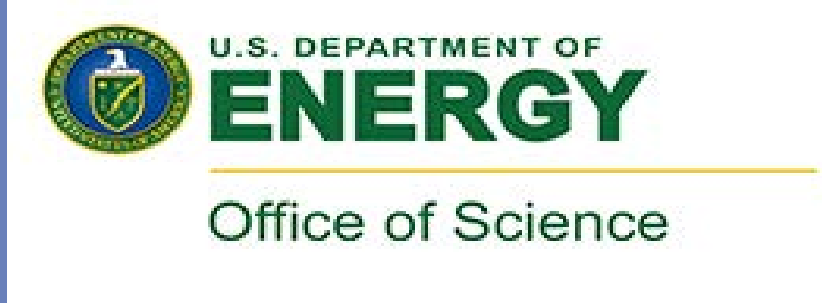


Fang Zhao



Aakash Dixit

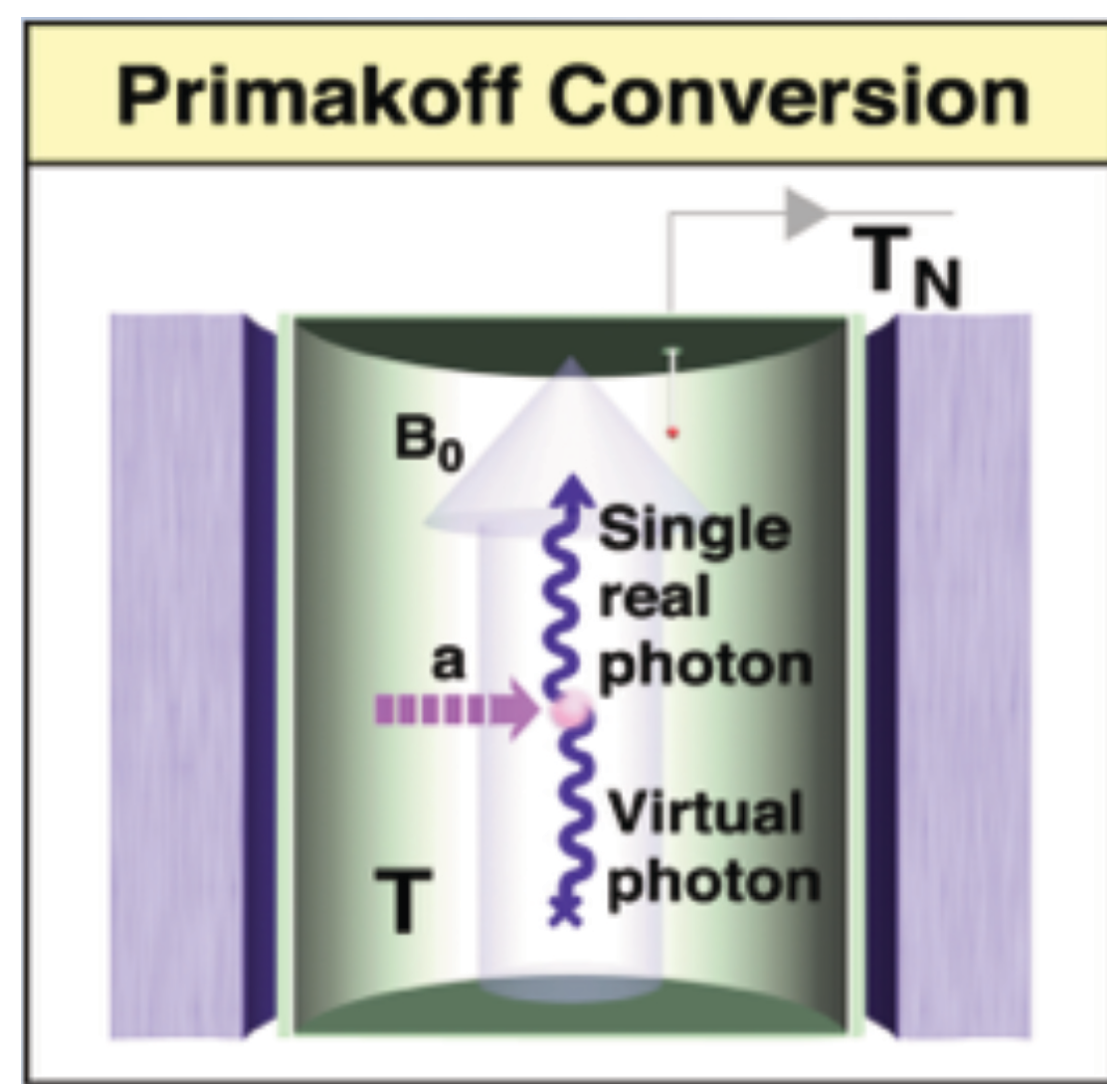
