

Dark Maller and Dark Sectors:

(unknown artist, PAB chalkboard, UW)

Masha Baryakhtar

Snowmass Summer Study July 20, 2022

Theory to Discovery in the

Thanks to Josh Foster, Tongyan Lin, Tien-Tien Yu, Ciaran O'Hare, CF01, CF02, TF09...







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Snowmass Summer Study 2022

adapted from Tongyan Lin TASI lectures <u>1904.07915</u> 2







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~90 orders of magnitude!

Snowmass Summer Study 2022

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Theoretical Directions

□ Theory development*

Production mechanism

D Can it be constrained or detected astrophysically

□ How do we detect and study it in the lab?

*bonus points for solving outstanding problems

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see Josh Foster's talk in this session!







Jungman, Kamionkowski & Griest (1996)

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Theorists Searching for New Physics

Weakly Interacting Massive Particles





Theorists Searching for New Physics







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Theorists Searching for New Physics

Sikivie: Experimental Tests of the "Invisible" Axion (1983)

Krauss, Moody, Wilczek, Morris: Calculations for Cosmic Axion Detection (1985)





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Goodman, Witten Detectability of Certain Dark Matter Candidates (1984)





Dark Matter in the Last Decade: Explosion of ideas





Dark Matter in the Last Decade: Some Examples



(2,3) axions coupled (1) c to photons media

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(I) dark photonmediated light dark



graphics by Tien-Tien Yu





Axions interacting with Photons $10 M_{\odot}$ $M_{\rm pl}$ GeV keV $100 \,\mathrm{TeV}$



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- Axions produced through misalignment and/or emission from cosmic strings
- Generically couple to photons, contributing new source terms in Maxwell's equations
- Generic light axions motivated from UV theories like string theory



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- photons relativistic while dark matter is almost at rest •
- Use material lattice to introduce momentum scale in the problem

MADMAX collaboration MB, J. Huang, R. Lasenby, PRD 2018 Chiles, Charaev, Lasenby, MB, Huang, Roshko, Burton, Colangelo, Van Tilburg, Arvanitaki, Nam, Berggren 2110.01582 Manenti, Mishra, Bruno, Di Giovanni, Millar, Morå, Roberts, Oikonomou, Sarnoff, Arneodo 2110.10497

Heavier Axion Dark Matter

$M_{\rm pl}$ GeV $100 \,\mathrm{TeV}$





- Convert to photons, which are relativistic while dark matter is almost • at rest
- Use material lattice to introduce momentum into the setup

MADMAX collaboration MB, J. Huang, R. Lasenby, PRD 2018 Chiles, Charaev, Lasenby, MB, Huang, Roshko, Burton, Colangelo, Van Tilburg, Arvanitaki, Nam, Berggren 2110.01582 Manenti, Mishra, Bruno, Di Giovanni, Millar, Morå, Roberts, Oikonomou, Sarnoff, Arneodo 2110.10497

Heavier Axion Dark Matter







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Axion field modifies Maxwell's equations

$$\nabla \times \mathbf{B} = \mathbf{J} + g_{a\gamma\gamma} \frac{\partial a}{\partial t} \mathbf{B}$$

- Time-varying effective current, induces time-varying magnetic flux
 - Measuring the flux = measuring the local axion field

Lighter Axion Dark Matter



Y. Kahn, B.R. Safdi, J. Thaler, Phys. Rev. Lett. 117 (2016) 14, 141801









ABRA-10cm Prototype Detector from Josh Foster









 $10 \, M_{\odot}$

Challenges and Opportunities at Low Masses



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nucleus adapted from Tien-Tien Yu



Challenges and Opportunities at Low Masses



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Challenges and Opportunities at Low Masses



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Inherently different and messy, but also fun and full of potential!

Dark matter energies and length scales become comparable to interatomic spacing, vibrational modes in materials, people, labs, planets...





Dark Matter in the Next Decade: New Materials, New Thresholds, New Interactions

 $\frac{10^{-15}}{80}$ $\frac{10^{-16}}{10^{-16}}$

Ultralight scalar modifications of SM couplings

Axion interactions with electrons, nuclei, incl. `fundamental coupling'



+new models and observables





Mount Rainier as seen from Paradise Meadow NPS photo



AXION DIRECT DETECTION WITH ABRACADABRA



Salemi, **JWF** et al., arXiv:<u>2102.06722</u> [hep-ex]

