

A Tour of International Underground Facilities & Science: Today and Tomorrow

Clarence Virtue Interim Executive Director





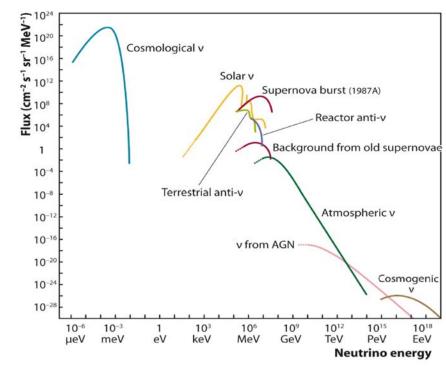
Why create deep underground facilities?

To access the science with improved signal to background

- Primarily by reducing cosmic ray muon flux
- Reducing spallation products and avoiding cosmogenic activation of materials
- Making highly sensitive searches for new physics possible, and
- Improving precision on low signal measurements

Depending on the objective, signals (or backgrounds) have been

- Cosmic rays
- Neutrinos from multiple natural and terrestrial sources
- Dark matter
- Radioactive decay including $0\nu\beta\beta$ decay
- Ambient neutron and gamma fields, and radon concentrations







Physical

- Depth, size (floor space and volume), layout
- Access horizontal or vertical
- Overburden flat or not
- Proximity to reactors, coupling to accelerator beams
- Limitations on component size and weight
- Local geology rock quality, U, Th, K concentrations, water ingress

Operational

- Scientific Support laboratory systems QA and development
- Access support, logistics
- Utilities power, water, internet, LN2, ventilation, chilling
- Integration support personnel, safety, machine shops
- Project management support, engineering



Broad Scientific Programme / Diverse Facilities

Diversity in

- Physical and operational attributes
- Challenges associated with specific facilities
 - Material handling logistics
 - Seismicity engineering constraints
 - Rock mechanics
 - Local support
 - Remoteness / travel logistics

N.B.

- Excluding mega-scale water Cherenkov detectors for neutrino astronomy
- Concentrating on deep, operating, or planned / in construction facilities
- Focus on facilities and their scientific programmes but not generally surveying scientific results
- Drawing on recent (LRT 2022) overviews of facilities presented by Jaret Heise, Sean Paling and Douglas Leonard with input from multiple lab directors

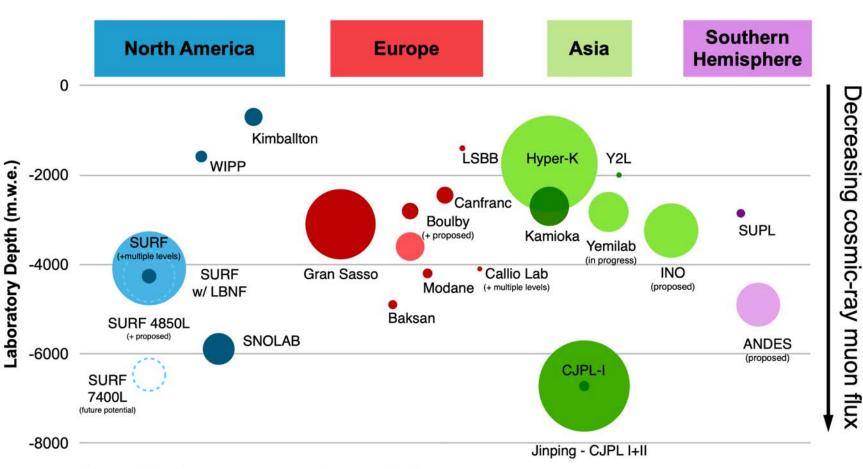
International Deep Underground Facilities

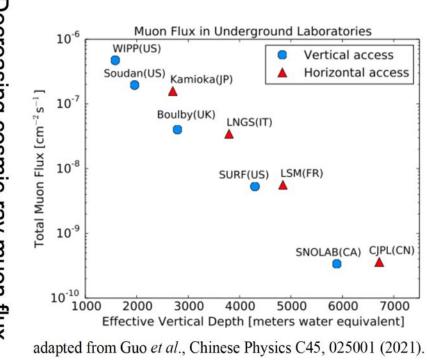






Some visual attributes





Note: Circles represent volume of science space

Baksan (Russia)

- Celebrated 50 years in 2017
- New results from BEST (Baksan Experiment on Sterile Transitions) V. V. Barinov et al., Phys. Rev. C 105, 065502, June 2022.



Underground laboratories of the Baksan Neutrino Observatory, at increasing distances from the tunnel entrance (right). Being under a mountain the shielding increases with distance along the tunnel.

Image credit: Y Gavrilyuk.







Photos from CERN Courier, May 2017



The assembled gallium sterile-neutrino experiment (BEST). Image credit: S Demidov.

Boulby Underground Laboratory (UK)



Boulby Underground Laboratory



Boulby Facility Details...



- The UK's deep underground science facility. One of 5 in Europe, <15 in the world.
- Supports work of >10 collaborative projects (astrophysics to climate, geology, environment etc), >40 institutions, >170 scientists & students.
- Facility funded and operated by the Science & Technology Facilities Council (STFC).
- Operations, H&S & science programme managed by 10 (+2) onsite staff and supported by Rutherford Appleton Lab (PPD).
- · Mine operators ICL-UK provide wide-ranging operational & high level support.