Aaron Chou, David Schuster



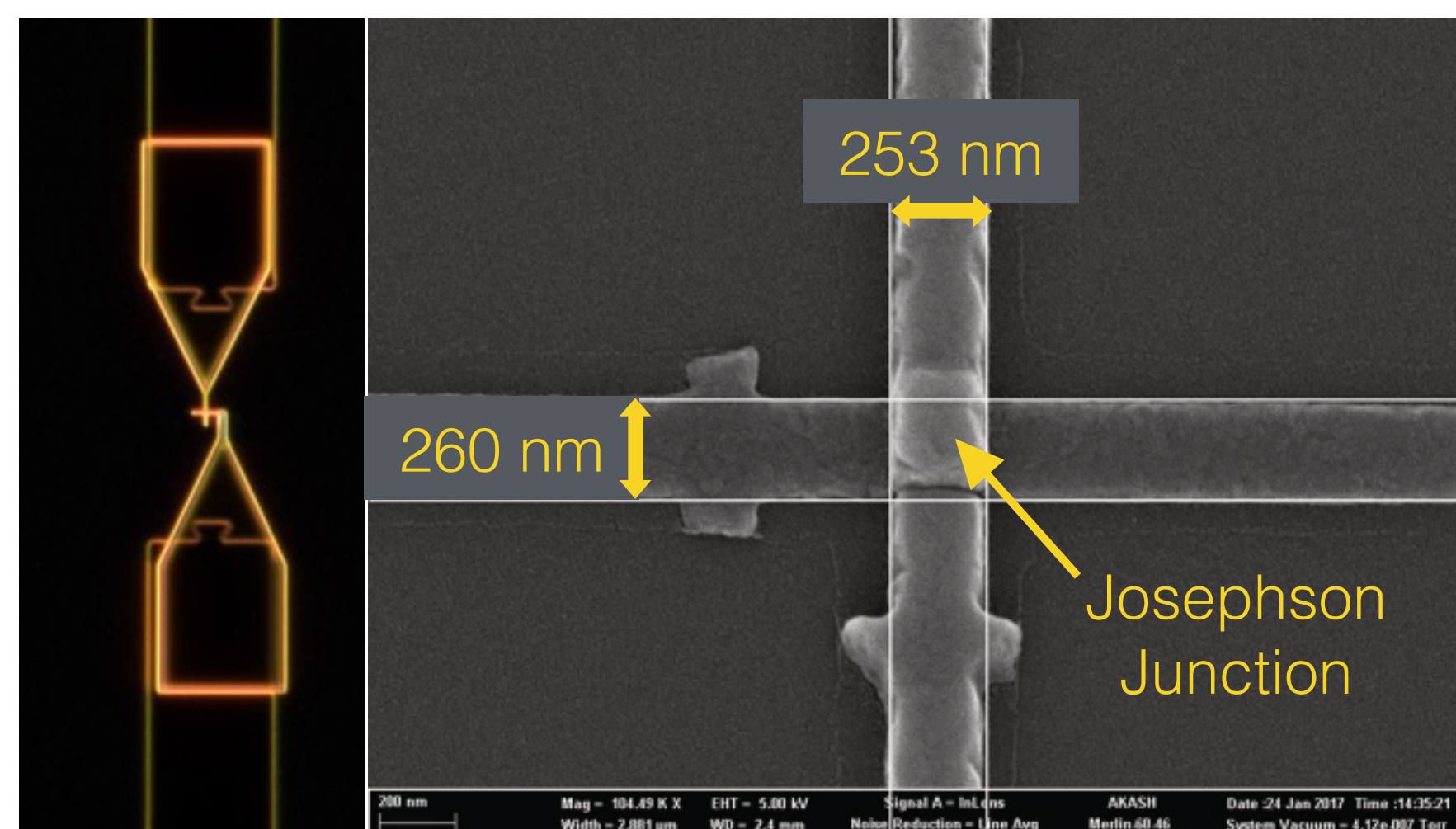
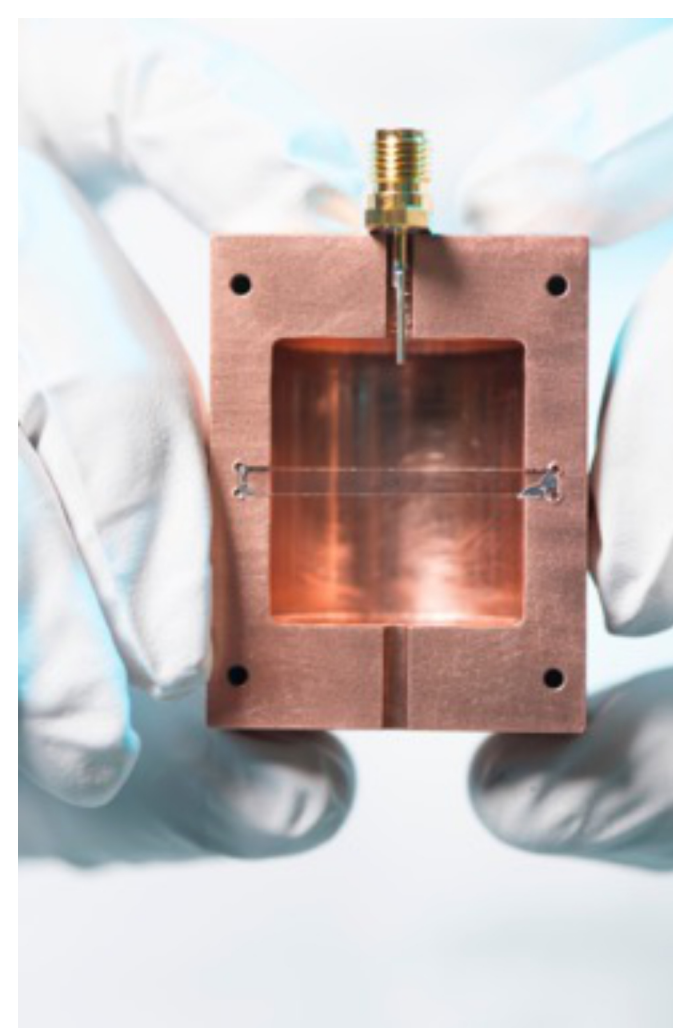
