

Skipper-CCDs

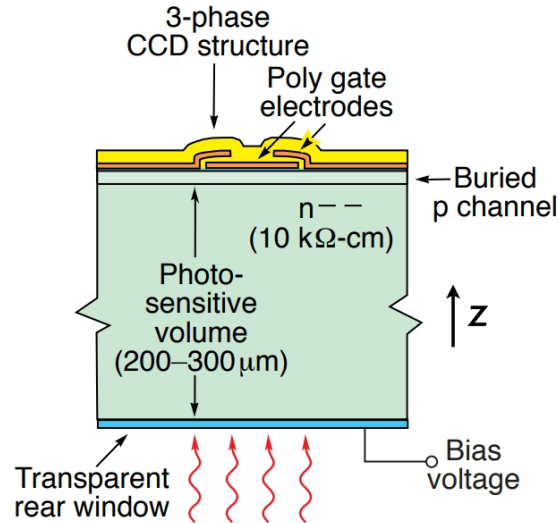
Brenda Aurea Cervantes Vergara

Snowmass 2022

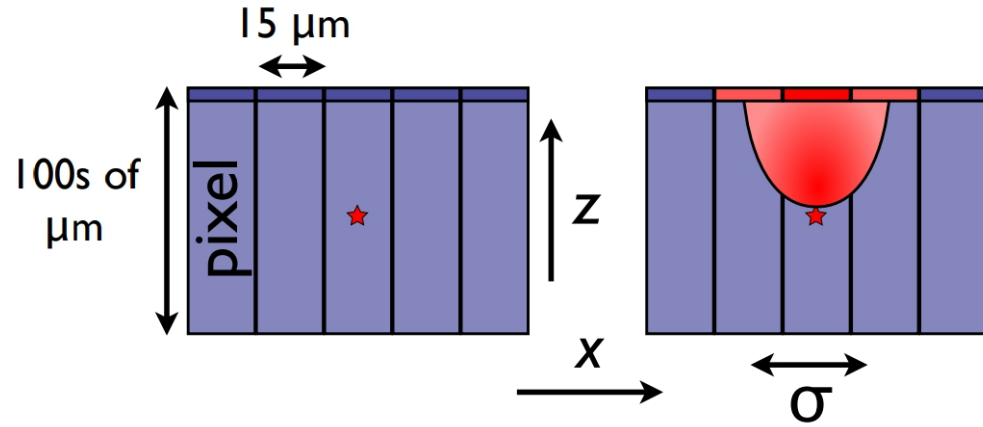
July 18, 2022

Scientific Charge-Coupled Devices: structure and operation

CCDs are essentially an array of Metal-Oxide-Semiconductor capacitors

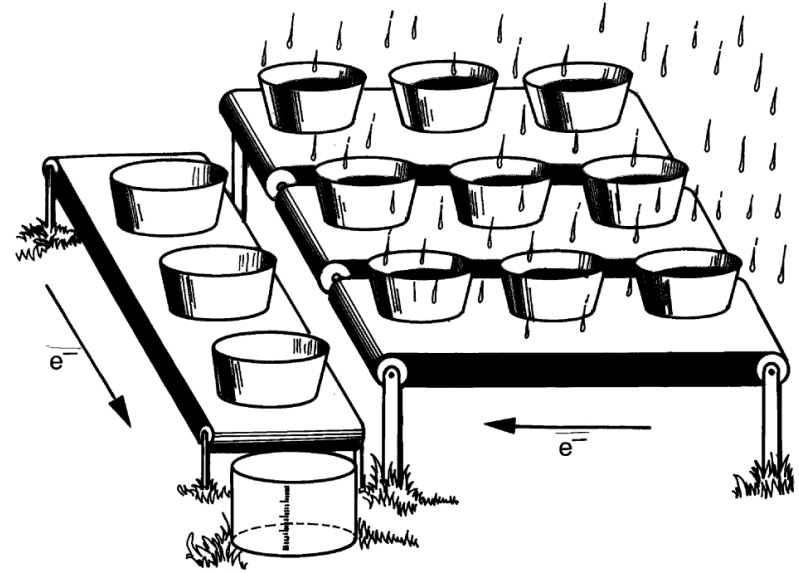
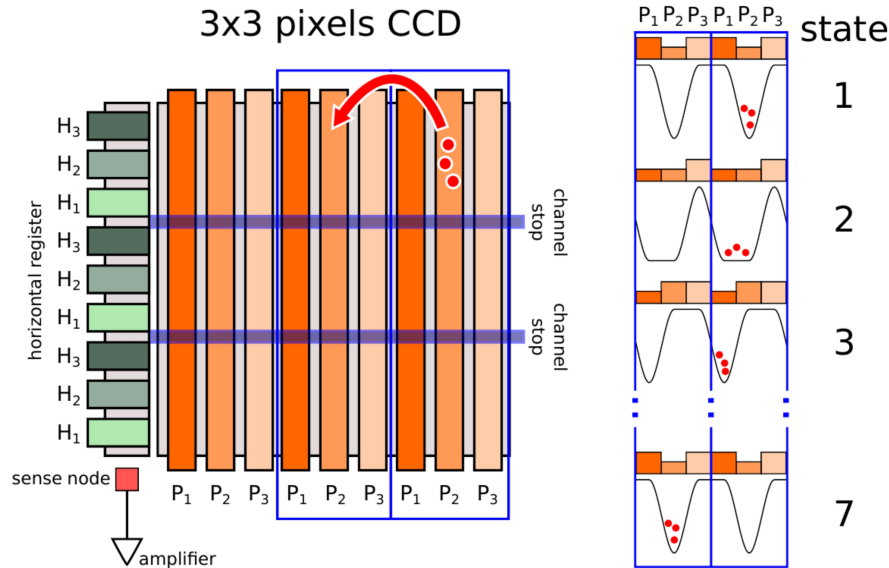


