

# Beyond Three-Flavor $\nu$ Oscillations with DUNE

Alexandre Sousa

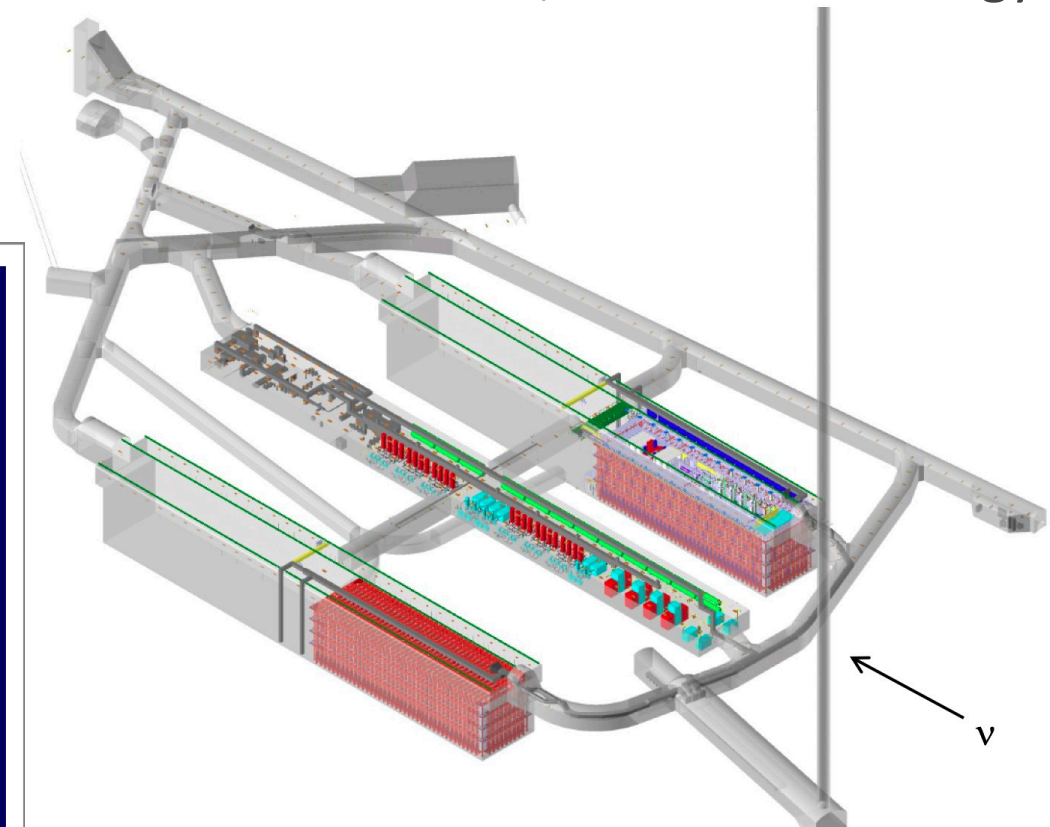
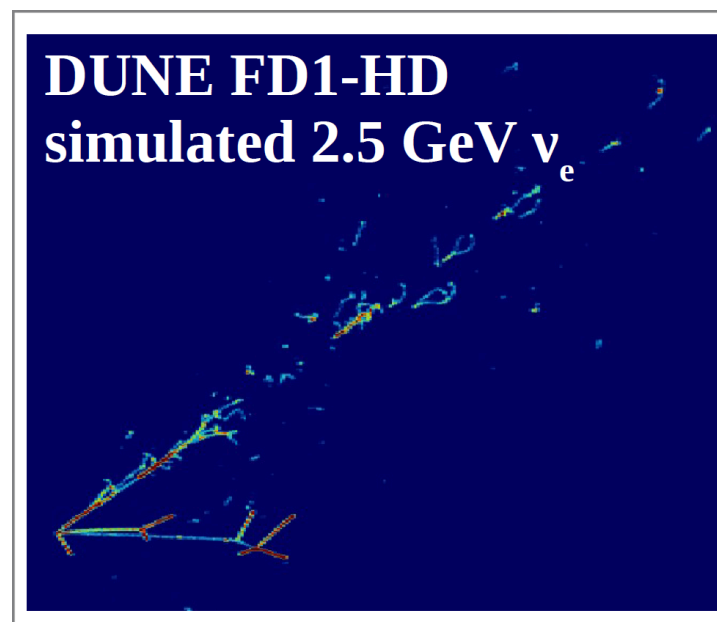
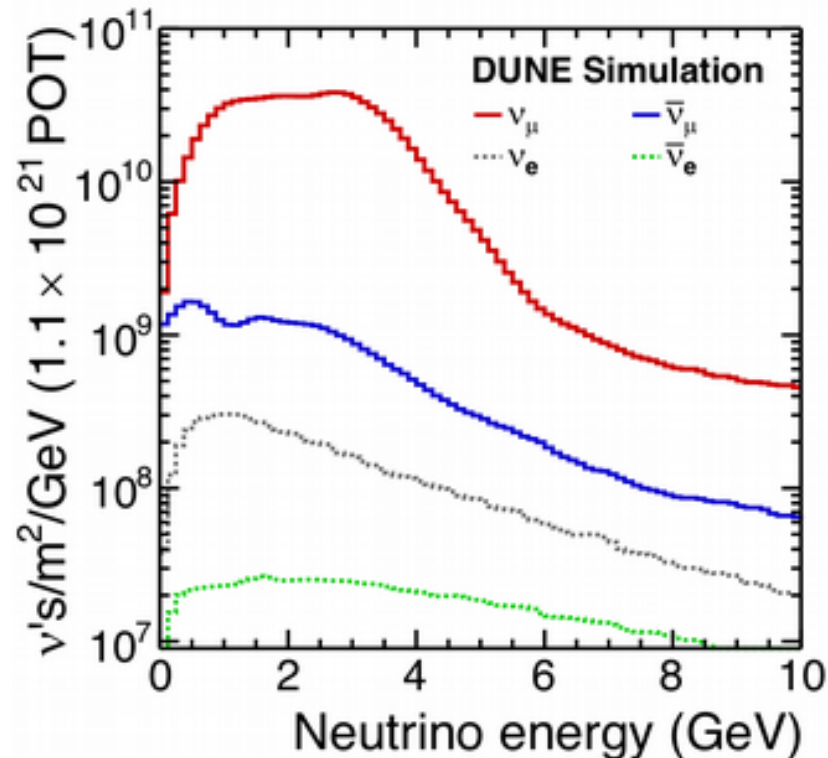
University of Cincinnati, on behalf of the DUNE Collaboration

Snowmass 2021 Community Summer Study, Seattle

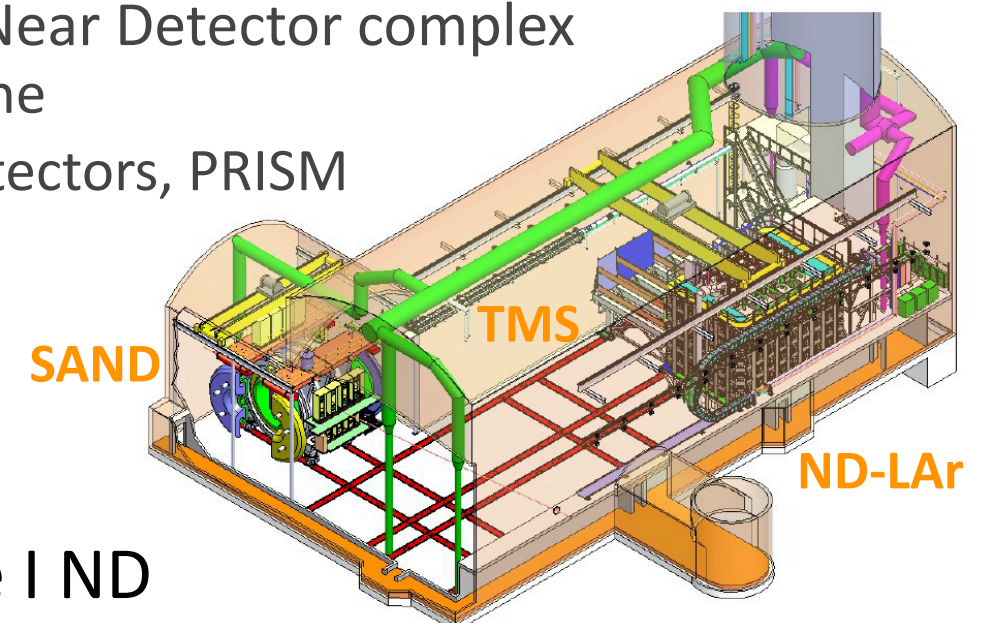
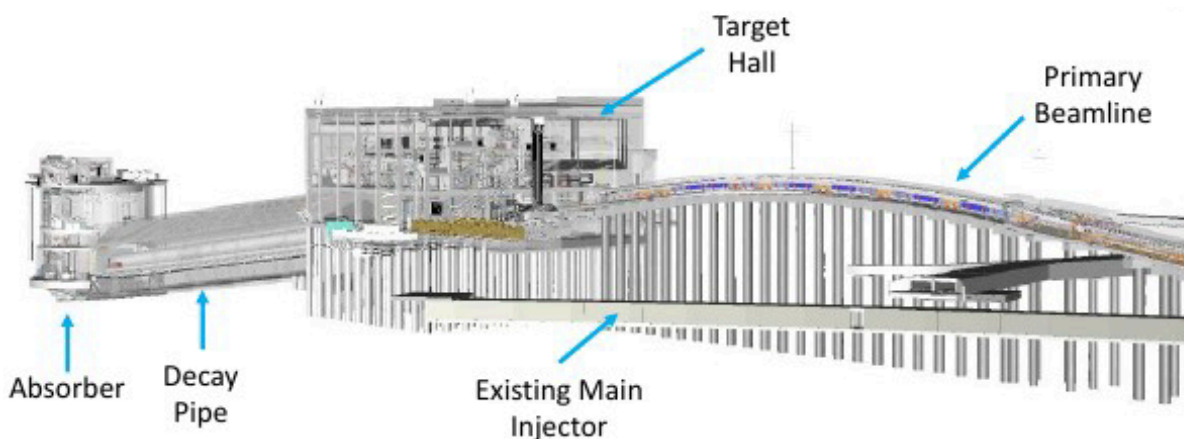
July 19, 2022

# DUNE is a Machine for Discovery!

- ▶ High-intensity wide-band LBNF neutrino beam
  - 1.2 MW upgradeable to 2.4 MW



- ▶ Highly-capable Near Detector complex
  - 574 m baseline
  - High-Res. Detectors, PRISM

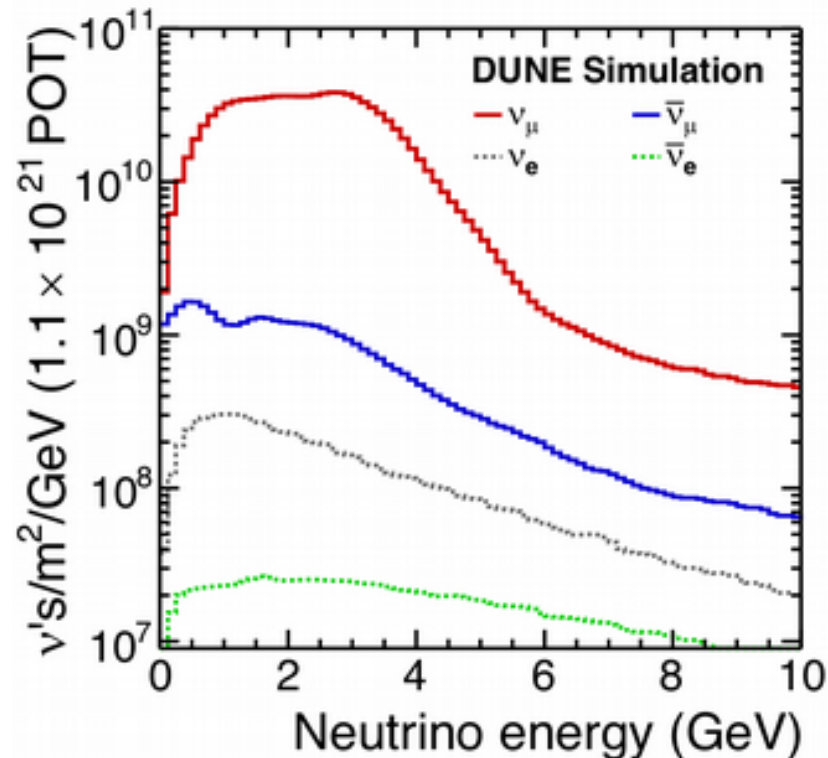


Phase I ND

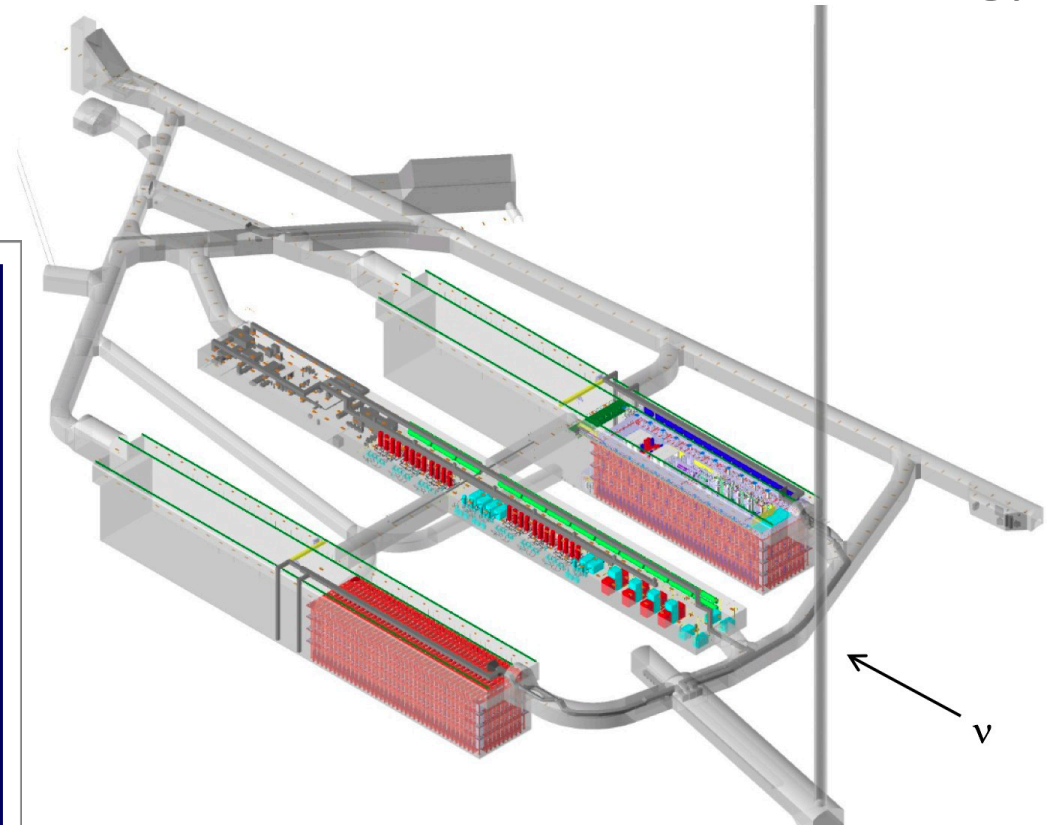
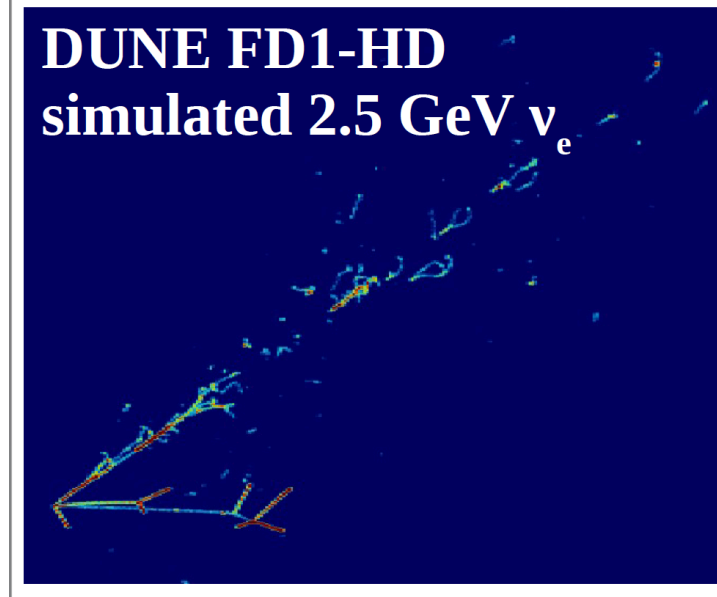


# DUNE is a Machine for Discovery!

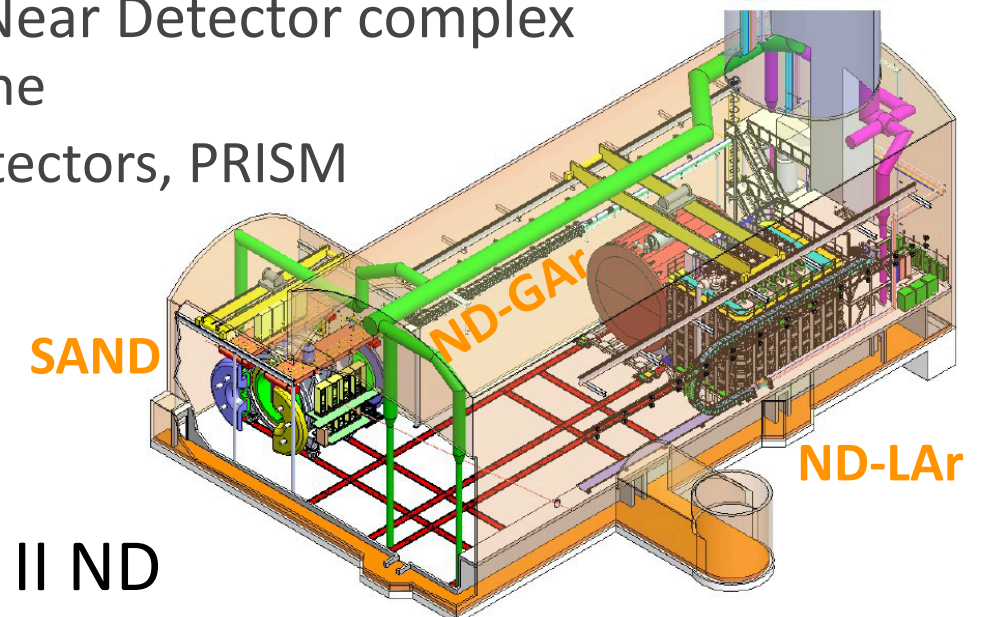
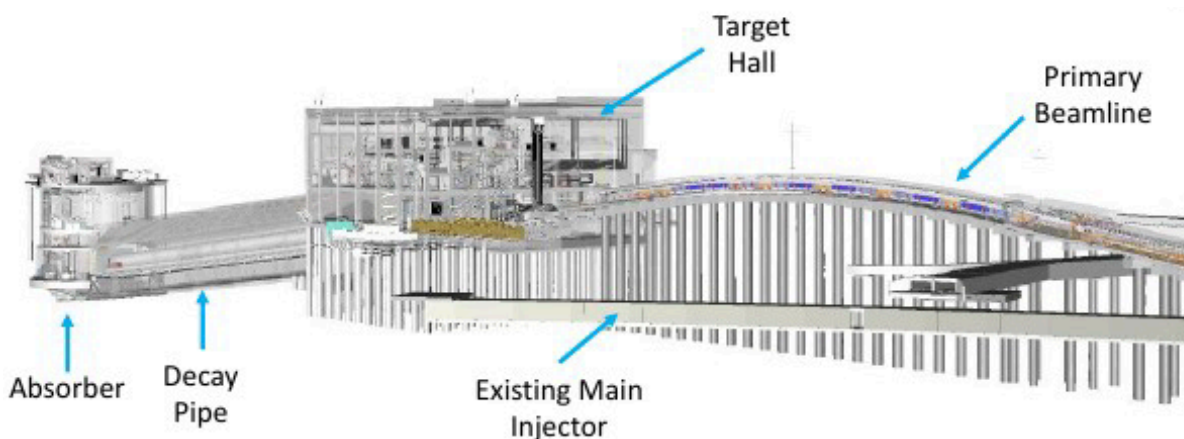
- ▶ High-intensity wide-band LBNF neutrino beam
  - 1.2 MW upgradeable to 2.4 MW



**DUNE FD1-HD**  
simulated  $2.5 \text{ GeV } \nu_e$



- ▶ Highly-capable Near Detector complex
  - 574 m baseline
  - High-Res. Detectors, PRISM



Phase II ND

# DUNE is a Machine for Discovery!

## ► Non-standard short-baseline and long-baseline oscillation phenomena

- Mixing with light sterile neutrinos
- Large extra-dimensions
- Non-unitarity of the mixing matrix
- Non-standard neutrino interactions
- Violation of CPT Symmetry

In this talk!

