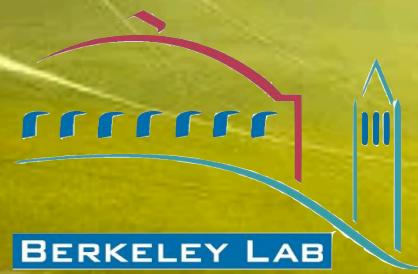


# THEIA

*Let there be  
light*

*Gabriel D. Orebi Gann*  
UC Berkeley & LBNL  
Snowmass  
20th July, 2022



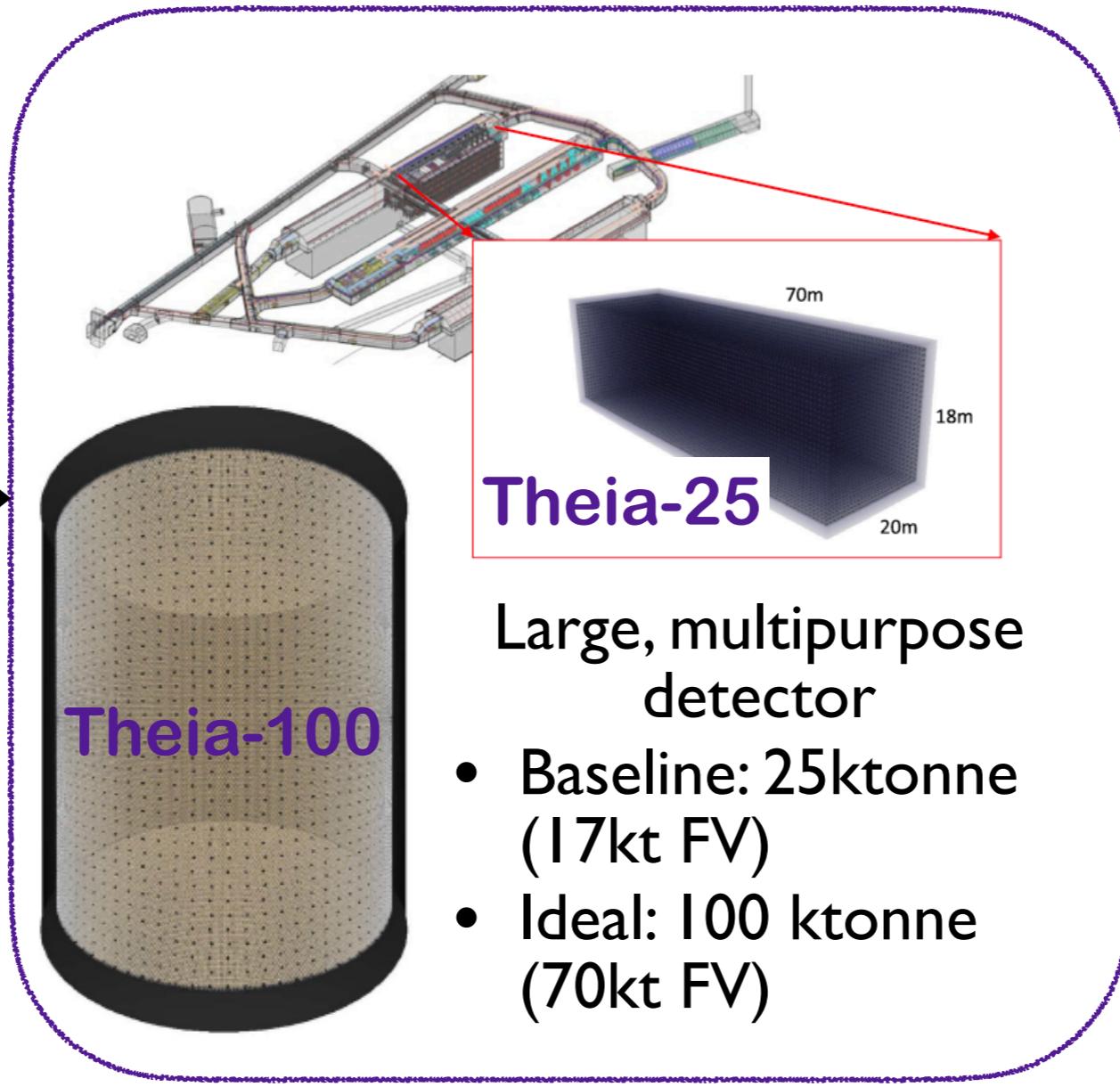
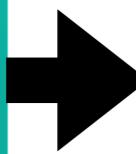
[MyGodPictures.com](http://MyGodPictures.com)

# Theia as a DUNE module in Phase II

*Long-baseline sensitivity comparable to a LAr DUNE module*

*Complementary supernova sensitivity (primarily anti- $\nu$ , fast response: can act as trigger)  
+ broad additional physics program*

Cutting-edge developments in target material and photon detection



**Theia-25**

Large, multipurpose detector

- Baseline: 25 ktonne (17 kt FV)
- Ideal: 100 ktonne (70 kt FV)

Broad physics program:  
Studying neutrino fundamental properties and astrophysical objects

THEIA: An advanced optical neutrino detector  
Eur. Phys. J. C 80, 416 (2020)

# THEIA Collaboration



White paper - Eur. Phys. J. C 80, 416 & arXiv:2202.12839 [hep-ex]



## Canada

Alberta  
Laurentian  
Queens  
SNOLAB  
Toronto

## China

Tsinghua

## Finland

Jyvaskyla

## Germany

Aachen  
Dresden  
Hamburg  
Jülich  
Mainz

## TU Munich Tübingen

## Italy

SISSA/INFN

## Korea

CUP

## Portugal

LIP  
Lisbon

## Turkey

Erciyes

## UK

King's College  
Sheffield

## US

BNL  
Boston  
Chicago  
Colorado  
Cornell  
FNAL  
U. Hawaii  
Iowa

## Iowa State LBNL

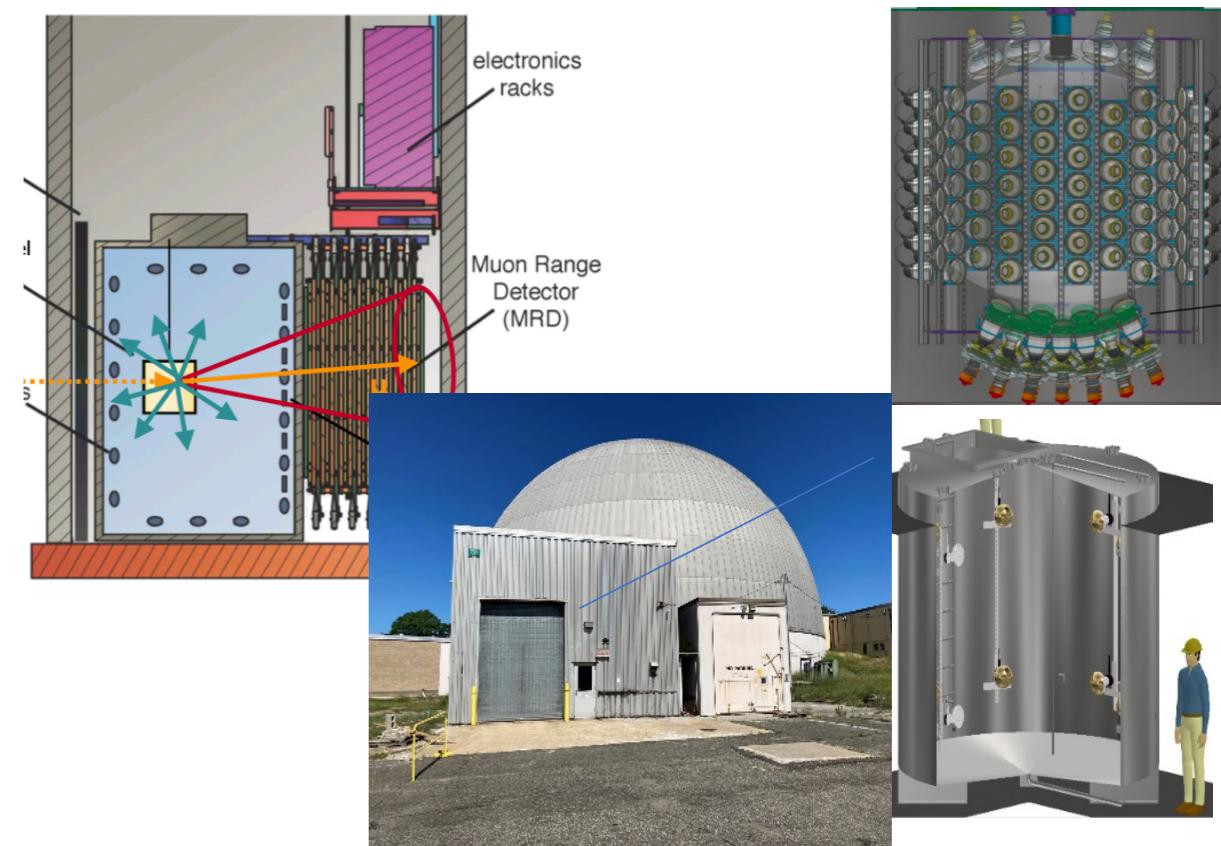
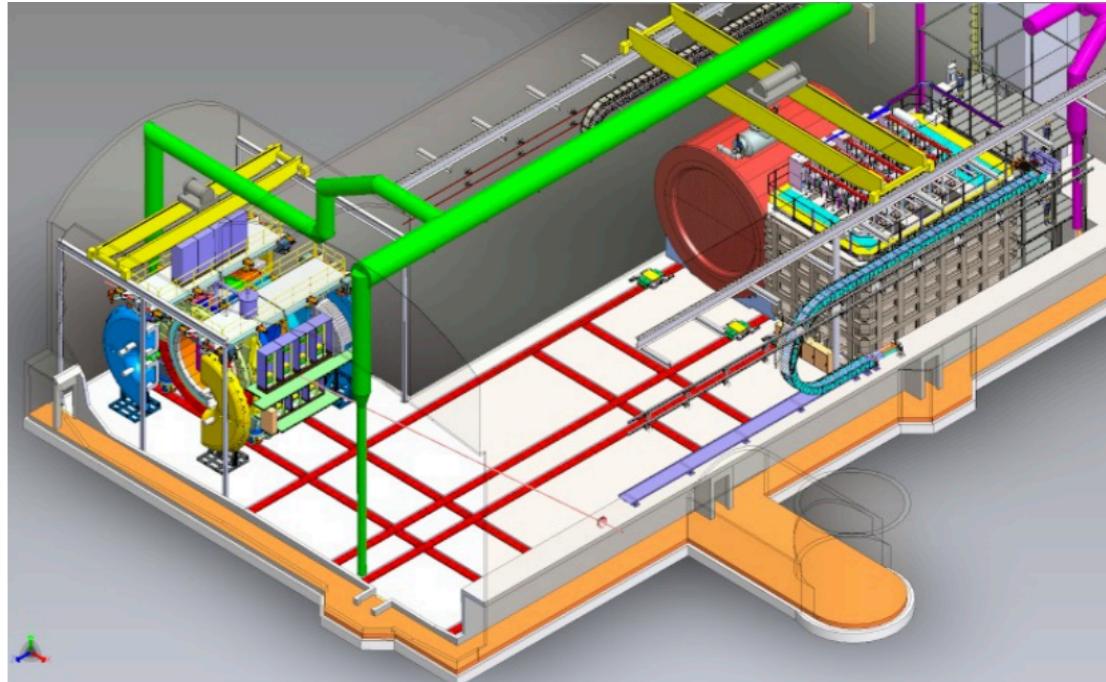
LLNL  
LSU  
MIT  
U. Penn  
PNNL  
Rutgers  
SD SMT

## Stony Brook SURF

Temple  
UC Berkeley  
UC Davis  
UC Irvine  
UCLA

# Challenges

- Near detector: many options, thanks to the flexible design of this highly sophisticated facility
  - SAND: planned C & H targets
  - Potential alternative target options for ND-GAr?
  - Additional module for DUNE-PRISM?
- Cavern shape: letterbox suboptimal
- Utilities: UG deployment, sacrifice mass?
- Program of prototypes for necessary technological demonstrations

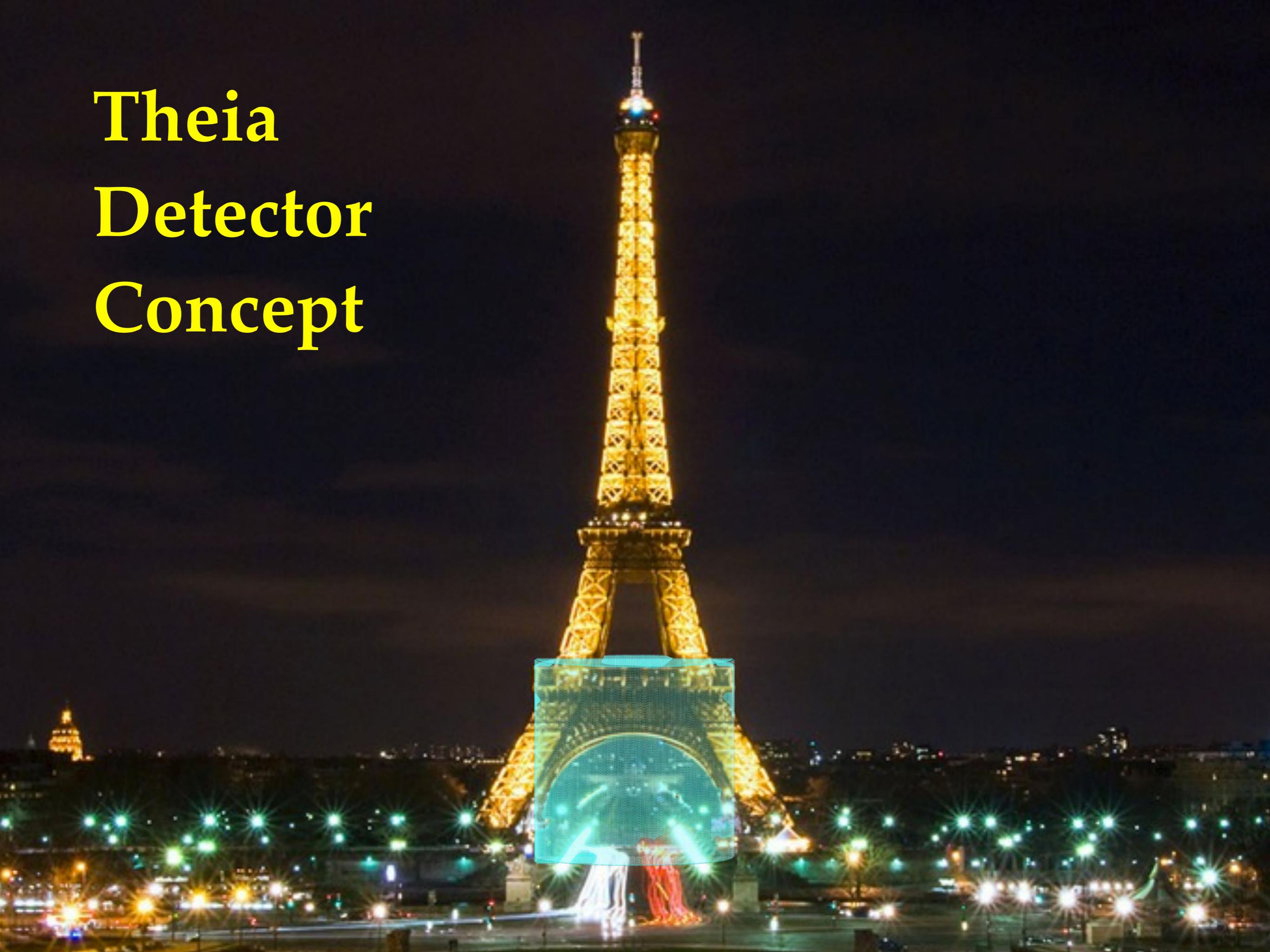


# Overview

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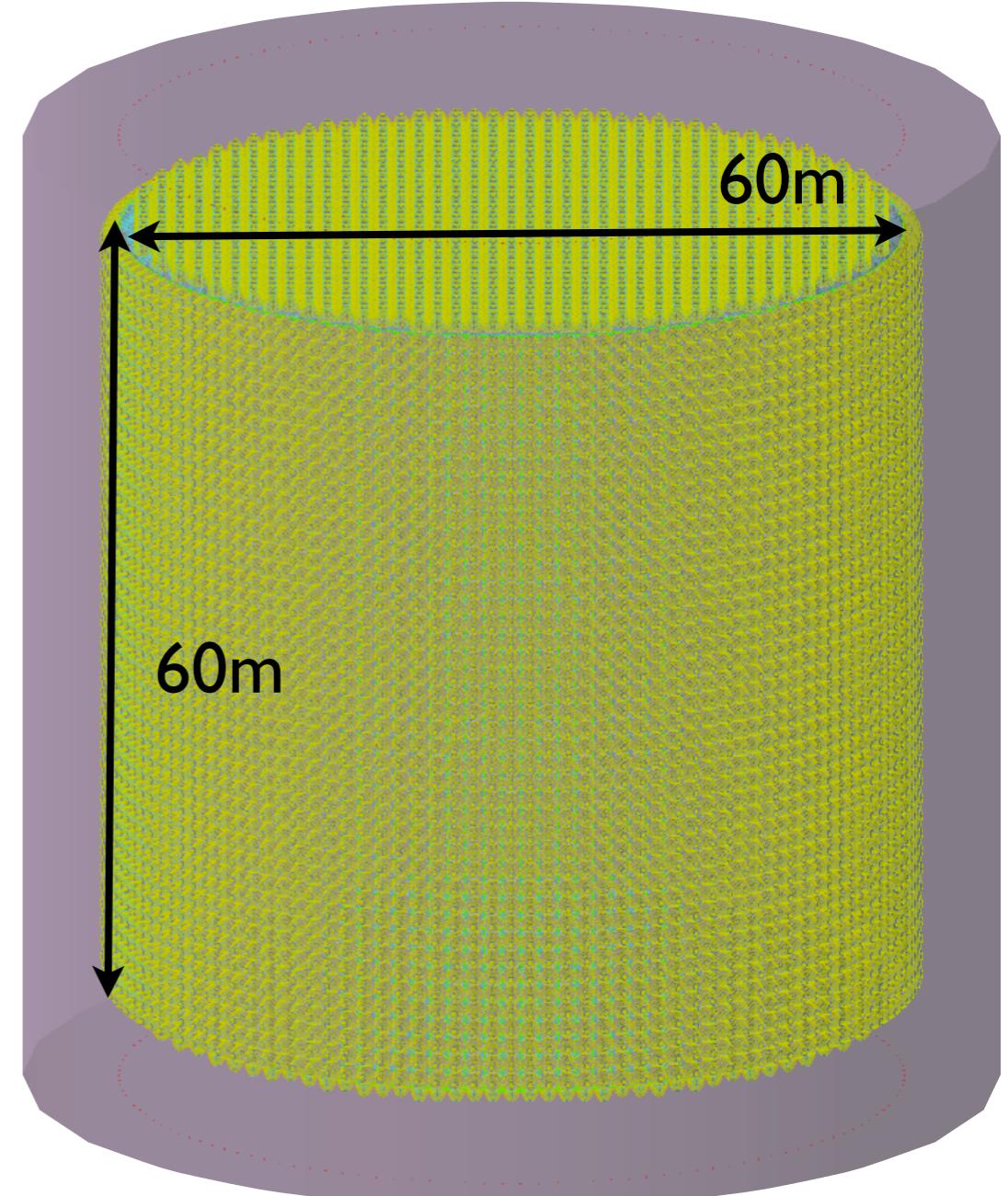
- THEIA detector
- Hybrid neutrino detection concept
- Detector development
- Prototypes (*T. Kaptanoglu, 7/24 @ 10.48am*)
- Physics program (*Z. Bagdasarian, L. Pickard, 7/24 @ 10.00, 10.20am*)

# Theia Detector Concept



# THEIA

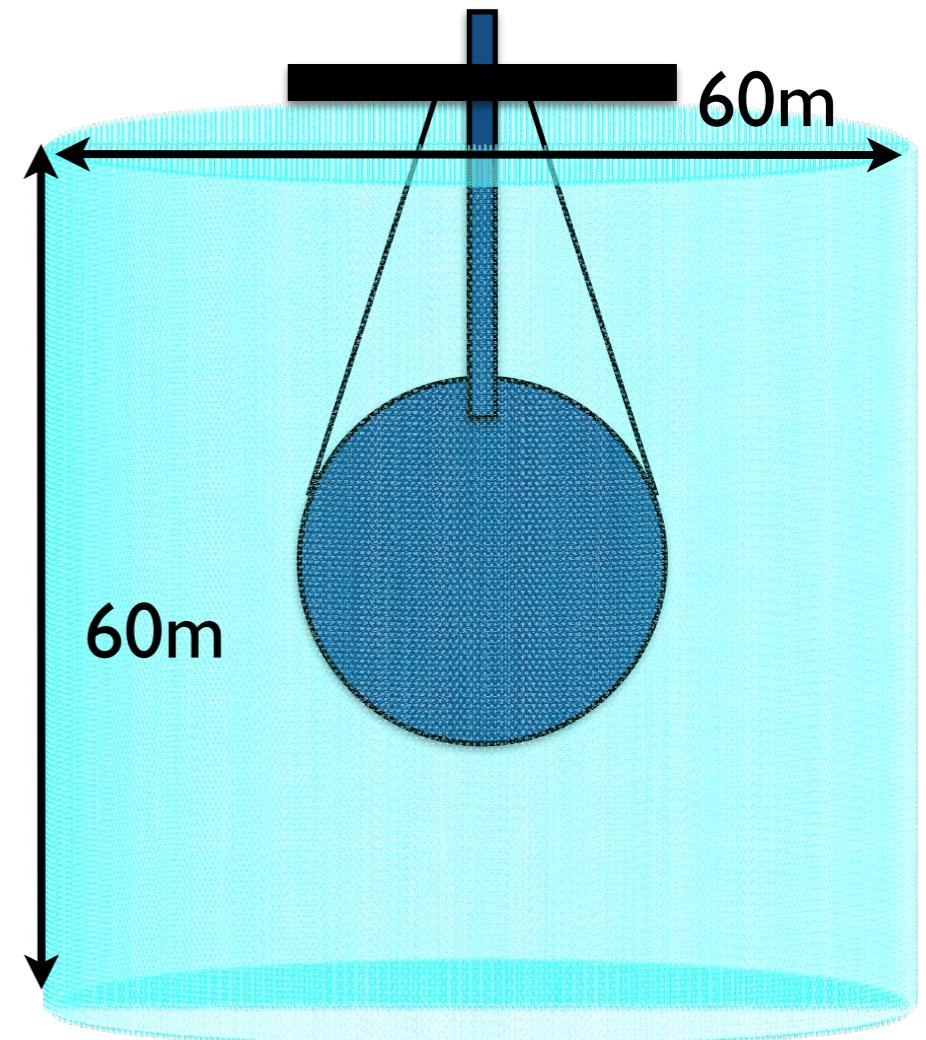
- Large-scale detector (25-100 kton)
- Novel LS target e.g. WbLS
- Fast, high-efficiency (spectrally sensitive?) photon detection with high coverage
- Deep underground
- Isotope loading (Gd, Te, Li...)
- *Flexible!* Target, loading, configuration  
→ Broad physics program!



White paper - [Eur. Phys. J. C 80, 416 \(2020\)](#)  
[& arXiv:2202.12839 \[hep-ex\]](#)

# THEIA

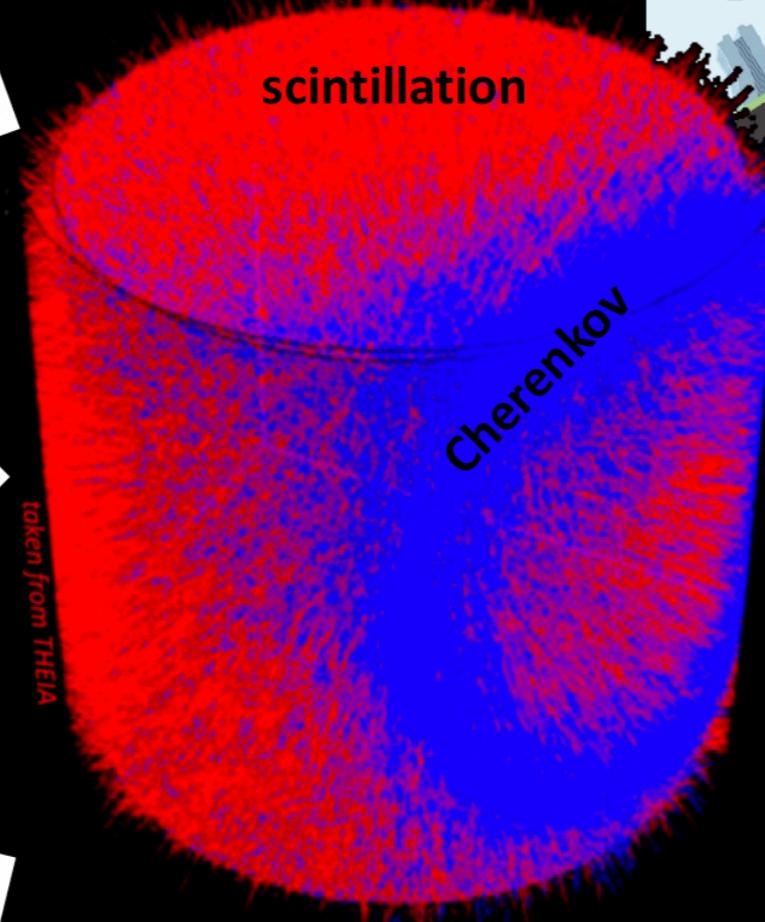
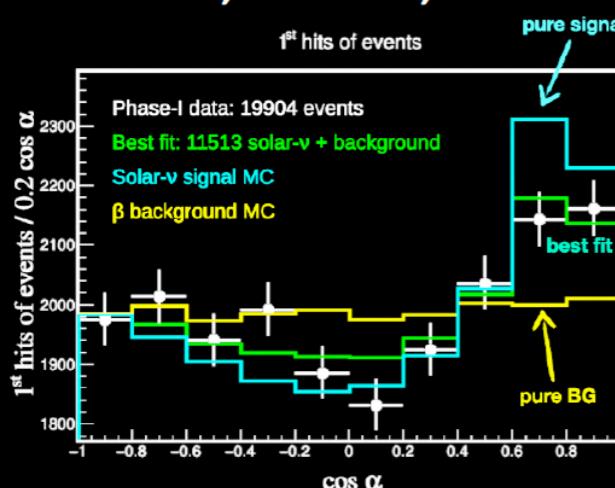
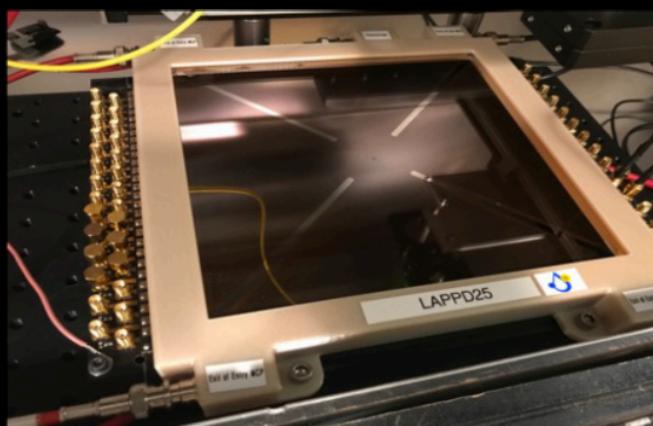
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→ Broad physics program!