- Radiation interacting in the substrate produces e-h pairs (in Si, 1 e-h pair corresponds to $\sim 3.8 \text{ eV}$)
- Charge is collected near the surface by applying V_{sub}



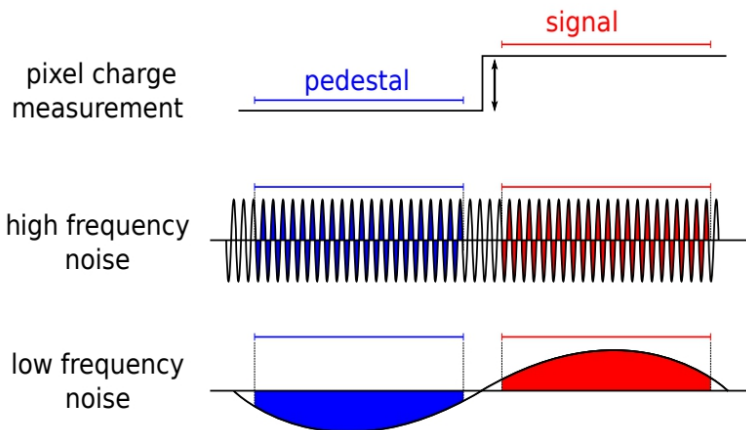
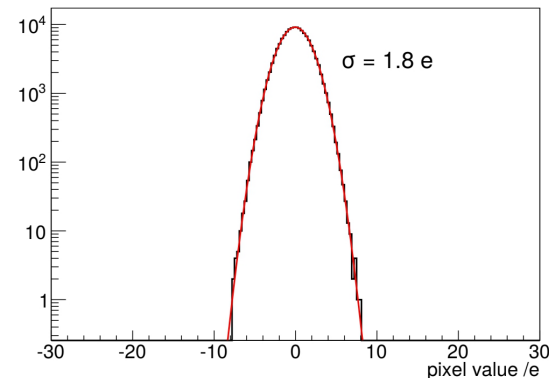
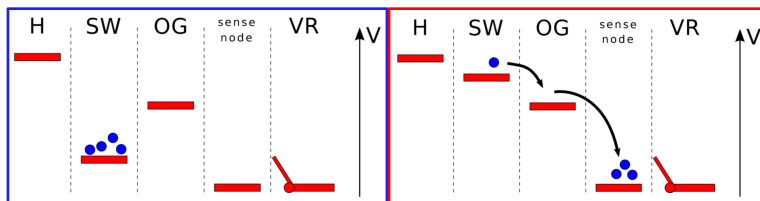
Scientific Charge-Coupled Devices: structure and operation

- Collected charge is then transferred along the surface by varying potential wells until reaching the readout stage



Standard CCDs: readout

- “Charge packets” are moved to the sense node, read and then discarded



Correlated Double Sampling to measure charge:

1. Pedestal integration
2. Signal integration
3. Charge = Signal - Pedestal

Sensitive to low frequency noise

Skipper CCDs

[1990ASPC....8...18J]

→ Very old idea!

