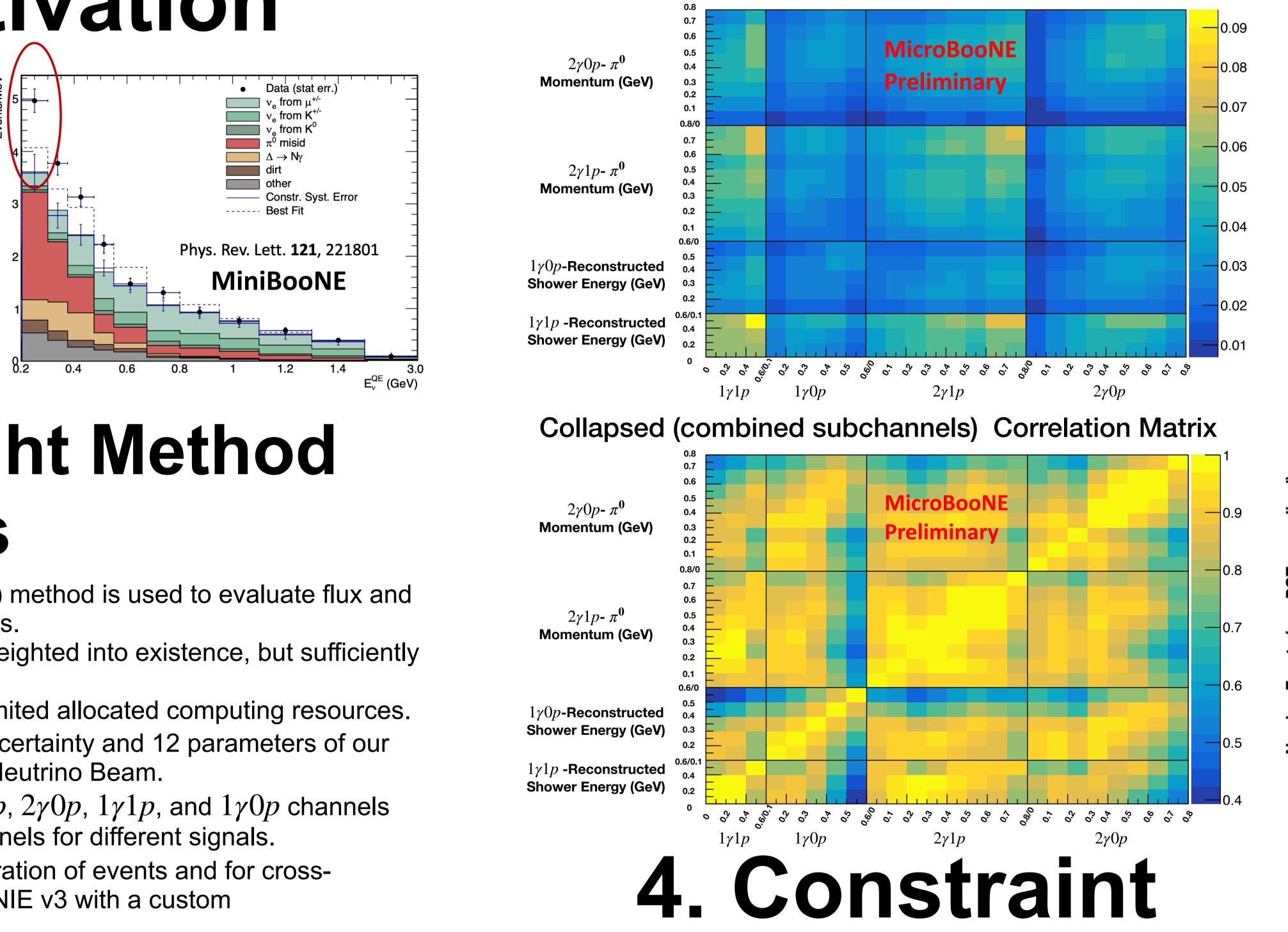
## Systematic Studies in the Single Photon Low Energy **Excess Search at MicroBooNE** uBooNE

### Gray Yarbrough on behalf of the MicroBooNE Collaboration

## 1. Motivation

- MiniBooNE Cherenkov detector along the booster neutrino beam
- Observed low energy excess (LEE) of neutrinos.<sup>1</sup>
- Use data from MicroBooNE, a liquid argon time projection chamber (LArTPC), to test an explanation as neutral current (NC)  $\Delta$  radiative decay
- 3x Standard Model prediction for  $\Delta \rightarrow N\gamma$  would explain the excess



