Big Ideas 3:

New flavors and rich structures in dark sectors

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Logistics

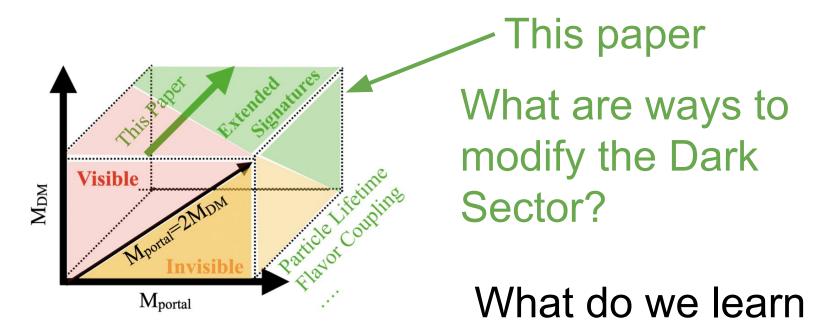
- Current draft of the paper is located here:
 https://www.dropbox.com/s/gksd3y43k0vtpyw/Snowmass_RF6_Big_Idea_3.pdf?dl=0

 https://arxiv.org/abs/2207.08990
- Please send any comments to :
 - <u>pcharris@mit.edu</u> <u>schuster@slac.stanford.edu</u> <u>zupanje@ucmail.uc.edu</u>
- In the next set of slides we will review the paper
 - Focus will be on the plots and table
 - You can read the paper if you want to look at the text

This talk is a summary of what we have finally ended up with

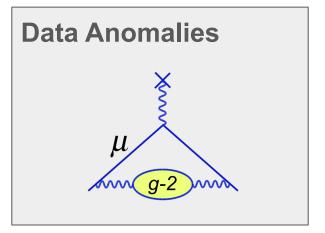
Framing the

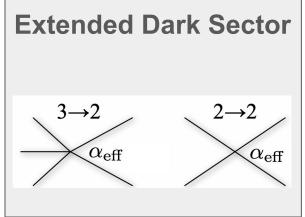
- Goal of this paper: emphasize beyond baseline Dark Sectors
- Many reasons to modify the dark sector to resolve other problems

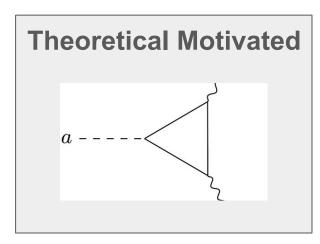


Organization of the paper

3 good reasons to motivate an extended dark sector







Well motivated extensions of baseline dark sector models

Well motivated data anomalies

- There are a number of anomalies that could explain DM
- (g-2)_µ
 - B Physics anomalies
 - RK and other flavor anomalies
- Xenon 1T
- MiniBoone/µBoone Excess
- Neutron lifetime anomaly
- KOTO

Our Highlight

Brian Batell

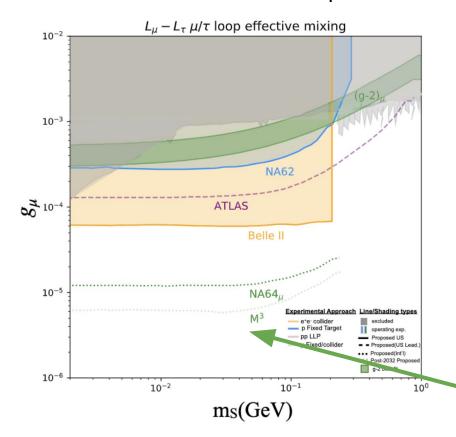
How do you search for (g-2)_µ

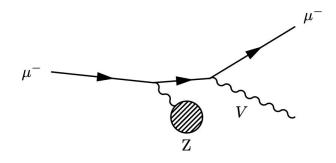
	Invisible			Visible					
Final State/	Long-	Neutrinos	DM	Photons	Electrons	Muons	Hadrons		
Mediator	Lived	(u u)	$(\chi\chi)$	$(\gamma\gamma)$	$(e^{+}e^{-})$	$(\mu^+\mu^-)$	$(\pi^+\pi^-)$		
Vector	No	Yes	Yes	No	No	Yes	No		
						$m_V>2m_\mu$			
	 L_μ - L_τ gauge boson: UV complete, automatic coupling to neutrinos, easy to couple to DM. Challenging to build viable models with sizable couplings of vector mediator to electrons or hadrons (gauge anomalies, constraints from neutrino physics). 								

How do you search for (g-2),

Invisible			Visible				
Long-	Neutrinos	DM	Photons	Electrons	Muons	Hadrons	
Lived	(u u)	$(\chi\chi)$	$(\gamma\gamma)$	$(e^{+}e^{-})$	$(\mu^+\mu^-)$	$(\pi^+\pi^-)$	
Yes	Yes	Yes	Yes	Yes	Yes	Yes	
$m_S < 2m_u$			$m_S < 2m_u$	$m_S < 2m_u$	$n_S > 2m_{\scriptscriptstyle H}$	$m_S>2m_\pi$	
 All minimal signatures can be realized in scalar simplified models. UV complete models require new SM-charged states above weak scale with special flavor structure (such states can in principle also affect (g - 2)_μ). More phenomenological studies needed to chart the parameter space. 							
Miss	sing Energy		Pror	mpt or Displaced Resonance			
	Long- Lived Yes ms < 2mu All r plifie UV above state More para	Long- Lived Neutrinos $(\nu\nu)$ Yes $m_S < 2m_u$ Yes • All minimal sign plified models. • UV complete m above weak scal states can in pri • More phenomen	Long- Lived Neutrinos $(\nu\nu)$ DM $(\chi\chi)$ Yes $m_S < 2m_u$ Yes Yes • All minimal signatures plified models. • UV complete models reabove weak scale with states can in principle at • More phenomenological parameter space.	Long-Lived Neutrinos $(\nu\nu)$ DM $(\chi\chi)$ Photons $(\chi\chi)$ Yes Yes Yes Yes $ms < 2m_u$ • All minimal signatures can be real plified models. • UV complete models require new above weak scale with special flav states can in principle also affect (grammater space). • More phenomenological studies ne parameter space.	Long-Lived Neutrinos $(\nu\nu)$ DM $(\chi\chi)$ Photons (e^+e^-) Yes Yes Yes Yes $m_S < 2m_u$ Yes Yes Yes $m_S < 2m_u$ $m_S < 2m_u$ $m_S < 2m_u$ • All minimal signatures can be realized in scalar plified models. • UV complete models require new SM-charged above weak scale with special flavor structure states can in principle also affect $(g-2)_{\mu}$). • More phenomenological studies needed to charparameter space.		

Vector mode for $(g-2)_{\mu}$

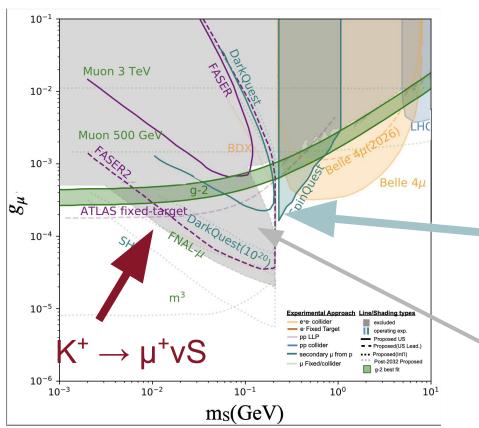


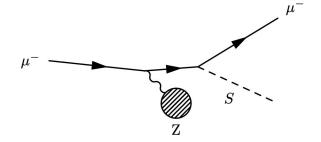


Belle II can close the story by next snowmass

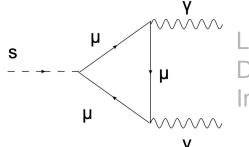
Muon beams drive these searches

Data Motivated Anomalies => (g-2)_u





Secondary Muons from proton beam dumps to appear here too

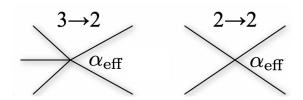


Light mediators
Decay to photons
In a look

Simple extensions of the standard model

SIMP:

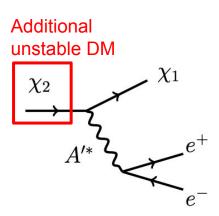
+3→2 Dark Sector interactions



Scenario has dark mesons (rich dark sector)

