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# Neutrino quantum decoherence at current and future reactor experiments

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# Introduction

Neutrino oscillations provide the first laboratory evidence for New physics

They arise as a consequence of neutrino mixing

The QM uncertainty principle implies that when neutrinos are produced at some source they must be in a superposition of different momentum states. The neutrino wave function must be a **wave packet**

Coherence is essential for neutrino oscillations!

# Introduction

Physics that leads to decoherence: the wave packets corresponding to different neutrino mass eigenstates propagate with different speeds and, given enough time, the wave-packets ultimately separate.

Eventually, the wave packets of the  $j$ -th and  $k$ -th mass eigenstate will have no significant overlap any more and their coherence will be lost, leading to a suppression of neutrino oscillations

Nuclear reactors are excellent laboratories to study neutrino coherence!

# Neutrino oscillations with decoherence

When treating neutrinos as wave packets, there is a correction due to the wave packet size

$$P^{\text{dec}}(\bar{\nu}_e \rightarrow \bar{\nu}_e) = \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \exp[-i\Delta_{jk} - \xi_{jk}]$$

where

$$\xi_{jk}(L, E) = \left( \frac{L}{L_{jk}^{\text{coh}}} \right)^2, \quad L_{jk}^{\text{coh}} = \frac{4\sqrt{2}E^2}{|\Delta m_{jk}^2|} \sigma$$

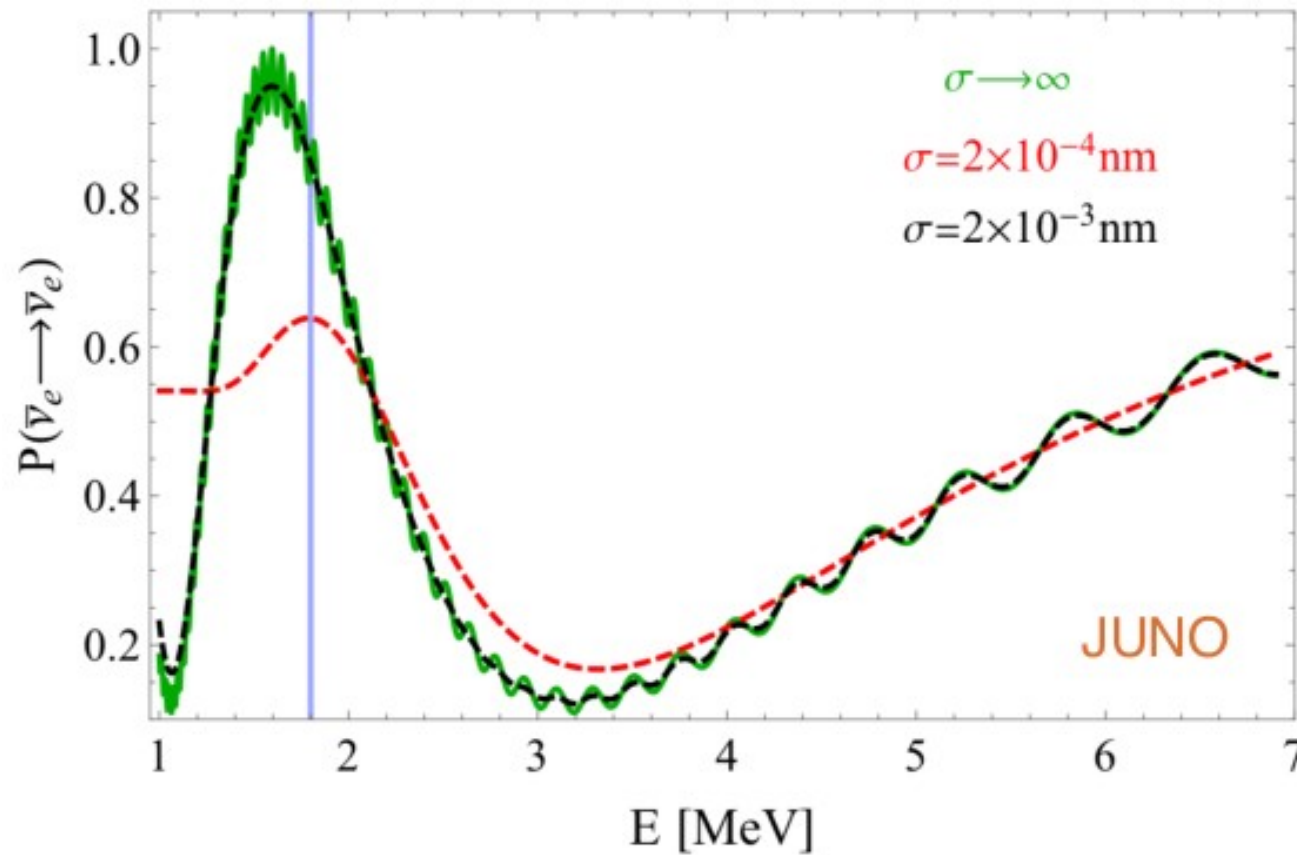
Giunti, Kim, Lee, PLB274 (1992)

