

# Neutrino Theory and Particle Physics

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# A wealth of neutrino data...



(image credit: C. Wiens)

#### Some highlights:

- 1998: atmospheric \(\nu\_\mu\) disappearance (SK) 2011
  2002: solar \(\nu\_e\) disappearance (SK) 2012
  2002: solar \(\nu\_e\) appear as \(\nu\_\mu\), \(\nu\_\pi\) (SNO)
  2004: reactor \(\nu\_e\) oscillations (KamLAND) 2014
  2004: accelerator \(\nu\_\mu\) disappearance (K2K) 2015
  2006: accelerator \(\nu\_\mu\) disappearance (MINOS) 2016
  - 2011: accel.  $\nu_{\mu}$  appear as  $\nu_{e}$  (T2K,MINOS) 2012: (Daya Bay, reactor  $\overline{\nu}_e$  disappear **RENO,...)** reactor angle measured! CP violation hint? (T2K) 2014: 2015: normal hierarchy hint? (SK, T2K, NOvA) 2016: non-maximal atm hint? (NOvA) 2018: CP cons disfavored at  $2\sigma$  (T2K) 2021: improved direct mass limit (KATRIN)

# The emergent picture...

# a (seemingly) robust 3-neutrino mixing scheme



$$\sum m_i < 0.12 \; \mathrm{eV}$$
 (cosmology)

 $m_{
m 
u} < 0.8 \ {
m eV}\,$  KATRIN (2021)

**Global Fits:** 

Esteban et al. '20, (www.nu-fit.org)

#### Capozzi et al. '21, deSalas et al. '20

NuFIT 5.1 (2021)

		Normal Ord	dering (best fit)	Inverted Ordering $(\Delta \chi^2 = 2.6)$	
without SK atmospheric data		bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range
	$\sin^2 \theta_{12}$	$0.304^{+0.013}_{-0.012}$	$0.269 \rightarrow 0.343$	$0.304^{+0.012}_{-0.012}$	$0.269 \rightarrow 0.343$
	$\theta_{12}/^{\circ}$	$33.44_{-0.74}^{+0.77}$	$31.27 \rightarrow 35.86$	$33.45_{-0.74}^{+0.77}$	$31.27 \rightarrow 35.87$
	$\sin^2 \theta_{23}$	$0.573^{+0.018}_{-0.023}$	$0.405 \rightarrow 0.620$	$0.578\substack{+0.017\\-0.021}$	$0.410 \rightarrow 0.623$
	$\theta_{23}/^{\circ}$	$49.2^{+1.0}_{-1.3}$	$39.5 \rightarrow 52.0$	$49.5^{+1.0}_{-1.2}$	$39.8 \rightarrow 52.1$
	$\sin^2  heta_{13}$	$0.02220^{+0.00068}_{-0.00062}$	$0.02034 \rightarrow 0.02430$	$0.02238^{+0.00064}_{-0.00062}$	$0.02053 \rightarrow 0.02434$
	$ heta_{13}/^{\circ}$	$8.57^{+0.13}_{-0.12}$	$8.20 \rightarrow 8.97$	$8.60^{+0.12}_{-0.12}$	$8.24 \rightarrow 8.98$
	$\delta_{ m CP}/^{\circ}$	$194^{+52}_{-25}$	$105 \rightarrow 405$	$287^{+27}_{-32}$	$192 \rightarrow 361$
	$\frac{\Delta m_{21}^2}{10^{-5} \ {\rm eV}^2}$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \ {\rm eV}^2}$	$+2.515^{+0.028}_{-0.028}$	$+2.431 \rightarrow +2.599$	$-2.498^{+0.028}_{-0.029}$	$-2.584 \rightarrow -2.413$
		Normal Ordering (best fit)		Inverted Ordering $(\Delta \chi^2 = 7.0)$	
		bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range
	$\sin^2 \theta_{12}$	$0.304^{+0.012}_{-0.012}$	$0.269 \rightarrow 0.343$	$0.304^{+0.013}_{-0.012}$	$0.269 \rightarrow 0.343$
with SK atmospheric data	$\theta_{12}/^{\circ}$	$33.45_{-0.75}^{+0.77}$	$31.27 \rightarrow 35.87$	$33.45_{-0.75}^{+0.78}$	$31.27 \rightarrow 35.87$
	$\sin^2 \theta_{23}$	$0.450^{+0.019}_{-0.016}$	$0.408 \rightarrow 0.603$	$0.570^{+0.016}_{-0.022}$	$0.410 \rightarrow 0.613$
	$ heta_{23}/^{\circ}$	$42.1^{+1.1}_{-0.9}$	$39.7 \rightarrow 50.9$	$49.0^{+0.9}_{-1.3}$	$39.8 \rightarrow 51.6$
	$\sin^2  heta_{13}$	$0.02246^{+0.00062}_{-0.00062}$	$0.02060 \to 0.02435$	$0.02241^{+0.00074}_{-0.00062}$	$0.02055 \to 0.02457$
	$ heta_{13}/^{\circ}$	$8.62_{-0.12}^{+0.12}$	$8.25 \rightarrow 8.98$	$8.61_{-0.12}^{+0.14}$	$8.24 \rightarrow 9.02$
	$\delta_{ m CP}/^{\circ}$	$230^{+36}_{-25}$	$144 \rightarrow 350$	$278^{+22}_{-30}$	$194 \rightarrow 345$
	$\frac{\Delta m_{21}^2}{10^{-5} \ {\rm eV}^2}$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$	$7.42^{+0.21}_{-0.20}$	$6.82 \rightarrow 8.04$
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.510^{+0.027}_{-0.027}$	$+2.430 \rightarrow +2.593$	$-2.490^{+0.026}_{-0.028}$	$-2.574 \rightarrow -2.410$