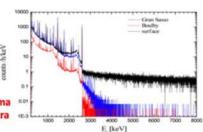


How does Boulby Compare?

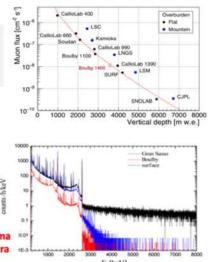
- Low Radon levels (3 Bg/m³) · Diverse science programme.
- Science and Industry partnership









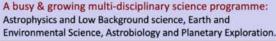




Underground Science @ Boulby Mine

High level suppor

- DRIFT/CYGNUS: Directional Dark Matter
- · Spherical Proportional Counters (NEWS-G) R&D
- · BUGS: Ultra-low background material screening (for LUX-ZEPLIN and Super-K-Gd and more)
- AWE(Ge): Atmospheric gamma spectroscopy
- RESOURCE: Salt cavity energy storage study
- BISAL: Geo-microbiology / Astrobiology studies
- · MINAR: Space Exploration Tech. Development
- Misc. Low Background & Geoscience...
- Etc... (More to come).







ULB screening of LZ PMTs



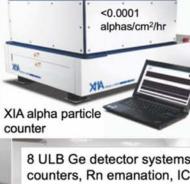
Boulby Science Now & Future

Particle physics and ultra-low background studies

BUGS



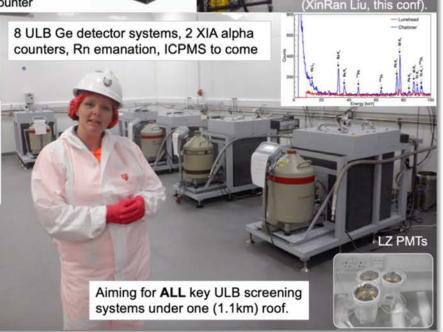
XIA alp counte

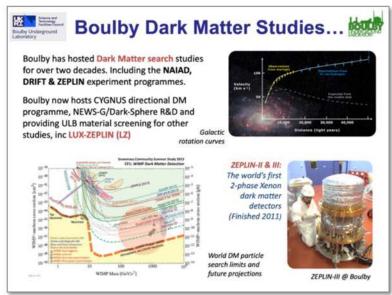


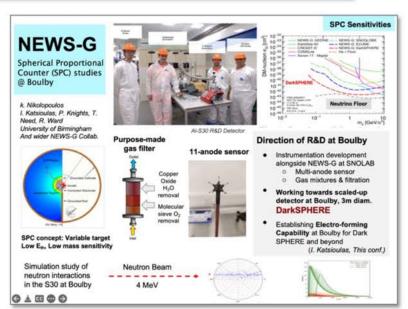


BUGS (Boulby UnderGround Screeening). World-class material screening for current and future ULB experiments. Towards PPT sensitivity for G3 DM and Neutrino experiments











Underground Centre for Science and R&D



LOCATED AT THE 1.4 KM (4100 MWE) DEEP PYHÄSALMI MINE, PYHÄJÄRVI, FINLAND

UNIQUE UNDERGROUND RESEARCH NETWORK AND INFRASTRUCTURE -ACCESS, DEPTH, FACILITIES

CURRENTLY SIX UNDERGROUND HALLS OR TUNNEL NETWORKS HAVE BEEN TURNED INTO MINE RE-USE FACILITIES: LABS.

MINING ENDS IN JUNE 2022. POST-MINING ACTIVITIES COORDINATED BY CALLIO PYHÄJÄRVI – BUSINESS CONCEPT

MULTI-DISCIPLINARY STEERING GROUP ESTABLISHED 2020



FACILITIES







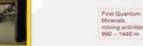


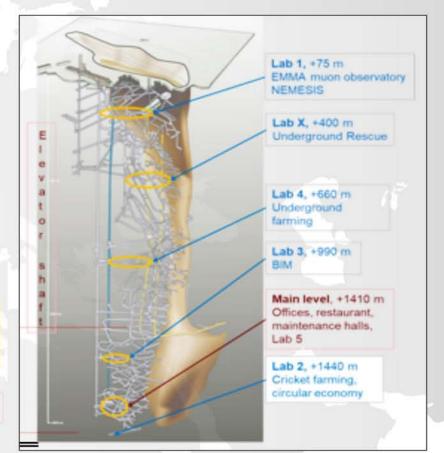












Web: calliolab.com





Underground Centre for Science and R&D

ACTIVITIES

- EMMA: Experiment with a MultiMuon Array, cosmicray induced bkgds using drift chambers
- NEMESIS: New Emma MEasurement with neutronS In cosmic Showers, study neutron bursts in shielding materials (also performed at other Eurpoean labs)
- C14: Measure 14C isotope in oil-based liquid scintillators (goal of 10e-20 or lower)
- BSUIN: Member of the Baltic Sea Underground Innovation Network (13 partners from 8 countries), incl bkgd measurements, develop best practices, etc
- Goldeneye: Test site for remote sensing technology (safety and environmental monitoring)
- Occupational Health: Intelligent, adaptive lighting studies for UG workers
- · Biology & food production, geology & hydrogeology
- UG Rescue & mining training





Coordinator, Jari Joutsenvaara (jari.joutsenvaara@oulu.fi)