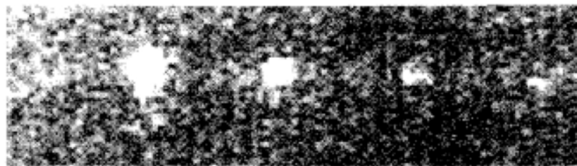
The CCD Skipper was invented to circumvent the $1/f$ noise problem and realize a square-root reduction in noise with increasing sample time thereby allowing sub-electron noise floors to be achieved. The principal function of Skipper technology is to allow the user to nondestructively measure the charge contained in a pixel multiple times (similar to CID operation) using a "floating gate" amplifier. The samples collected for a given pixel are then averaged together off-chip reducing the random noise of the on-chip amplifier by the square-root of the number of samples taken. For example, if a pixel is sampled 100 times, the random noise associated with the on-chip amplifier is diminished by a factor of ten.

→

- Multiple (N) non-destructive measurements of same "charge packet"
- The only modification is done in the readout stage! (floating sense node)



READ NOISE = 7.6 e⁻ rms

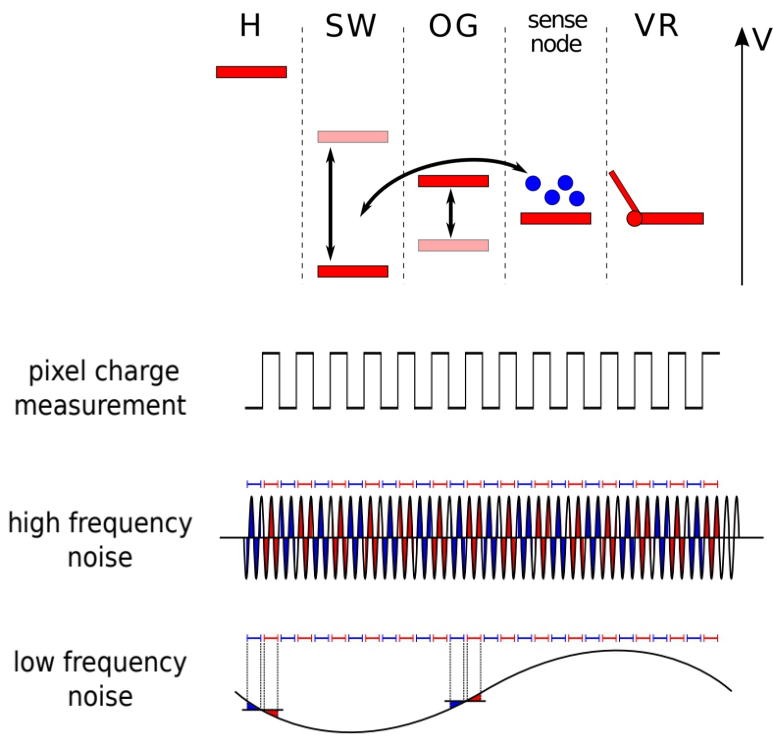


READ NOISE = 0.97 e⁻ rms

$N_s = 64$

The ultimate test for the Skipper CCD, yet to be achieved, is to detect the single photo-electron. Skipper cameras have been constructed which employ ultra high gain, in excess of 100 ADC counts per electron. Assuming that the noise can be lowered to 0.2 e⁻ rms using multiple sampling, there is no fundamental reason why the photo-electron can't be detected. It will be interesting to see if the CCD can accomplish this feat in the near future.

Skipper CCDs: readout



- “Charge packets” are moved back and forth without being corrupted nor destroyed

1. Pedestal integration
2. Signal integration
3. Charge = Signal - Pedestal
4. Repeat N times
5. Pixel value = average of all samples

