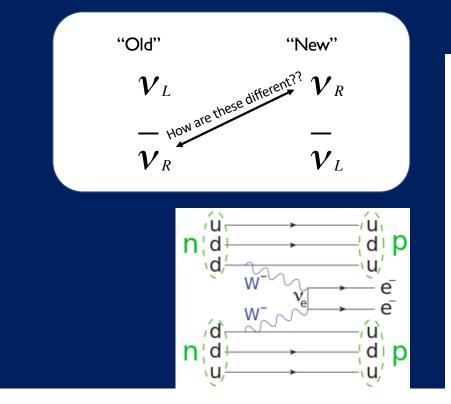
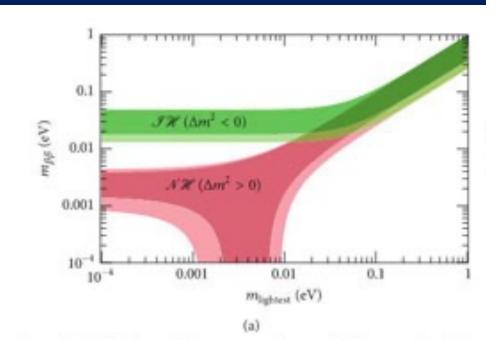
# Tonne-Scale $\beta\beta$ Overview and Future Large Scintillation Detectors

Josh Klein University of Pennsylvania

## $0\nu\beta\beta$ Decay

#### • Most critical, contextual question here at Snowmass





### **US Tonne-Scale Process**

#### $0\nu\beta\beta$ experiments are under the rubric of "nuclear physics"

#### REACHING FOR THE HORIZON

#### The 2015 LONG RANGE PLAN for NUCLEAR SCIENCE



We recommend the timely development and deployment of a U.S.-led ton-scale neutrinoless double beta decay experiment.

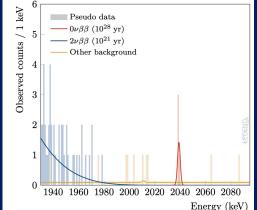
### **US Tonne-Scale Process**

#### Since 2015

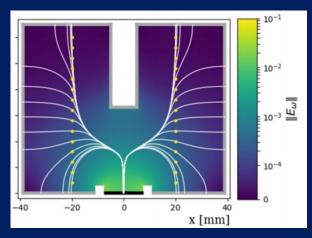
- R&D Review of many different programs
- Funding for several of these
- "Portfolio Review" of tonne-scale possibilities deemed ready, Summer 2021
  - LEGEND1000
  - nEXO
  - CUPID
- International Funding agency discussions in Fall of 2021
- Strong vocal support for a multi-isotope strategy
- Funding from DOE/NP typically quoted as "\$200-\$250 M"
- Full program will require significant international contributions