Project engineer, Julia Puputti (julia.puputti@oulu.fi)

Future: Globally recognised underground research network and infrastructure



Education and training



Future food & Underground farming



Mining & tunnelling



SpaceLab



Mine reuse



Earth Observation and remote sensing



Geothermal research



Deep underground low background facility



Working environment



Particle physics & muography



Underground H&S

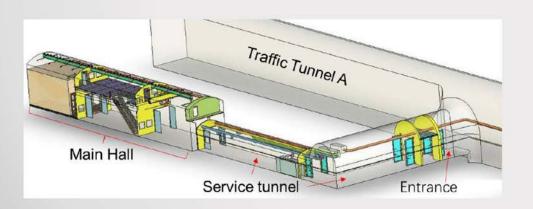


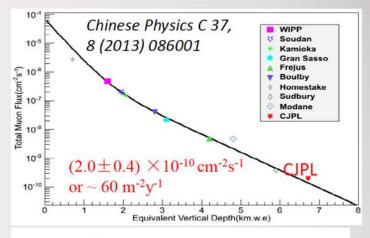
Something new?

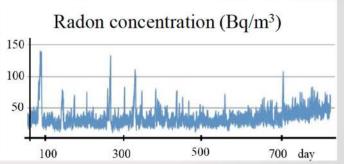


CJPL Features

- Deepest underground laboratory with a rock overburden of 2400m.
- Open on Dec. 12, 2010, extended from hydropower facility.
- Total space: ~4000 m³.
- Main Hall: 6.5m(W) × 6.5m(H) × 42m(L).
- Low Muon flux and environmental background.







Rock sample(Marble) measurement by Germanium

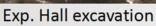
(Unit: Bq/kg)	K-40	Ra-226 (609keV)	Th-232 (911keV)	
Rock Sample	< 1.1	1.8 ± 0.2	< 0.27	
Ground Level(Beijing)	~600	~25	~50	



CJPL-II Project

- Construction of CJPL-II started on Nov. 25, 2014
- Dec. 2015: The rock excavation of all exp. halls completed.
- May 2016: Expansion of two exp. pits finished.
- Dec. 2016: Installation of ventilation system started.
- Total volume: 300k m³ with 4 main halls of 14x14x130 m (4000m³ of CJPL-I) containing 2° large caverns/pits.







Exp. Pit



Ventilation tubes



PHYSICS

China supersizes its underground physics lab

Planned expansion could pave way for "ultimate dark matter experiment"

he world's deepest physics laboratory is about to become one of its largest.

Early next year, workers will start can ha physics laboratory ince. One is about to become gundergreas well, physics," says Nigel Smith, director of the underground SNOLAB in Sudbury, Canada.

Science, Nov. 30, 2014



Science at CJPL





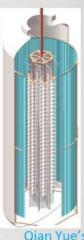
Hao Ma et al 2021 J. Phys.: Conf. Ser. 2156 012170

CJPL-I

- CDEX-1 p-type Ge WIMP search in PE room. Excludes DAMA/LIBRA region.
- CDEX-10 Ge in LN2. WIMP and dark photon search
- PANDAX-II 580kg Xe TPC.

CJPL-II

- CDEX-50 (DM-search), WIMP-nucleon SI coupling cross section goal: 10⁻⁴⁴ cm² at WIMP mass <10 GeV/c².
- CDEX-300v $0v\beta\beta$ with p-type point contact (PPC) ⁷⁶Ge in purified LAr in LN2 tank, half-life goal > 10^{27} years.
- PandaX-4T: 3.7-tonne of liquid Xenon, in B2 hall (water tank), SI DM CS limit results: 3.8×10⁻⁴⁷cm² at 30 GeV/c2. Yue Meng et al., Phys. Rev. Lett. 127, 261802
- JUNA in A1 hall, accelerator-based nuclear astrophysics, accelerator installed in 2020, with some early results. Liu, W.P., Li, Z.H., He, J.J. et al. Few-Body Syst 63, 43 (2022)
- Full commissioning of CJPL-II → March 2023.



Qian Yue's CDEX-300v talk, TAUP 2021.

Kamioka underground facilities (Mt. Nijugo-yama)

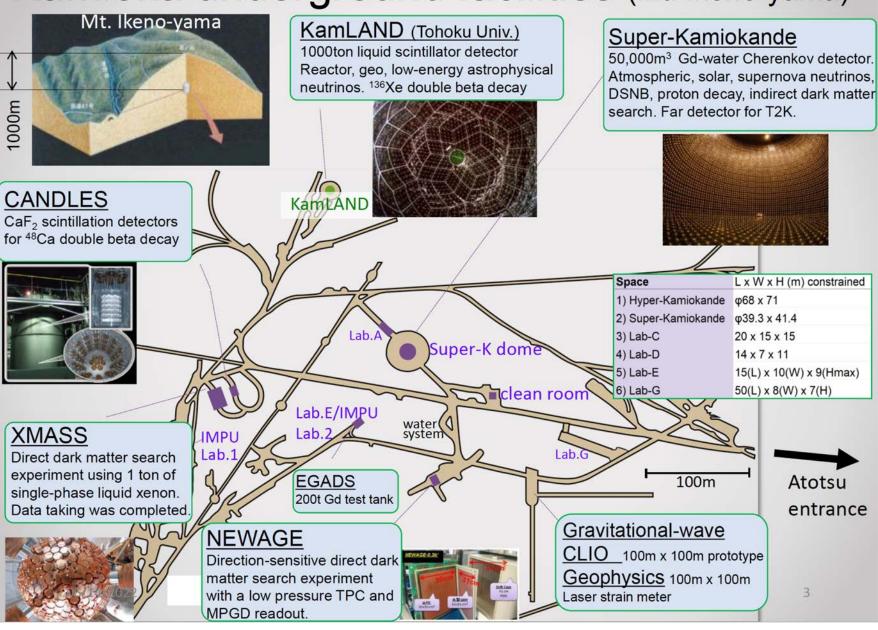
Following SK slides collected by Hiroyuki Sekiya (slightly modified)