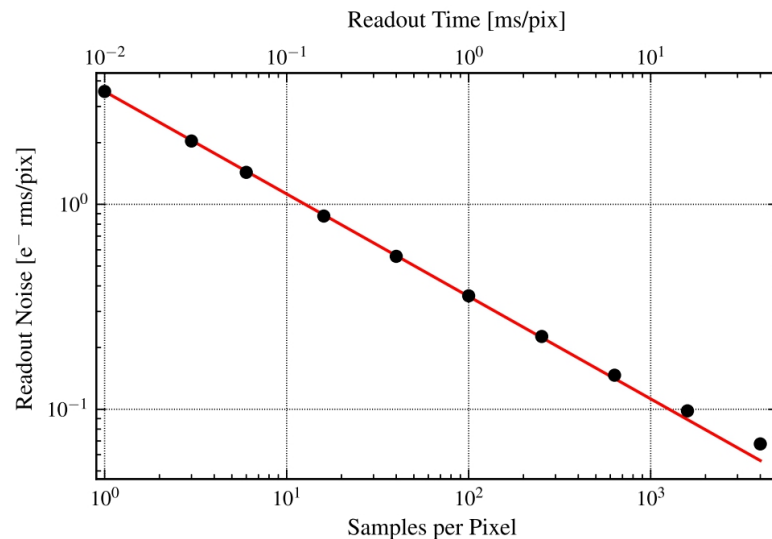
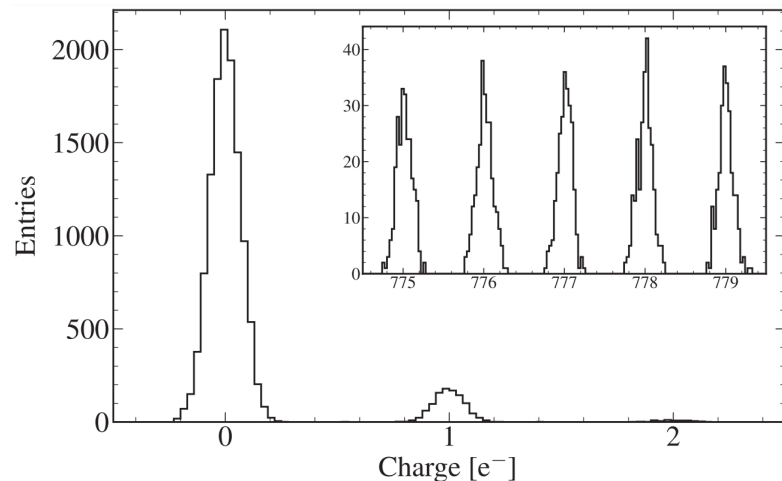
Noise is reduced as $\sigma = \frac{\sigma_1}{\sqrt{N}}$

Low-frequency noise can also be reduced!

Skipper CCDs

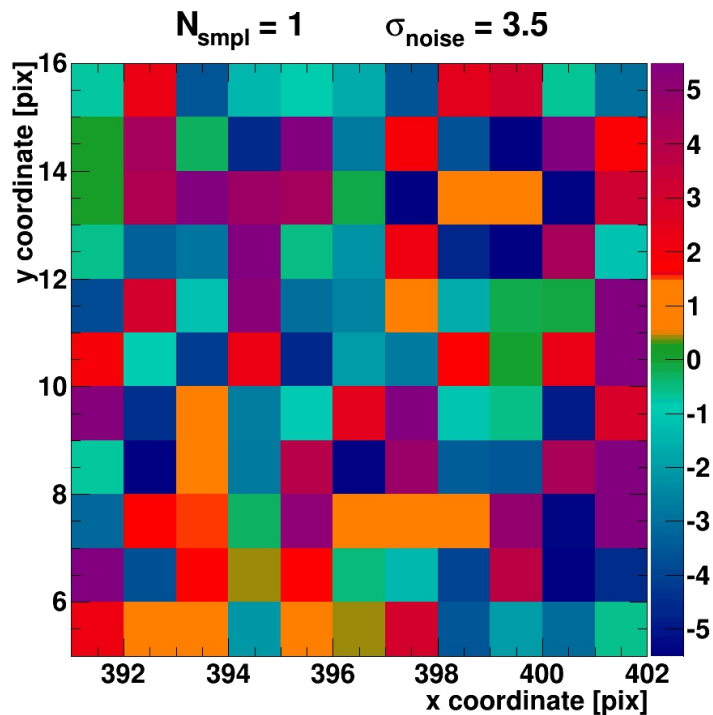
The technology first demonstrated a good performance in 2017 using a detector designed by Stephen Holland (LBNL) allowing to count electrons in a wide dynamic range!

[10.1103/PhysRevLett.119.131802]

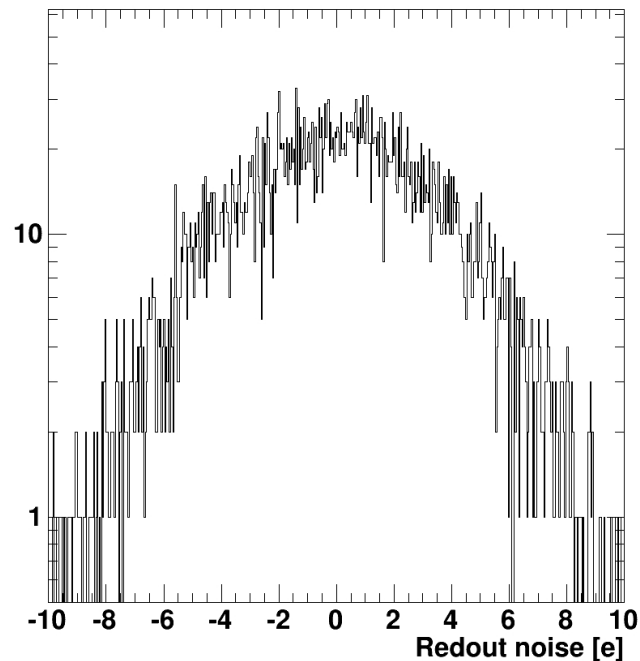


The price to pay is readout time, but this can be optimized depending on your interests*

Skipper CCDs



Taken from real data!