### LEGEND1000

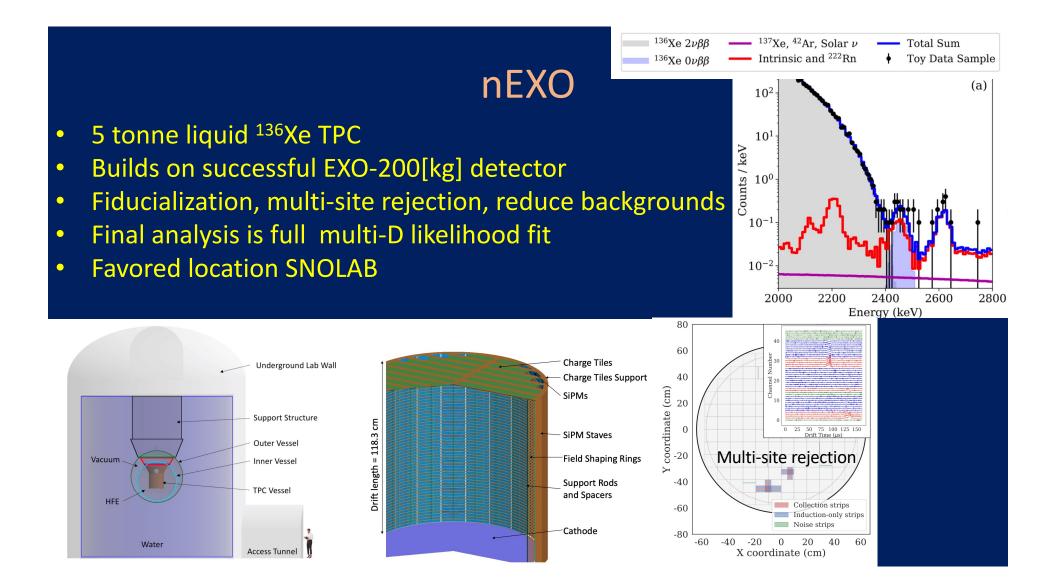
- Merger of GERDA and MAJORAN Collaboration for high-purity <sup>76</sup>Ge experiment
- LAr veto helps with cosmogenic and other backgrounds
- Excellent energy resolution
- New Ge technology for particle ID
- LEGEND200 (200 kg) already underway
- Zero background goal for 10<sup>28</sup> y lifetime





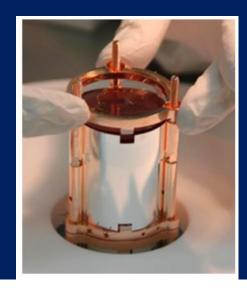


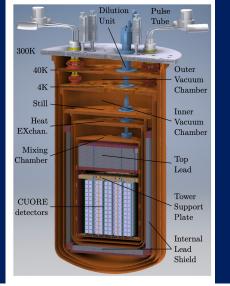


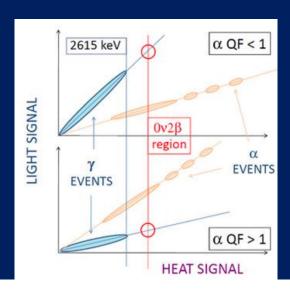


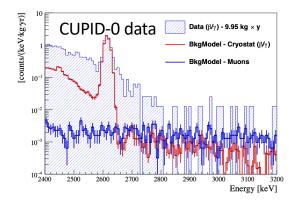
### CUPID

- "CUORE+Particle ID" --> scintillating bolometers
- Bkgds rejected via heat/light ratio and high endpoint
- Initially 450 kg of <sup>100</sup>Mo; plan to move toward 1 tonne
- Re-use CUORE cryostat for big cost savings
- Located at LNGS







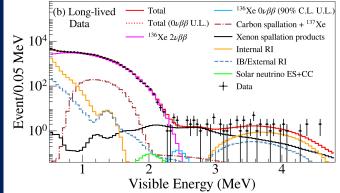


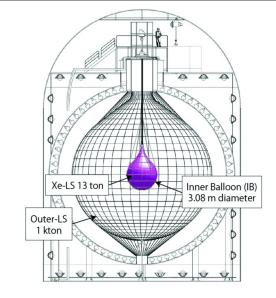
#### Liquid Scintillator Approaches

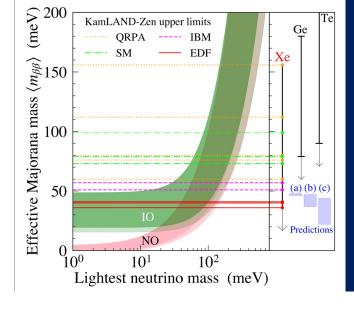
- Break the detector=isotope scaling rules
- (Many backgrounds stay fixed as isotopic mass goes up)
- (But this means solar neutrinos become a background source)
- Leverages large size to eliminate "external" backgrounds
- Broad program of "signal" physics (e.g. solar, reactor, supernova vs...)
- In some cases re-use existing large-scale detectors to save cost

### KamLAND-Zen

- 745 kg <sup>136</sup>Xe gas dissolved in LS
- Fiducialization and ML approaches for PID
- Fast timing rejects coincident backgrounds
- Currently leading world in sensitivity