Hyper-Kamiokande is under construction in Mt. Nijugo-yama 8km south from Super-K

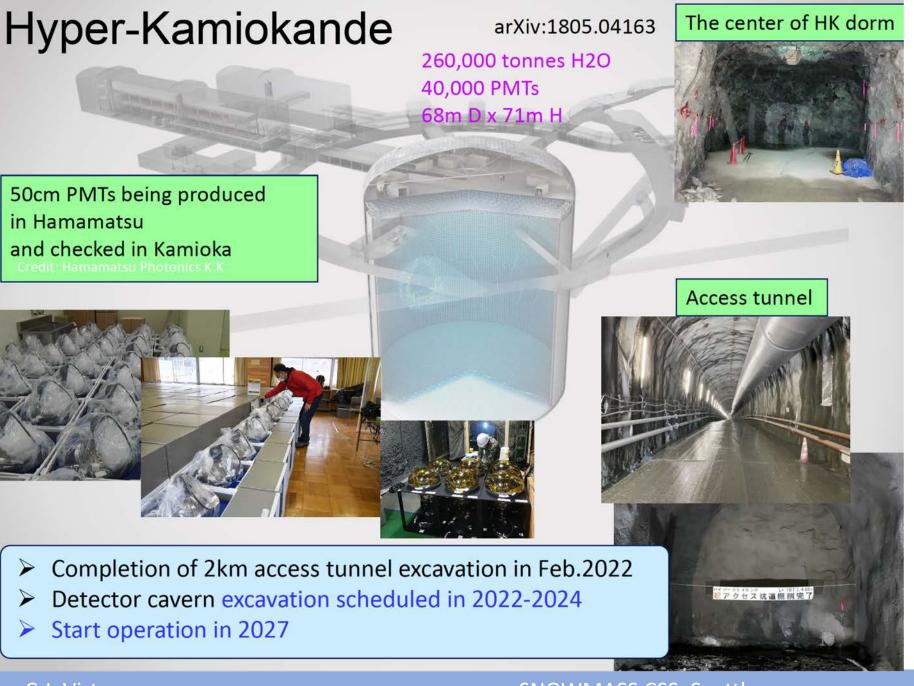




Kamioka underground facilities (Mt. Ikeno-yama)





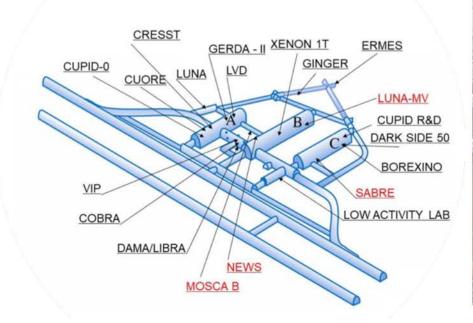


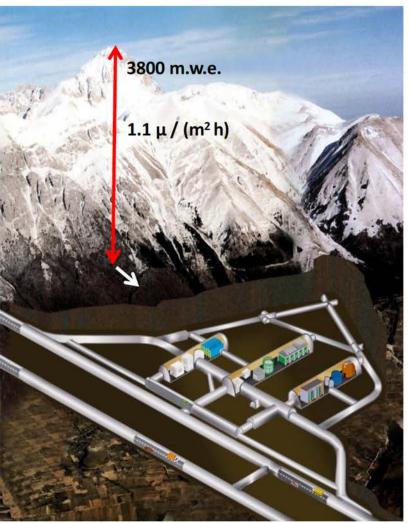


LNGS / Gran Sasso Deep Underground Lab

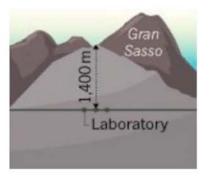


- Shielded by 1400 m (3800 m.w.e.) of rock (Gran Sasso Mountains)
- Total Muon flux 3 x10⁻⁸ cm⁻² s⁻¹
- Radon ~100 Bq/m3 with 5-8 air changes/day
- 3 main experimental halls: 100 m long, 20 m width and 18 m height (Vol = 180,000m³)
- 22 experiments data taking or under construction
- Laboratory for very low radioactivity measurements





Area: 17.800 m² Volume: 180.000 m³





E. Previtali

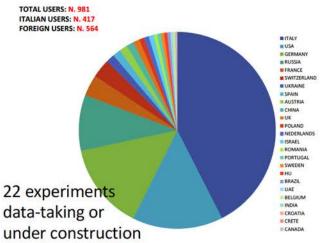


Gran Sasso Science











..... but also

- Test on quantum mechanics
 - Study on Planck invariance
 - Electron decay
- Radiobiology
 - Biological effects of low radioactive environment
- Geophysics
 - Earthquacke monitoring and study
 - Analysis of water resources
- Ultra Trace elemental analysis
 - Low radioactivity tests and measurements
 - Cultural Heritage analysis
 - Advanced additive manufacturing