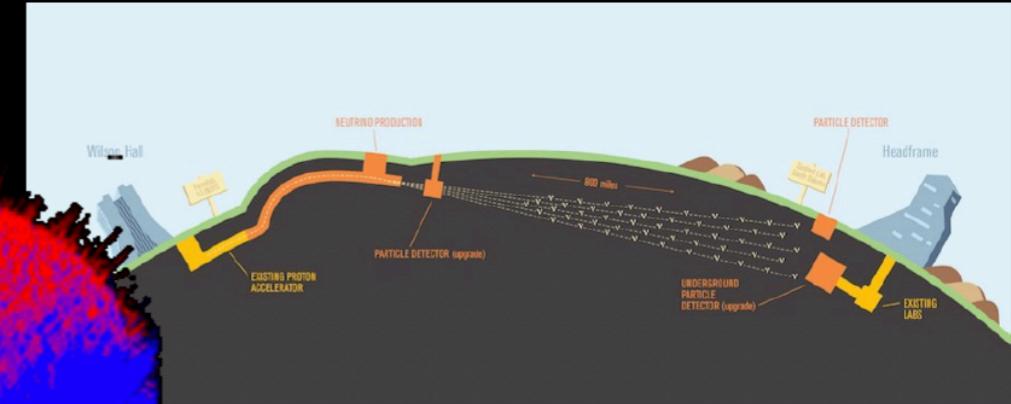
White paper - [Eur. Phys. J. C 80, 416 \(2020\)](#)  
[& arXiv:2202.12839 \[hep-ex\]](#)



# Hybrid neutrino detection

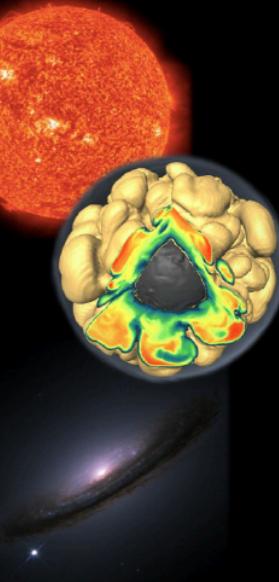


→ Enhanced sensitivity to broad physics program



→ Long-Baseline Oscillations

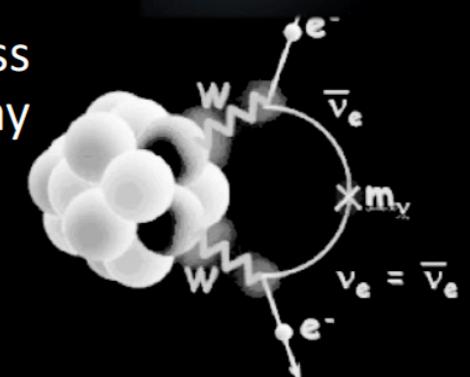
→ Solar neutrinos



→ Supernova neutrinos

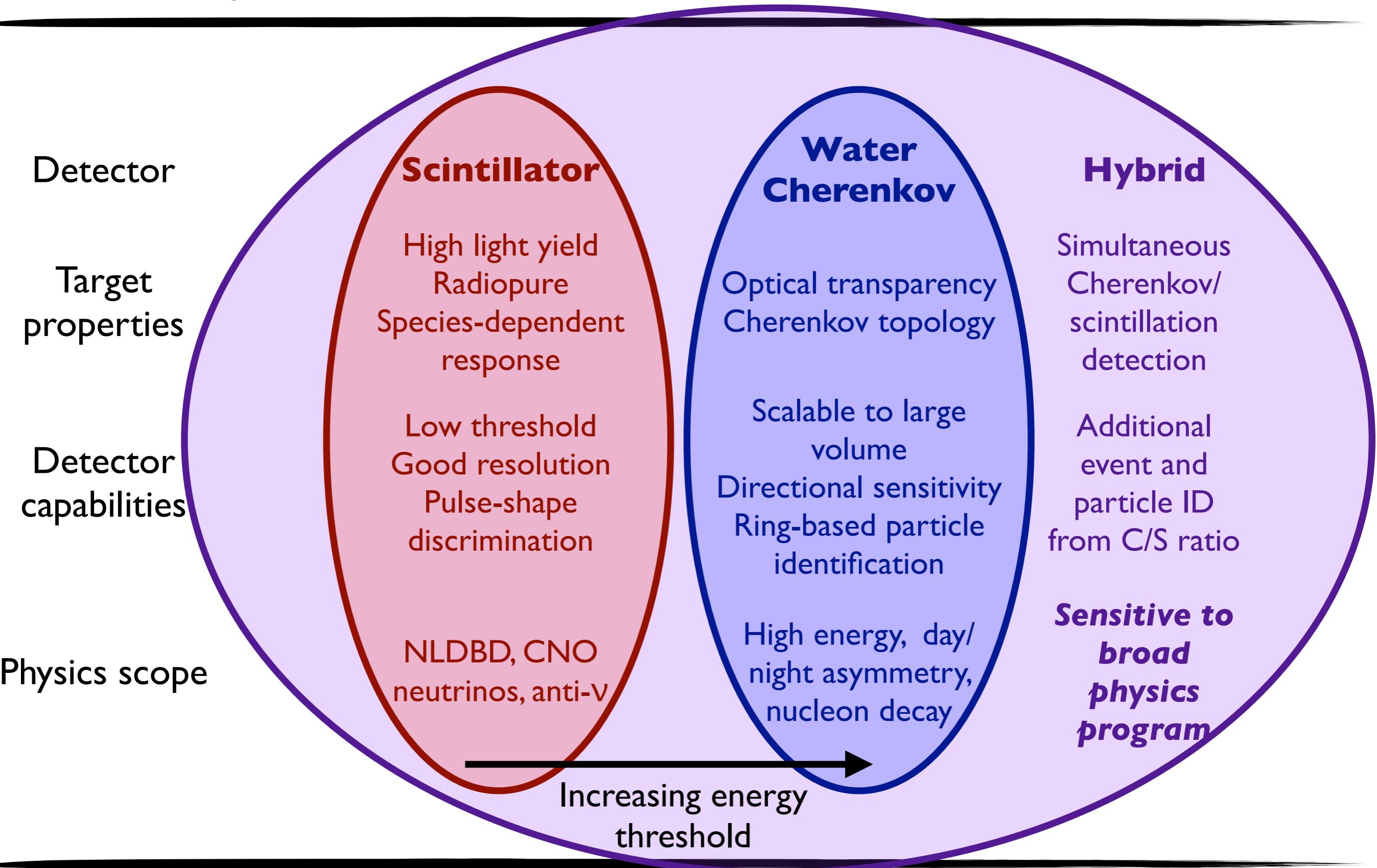
→ Diffuse SN neutrinos

→ Neutrinoless  
Double-Beta Decay



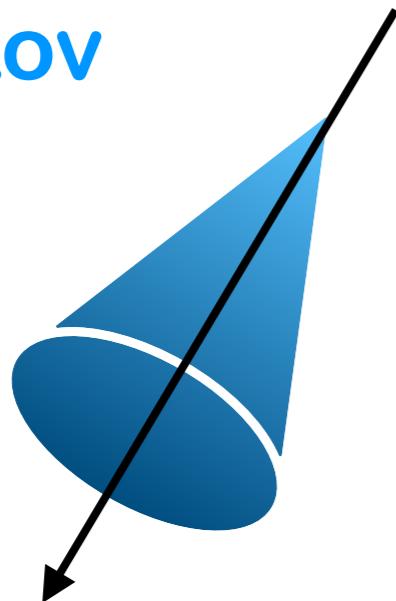
White paper - Eur. Phys. J. C 80, 416 & arXiv:2202.12839 [hep-ex]

# Hybrid neutrino detection

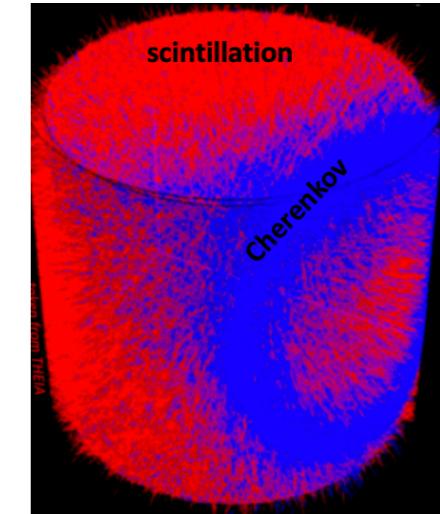
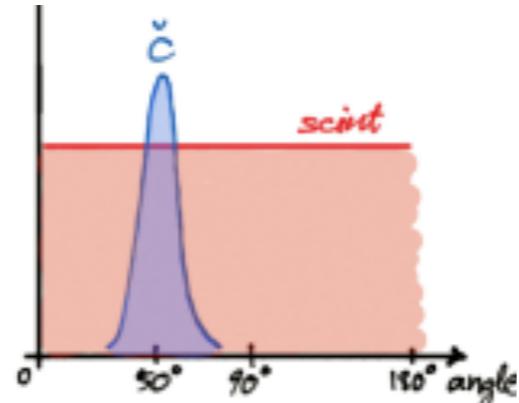


# How can it be done?

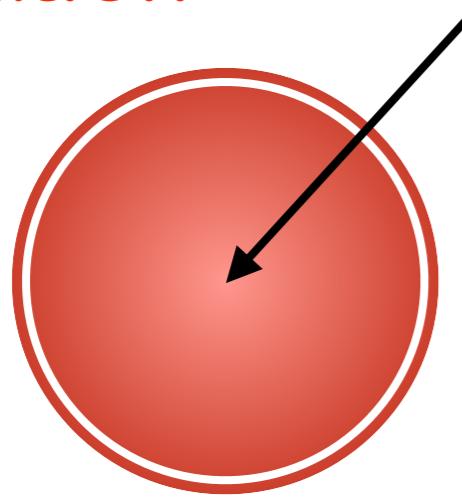
Cherenkov



Angular distribution

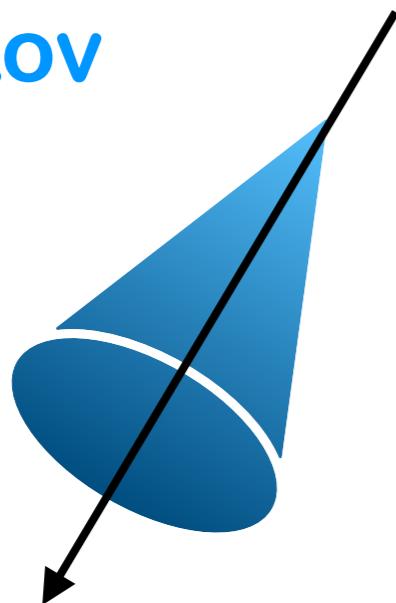


Scintillation

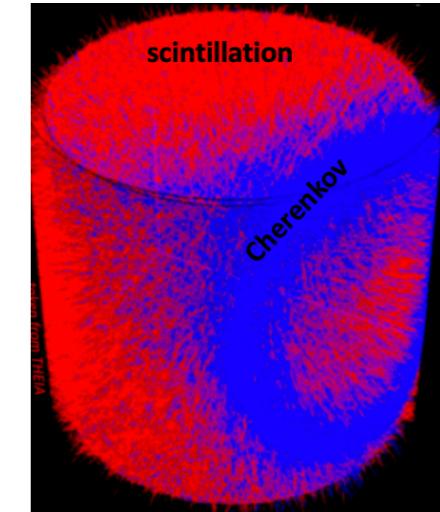
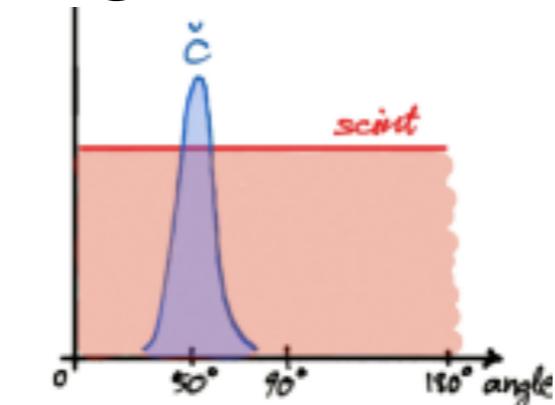


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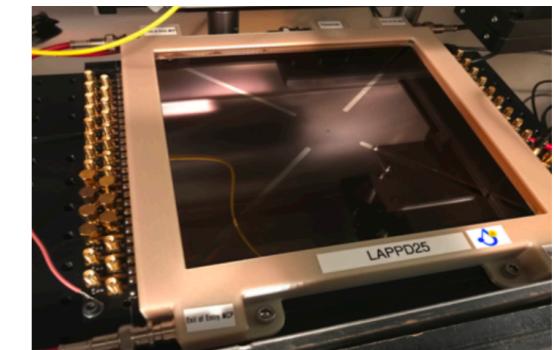
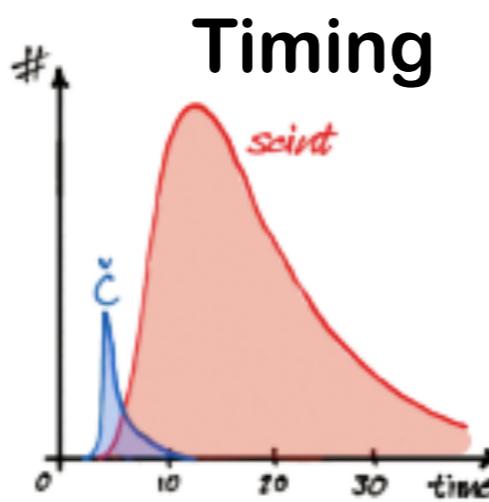
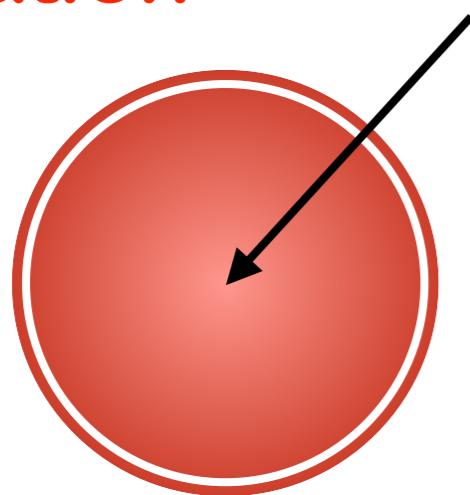
Cherenkov



Angular distribution



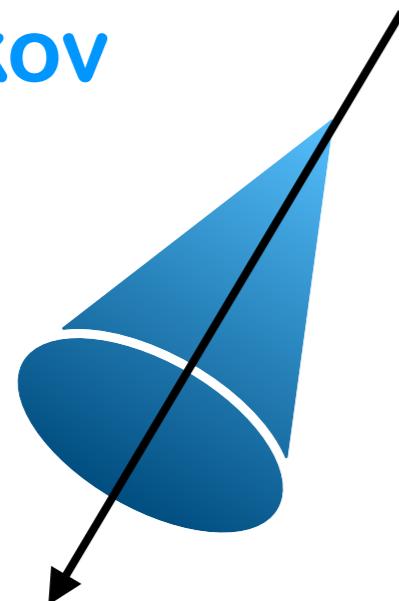
Scintillation



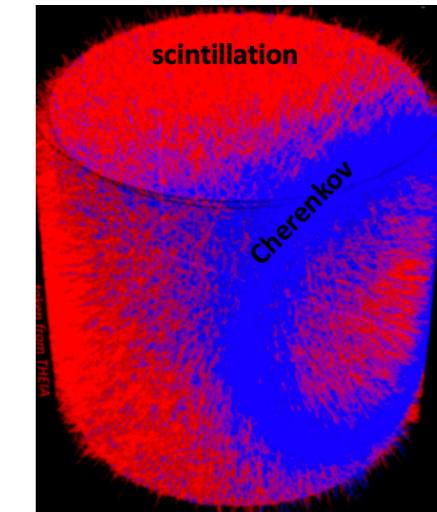
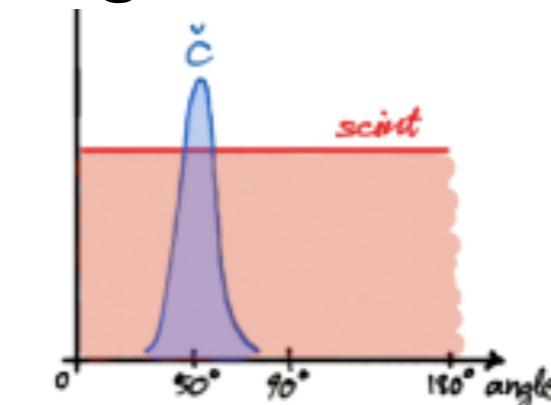
B.W.Adams et al. NIM A Volume 795, 1 (2015)

# How can it be done?

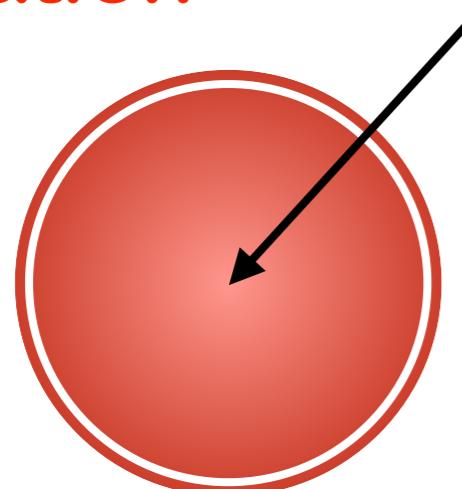
Cherenkov



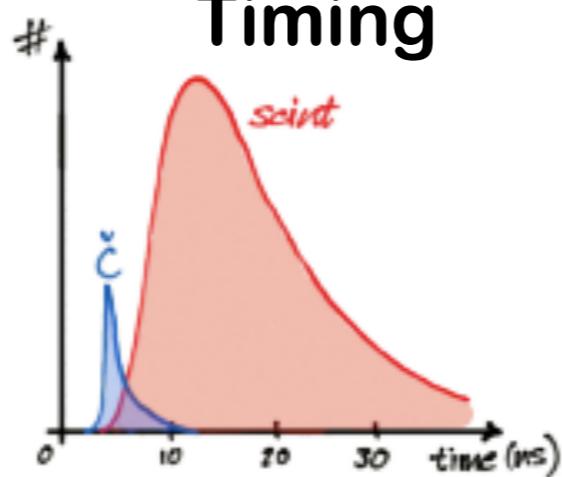
Angular distribution



Scintillation

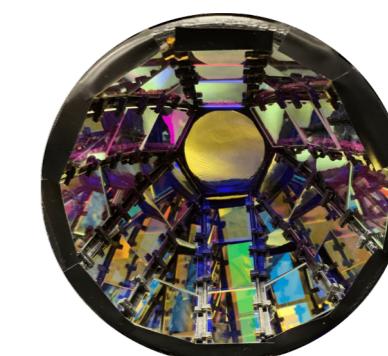
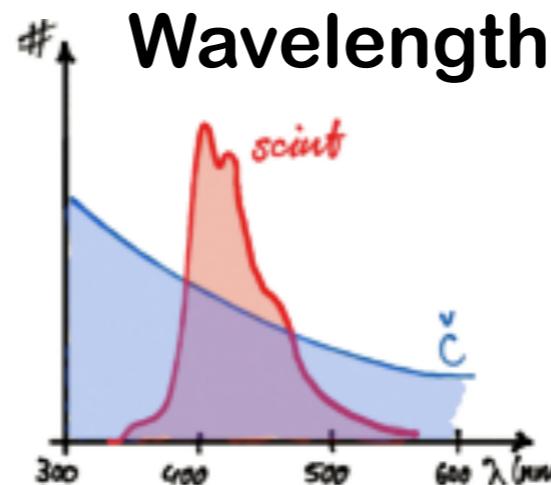


Timing



B.W.Adams et al. NIM A Volume 795, 1 (2015)

Wavelength

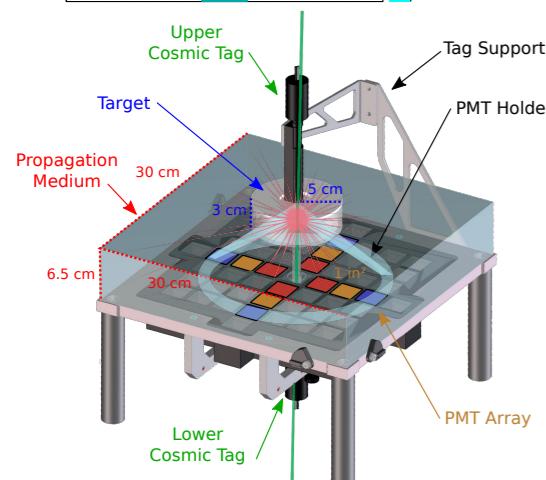
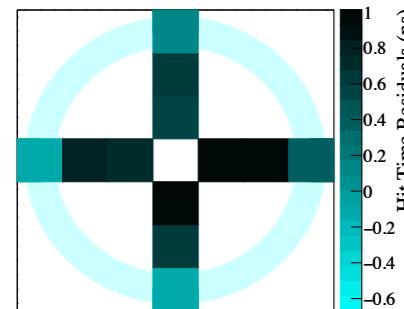


T. Kaptanoglu et al.  
Phys. Rev. D 101, 072002 (2020)

Builds on core  
(Wb)LS  
development at  
BNL (Yeh et al.)

# Detector development

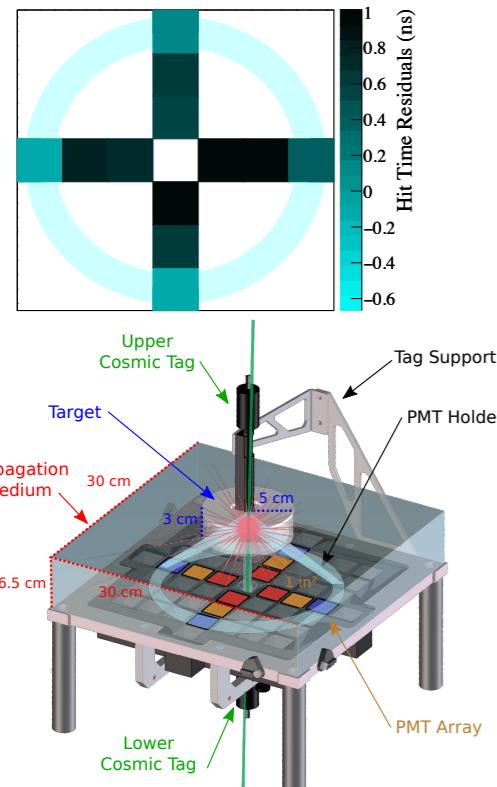
## CHESS detector: LBNL



*Phys. Rev. C 95 055801 (2017); Eur. Phys. Jour. C 80 867 (2020); Mat. Adv. I 71 (2020); Eur. Phys. Jour. C 82 169 (2022); NIMA 947, 162604 (2019); arXiv:1902.06912; JINST13 P07005 (2018); JINST9 P06012 (2014); NIMA 943 162420 (2019); Eur. Phys. Jour. C 77 811 (2017); arxiv:1908.03564; arXiv:1502.01132; arXiv:1707.08222; NIMA 972 164106 (2020); Astropart. Phys. 109 33 (2019); NIMA 852 15 (2017); NIMA 712 162 (2013); Phys. Rev. D 97 052006 (2018); JINST14 I (2019); Phys. Rev. D 101 072002 (2020); arXiv:2006.00173*

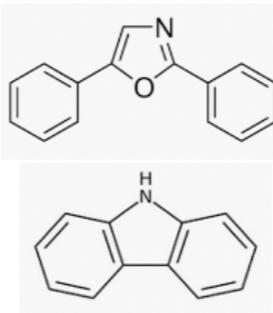
# Detector development

## CHESS detector: LBNL



## Engineering WbLS properties: Bourret (LBNL)

- Example:  
slowing down decay time
- Standard PPO → c1ccc(cc1)-c2cc(c3ccccc3)on2  
2ns
  - New carbazole → c1ccc2c(c1)[nH]c3ccccc23  
15ns

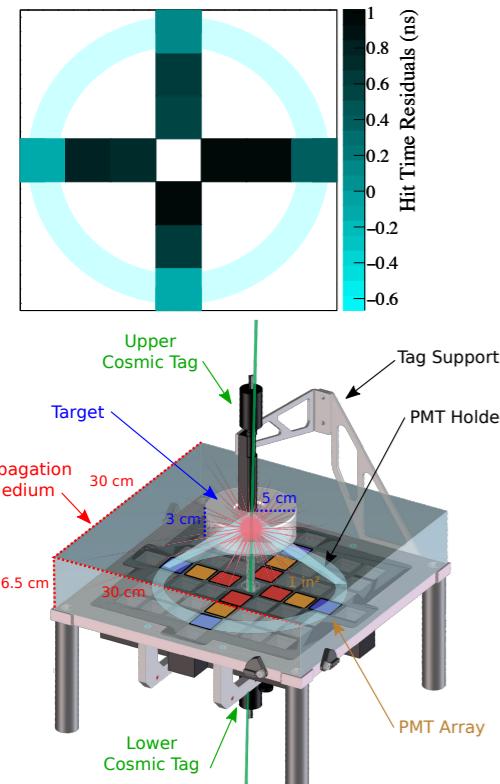


[Phys. Rev. C 95 055801 \(2017\); Eur. Phys. Jour. C 80 867 \(2020\); Mat. Adv. I 71 \(2020\); Eur. Phys. Jour. C 82 169 \(2022\); NIMA 947, 162604 \(2019\); arXiv:1902.06912; JINST13 P07005 \(2018\); JINST9 P06012 \(2014\); NIMA 943 162420 \(2019\); Eur. Phys. Jour. C 77 811 \(2017\); arxiv:1908.03564; arXiv:1502.01132; arXiv:1707.08222; NIMA 972 164106 \(2020\); Astropart. Phys. 109 33 \(2019\); NIMA 852 15 \(2017\); NIMA 712 162 \(2013\); Phys. Rev. D 97 052006 \(2018\); JINST14 I \(2019\); Phys. Rev. D 101 072002 \(2020\); arXiv:2006.00173](#)

Builds on core  
(Wb)LS  
development at  
BNL (Yeh et al.)

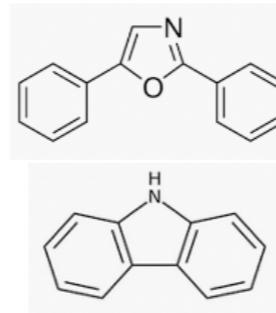
# Detector development

## CHESS detector: LBNL

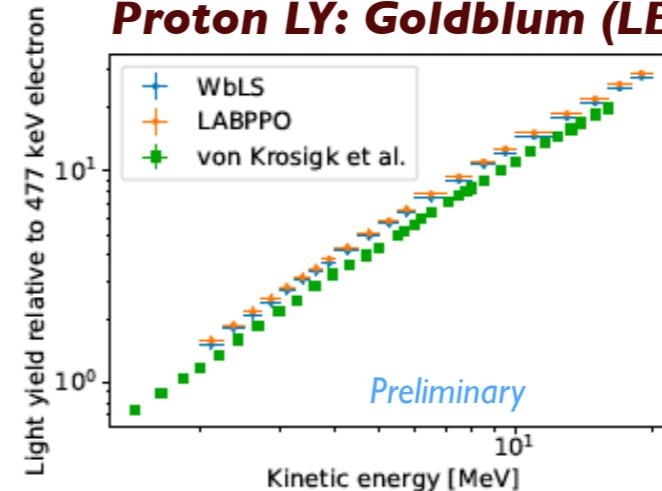


## Engineering WbLS properties: Bourret (LBNL)

- Example:  
slowing down decay time
- Standard PPO → 2ns
  - New carbazole → 15ns



## Proton LY: Goldblum (LBNL)

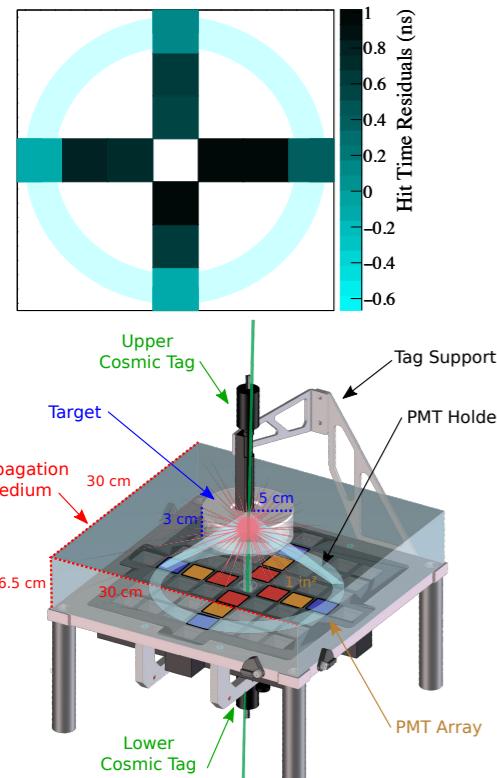


*Phys. Rev. C 95 055801 (2017); Eur. Phys. Jour. C 80 867 (2020); Mat. Adv. I 71 (2020); Eur. Phys. Jour. C 82 169 (2022); NIMA 947, 162604 (2019); arXiv:1902.06912; JINST13 P07005 (2018); JINST9 P06012 (2014); NIMA 943 162420 (2019); Eur. Phys. Jour. C 77 811 (2017); arxiv:1908.03564; arXiv:1502.01132; arXiv:1707.08222; NIMA 972 164106 (2020); Astropart. Phys. 109 33 (2019); NIMA 852 15 (2017); NIMA 712 162 (2013); Phys. Rev. D 97 052006 (2018); JINST14 I (2019); Phys. Rev. D 101 072002 (2020); arXiv:2006.00173*

Builds on core  
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development at  
BNL (Yeh et al.)