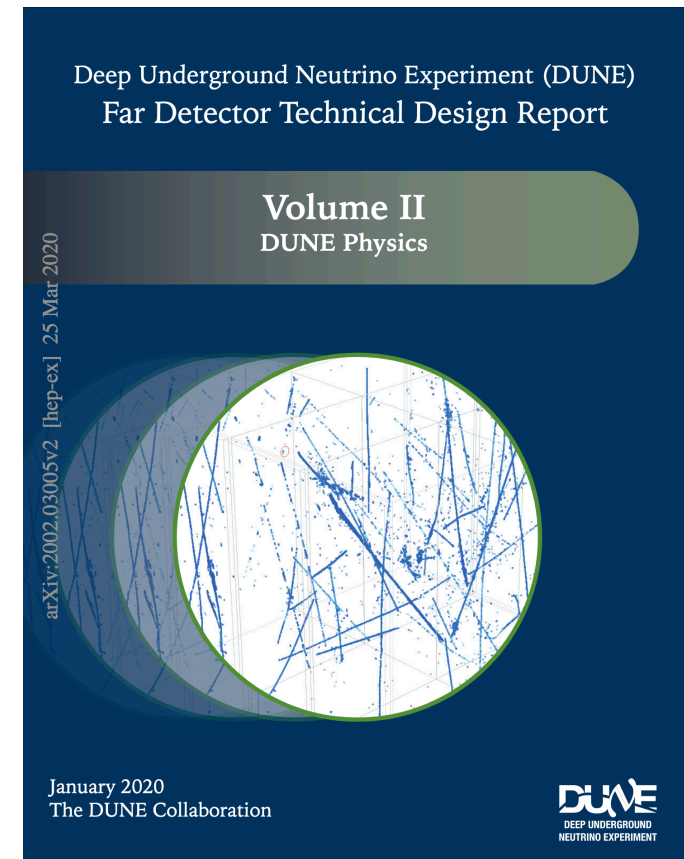
## ► Searches for new phenomena/particles at the ND

- Neutrino trident interactions
- Heavy neutral leptons
- Low-mass dark matter
- Axion-like particles

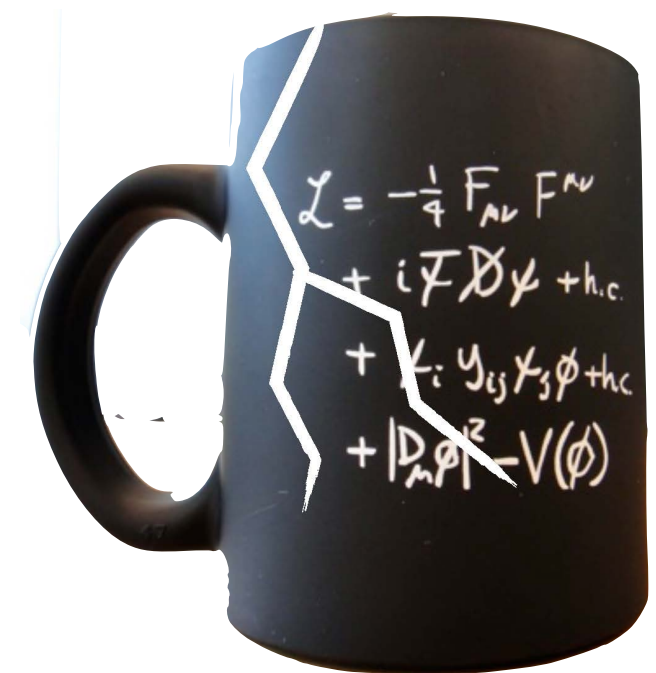
See next talk  
by Jae Yu!

## ► Searches for new phenomena at the FD benefitting from its large mass and high resolution

- Inelastic boosted dark matter from the galactic core
- Boosted dark matter from the Sun
- Nucleon decay



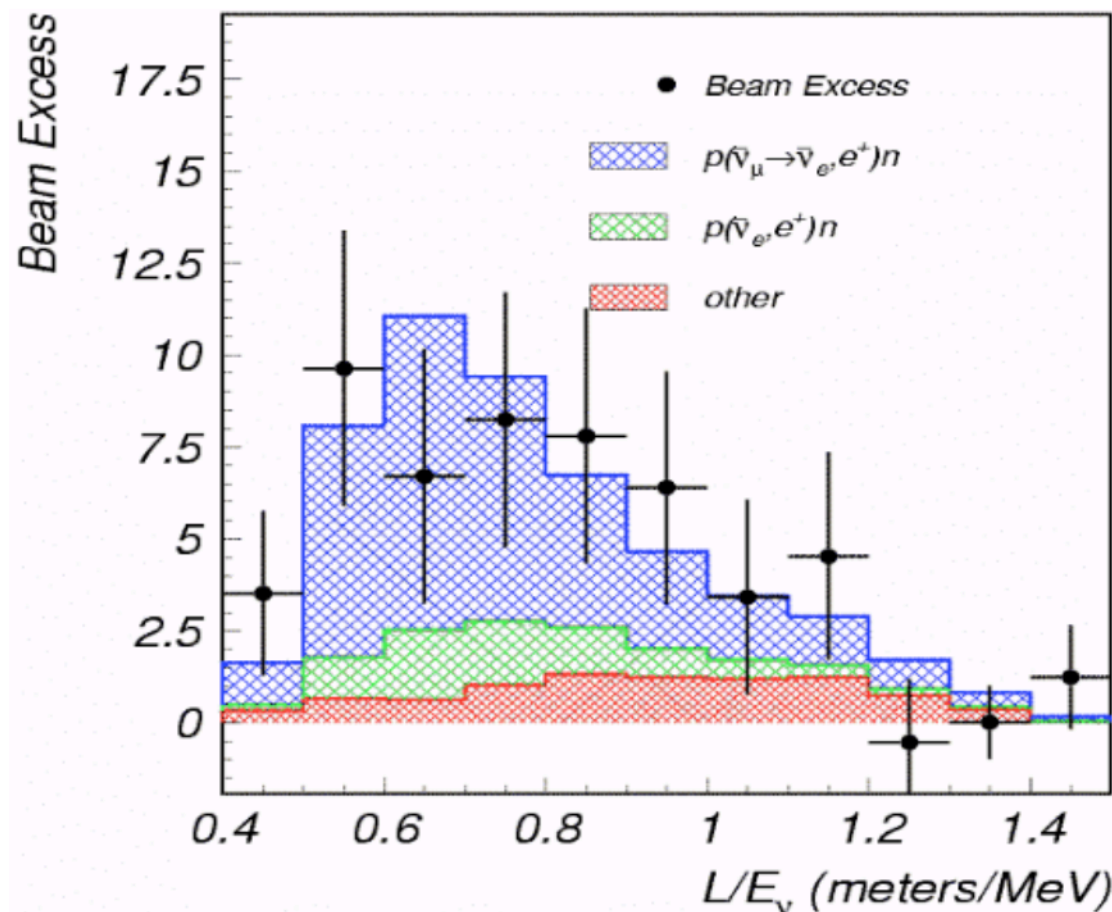
[\*Far Detector Technical Design Report\*](#)  
[\*DUNE BSM Paper, Eur. Phys. J. C 81, 322 \(2021\)\*](#)



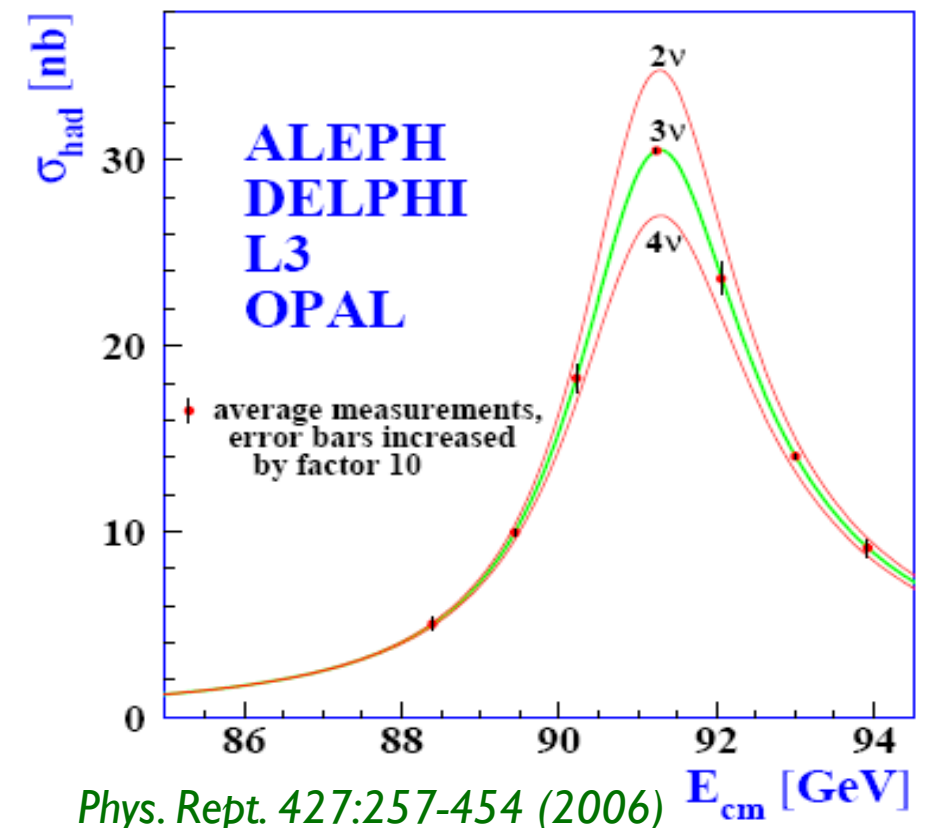
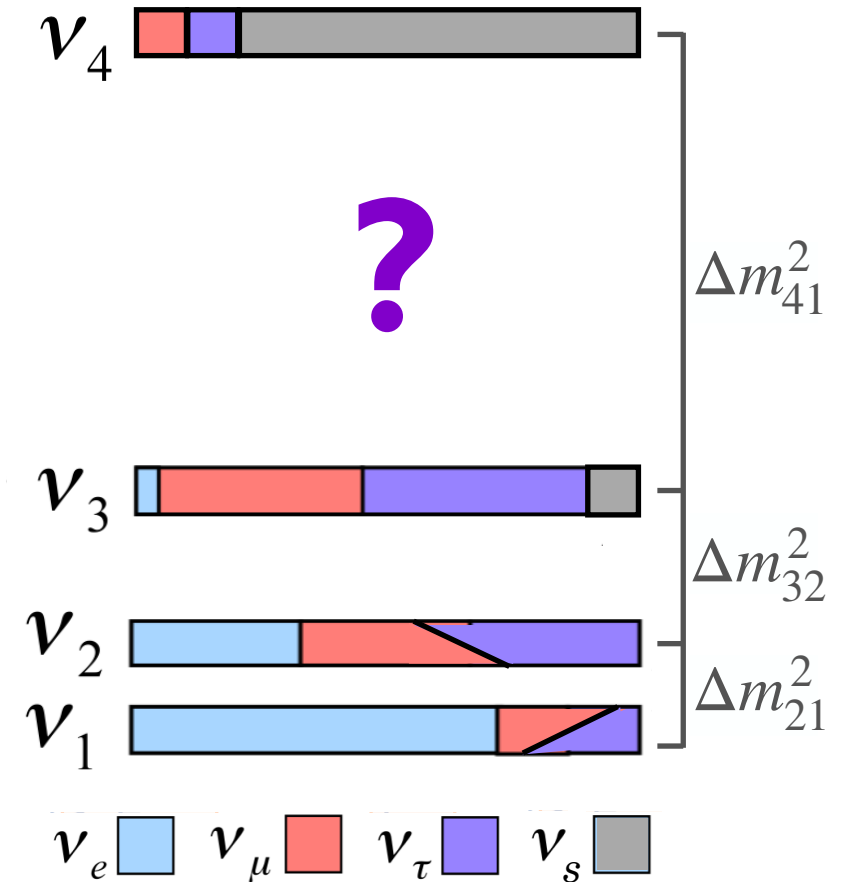


# More than Three Neutrinos?

- ▶ The LSND experiment measured a  $3.8\sigma$  excess of  $\nu_e$  in a DAR  $\nu_\mu$  beam over a very short baseline ( $\sim 30$  m).
- ▶ Oscillation explanation requires 4th neutrino state with  $\Delta m_{41}^2 \sim 1 \text{ eV}^2$ 
  - $Z^0$  width measured at LEP  $\Rightarrow$  only 3 light active neutrinos
  - 4th neutrino is very heavy or has no weak interactions  $\Rightarrow$  **Sterile neutrino ( $\nu_s$ )**



LSND, Phys. Rev. D 64, 112007 (2001)



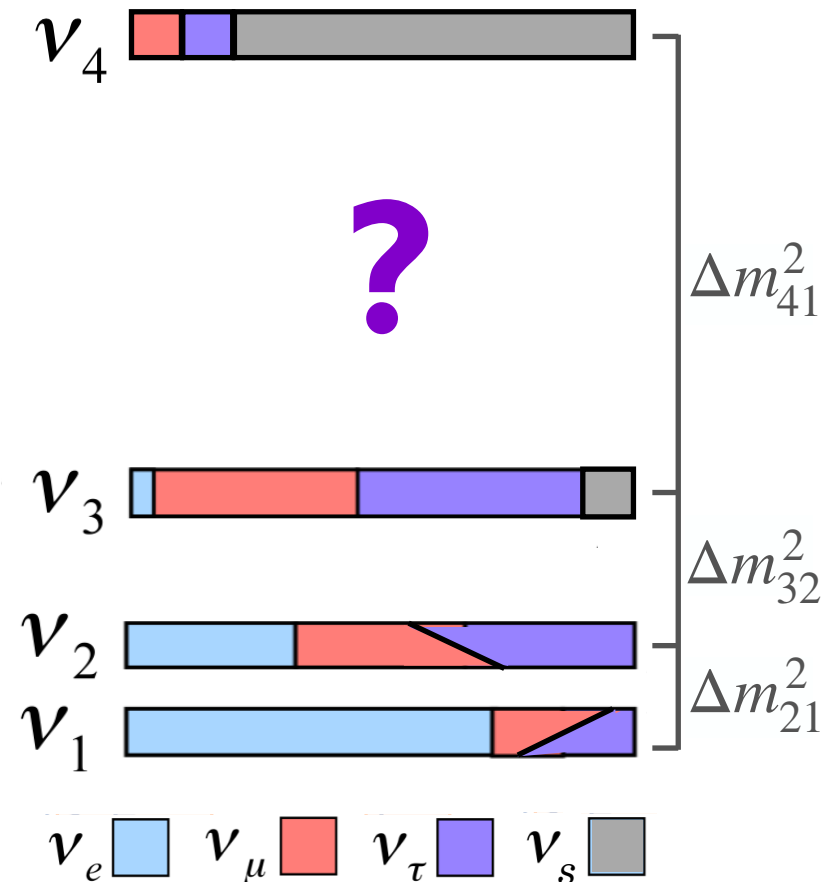
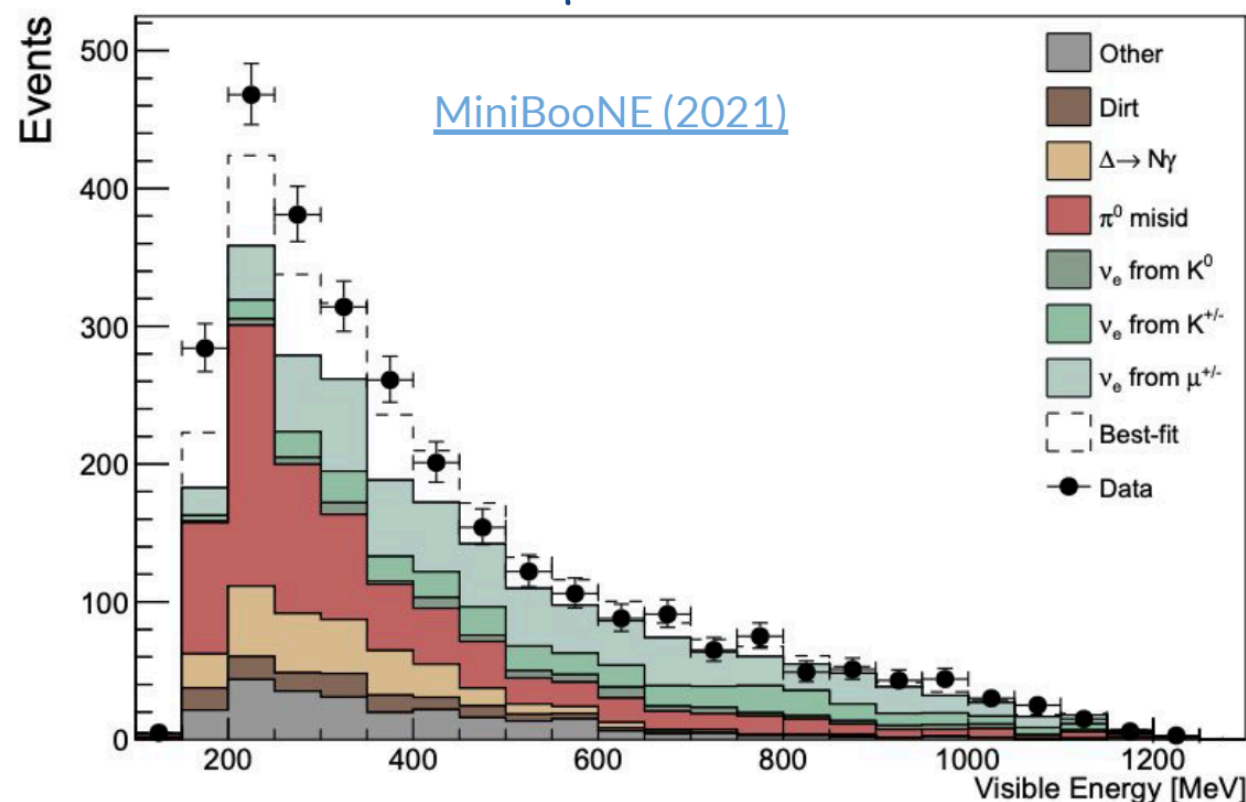
Phys. Rept. 427:257-454 (2006)