E. Previtali



Canfranc - LSC





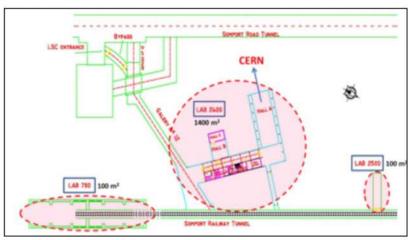
Depth: 2450 mwe

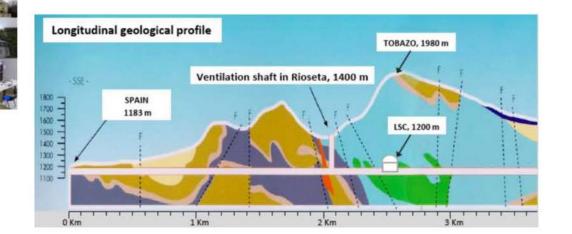
Volume: 10,000 m³

Radon: 50-80 Bq/m³

Access: Horizontal

Rail tunnel between France & Spain





1986 - First experiments in train tunnel 2003-2006 - new lab built 1600m² 2007-2010 - refurbishment works Since 2010 - re-start experimental activities Previous Directors: A. Bettini, A. lanni Inlet air flux: 20000 m³/h Radon: 50-80 Bq/m³ Radon-free: ImBq/m³, 220 m³/h Muons: (5.3+-0.2) ·10-³ m⁻²s-¹ Neutrons: 3.5 ·10-⁶ cm⁻²s-¹







Experiments

Neutrino physics

- NEXT ¹³⁶Xe high-pressure gas TPC (data and construction)
- CROSS Cryogenic Observatory with Surface Sensitivity for CUPID
- SK-Gd & Hyper-K Screening and R&D (e.g., PMT covers)

Dark matter searches

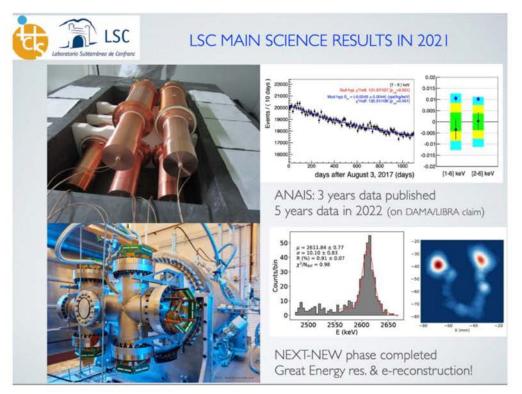
- ANAIS 112 kg Nal crystals (3-yrs data) to verify DAMA-LIBRA
- DaRT Argon activation detector for DarkSide (GADMC)
- TREX Nobel gas (Ne,Ar) high-pressure TPC

Lab Services

HP Ge detectors Electroformed Cu facilities Rn reduction (220 m³/h) Rn monitoring (1 mBq/m³) Rn-emanation detector







Future aims: NEXT: Ton-scale. Hyper-K construction, cryogenic technolgies and misc. low background R&D.



LSM Status & Plans



Laboratoire Souterrain de Modane (LSM)

 Depth: 4800 mwe (Deepest UG Lab in EU)

Volume: 3500 m³

Radon: ~15 Bq/m³

Access: Horizontal

Staff: ~ 13

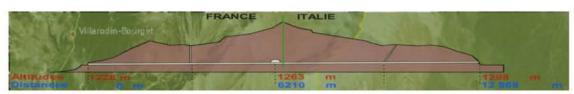
Projects: 7-9





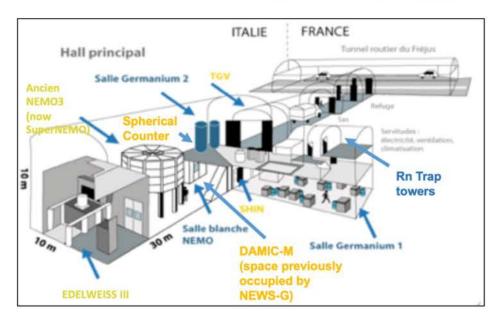












LSM current status and plans

June 2022



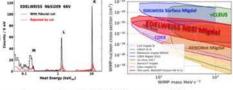


LSM current status and plans

Low-mass Dark Matter

Recent physics results:

 EDELWEISS [arXiv:2203.03993] Migdal limits for <35 MeV WIMPs with 200g Cryo with new NbSi TES phonon sensor



In preparation / coming results:

- CRYOSEL: Cryogenic Ge with single-e⁻ tag using μ-wire NbSi TES sensor: 40 g detector in BINGO cryostat @ LSM (2023)
- DAMIC-M: Search for low-mass Dark Matter with 1kg skipper CCDs
- Start of physics run with 2 CCDs (1 kg.d goal)





LSM current status and plans

Science program adapted to LSM size:

- Low-mass Dark Matter Experiments
- 0vBB demonstrators & technologies
 - HPGe array for low-radioactivity