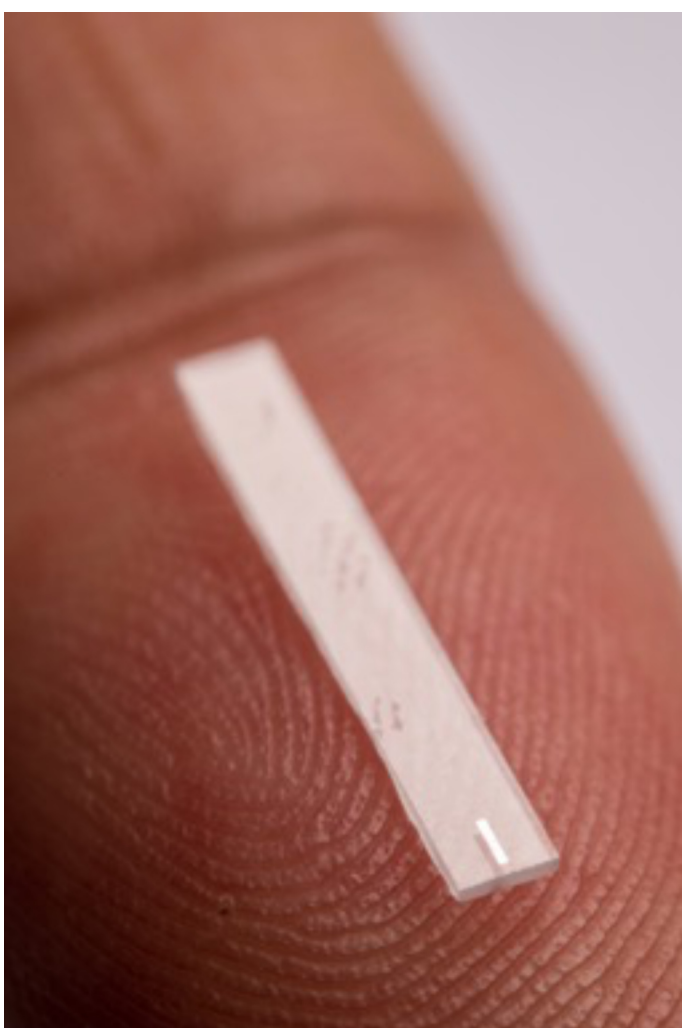
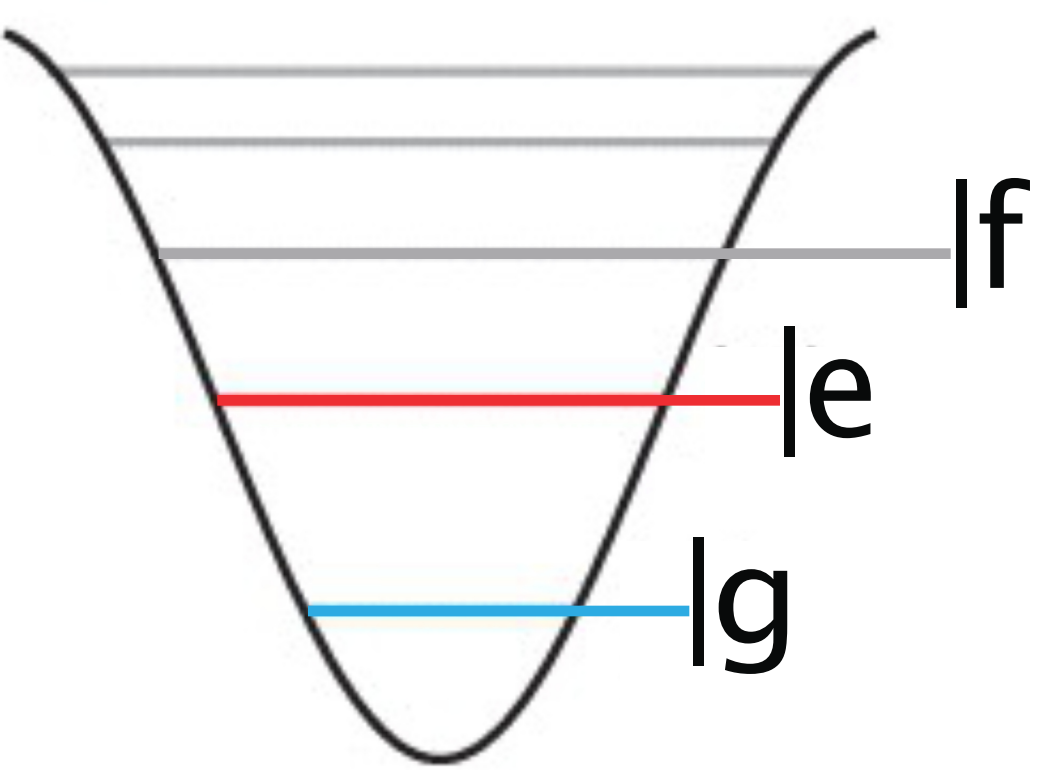
Axion Dark Matter