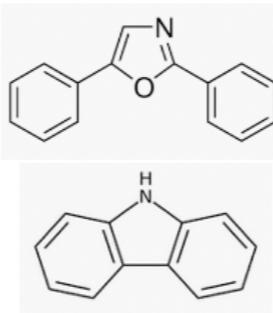
# Detector development

## CHESS detector: LBNL



## Engineering WbLS properties: Bourret (LBNL)

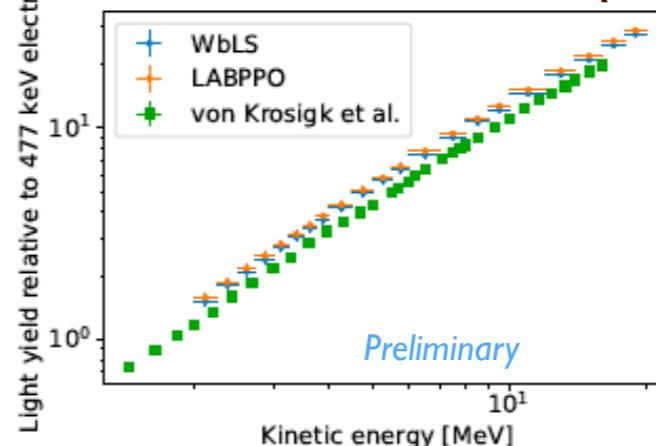
- Example:  
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- Standard PPO → 2ns
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## Scattering & attenuation: UC Davis, UC Berkeley+LLNL



## Proton LY: Goldblum (LBNL)

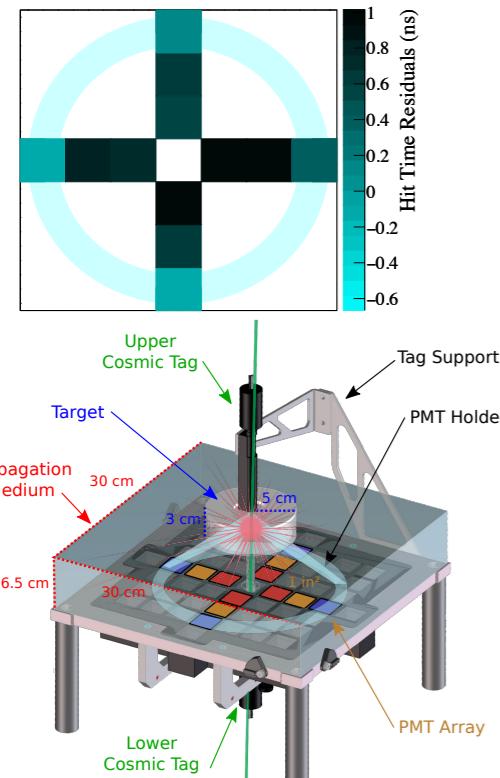


*Phys. Rev. C 95 055801 (2017); Eur. Phys. Jour. C 80 867 (2020); Mat. Adv. I 71 (2020); Eur. Phys. Jour. C 82 169 (2022); NIMA 947, 162604 (2019); arXiv:1902.06912; JINST13 P07005 (2018); JINST9 P06012 (2014); NIMA 943 162420 (2019); Eur. Phys. Jour. C 77 811 (2017); arxiv:1908.03564; arXiv:1502.01132; arXiv:1707.08222; NIMA 972 164106 (2020); Astropart. Phys. 109 33 (2019); NIMA 852 15 (2017); NIMA 712 162 (2013); Phys. Rev. D 97 052006 (2018); JINST14 I (2019); Phys. Rev. D 101 072002 (2020); arXiv:2006.00173*

Builds on core  
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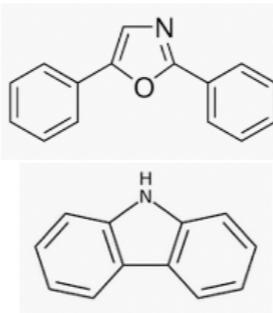
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## CHESS detector: LBNL



## Engineering WbLS properties: Bourret (LBNL)

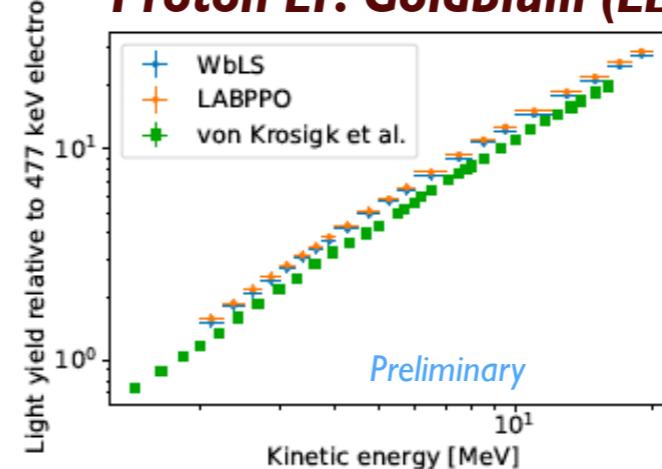
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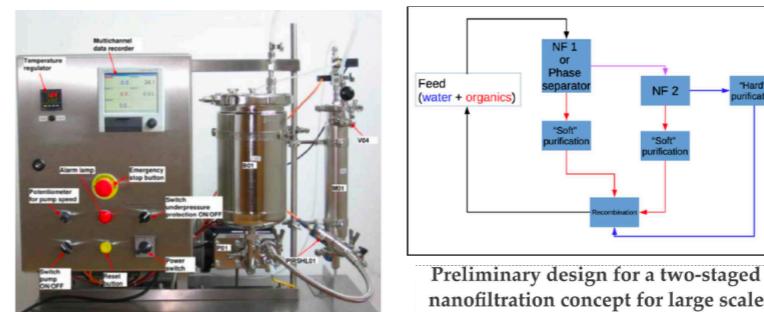
## Scattering & attenuation: UC Davis, UC Berkeley+LLNL



## Proton LY: Goldblum (LBNL)



## Nanofiltration: UC Davis

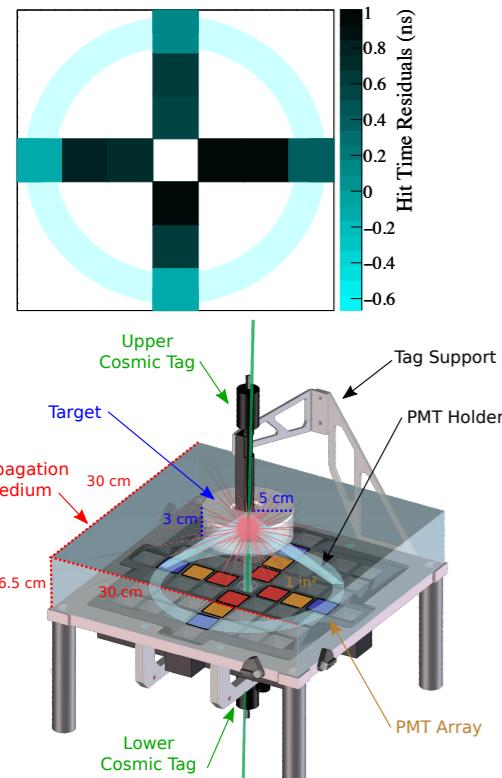


[Phys. Rev. C 95 055801 \(2017\); Eur. Phys. Jour. C 80 867 \(2020\); Mat. Adv. I 71 \(2020\); Eur. Phys. Jour. C 82 169 \(2022\); NIMA 947, 162604 \(2019\); arXiv:1902.06912; JINST13 P07005 \(2018\); JINST9 P06012 \(2014\); NIMA 943 162420 \(2019\); Eur. Phys. Jour. C 77 811 \(2017\); arXiv:1908.03564; arXiv:1502.01132; arXiv:1707.08222; NIMA 972 164106 \(2020\); Astropart. Phys. 109 33 \(2019\); NIMA 852 15 \(2017\); NIMA 712 162 \(2013\); Phys. Rev. D 97 052006 \(2018\); JINST14 I \(2019\); Phys. Rev. D 101 072002 \(2020\); arXiv:2006.00173](#)

Builds on core  
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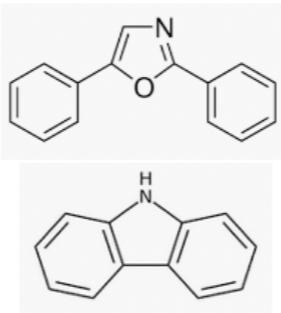
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## CHESS detector: LBNL



## Engineering WbLS properties: Bourret (LBNL)

- Example:  
slowing down decay time
- Standard PPO → 2ns
  - New carbazole → 15ns



## Scattering & attenuation: UC Davis, UC Berkeley+LLNL



1 top and 1 bottom PMT coupled to integrating spheres

5 side facing PMTs

1 in²

30 cm

3 cm

6.5 cm

30 cm

5 cm

Tag Support

PMT Holder

Target

Propagation Medium

Upper Cosmic Tag

Lower Cosmic Tag

1 in²

30 cm

3 cm

6.5 cm

30 cm

5 cm

Tag Support

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Upper Cosmic Tag

Lower Cosmic Tag

1 in²

30 cm

3 cm

6.5 cm

30 cm

5 cm

Tag Support

PMT Holder

Target

Propagation Medium

Upper Cosmic Tag

Lower Cosmic Tag

1 in²

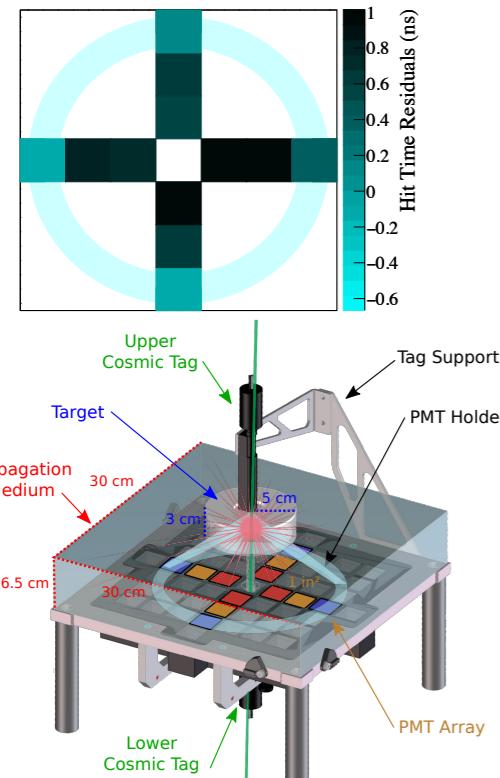
30 cm

3 cm

Builds on core  
(Wb)LS  
development at  
BNL (Yeh et al.)

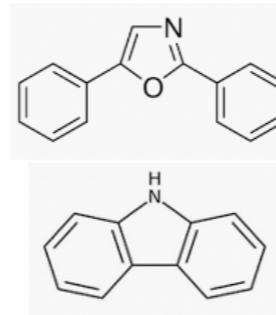
# Detector development

## CHESS detector: LBNL



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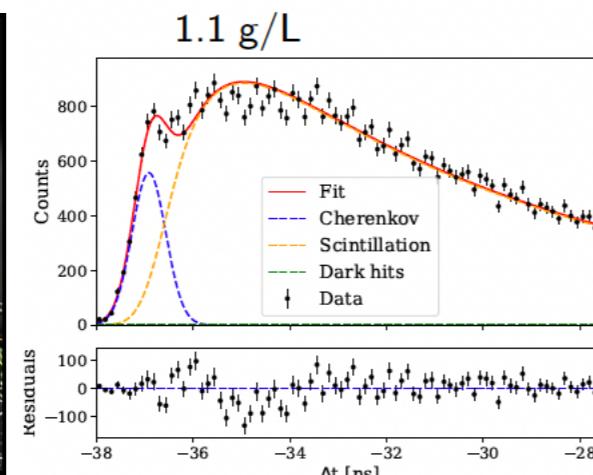
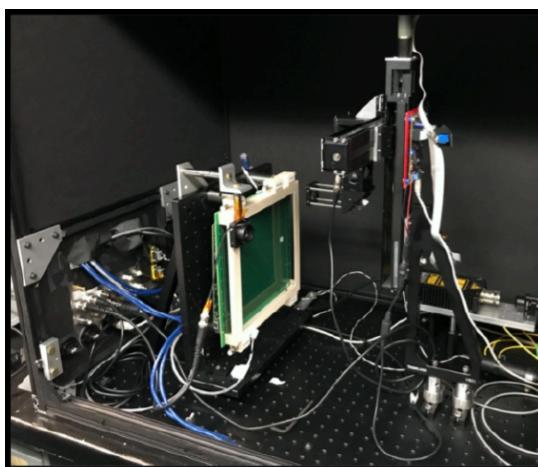
- Example:  
slowing down decay time
- Standard PPO → 2ns
  - New carbazole → 15ns



## Scattering & attenuation: UC Davis, UC Berkeley+LLNL

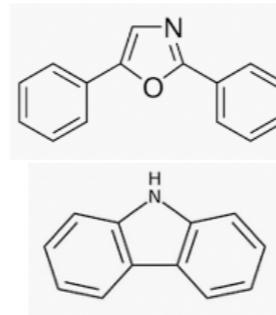


## LAPPDs: ANNIE, CHESS

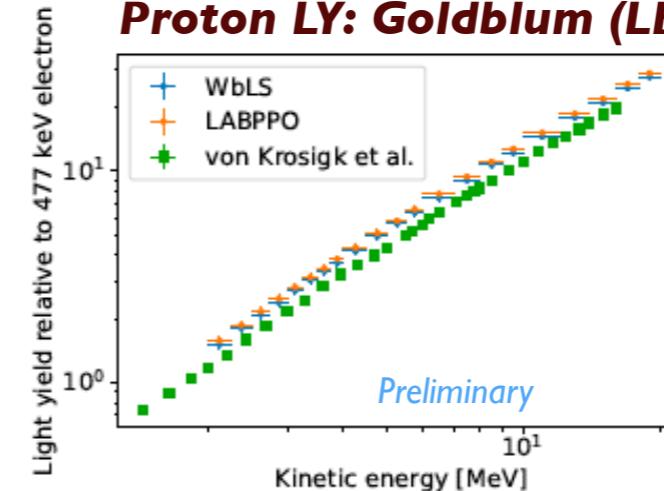


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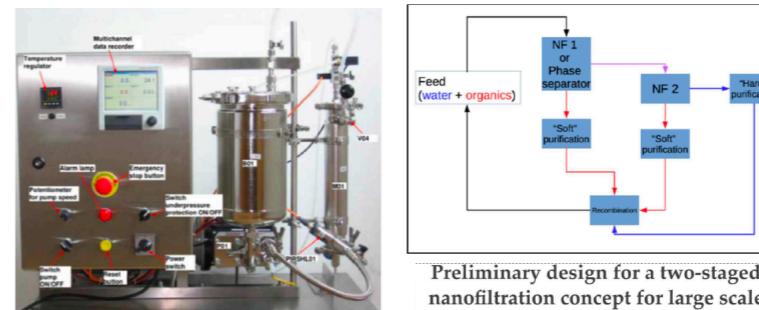
## Proton LY: Goldblum (LBNL)



Extensive international effort in Germany (Mainz, Munich), UK, China

Additional work on: slow LS, alternative fluors, alternative surfactants

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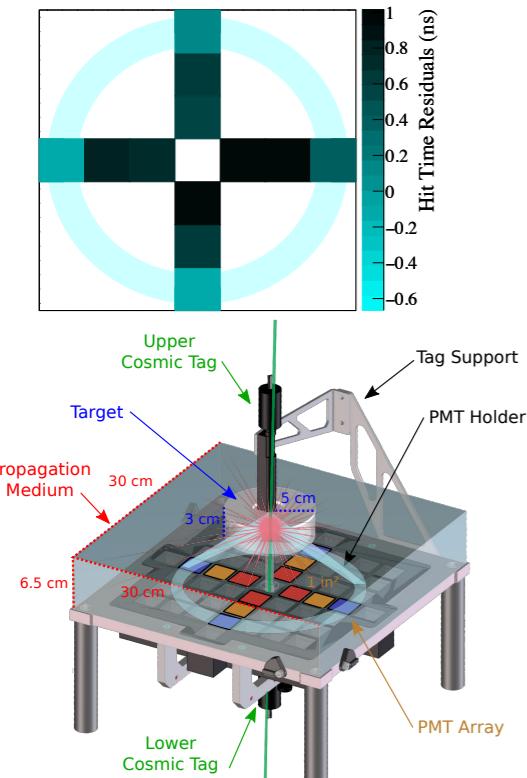
Preliminary design for a two-staged nanofiltration concept for large scale WbLS purification.

[Phys. Rev. C 95 055801 \(2017\); Eur. Phys. Jour. C 80 867 \(2020\); Mat. Adv. I 71 \(2020\); Eur. Phys. Jour. C 82 169 \(2022\); NIMA 947, 162604 \(2019\); arXiv:1902.06912; JINST13 P07005 \(2018\); JINST9 P06012 \(2014\); NIMA 943 162420 \(2019\); Eur. Phys. Jour. C 77 811 \(2017\); arXiv:1908.03564; arXiv:1502.01132; arXiv:1707.08222; NIMA 972 164106 \(2020\); Astropart. Phys. 109 33 \(2019\); NIMA 852 15 \(2017\); NIMA 712 162 \(2013\); Phys. Rev. D 97 052006 \(2018\); JINST14 I \(2019\); Phys. Rev. D 101 072002 \(2020\); arXiv:2006.00173](#)

Builds on core  
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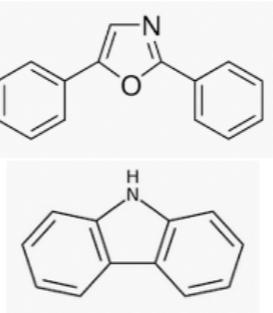
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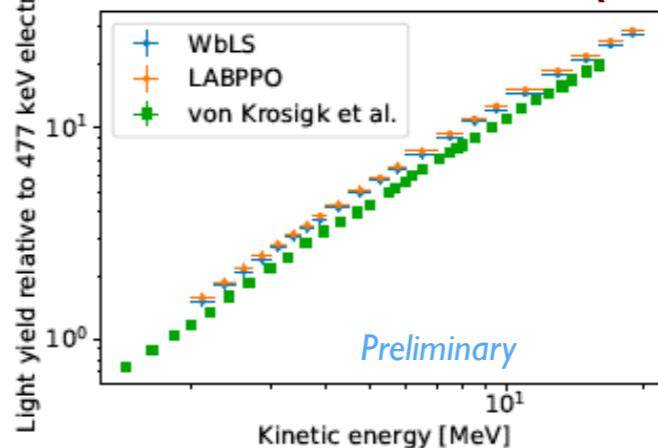
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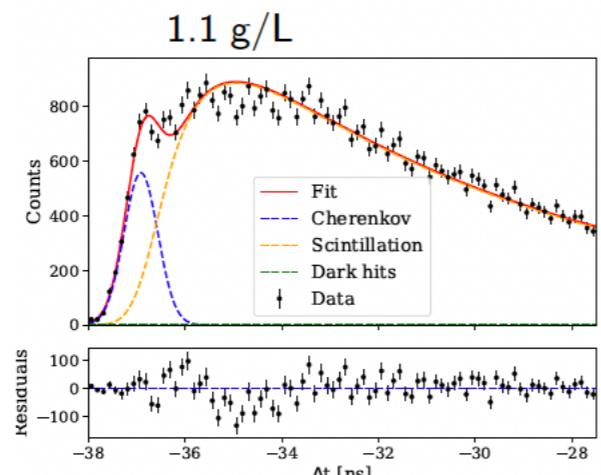
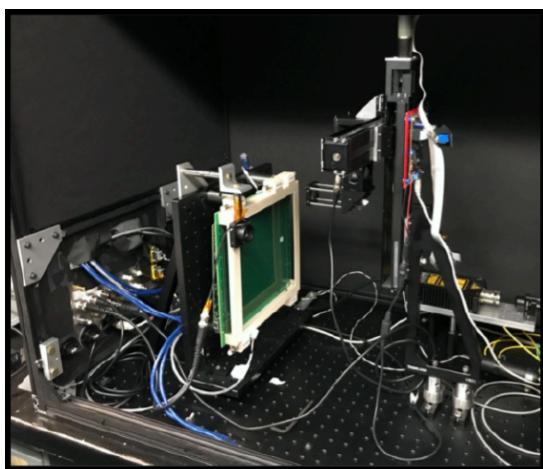
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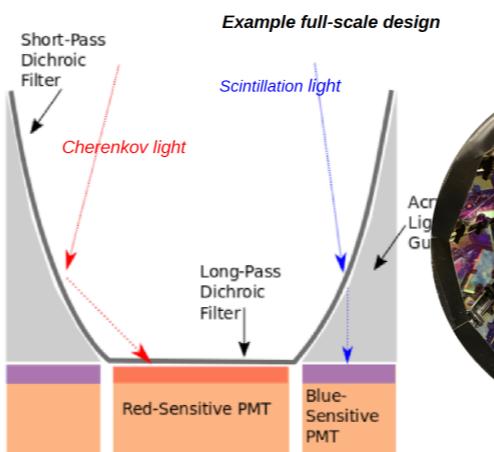
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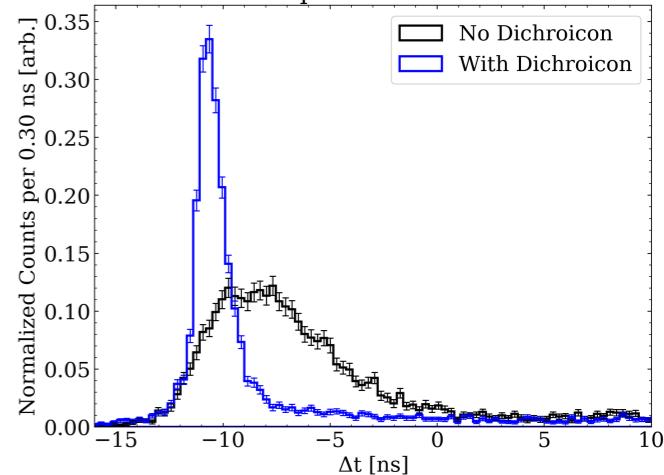
## Dichroicon: Penn, CHESS



T. Kaptanoglu, Nucl. Instrum. Meth. A889 (2018) 69-77  
T. Kaptanoglu, M. Luo, J. Klein, JINST 14 no. 05 T05001 (2019)

Phys. Rev. C 95 055801 (2017); Eur. Phys. Jour. C 80 867 (2020); Mat. Adv. I 71 (2020); Eur. Phys. Jour. C 82 169 (2022); NIMA 947, 162604 (2019); arXiv:1902.06912; JINST13 P07005 (2018); JINST9 P06012 (2014); NIMA 943 162420 (2019); Eur. Phys. Jour. C 77 811 (2017); arXiv:1908.03564; arXiv:1502.01132; arXiv:1707.08222; NIMA 972 164106 (2020); Astropart. Phys. 109 33 (2019); NIMA 852 15 (2017); NIMA 712 162 (2013); Phys. Rev. D 97 052006 (2018); JINST14 1 (2019); Phys. Rev. D 101 072002 (2020); arXiv:2006.00173

Atmospheric Muons Incident on LABPPO Target Aperture PMT



# Prototypes

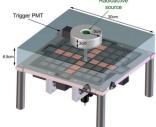
## Data-driven demonstration

- ✓ Demonstrate *Cher+scint* reconstruction
- ✓ Demonstrate *Cher+scint* particle ID
- ✓ Enable broad, world-leading physics & nonproliferation program

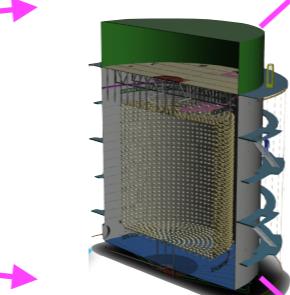
**Demonstration**  
of next-generation  
detector capabilities

Develop **R&D infrastructure** —  
testbed for future  
programs

Broad R&D  
program



Ch/S separation,  
microphysical  
parameter  
measurements



ktonne-scale  
demonstration of  
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monitoring using  
antineutrinos



THEIA  
10s of ktonne  
LBNL-led  
international  
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(US, UK, Germany,  
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Broad physics  
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# Prototypes

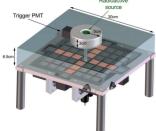
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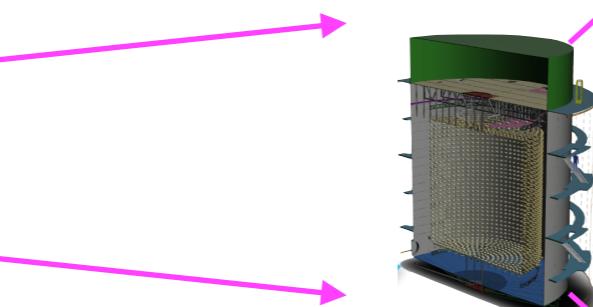
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Ch/S separation,  
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Critical ton — 10s  
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*Importance for  
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+ ONP, NNSA  
& international partners*



ktonne-scale  
demonstration of  
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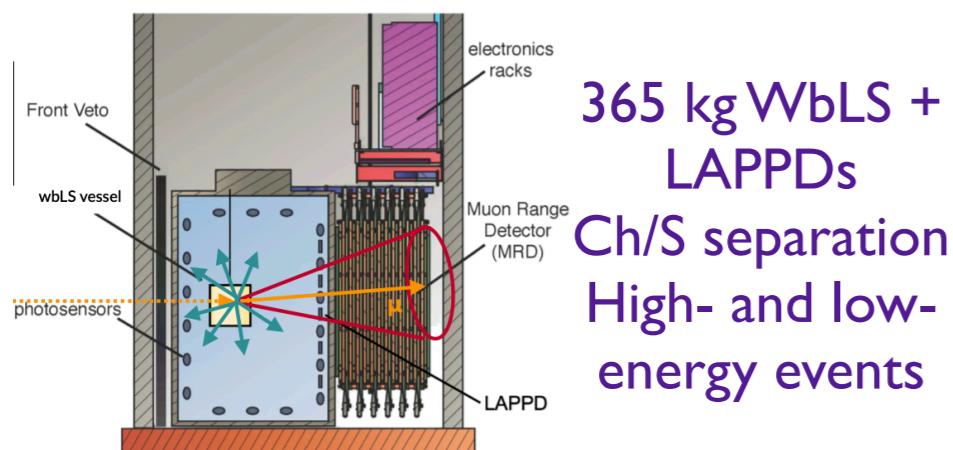


THEIA  
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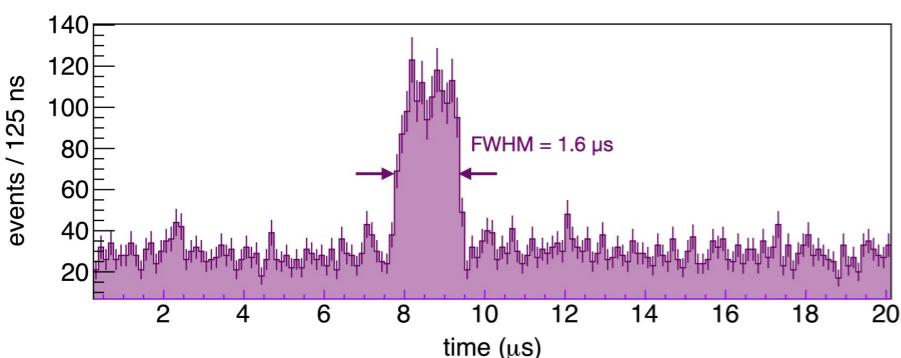
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# ANNIE, NuDot, Eos, BNL

## ANNIE:WbLS in a $\nu$ beam

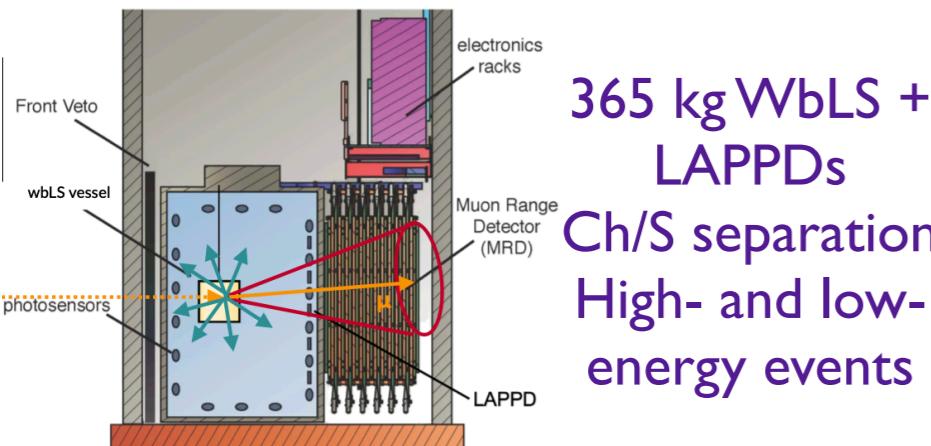


**First neutrinos detected with an LAPPD!**



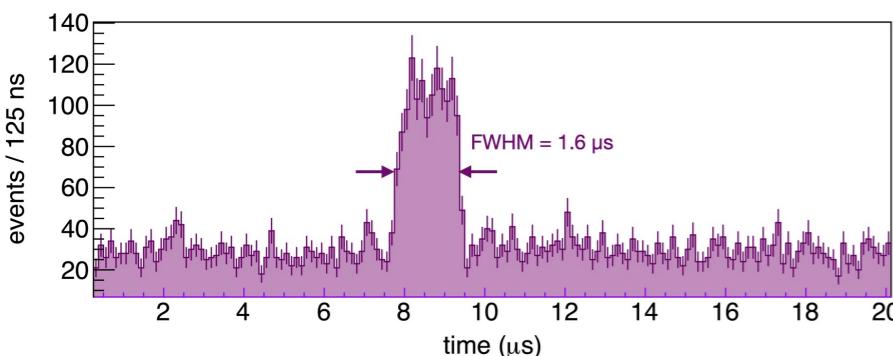
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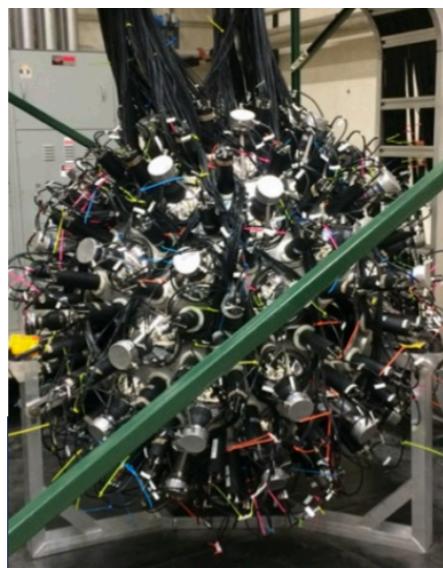


365 kg WbLS +  
LAPPDs  
Ch/S separation  
High- and low-  
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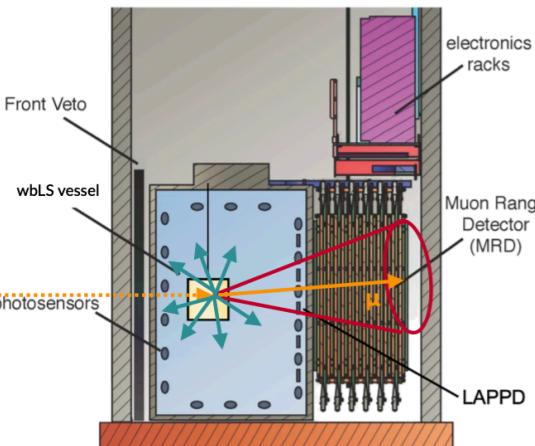