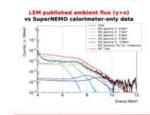
ββ0v demonstrators & technologies

Recent physics results:

- CUPID-Mo 20 x 0.2 kg scintillating
 Bolometers. Mo-100 test, CUPID demo
- [arXiv:2202.08716] New leading
 limits on ¹⁰⁰Mo 0vββ & beta decays

In preparation / coming results:

- BINGO: Development of next generation cryogenic ββ0v technologies with reduced support mass, Neganov-Luke light detectors and active shield
- SuperNEMO installation and commissioning



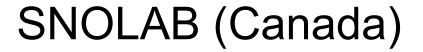
Tracko-Calorimeter detection of bb0n decays with identification of the two electron tracks





June 2022 LSM current status and plans

June 2022







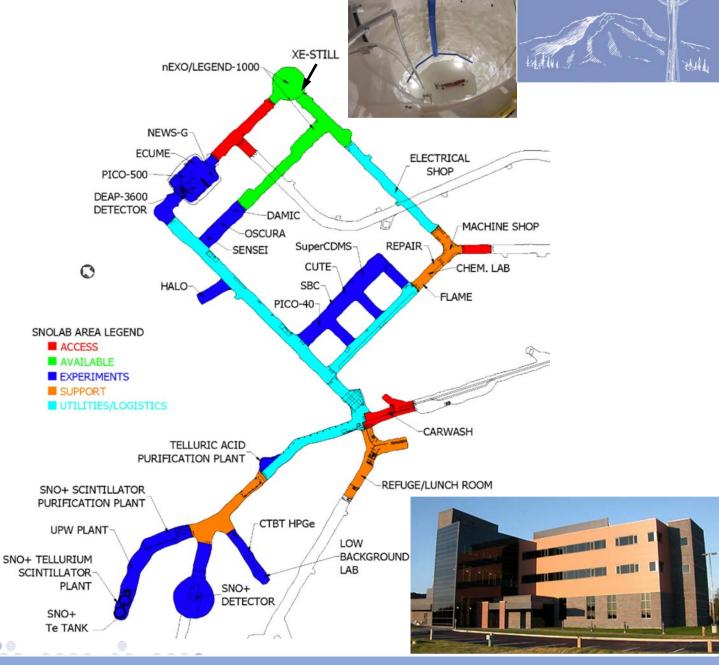


6800' lab layout

 $(6000 \text{ m.w.e.} / 0.3 \mu \text{ m}^{-2} \text{ day}^{-1})$

- Cryopit is allocated to the tonne-scale $0
 u\beta\beta$ program
- Otherwise all available space is occupied and actively managed under SNOLAB's project lifecycle program
- A conceptual design for a lab expansion exists...

Map of the underground facility showing locations of experiments and various ancillary areas.





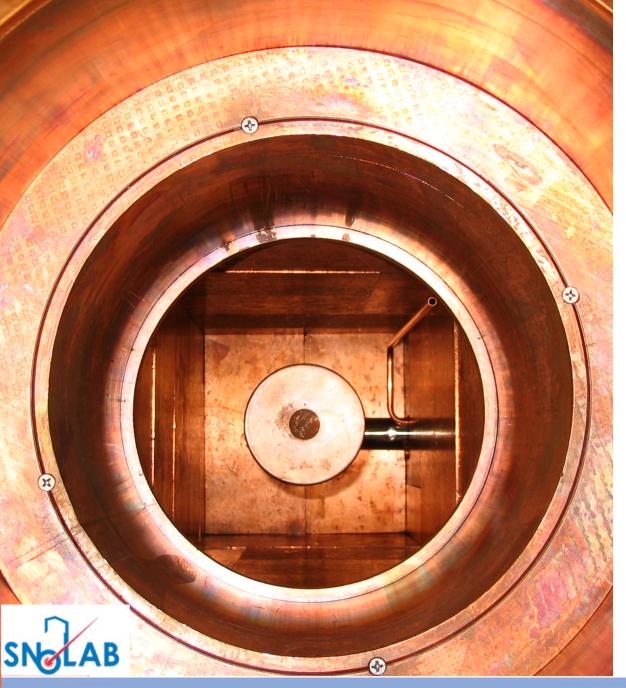
Experiment	Research Focus				Status		
	Dark Matter	Neutrino 0vBB	Neutrino Other	Supernova	Other	Space allocated	Phase
Ar2D2	√				Low-rad argon	LBL	Letter of Intent
ARGUS-I	√	V			Ar storage	4000L	Letter of Intent
COUPP-4	√					Ladder Labs	Completed
CTBT-HPGe					Low background	LBL	In construction
CUTE	√				Test facility	Ladder Labs	Operational
DAMIC	√					J-Drift	Operational
DEAP-1	√					J-Drift	Completed
DEAP-3600	√					Cube Hall	Completed
DEAP-3600-II	√					Cube Hall	In construction
ECUME					Cu electroform	Cube Hall	In design
FLAME					Genomics	Bio/chem lab	Operational
HALO	7			V		Exp. Stub	Operational
LEGEND-1000	3	V				Cryopit	In design
LNG-CTF					Cryogenics	Surface Facility	In design
MiniCLEAN	√					Cube Hall	Completed
MODCC					Mining data	Surface Facility	Completed
nEXO		1				Cryopit	In design
NEWS-DM	√					Cube Hall	Discontinued
NEWS-G	√					Cube Hall	Operational
OSCURA	√					J-Drift	Letter of Intent
PICASSO-III	√					Ladder Labs	Completed
PICO-2L	√					J-Drift	Completed
PICO-60	√					Ladder Labs	Completed
PICO-40L	√					Ladder Labs	Operational
PICO-500	√					Cube Hall	In preparation
PUPS					Seismicity	External drift	Completed
SBC	√					Ladder Labs	In preparation
SENSEI	√		,			J-Drift	Operational
SNO+ (H2O)			√	√		SNO Cavity	Completed
SNO+ (LAB)			V	√		SNO Cavity	Operational
SNO+(Te)		V	√	√		SNO Cavity	In construction
SuperCDMS	V	0	010			Ladder Labs	In construction
REPAIR					Genomics	Bio/chem lab	Operational
Xe-Still					Cryogenics	Cryopit	Operational