Light mediator naturally exists w/SIMP for stable cosmology

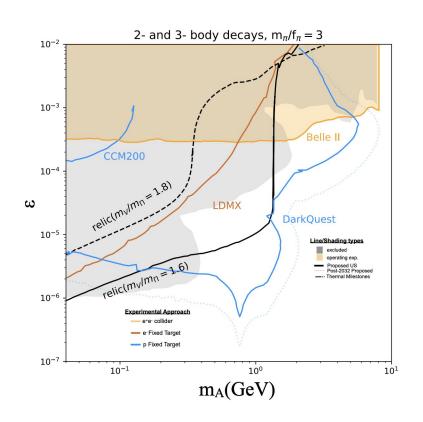
Inelastic DM

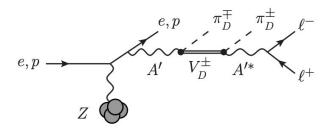


Inelastic DM (Simple extension)

Small mass splitting added for two dark sector particles Leads to unstable (LL) χ_2

Simple extensions of the standard model: SIMP



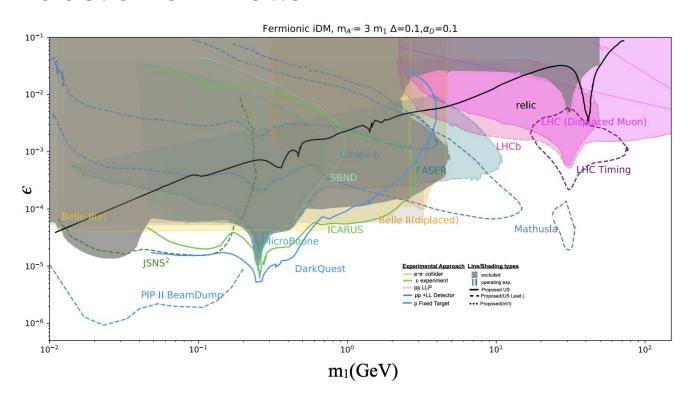


Final state is a displaced dilepton w/intermeidate dark sector particles

SIMP models benefit from small beam dumps (lifetime vs intensity)

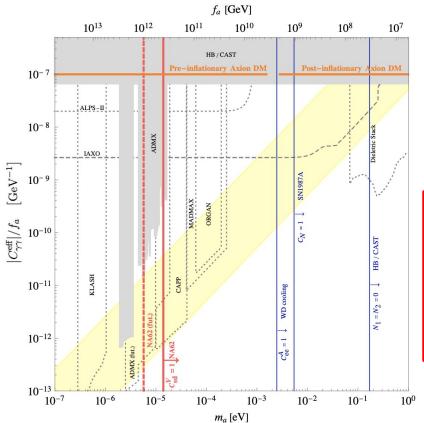
Relevant cosmologies are within reach

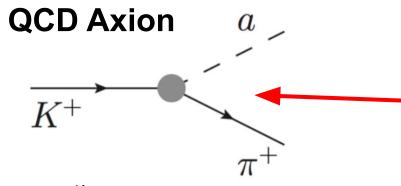
Inelastic Dark Matter



Relic density yields a slightly larger coupling compared to dark photon

Theoretical Motivations





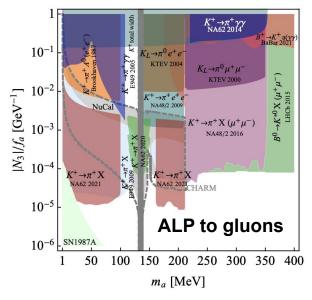
$$\mathcal{L}_{\text{ALP-}f} = \frac{\partial_{\mu} a}{2f_a} \, \bar{f}_i \gamma^{\mu} \left(C_{f_i f_j}^V + C_{f_i f_j}^A \gamma_5 \right) f_j$$

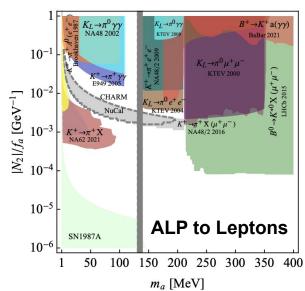
$$C_{d_id_i}^V - C_{d_id_i}^A = V_{u_kd_i}^* (C_{u_ku_l}^V - C_{u_ku_l}^A) V_{u_ld_j}$$
 $V_{u_ku_{d_i}} = \begin{pmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix}
ightarrow ext{for } \begin{pmatrix} u \\ d \\ \ell \end{pmatrix}$

Adding an non-flavor diagonal CKM matrix

More General Axions

$$\mathcal{L}_{\text{ALP-gauge}} = \frac{N_3 \alpha_s}{8\pi f_a} a G^a_{\mu\nu} \tilde{G}^{a\mu\nu} + \frac{N_2 \alpha_2}{8\pi f_a} a W^i_{\mu\nu} \tilde{W}^{i\mu\nu} + \frac{N_1 \alpha_1}{8\pi f_a} a B_{\mu\nu} \tilde{B}^{\mu\nu}$$





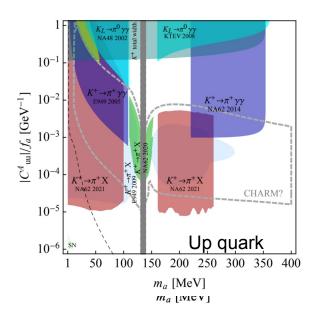
Kaon decays drive the constraints

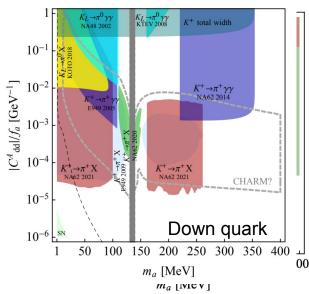
Probing a scale fo 10 TeV

More General Axions

$$\mathcal{L}_{\text{ALP-}f} = \frac{\partial_{\mu} a}{2f_a} \, \bar{f}_i \gamma^{\mu} \left(C_{f_i f_j}^V + C_{f_i f_j}^A \gamma_5 \right) f_j$$

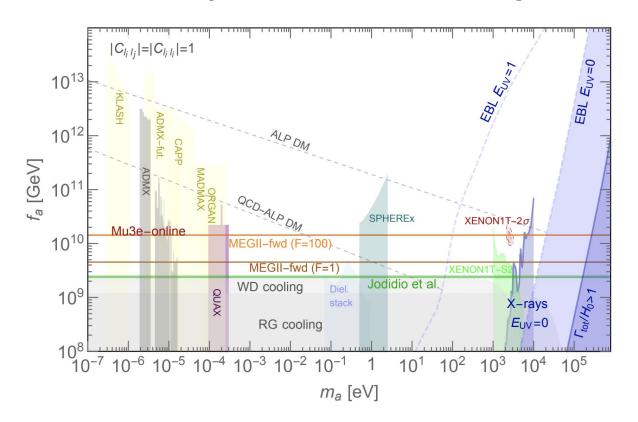
Loop induced di-photon final state

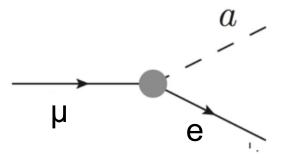




Flavor diagonal bounds on coupling (less constraints)