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$\Rightarrow$  **Sterile neutrino ( $\nu_s$ )**

excess( $\bar{\nu}_e$ ) in a ( $\bar{\nu}_\mu$ ) dominated beam,  $4.8\sigma$



- ▶ In a 3+1 model, have 1 new mass scale,  $\Delta m_{41}^2$ , 3 new mixing angles,  $\theta_{14}$ ,  $\theta_{24}$ ,  $\theta_{34}$ , and 2 new CP phases  $\delta_{14}$ ,  $\delta_{24}$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix}$$

- ▶ Further motivated by MiniBooNE, Reactor, Gallium anomalies



# Looking for Light Sterile Neutrinos

- ▶  $\nu_e, \bar{\nu}_e$  CC appearance (LSND, KARMEN, mBooNE,  $\mu$ BooNE, etc.):  $\theta_{14}, \theta_{24}$

$$P_{\mu e} \approx 2 \sin^2 2\theta_{14} \sin^2 \theta_{24} \times \sin^2 \frac{\Delta m_{41}^2 L}{E} \quad 4|U_{e4}|^2 |U_{\mu 4}|^2$$

- ▶  $\bar{\nu}_e$  CC disappearance (Reactor experiments):  $\theta_{14}$

$$P_{ee} \approx 1 - 2 \sin^2 2\theta_{14} \times \sin^2 \frac{\Delta m_{41}^2 L}{E} \quad |U_{e4}|^2$$

- ▶  $\nu_\mu, \bar{\nu}_\mu$  CC disappearance (mBooNE, MINOS(+), NOvA, T2K, IceCube, etc.):  $\theta_{24}$

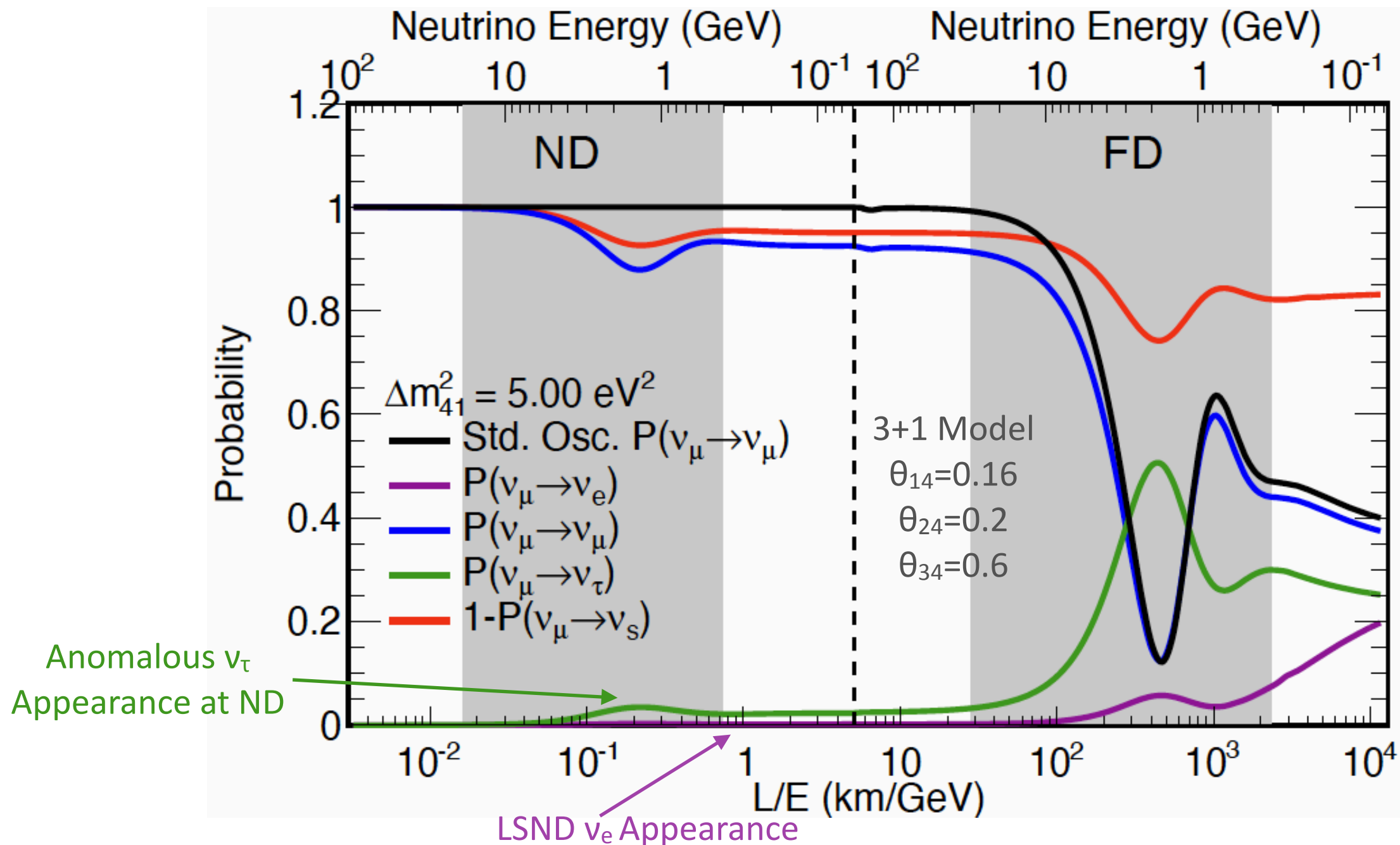
$$P_{\mu\mu} \approx 1 - 2 \sin^2 2\theta_{24} \times \sin^2 \frac{\Delta m_{41}^2 L}{E} \quad |U_{\mu 4}|^2$$

- ▶ NC disappearance (MINOS(+), NOvA, T2K, etc.):  $\theta_{24}, \theta_{34}$

$$1 - P_{\mu s} \approx 1 - \underbrace{\sin^2 2\theta_{24} \sin^2 \frac{\Delta m_{41}^2 L}{E}}_{\text{SBL}} - \underbrace{\sin^2 \theta_{34} \sin^2 2\theta_{23} \sin^2 \frac{\Delta m_{31}^2 L}{E}}_{\text{LBL}} \quad |U_{\mu 4}|^2, |U_{\tau 4}|^2$$

- ▶ Thanks to the intense LBNF neutrino beam and exquisite detector spatial resolution of its detectors, DUNE is sensitive to all of these channels in a single experiment!
  - ◉ Also sensitive to atmospheric measurements due to large FD, 1500 m deep at SURF

# Looking for Light Sterile Neutrinos



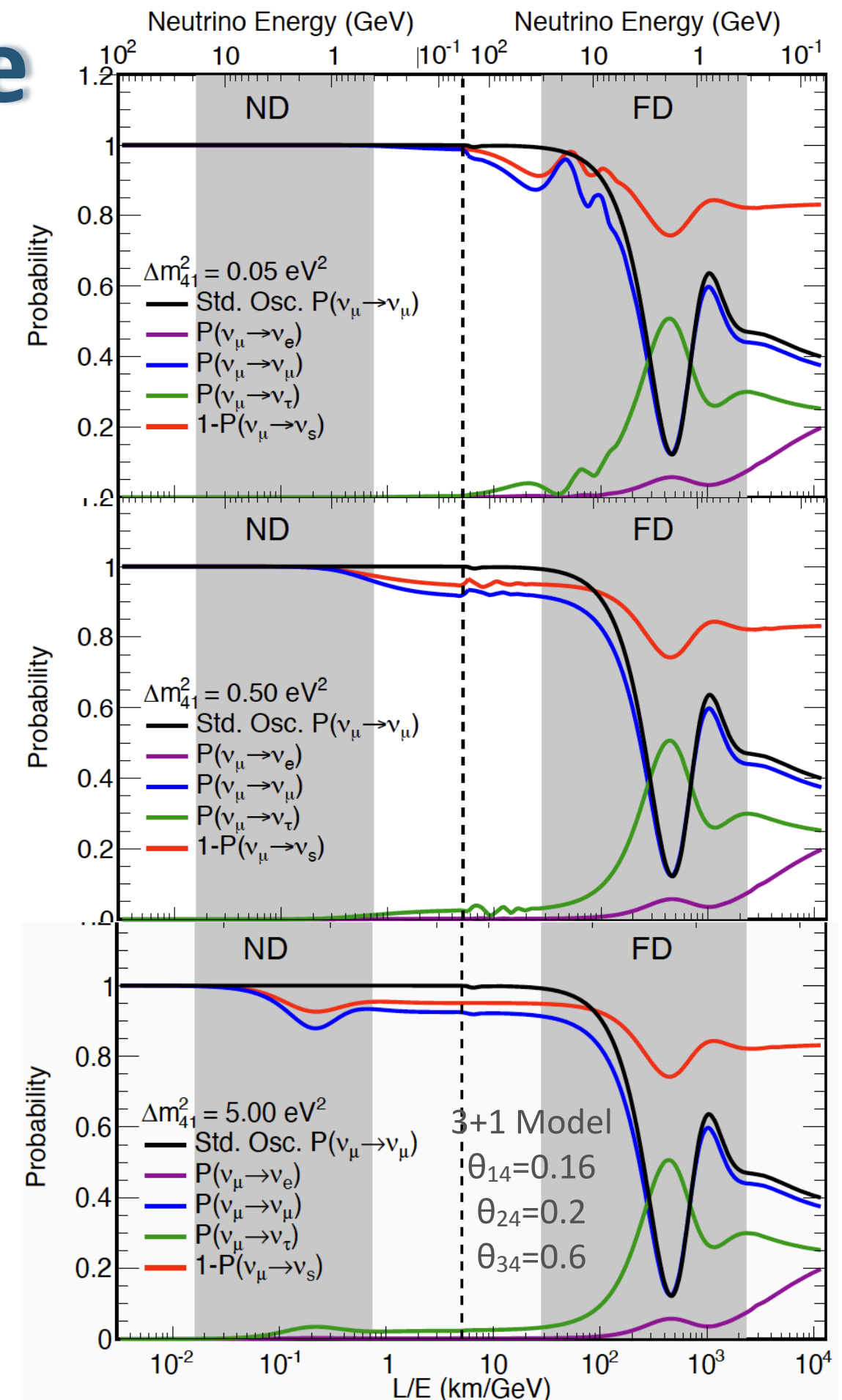
## ► Active-sterile mixing would distort standard oscillation probabilities

- DUNE will be sensitive to this effect through both the Near and Far detectors
- Wide-band LBNF beam enables probes over large regions of parameter space
- Plot shows distortion of standard oscillation probabilities for  $L/E$  or  $\nu$  energies at ND and FD



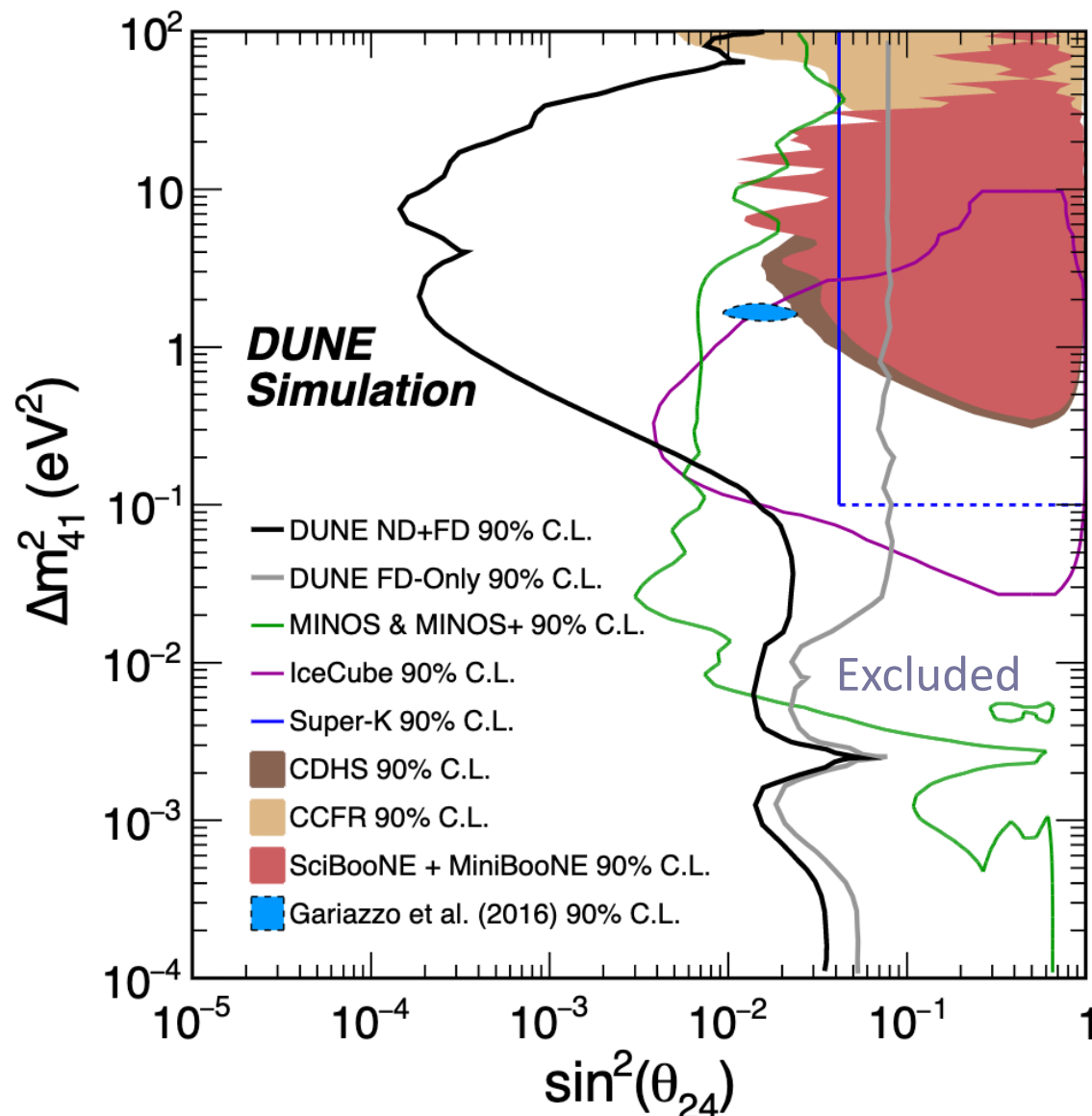
# Looking for Light Sterile Neutrinos

- ▶ Distortions of standard oscillation probabilities change for different values of  $\Delta m_{41}^2$ 
  - **Small  $\Delta m_{41}^2$ :** slow oscillations visible at FD only (FD-dominated region)
  - **Intermediate  $\Delta m_{41}^2$ :** rapid oscillations average out at FD but still not visible at ND (Counting experiment)
  - **Large  $\Delta m_{41}^2$ :** oscillations average out at FD and distortions are visible at the ND (ND-dominated region)

