CCDs for DM and ν Physics





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- Charge-coupled devices (CCDs).
- Search for light DM with DM-e⁻ scattering.
- DAMIC and other CCD dark matter experiments.
- CONNIE: search for CE_VNS with CCDs.
- Low-energy spectroscopy with CCDs.
- Radioactive backgrounds in CCDs.
- Spin-off: The Selena Neutrino Experiment.



Charge-coupled devices



 Depth (z) reconstructed from distribution of charge on pixel array.

 Device is "exposed," collecting charge until user commands readout.

Readout can be slow : low noise (few e-).

 Standard fabrication in semiconductor industry and easy cryogenics (~100 K).

Sample CCD image (~15 min exposure) segment in the surface lab.

Cosmic muon –

CU

 \overline{i}

1.

.

Point-like

 β particle

Zoom

. .



.

50 pixels

5 10 15 20 25 Energy measured by pixel [keV]



Cosmic muon

CG

 $\overline{\mathbf{y}}$

7

 β particle

Point-like







DM-e scattering

DAMIC at SNOLAB

- CCDs for direct dark matter search.
- Multi-CCD array operating at SNOLAB since 2012.
- First DM search results from ~eV ionization signals:
- WIMP search:

PRL125(2020)241803

- 11 kg-day of data from seven-CCD array.
- First full background model in CCDs.
- Recently upgraded with skipper CCDs.

SENSE

- First DM-search with skipper CCDs at Fermilab. PRL125(2020)171802
- Experimental / simulations studies on single e⁻ backgrounds:
- 10 skipper CCDs (~25g) deployed at SNOL
- Science goal: DM-e⁻ search with 100 g-year.



PRL118(2017)141803

Full details:

PRD105(2022)062003

24 Mpix skipper CCD @ SNOLAB



SENSEI @ SNOLAB



AB.	PRX12(2022)011009

PRAppl17(2022) 014022



DAMIC-M

- ► 52 skipper-CCD modules in LSM (France) for kg-year target exposures.
- 50x background reduction from DAMIC at SNOLAB.
- Besides ER searches, NR result may have comparable sensitivity to HV detectors of SuperCDMS SNOLAB.
- Commissioning in early 2024.
- **LBC**: small prototype detector at LSM just released best exclusion limits of DM-e⁻ scattering.

Oscura

- R&D: scale the existing technology towards a 10 kg experiment.
- ► **Goal**: DM-*e*⁻ scattering search with 30 kg-yr exposure.

DAMIC-M 4-CCD Module @ UW









IDM2022 Parallel Session 1A

14:00	The DAMIC-M Experiment: Status and First Results	Danielle Norcini
	E17	14:00 - 14:20
	The low-energy spectrum in DAMIC at SNOLAB	Alvaro Chavarria
	E17	14:20 - 14:40
	SENSEI: Sub-GeV Dark Matter Search with Skipper CCDs	Mariano Cababie
	E17	14:40 - 15:00
15:00	The Oscura experiment – searching for low-mass dark matter with a very-large array of skipper-CCDs	Nathan Saffold
	E17	15:00 - 15:20
	First 100 eV nuclear recoil ionization yield measurement in silicon	Dr Valentina Novati
	E17	15:20 - 15:40
	Measurement of low-energy Compton and neutron scattering in Si CCDs for dark matter searches	R Smida
	E17	15:40 - 16:00

16:00

• One parallel session at latest IDM Conference dedicated to CCDs!









Outlook

10

CONNIE:

COherent Neutrino-Nucleus Interaction Experiment

- Flux: ~ $10^{12} \ \bar{\nu_e} \ \mathrm{cm}^{-2} \ \mathrm{s}^{-1}$
- ◆ Energy threshold ~50-70 eV_{ee}
- ◆ Reactor ON (2.1 kg-day) vs. OFF (1.6 kg-day) spectral comparison leads to



Slides from C. Bonifazi 11

Experiment @ 30 m from the 3.9 GW reactor core

◆ 2016-2018: CCD array with 47.6 g active mass.





First competitive BSM constraints from CEvNS at reactors:

CONNIE, JHEP 04, 054 (2020); arXiv:1910.04951

CONNIE 2019





- analysis.

