

BREAD: Broadband Reflector Experiment for Axion Detection



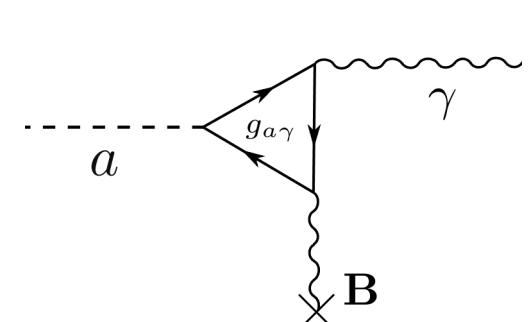
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Motivation

Axions are excellent dark matter candidates [36-41], initially introduced to explain the smallness of the neutron electric dipole moment [29-31]. Axion direct detection experiments like ADMX [52] are typically only sensitive to axions of about $\sim \mu\text{eV}$ mass but unsuitable around $\sim \text{meV}$, favored by a cosmological scenario where most axions are effectively produced after inflation [65-67].

Concept

- 1** Axions can convert to photons on metallic surfaces of area A_{dish} under a strong parallel magnetic field B_{ext} ("dish antenna" concept [68]) giving a signal power of



$$P_{\text{sig}} = 1 \cdot 10^{-25} \text{ W} \cdot \left(\frac{A_{\text{dish}}}{10 \text{ m}^2} \right) \left(\frac{B_{\text{ext}}}{10 \text{ T}} \right)^2 \times \left(\frac{\rho_{\text{DM}}}{0.45 \text{ GeV cm}^{-3}} \right) \left(\frac{g_{a\gamma\gamma}}{3.9 \cdot 10^{-16} \text{ GeV}^{-1}} \right)^2 \left(\frac{1 \mu\text{eV}}{m_a} \right)^2$$

ρ_{DM} : Local Dark Matter Density, $g_{a\gamma\gamma}$: Axion-Photon-Coupling, m_a : Axion Mass