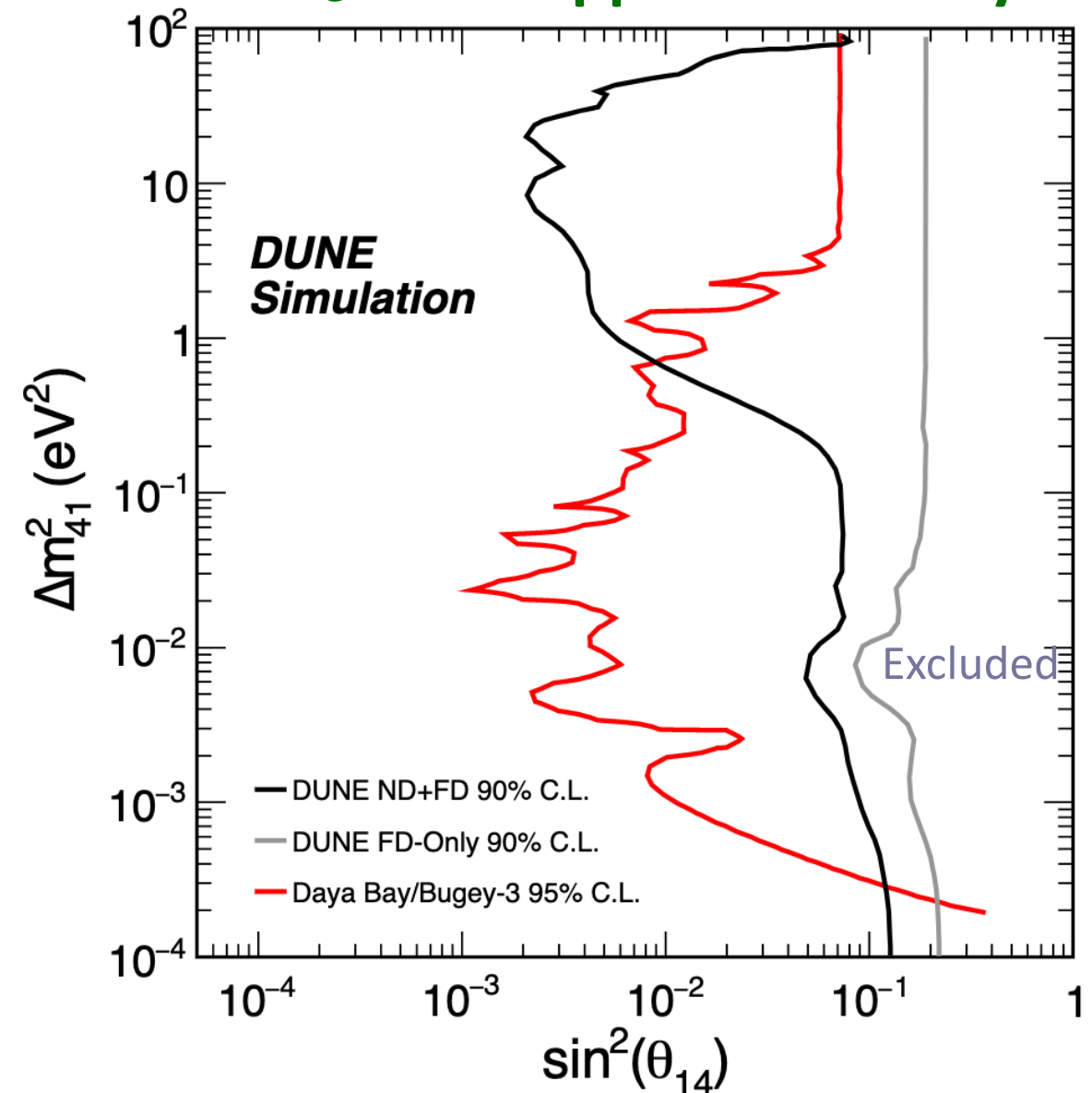


# DUNE Sensitivities to Sterile Mixing

$\nu_\mu$  CC + NC Disappearance



$\nu_e$  CC Disappearance Only

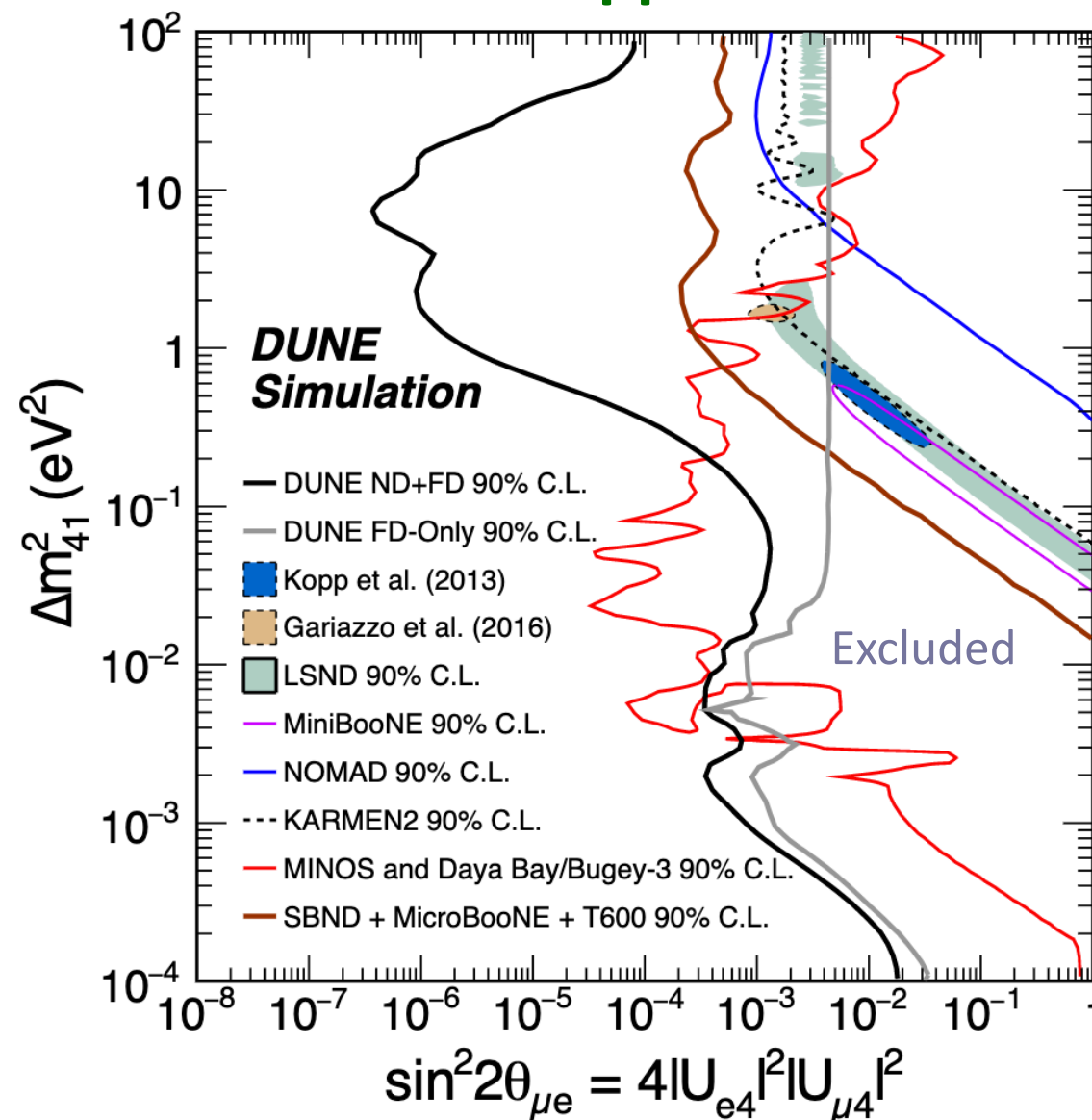


- Assuming 300 kton.MW.year exposure (Phase I ND) for 3+1 model with simultaneous fit to oscillations at ND and FD
  - GLOBES sensitivities include normalization-only systematics, so the two DUNE lines represent best (black) and worst (gray) scenarios
  - On its own, DUNE can potentially probe the sterile mixing parameter space at same level or better than present and future experiments, depending on  $\Delta m^2_{41}$



# DUNE Sensitivities to Sterile Mixing

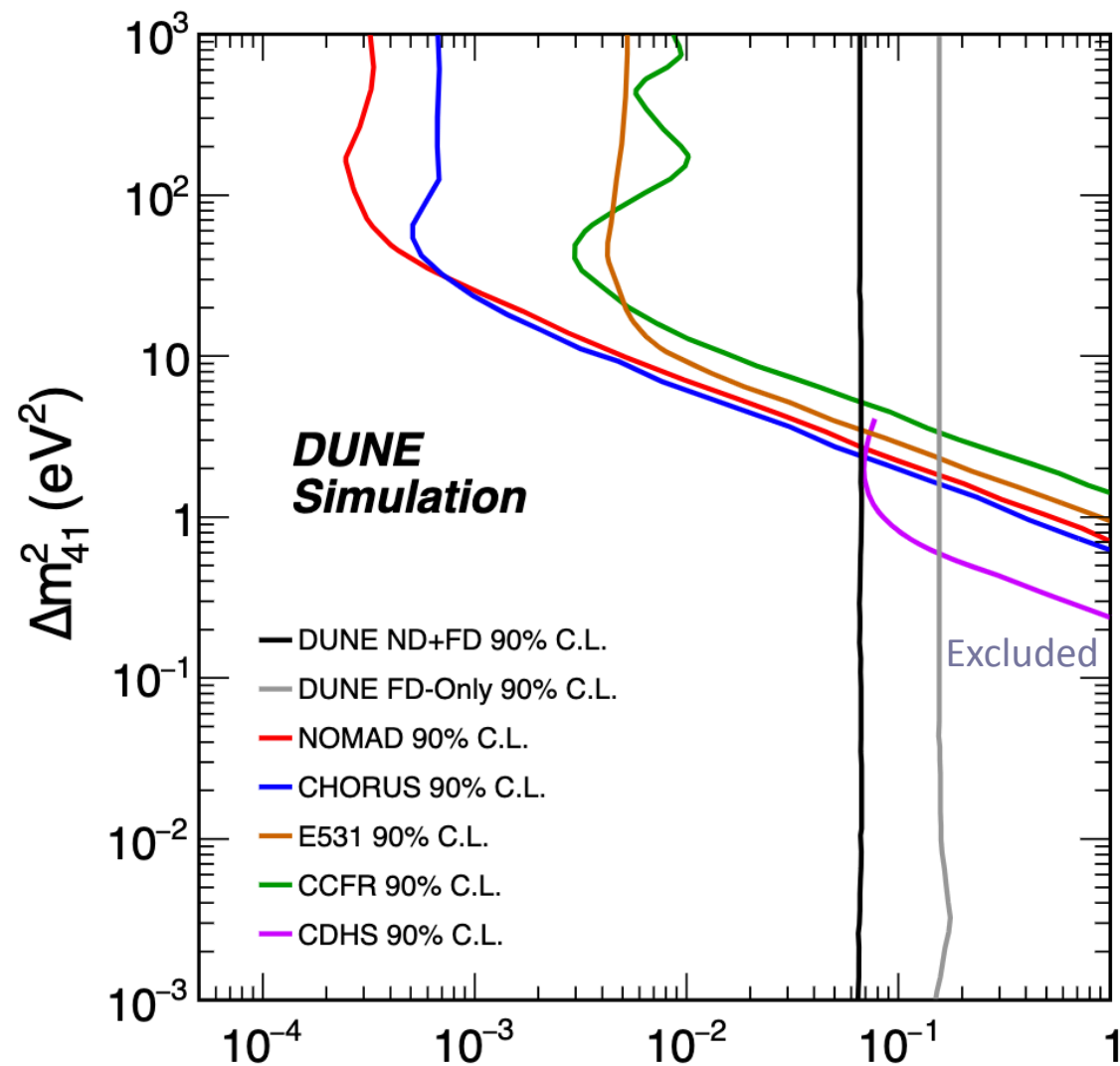
$\nu_e$  CC Appearance



- ▶ Assuming 300 kton.MW.year exposure (Phase I ND) for 3+1 model with simultaneous fit to oscillations at ND and FD
  - ⦿ PRISM and Phase II ND will help control systematics in ND-dominated region,  $\Delta m_{41}^2 \gtrsim 0.5 \text{ eV}^2$
  - ⦿ Strong complementarity with SBN program, thanks to ND measurement, while extending probes to lower values of  $\Delta m_{41}^2$  via the FD measurement

# DUNE Sensitivities to Sterile Mixing

$\nu_\mu$  CC + NC Disappearance



3+1 limits at  $\Delta m^2_{41} = 0.5 \text{ eV}^2$

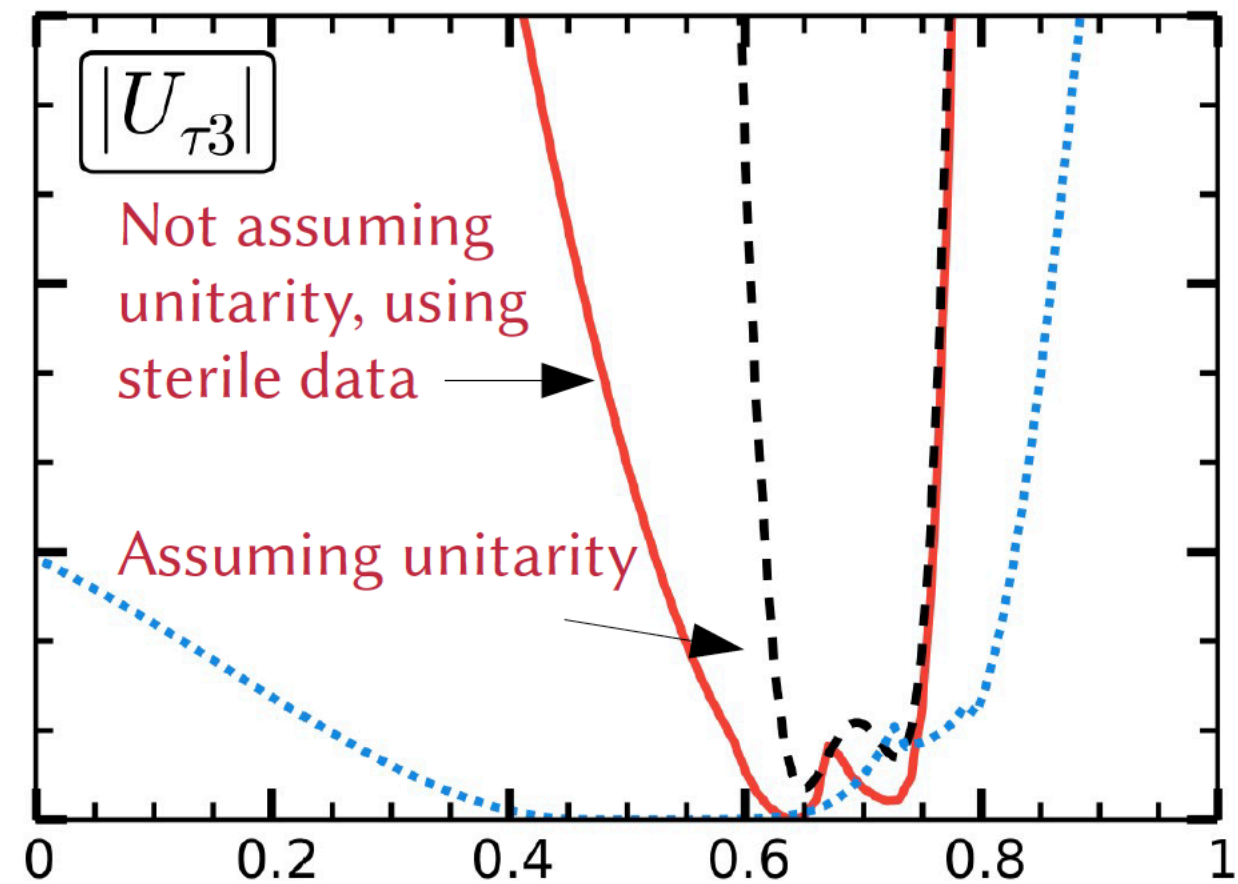
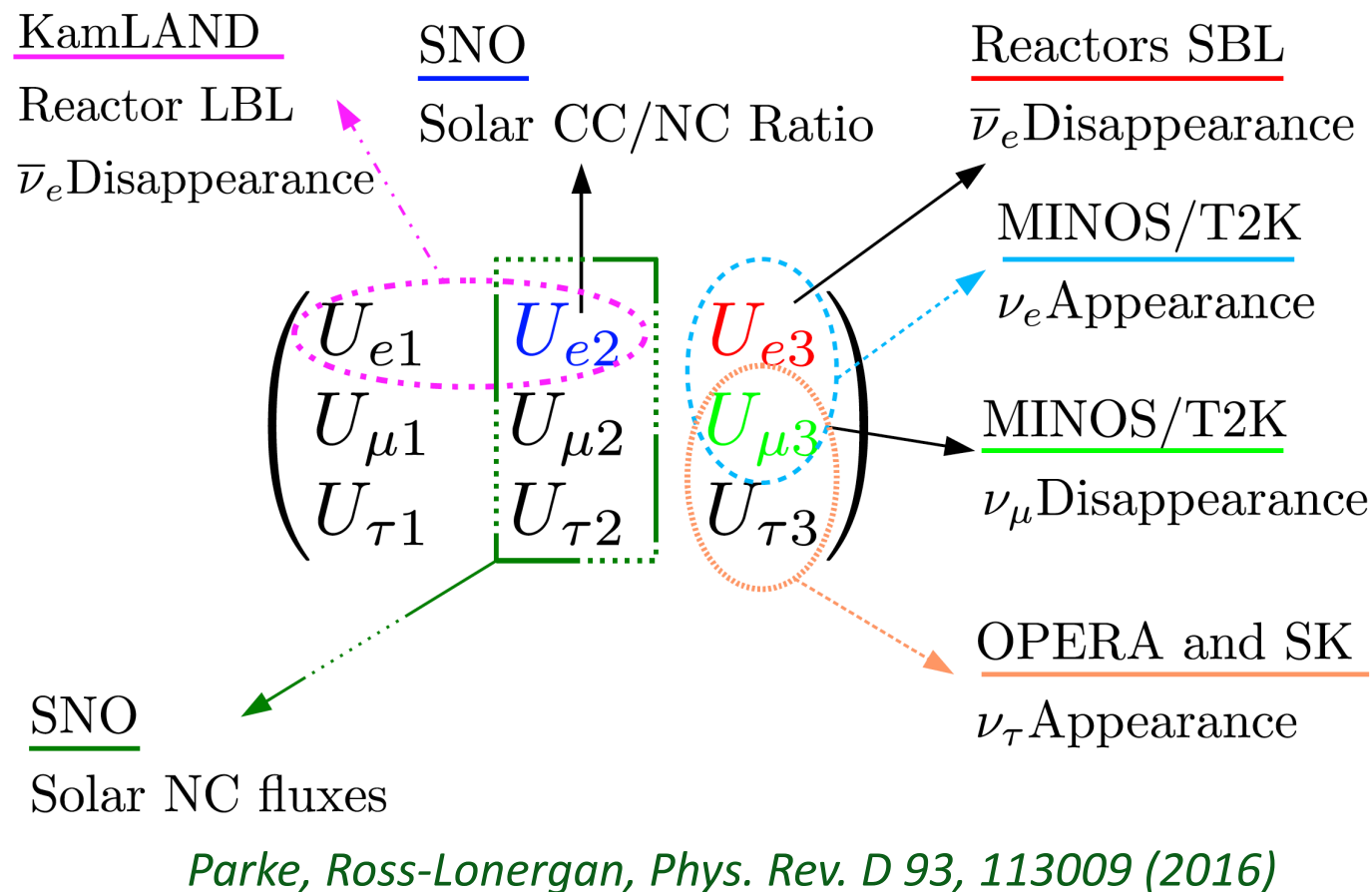
	$\theta_{24}$	$\theta_{34}$	$ U_{\mu 4} ^2$	$ U_{\tau 4} ^2$
DUNE Best-Case	1.8°	15.0°	0.001	0.067
DUNE Worst-Case	15.1°	25.5°	0.068	0.186
NOvA	20.8°	31.2°	0.126	0.268
MINOS/MINOS+	4.4°	23.6°	0.006	0.16
Super-Kamiokande	11.7°	25.1°	0.041	0.18
IceCube	4.1°	-	0.005	-
IceCube-DeepCore	19.4°	22.8°	0.11	0.15

$$\sin^2(2\theta_{\mu\tau}) \equiv 4|U_{\mu 4}|^2|U_{\tau 4}|^2 = \cos^4 \theta_{14} \sin^2 2\theta_{24} \sin^2 \theta_{34}$$

- Assuming 300 kton.MW.year exposure (Phase I ND) for 3+1 model with simultaneous fit to oscillations at ND and FD
  - DUNE can extend probes of  $\nu_\tau$ -sterile mixing to lower values of  $\Delta m^2_{41}$  and improve the limits on  $\theta_{34}$ , the least constrained sterile mixing angle
  - Can be further improved by combining beam + atmospheric measurements



# DUNE and $\nu_\tau$ Probes

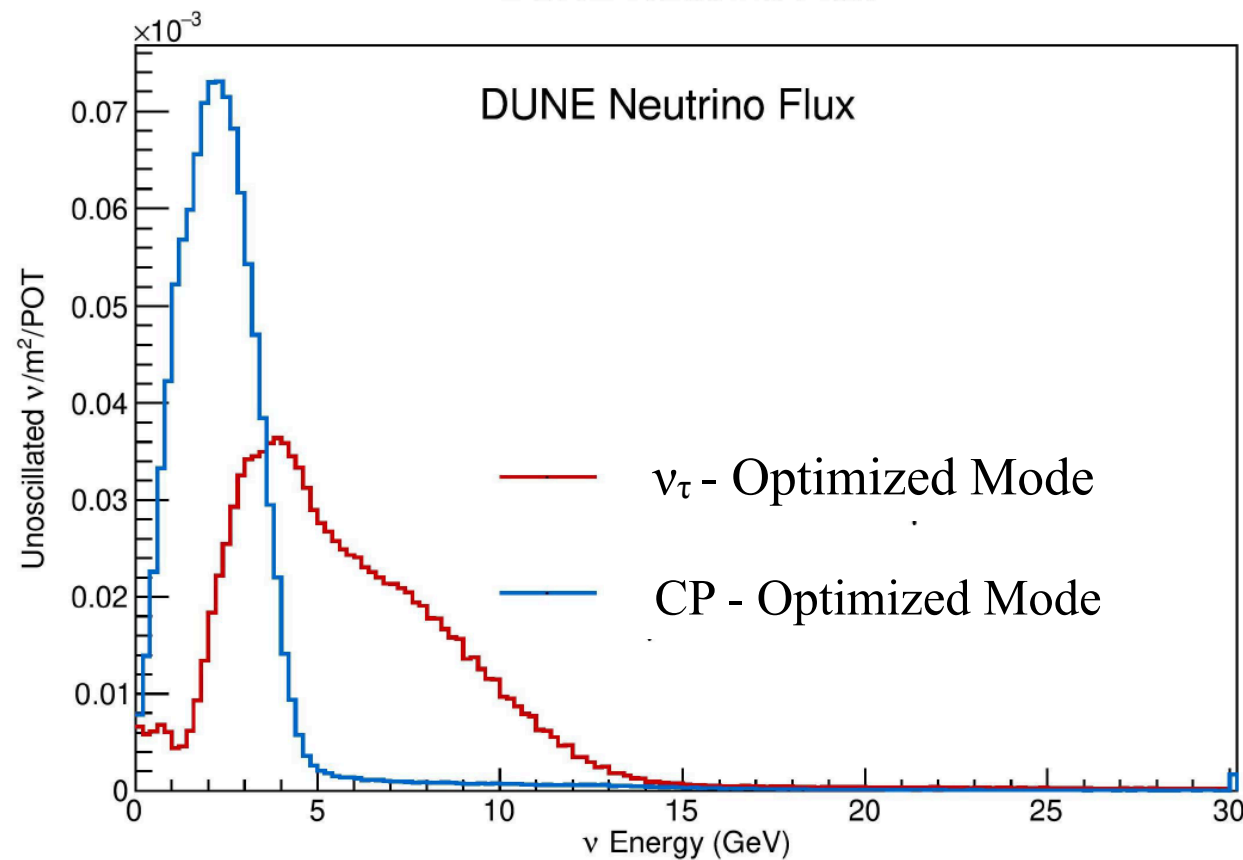


- ▶ Most of our knowledge of  $\nu_\tau$  sector results from assuming unitarity of PMNS matrix and lepton universality for cross sections
  - $\nu_\tau$  CC production threshold,  $E_\nu > 3.5$  GeV, is very challenging for signal accumulation
- ▶ DUNE is in a unique position to probe the  $\nu_\tau$  sector using:
  - High-energy tail of CP-optimized beam during Phases I and II (130  $\nu_\tau$  CC, 30  $\bar{\nu}_\tau$  CC/year in FD)
  - **Dedicated  $\nu_\tau$  - optimized beam run during Phase II (800  $\nu_\tau$  CC/year in FD)**
  - Also atmospheric neutrino measurements with FD, like SuperK and IceCube