Axion model provides a solution to the Strong CP problem and accounts for the observed dark matter density

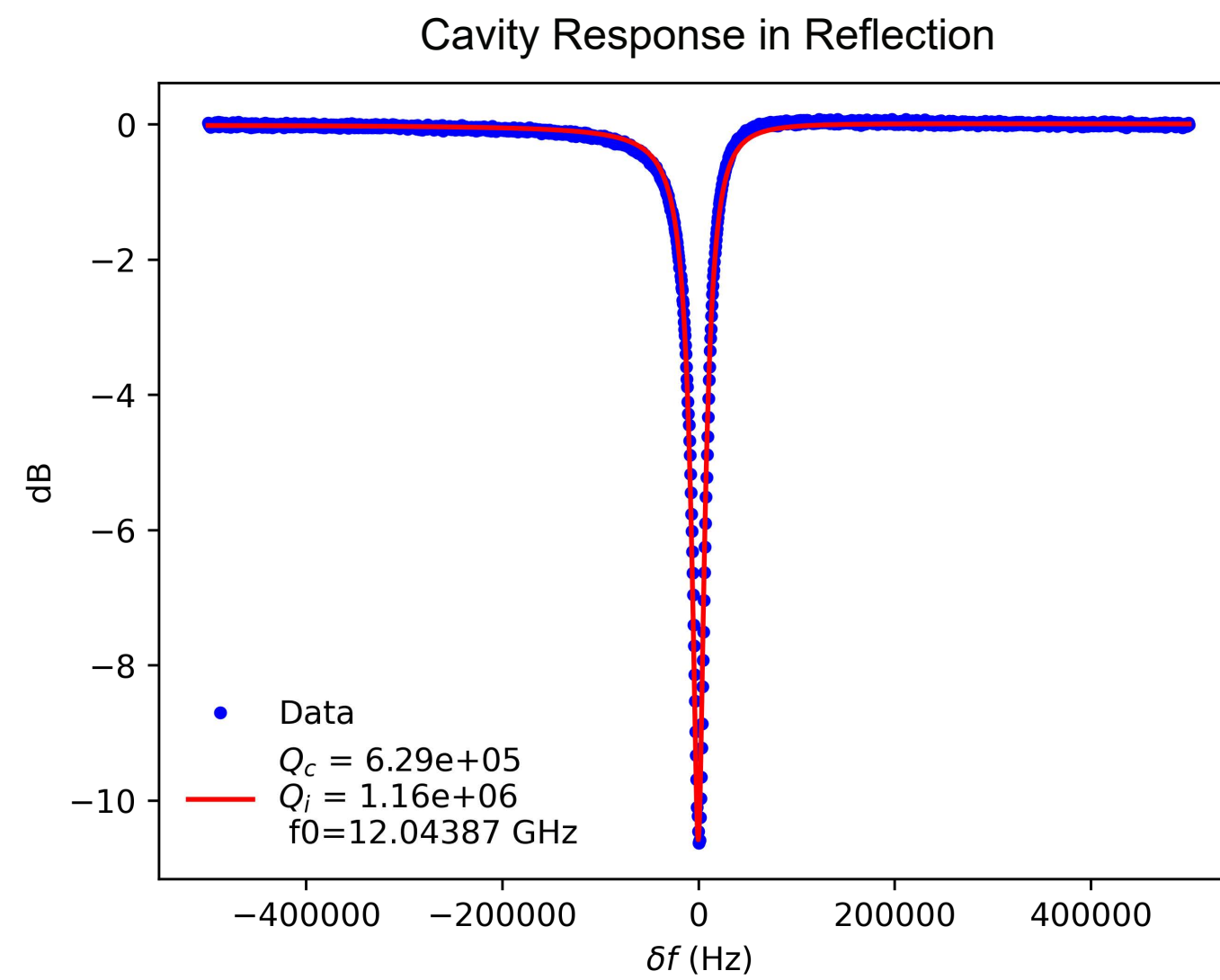
$$\mathcal{L} = g_{a\gamma\gamma} a \vec{E} \cdot \vec{B}_0$$



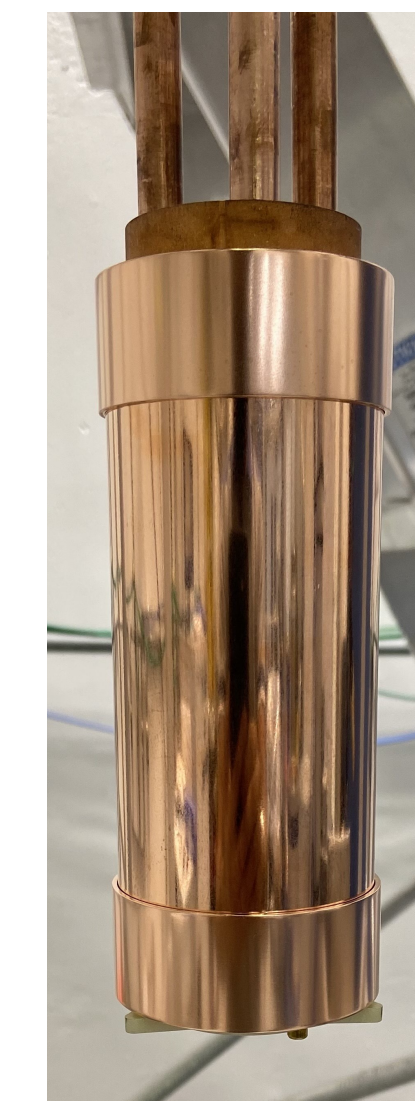
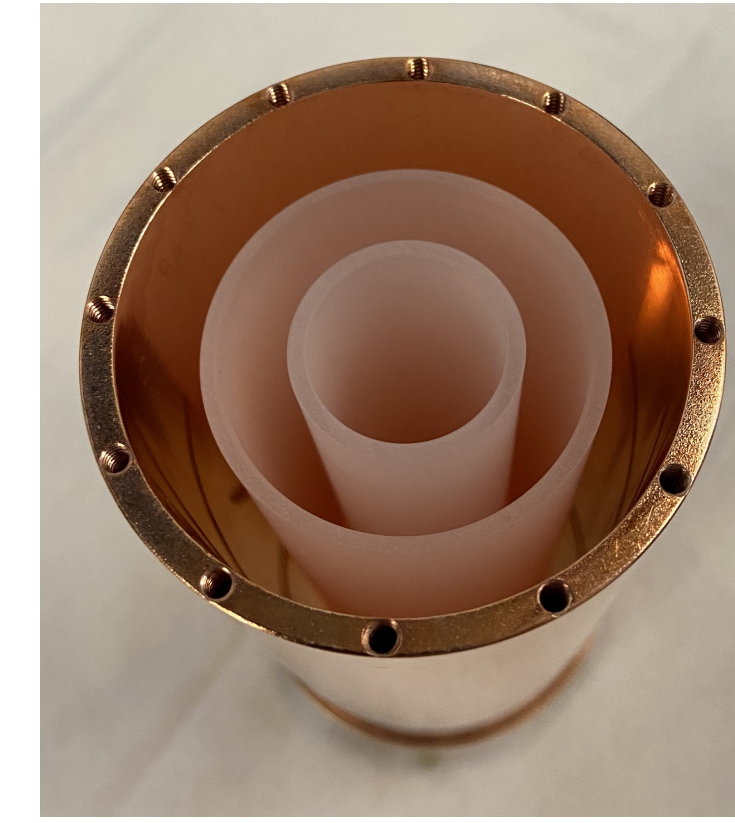
Use a superconducting qubit to enhance and detect dark matter