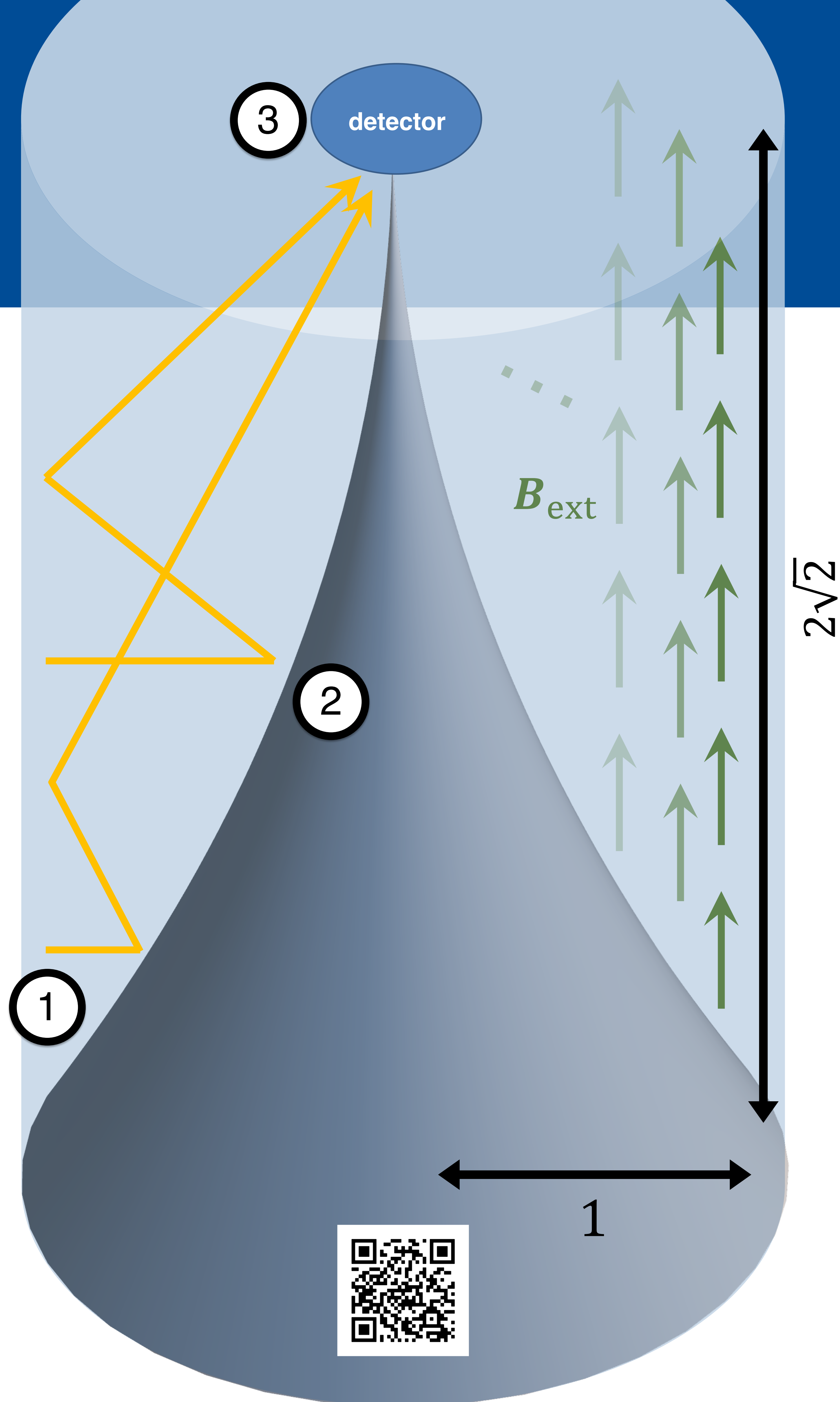
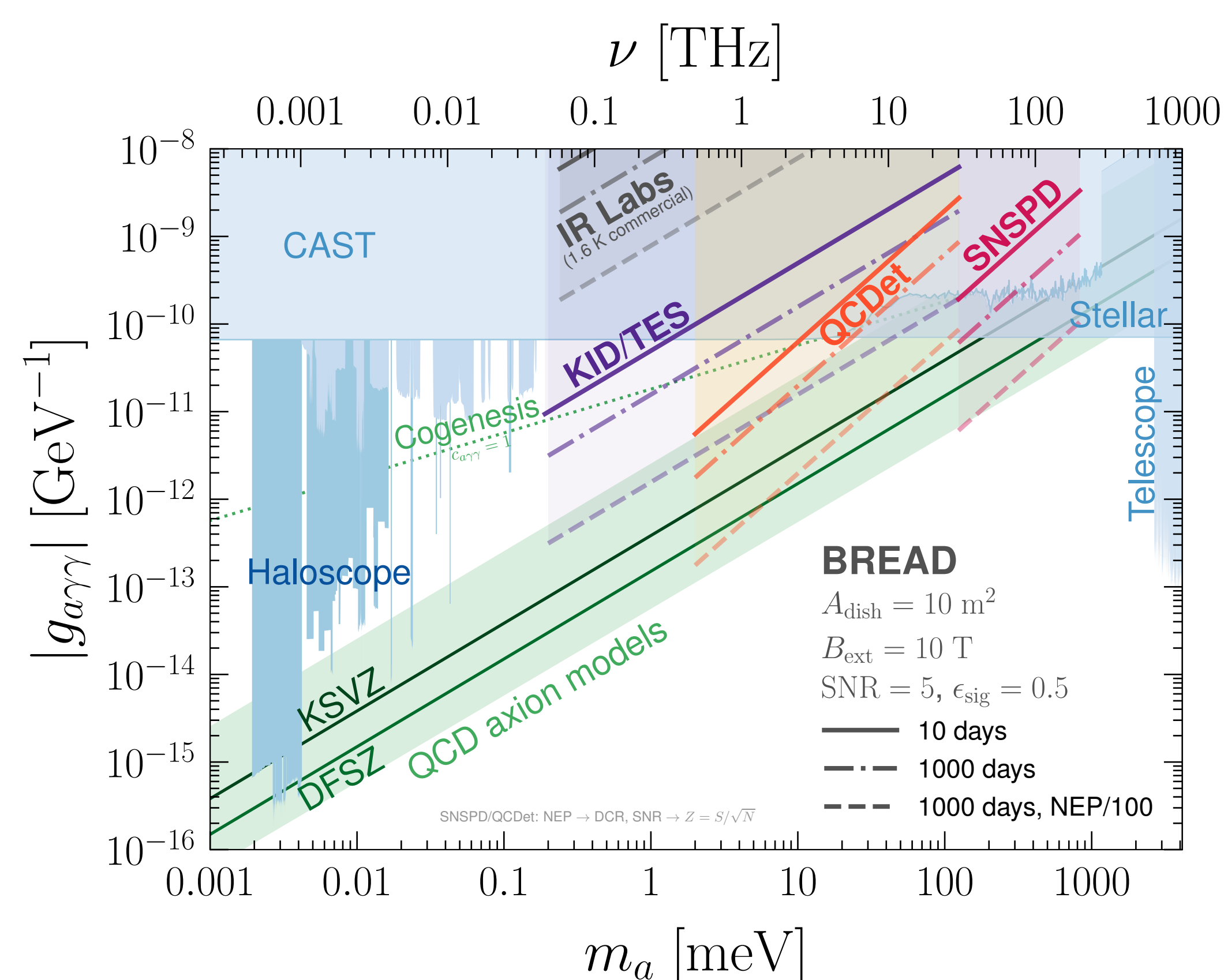
- 2** Our design uses a cylindrical metallic surface and a **parabolic surface of revolution (reflector)** to focus the generated radiation. This allows using a **solenoid magnet**, drastically simplifying magnet design and reducing cost compared to other designs [68-73].