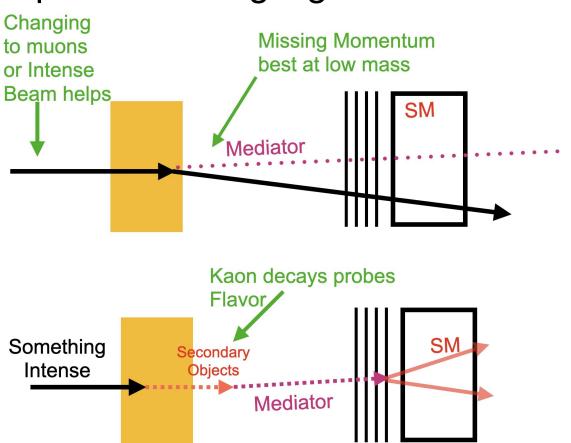
Axions with just lepton couplings



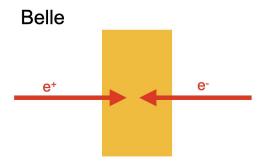


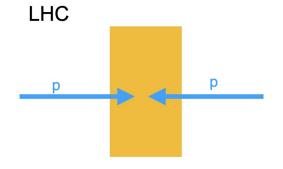
Muon to axion decays

Experimental highlights

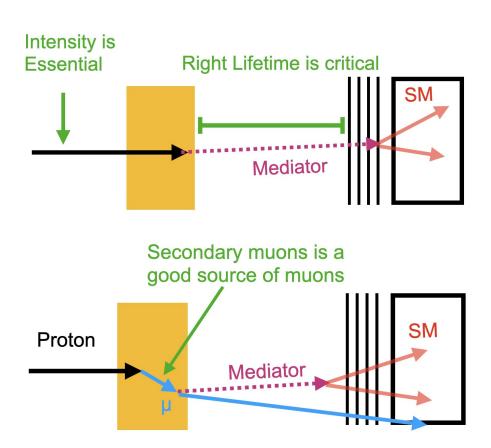


Higher Energy

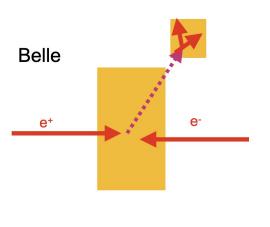


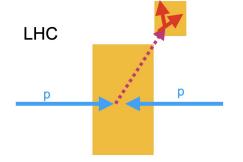


Experimental highlights



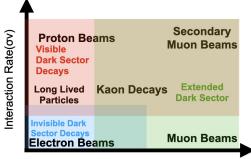
Higher Energy





Conclusions

- Extending the dark sector highlights specific approaches
 - Our focuses:
 - g-2 : (Secondary) Muon beams
 - Extend Dark sector : beam dumps (LLPs)
 - Axions : Kaon decays
- Keep in mind Rich Dark sector
 - Highlights certain tech
- Many opportunities this decade!
 - Modest upgrades probe critical space



Flavor Specific

Backup

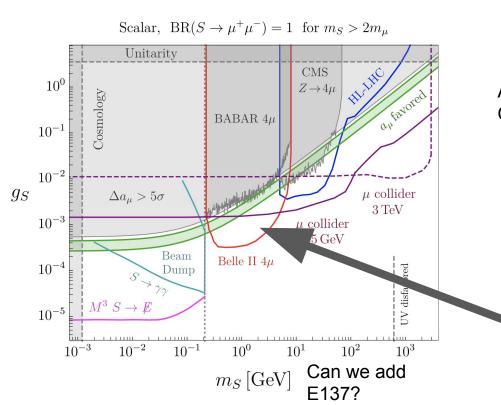
Experimental highlights

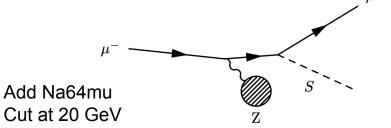
- A few general conclusions follow in what is sensitive
 - Long Lived particles enhance these models
 - Beam dump decay size is critical
 - Increased importance of non-resonant processes
 - Importance of muon beams
 - Direct and Secondary enhance rich models
 - Increased importance of Heavy Flavor
 - Kaon decay play an important role

Strategy

- Aimed to cover 3 major motivations for extended dark sectors
 - Data based anomalies
 - Simple Motivated extensions
 - Theoretical motivations
- Picked a few examples that highlight these
 - No intention of being exclusive
 - Critical point is what we chose to highlight
- We should aim to harmonize plotting styles across papers
 - Not clear which main plots should come from BI3

Data Motivated Anomalies => $(g-2)_{\mu}$

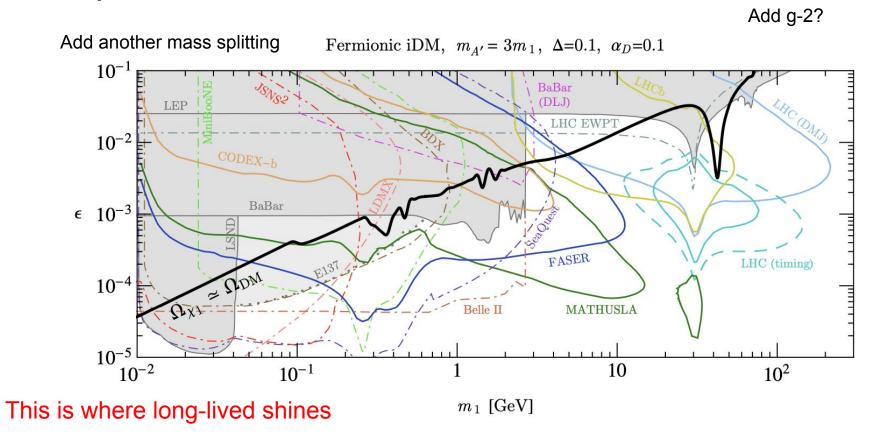




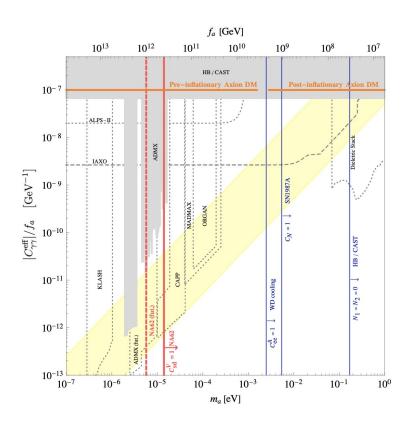
General Message, this is where Muon beams shine $(L_{_{\text{\tiny II}}}-L_{_{\text{\tiny T}}})$ in extended

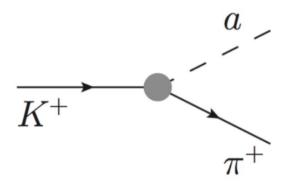
Secondary Muons from proton beam dumps to appear here too

Simple extensions of the standard model



Theoretical Motivations

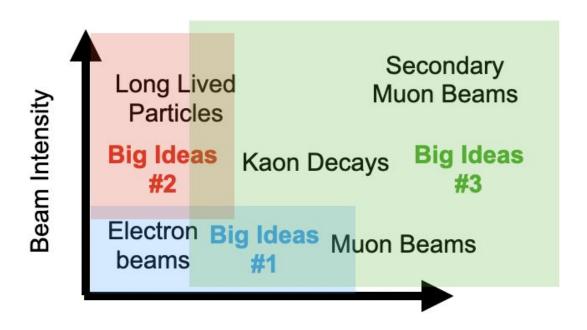




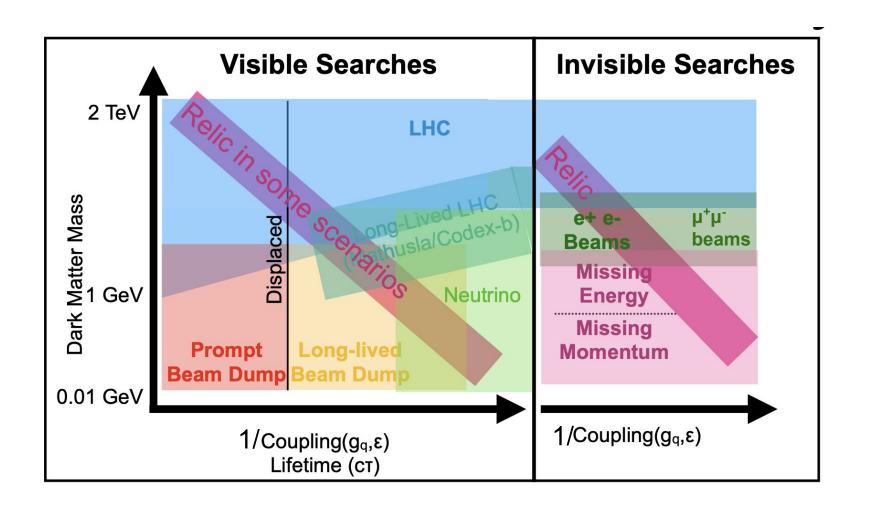
Connecting Kaon decays with the standard Model Axion

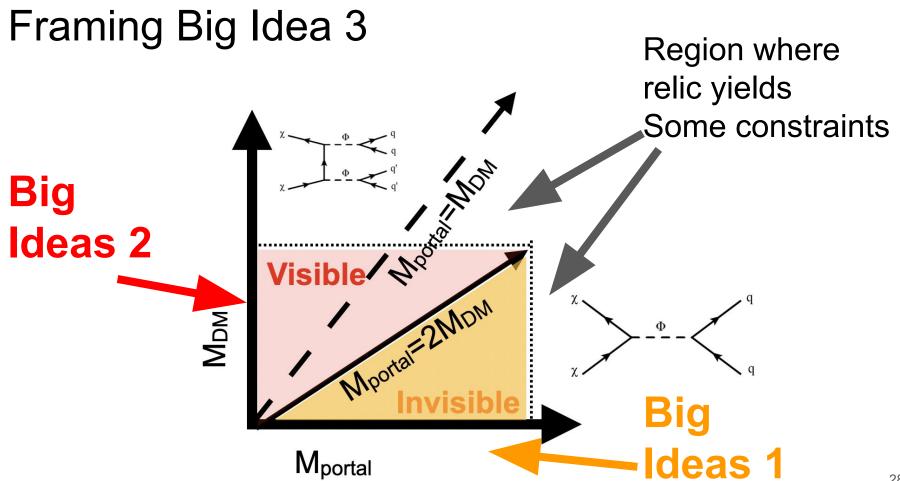
(Not really like the other plots)
Missing a huge number of lines