High Q Bandgap Cavity



Bandgap produced by arrangement of dielectric structures prevents radiation from leaving structure and can support high Q (~1e6) electromagnetic modes.



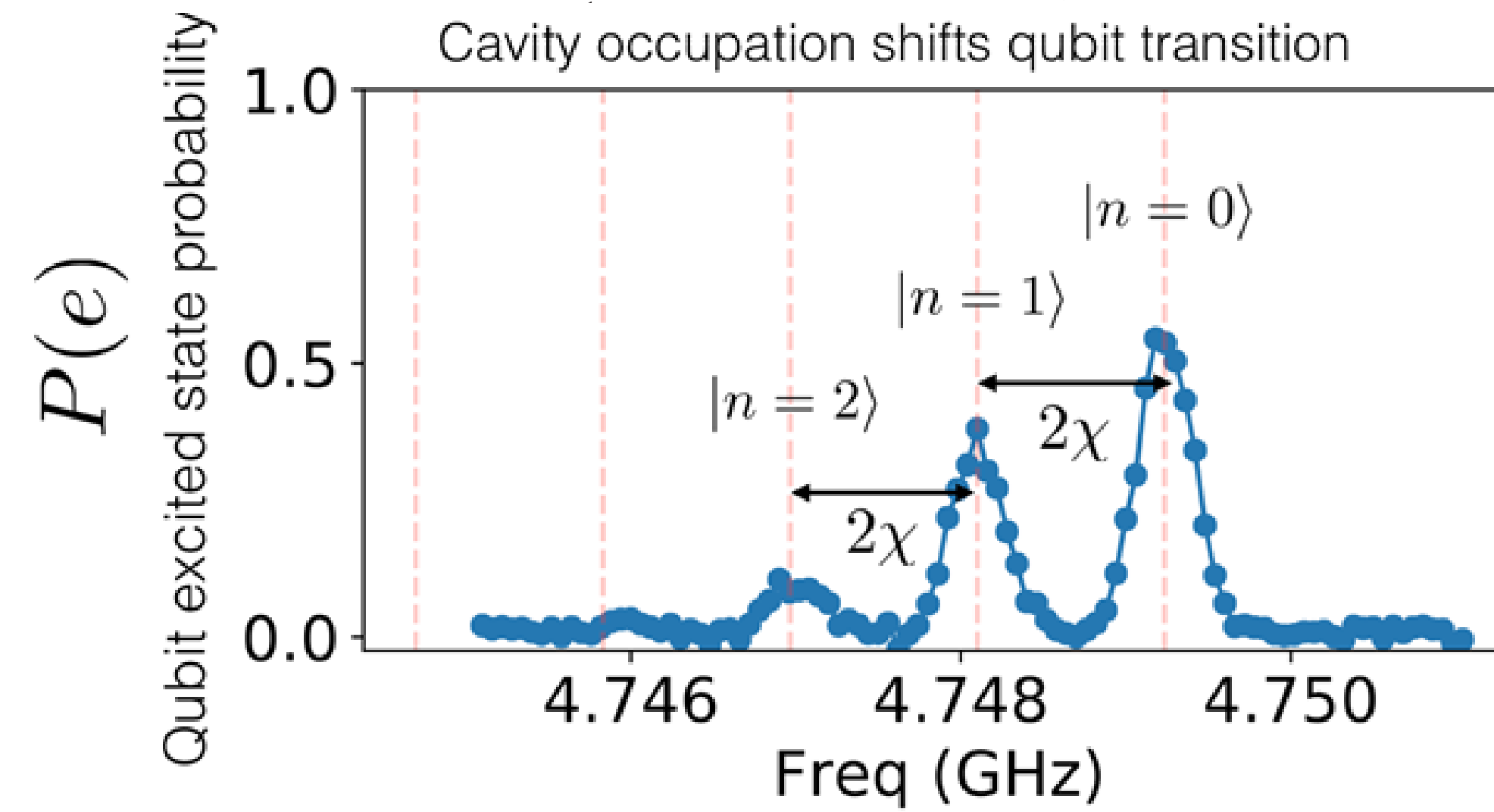
Photon counting with a qubit

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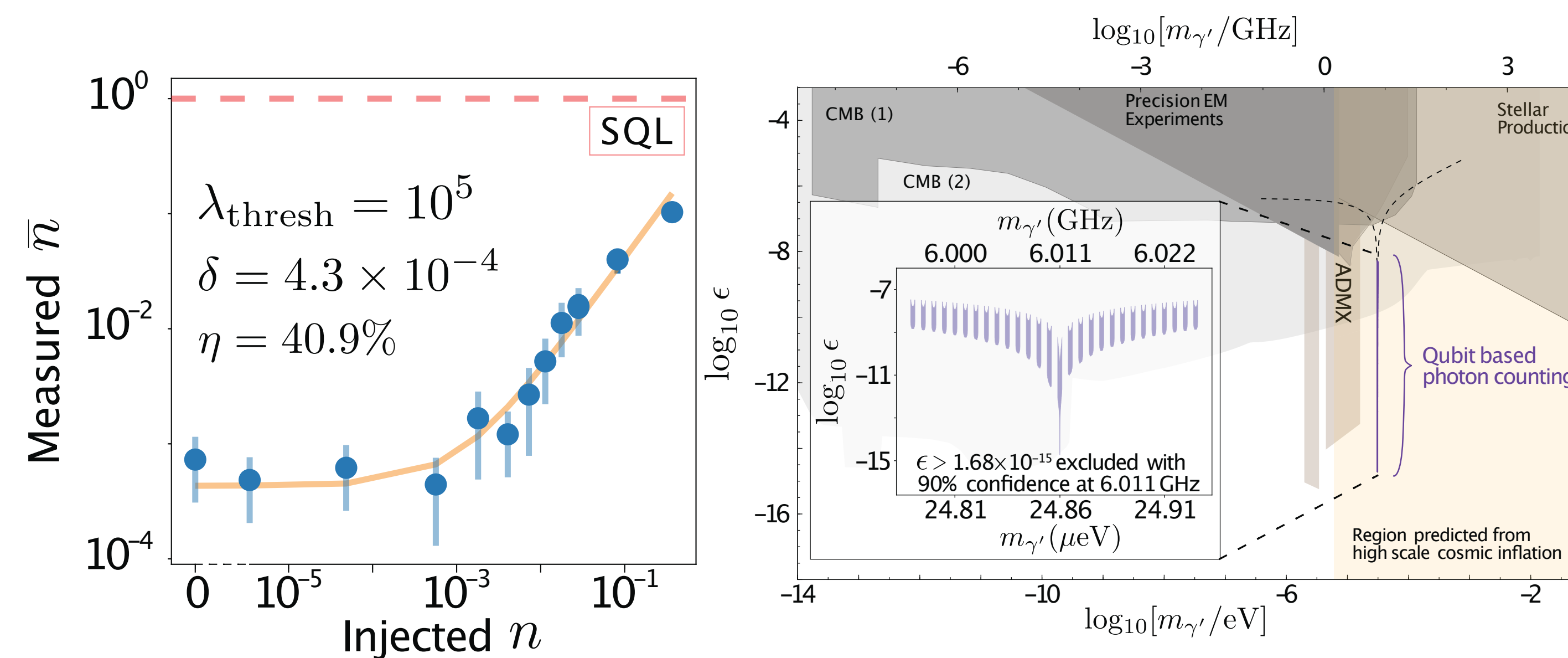
Searching for Dark Matter with a Superconducting Qubit

Aakash V. Dixit^{1,2,3,*}, Srivatsan Chakram^{1,2,4}, Kevin He^{1,2}, Ankur Agrawal^{1,2,3}, Ravi K. Naik^{1,2,3}, David I. Schuster^{1,2,6} and Aaron Chou⁷