Skipper CCDs: smart readout

[10.1103/PhysRevLett.127.241101]

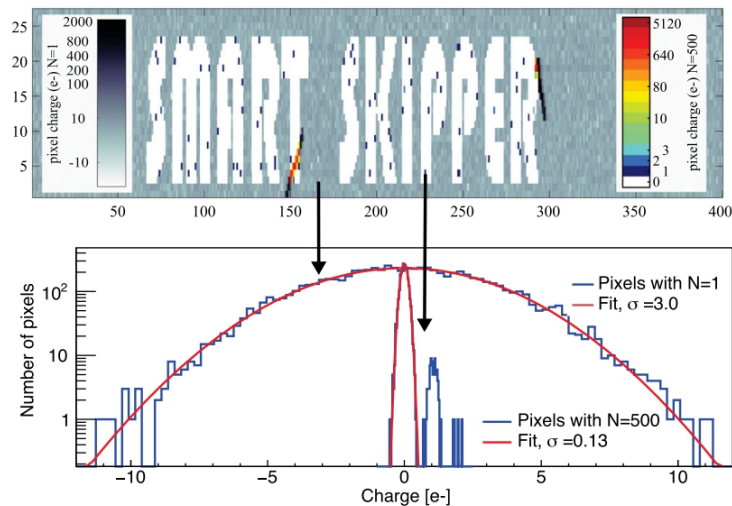


FIG. 3. Measurement using ROI technique. Pixels in the words have $N = 500$ (right scale); pixels outside the words have $N = 1$ (left scale). s_f was zero in most pixels, with some pixels having $s_f = 1, 2, 3$ or very large values for the two muon tracks that are observed.

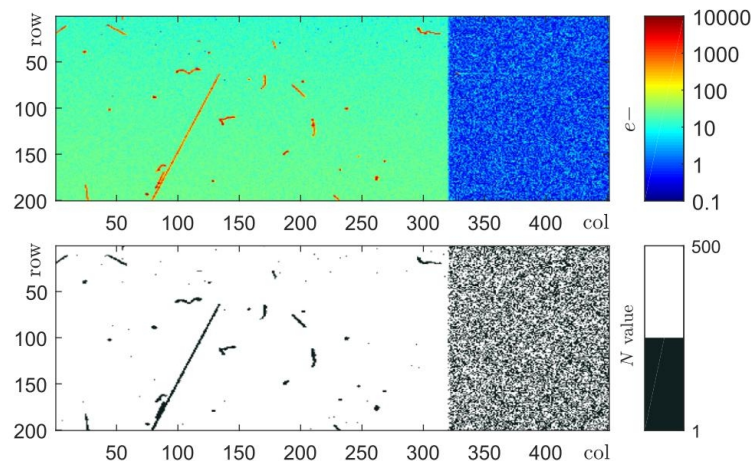


FIG. 4. (Top) Image using EOI technique. (Bottom) N for each pixel.

Skipper CCDs: applications

These are great detectors to study physical processes where a low-energy deposition takes place!

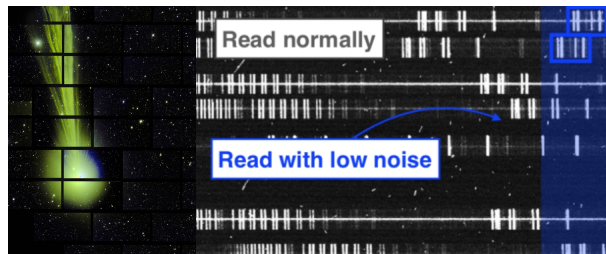
Plus, they have all the attractive features of CCDs, e. g., high quantum efficiency (>90%), great spatial resolution (15 μm x 15 μm pixels), high charge transfer efficiency, low dark current, etc.

Particle physics



Direct dark matter search
Low-energy neutrinos ($\text{CE}\nu\text{NS}$)

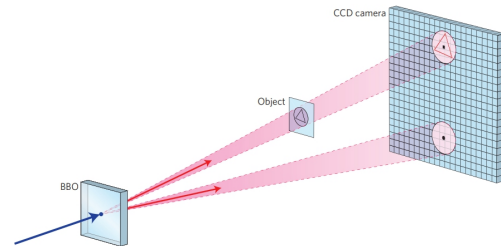
Astronomy



Faint sources
Short duration transients

[10.1117/12.2562403]

Quantum imaging



Agustina Magnoni talk!