Experimental Front



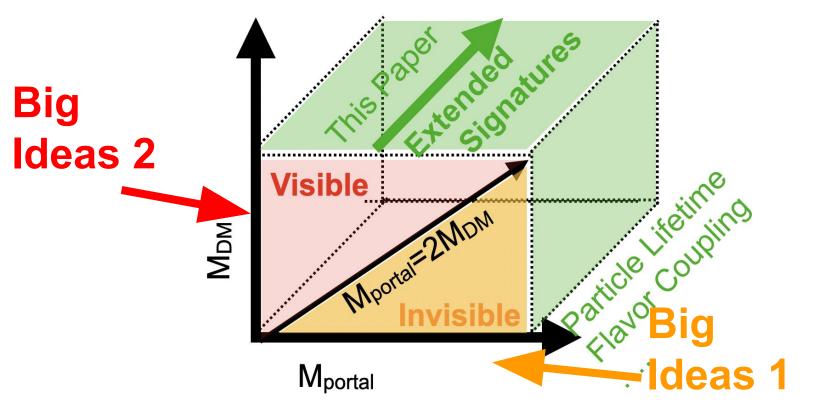
Flavor Specific





Framing Big Idea 3

Big Ideas 3



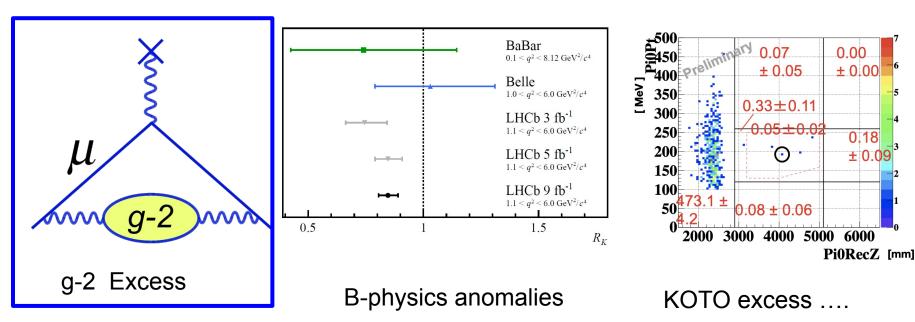
Flavor Violating axions

- One can envision the QCD axion as a dark matter candidate
 - Axion couplings are very sensitive in Kaon decays
 - Can probe a very high scale in axions (Dim 5)
 - Lead to a probe of the QCD Axion
- X

$$\mathcal{L}_{\text{ALP-photon}} = \frac{\alpha}{8\pi f_a} C_{\gamma\gamma}^{\text{eff}} a F_{\mu\nu} \tilde{F}^{\mu\nu}$$

$$C_{\gamma\gamma}^{\text{eff}}(m_a) = N_1 + N_2 + \sum_q 6 Q_q^2 C_{qq}^A(\mu_0) B_1(\tau_q) + 2 \sum_\ell C_{\ell\ell}^A B_1(\tau_\ell)$$

Big Idea 3: Data-driven Motivations

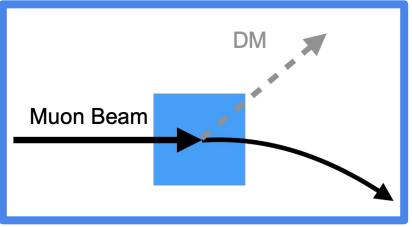


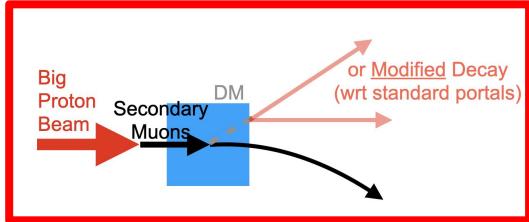
Our Anomaly of choice!

Recent list of many anomalies

How do you explain g-2_µ excess?

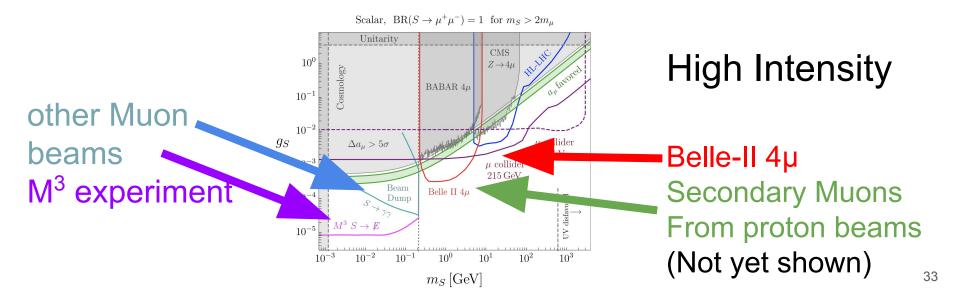
- Safest way to probe g-2_µ excess is with muons
- 3 Good sources of muons
 - Muon beams, secondary muons, or intense e⁺e⁻ collider(Belle)





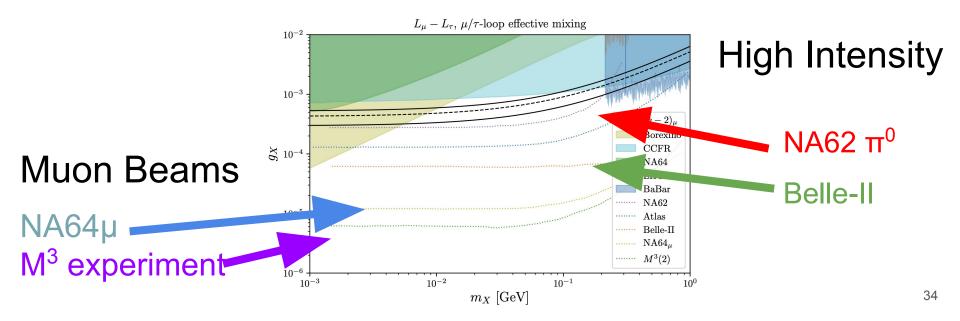
How do you explain g-2 excess?

- In minimal scenario consider just a muon coupling
 - Can explain this through a scalar coupling



How do you explain g-2 excess?

- With a more complicated scenario
 - Can consider an explanation with Lµ-LT

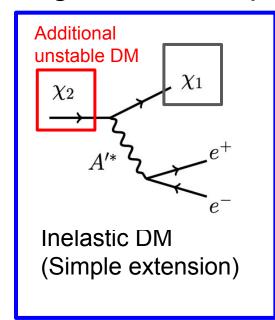


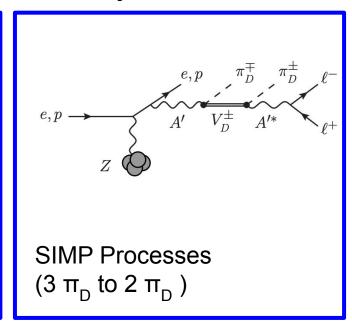
Other anomalies that we touched on

- $(g-2)_{\mu}$
- B Physics anomalies
 - RK and other flavor anomalies
- Xenon 1T
- MiniBoone Excess
- Beryllium
- Neutron lifetime anomaly
- KOTO

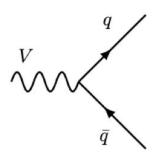
Discussion: Have we missed one? (W mass?)