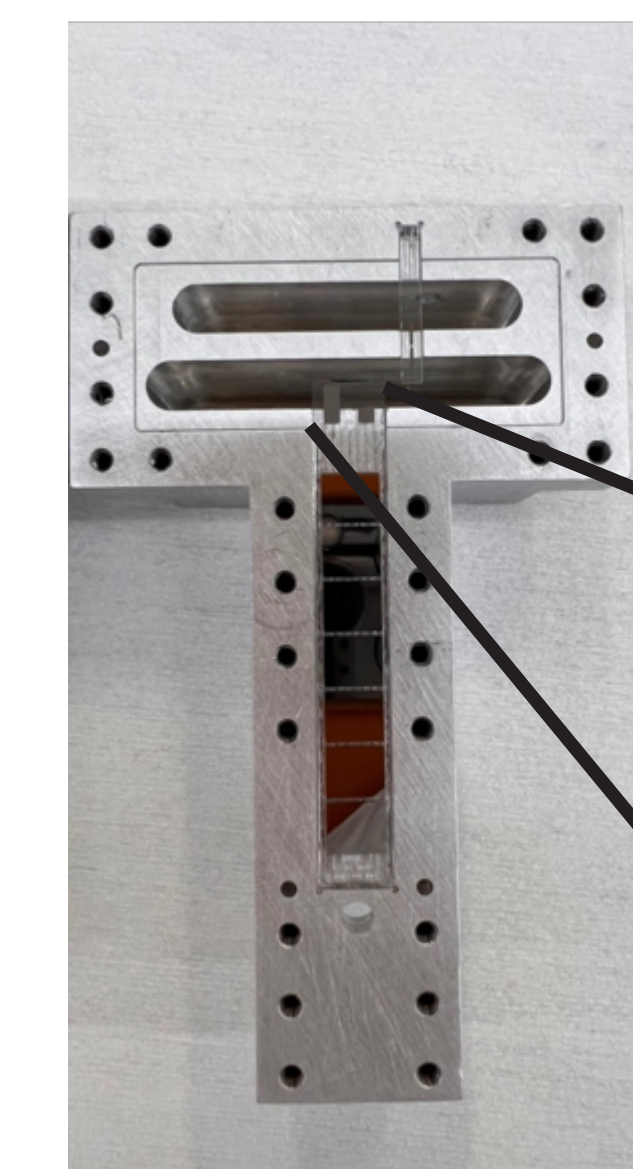


Cavity occupation imprinted on qubit transition frequency

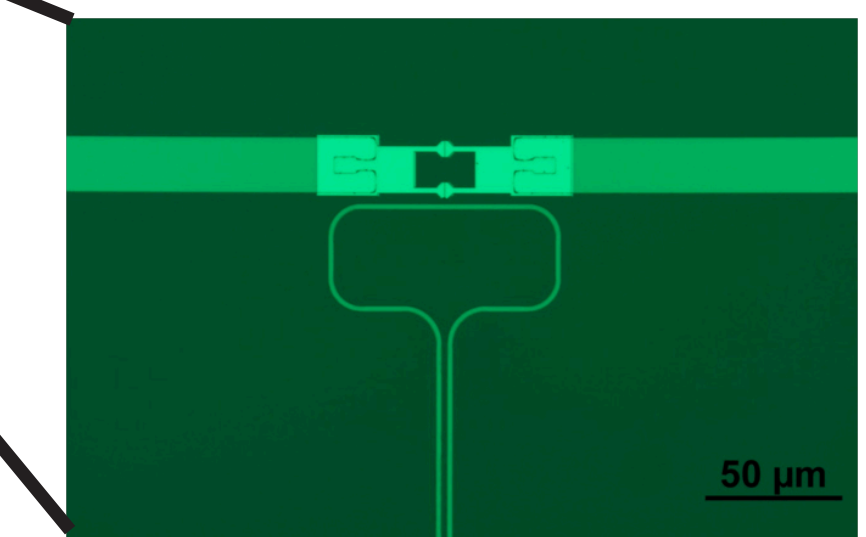
$$\mathcal{H} = \omega_c a^\dagger a + \frac{1}{2} (\omega_q + 2\chi a^\dagger a) \sigma_z$$