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Ton scale  
Highly  
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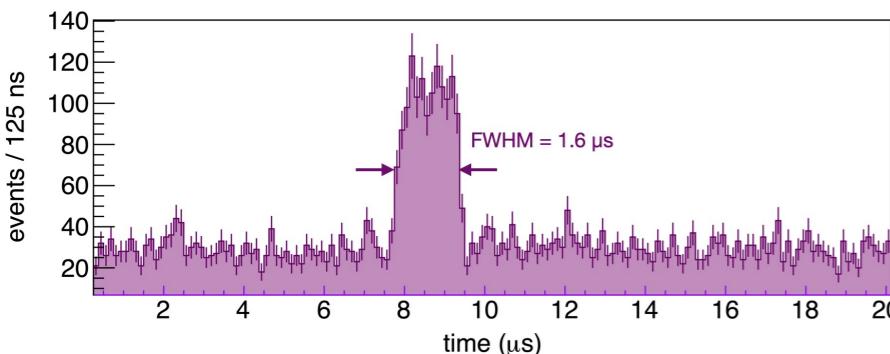
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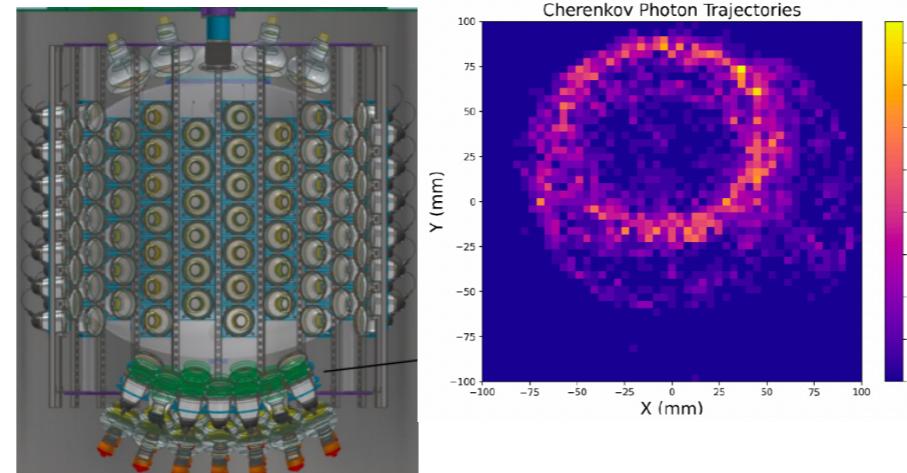
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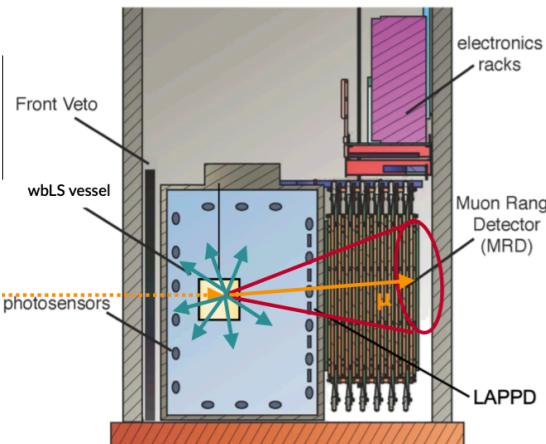


- 4-ton target mass
- 200 8" PMTs
- Dichroicon deployment for spectral sorting
- Vertex, energy, direction, PID



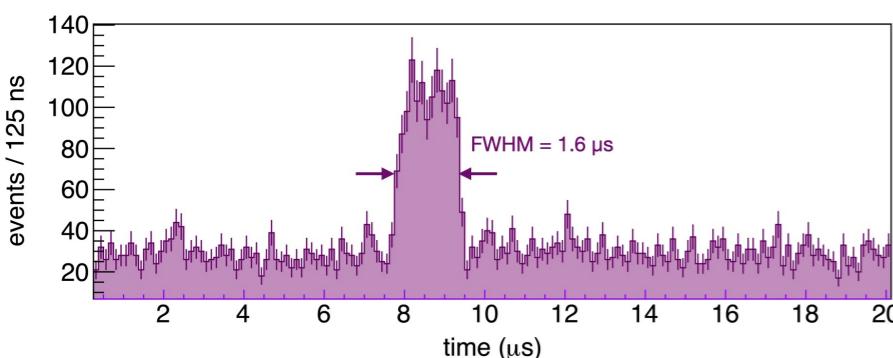
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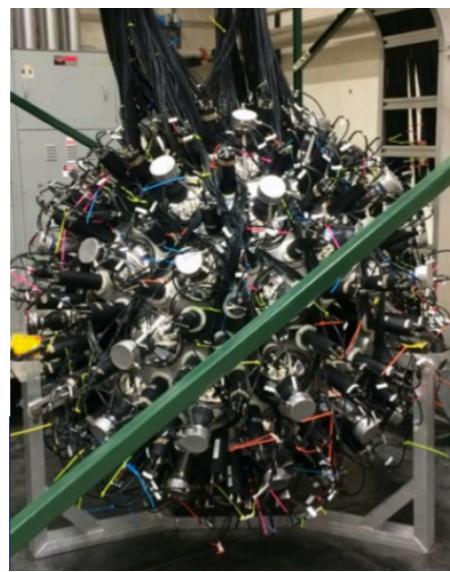


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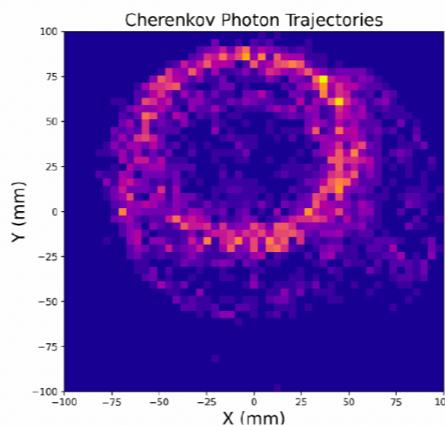
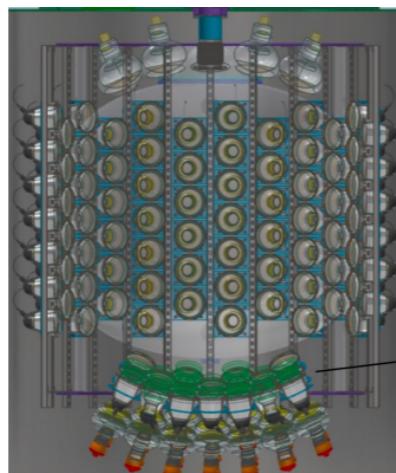
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Ton scale  
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## Eos: Low-energy event reconstruction and model validation



## BNL: 1- and 30-ton

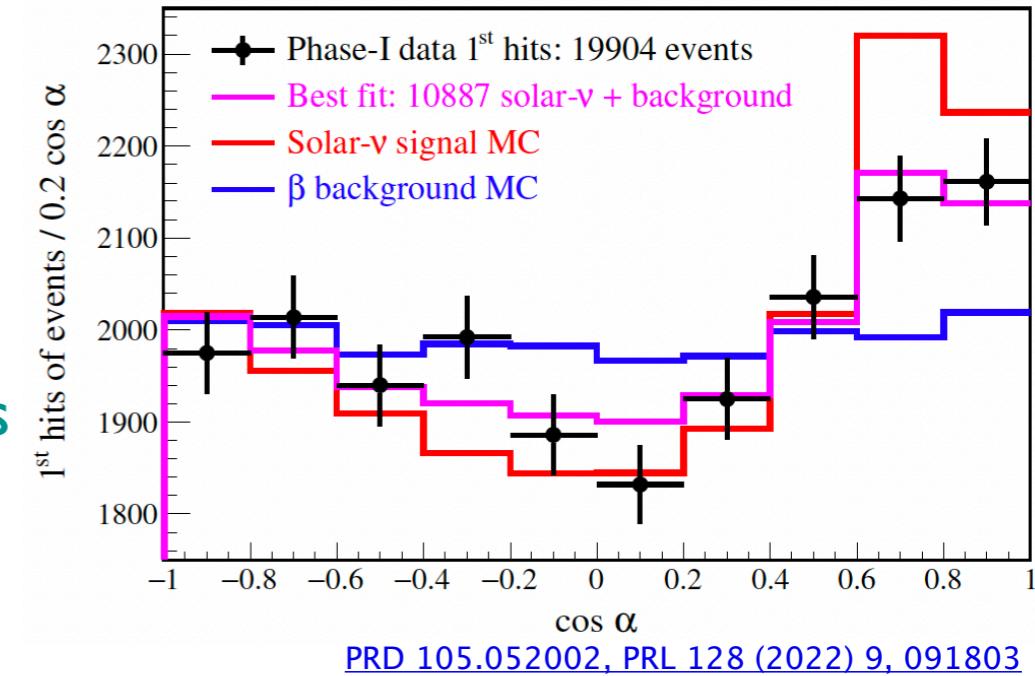


- First ton-scale deployment
- Optical transparency in an operating detector
- Optical stability over time
- Recirculation of WbLS (nanofiltration)



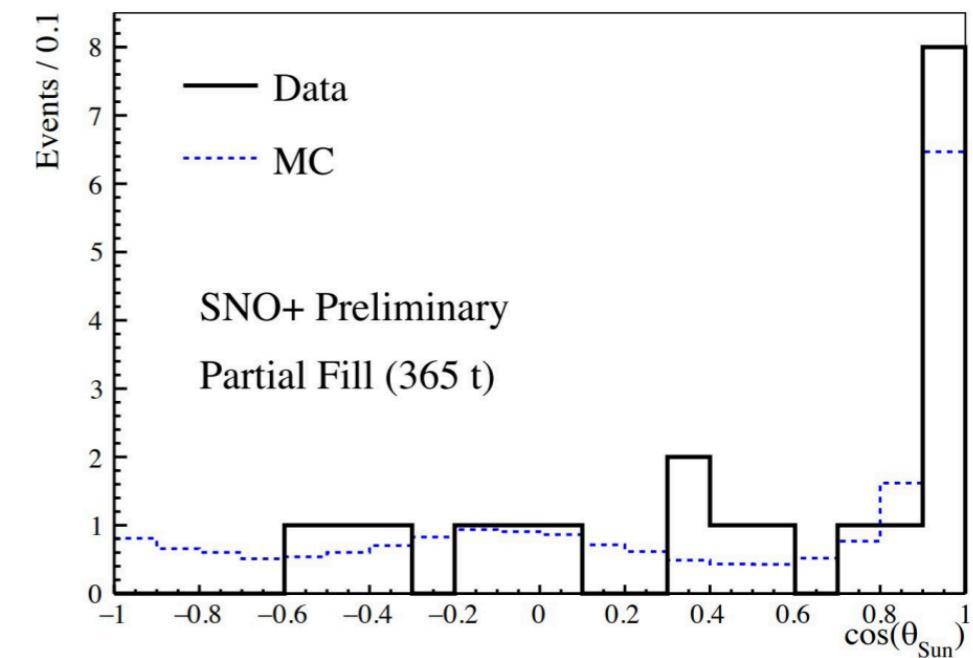
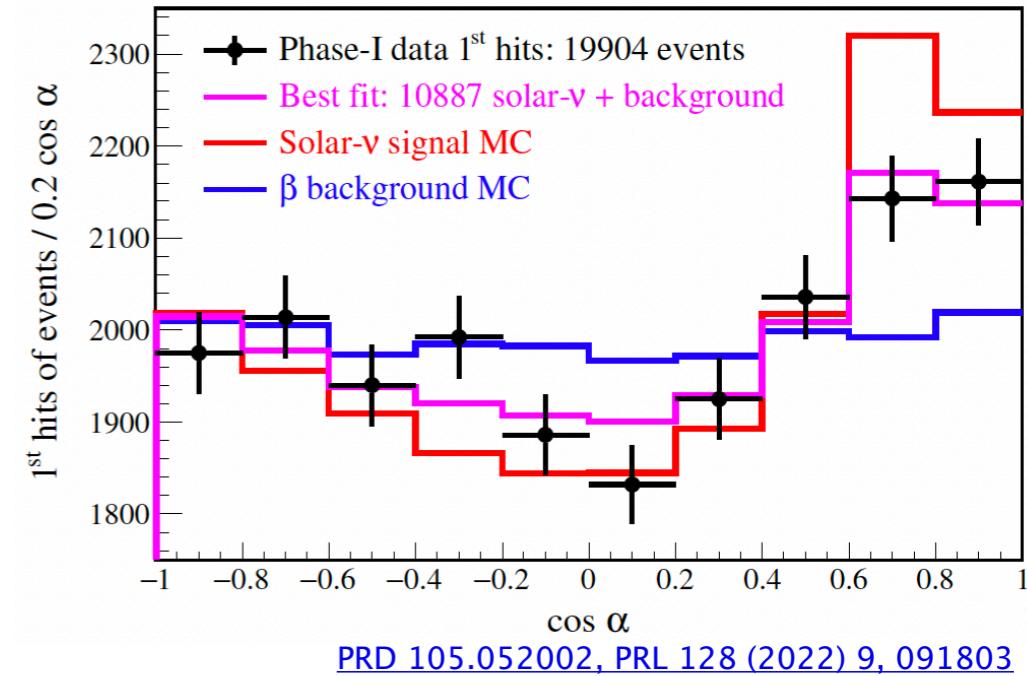
# Full-scale demonstrations

- Integrated directionality at Borexino:
  - consider earliest photons in the event
  - take angle between early photons and solar direction
  - $6\sigma$  angular excess caused by Cherenkov photons
- Measurement of primarily  ${}^7\text{Be}$   $\nu$  demonstrates first directional detection of sub-MeV neutrinos



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- Measurement of primarily  $7\text{Be}$   $\nu$  demonstrates first directional detection of sub-MeV neutrinos
- Event-level directionality at SNO+:
  - Partial-filled detector (365 t LAB + 0.6 g/L PPO)
  - ToF and angular reconstruction
  - Demonstration with  $> 5\text{MeV}$  8B  $\nu$
- First event-by-event demonstration of directional reconstruction for 8B solar  $\nu$  in slow LS



SNO+, L. Lebanowski APS talk

# Theia Physics Program



# Physics Program

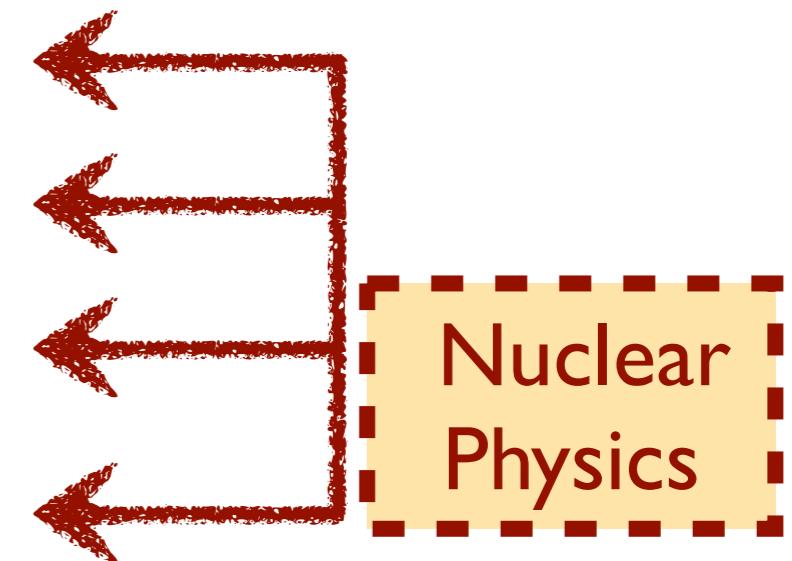
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1. Neutrinoless double beta decay
2. Solar neutrinos (solar metallicity, luminosity)
3. Geo-neutrinos (& reactor neutrinos)
4. Supernova burst neutrinos & DSNB
5. Source-based sterile searches
6. Nucleon decay
7. Long-baseline physics (mass hierarchy, CP violation)

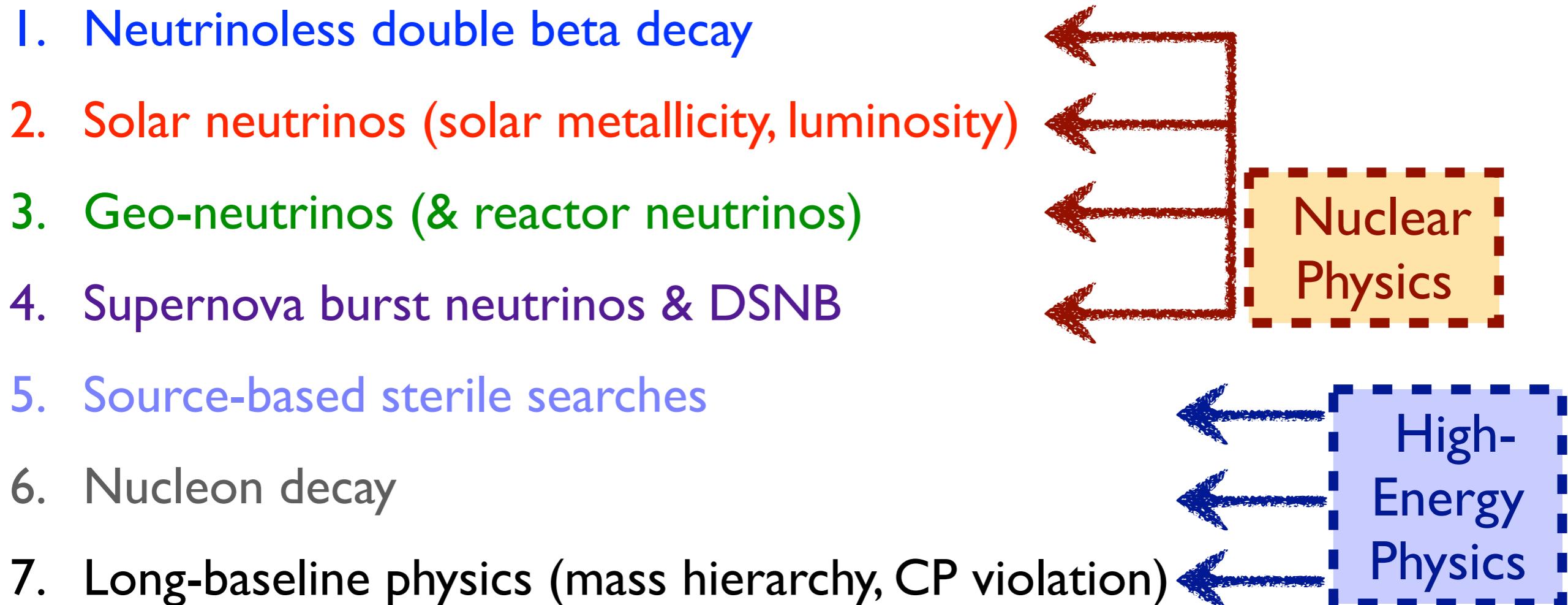
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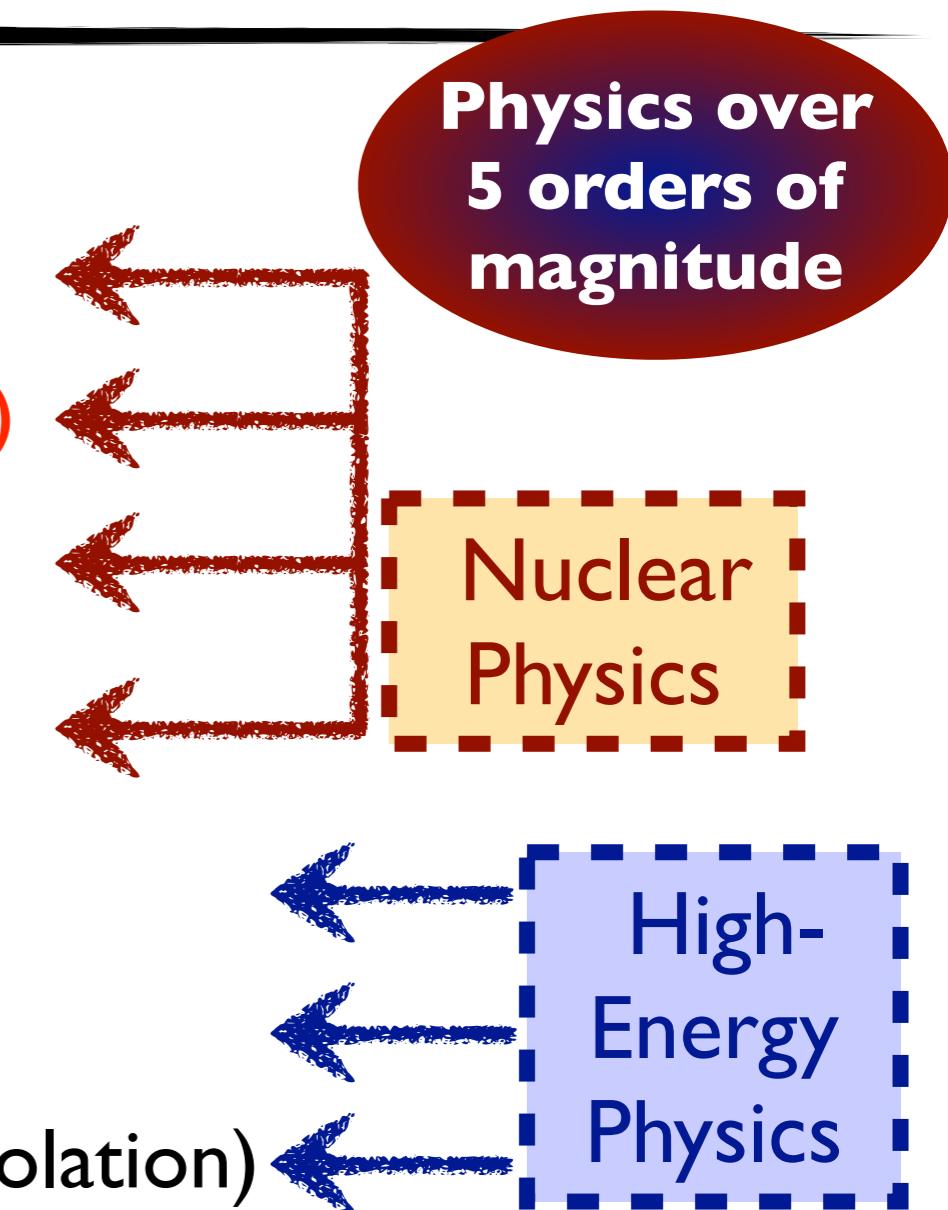


# Physics Program



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# Physics Program



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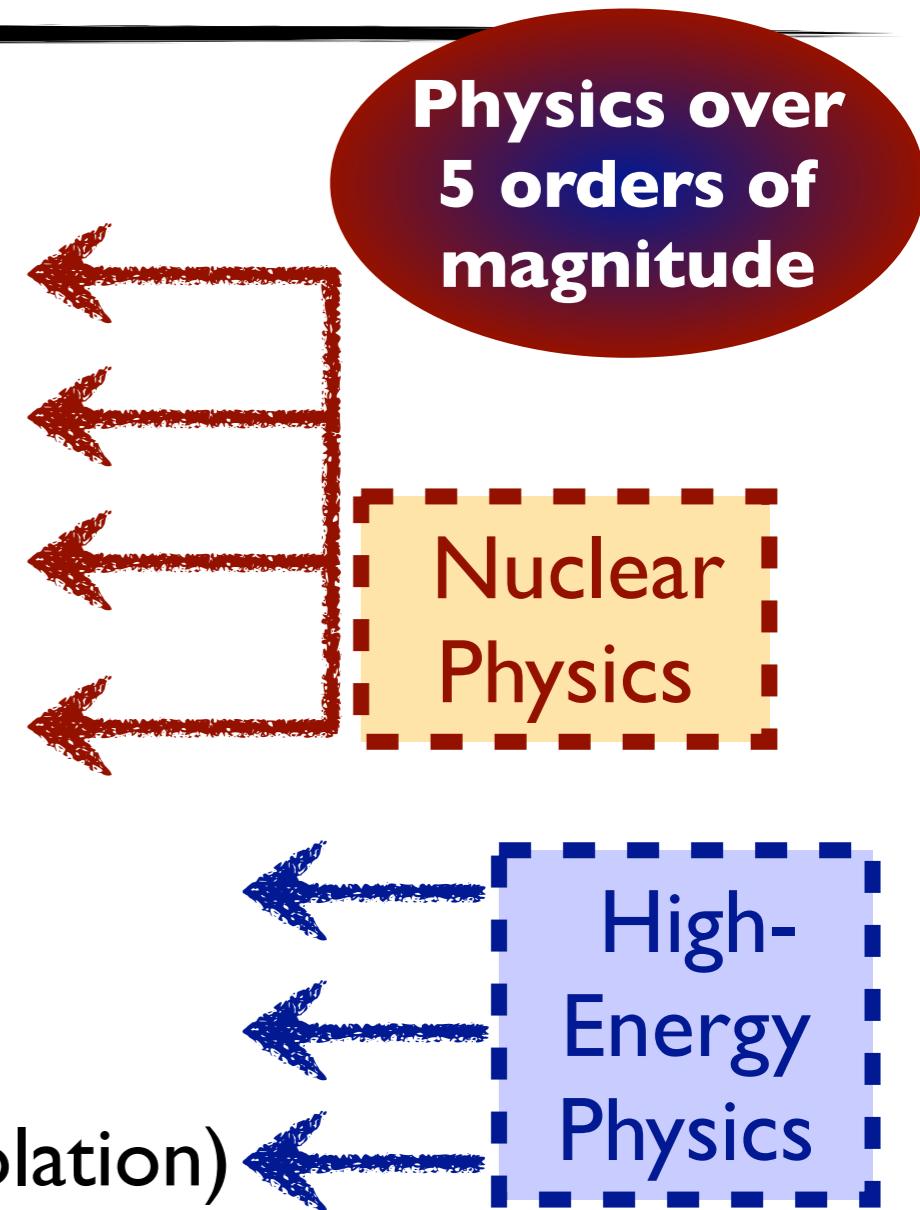
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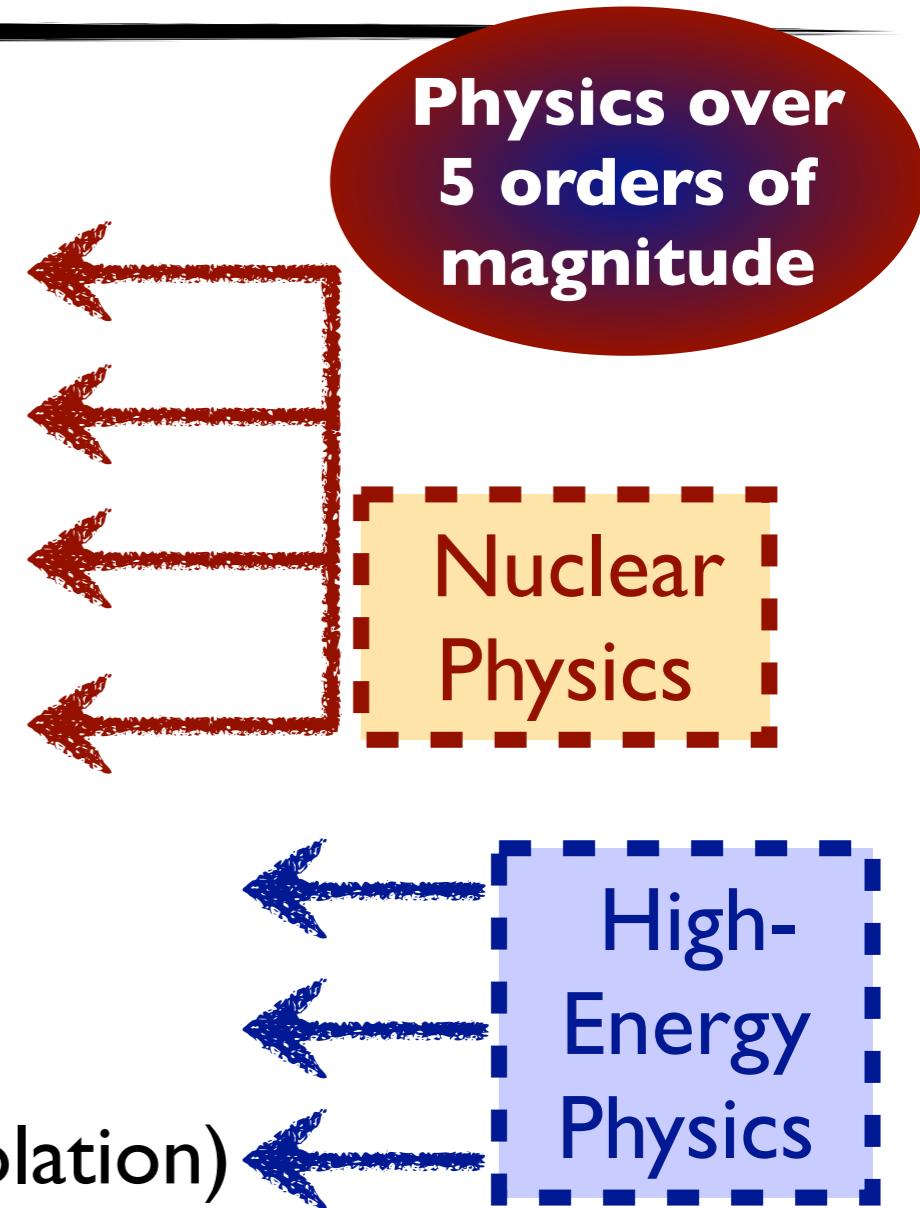