Big Idea 3: Simple Theory Extensions





 $V \to \text{hadrons}$



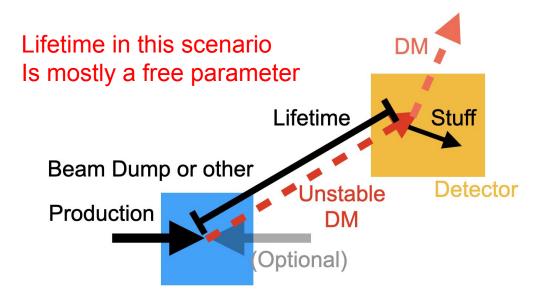
Hadrophilic DM
To avoid existing bounds

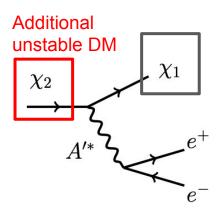
We highglight two models

Strategy is to consider simple well motivated models

Inelastic DM: Example of a minimal extension

- Start with the usual portal
 - Add an unstable DM candidate





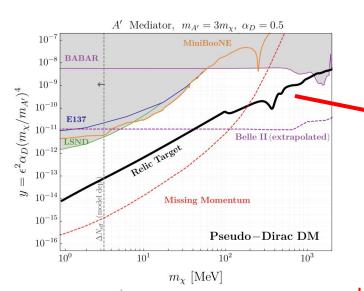
Inelastic DM generally does not change the relic bounds much, but adds new signatures

Inelastic DM: Example of a minimal extension

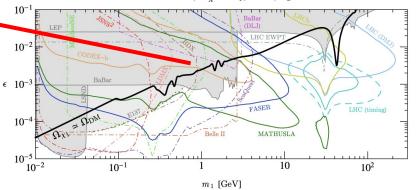
By adding new signatures we make it easier to detect

Simple extension of the model changes bounds a lot

A case where many detectors can play a role



Relic Targets are roughly the same. However bounds from experiments differ greatly. Fermionic iDM, $m_{A'} = 3m_1$, $\Delta = 0.1$, $\alpha_D = 0.1$



Additional

 χ_2

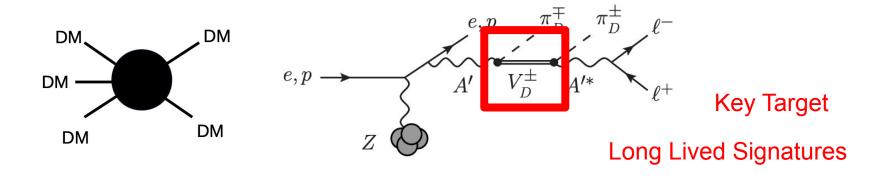
unstable DM

 χ_1

+Strong Bounds from LHC and Beam dump

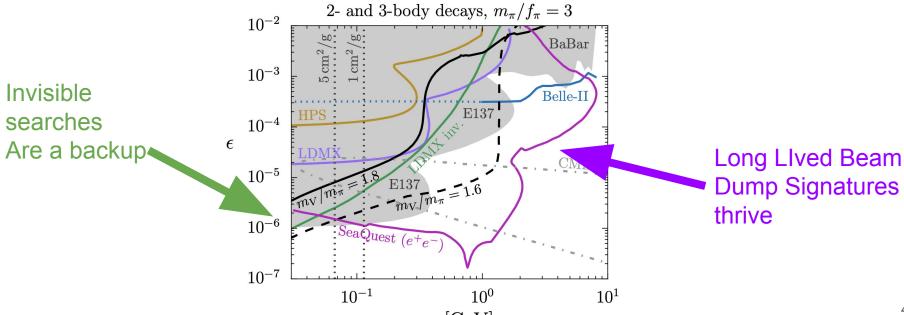
SIMP DM model

- Strongly Interacting Massive Particle (SIMPs)
 - Dark sector has a SU(N_c) dark that explains DM
 - Yields Dark QCD particles (π_D, V_D) akin to $(\pi, \omega, \varphi, ...)$
 - Originally motivated by 3 to 2 processes, which gives relic
 - Naturally forces a DM scale of 100 MeV
 - Also motivates long lived signatures naturally from model

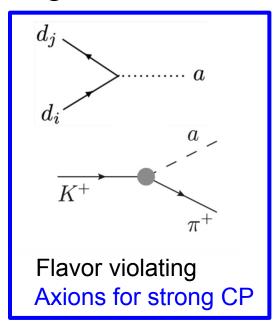


SIMP DM model: Bounds

- Core feature is the long lifetime signature
 - Potentially exploration with LHC based searches can help

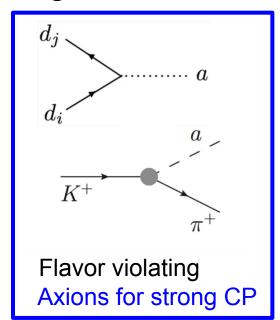


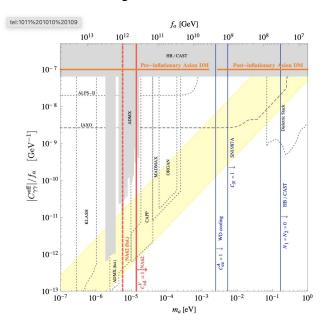
Big Idea 3: Motivated Theory Extensions



Can we explain core theory features with extended model

Big Idea 3: Motivated Theory Extensions





$$K \to \pi a$$

Measuring Kaon decays probes flavor structure

Can we explain core theory features with extended model

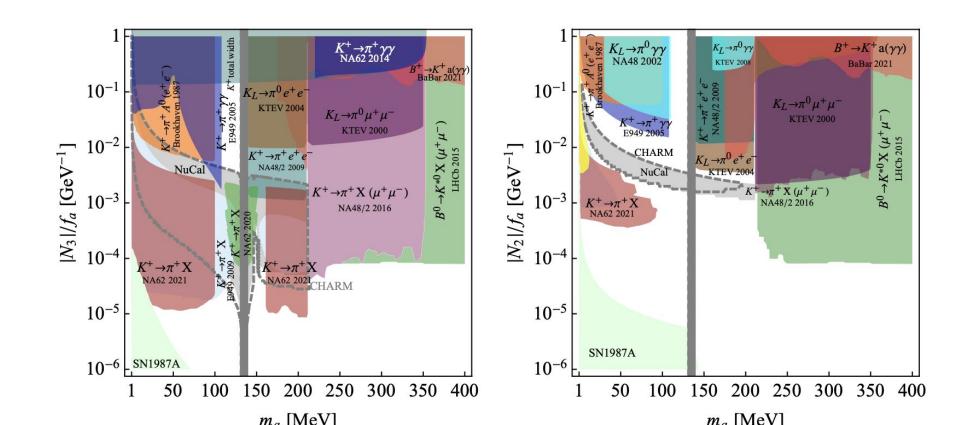
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More Axions



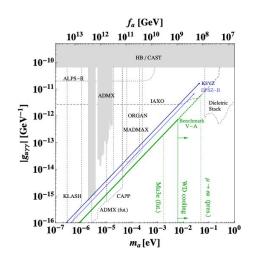
What needs to be done

- Add Mike's messaging suggestions:
 - exploiting capabilities of existing large multi-purpose detectors + near-future upgrades;
 - Investing in specialized small-scale experiments + Investing in facilities that can deliver high-intensity beams;
 - Supporting the theory community to further develop theory of dark sector
- Additional Suggestions from Stefania
 - Is there a way to highlight more BSM physics?
 - Highlight secondary muon beams

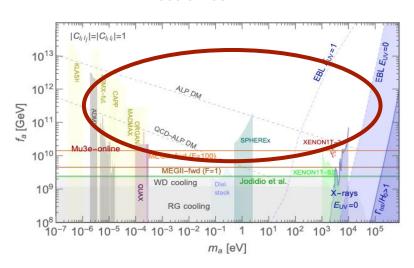
Backup

Lepton flavor violating ALP as dark matter candidate

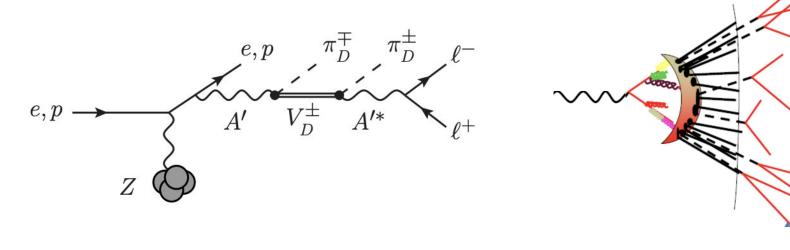
- If light enough ALP can be a DM candidate
 - If flavor violating couplings, can probe very high scales
 - For instance, for LFV couplings $\mu \rightarrow e$ a
 - Could be the QCD axion



2006.04795



Big Idea 3: Theory Motivations



Or just a full on question what happens when things get complicated?

Are we covering all signatures?