- - Data taking in ongoing
- to observe CEvNS at 90% C.L.
- ♦ Inside reactor dome.

Slides from C. Bonifazi

Higher efficiency, lower background, improved

Results compatible with previous analysis ◆ Expected limit in in the lowest-energy bin ~35 times the SM prediction (c.f. \sim 65 times from before). CONNIE, JHEP 05, 17 (2022); arXiv:2110.13033

Skipper-CCDs @ CONNIE since July 2021 ► 2 skipper-CCDs (1022 x 682 pixel) o Readout noise: ~0.15e- RMS Single electron rate: ~ 0.05 e-/pix/day ► Energy threshold 15 eV_{ee}: x2.2 CEvNS rate

CONNIE Skipper -CCDs

Considering 4 kdru of background (measured) and a future detector of 1 kg at the CONNIE site, 9 days (Lindhard) or 2 months (Chavarria)

Skipper CCD installed 12 meters from Atucha-2 2 GW reactor in Argentina.



co.vnie



Nuclear recoil response

- Detector response calibrated with 24 keV neutrons from ${}^{9}Be(\gamma,n)$ reaction.
- predicted by Lindhard model.



By comparing data and Monte Carlo spectra, ionization efficiency was measured to be lower than

Ionization efficiency for NR measured down to 60 eVee. On going efforts with skipper CCDs!



- decreasing threshold to 23 eV_{ee}:
- Confirmed softening of the L step, observed structure in the L step.
- Softening reproduced with *FEFF* code, which performs full QM treatment.
- Detector response model is good!

CCDs can resolve spectral features in the *ionization* spectrum near threshold!



Precision measurement with a skipper CCD improved energy resolution and arXiv:2207.00809



Radioactive backgrounds

- Particle classification (α , β , NR) by track topology (at high E>100 keV_{ee}).
- Spatial coincidence searches to identify decay sequences: JINST16(2021)P06019
 - Cosmogenic ³²Si: ³²Si (T_{1/2}= 150 y, β) \rightarrow ³²P (T_{1/2}= 14 days, β)

 $140 \pm 30 \ \mu Bq / kg$

- Also upper limits on every β emitter in the U/Th chain.
- Measurement of the cosmogenic activation of ³H in silicon by exposing a CCD to a neutron beam: PRD102(2020)102006
- Exhaustive radio-assay program: PRD105(2022)062003



112 ± 24 atoms / kg /day

Selena Neutrino Experiment

- and electron neutrinos in ⁸²Se.
- target layers.
- Concept paper with background estimates:
- Measurement of charge response:
- Snowmass 2021 white paper:
- First demonstration of imaging electron tracks in a hybrid CMOS/aSe device:



• Next generation neutrino experiment to perform zero-background spectroscopy of $\beta\beta$ decay

Low-noise CMOS charge readout sensor coupled to few mm-thick amorphous selenium (aSe)

JINST12(2017)P03022

JINST16(2021)P06018

arXiv:2203.08779



Selena Goals

- ► 100 ton-year exposure of ⁸²Se.
- Single vs. double β discrimination.
- Bulk backgrounds suppressed by α/β particle ID, spatial correlations.
- Background rate <6 x 10⁻⁵ /keV/ton/year!
- ► $T_{1/2} = 5 \times 10^{28} \text{ y}$: limit on ⁸²Se $0\nu\beta\beta$.
- Solar v spectroscopy from reaction: $v_e + {}^{82}Se \longrightarrow {}^{82}Br^* + C.E. (29 \text{ keV}) + e^-$
- Captures individually tagged from the spatial correlation with ⁸²Br* decay sequence. *Zero-background*.
- Constraints on solar luminosity, solar metallicity, solar core temperature, onset of matter effects in v oscillations, etc.



 $E_v - 172 \text{ keV}$







- CCD particle detectors have demonstrated unique capabilities for low-energy particle physics.
- Decade-long program in detector characterization / background studies.
- Low-energy threshold to probe light DM from DM-e-interactions.
- Skipper CCD kg-scale detectors (**DAMIC-M**) already under way.
- CONNIE plans a similar detector to measure CE_VNS from a reactor.
- Selena to leverage on the low-background capabilities developed for CCD detectors to do neutrino physics with a ⁸²Se target.

Conclusions