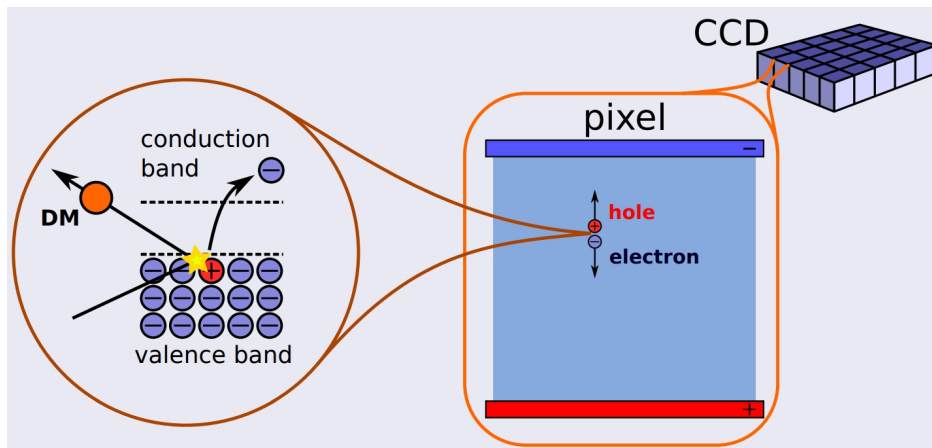
Direct dark matter search with skipper CCDs

DM-electron scattering:

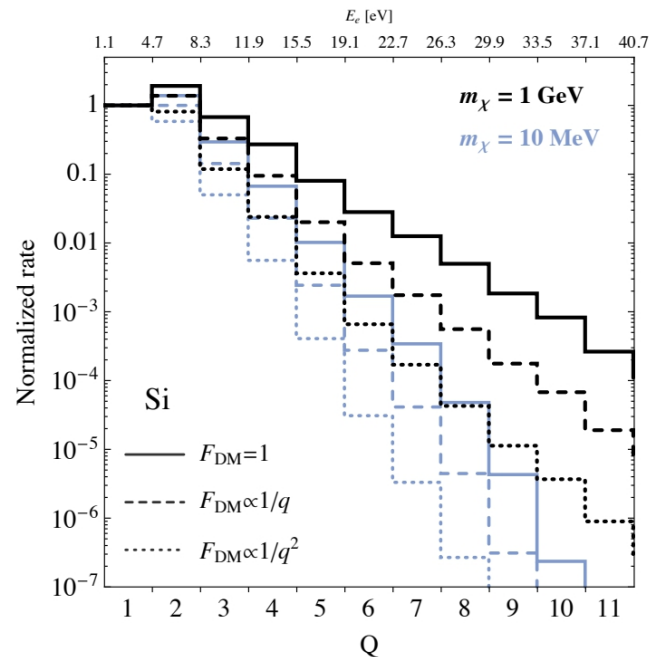
m_χ in the MeV-GeV range \rightarrow energy transfer is a few eV

DM absorption:

Bosonic DM at the eV scale \rightarrow energy transfer equals m_χ



Sensitivity depends entirely on $1e^-$ rate

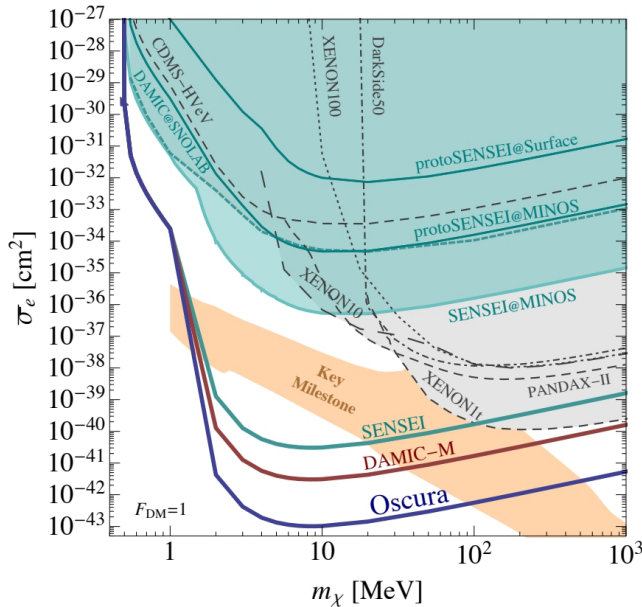


[10.1007/JHEP05(2016)046]

Direct dark matter search with skipper CCDs

[10.1103/PhysRevLett.125.171802]

World best limits for light dark matter candidates with this technology (SENSEI with ~2 g!) 