Big Idea 3

Benchmarks in Final State x Portal Organization

	DM Production	Mediator Decay Via Portal	Structure of Dark Sector
ecto		m _{A'} vs. € [decay-mode agnostic] m _{A'} vs. € [decays]	iDM m_χ vs. y [$m_{A'}/m_\chi$ =3, α_D =.5] (anom connection) SIMP-motivated cascades [slices TBD] U(1) _{B-L / μ-τ / B-3τ} (DM or SM decays)
llar	Schould ctill include)	m _S vs. sinθ [λ=0] m _S vs. sinθ [λ=s.t. Br(H→φφ ~10 ⁻²)]?	Dark Higgssstrahlung (w/vector) scalar SIMP models? Leptophilic/leptophobic dark Higgs?
Neutrino	e/μ/τ a la1709.07001?	$m_{ m N}$ v s. $U_{ m e}$ $m_{ m N}$ v s. $U_{ m \mu}$ $m_{ m N}$ v s. $U_{ m au}$ Think more about reasoanble flavor structures	Sterile neutrinos with new forces?
ALP	m_χ VS. fq/l [λ =0, fix m_a/m_χ , g_D] (thermal target excluded) What about f _y , f _G ?	$m_{\rm a}$ vs. $f_{\rm V}$ $m_{\rm a}$ vs. $f_{\rm G}$ $m_{\rm a}$ vs. $f_{\rm q}$ = $f_{\rm I}$ (separate?) Think more about reasoanble coupling relations including $f_{\rm W/Z}$	FV axion couplings

Neutron portal? Hidden valleys (or are these out-of-scope?)? See e.g. 2003.02270

What distinguishes Big Idea 3 from the rest?

- *Framing Big Idea 3:* Even if dark sectors have non-minimal structures, either in couplings to the standard model, or the dark sector spectra of states, they can still be efficiently searched for in high intensity experiments.
- **Framing Big Idea 1:** Dark matter particles can be observably produced at intensity-frontier experiments, and opportunities in the next decade will explore important parameter space motivated by thermal DM models, the dark sector paradigm, and anomalies in data.
- Tentative Framing of Big Idea 2: Light, weakly coupled mediators to a dark sector can be copiously produced in high-intensity experiments and detected through their decays to Standard Model particles. Existing, planned, and proposed experiments offer great potential to discover the mediator and discern the pattern of its interactions with ordinary matter.

Points from Stefania

- 1. Can we highlight more theory BSM anomalies
- 2. Point out that SIMPs introduce a scale of 100 MeV naturally
- 3. Make a stronger statement about iDM
- 5. SIMP Naturally predct LLPs
- 6. Highlight secondary muon beams from future proton beam dumps

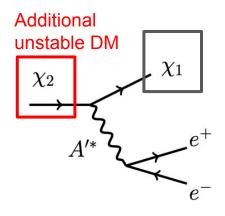
Points from Mike

* exploiting the capabilities of existing large multi-purpose detectors, including near-future upgrades;

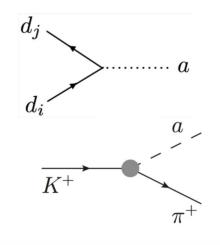
* investing in specialized small-scale experiments, and in facilities that can deliver high-intensity beams;

* and supporting the theory community to further develop the theory of dark-sector physics.

Big Idea 3: Theory Motivations

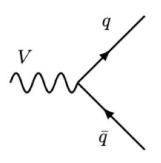


Inelastic DM (Simple extension)



Flavor violating
Axions for strong CP

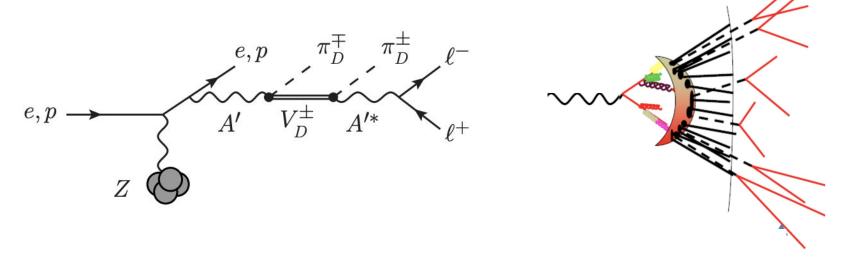
 $V \to \text{hadrons}$



Hadrophilic DM
To avoid existing bounds

Additionally theoretical motivations to make things more complicated

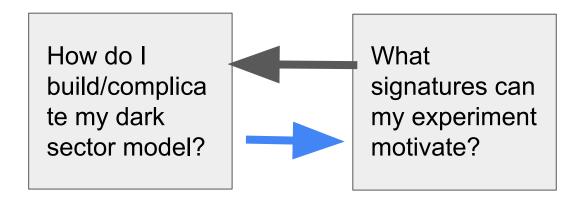
Big Idea 3: Theory Motivations



Or just a full on question what happens when things get complicated?

Are we covering all signatures?

Additional Backdrop



Small extensions of models can lead to many new signatures at experiments

Being experimentally prepared for well motivated extensions should be goal of next generation experiments

Proposed Outline of the paper

- Framing of Big Idea 3 amongst the other Big Ideas
- Big Idea 3 Motivations and Scope
 - Extended Dark Sector
 - Physics Anomalies
 - Additional Motivations
- Benchmark Models and Motivations
 - Review of the LOIs and how they fit into their papers
- Mapping of models to Experimental Approaches
 - Full discussion of experimental approaches will go in Big Ideas 4

Motivating Models for Big Idea 3

- Extended/non-minimal dark sectors
 - Minimality is good but may not correspond to reality (example: SM)
- Which models to present?
 - Strongly motivated theoretically (example: flavor violating QCD axion)
 - Highlighting experimental reach:
 - What do we learn about a model if nothing is found (example: closing the prompt decaying ALP window below kaon mass)
- Finding Benchmarks for all models can be a challenge
 - A key point of the Big Ideas 3 is to collect the motivations for extension
 - Allows us to have a benchmark

Data Anomalies that help promote Big Idea 3

- (g-2)_u
- B Physics anomalies
 - RK and other flavor anomalies
- Xenon 1T
- MiniBoone Excess
- Beryllium
- Neutron lifetime anomaly
- KOTO

Discussion: How complicated a model is needed to motivate these?

How complicated a detector is needed to find these models?

Examples

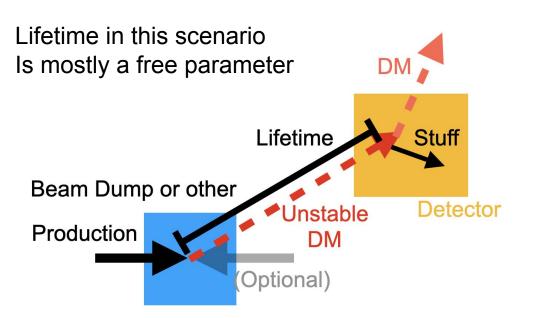
Questions to keep in mind for discussion

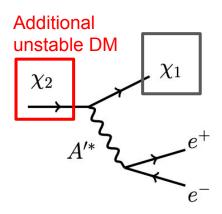
How do we prepare ourselves for newer, more complicated models?

Discussion: What do we define as complete or adequate signature coverage?