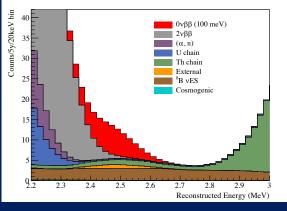


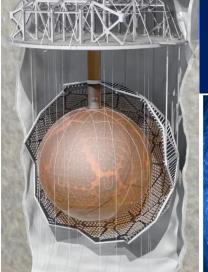




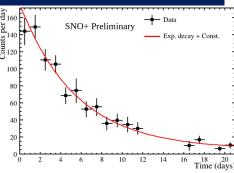
#### SNO+

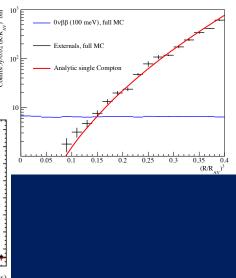
- Goal is ~3% loading by mass of *natural* Te in LS
- Underground purification of Te to eliminate cosmics
- Re-uses SNO PMTs and electronics to save cost
- 0.5% level of Te already u/g, cooling down for years
- Te purification and loading plants commissioning now
- Solar and reactor neutrino physics ongoing with scintillator











### JUNO

2000 days of data taking

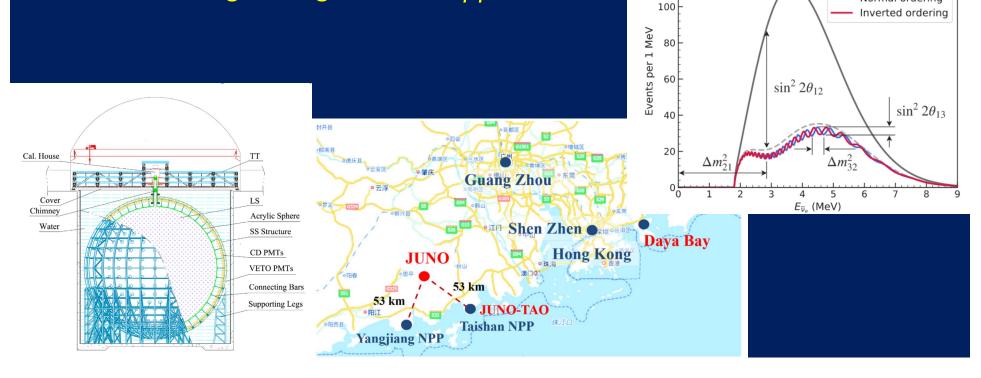
No oscillations

Only solar term

Normal ordering

120

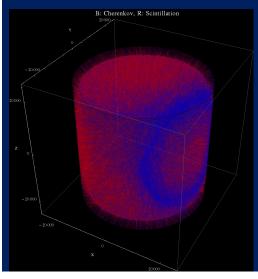
- 20 kt of LAB-PPO with very high PMT coverage (80%)
- Reactor antineutrino measurement of MO at distance of ~50 km
- Also precision on  $\theta_{12}$ ,  $\Delta m 2_{12}$  to < 1%
- Also considering loading Te to do  $0\nu\beta\beta$

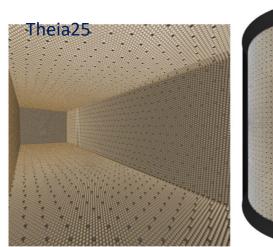


#### Theia

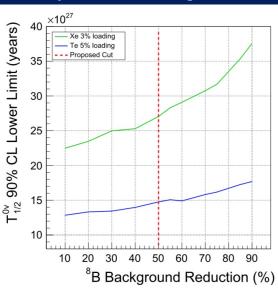
- Hybrid Cherenkov/scintillation detectors receiving a lot of interest
- 25 kt version could be DUNE Module 4 and broaden physics program
  - Solar neutrinos down to CNO or perhaps even below
  - Diffuse Supernova background vs
  - Long-baseline neutrinos
- At last Snowmass, detectors like this were just an idea
- Would require call for new space at SURF