Electronic tuning of a cavity

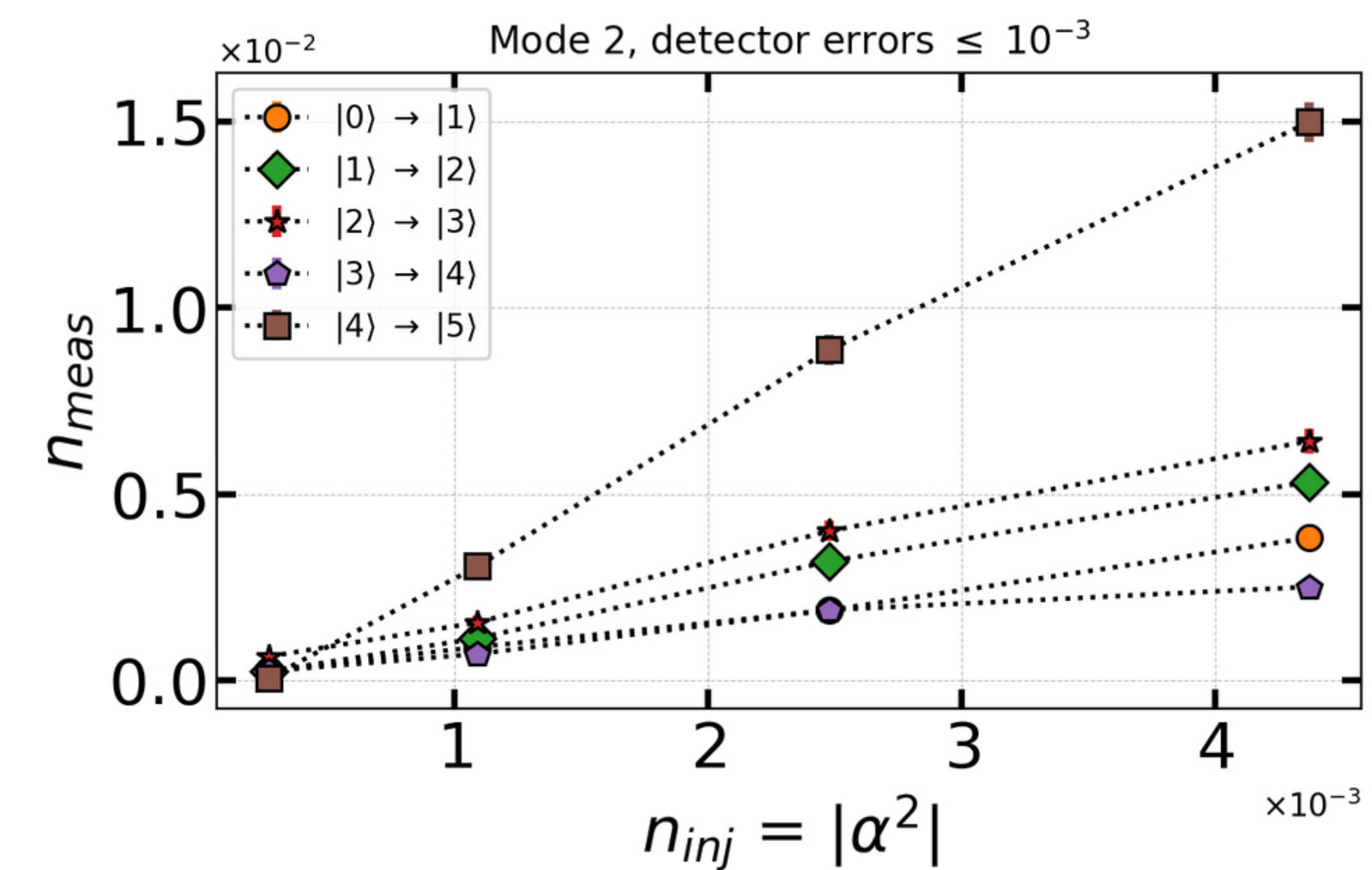
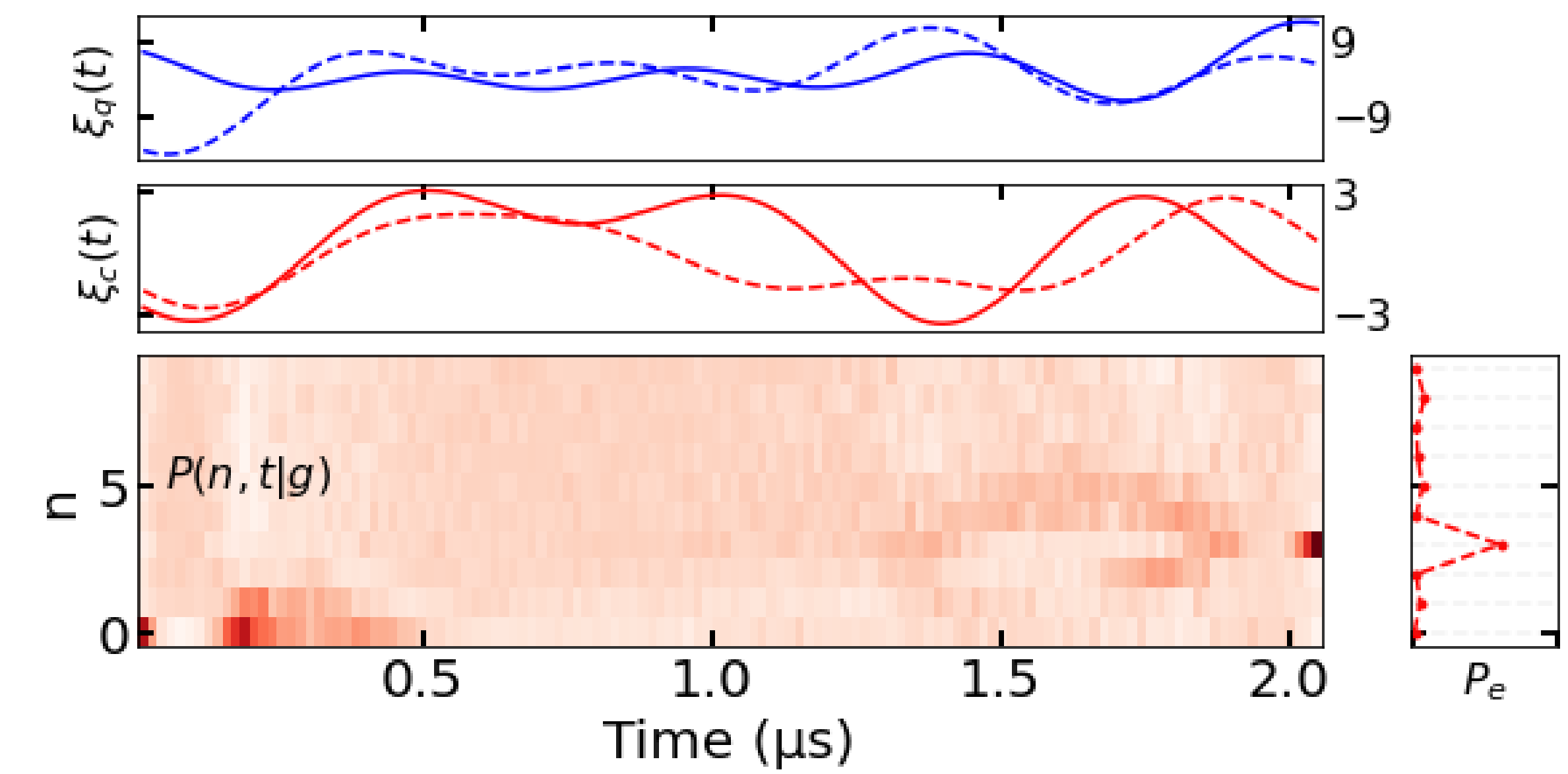


Tuning is required to search a range of frequencies. Applying magnetic field to loop with two Josephson junctions tunes frequency.



Enhancement by state preparation

$$|\langle n+1 | \hat{D}(\alpha) | n \rangle|^2 \propto (n+1)$$



Starting cavity with more photons results in a more efficient axion search