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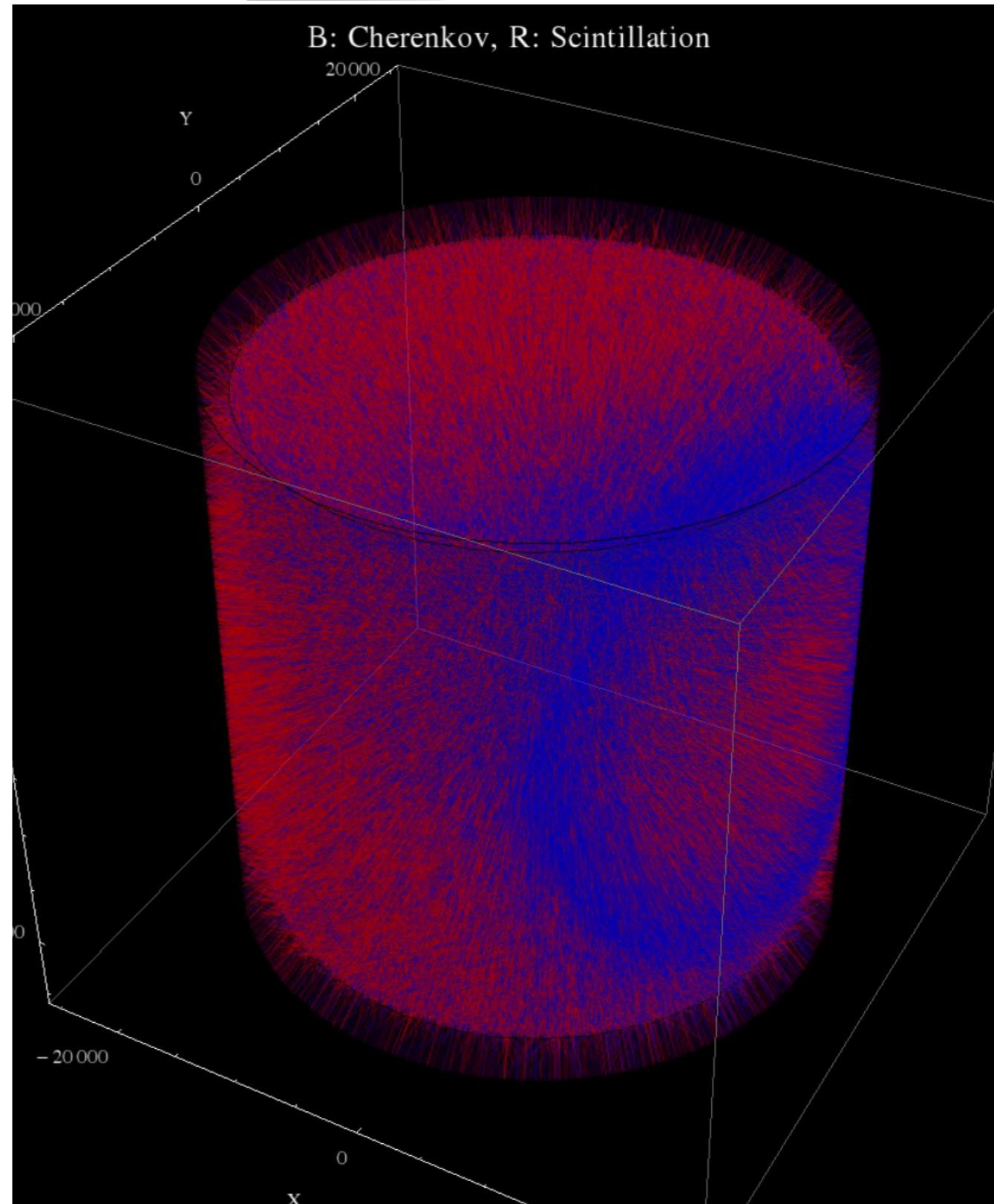
**Matter-dominated universe**

# (Phased) Physics program

Primary physics goal	Reach	Exposure/assumptions
<b>Long-baseline oscillations</b>	$>5\sigma$ for 30% of $\delta_{CP}$	524kt-MW-year
<b>Nucleon decay</b> $p \rightarrow \bar{\nu} K^+$	$T > 3.8 \times 10^{34}$ year	800 kt-year
<b>Supernova burst</b>	$< 1(2)^\circ$ pointing 20K(5K) events	100(25)kt, 10kpc SN
<b>Diffuse Supernova Neutrino</b>	$5\sigma$	125kt-year
<b>CNO neutrinos</b>	$< 5(10)\%$	300(62.5)kt-year
<b>Geoneutrinos</b>	$< 7\%$	25 kt-year
<b><math>0\nu\nu\beta</math></b>	$T_{1/2} > 1.1 \times 10^{28}$ year (90% C.L.)	800 kt-year (Multi-tonne loaded LS in suspended vessel search )

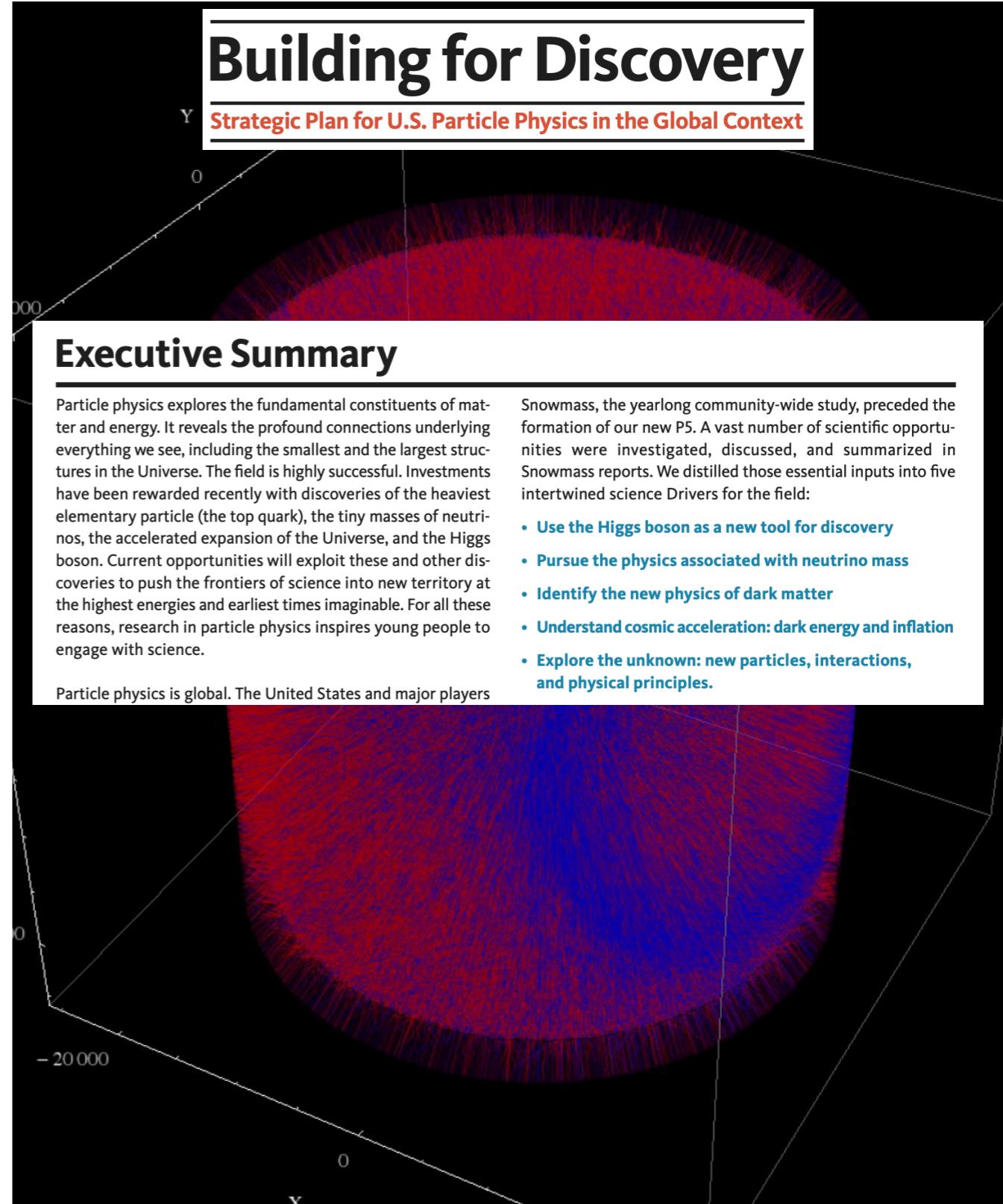
# Take-home message

- A hybrid detector module would add to the LBL program at DUNE and bring a broad program of additional physics
- Major technological developments have been achieved since last Snowmass
- Results from existing large detectors demonstrate the feasibility of this concept
- Prototypes underway will demonstrate the full range of capabilities
- Brings a new community of physicists to the facility and the program
- Timely to pursue serious engineering design studies to evolve the concept
- Depends on an open process for determining the future of Module 4



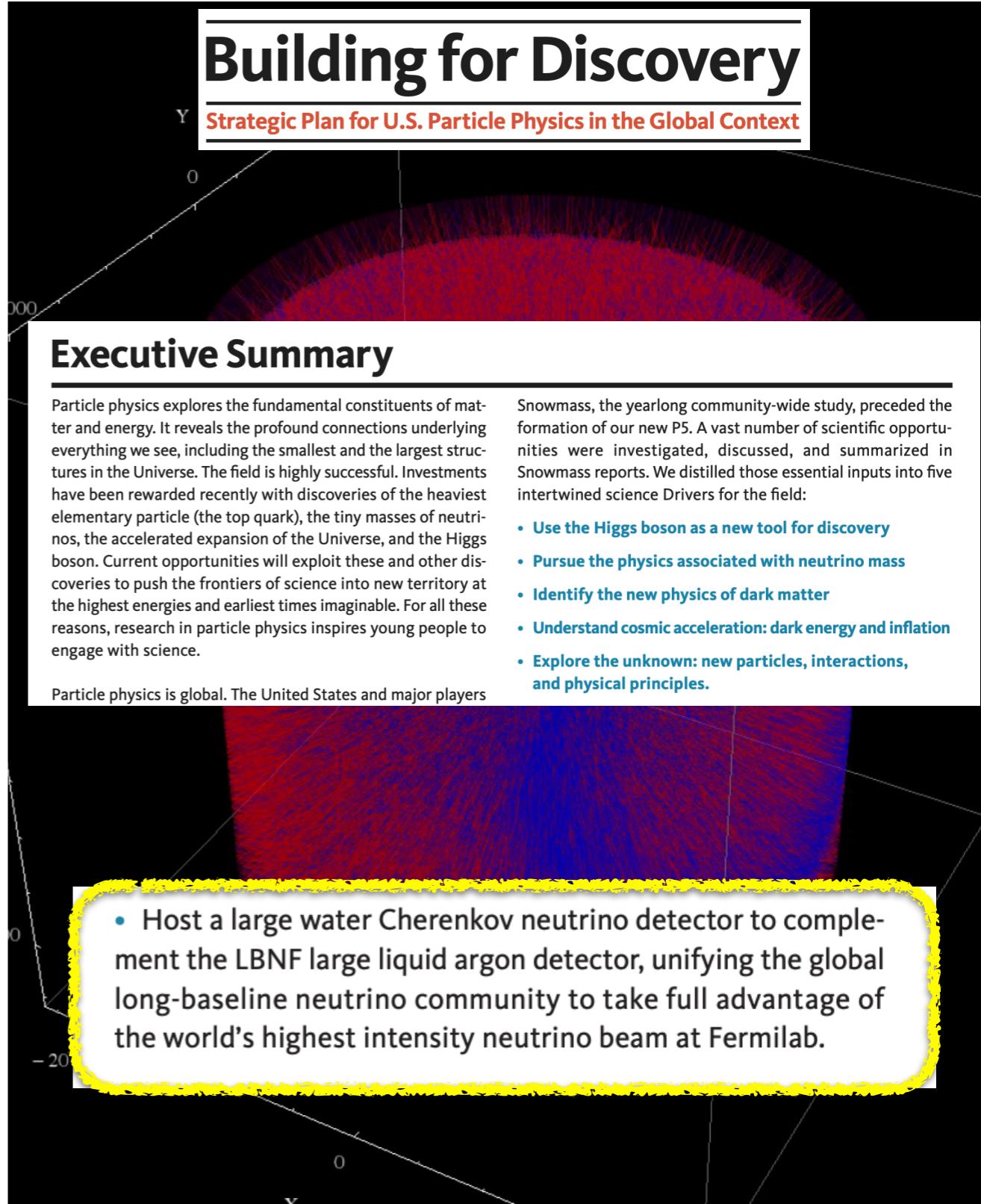
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# Thank you



This material is based upon work supported by the U.S. Department of Energy, Office of Science, Office of High Energy Physics, under Award Number DESC0018974. Work conducted at Lawrence Berkeley National Laboratory was performed under the auspices of the U.S. Department of Energy under Contract DEAC02-05CH11231. The project was funded by the U.S. Department of Energy, National Nuclear Security Administration, Office of Defense Nuclear Nonproliferation Research and Development (DNN R&D).

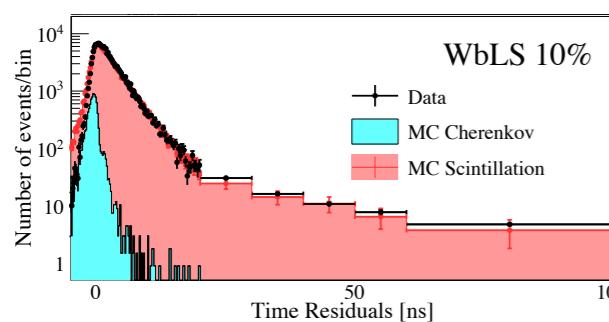
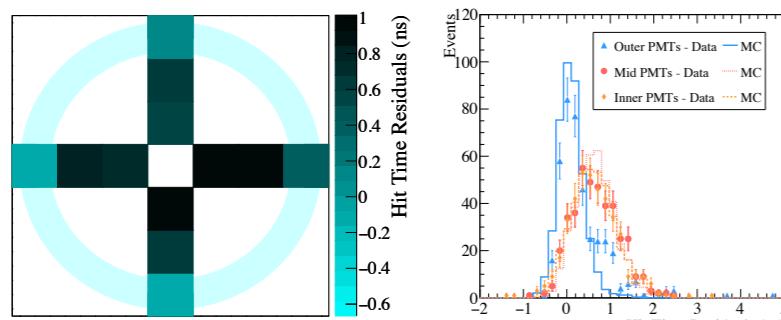
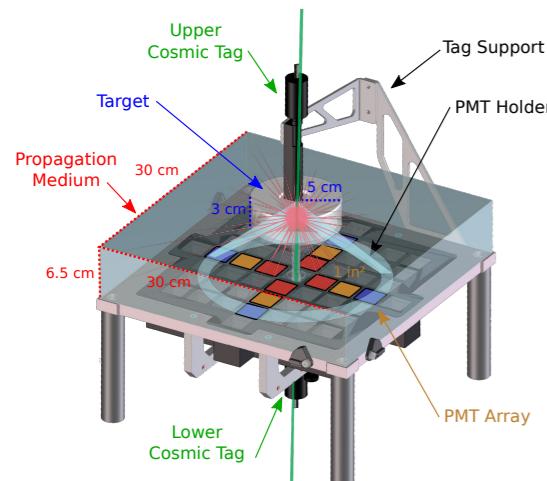


- 
- Backup
-

# (Wb)LS development

Primary development at BNL (Yeh et al)

## CHESS detector, LBNL

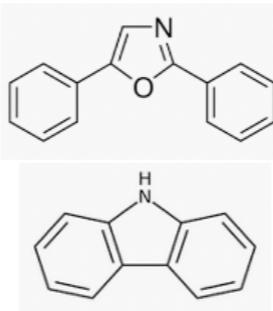


Phys. Rev. C **95** 055801 (2017)  
Eur. Phys. Jour. C **77** 811 (2017)  
Eur. Phys. Jour. C **80** 867 (2020)

## Engineering WbLS properties, LBNL

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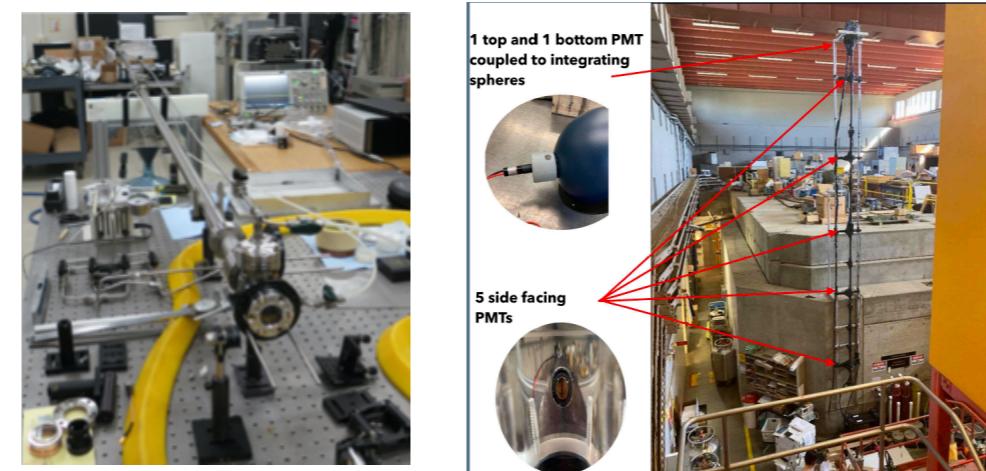
*Mat. Adv.* **1** 71 (2020)



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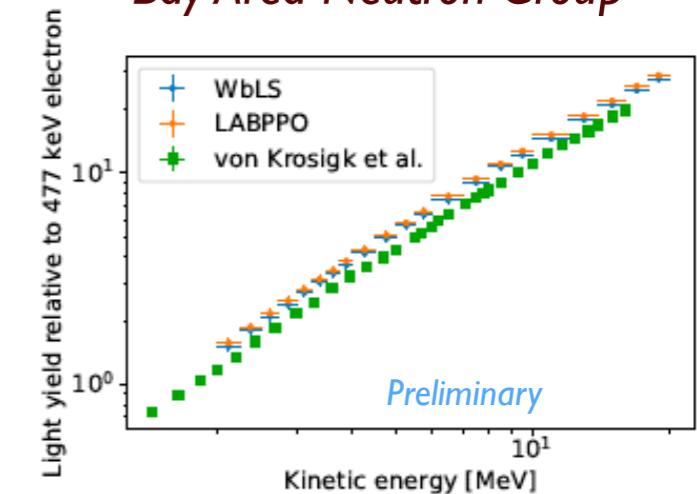
Additional work on: slow LS, alternative fluors, alternative surfactants

## Scattering & attenuation, UC Davis, UC Berkeley+LLNL

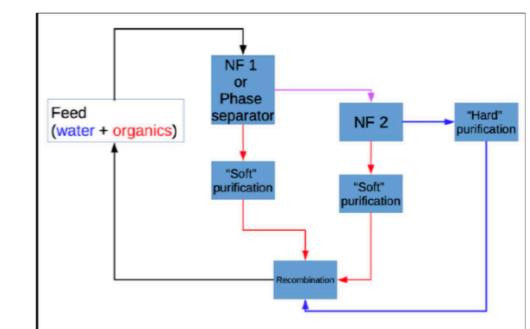
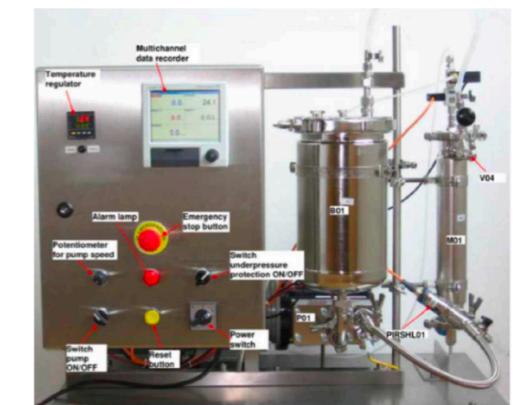


## Proton LY, LBNL

### Bay Area Neutron Group



## Nanofiltration, UC Davis



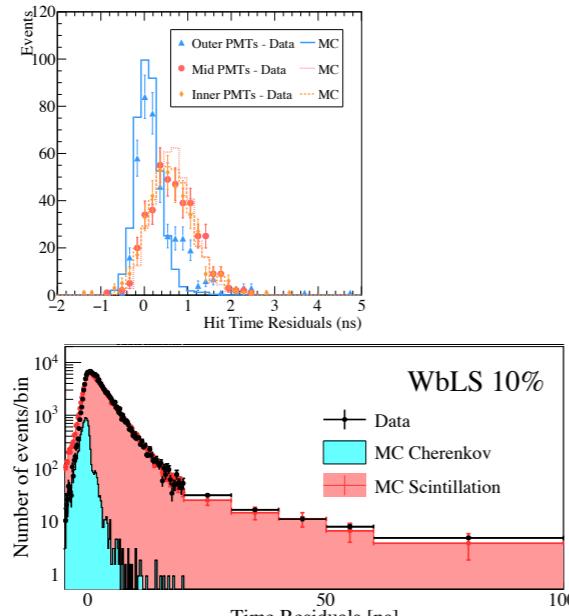
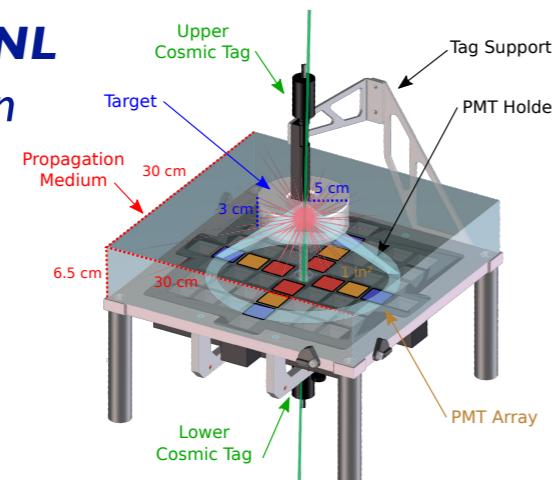
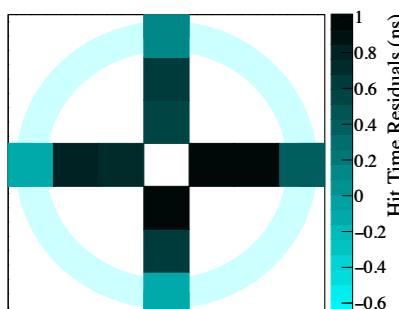
Preliminary design for a two-staged nanofiltration concept for large scale WbLS purification.

# (Wb)LS development

## CHESS detector, LBNL

Microphysical extraction of intrinsic light yield and time profile; demonstration of CIS separation via timing and topology

Phys. Rev. C **95** 055801 (2017)  
 Eur. Phys. Jour. C **77** 811 (2017)  
 Eur. Phys. Jour. C **80** 867 (2020)



## Engineering WbLS properties, LBNL

Example: slowing down decay time

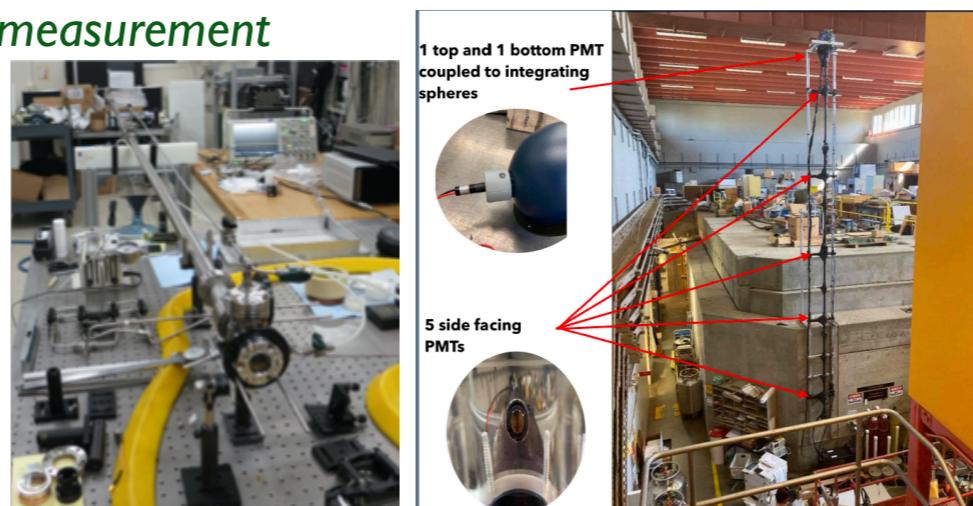
- Standard PPO → O=C1C=CC=CC=C1c2ccccc2 2ns
- New carbazole → c1ccc2c(c1)nc3ccccc23 15ns

+ optical & x-ray excitation for pure scintillation characterisation; independent check of CHESS MeV-scale physics events

Mat. Adv. I **71** (2020)

## Scattering & attenuation, UC Davis, UC Berkeley+LLNL

Long-arm (7.5m) scattering and attenuation measurement device; attenuation measured as a function of wavelength; side ports for scattering measurement



Extensive international effort in Germany (Mainz, Munich), UK, China

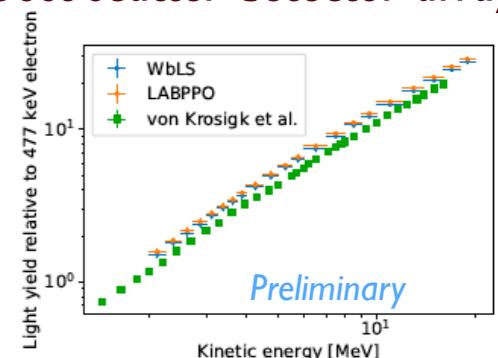
Additional work on: slow LS, alternative fluors, alternative surfactants



## Proton LY, LBNL

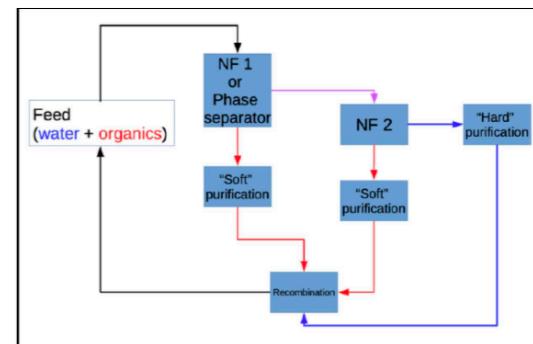
Bay Area Neutron Group

Double time-of-flight method, pulsed deuteron beam on Be breakup target; PID-capable post-scatter detector array



## Nanofiltration, UC Davis

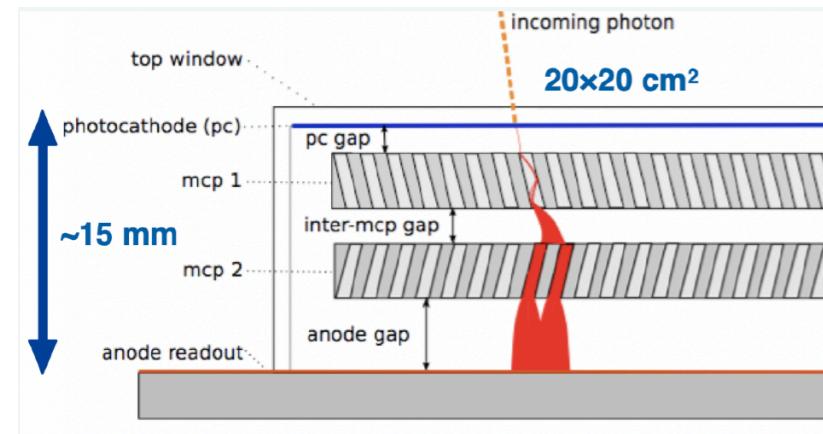
Membrane filtration process to separate water & ionic compounds from LS before deionisation. Critical to maintain optical transparency. Flow rate determined by detector size.



Preliminary design for a two-staged nanofiltration concept for large scale WbLS purification.