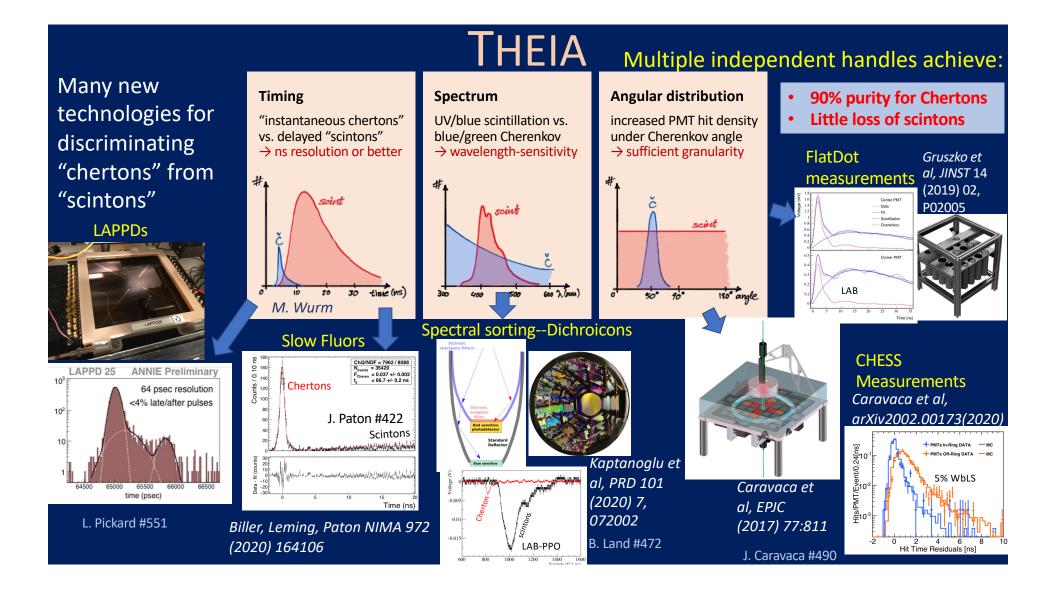
### Cherenkov light allows rejection of <sup>8</sup>B background







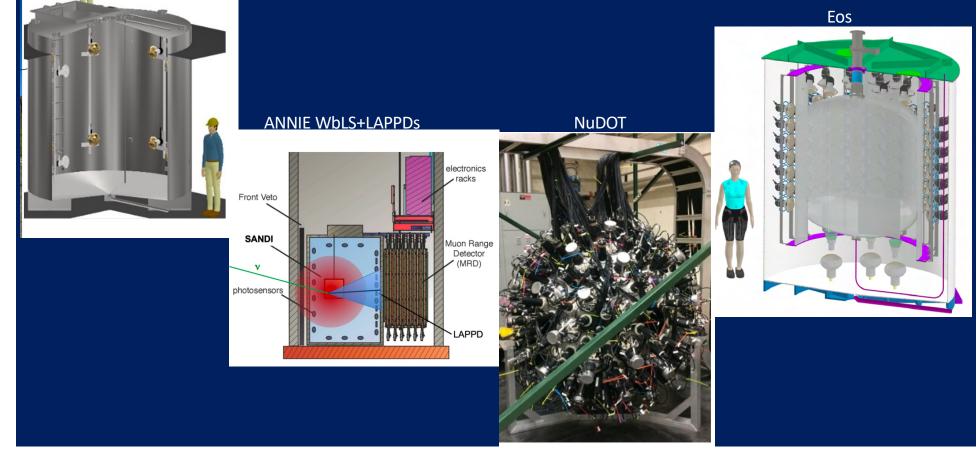




### Hybrid Cherenkov/Scintillation Detectors

#### BNL 30 t WbLS





#### Sensitivity Comparisons