# **Caveat: light sterile neutrino(s)?**

#### (image credit: ParticleBites)

#### **Anomalies:**

1995:	$\overline{\nu}_e$	appearance	(LSND)
2007:	$\overline{\nu}_e$	appearance	(MiniBooNE)
2012:	$\nu_e$	appearance	(MiniBooNE)
1995:	$\nu_e$	disappearanc	e (Gallium)
2011:	$\nu_e$	disappearance	ce (Reactor)

[well-documented tension between appearance and disappearance data]



**Recent update: Kopp's talk at Neutrino 2022** 

(see also: Arguelles' talk at Neutrino 2022)

status remains unclear!

**Restrict focus here to 3 active light families only** 

# **Physics Beyond the Standard Model!**

suppression of overall neutrino mass scale

window to new physics, possibly at very high energy scales!

- implications for the flavor puzzle of the Standard Model
  - ✓ two large lepton mixing angles (or more\*\*)
  - ✓ mass pattern: nature of mass hierarchy, neutrino mass ordering



connections to cosmology, astrophysics (matter-antimatter asymmetry,...)

# **Neutrino Masses**

- Minimal extensions of SM:
  - ✓ include NR terms  $\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_5 + \dots$
  - ✓ include RH neutrinos  $\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_4 + \dots$

which (if any) global symmetries are preserved in the truncation...

main question: are neutrinos Dirac or Majorana?

Dirac

 $\Delta L = 0$ 







Critically important question, to be settled by experiment!



# Type I seesaw

Minkowski '77, Gell-Mann, Ramond, Slansky '79, Yanagida '79 Mohapatra, Senjanovic '80, Schecter, Valle '80...