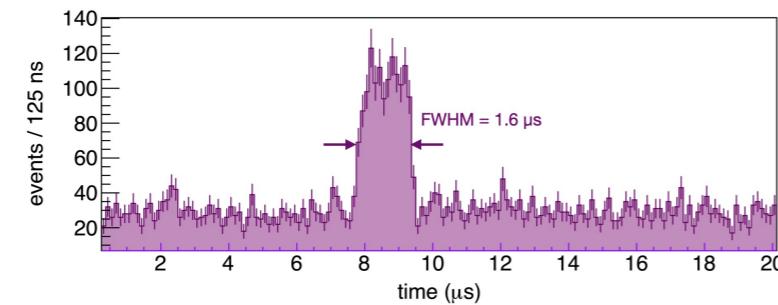
# LAPPDs

- Layered micro channel plates → sub-cm spatial and < 70ps time resolution for SPE

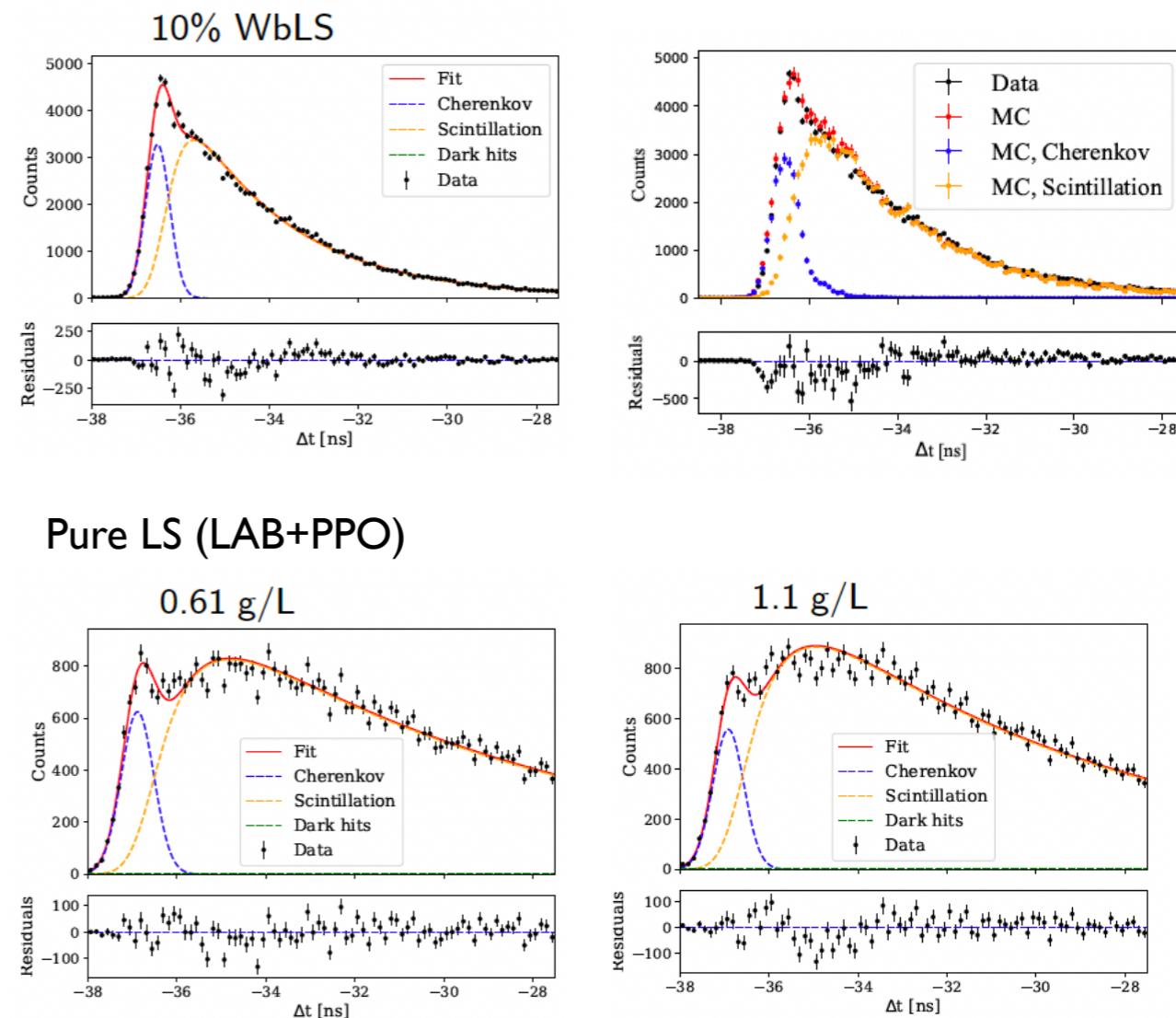


- LAPPD operating in ANNIE: world-first demonstration in running neutrino experiment (+4 planned for next beam year)

**First neutrinos detected with an LAPPD!**



- Deployment at CHESS with LS and WbLS demonstrate clear Cherenkov peak even in high LY scintillators



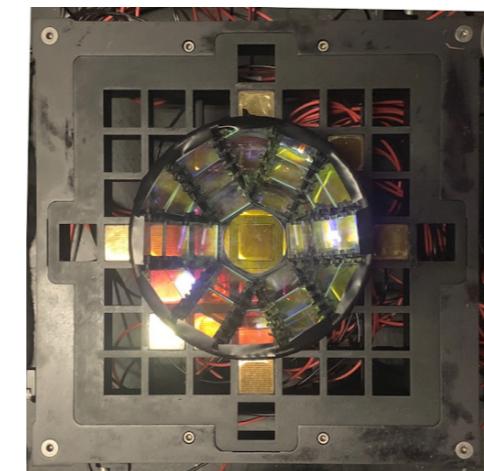
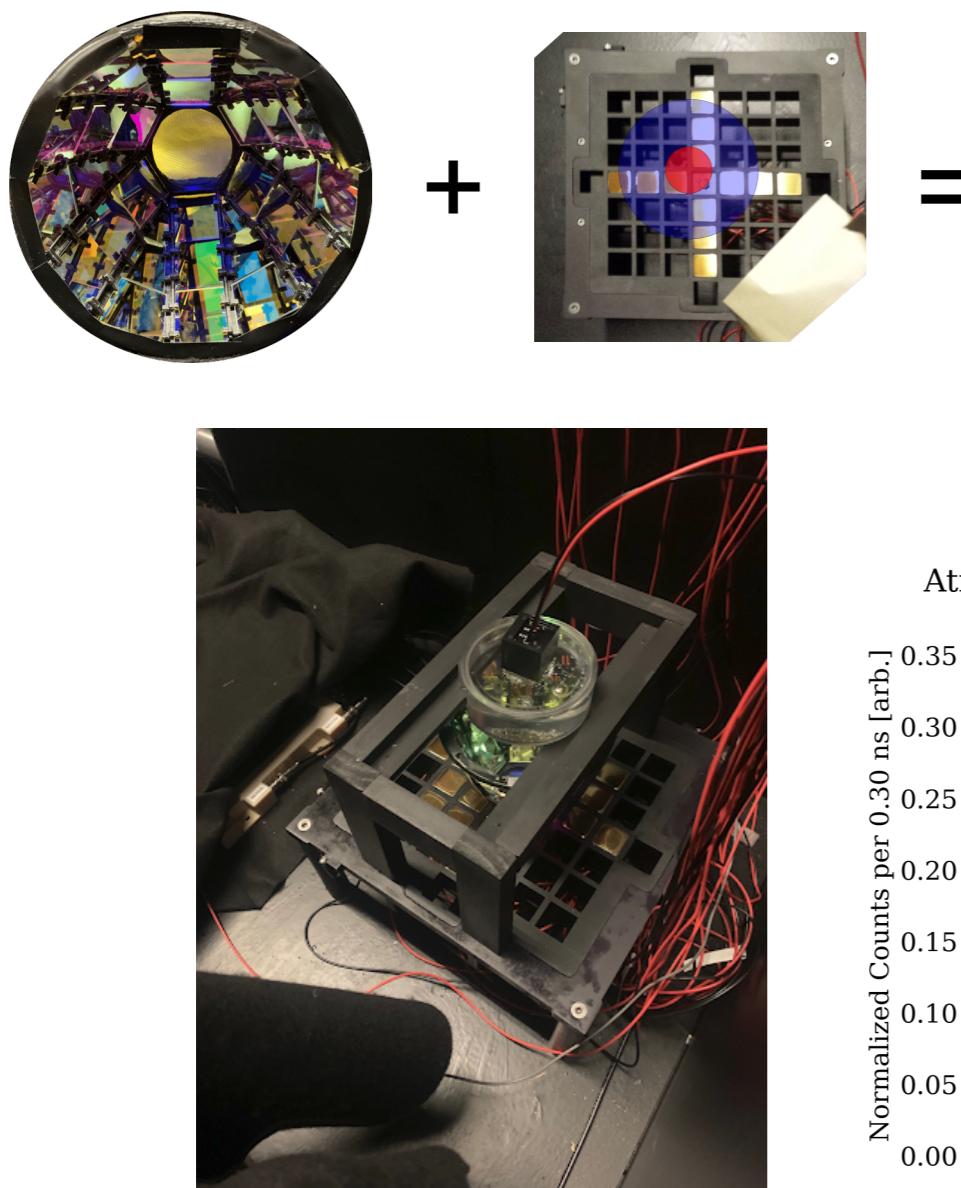
Eur. Phys. Jour. C 82 169 (2022)

# Dichroicons

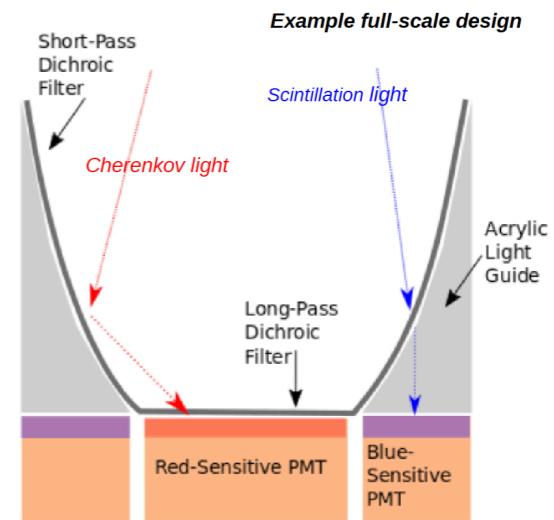
Concept developed by  
Penn team  
(Klein, Kaptanoglu et al.)

Dichroicons sort incident photons by wavelength  
Achieves spectral separation of Cherenkov and scintillation.  
Dichroicon deployment at CHESS demonstrates capabilities

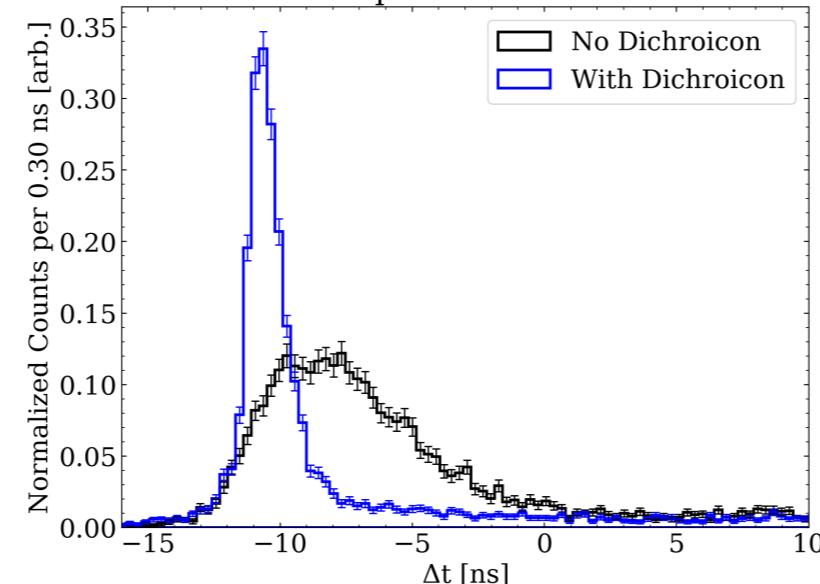
T. Kaptanoglu, Nucl. Instrum. Meth. A889 (2018) 69-77  
T. Kaptanoglu, M. Luo, J. Klein, JINST 14 no. 05 T05001 (2019)



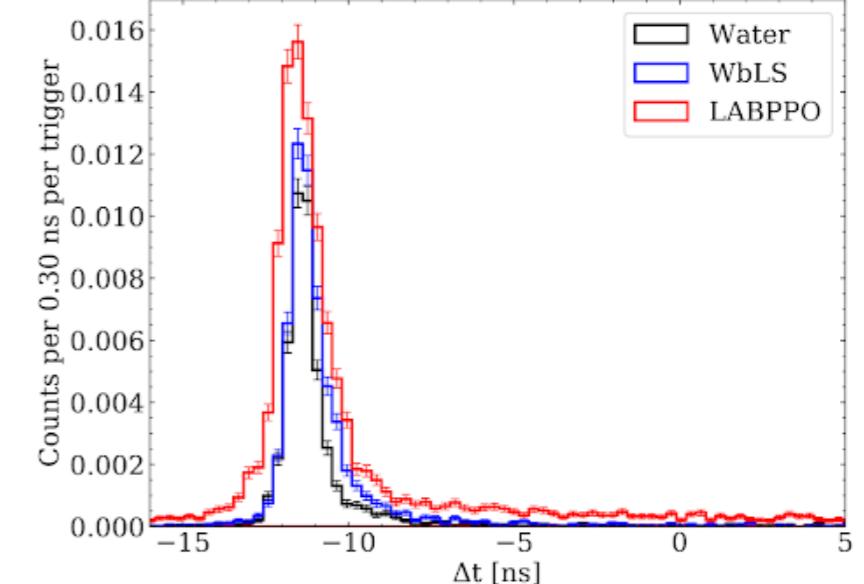
“Standard” design  
Red-focusing



Atmospheric Muons Incident on LABPPO Target Aperture PMT



Atmospheric Muons Incident on All Targets Standard Dichroicon



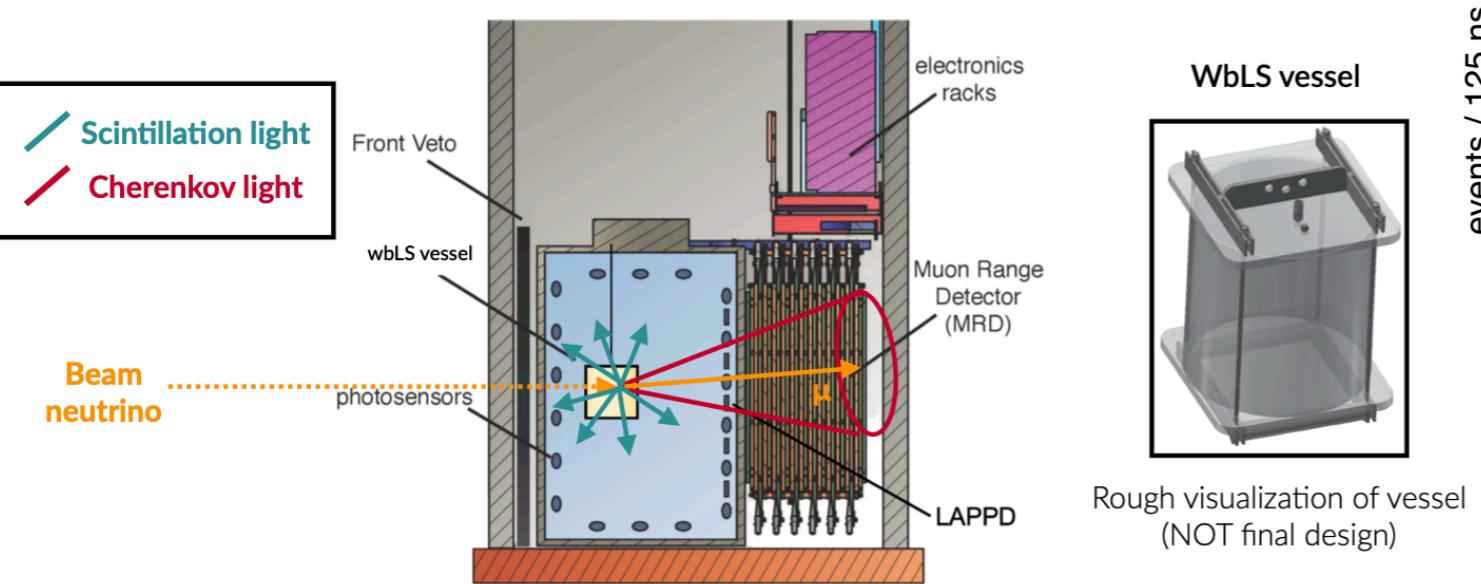
# WbLS @ ANNIE

Show feasibility of WbLS in a neutrino-beam

*Understanding neutrino-nucleus interactions,  
focusing on production and multiplicity of  
final-state neutrons*

Interest to Theia:

- First deployment of LAPPDs
- Deployment of 365 kg WbLS
- C/S separation in a large-scale experiment
- High- and low-energy event reconstruction
- Use of sub-Chr t/h scintillation for recoil p & hadron detection
- Neutron detection



## WbLS data taking:

Planned for summer/fall 2022

**Location:** Booster Neutrino Beam @ FermiLab (USA)

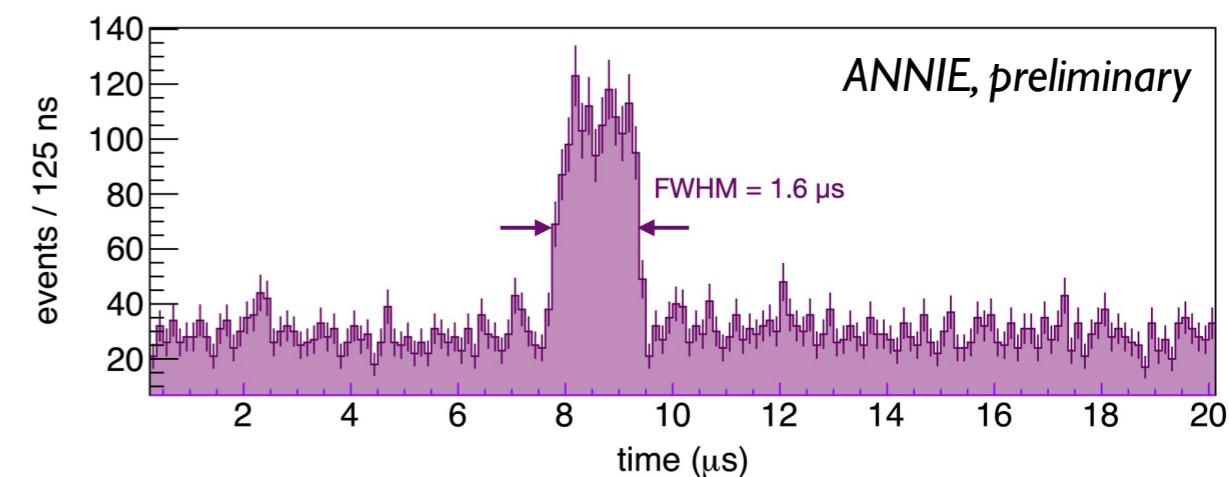
**Water-volume:** 26t

**WbLS volume:** 0.5t

**Interests:** neutrino-nucleus interactions



*First neutrinos detected with an LAPPD!*



A.R. Back et al JINST **15** P03011 (2020)  
M.J. Minot et al NIM A **787**, 78 (2015)

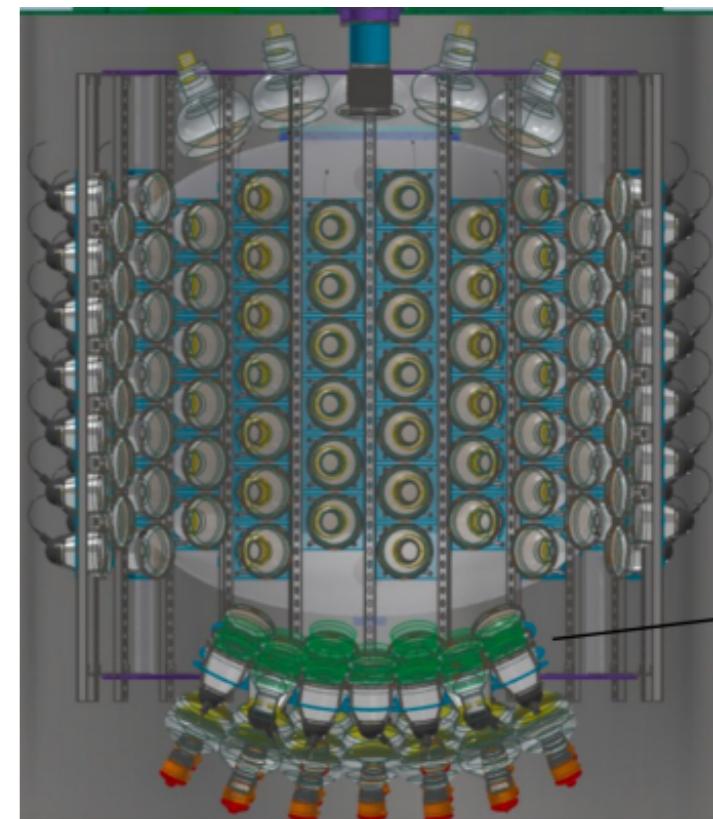
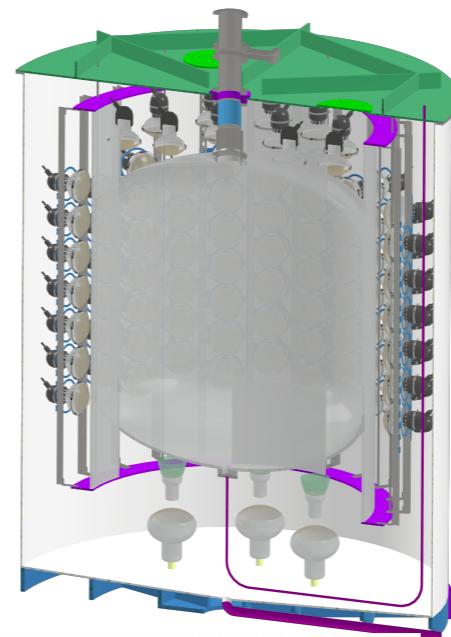
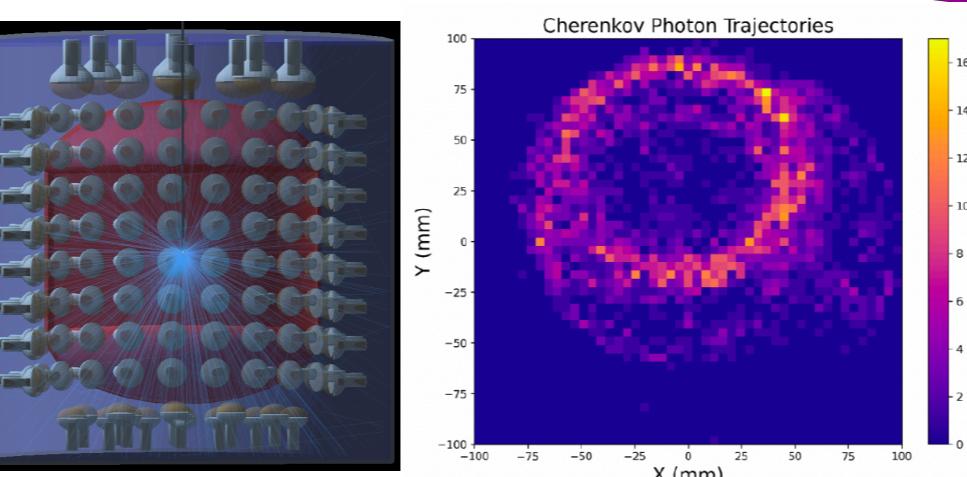
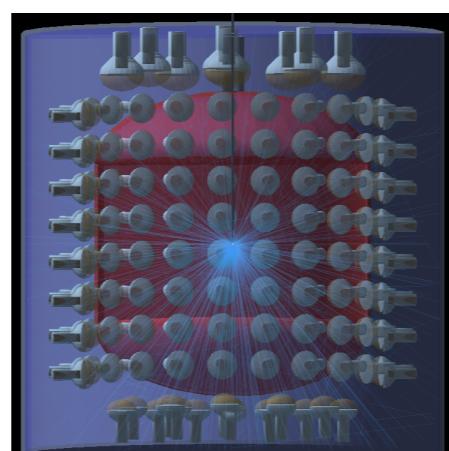


# Eos: performance demonstrator for next-generation neutrino experiments



- Demonstrate event reconstruction using hybrid Cherenkov + scintillation signatures
- Validate models to support large-scale detector performance predictions
- Provide a flexible testbed to demonstrate impact of novel technology: targets, photon detectors, readout, reconstruction algorithms

- 4-ton target mass water, WbLS, LS
- 200 8-” PMTs *R14688-100, 900ps FWHM*
- CAEN VI730 readout
- Dichroicon deployment for spectral sorting
- Range of deployable sources for studies of vertex, energy, direction resolution and PID



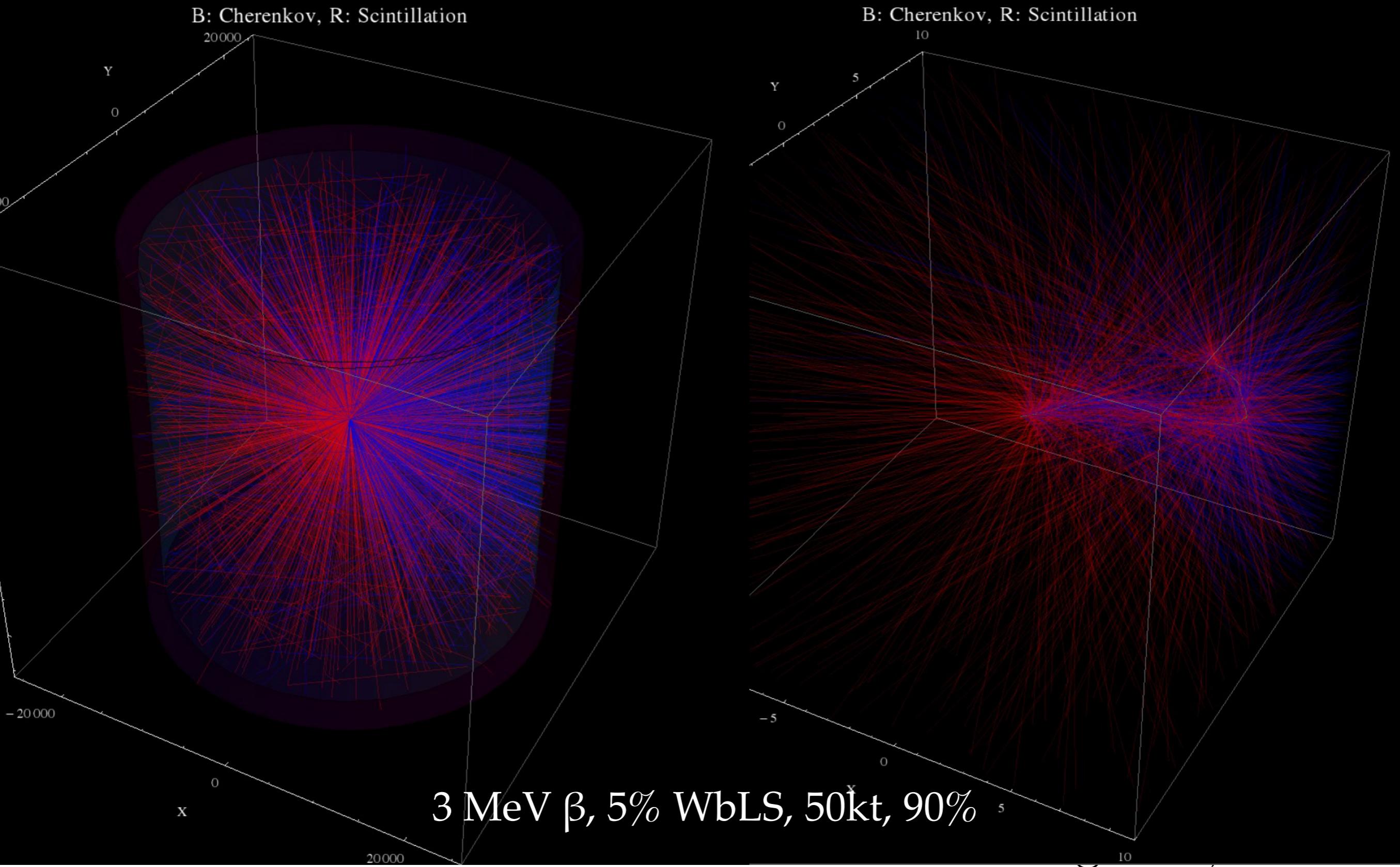
# BNL 1- & 30-ton demonstrators

- First ton-scale deployment of WbLS
- First demonstration of:
  - Ton-scale production
  - Optical transparency in an operating detector
  - Optical stability over time
  - Recirculation of WbLS (nanofiltration)

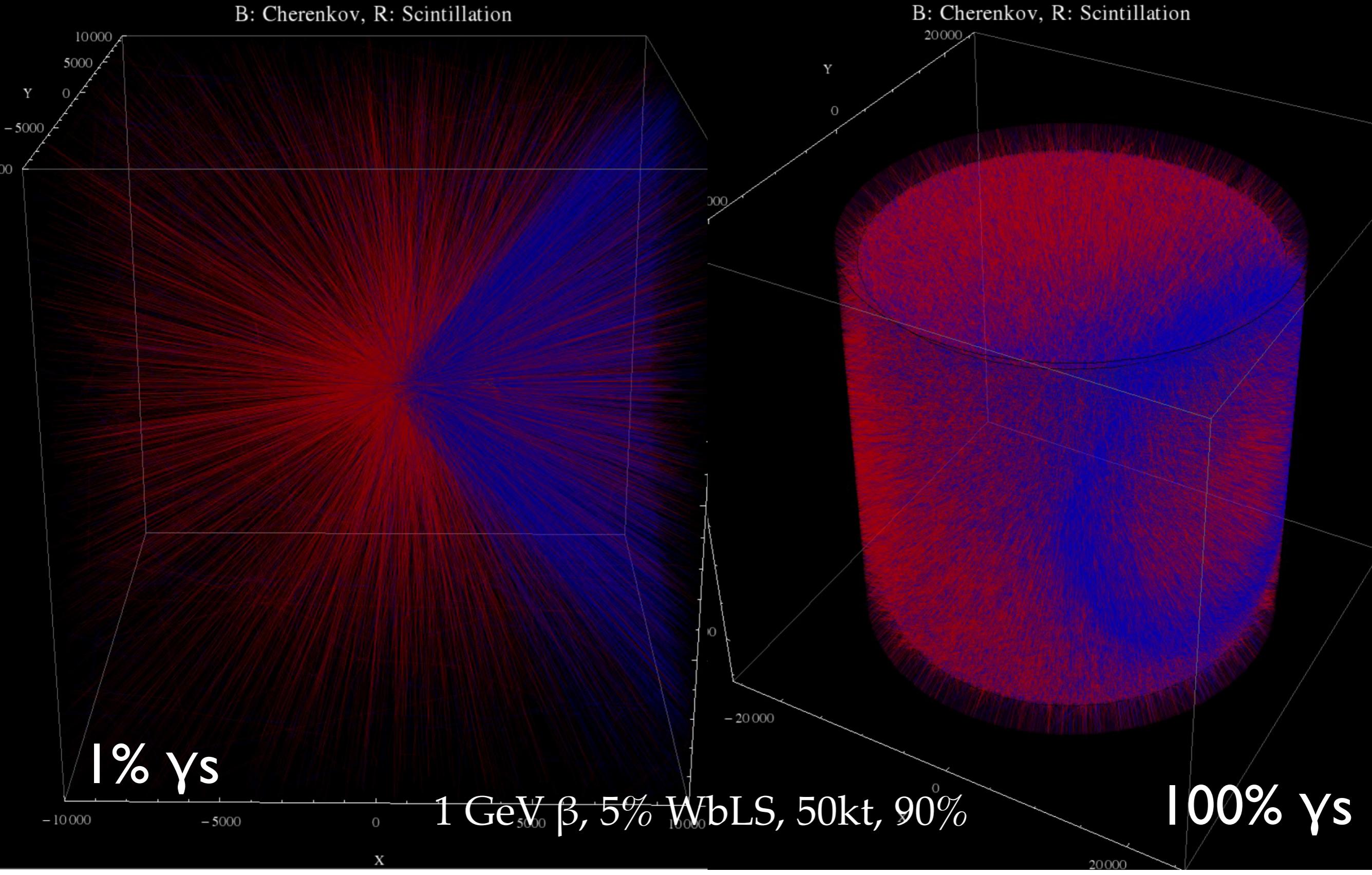


Together these prototypes will demonstrate the feasibility and capabilities of hybrid detectors for nonproliferation and fundamental physics