Inelastic DM: Example of a minimal extension

- Strategy start with one of our usual portals:
 - Modify the Dark Matter to have an unstable candidate





Inelastic DM generally does not change the relic bounds, but adds new signatures

Inelastic DM: Example of a minimal extension

 $m_{\chi} [\text{MeV}]$

Additional By adding new signatures we make it easier to detect unstable DM Simple extension of the model changes bounds a lot χ_1 χ_2 A case where many detectors can play a role Relic Targets are roughly A' Mediator, $m_{A'} = 3m_{\chi}$, $\alpha_D = 0.5$ the same. However 10^{-7} MiniBooNE bounds from experiments 10^{-8} BABAR differ greatly. Fermionic iDM, $m_{A'} = 3m_1$, $\Delta=0.1$, $\alpha_D=0.1$ 10^{-9} $\epsilon^2 \alpha_D (m_\chi/m_{A'})^4$ E137 Belle II (extrapolated) 10^{-12} € 10⁻³ FASER LHC (timing) Missing Momentum 10^{-4} 10^{-15} MATHUSLA Pseudo-Dirac DM 10^{-16} 10 10^{-1} 10^{2} 10^{-2} 10^{0} 10^{2} 10^{3}

 m_1 [GeV]

Long Lived particles as an extension of dark sector

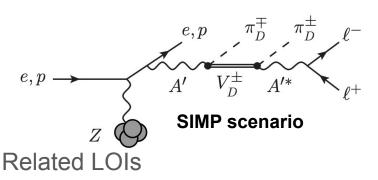
- Adding an unstable DM is a simple extension to dark sector
 - Leads to a wealth of other new final states
- More flexibility of signatures than just portals with small couplings

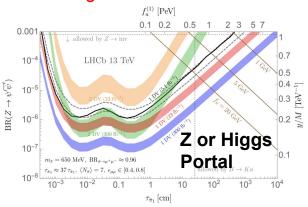
Related LOIs that cover this

- BDF/SHiP facility AF5_AF0-RF6_RF0-163
- LLP at Energy Frontier EF9_EF10-RF6_RF0-TF7_TF8_James_Beacham-201
- LLP at FCC-ee EF8_EF9-RF6_RF0_Rebeca_Gonzalez_Suarez-147
- New light particles at ILC main beam dump EF9_EF0-RF0_RF6-086.
- Codex-b EF9_EF0-RF6_RF0-034
- Electron fixed target spectrometer : HPS RF6_RF0_Nelson-078
- Proton fixed target spectrometer : DarkQuest RF6_RF0_Nhan_Tran-025

More complex Dark sector: Dark Pion

- Dark Pions, Dark Vector Mesons as a further extension
 - More complicated Dar Sector Scenario
 - Broad range of complex decays many of them can be long lived



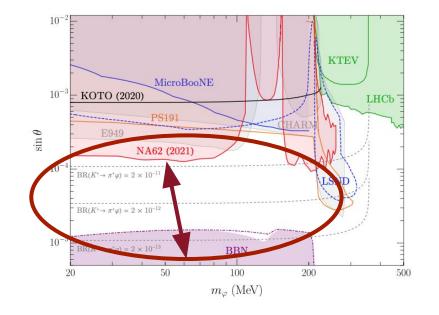


- Dark Pion Searches at LHC and High Intensities EF9_EFU-RF6_RFU-075
- LDMX RF6_RF0-EF10_EF0-CF1_CF0_Andrew_Whitbeck-104
- Electron fixed target spectrometer: HPS RF6_RF0_Nelson-078
- Proton fixed target spectrometer : DarkQuest RF6_RF0_Nhan_Tran-025

Connecting Portals with bounds: Higgs mixed scalar

Higgs Portal: one of the minimal portals (Big Idea 1 and 2)

Next generation(s) of charged kaon experiments can close low mass allowed region

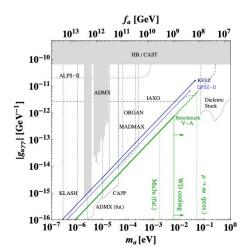


Bridge the gap with next generation Kaon experiments

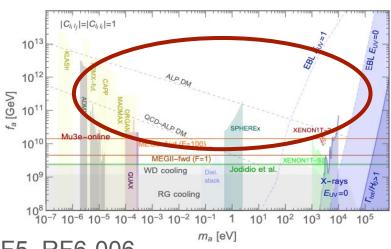
2201.07805

Lepton flavor violating ALP as dark matter candidate

- If light enough ALP can be a DM candidate
 - If flavor violating couplings, can probe very high scales
 - For instance, for LFV couplings $\mu \rightarrow e$ a
 - Could be the QCD axion



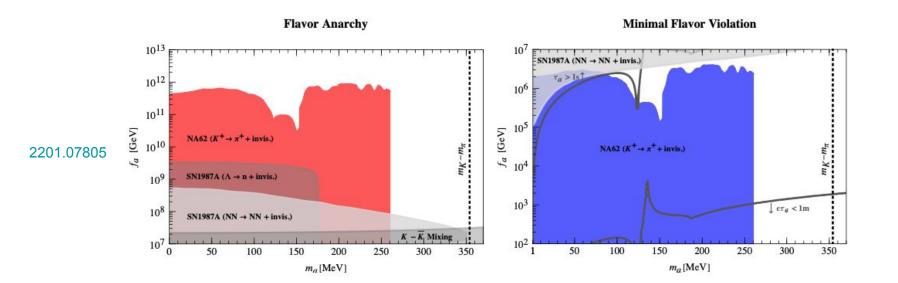
2006.04795



LOI Physics potential with MegII-fwd — RF5_RF6-006

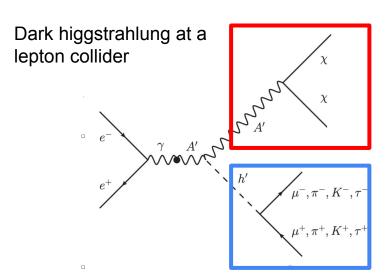
ALP couplings to quarks

- Flavor violating couplings can translate to drastically different bounds on the parameters of the model
- Example: bounds on fa for ALPs



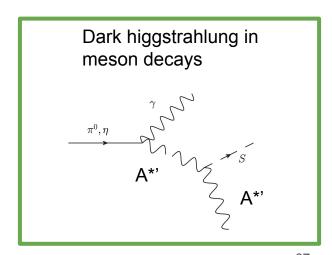
Adding Both a Dark Higgs and a Dark Photon

- Dark Higgs is one way to give the dark photon a mass
- The addition of the Dark Higgs introduces lots of other possibilities
 - Effectively yields both light scalar and vector interactions
- Now have the possibility of multi light boson production



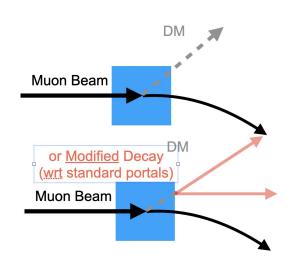
Dark Photon that can decay either visibly or *invisibly*

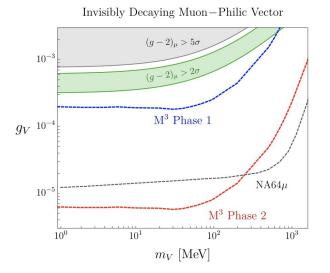
Dark Higgs with yukawa enhanced decays



Models motivated by Physics anomalies

- (g-2)₁₁ is a major motivation for modified physics models
- Overall Strategy is to modify existing portal models
- Commonly used model for this is U(1)_{u-т}
- Collection of other possible models will go in paper





Motivated by anomalies

- A number of LOIs focus on modified models to explain g-2
- Additionally there is interest in models that are motivated by heavy flavor
- Another, motivated modification of the couplings also exist
 - Baryophilic
 - Hadrophilic
 - O ...

Related LOIs

- Light mediators and flavor anomalies (theory) RF6_RF1_Alakabha_Datta-01
- LFV at FCCee RF6_RF4-EF3_EF4_Mogens_Dam-119

Other Models

Sexaquark Dark Matter

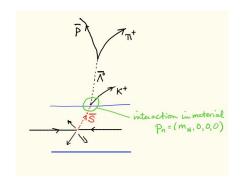
- Using Quarks to motivate Dark Matter, Distinct, rare signatures
- LOI: Delayed electroweak phase transition (theory)
 CF1_CF0-EF7_EF10- RF3_RF6_Glennys_Farrar-198
- LOI: Accelerator search for color-flavor-spin singlet uuddss bound state DM CF1_CF0-EF7_EF10- RF3_RF6_Glennys_Farrar-198

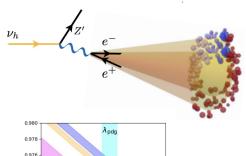
Non minimal HNL models

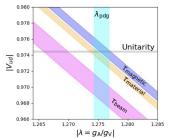
- Modified HNL which can potentially explain excess like MiniBoone
- LOI: NF2_NF3-EF9_EF0-RF4_RF6-CF1_CF0-TF8_TF11_Matheus_Hostert-041

Neutron Portal Dark Matter

- \circ **LOI**: $\triangle B = 2 RF4_RF6-NF3_NF10-TF2_TF5_Joshua_Barrow-105$
- LOI: Sterile neutrons at ORNL and ESS RF6_RF3_Joshua_Barrow-115



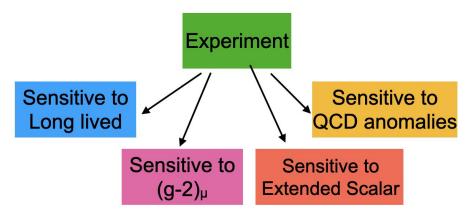




Experimental Signatures (and Motivations)