# Anomalous $\nu_\tau$ Appearance at DUNE ND

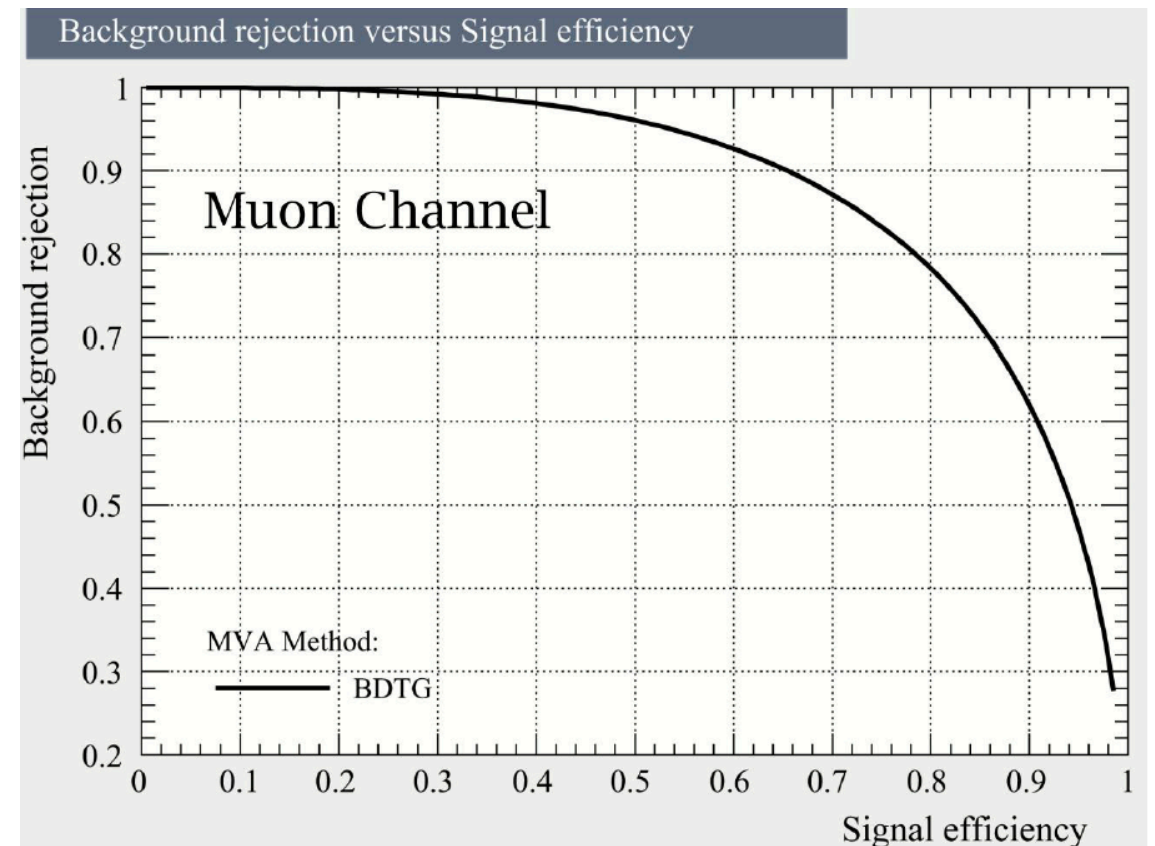
$\tau$  decay modes

Decay mode	Branching ratio (%)
$\pi^- \pi^0 \nu_\tau$	25.49
$e^- \bar{\nu}_e \nu_\tau$	17.82
$\mu^- \bar{\nu}_\mu \nu_\tau$	17.39
$\pi^- \nu_\tau$	10.82
$\pi^- 2\pi^0 \nu_\tau$	9.26



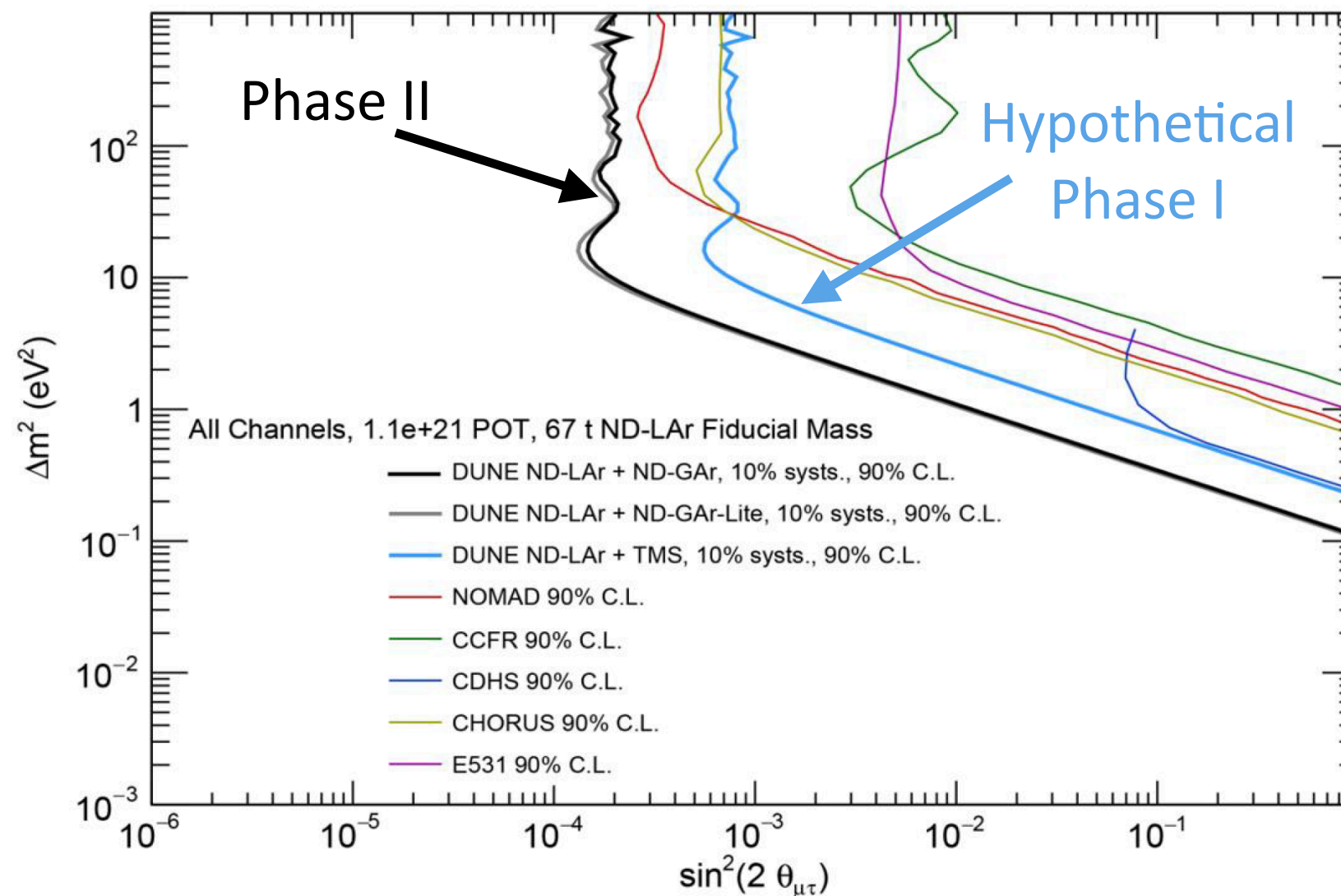
$$P_{\mu\tau} \approx \sin^2 2\theta_{\mu\tau} \sin^2 \frac{\Delta m_{41}^2 L}{E}$$

$$\sin^2 2\theta_{\mu\tau} \equiv 4|U_{\mu 4}|^2 |U_{\tau 4}|^2 = \cos^4 \theta_{14} \sin^2 2\theta_{24} \sin^2 \theta_{34}$$



- Higher energy of  $\nu_\tau$  - optimized beam, BDT based on kinematic variables, and measurements of a wide range of muon momenta from curvature in DUNE ND Phase II's ND-GAr, enable good statistics in selection of  $\nu_\tau$  CC sample from  $\nu_\mu$  CC backgrounds.

# Anomalous $\nu_\tau$ Appearance at DUNE ND



- ▶ Sensitivities for 1 year of running with  $\nu_\tau$  - optimized beam (not expected during Phase I)
  - ⦿ Including an overall 10% syst. uncertainty; smearing according to each detector's resolution
- ▶ Potential leading sensitivity in difficult to probe parameter space during DUNE Phase II
  - ⦿ Further improvements possible by using the SAND detector in the ND complex
  - ⦿ Complementary to  $\nu_\tau$ -sterile mixing probes with FASERnu and FPF@LHC [Bai et al., arXiv:2002.0301](#)

*See H. Razafinime's talk, Saturday NF EC parallel*



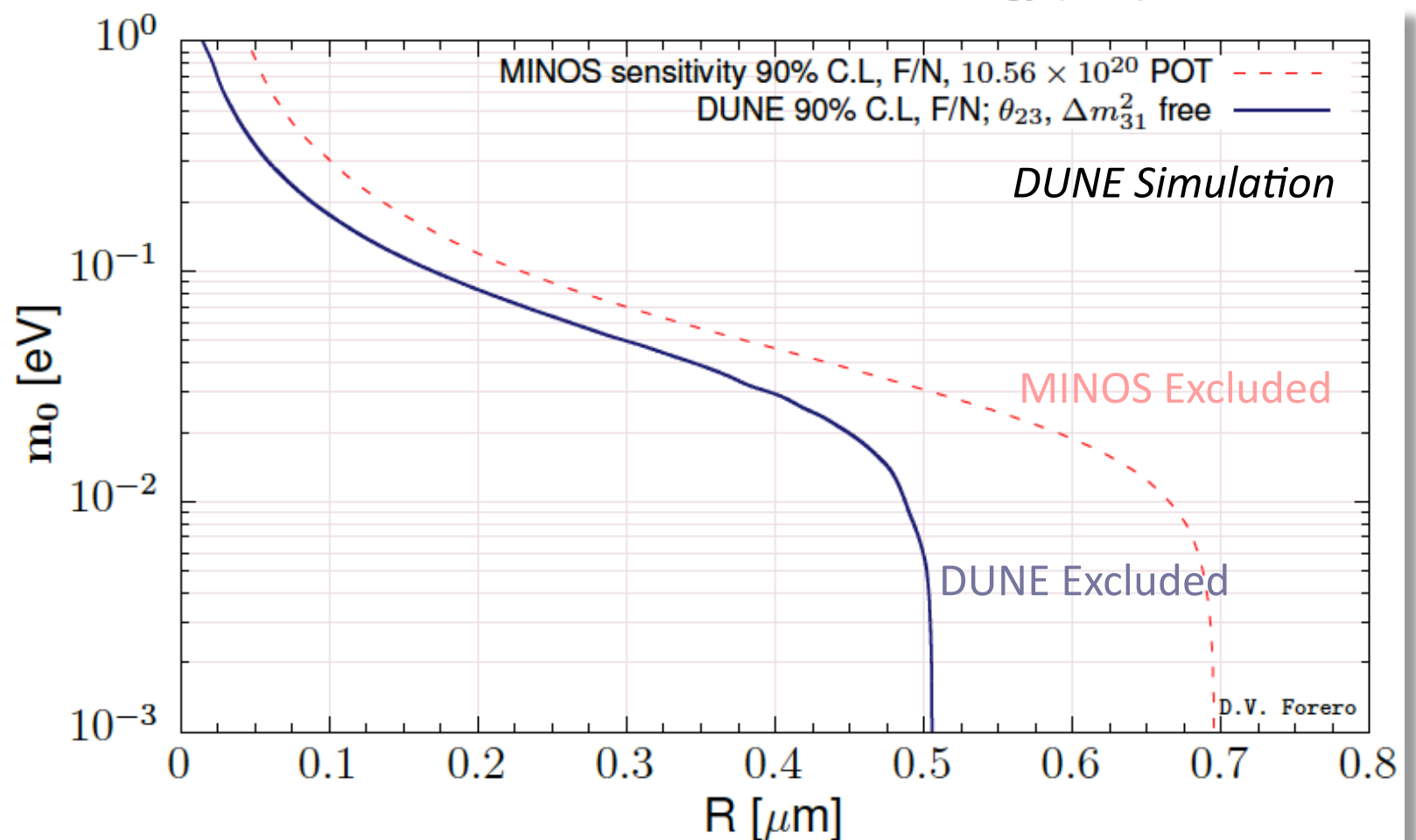
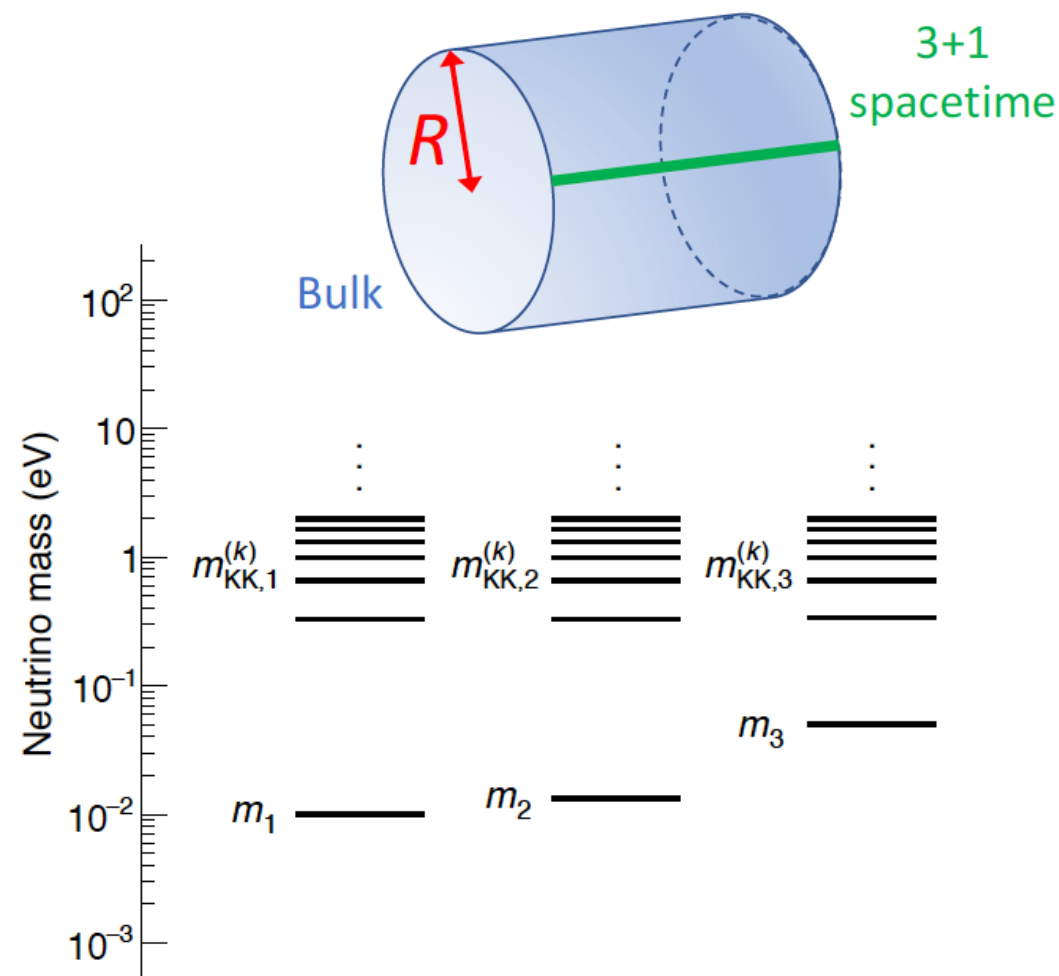
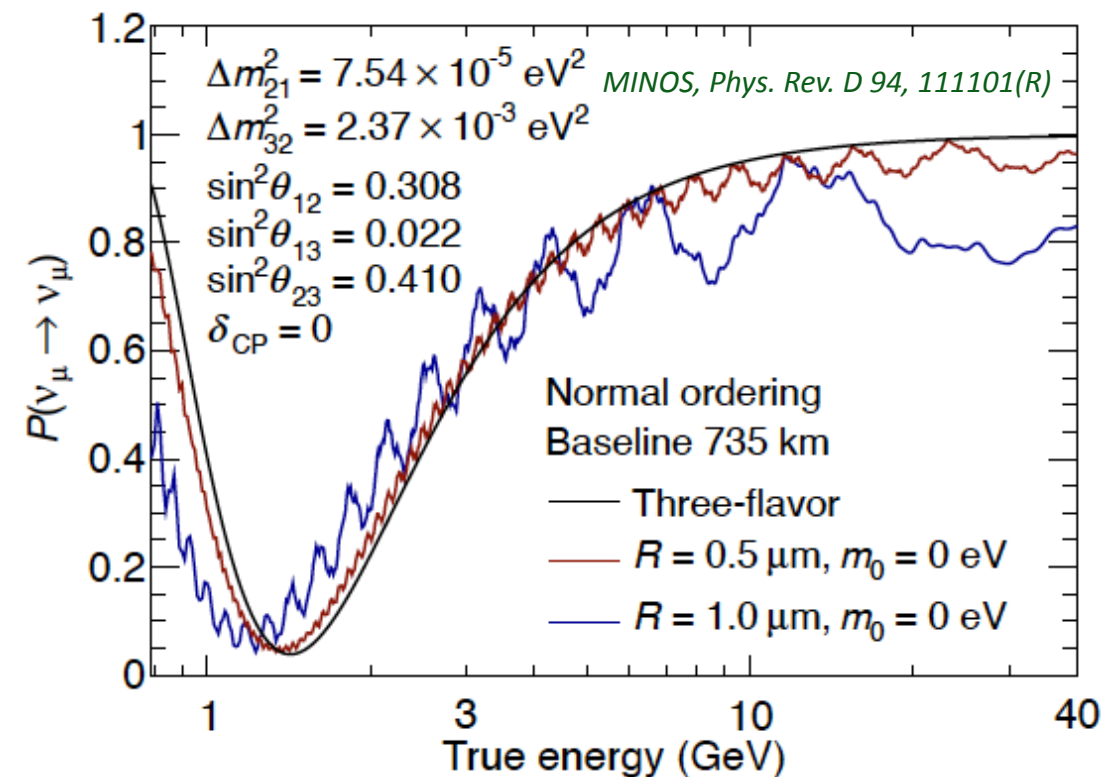
# DUNE and the Experimental v Anomalies