# Signal Separation in Theia



# Ring Imaging

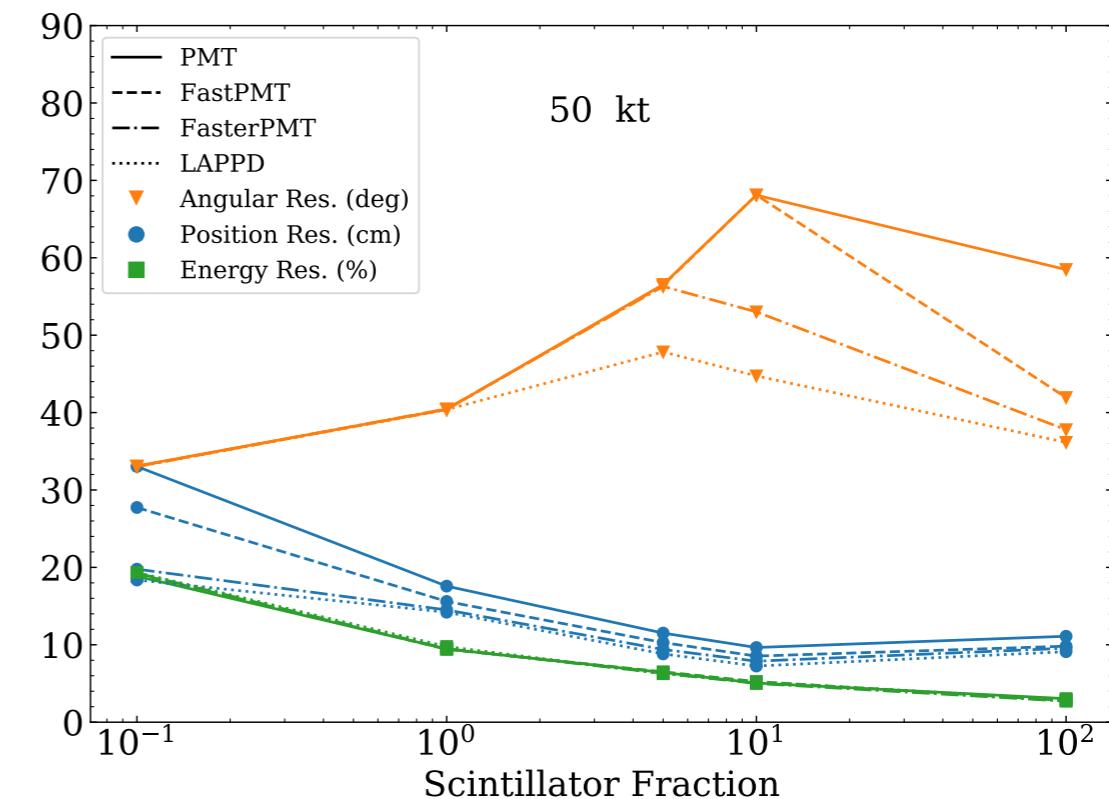
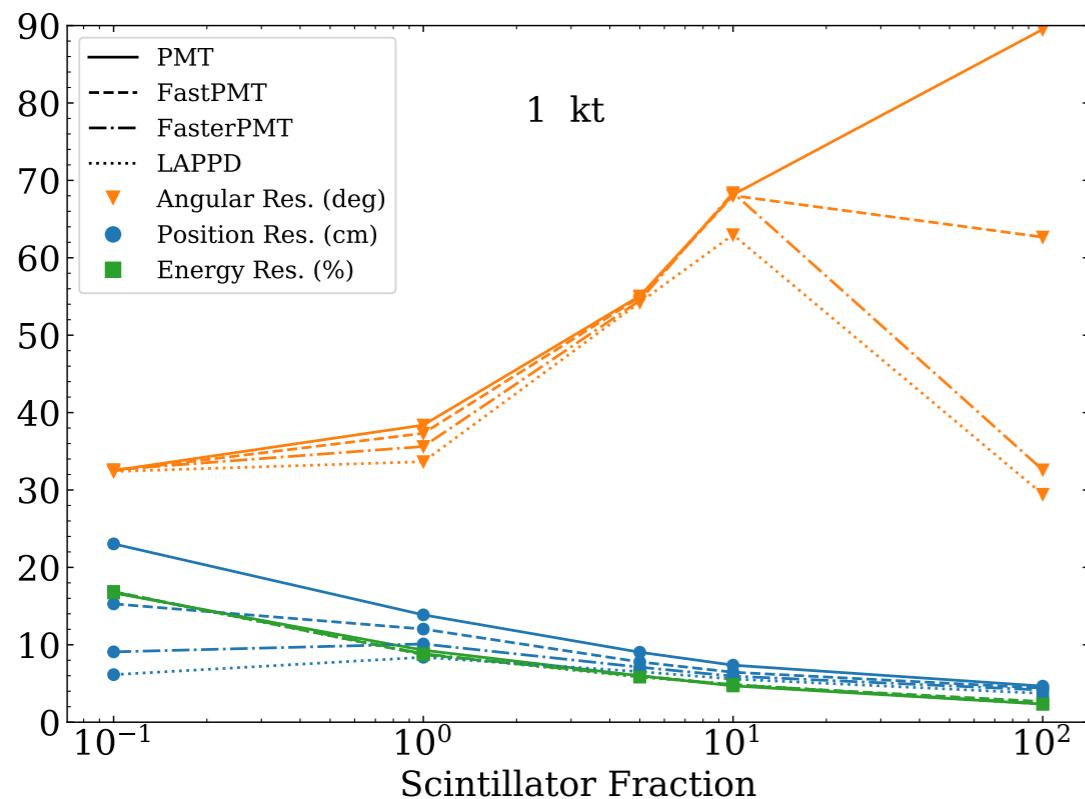


# Model validation

A number of metrics are considered for detector performance:

- |   |   |
|---|---|
| 1. Energy resolution                            | <i>Reduce flux uncertainty, increase background rejection</i> |
| 2. Vertex resolution                            | <i>Reduce flux uncertainty, increase background rejection</i> |
| 3. Angular resolution                           | <i>Elastic scattering event ID, physics scope</i>             |
| 4. Cherenkov (C) / scintillation (S) separation | <i>Particle &amp; event ID</i>                                |

These tools can be used to define “desired” properties for WbLS

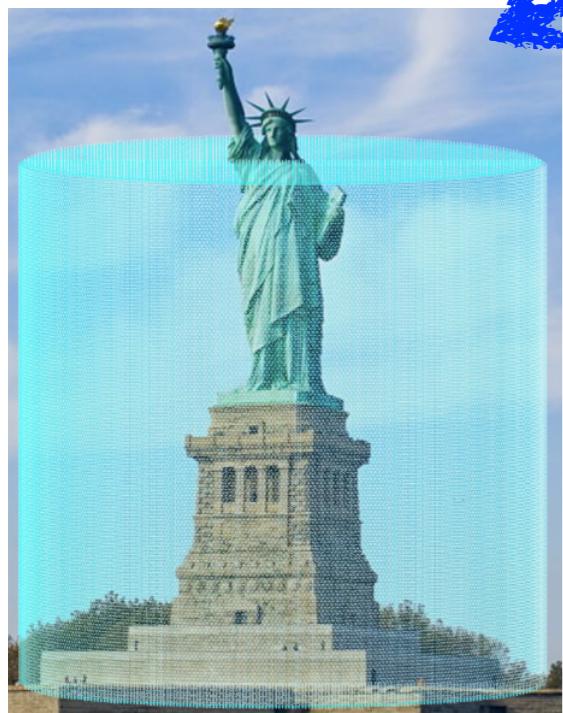


*Impact of target properties and photon detector response on detector performance for (left) 1-kton and (right) 50-kton detectors*

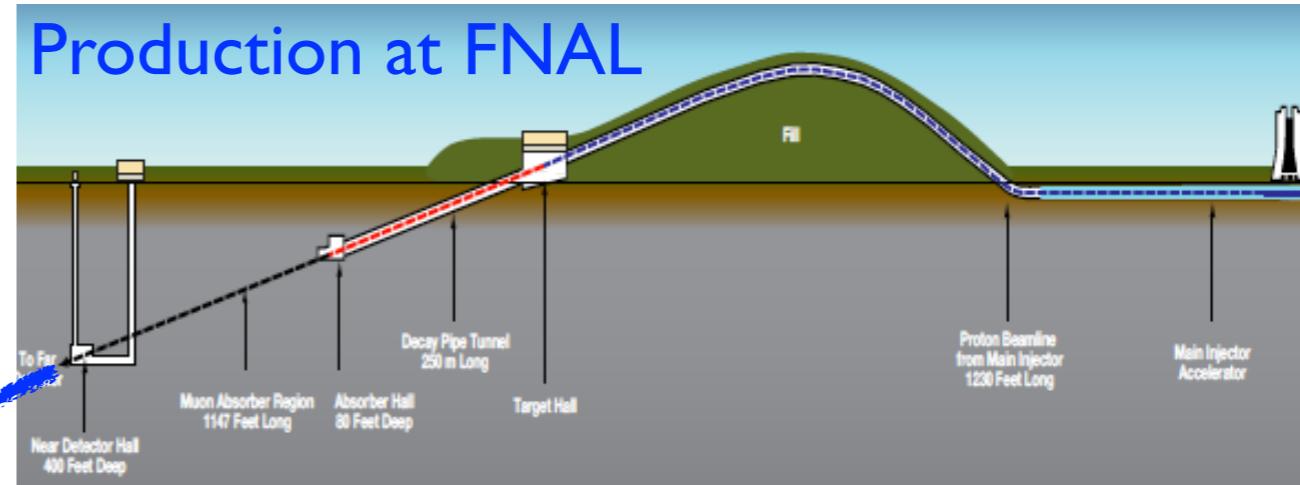
Phys. Rev. D 103 052004 (2021)

# Long-Baseline Program

- Large-scale detector at Homestake, in the LBNF beam
- Complementary program to LArTPC (DUNE)
- Build on WCD studies  
(arXiv:1204.2295)

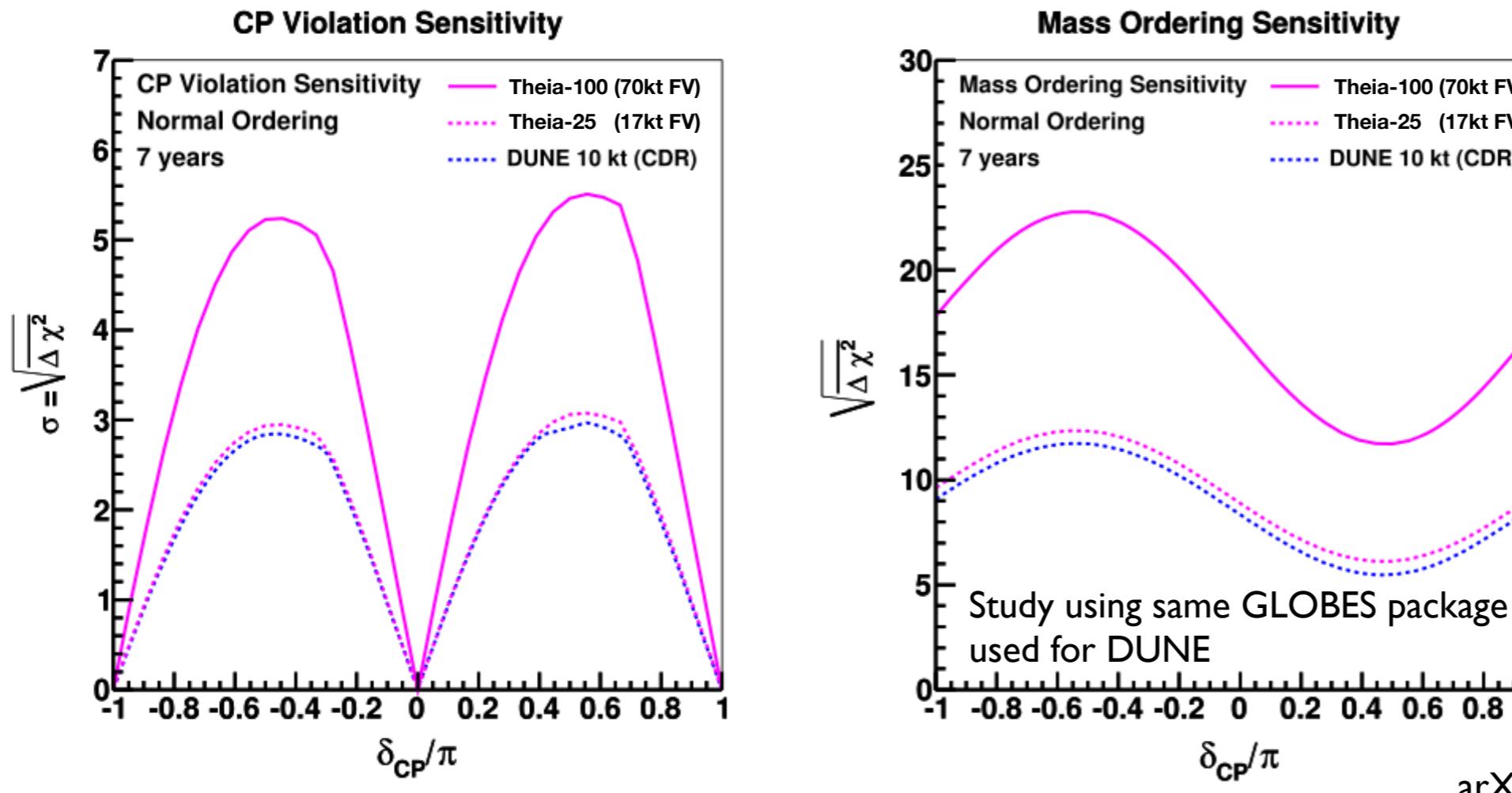


1300km



- Ring-imaging of a water Cherenkov detector
- Particle ID from Cher/scint separation
- n and low-E hadron detection (low threshold)
  - ▶ reduce wrong-sign component ( $\nu$  vs anti- $\nu$ )
  - ▶ reduce NC background by detecting  $\pi^0 \rightarrow \gamma\gamma$
- Large size → sensitivity to 2nd oscn max

# Long-Baseline Sensitivity

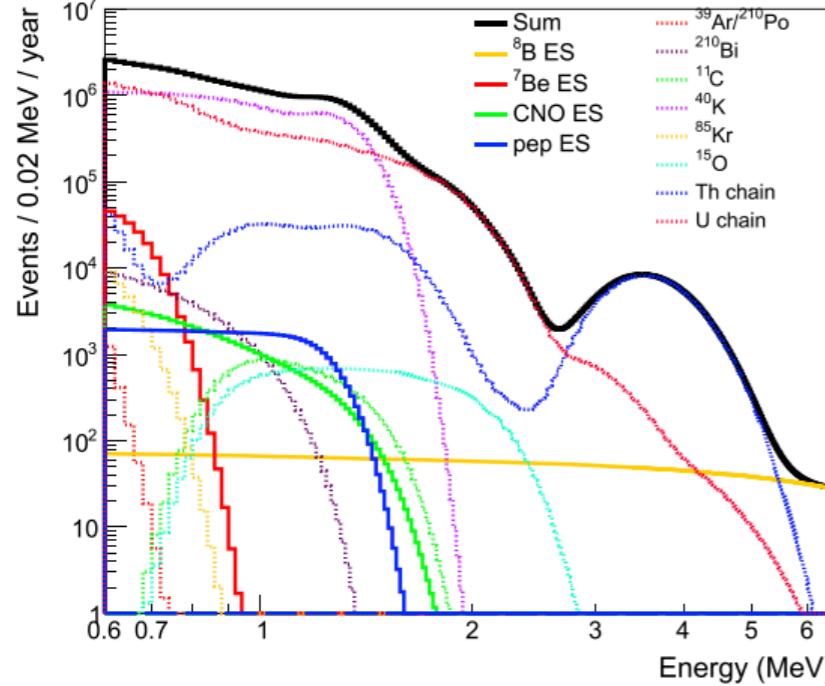


Synergy with LAr TPC  
Independent systematics  
High-energy events

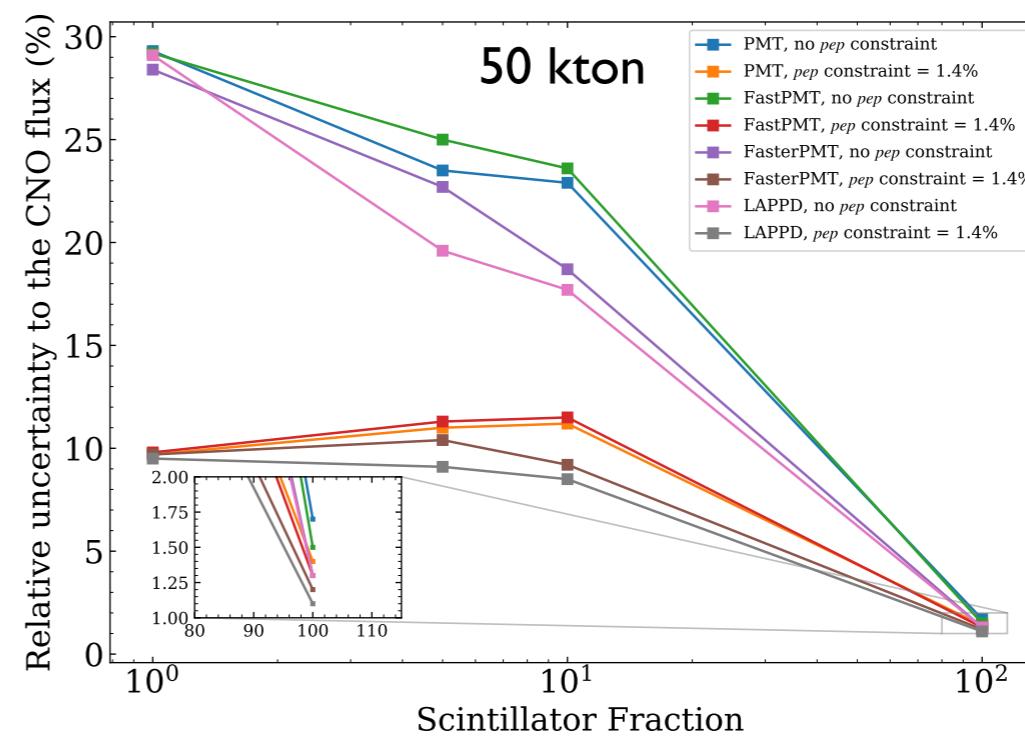
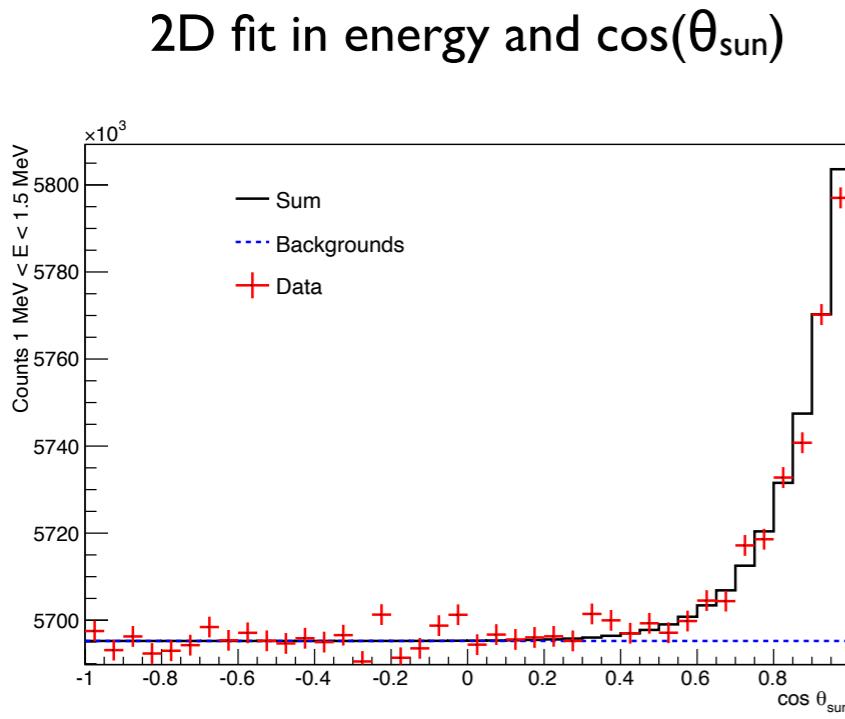
arXiv: 1606.09550

Performance of small (25kt) Theia module  
competitive with 10kt LAr TPC

# Solar Neutrinos with Theia



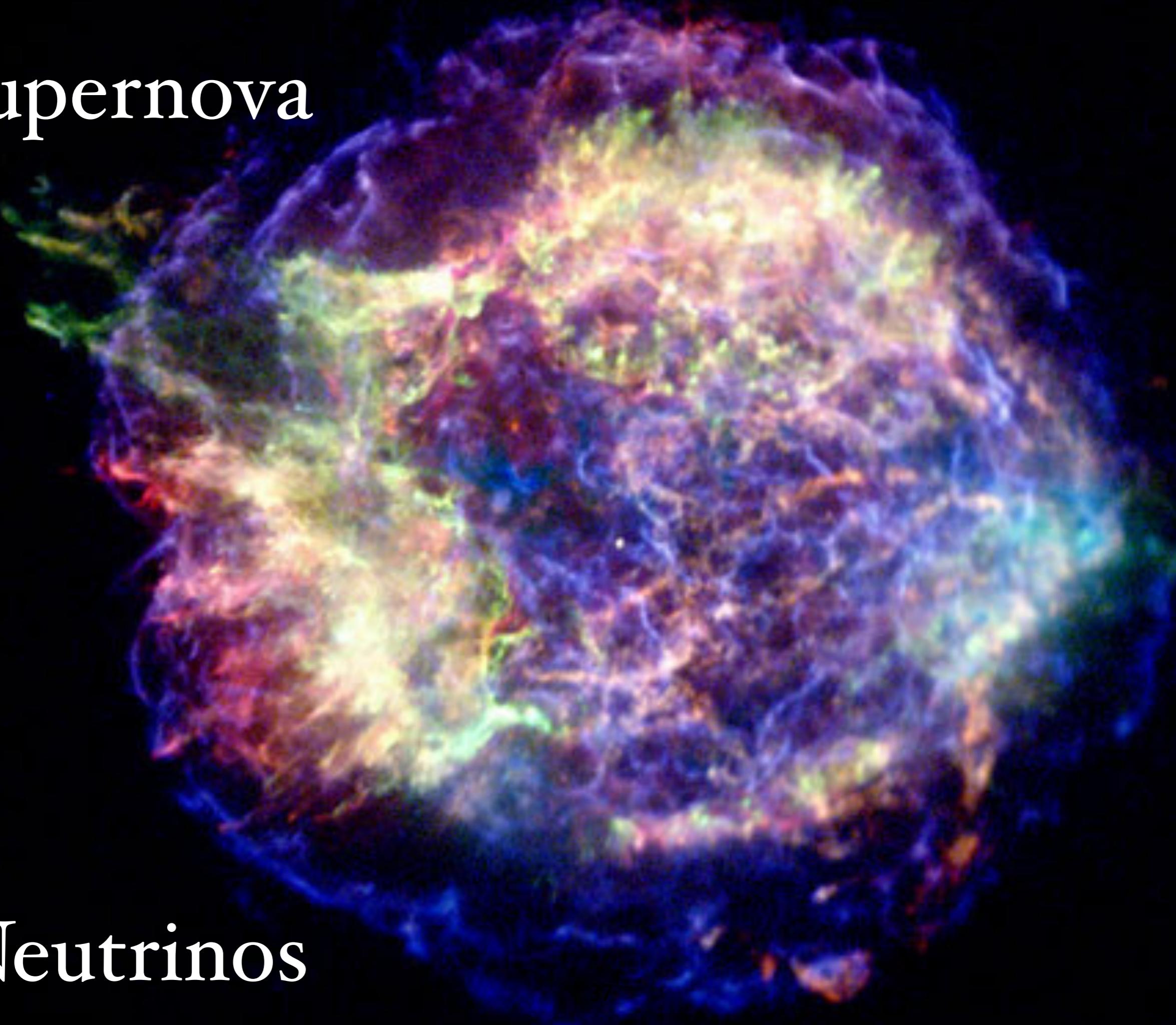
- Dominant background to CNO  $\nu$  measurement:  $^{210}\text{Bi}$
- Theia offers unique low-threshold, directional detection



Sensitivity based on MC study of detector response, including full vertex and direction reconstruction

Phys. Rev. D 103, 052004 (2021)

# Supernova



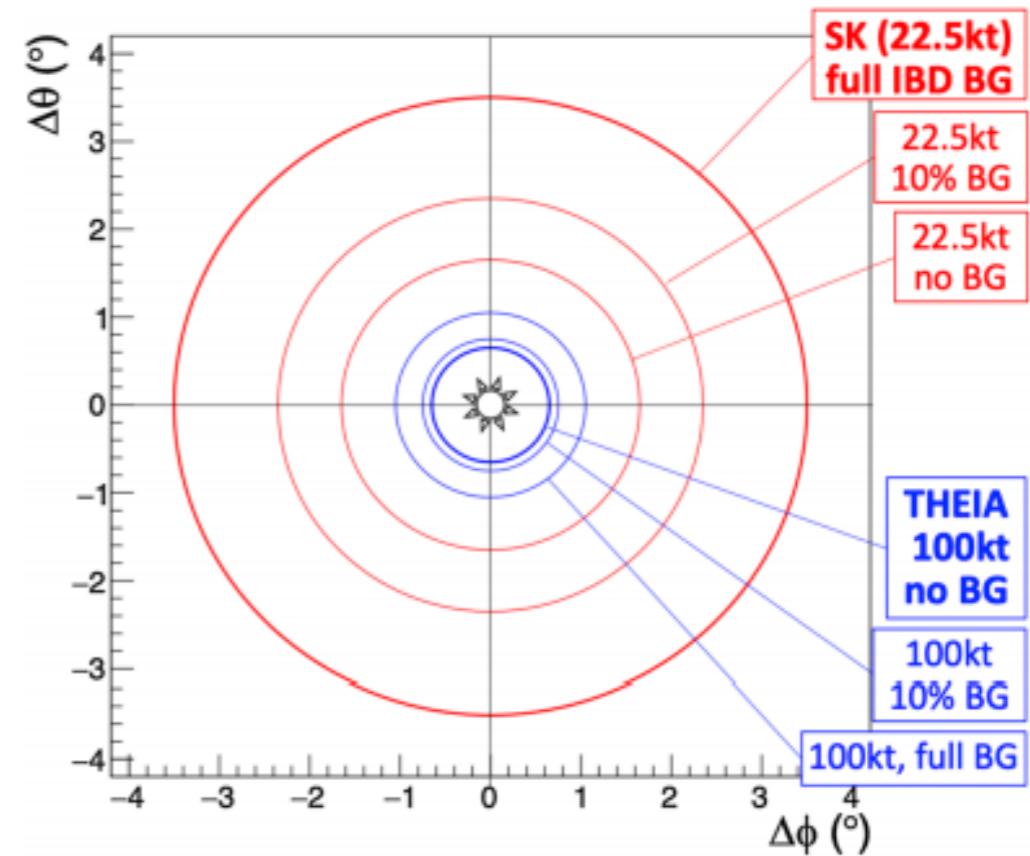
# Neutrinos

# Supernova Detection

- ~90% events are IBD  
*Highly complementary to  $\nu_e$  LAr signal*
- ES  $\Rightarrow$  pointing accuracy  $< 1^\circ$
- CC & monoE  $\gamma$  from NC  $\Rightarrow$  burst T & subsequent mixing
- Flavour-resolved neutrino spectra
- High-stats, low-threshold signal with good resolution
- Pre-supernova neutrino sensitivity

Event rate in 100-kt WbLS, SN at 10 kpc

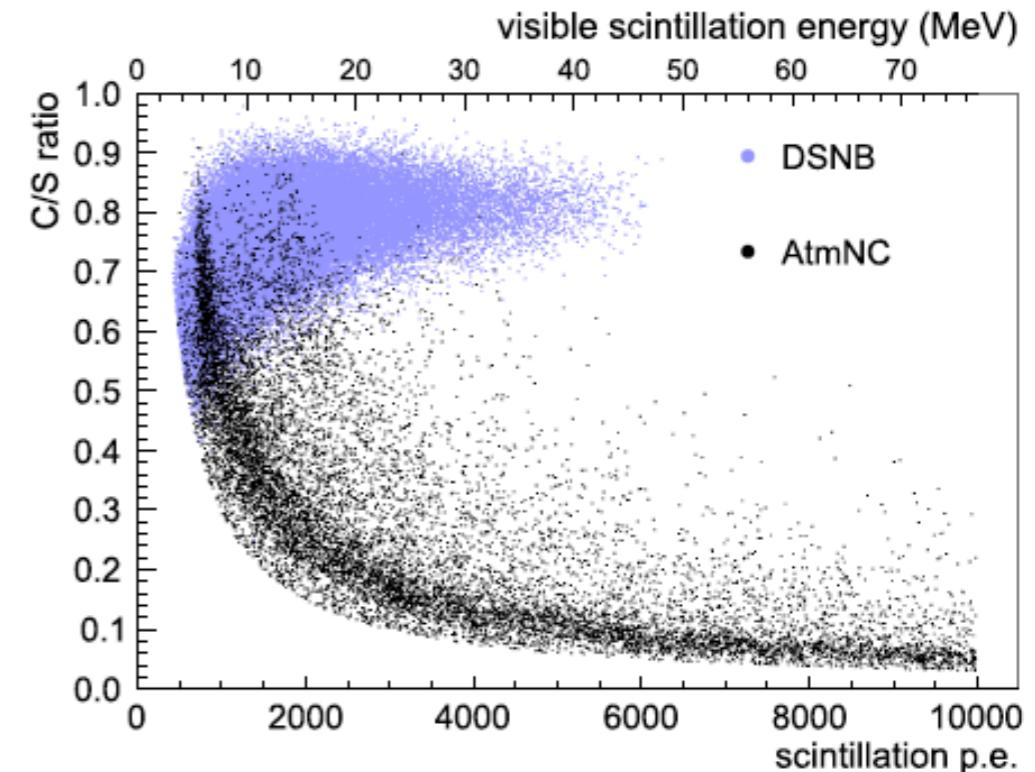
Reaction	Rate
(IBD) $\bar{\nu}_e + p \rightarrow n + e^+$	19,800
(ES) $\nu + e \rightarrow e + \nu$	960
( $\nu_e$ O) $^{16}\text{O}(\nu_e, e^-)^{16}\text{F}$	340
( $\bar{\nu}_e$ O) $^{16}\text{O}(\bar{\nu}_e, e^+)^{16}\text{N}$	440
(NCO) $^{16}\text{O}(\nu, \nu)^{16}\text{O}^*$	1100