M. Beuthe, PRD66 (2002)

Kayser, Kopp, arXiv:1005.4081

# Neutrino oscillations with decoherence

The decoherence flattens out the neutrino oscillations



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

# Bounds from current experiments

## RENO

6 power plants

2 identical detectors

We include 2900 days of data

RENO, private communication

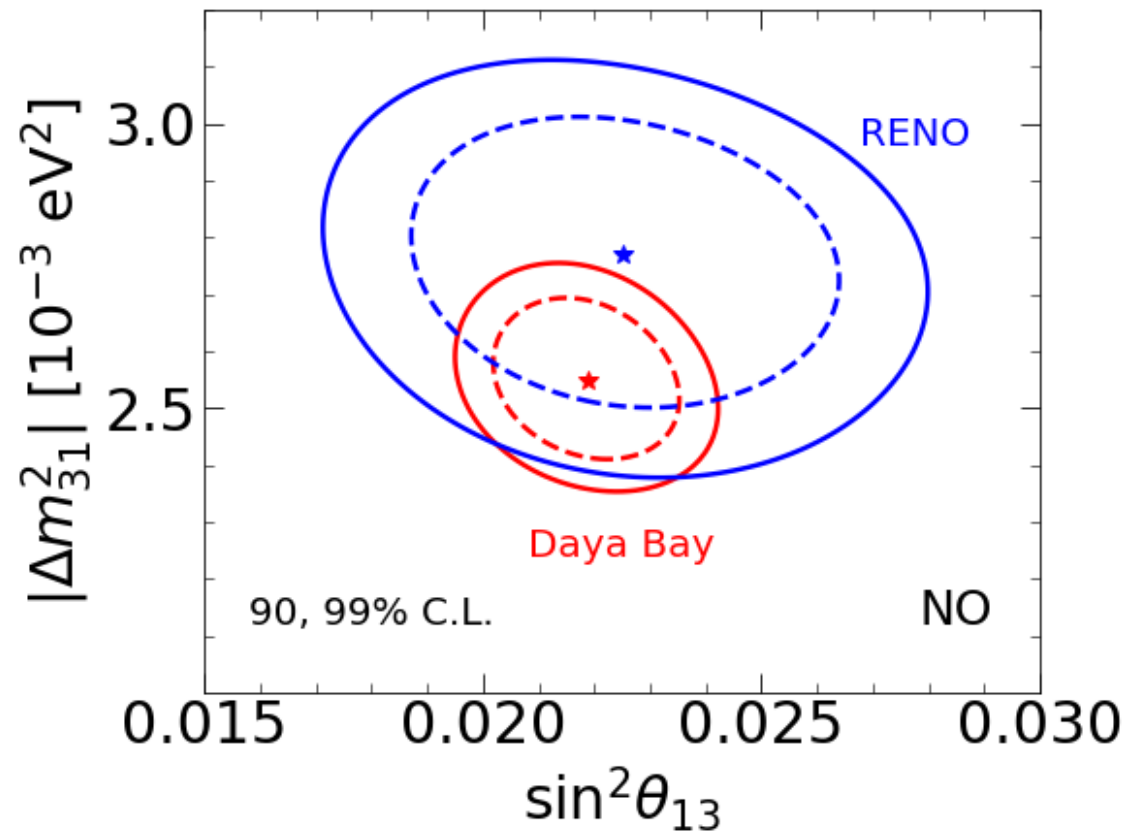
## Daya Bay

6 power plants

8 identical detectors at 3  
experimental halls, 2 function as  
near detectors, 1 as far detector