Source	Flavor Conversion: 3+N Oscillations	Flavor Conversion: Anomalous Matter Effects	Flavor Conversion: Lepton Flavor Violation	Dark Sector: Decays in Flight	Dark Sector: Neutrino-induced Up-scattering	Dark Sector: Dark-particle-induced Up-scattering
Reactor	DANSS Upgrade, JUNO-TAO, NEOS-II, Neutrino-4 Upgrade, PROSPECT-II					
Radioactive Source	BEST-2, IsoDAR, THEIA, Jinping					
Atmospheric	IceCube Upgrade, KM3NET, ORCA and ARCA, DUNE, Hyper-Kamiokande, THEIA				IceCube Upgrade, KM3NET, ORCA and ARCA, DUNE, Hyper-Kamiokande, THEIA	
Pion/Kaon Decay-At-Rest	JSNS <sup>2</sup> , COHERENT, Coherent-Captain-Mills, KPIPE		JSNS <sup>2</sup> , COHERENT, Coherent-Captain-Mills, KPIPE, PIP2-BD			COHERENT, Coherent-Captain-Mills, KPIPE, PIP2-BD, SBN-BD
Beam Short Baseline	SBN			SBN, FASER $\nu$ , SND@LHC, FLArE		
Beam Long Baseline	DUNE, Hyper-Kamiokande, ESSnuSB			DUNE, Hyper-Kamiokande, ESSnuSB		
Muon Decay-In-Flight	nuSTORM				nuSTORM	
Beta Decay and Electron Capture	KATRIN/TRISTAN, Project-8, HUNTER, BeEST, DUNE ( <sup>39</sup> Ar), PTOLEMY, $2\nu\beta\beta$			<a href="#">NF02 White Paper, hep-ex/2203.07323</a> <b>NF02 Session on Expt. v Anomalies, Thurs, 10 am</b>		

- ▶ Pure 3+N sterile mixing disfavored as explanation for global neutrino data due to tension between appearance signals and null disappearance results
- ▶ With its strong multi-channel sensitivity and wide-band beam, DUNE will test more exotic scenarios

# Large Extra-Dimensions

- Large Extra-Dimensions would cause distortions of 3-flavor oscillations from mixing of neutrinos with Kaluza-Klein (KK) modes
  - For LED model, *Davoudiasl et al., PRD 65, 105015 (2002)*, assuming one LED in the bulk, KK modes in 3+1 spacetime brane behave like sterile neutrinos
  - Showing DUNE sensitivity for 300 kton.MW.year (Phase I) compared to MINOS published results

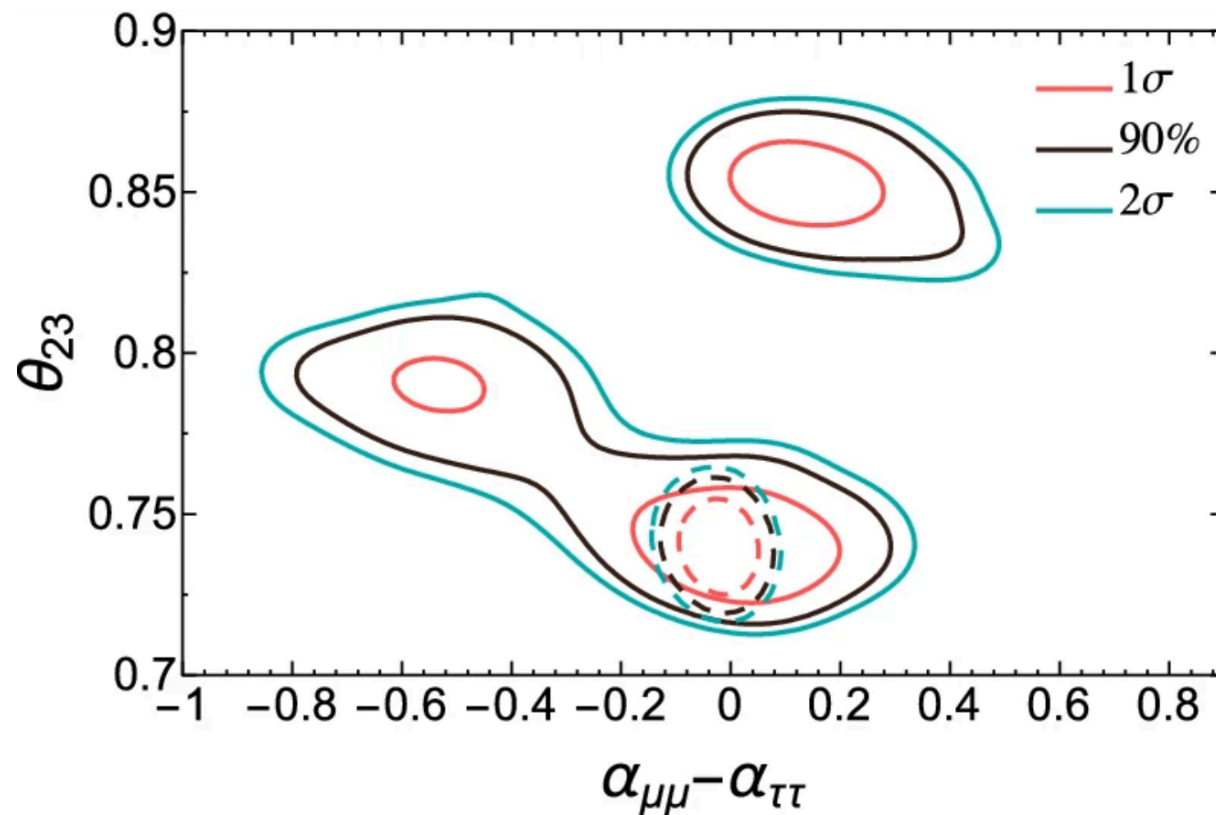


See D. V. Forero's talk, Saturday NF EC parallel

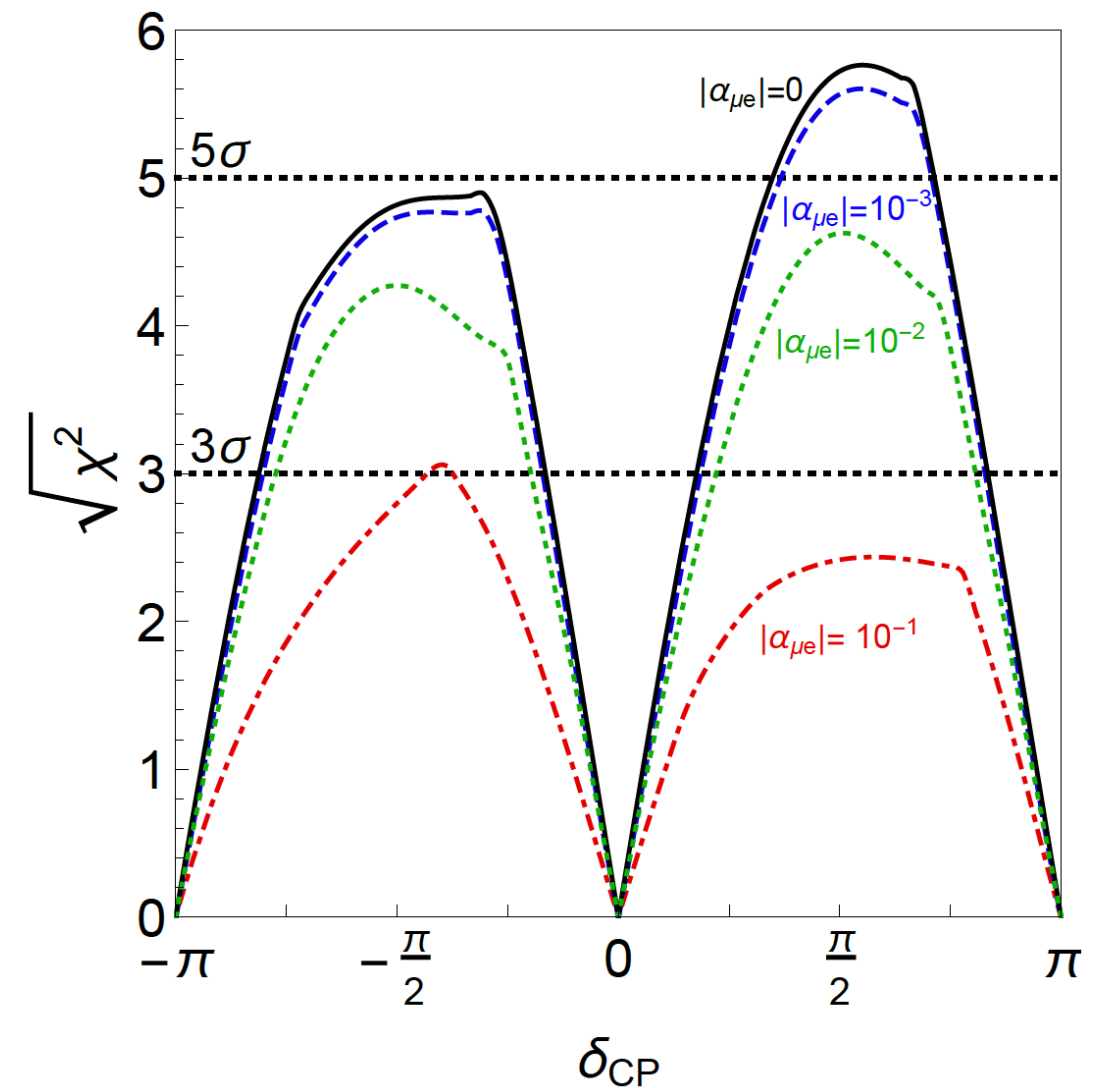
# Non-Unitary Mixing

- If new heavy states mix with active neutrinos (*e.g.* if neutrinos acquire mass through a type I seesaw mechanism), the mixing matrix need not be unitary

$$N = \begin{pmatrix} 1 - \alpha_{ee} & 0 & 0 \\ \alpha_{\mu e} & 1 - \alpha_{\mu\mu} & 0 \\ \alpha_{\tau e} & \alpha_{\tau\mu} & 1 - \alpha_{\tau\tau} \end{pmatrix} U^{3 \times 3}$$



- Allowed regions at the 1 $\sigma$ , 90%, and 2 $\sigma$  CL for non-unitary mixing parameters for DUNE-only (solid), and DUNE + existing constraints (dashed)
  - ◉ Assuming 300 kton.MW.year (Phase I ND)

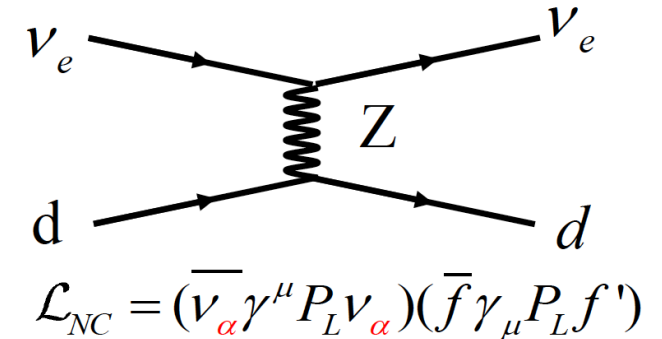
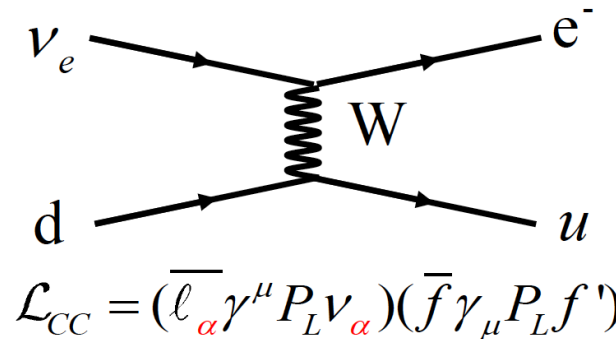


- Potential impact of non-unitarity on the DUNE CP violation discovery potential



# Non-Standard Neutrino Interactions (NSI)

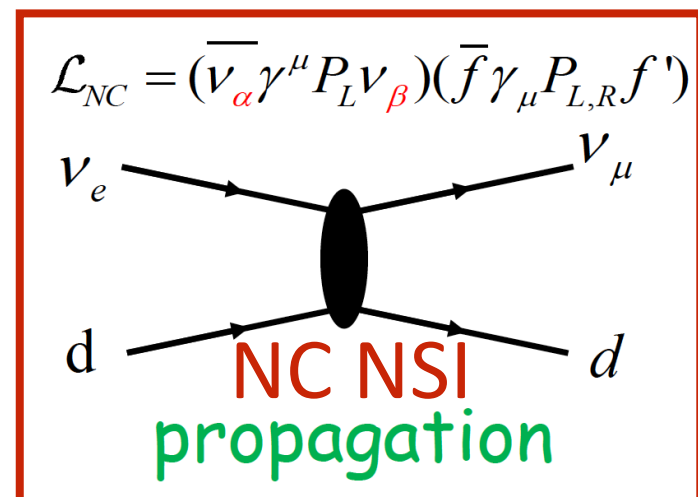
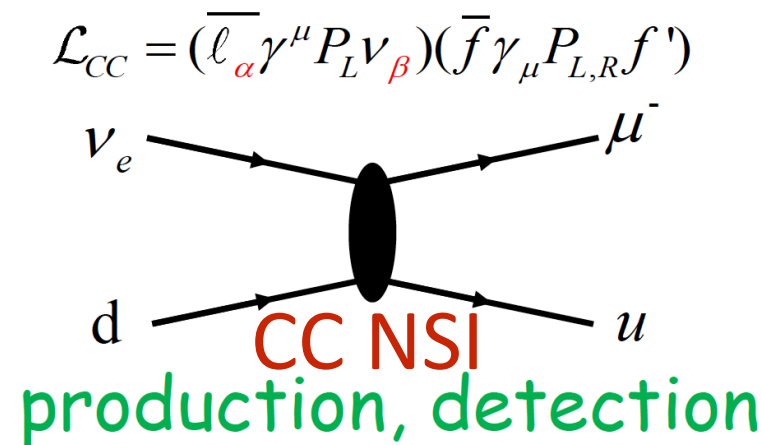
- In the Standard Model,



- ▶ New neutral-current-like interactions during neutrino propagation between the Near and Far detectors can be described as new contributions to the neutrino matter effect (MSW)

- These contributions are encoded by the new coefficients  $\epsilon_{ij}$

- With new physics, we could have

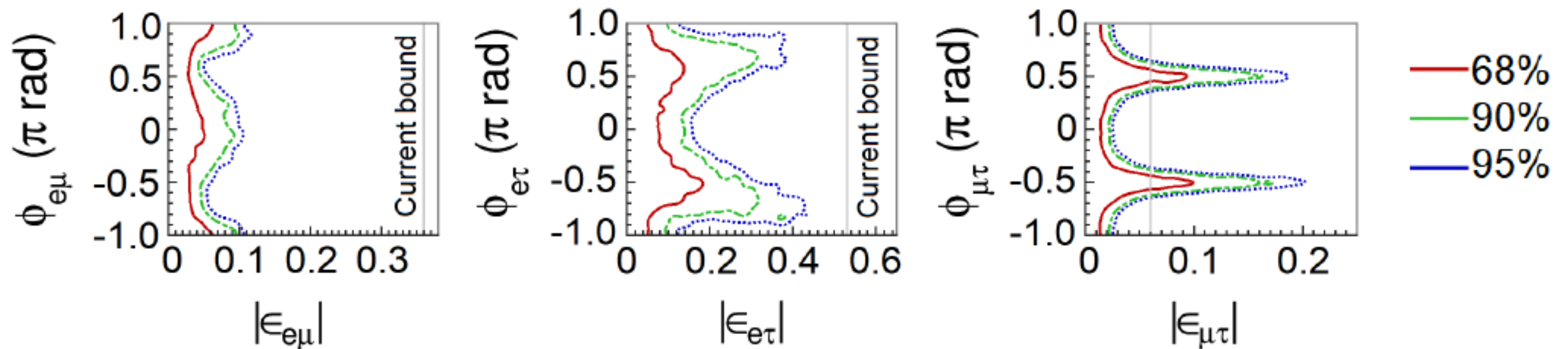


$$H = U \begin{pmatrix} 0 & & \\ & \Delta m_{21}^2/2E & \\ & & \Delta m_{31}^2/2E \end{pmatrix} U^{\dagger} + \tilde{V}_{\text{MSW}}$$

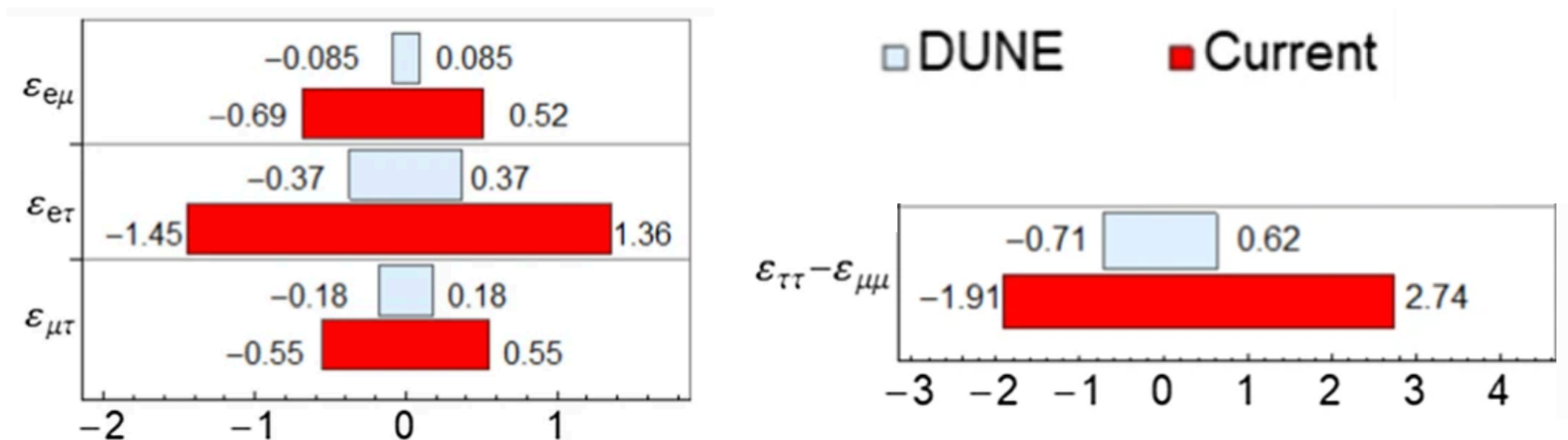
$$\tilde{V}_{\text{MSW}} = \sqrt{2} G_F N_e \begin{pmatrix} 1 + \epsilon_{ee}^m & \epsilon_{e\mu}^m & \epsilon_{e\tau}^m \\ \epsilon_{e\mu}^{m*} & \epsilon_{\mu\mu}^m & \epsilon_{\mu\tau}^m \\ \epsilon_{e\tau}^{m*} & \epsilon_{\mu\tau}^{m*} & \epsilon_{\tau\tau}^m \end{pmatrix}$$

# Non-Standard Neutrino Interactions (NSI)

- DUNE can improve current constraints on  $|\epsilon_{e\mu}|$  and  $|\epsilon_{e\tau}|$  by a factor of  $\sim 2$



- Allowed regions for an exposure of 300kt.MW.year. Current bounds from [Gonzalez-Garcia, Maltoni, arXiv:1307.3092](#)