## 2. Eventweight Method and Samples

- The EventWeight (event reweighting) method is used to evaluate flux and cross-section systematic uncertainties.
- Limitation in that events cannot be weighted into existence, but sufficiently large samples should negate this
- This method is most optimal given limited allocated computing resources.
- We vary 56 parameters of GENIE uncertainty and 12 parameters of our source of neutrino flux: the Booster Neutrino Beam.
- The final stage samples for our  $2\gamma 1p$ ,  $2\gamma 0p$ ,  $1\gamma 1p$ , and  $1\gamma 0p$  channels are examined and split into sub channels for different signals.
- GENIE is used for monte-carlo generation of events and for crosssection uncertainties. We utilize GENIE v3 with a custom MicroBooNE central value tune.<sup>2</sup>
- A combined GENIE variation is used including a set of approved GENIE uncertainties including quasi-elastic, meson exchange current, resonant, non-resonant, hadronization and final state interaction processes.
- Delta resonance signal uncertainties were excluded as they correspond to our final measurement.
- Individual GENIE parameters are also studied but are not included in fits or final covariance matrices.

# **3. Fitting Module**

- SBNFit (short baseline neutrino fit) is a module built for the SBN project enabling for fitting which utilizes fractional covariance matrices
- Fractional covariance matrices contain the statistical and systematic uncertainties and systematic correlations among different bins and samples
- Correlation matrices are the covariance of a bin divided by the standard deviation of the two components, and are useful for visualization

#### References

- 1. Phys. Rev. Lett. 121, 221801 (2018),



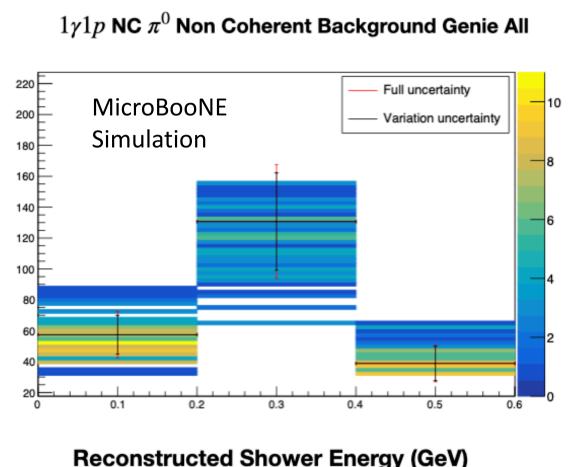
- We evaluate the level of constraint on the uncertainty of the final  $1\gamma 1p$  signal by considering the NC (neutral current)  $\pi^0$  sideband ( $2\gamma 1p$ ) measurement using the following the method.<sup>3</sup>
- Form a matrix of  $1\gamma 1p$  backgrounds and  $2\gamma 1p$  selection ( $M_{ii}$ ).
- For the  $2\gamma 1p$  portion assume  $\sigma_i^{\text{data}} = \sqrt{N_i^{\text{data}}}$ , and  $N_i^{\text{data}}$
- Calculate a new matrix  $(M_{ii}^{-1})^{new} = M_{ii}^{-1} + 1/N_i^{MC}$  and re-invert for constrained uncertainty on the  $1\gamma 1p$  bins.