SNOLAB Experimental Programme

- 10 operational
- 4 in construction
- 9 in pipeline in various stages
 - letter of intent
 - in design
 - in preparation





SNOLAB is increasing low background screening capabilities





HPGe Detectors:

PGT Coaxial Detector , Canberra Coaxial Detector (**Lively**), Canberra **Well** Detector, Eurisys Mesures Coaxial Detector (**Vue Des Alpes**), Canberra Coaxial Detector (**Gopher**), Canberra Dual HPGe (**CTBT**)

Alpha Counting:

XIA UltraLo-1800

Radon Emanation Measurements:

Electrostatic Counter (ESC)

Radon Emanation Studies using Bronze and Xeolite/charcoal

traps

Radon Board on Water System

77



New Underground LN2 Plant

The LN2 Plant is operational since May 2022, creating high purity LN2 (99.999%) at 3,000 L/week.

It is being tied in to supply SNO+/UPW Plant. It is currently used for cooling down the HPGe detectors and will soon provide it GN2 through the LBL gas distribution system. It is currently supplying the CUTE facility, and soon all users at SNOLAB will use this LN2 system.







SUPL / Yemilab

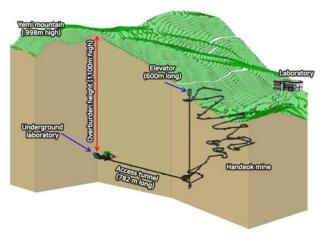
SUPL (Stawell Underground Physics Lab – Australia)

- Being built at depth of 1000 m
- First UG lab in southern hemisphere
- To participate in DM searches

Yemilab (Korea)

- New, being constructed at 1100 m depth
- Deeper than Kamioka but smaller







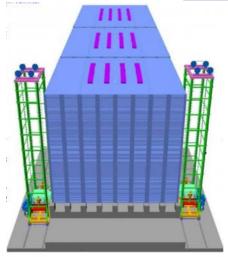
INO (India-based Neutrino Observatory)

- Planned for 1200 m cave
- Construction delayed
- National with international involvement

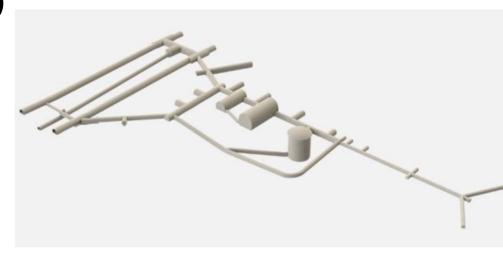
ANDES (Agua Negra Deep Experiment Site – Argentina / Chile)

- Deep: 1750 m rock overburden
- Large: 4000 m², 70000 m³
- Planned as an international laboratory run by a consortium
- Agua Negra tunnel situation unclear





Schematic of a detector at the observatory. Photo: INO website







G3 Dark Matter

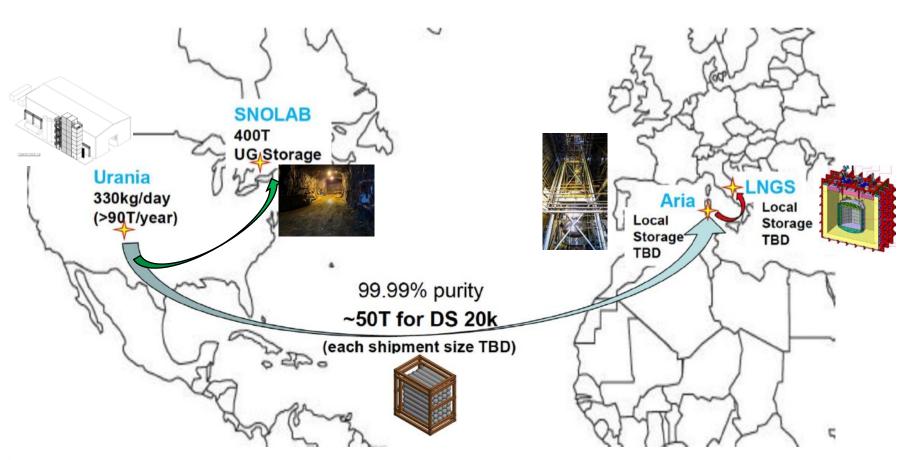
- Liquid Argon-based (GADMC Global Argon Dark Matter Consortium)
 - ARGO LAr 300 tonne to follow Darkside programme
 - SNOLAB is target location
- Liquid Xenon-based (XLZD XENON LUX-ZEPLIN Dark matter consortium)
 - DARWIN LXe 40 tonne-scale
 - Multi-purpose detector

