Physics at the MeV-Scale in Neutrino LArTPCs



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Snowmass Community Summer Study University of Washington, Seattle July 24, 2022

Liquid argon time projection chambers







LArIAT Data





MeV-scale activity from de-excitation γ's and neutrons



Energy scales in LArTPCs

- Most neutrino reconstruction tools tailored for higher-energy, GeV-scale tracks and EM showers found in v-Ar final-states
- But LArTPCs are sensitive to sub-MeV-scale physics too!

Low-Energy Physics in Liquid Argon (LEPLAr)

- Snowmass workshop held in 2020
 - Identify opportunities for DUNE in the <100 MeV regime
 - Develop standard set of signal/background assumptions
 - Exchange ideas between DUNE technical working groups







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Reconstruction in a nut-shell





- Signals time-matched between wire readout planes
 - Wire intersections \rightarrow YZ coordinate
- Easy with extended (multi-hit) signals
- More challenging at lower energy
 - Hit-finding thresholds
 - Noise hits create ambiguous fake matches
- Standardized toolset/algorithms under development for broad LArTPC use



Demonstrations of MeV-scale signal sensitivity



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See talk by Thiago right after this





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Simulation studies

Generic 'DUNE-like' MC used to explore some of these benefits

Supernovae neutrinos

- 50% more v energy recovered when including isolated blips
- **Channel ID capabilities**
- Tagging neutron-producing evts
 - Recover 7.8 MeV of lost v energy (n-separation energy in K⁴⁰)

30

True Neutrino Energy (MeV)

40

Phys Rev D 102, 092010 / arXiv:2006.14675

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[Submitted on 25 Jun 2020]

Benefits of MeV-Scale Reconstruction Capabilities in Large Liquid Argon Time Projection Chambers

W. Castiglioni, W. Foreman, I. Lepetic, B. R. Littlejohn, M. Malaker, A. Mastbaum



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New Reconstructed Energy (MeV)

Simulation studies



n



Summary

- MeV-scale reconstruction is the 'next frontier' in neutrino LArTPC physics!
- The <u>LEPLAr Snowmass White</u> <u>Paper</u> presents a broad and detailed overview of the applications and challenges ahead
 - Much more than I could include in this talk

Thank you!





Backup

Related references

<u>Study of reconstructed 39Ar Beta Decays at the MicroBooNE Detector</u>

Snowmass contributions:

- Low-Energy Electron-Track Imaging for a Liquid-Argon Time-Projection Chamber using Probabilistic Deep Learning (poster), Micah Buuck et. al., <u>arXiv:2207.07805</u>
- Improving LArTPC Performance with Photo-Ionizing Dopants, Joseph Zennamo
- Low-Energy Physics Opportunities with DUNE, Daniel Pershey
- <u>DUNE-Beta: Searching for Neutrinoless Double Beta Decay with a Large LArTPC</u> (LOI), Joseph Zennamo, Fernanda Psihas, Andy Mastbaum
- LArTPC Pixelated Readout, Brooke Russel



Photo-ionizing dopants in LAr for improved response at low energies



Improving LArTPC Performance with Photo-Ionizing Dopants, Joseph Zennamo



LArTPC example event display





Calculating energy from charge



For *tracks,* we know the length and therefore dQ/dx for each hit.



For *blips*, no spatial extent: we lose that "dx" information!



Calculating energy from charge

- MicroBooNE (<u>arXiv:1704.02927</u>) and LArIAT (<u>arXiv:1909.07920</u>)
 - Analyses of Michel electron showers
 - For blips, assumed constant dE/dx (i.e., constant recombination)
- ArgoNeuT (<u>arXiv:1810.06502</u>)
 - Nuclear de-excitation γ analysis
 - Used NIST data on low-E electrons, together with recombination, to directly relate measured Q to energy





Ar39 contamination in a large LArTPC

- Produced from cosmic ray exposure, present at 1 Bq/kg
- β decay, Q value of 0.565 MeV
- Randomly distributed background of blips

Simulated decays in a DUNE-sized drift region (2.2ms data acquisition window).

For randomly-selected point in fiducialized active volume, using 75 keV blip threshold, contribution from Ar³⁹ in 30cm sphere:

- Energy ~ 0.08 MeV
- Energy RMS spread ~ 0.15 MeV

N_{blips} ~ 0.3



Effect of electronic noise



FIG. 20. The resolution of the full-energy peak for simulated 1.46 MeV γ -rays, over a range of different blip smearing levels, for both 75 keV and 150 keV energy thresholds. A proximity requirement of 30 cm is used. Resolution is calculated based on the FWHM of the peak using the relationship to standard deviation: $\sigma = FWHM/(2\sqrt{2 \ln 2})$.



Final-state neutron ID and calorimetry

Adding up "blips" within 60 cm of neutron production point...



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