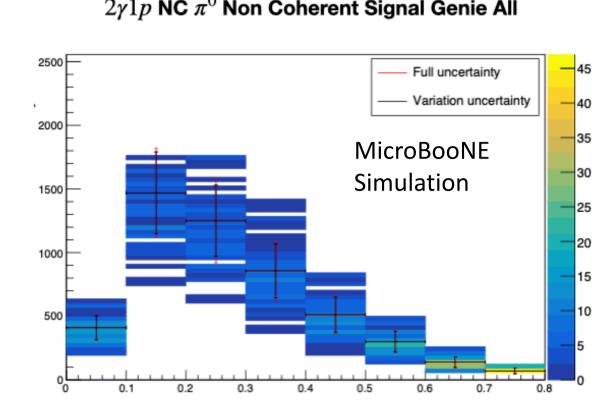
Variation Description	Unconstrained	Constrained	Unconstrained	Constrained
	Uncertainty $1\gamma 1p$	Uncertainty $1\gamma 1p$	Uncertainty $1\gamma 0p$	Uncertainty $1\gamma 0p$
All genie variables combined	22.64%	7.21%	13.82%	4.48%
Fractional cross section for N charge exchange	9.58%	6.69%	1.58%	1.10%
Axial mass for NC resonance $\nu$ production	18.94%	5.45%	10.44%	3.01%
Variation of angle of $\pi$ with respect to detector z axis	7.83%	4.91%	0.98%	0.62%
Vector mass for NC resonance $\nu$ production	8.06%	4.77%	4.41%	2.61%
Fractional cross section for N charge exchange	9.32%	4.36%	4.18%	1.96%
Skin Depth-electric currents penetrate conductor	4.93%	3.41%	4.01%	2.77%
$\pi$ absorption probability	5.12%	3.26%	3.33%	2.13%
Primary Hadron SW Central Spline Variation for $\pi^+$	4.51%	3.23%	3.86%	2.76%
N absorption probability.	4.91%	3.21%	4.61%	3.02%

**Table:** Top 10 highest uncertainties after constraint in  $1\gamma 1p$ 

2. C. Andreopoulos, "Genie physics & users manual v3 prelim." https://genie-524docdb.pp.rl.ac.uk/DocDB/0000/000002/006/man.pdf, 2019 3. M. Shaevitz, "Constraining the nue background using the observed numu events." MicroBooNE DocDB: 7635, 2017

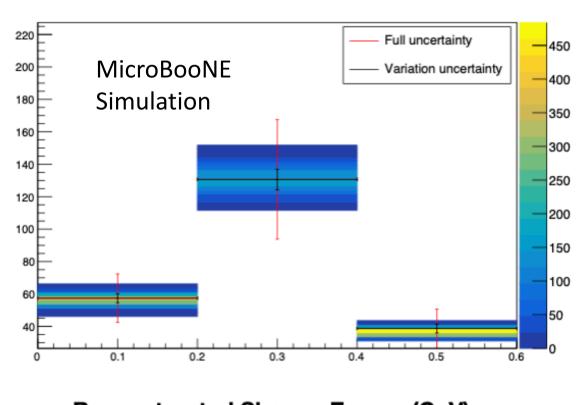




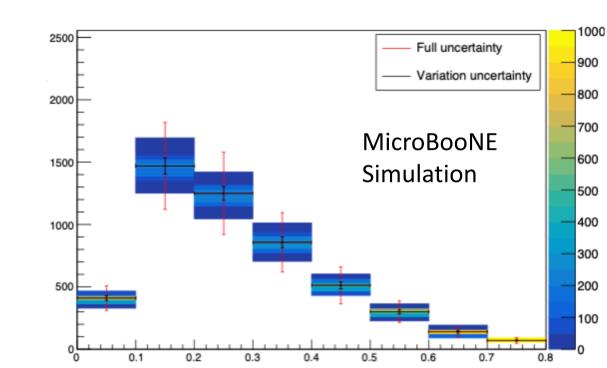


**Reconstructed Shower Energy (GeV)** 

Figure:  $2\gamma$  channel variation plots using the Eventweight Method for combined GENIE variations on NC  $\pi^0$  non-coherent signal subchannels. (majority of all NC  $\pi^0$ ). Y axis correlates to number of events but is not normalized to the POT of the detector. Coloration depicts the number of "multisims" that are in each bin of x and v.