General View of Experimental Connections

- When trying to construct a model or explain an anomaly
 - Certain experiments are emphasized over other experiments
- This can help to further motivate a specific experimental emphasis
- For Big Idea 3, we would like to highlight these extended motivations
 - And as a consequence, further highlight certain final states



The full details of the experiments is left for Big Idea 4

Photon: Photon beam experiments RF6_RF0-112

Neutrino beams: SNOWMASS21-NF3_NF0-RF6_RF0-CF1_CF3-TF9_TF11-148

Electron: LDMX RF6_RF0-EF10_EF0-CF1_CF0_Andrew_Whitbeck-104

Electron fixed target spectrometer: HPS RF6_RF0_Nelson-078

Muon: Muon missing momentum RF6_RF0-EF10_EF0-CF1_CF0_Andrew_Whitbeck-111

Proton fixed target spectrometer: DarkQuest RF6_RF0_Nhan_Tran-025

Proton: 1GeV proton beam dump at Fermilab RF6_RF0-NF2_NF3-AF2_AF5-099

Proton: 10GeV proton beam dump at Fermilab RF6_RF0-NF3_NF0-AF5_AF0-084

Kaon: Dark Sectors at NA62 & KLEVER — RF6_RF0-011

Kaon: Dark Sectors at KOTO — RF6_RF0_KOTO-050

Kaon: Dark sectors at kaon factories (theory) — RF6_RF0-034

Eta: Redtop RF2_RF6-IF6_IF3_REDTOP_Collaboration_-_new-083

Eta: Eta-Eta' factories RF6_RF2_Sean_Tulin-117

Photon: Photon beam experiments RF6_RF0-112

Neutrino beams: SNOWMASS21-NF3_NF0-RF6_RF0-CF1_CF3-TF9_TF11-148

Electron: LDMX RF6_RF0-EF10_EF0-CF1_CF0_Andrew_Whitbeck-104

Electron fixed target spectrometer: HPS RF6_RF0_Nelson-078

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Enhanced by Dark Mesons/inelastic DM/Neutrino dipole

Photon: Photon beam experiments RF6_RF0-112

Neutrino beams: SNOWMASS21-NF3_NF0-RF6_RF0-CF1_CF3-TF9_TF11-148

Electron: LDMX RF6_RF0-EF10_EF0-CF1_CF0_Andrew_Whitbeck-104

Electron fixed target spectrometer: HPS RF6_RF0_Nelson-078

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Enhanced by Dark Mesons/inelastic DM/Neutrino dipole

Enhanced by Extended Scalar/ Heavy Flavor

Photon: Photon beam experiments RF6_RF0-112

Neutrino beams: SNOWMASS21-NF3_NF0-RF6_RF0-CF1_CF3-TF9_TF11-148

Electron: LDMX RF6_RF0-EF10_EF0-CF1_CF0_Andrew_Whitbeck-104

Electron fixed target spectrometer: HPS RF6_RF0_Nelson-078

Muon : Muon missing momentum RF6_RF0-EF10_EF0-CF1_CF0_Andrew_Whitbeck-111 $(g-2)_{II}$

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Proton: 1GeV proton beam dump at Fermilab RF6_RF0-NF2_NF3-AF2_AF5-099

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Eta: Eta-Eta' factories RF6_RF2_Sean_Tulin-117

Enhanced by Dark Mesons/inelastic DM/Neutrino dipole

Enhanced by (g-2)_µ explanation

Enhanced by Extended Scalar/ Heavy Flavor

Enhanced by explanation of QCD Anomalies

76

Electron: LLP at Belle II RF6_RF0_Torben_Ferber-020

Electron: ILC beam dump: EF9_EF0-RF0_RF6-086

Electron: LLP at FCCee EF8_EF9-RF6_RF0_Rebeca_Gonzalez_Suarez-147

Proton: BDF/SHiP facility AF5_AF0-RF6_RF0-163

Proton: Forward Physics Facility: EF9_EF6_EF10_EF5-NF6_NF3_NF10-RF6_RF0-CF7_CF0-AF5_AF0-UF1_UF2

Proton: Codex-b: EF9_EF0-RF6_RF0-034

FASER: EF9_EF6-NF3_NF6-RF6_RF0-CF7_CF0-AF5_AF0_FASER2

Mathusla: EF9_EF10-NF3_NF0-RF6_RF0-AF5_AF0-IF3_IF7_MATHUSLA_(David_Curtin)-184

Electron: LLP at Belle II RF6_RF0_Torben_Ferber-020

Electron: ILC beam dump: EF9_EF0-RF0_RF6-086

Enhanced by Dark Mesons/inelastic

DM/Neutrino dipole

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Proton: BDF/SHiP facility AF5_AF0-RF6_RF0-163

Proton: Forward Physics Facility: EF9_EF6_EF10_EF5-NF6_NF3_NF10-RF6_RF0-CF7_CF0-AF5_AF0-UF1_UF2

Proton: Codex-b: EF9_EF0-RF6_RF0-034

Proton: FASER: EF9_EF6-NF3_NF6-RF6_RF0-CF7_CF0-AF5_AF0_FASER2

Proton: Mathusla: EF9 EF10-NF3 NF0-RF6 RF0-AF5 AF0-IF3 IF7 MATHUSLA (David Curti

Electron: LLP at Belle II RF6_RF0_Torben_Ferber-020

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Enhanced by Dark Mesons/inelastic

DM/Neutrino dipole

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Proton: BDF/SHiP facility AF5_AF0-RF6_RF0-163

Proton: Forward Physics Facility: EF9_EF6_EF10_EF5-NF6_NF3_....

Proton: Codex-b: EF9_EF0-RF6_RF0-034

Enhanced by

Extended Scalar/

Heavy Flavor

Proton: FASER: EF9_EF6-NF3_NF6-RF6_RF0-CF7_CF0-AF5_AF0_FASER2

Proton: Mathusla: EF9_EF10-NF3_NF0-RF6_RF0-AF5_AF0-IF3_IF7_MATHUSLA_(David_Curti

Electron: LLP at Belle II RF6_RF0_Torben_Ferber-020

Electron: ILC beam dump: EF9_EF0-RF0_RF6-086

Enhanced by Dark Mesons/inelastic DM/Neutrino dipole

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Proton: Codex-b: EF9_EF0-RF6_RF0-034

Proton: FASER: EF9_EF6-NF3_5_AF0_FASER2..

Proton: Mathusla: EF9_EF10-NF3_MATHUSLA_(David_Curti...

Enhanced by

Extended Scalar/

Heavy Flavor

Enhanced by

 $(g-2)_{\mu}$

explanation≈

Discussion Points

How much can we use existing anomalies to drive this paper?

- How exotic of a model should we include?
 - What is the right level to motivate DM measurements beyond portals

- How do we handle overlap with BI 1 and BI 3
 - It can be good to have some redundancy

Are we missing something?

Come and Contribute

- Please let us know (by email/...) if we have missed your stuff
- There have been many developments since the LOIs
 - A number of new interesting channels have emerged
 - We are eager to update this with the many ongoing developments
- Please let us know if you plan on contributing to this white paper

Send an email to us with you contribution topic

zupanje@ucmail.uc.edu,schuster@slac.stanford.edu,pcharris@mit.edu

Contributed white papers due by March 15th, 2022

Solicited whitepapers due by April 15th, 2022

Thanks!

Visualizing The Big Ideas

Big idea 2
Over here we look for visible decays from the dark matter

For a give model we can vary dark matter mass (M_{DM}) and mediator mass (M_{med})

Big idea 1
Over here we
look for
invisible decays
or DM scatters

We still get visible states here(sometimes Musuppressed), also big idea 2

Big Ideas 2

Big Ideas 1

Motivating The Big Ideas

 M_{DM}

Big idea 2
For DM that is lighter than the mediator mass Can find a minimum coupling otherwise no

minima

For each model we can compute the minimal coupling For a thermal relic target

Big idea 1

Minimum

coupling

present

(usually take

1/3=M_{DM}/M_{med}

Below the minimum, hard to get a thermal target

mum∞eốupling

Unconstrained

Big Idea 3

Big idea 3
What is another way to extend the model to — enhance the physics performance

Мом Big Ideas 2 New **Axis** Big Ideas 1

Visible and Invisible can play a role in all space once model is more completed

With Big Idea 3 we aim to make models that are **slightly** more complicated

Motivations

The extensions of the models need to be well motivated to ensure lack of simplicity