✓ introduce right-handed neutrinos

$$\mathcal{L} \simeq Y_{ij} L_i N_j H + M_{R\,ij} N_i N_j$$

$$\mathcal{M}_{\nu} \sim \langle H \rangle^2 Y M_R^{-1} Y^T$$



 $\mathcal{M}_{\nu} = \left( egin{array}{ccc} 0 & m \\ m & M \end{array} 
ight) \qquad m_1 \sim rac{m^2}{M} \qquad m_2 \sim M \gg m_1 \quad \mbox{(canonical version)}$ 

"phenomenological" Type I seesaw: active - (heavy) sterile mixing



Type II seesaw

Magg, Wetterich '80, Lazarides, Shafi, Wetterich '81, Cheng, Li '80, Schechter, Valle '80, Mohapatra, Senjanovic '81

 $\checkmark~$  introduce EW triplet Higgs scalar  $~\Delta\sim({\bf 3},2)$ 

 $\mathcal{L} \simeq Y_{\Delta ij} L_i L_j \Delta + \mu_\Delta H H \Delta$ 

$$\mathcal{M}_{\nu} \sim \langle H \rangle^2 Y_{\Delta} \mu_{\Delta} / M_{\Delta}^2$$

$$\begin{array}{c}
H \\ \searrow \\ H \\ & \swarrow \\ & \downarrow \\ & \downarrow \\ L_i \\ & \downarrow \\ & L_j
\end{array}$$

$$\sqrt{2}\Delta = \begin{pmatrix} \Delta^+ & \sqrt{2}\Delta^{++} \\ \sqrt{2}\Delta^0 & -\Delta^+ \end{pmatrix}$$

couplings to SM leptons directly connected to PMNS parameters!

Han et al., 2203.06131 Snowmass TF11

Fileviez Perez et al. '08, Gavela et al. '09,... Nemeviesk et al. '16, Fuks et al. '19 ,...



# **Type III seesaw**

Foot, He, Joshi '89, Ma '98,...

✓ introduce electroweak triplet fermions  $\Sigma \sim (\mathbf{3}, 0)$ 

 $\mathcal{L} \simeq Y_{\Sigma ij} L_i \Sigma_j H + M_{\Sigma ij} \Sigma_i \Sigma_j$ 

$$\mathcal{M}_{\nu} \sim \langle H \rangle^2 Y_{\Sigma} M_{\Sigma}^{-1} Y_{\Sigma}^T$$



triplet mass can be at high or low scales

"phenomenological" Type III seesaw:

mixing w/charged leptons

 $M_{\Sigma} \sim O(\text{TeV})$ 

Han et al., 2203.06131 Snowmass TF11

del Aguila et al. '08, Franchesini et al. '08,... Abada et al. '07, Gavela et al. '09,...



"radiative Type I

seesaw"

Canonical example: "scotogenic" model Ma '06,...

 $\mathcal{M}_{\nu} \sim \lambda \frac{\langle H \rangle^2}{16\pi^2} Y M_R^{-1} Y^T$ 

 $\nu$ SM

 $\begin{array}{c} & \langle H \rangle \\ & & \langle H \rangle \\ & & & \langle H \rangle \\ & & & & \\ & & & \\ & & & & & \\ & & & & & \\ & & & & \\ & & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\$ 

can also have radiative Type III, Type II hybrid versions: "scoto-seesaw,"...

Ma, Suematsu '09,... Fraser et al. '15,... Rojas et al. '18, Valle et al., '20,...

 $N_k$ 

```
\Delta L = 2
   Radiative models:
                                       beyond \nuSM
                                                                                            Zee '86, Babu '88, Ma '98,...
                                                                                   Babu et al. '01, de Gouvea et al. '07,
                                                                                                    Bonnet et al. '12 ...
             d=7
                                                    d=9
                                                                                             An excellent review:
 LLLe^{c}H \ LLQd^{c}H
                                                LLLe^{c}Le^{c} (Zee-Babu)
                                                                                               Cai et al. 1706.08524
LL\overline{Q}\overline{u}^{c}H \quad L\overline{e}^{c}\overline{u}^{c}d^{c}H
                                               LLQd^{c}Qd^{c}
                            possible connections to dark matter
                  \checkmark
                                                                                         (discrete symmetries)
                            possible connections to |flavor physics anomalies
                  \checkmark
                                                                                                       muon g-2
```

many authors!