# Anti- $\nu$ Detection

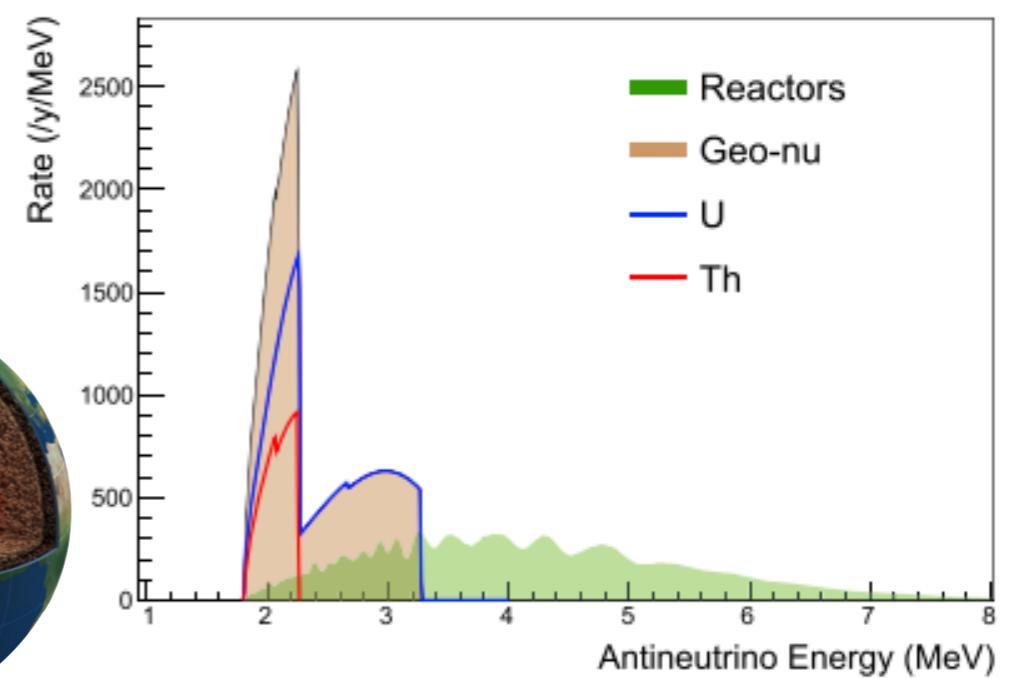
## DSNB

- Diffuse  $\nu$  “glow” from past core-collapse supernovae
- Cherenkov/scintillation ratio gives a powerful handle to discriminate atmospheric NC background
- $5\sigma$  in 125 kton-yrs
- Astrophysics of SNe



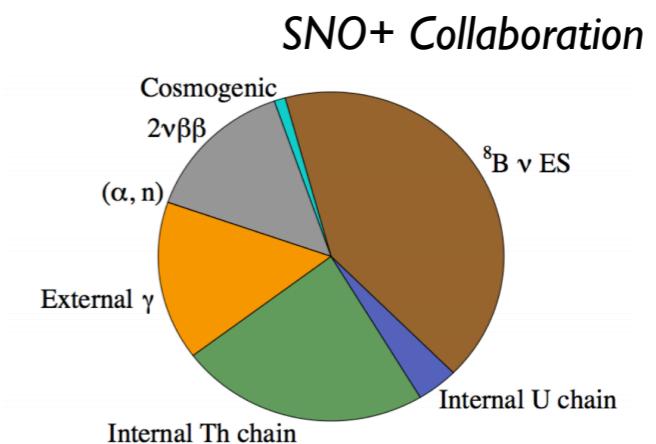
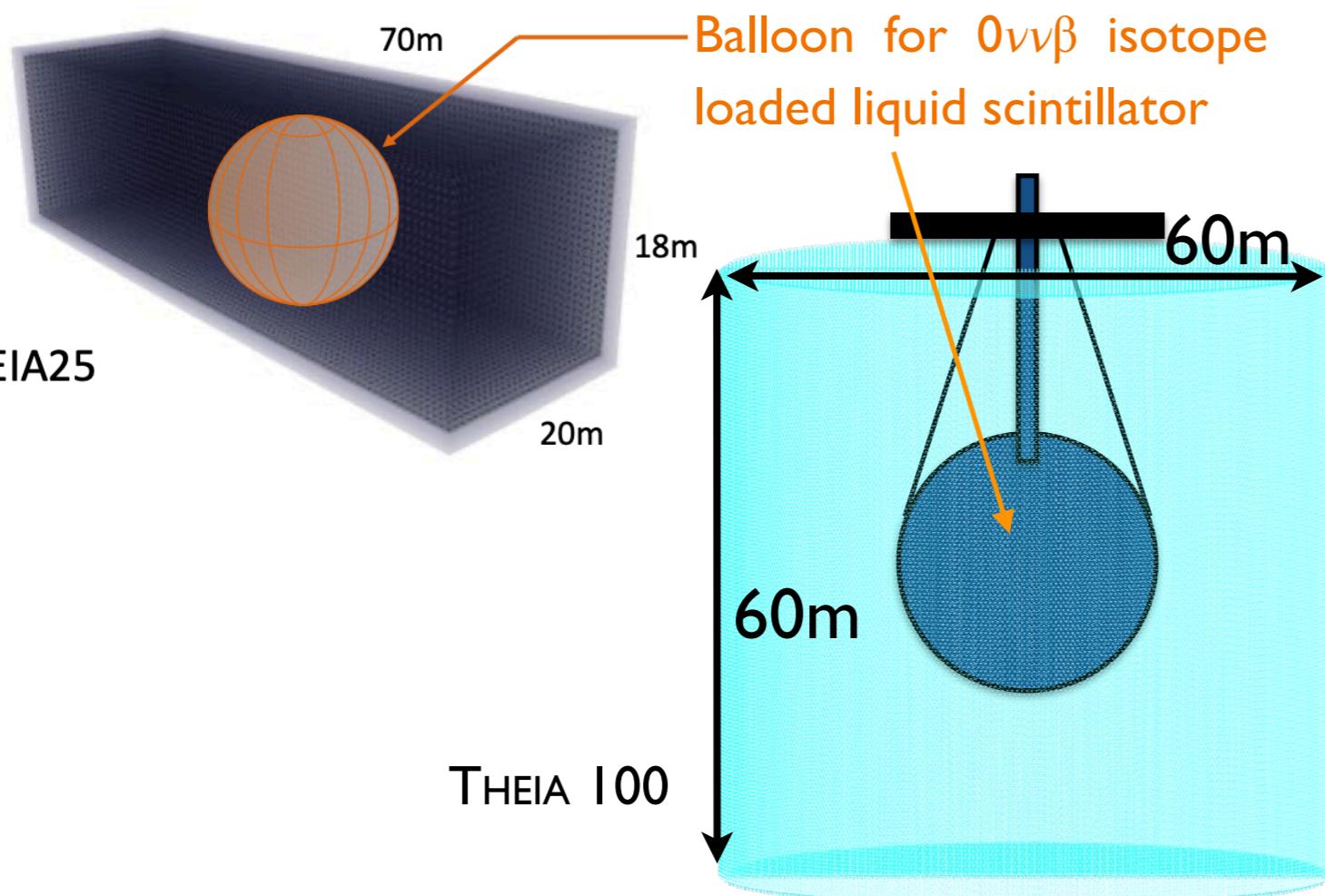
## Geo- and reactor

- Current geo- $\nu$  exposure < 10kt-yr (KL + Borexino)
- **THEIA:** large statistics in a complementary geographical location: 26.5 ev/kt-yr
- Could offer first evidence for surface variation
- ~ 20 reactor ev/kt-yr  
Demonstrate techniques for remote reactor monitoring  
Range & direction at > 1000-km standoff



# NLDBD with Theia

25-100 kton hybrid optical neutrino detector  
8-m radius balloon with high-LY LS and isotope  
7-m fiducial, 3%  $^{nat}Te$  (or  $^{enr}Xe$ ), 10 years



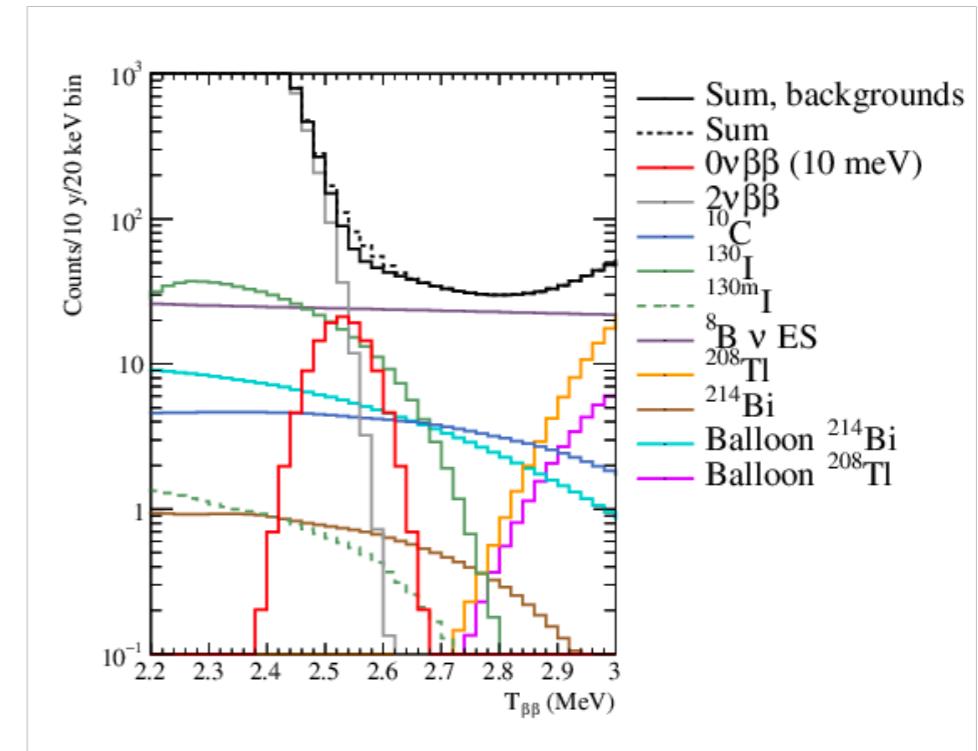
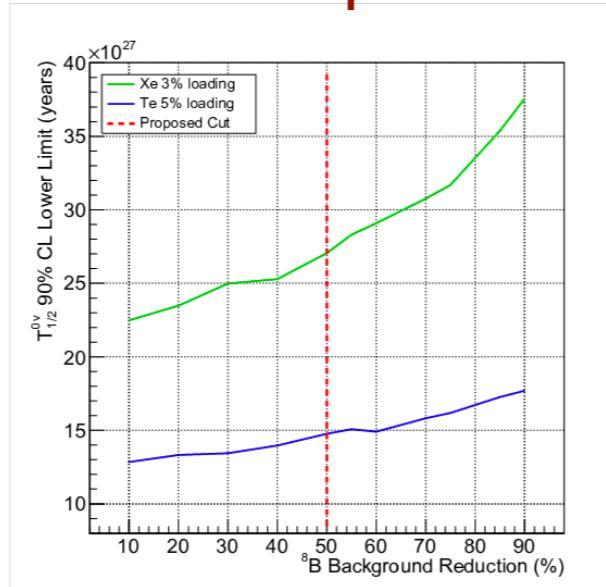
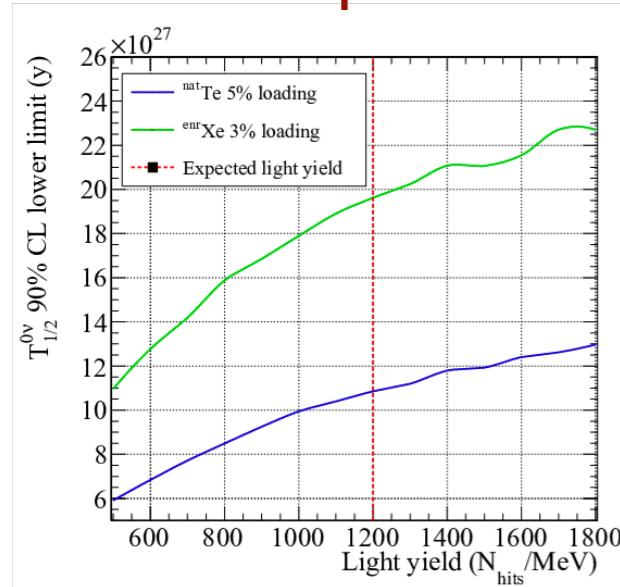
Directionality, PID and  
multi-site discrimination  
reduces backgrounds

Builds on critical developments by KLZ & SNO+ collaborations

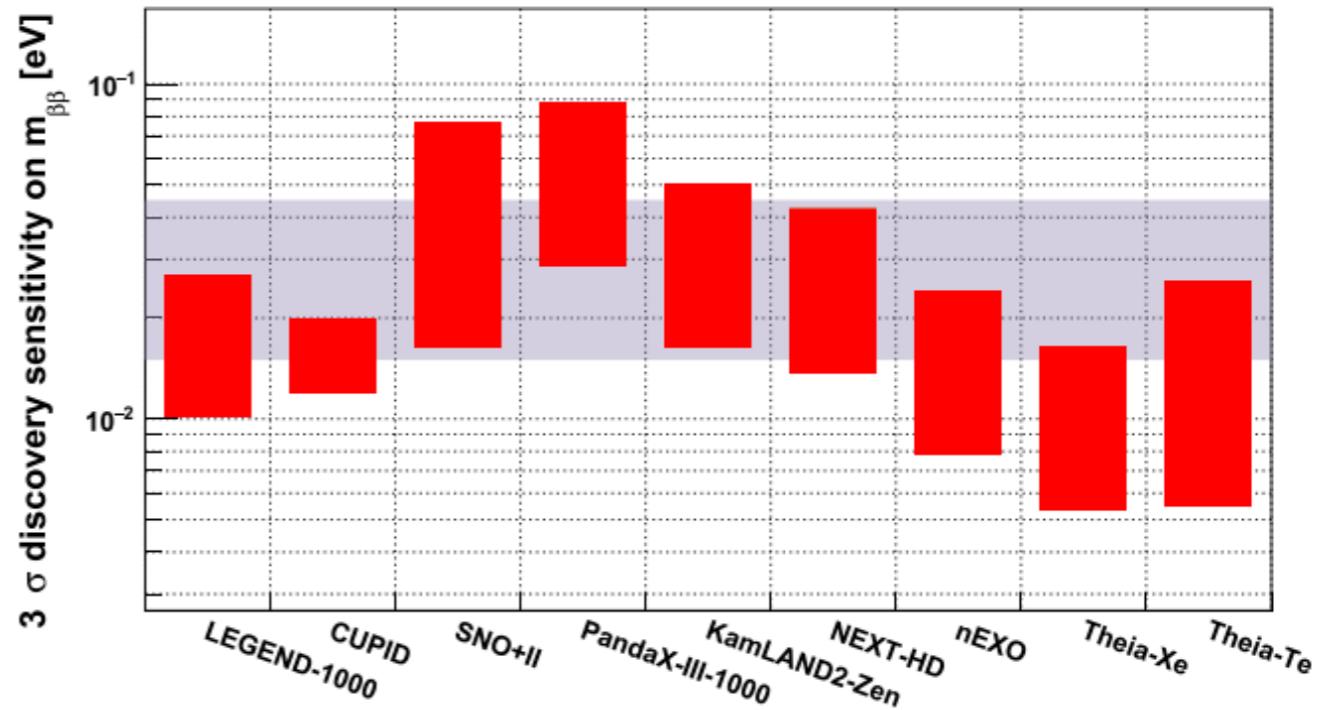
Phys.Rev.Lett. 110 : 062502 (2013); Adv.High Energy Phys. 2016 (2016) 6194250; Phys. Rev. D 87 no. 7 : 071301 (2013)

# NLDBD with Theia

## Dependence on detector response



$T_{1/2} > 1.1 (2.0) \times 10^{28}$  yrs  
 90% CL for Te (Xe)  
 $m_{\beta\beta} < 6.3 (5.6)$  meV



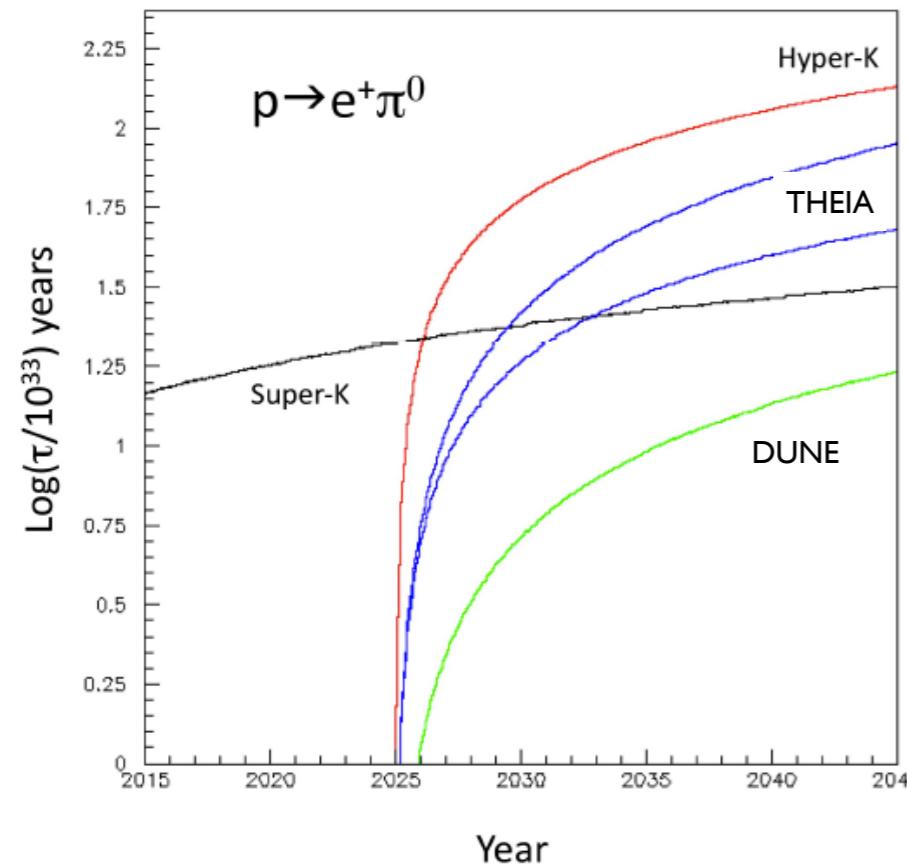
# Nucleon Decay

Testing the existence of GUTs with THEIA:

- Large size (statistics), deep location, very clean
- n tagging (low threshold plus potential isotope loading)
- Sub-Cherenkov threshold detection

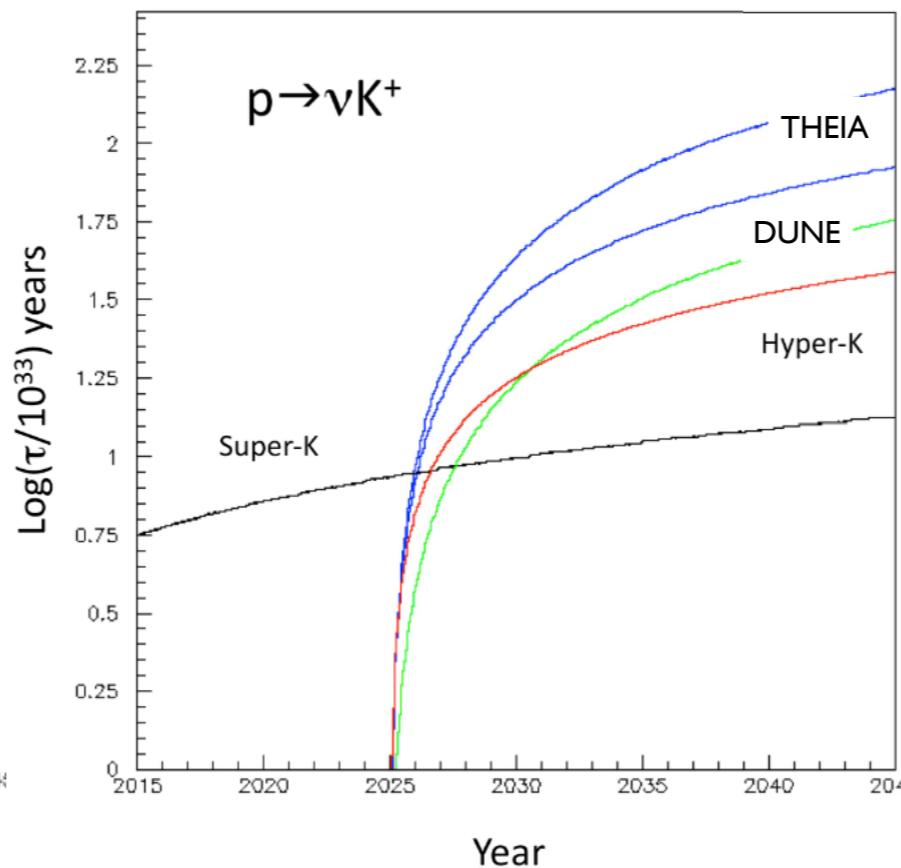
Figs from arXiv:1409.5864, assume  
100t FV; studies based on Phys. Rev. D  
72, 075014 (2005); LAr from JHEP  
0704:041, 2007

Heavy X boson exchange



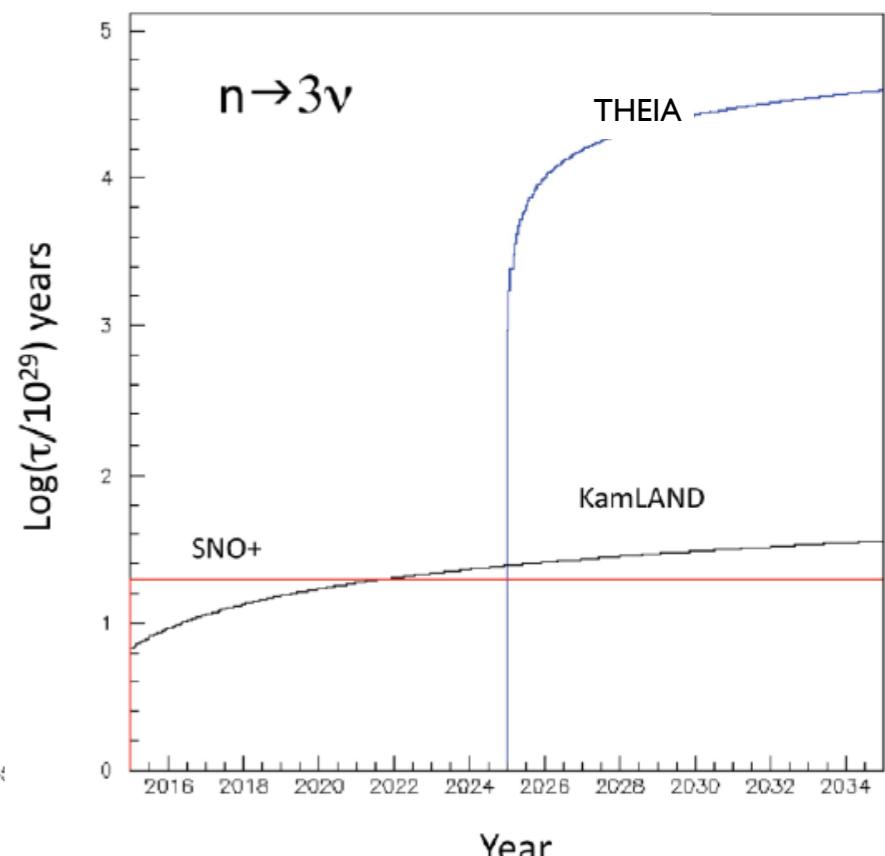
Enhanced n tag  
Reduced atmos. ν bkg

2nd order processes



Sub-Chr t/h detection  
⇒ Directly visible K⁺

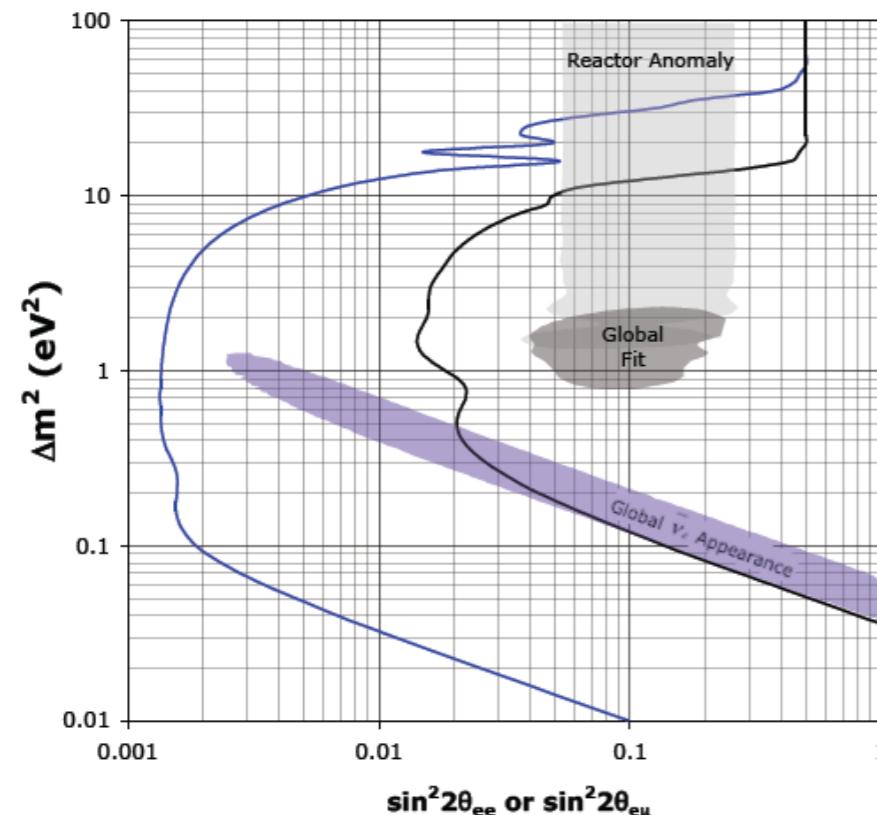
Extra dimensions



Deep, low threshold  
Directionality + n tag

# Sterile Neutrinos

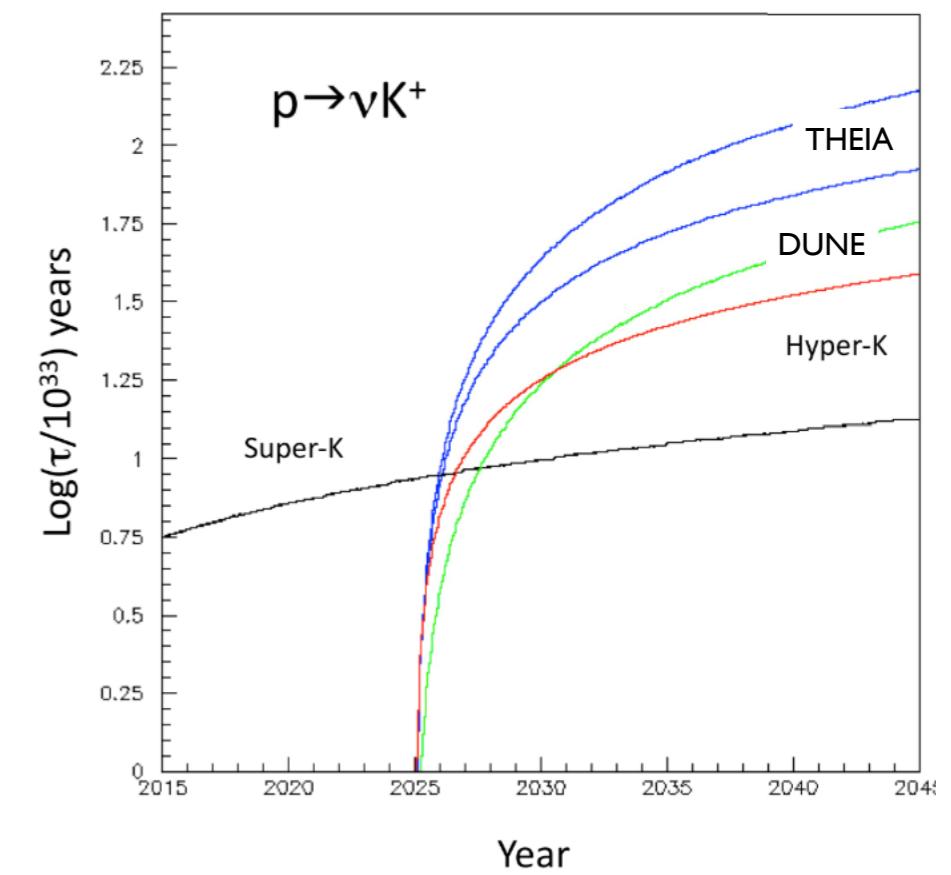
- Deploy  ${}^8\text{Li}$  decay-at-rest (IsoDAR)
  - 13MeV endpoint (above r/a)
  - Required detector response: 15% (E) & 50cm (R)
  - 5 yrs, 1kt (black) / 20kT fid. (blue)



- Heavy-water based LS: 2n tag  
(reduce bkg in IBD searches)

# Nucleon Decay

- Large, deep, very clean
- Enhanced n tag
- Sub-Cherenkov threshold detection
- Sensitive to several modes



Sub-Chr t/h detection  
⇒ Directly visible  $K^+$

# Photon propagation

Figs from Ben Land

5-MeV electron, 7-m (radius/half-height) detector