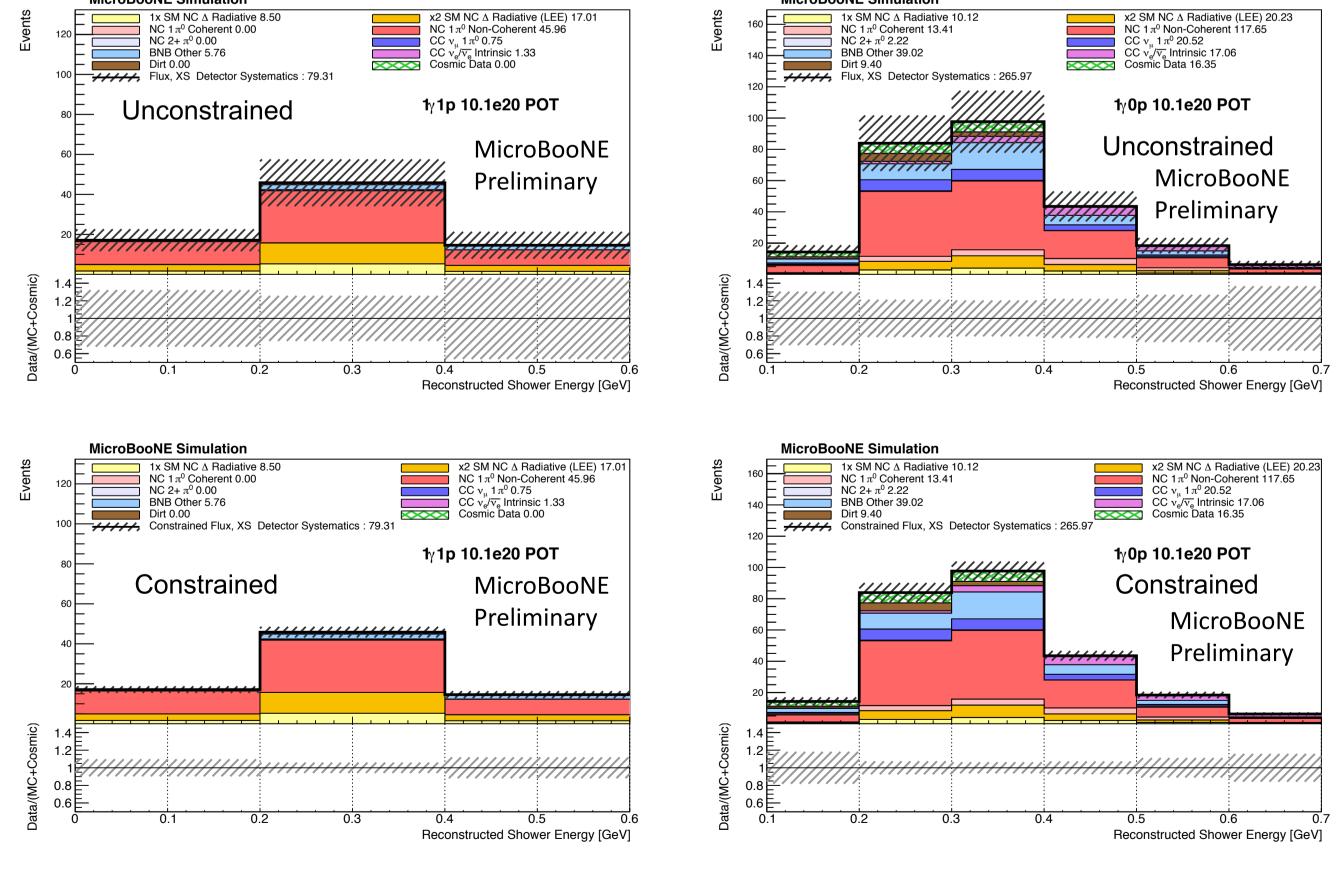
p NC  $\pi^0$  Non Coherent Background Piplus Spline Variation



**Reconstructed Shower Energy (GeV)** 

**Figure:** A variation plot illustrating the central Sanford Wang  $\pi^+$  flux uncertainty effect on the  $NC\pi^0$  non-coherent signal in the final  $2\gamma 1p$  and  $2\gamma 0p$  selection. As the primary signal for neutrino production it has a comparably large uncertainty compared to other flux variations.

$$a = N^{MC}_{i}$$



The per bin constraint on the  $1\gamma 1p$  and  $1\gamma 0p$  final stage samples showing a ~3 x reduction factor uncertainty for a promising evaluation of systematic reduction

- the single photon analysis.











5. Results

 $2\gamma 1p$  NC  $\pi^0$  Non Coherent Signal Genie All

"GENIE All" (the combined GENIE variation) includes a number of highly correlated parameters for our backgrounds i.e. NC Resonance Axial Mass, so the constraint is powerful (3.11 reduction factor)

Pi0 Momentum (GeV)

 $2\gamma 1p$  NC  $\pi^0$  Non Coherent Signal Piplus Spline Variation

Flux variations have smaller uncertainties but aren't specific to our signal so the constraint is only somewhat effective (1.40 reduction factor)

#### Pi0 Momentum (GeV)

• A complete flux and cross section systematic study has been performed for

• Estimated constraint is promising for uncertainties on our final measurement