Many other ideas for Majorana masses

✓ **inverse seesaw:** 3 RH neutrinos and 3 new singlet fermions

 $\mathcal{L} \simeq Y_{ij} L_i N_j H + M_{ij} S_i N_j + \mu_{ij} S_i S_j$ 

Mohapatra, Valle '86,...

$$M_{\nu} \sim \langle H \rangle^{2} (Y^{T} (M^{T} \mu^{-1} M)^{-1} Y) \qquad \qquad \mathcal{M}_{\nu} \sim \begin{pmatrix} 0 & m & 0 \\ m & 0 & M \\ 0 & M & \mu \end{pmatrix} \qquad \mu, m \ll M$$

**Recent examples:** 

Pati-Salam + conn. to B anomalies, muon g-2Fileviez Perez et al. '21flavor+CP model with $\Delta(3n^2), \ \Delta(6n^2)$ Hagedorn et al. '21

#### ✓ SUSY with R-parity violation

Aulakh, Mohapatra '82, Hall, Suzuki '84,... Borzumati et al. '96, Grossman et al. '03,...



#### Lepton mixing

 $U_{\rm PMNS} = \mathcal{R}_1(\theta_{23})\mathcal{R}_2(\theta_{13},\delta)\mathcal{R}_3(\theta_{12})\mathcal{P}$ 

 $|U|_{3\sigma}^{\text{w/o SK-atm}} = \begin{pmatrix} 0.801 \to 0.845 & 0.513 \to 0.579 \\ 0.232 \to 0.507 & 0.459 \to 0.694 \\ 0.260 \to 0.526 & 0.470 \to 0.702 \end{pmatrix}$  $0.143 \rightarrow 0.156$  $0.629 \rightarrow 0.779$  $0.609 \rightarrow 0.763$  $|U|_{3\sigma}^{\text{with SK-atm}} = \begin{pmatrix} 0.801 \to 0.845 & 0.513 \to 0.579 \\ 0.244 \to 0.499 & 0.505 \to 0.693 \\ 0.272 \to 0.518 & 0.471 \to 0.669 \end{cases}$  $0.144 \rightarrow 0.156$  $0.631 \rightarrow 0.768$  $0.623 \to 0.761$ 

Two large mixing angles:  $\theta_{23}, \ \theta_{12}$ 

Is  $\theta_{13}$  "large" or "small"?

small reactor angle large reactor angle VS. anarchy structure



NuFIT 5.1 (2021)

# Anarchy



(character: Watterson)

 $\mathcal{U}_{\nu}$  : random draw of unbiased distribution of 3x3 unitary matrices statistical tests: lower bound on  $|\mathcal{U}_{e3}|^2$ 

### basis independence:

distribution invariant upon unitary transformations

Haba, Murayama '00

## **Post-reactor angle measurement: renewed focus**

de Gouvea, Murayama '12 Altarelli et al. '12,...

Some recent highlights:

RG analysis model-building + quark sector CP violation

Brdar, Konig, Kopp '15 Babu et al. '16,... Fortin et al. '17,...

Anarchy hypothesis alone does not provide information on  $\Delta m^2$ 

# Structure

Achieve special structures via symmetries:

• Postulate family symmetry  $G_f$ 

spontaneously broken by "flavon" fields  $\{\varphi_a\}$ 

**Recall often-used quark paradigm:** 

$$\epsilon = \left(rac{\langle arphi 
angle}{\Lambda}
ight) \qquad \Lambda > \langle arphi 
angle > \langle H 
angle \qquad \epsilon \ll 1$$
 Froggatt, Nielsen '79

**natural identification:**  $\epsilon \sim O(\lambda)$   $\lambda = \sin \theta_c$ 

#### (typically related to ratios of quark masses)

Unique theoretical starting point:  $U_{CKM} \sim 1 + O(\lambda)$ 

#### Now, for leptons:



 $\theta_{13}^{
u}=0 \qquad heta_{13}=\lambda/\sqrt{2}$  Vissiani '98, '01

Ramond '04



Many papers and authors! Some authors (not comprehensive):