G2 $0\nu\beta\beta$ Decay

- Tonne-scale neutrinoless double-beta decay with enriched isotopes
 - nEXO 5T ¹³⁶Xe
 - LEGEND-1000 1T ⁷⁶Ge
 - Both have SNOLAB as preferred location; LEGEND-1000 has LNGS as alternate site

Storage of Low Radioactivity Argon



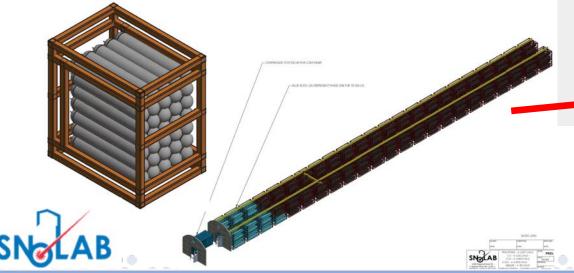


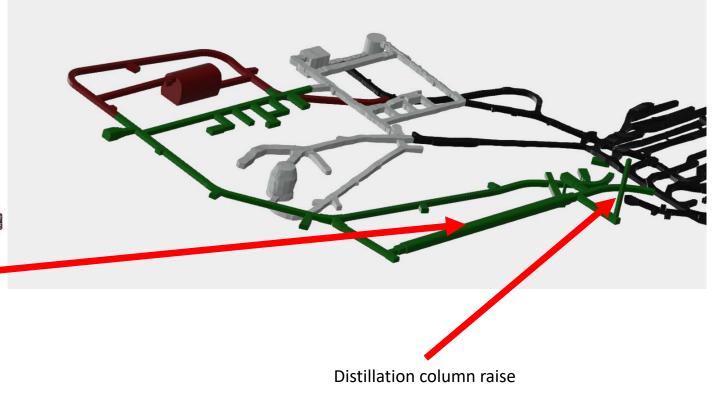
- The GADMC and DarkSide collaborations have build the Urania separation plant to produce argon with very low in ³⁹Ar activity from an underground CO₂ production source.
- With the initial production planned to be shipped to Italy for DS-20k, the plant can then continue production for future requirements.
- There is a unique global opportunity to harvest and store this radio-pure argon for future experiments such as LEGEND, COHERENT, Argo, and DUNE.

SNOLAB Expansion Conceptual Design



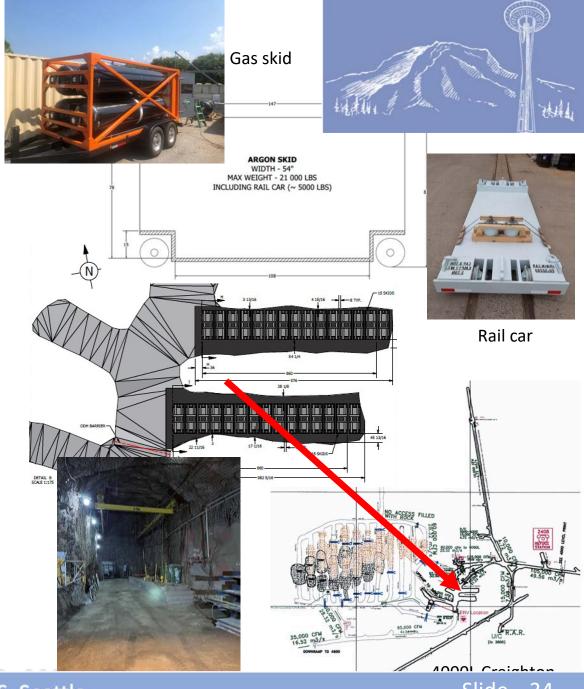
- SNOLAB has completed a conceptual design for the underground lab expansion to create new experiment spaces and underground facilities. SNOLAB is now seeking a funding opportunity for this expansion.
- Part of the expansion design includes a gas storage terminal for up to 400 tonnes noble gas, e.g. argon.
- The design also includes a raise for a noble gas distillation column.





ARGUS

- ARGUS is a project within the GADMC collaboration with the goal to design and procure high pressure argon gas storage and transportation skids, and to store the gas underground at SNOLAB.
- The goal is to store up to 400 tonnes of the low radioactivity argon production from Urania.
- The gas will be stored underground to shield from cosmogenic activation and to be available for next generation experiments (e.g. LEGEND, Argo).
- High pressure gas (500 bar) storage is preferred as storage is compact and low maintenance compared to cryogenic storage.
- Prior to the SNOLAB expansion, the gas skids can be stored at other locations in the mine.
- Location at 4000L level Creighton Mine suitable for ~100 tonnes with minimal development, and can be developed for 400T capacity if required.









- Direct dark matter and $0\nu\beta\beta$ decay searches
 - Increasing sensitivity
 - Larger targets / more isotope
 - Must have commensurate reduction in backgrounds
 - Purer materials / underground manufacturing and storage
 - Advances in purification technologies
 - Advances in low background material assay capabilities
 - Larger shielding / increasing space requirements
 - Improved vetoing / tagging of cosmic rays and spallation products
 - When needed greater depth





• Questions?