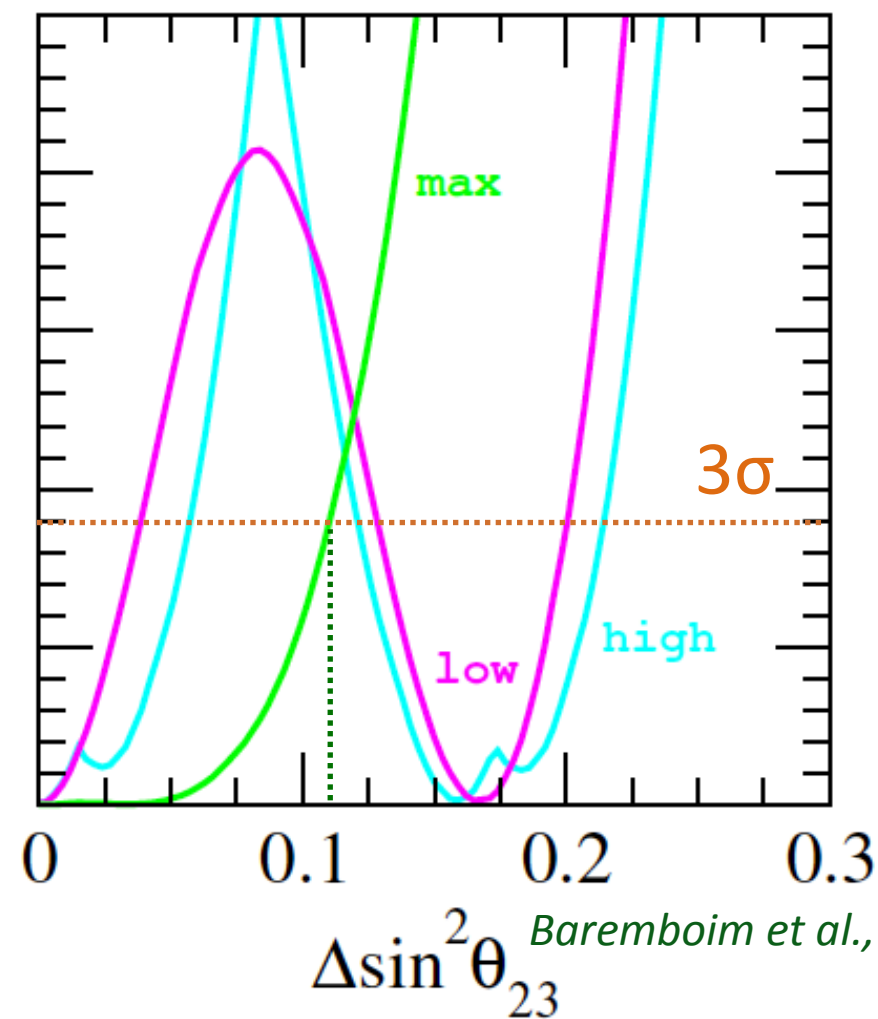
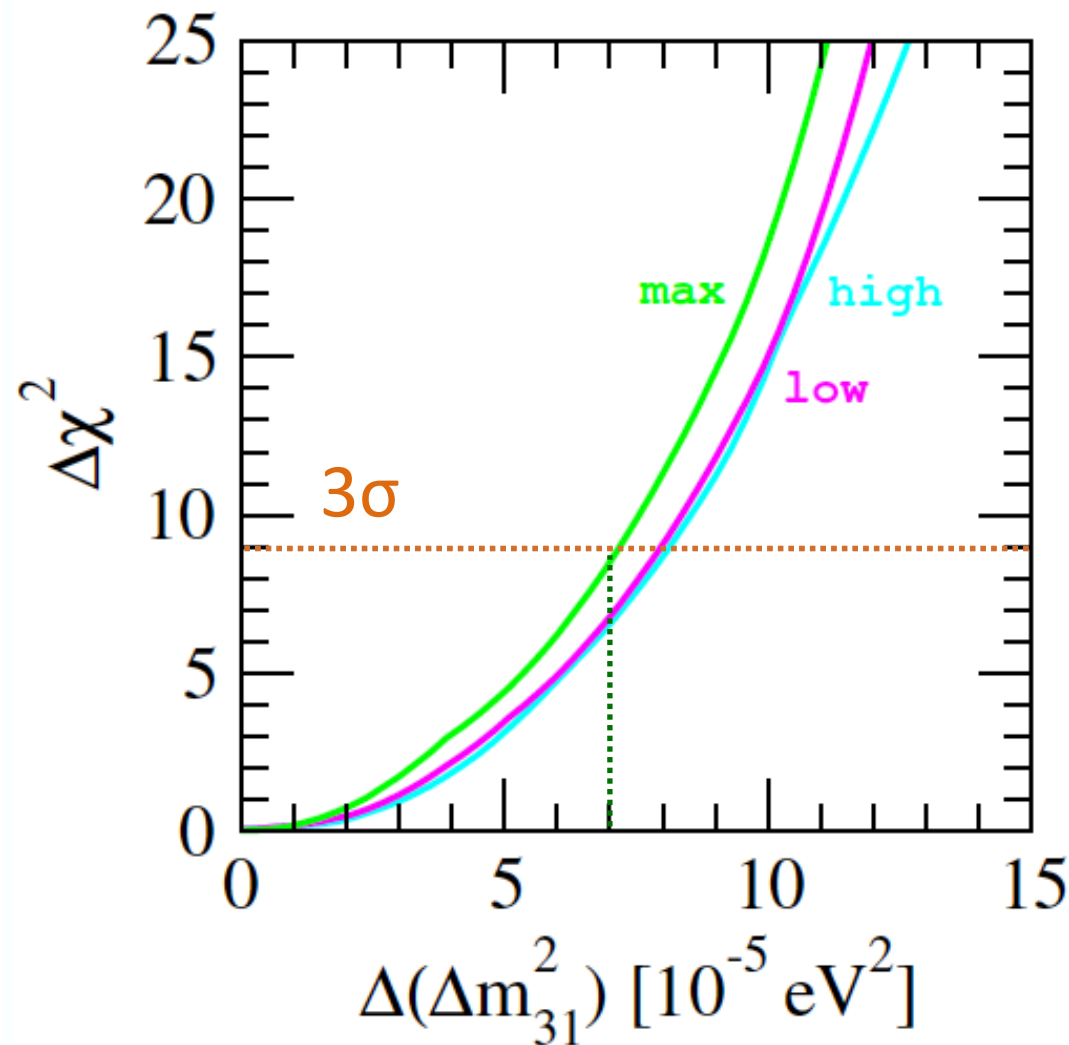
- 90% C.L. 1-dim. DUNE constraints compared with bounds from [Gonzalez-Garcia, Maltoni, arXiv:1307.3092](#)

# Violation of CPT symmetry

- DUNE can search for CPT violation by comparing  $\nu_\mu$  and  $\bar{\nu}_\mu$  disappearance between ND and FD

$$P(\nu_\mu \rightarrow \nu_e) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \Rightarrow \text{CP violation}$$

$$P(\nu_\mu \rightarrow \nu_\mu) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu) \Rightarrow \text{CPT violation}$$



*Bareimboim et al., Phys. Lett. B 780, 631 (2018)*

- With 300 kton.MW.year running, DUNE can improve limits on  $\Delta(\Delta m_{31}^2)$  by over a factor of 5 and, depending on the octant of  $\theta_{23}$  (low, high, maximal), the limits on  $\Delta(\sin^2 \theta_{23})$

## Present limits

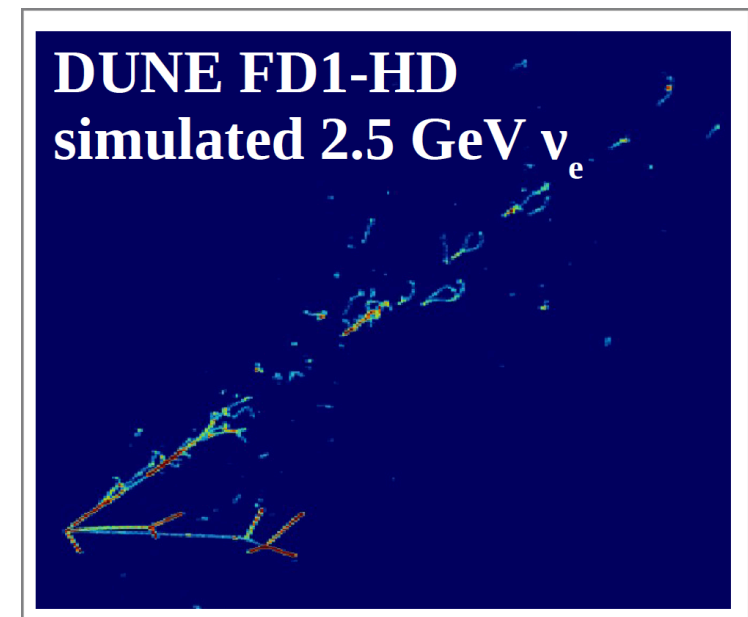
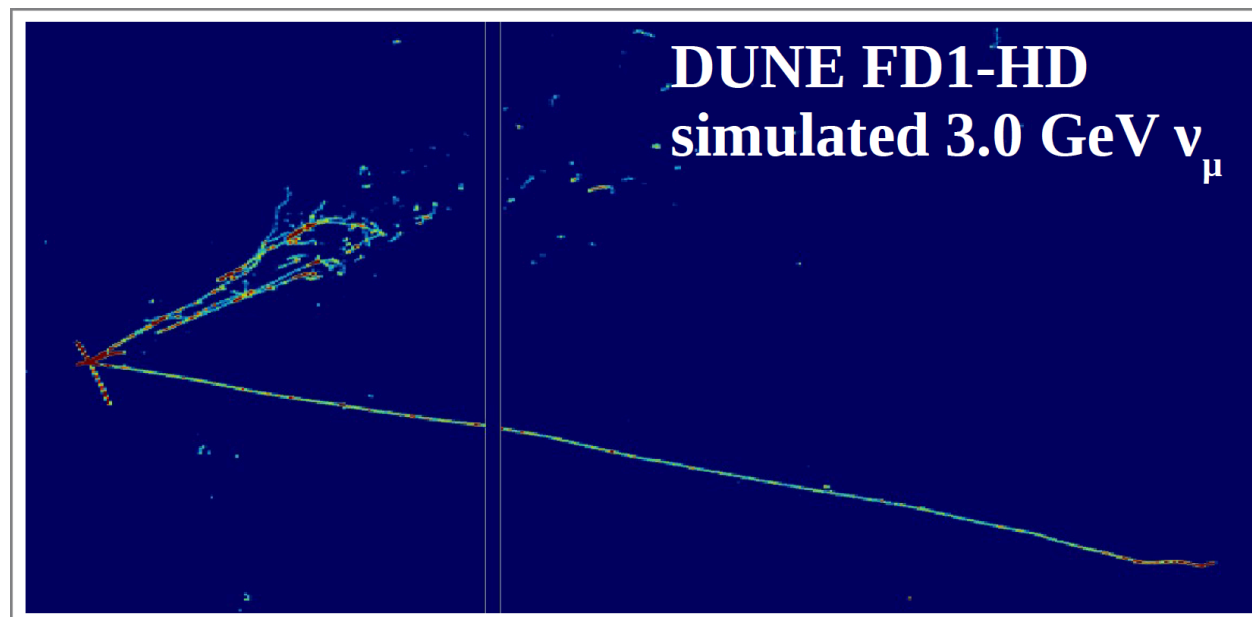
$$\Delta(\Delta m_{31}^2) \equiv |\Delta m_{31}^2 - \Delta \bar{m}_{31}^2| < 3.7 \times 10^{-4} \text{ eV}^2$$

$$\Delta(\sin^2 \theta_{23}) \equiv |\sin^2 \theta_{23} - \sin^2 \bar{\theta}_{23}| < 0.32$$



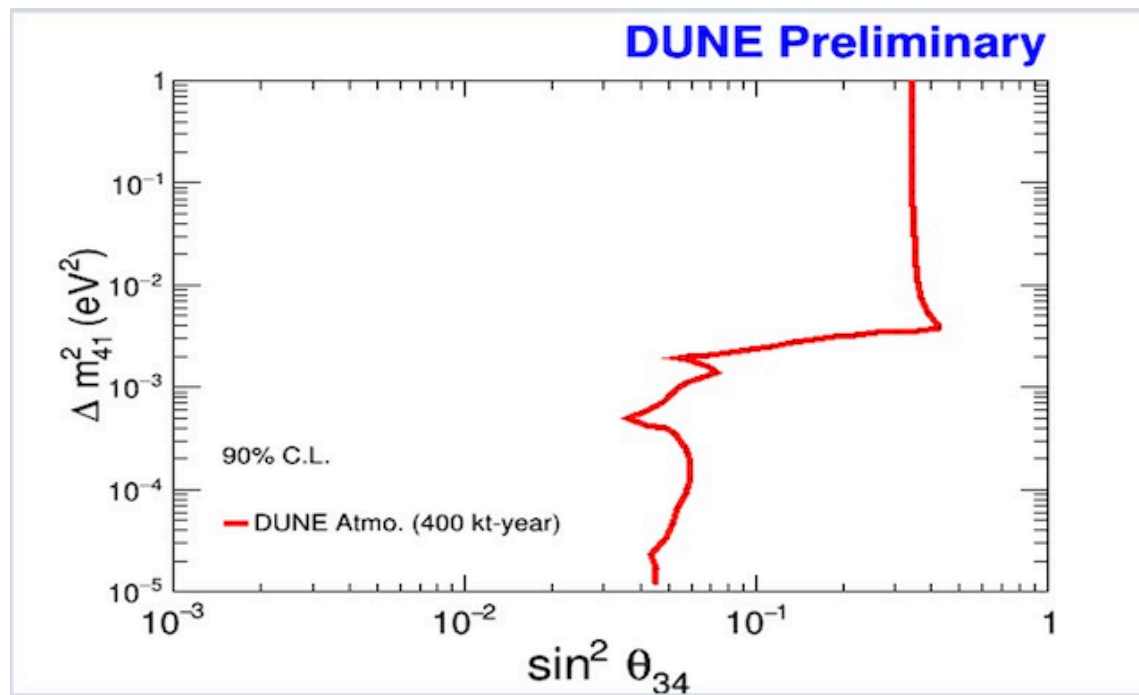
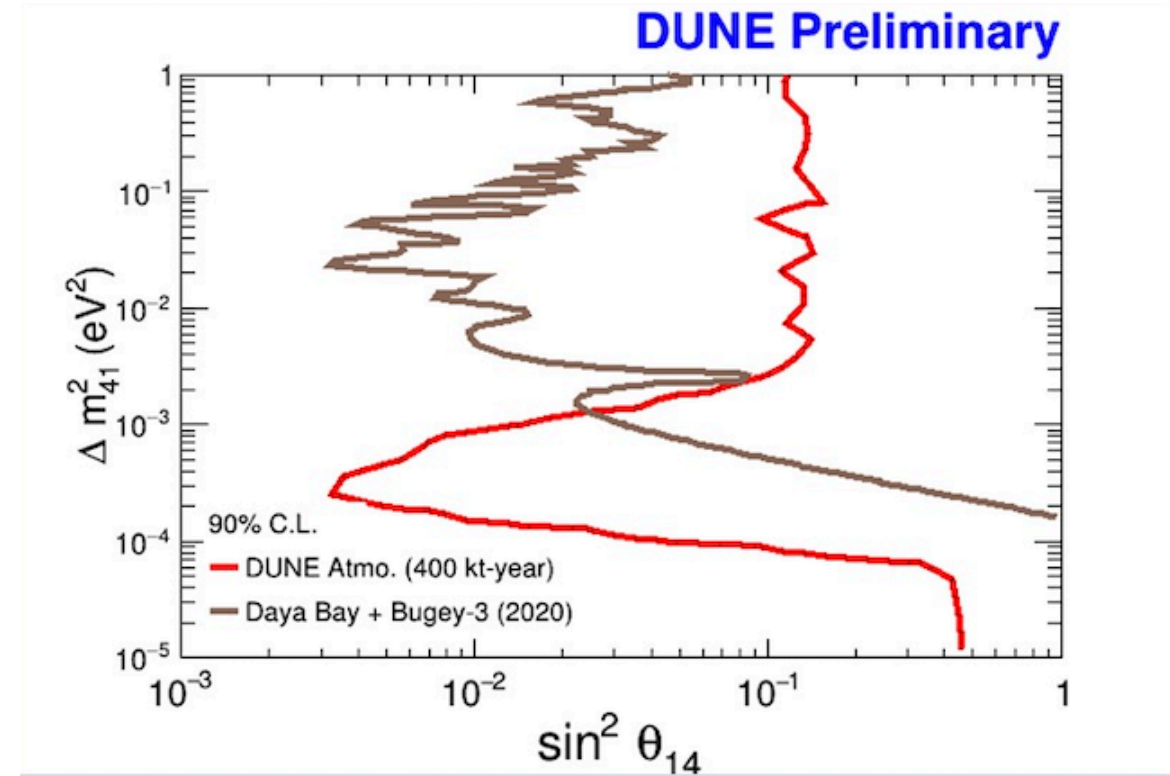
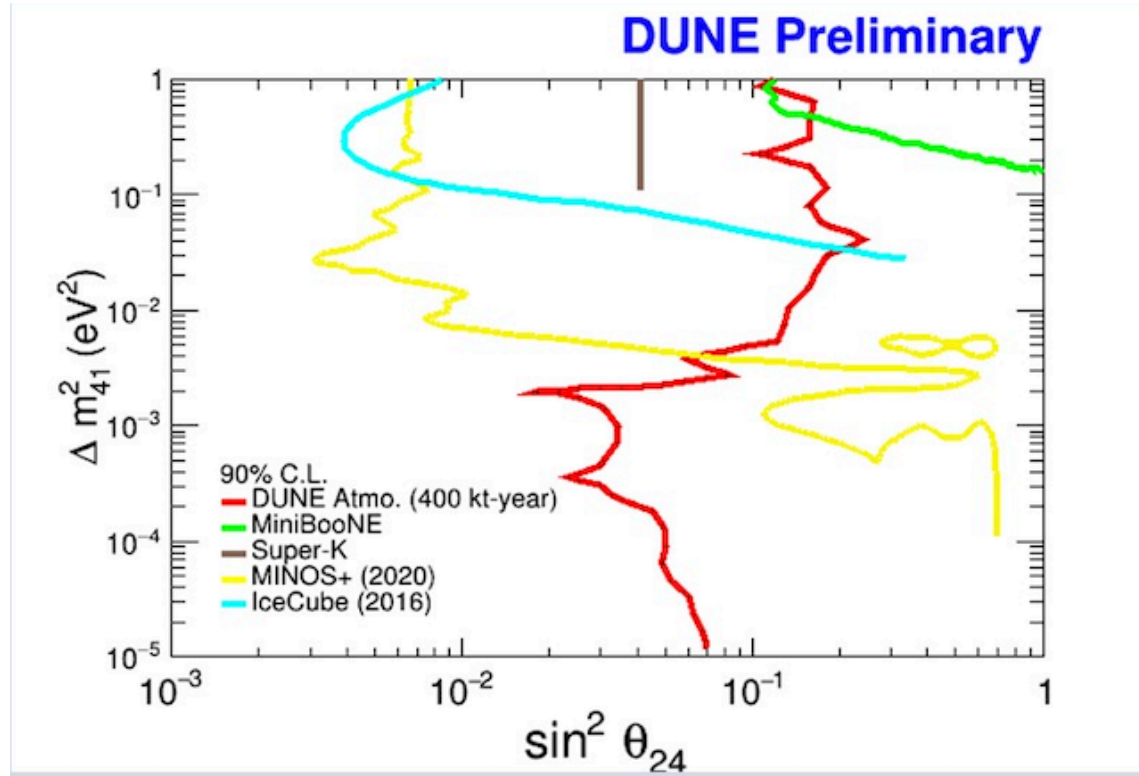
# Summary and Outlook

- ▶ The highly-capable DUNE detectors and the powerful LBNF beam will enable a very rich and diverse program for New Physics probes in the next decades
- ▶ DUNE has powerful physics reach for a broad range of beyond three-flavor neutrino mixing models
  - Can probe most flavor transition channels over wide energy spectrum within single expt.
  - Deployment of Phase II essential for DUNE to achieve its full BSM physics potential
  - Highly complementary to other efforts, ongoing or projected for the next decade
- ▶ As a machine for discovery, DUNE will provide leading guidance to experimental and theoretical efforts involving neutrinos and/or new particles/interactions
- ▶ Stay tuned for the next talk on further exciting DUNE BSM Physics opportunities!