Babu, Chen, Ding, L.E., Feruglio, Grimus, Hagedorn, King, Lam, Luhn, Ma, Merle, Ohlsson, Rodejohann, Stuart,...

see also the very nice recent review of Feruglio and Romanino '19



All can be obtained via SSB of discrete non-Abelian family symmetries

### **Residual Symmetries**

model-independent approach:

Lam '08, '09, Grimus et al. '09, ... many authors!

determine rows and columns in  $U_{MNSP}$ as pure numbers, independent of masses, depending on preserved subgroups of finite group  $G_f$ 



Very different from quark paradigm (CKM mixing angles as ratios of quark masses)!

#### Corrections





• spontaneous CP violation

**Generalized CP** 

$$X^T \mathcal{M}_{\nu} X = \mathcal{M}_{\nu}^* \qquad Y^{\dagger} \mathcal{M}_e \mathcal{M}_e^{\dagger} Y = (\mathcal{M}_e \mathcal{M}_e^{\dagger})^*$$

"ordinary" CP has X = Y = 1 Branco, Lavoura, Rebelo '86...

Grimus, Rebelo '95

automorphisms of discrete family symmetry







$$Q_i, L_i \sim \mathbf{3}$$



requires intricate sector of flavons+ driving fields (characteristic)

✓ recent proposal:

flavor model-building based on modular groupsFeruglio '17,...<br/>many authors!!modular symmetry:modulus $\tau \rightarrow \frac{a\tau + b}{c\tau + d}$ SL(2, Z)Feruglio '17,...<br/>many authors!! $S: \tau \rightarrow -1/\tau$  $T: \tau \rightarrow \tau + 1$  $S^2 = (ST)^3 = 1$ generates discrete group $\Gamma_N$  $S^2 = (ST)^3 = T^N = 1$ finite discrete group, level N

(holomorphic) Yukawa couplings as modular forms:

$$Y(\tau) \to Y(\gamma(\tau)) = (c \tau + d)^k \rho(\gamma) Y(\tau)$$
 group rep matrix

Allows for simplification of flavon sector:

e.g. triplet flavon  $\phi^{\nu} = (\phi_1^{\nu}, \phi_2^{\nu}, \phi_3^{\nu})^T \longrightarrow (Y_1(\tau), Y_2(\tau), Y_3(\tau))^T$  known functions!  $(A_4)$ 

see also recent talks of Feruglio, Tanimoto at FLASY 2022



modular family symmetries — new sum rules





# **Beyond Modular Symmetries**

Modular symmetries: perhaps a hint at top-down physics

#### String constructions:

higher-dimensional operators (field theoretic) see e.g. Creating geometric suppression (braneworlds)
 worldsheet instantons (nonperturbative)

see e.g. Cremades et al. '03, '04 Blumenhagen et al. '06,... Langacker '11 (review)

Buchmüller et al. '07

many fields: many possible RH neutrino candidates
 can have anarchy or family symmetry

discrete symmetries: flavor sector

see e.g. Kobayashi et al. '06 Abe et al. '09,... Cvetič et al. '18

**Bottom-up approach:** 

modular symmetry and further "stringy" ingredients

further reduction of ad hoc parameters (modular weights,...)

connection to explicit top-down constructions

**Recent examples:** 



see also recent talks of Chen, Ratz, Trautner at FLASY 2022

Neutrino data has led to a renaissance for beyond SM model-building!

Many mechanisms for suppressing the neutrino mass scale many TeV-scale scenarios, testable at LHC

For 3 active neutrinos only:

mixings: anarchy or symmetry

(spontaneously broken) symmetries:

discrete non-Abelian groups

modular symmetries and generalizations

More data (atmospheric angle, Dirac CP phase,...) will help enormously!

**If sterile neutrinos confirmed:** paradigm shifts again!

Stay tuned!