- 3** A **photosensor** is placed at the focus. Candidates for different photon energy ranges E are

Photosensor	$\frac{E}{\text{meV}}$	$\frac{T_{\text{op}}}{\text{K}}$	$\frac{\text{NEP}}{\text{W}/\sqrt{\text{Hz}}}$	$\frac{A_{\text{sens}}}{\text{mm}^2}$
Bolometers				
GENTEC [97]	[0.4, 120]	293	$1 \cdot 10^{-8}$	$\pi 2.5^2$
IR LABS [98]	[0.24, 248]	1.6	$5 \cdot 10^{-14}$	1.5^2
KID/TES [99, 100]	[0.2, 125]	0.3	$2 \cdot 10^{-19}$	0.2^2
Single Photon Counters				
QCDet [101, 102]	[2, 125]	0.015	$\frac{\text{DCR}}{\text{Hz}} = 4$	0.06^2
SNSPD [103, 104]	[124, 830]	0.3	$\frac{\text{DCR}}{\text{Hz}} = 10^{-4}$	0.4^2

T_{op} : Operating Temperature, NEP: Noise Equivalent Power, A_{sens} : Sensor Area, DCR: Dark Count Rate

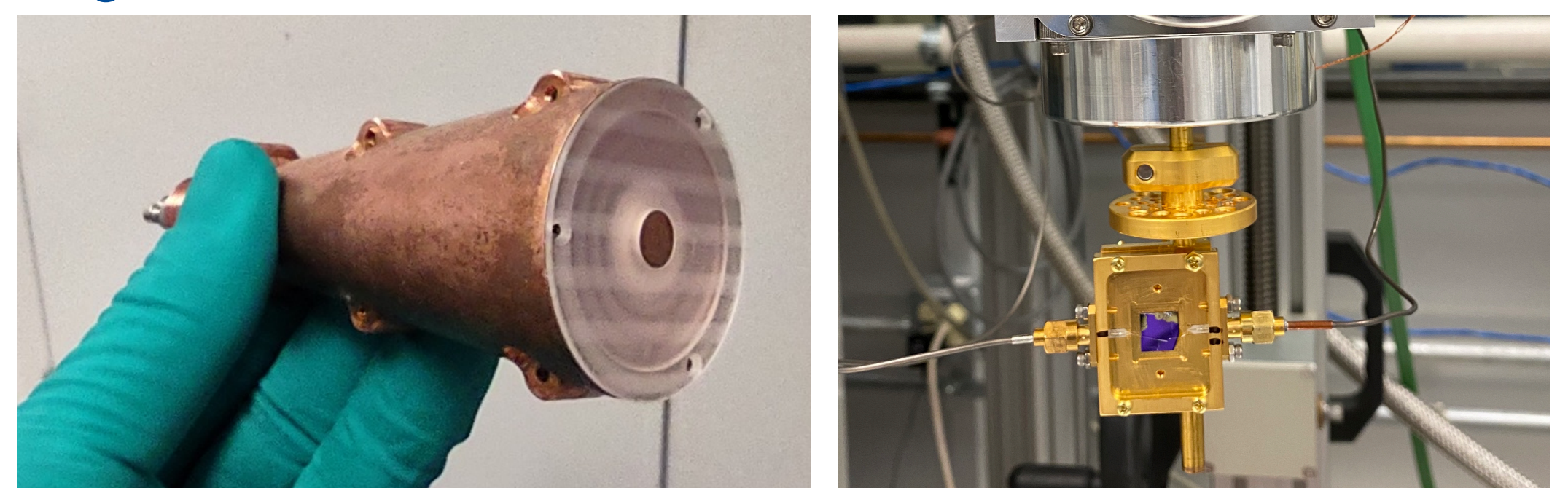
For a 3-year experiment with a signal-to-noise ratio (SNR) of 5 and 50% efficiency the single photon detectors are **sensitive to QCD axions above $\sim \text{meV}$** .



Pilot Experiments

Pilot Experiments in the microwave (**GigaBREAD**) and infrared (**InfraBREAD**) ranges are in preparation at the University of Chicago and Fermilab.

GigaBREAD antenna horn: **InfraBREAD SNSPD**:



Conclusions

BREAD is a promising concept to discover **axions** in a **broad mass range around $\sim \text{meV}$** within the next decade. First pilot experiments are in preparation.

Acknowledgements & References: For detailed authorlist and references please see <https://doi.org/10.1103/PhysRevLett.128.131801> or <https://arxiv.org/abs/2111.12103>.



Full-wave simulation of GigaBREAD featured on the 2022 April issue of Physical Review Letters.