# Supplements

# Preliminary Atmos. Sterile Sensitivities



- ▶ Assuming 400 kton.years exposure

	Track-like events	Shower-like events
Reconstruction efficiency (CC)	80%	80%
Reconstruction efficiency (NC)		0.5%
Neutrino energy resolution	18%	13%
Neutrino direction resolution	10 degrees	10 degrees

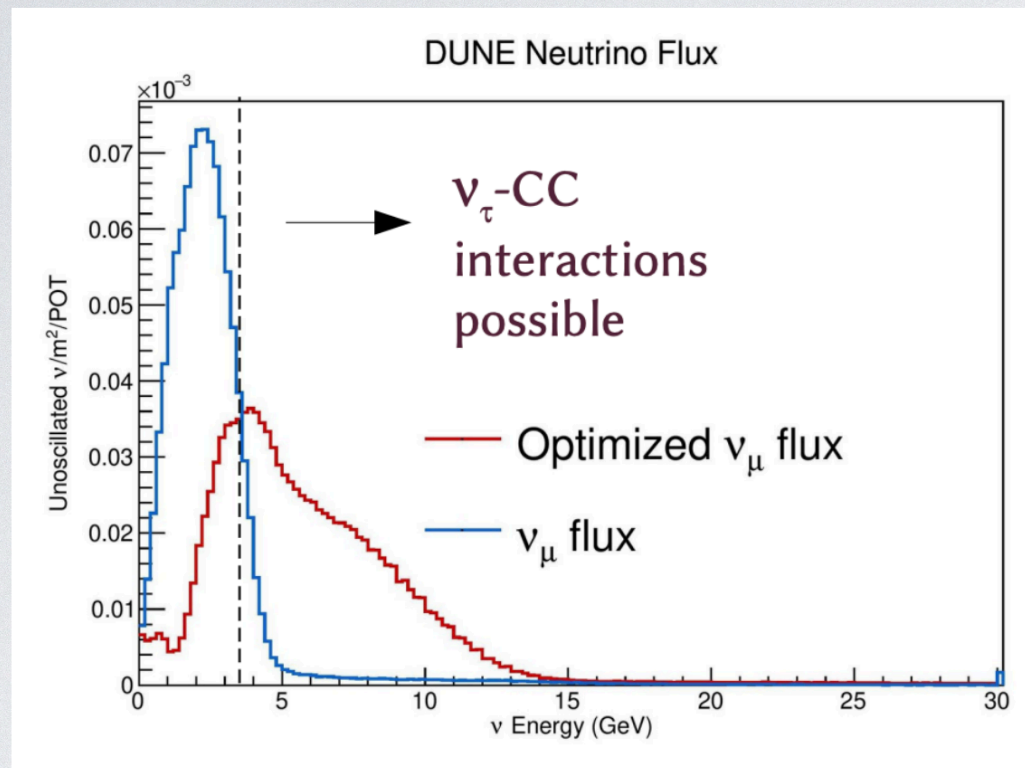
*T. Thakore, Neutrino '22*



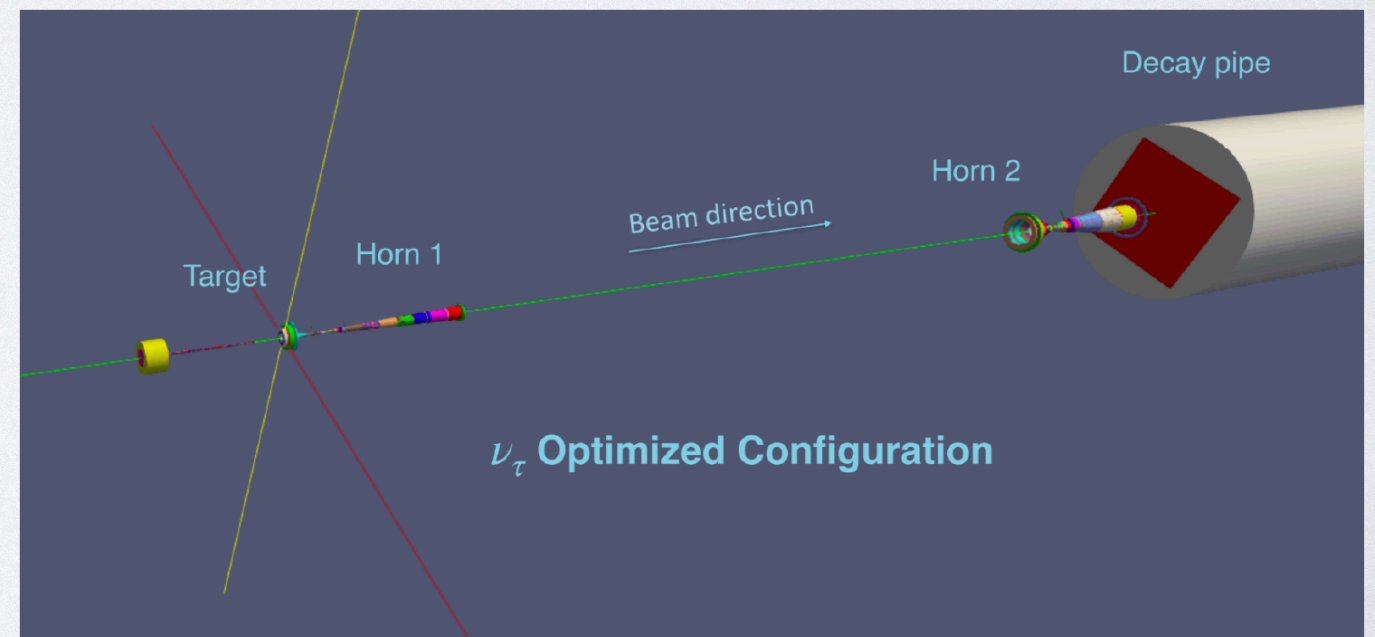
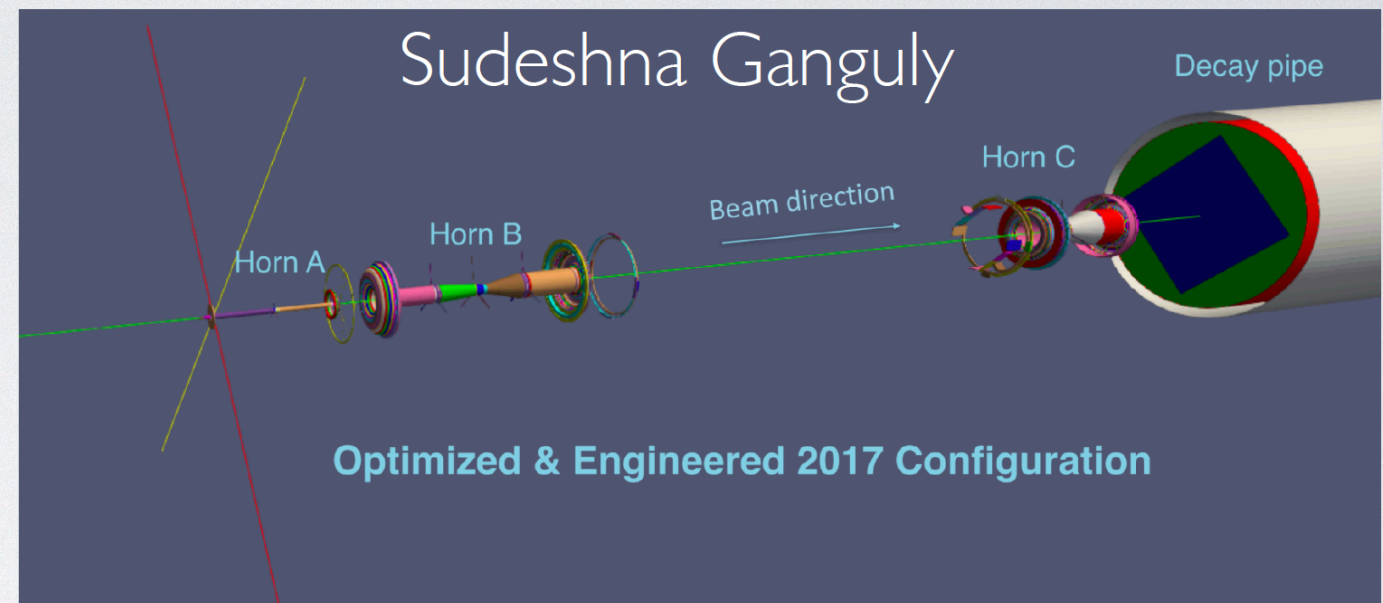
# $\nu_\tau$ -Optimized Beam

Laura Fields, NuTau 2021

## FUTURE: DUNE



- Beamline **can be tuned to higher energy** by using two NuMI horns and increasing horn separation
- Fairly **simple optimization**; can probably be improved on, but not dramatically



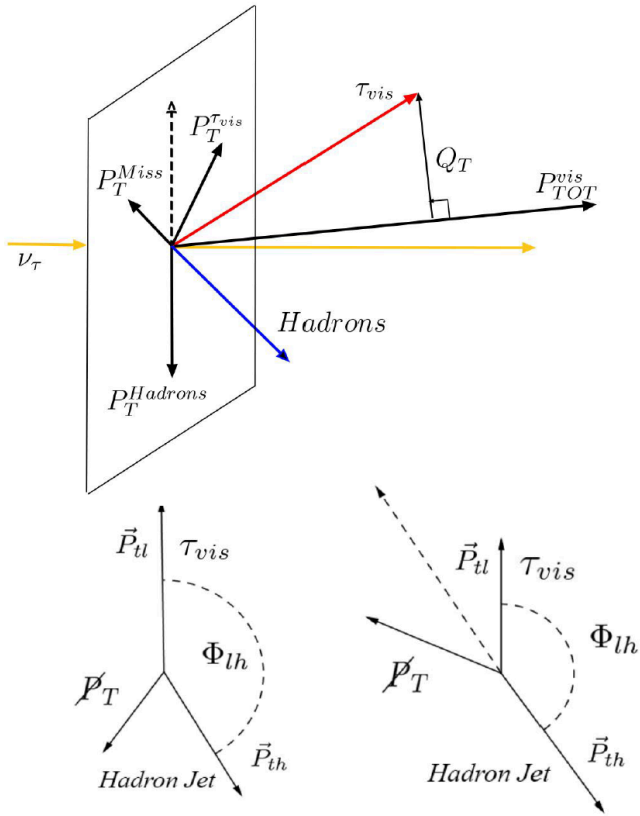


# $\nu_\tau$ CC Selection Variables

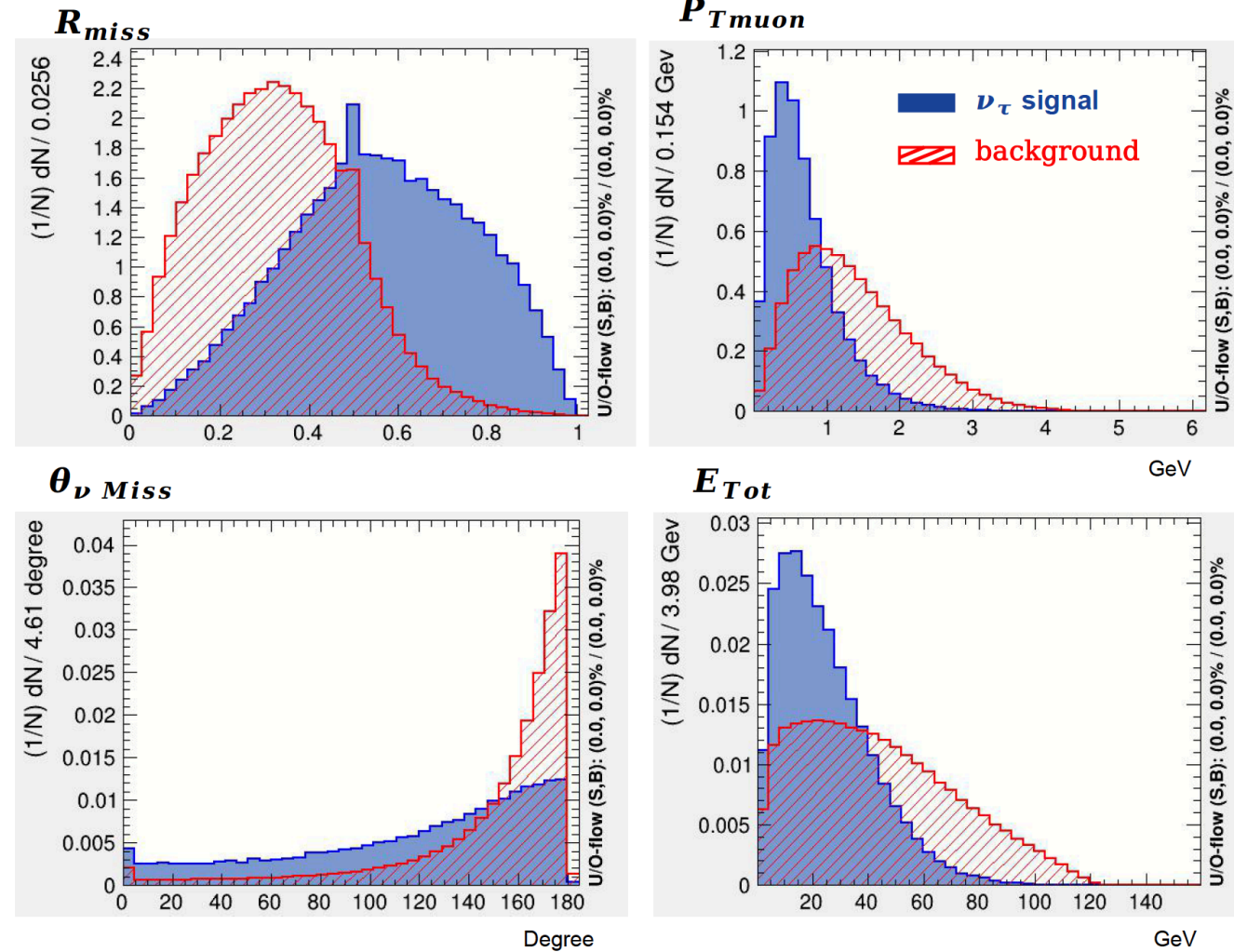
- The signal and background separation is based on kinematic differences. Used a total of 18 variables.

Table shows 6 variables providing highest signal/bg separation

Ranking	Kinematic variables
1	$R_{miss} = \frac{P_{T Miss}}{P_{T Miss} + P_{T muon}}$
2	$P_{T muon}$ : transverse lepton momentum
3	$\theta_{\nu Miss}$ : angle between beam direction and missing transverse momentum
4	$E_{Tot}$ : total visible energy
5	$\Phi_{muon hadron}$ : angle between transverse muon and hadron momentum
6	$P_{Tot}$ : total transverse momentum



Background interaction products in the transverse plane  $\nu_\tau$  CC interaction products in the transverse plane



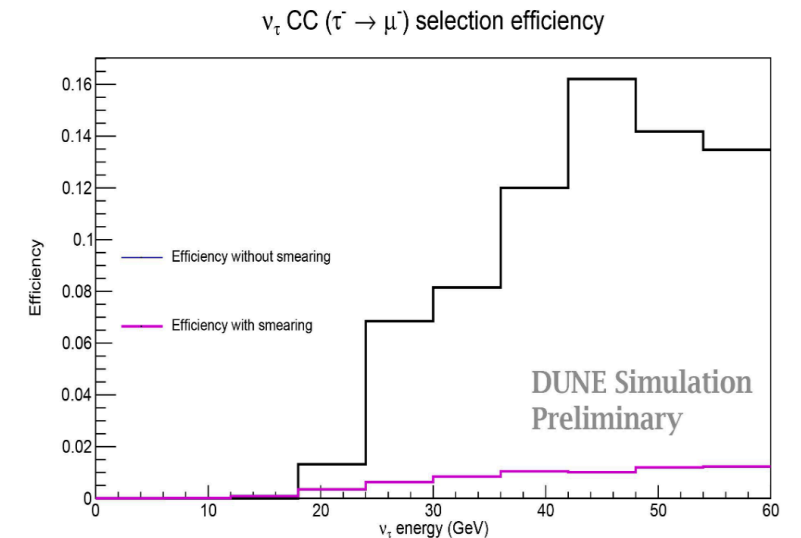
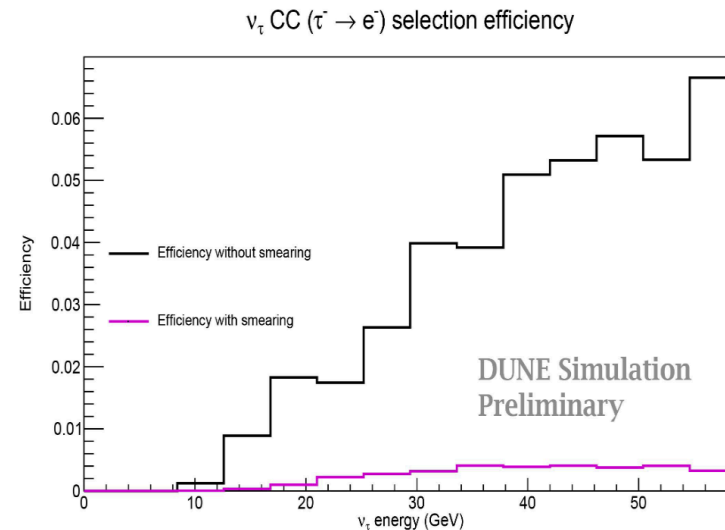
Examples of  $\nu_\tau$  signal (blue) and background (red) kinematic variables distribution.

► Electron channel

► Muon channel

H. Razafinime, Neutrino '22

M. Rajaolisoa, APS '21



# Proton Decay

- ▶ Proton decay expected to occur in Grand Unified Theories with a lifetime of  $\sim 10^{34} - 10^{36}$  years
- ▶ DUNE is most sensitive to proton decay in the  $p \rightarrow K^+ + \bar{\nu}$  channel
  - Excellent calorimetric capabilities of LArTPCs enable good kaon identification, as well as high kaon/muon tracking efficiency
  - A lower limit on the proton lifetime of  $1.3 \times 10^{34}$  yrs@90%CL is expected if no signal is observed in 10 years

- ▶ World's Best Limits from SuperK:
  - $\tau(p \rightarrow e^+ \pi^0) > 2.4 \times 10^{34}$  years
  - $\tau(p \rightarrow K^+ \bar{\nu}) > 8.2 \times 10^{33}$  years