DM-e- scattering (heavy mediator)

Ongoing program:

Experiment	Mass	Background	Year commissioning
DAMIC @ SNOLAB	~20 g	5 dru	12/2021
DAMIC-M (proto)	~10 g	1 dru	11/2021
SENSEI-100	~100 g	5 dru	2022
DAMIC-M	~1 kg	0.1 dru	2023
OSCURA	10 kg	0.01 dru	~2027

We know how to build experiments with skipper CCDs.
The challenge is going from 100 g to 10 kg (50 CCDs → 24,000 CCDs)

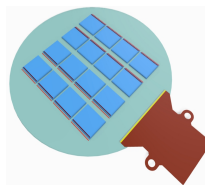
Multi-kg experiments with skipper CCDs

New ideas on sensor packaging, cryogenics and electronics are needed.

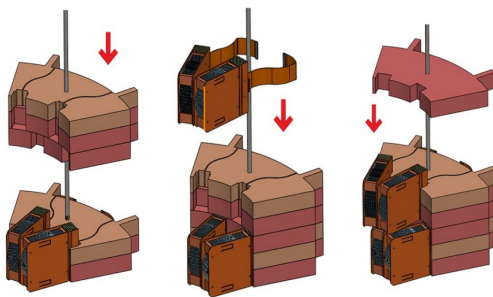
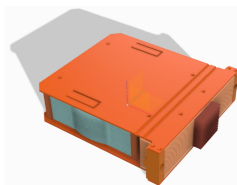
OSCURA's overall envisioned design

[arXiv:2202.10518]

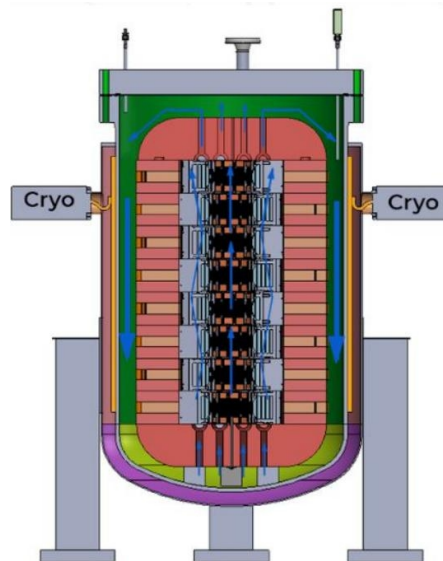
Multi-Chip Module
(16 skipper CCDs)



Super Module
(16 MCMs)

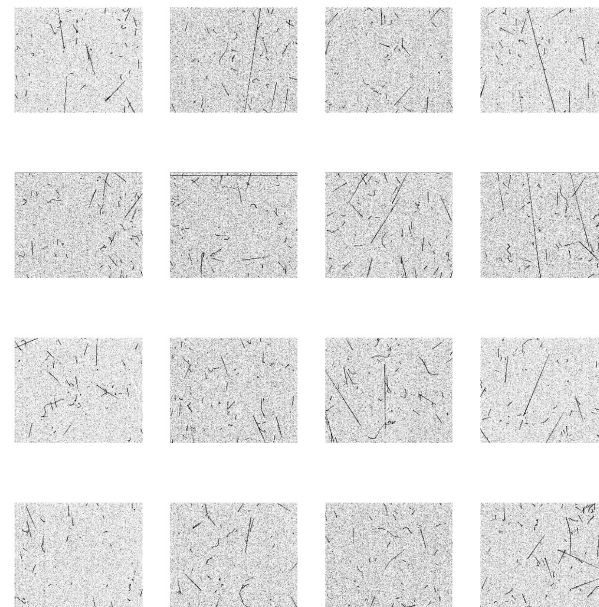
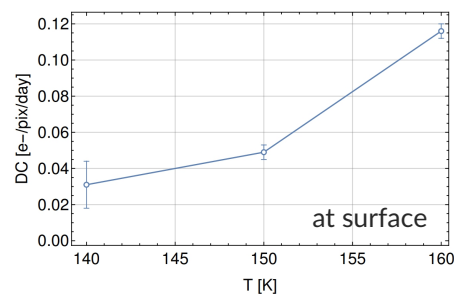
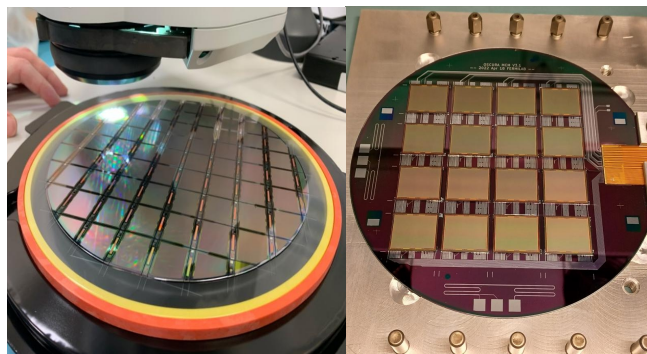


LN₂ pressure vessel



Multi-kg experiments with skipper CCDs

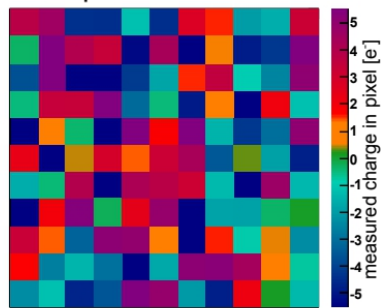
Great progress has been achieved during last year... stay tuned!



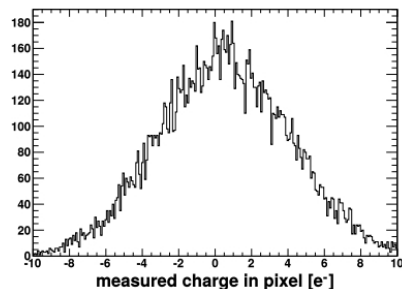
THANK YOU!

Standard vs skipper

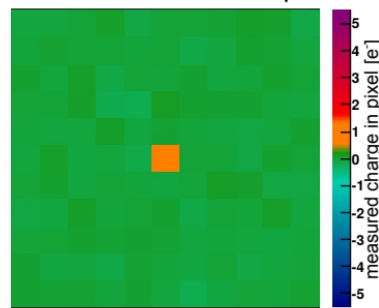
Standard CCD mode: charge in each pixel is measured once



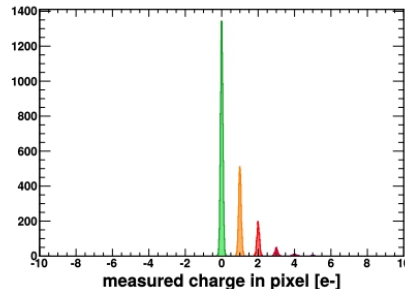
Readout-noise: 3.5 e RMS



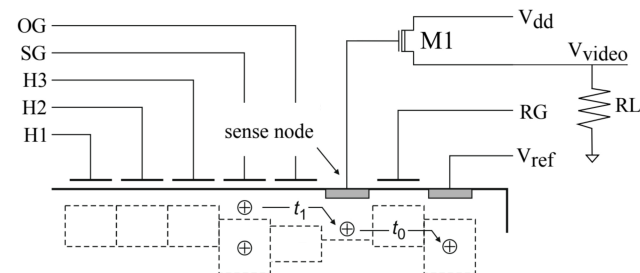
New Skipper CCD: charge in each pixel is measured multiple times



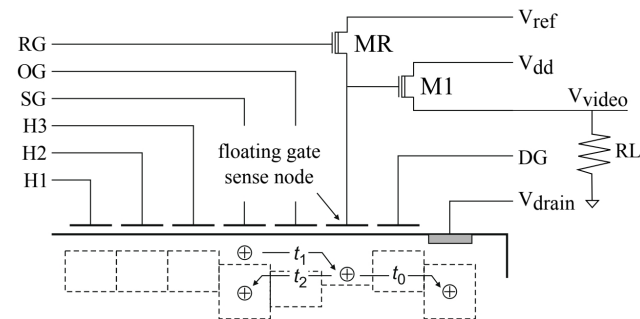
Readout-noise: 0.06 e RMS



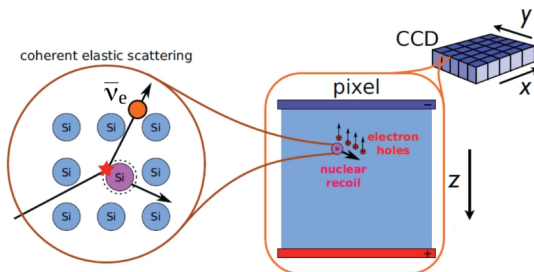
Standard CCD



Skipper CCD



Coherent Elastic Neutrino Nucleus Scattering with skipper CCDs



In Si, for $E_\nu = 2$ MeV, $E_R \sim 300$ eV



Reactor neutrinos are a great source!

- $E_\nu < 4$ MeV and high flux $\sim 10^{20}$ ν/s

- Standard CCDs have demonstrated to be competitive constraining BSM physics (CONNIE) [10.1007/JHEP04(2020)054]
- Skipper CCDs can extend the reach of these searches (ν IOLETA)