We include 1958 days of data

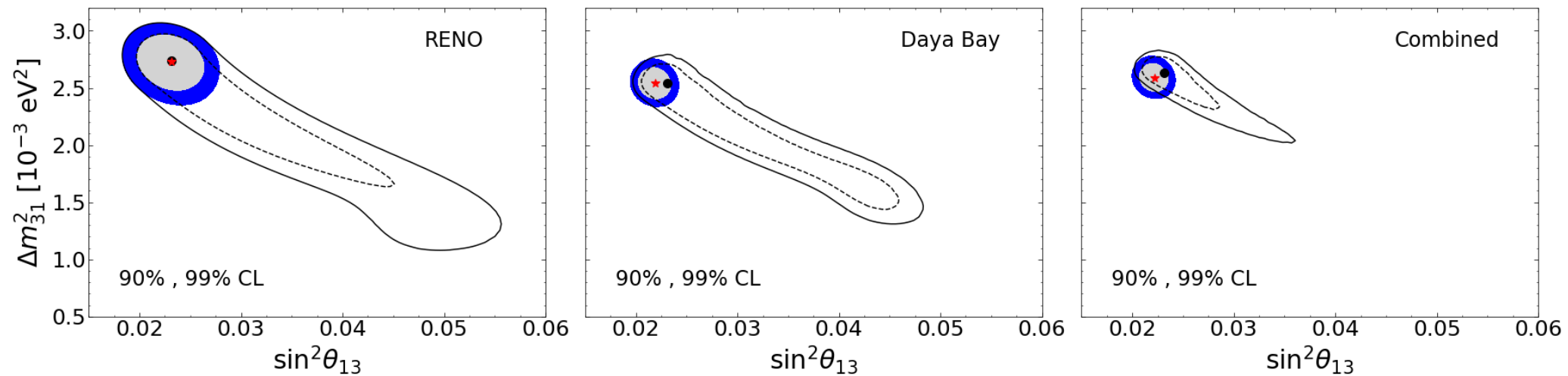
Daya Bay, PRL, 1809.02261



# Bounds from current experiments

Sensitivity is reduced when including the wave packet width in the analysis

The role of RENO is more important here than in the standard analysis

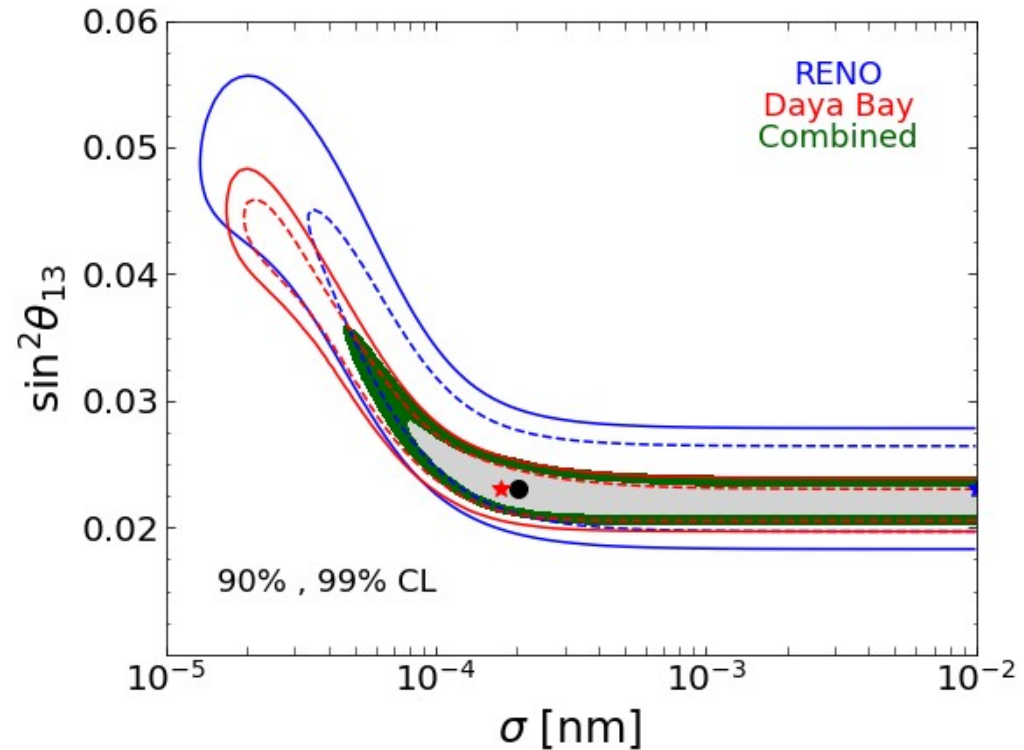
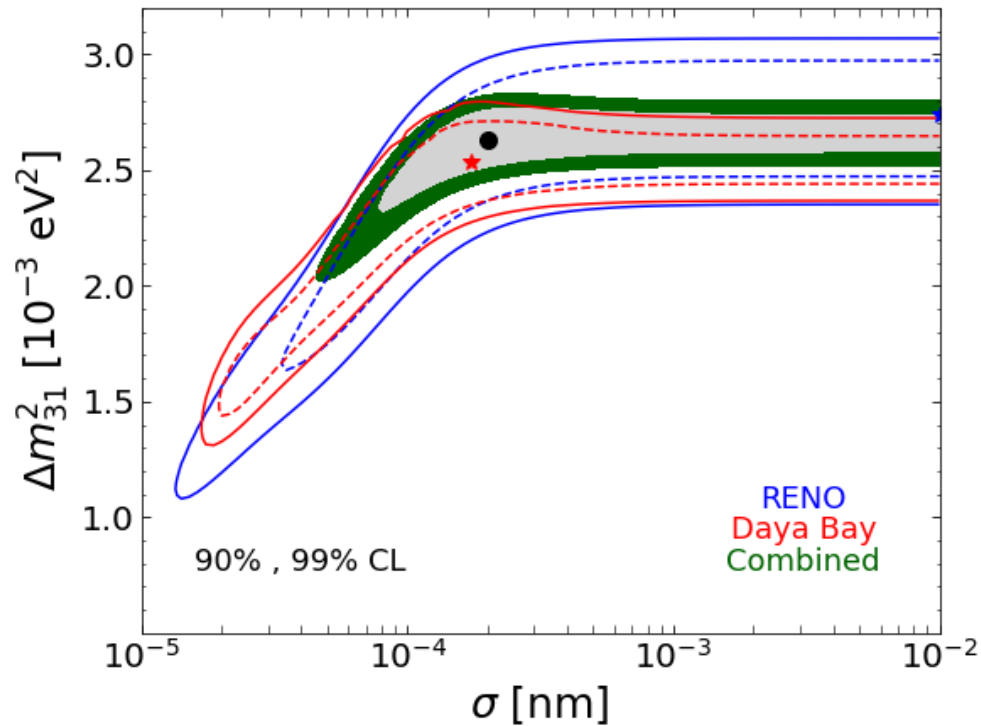


arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

# Bounds from current experiments

The reduction of sensitivity is due to a new correlation between the standard parameters and the wave packet width

Small values of sigma a correlated with large (small) values of the mixing angle (mass splitting)



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

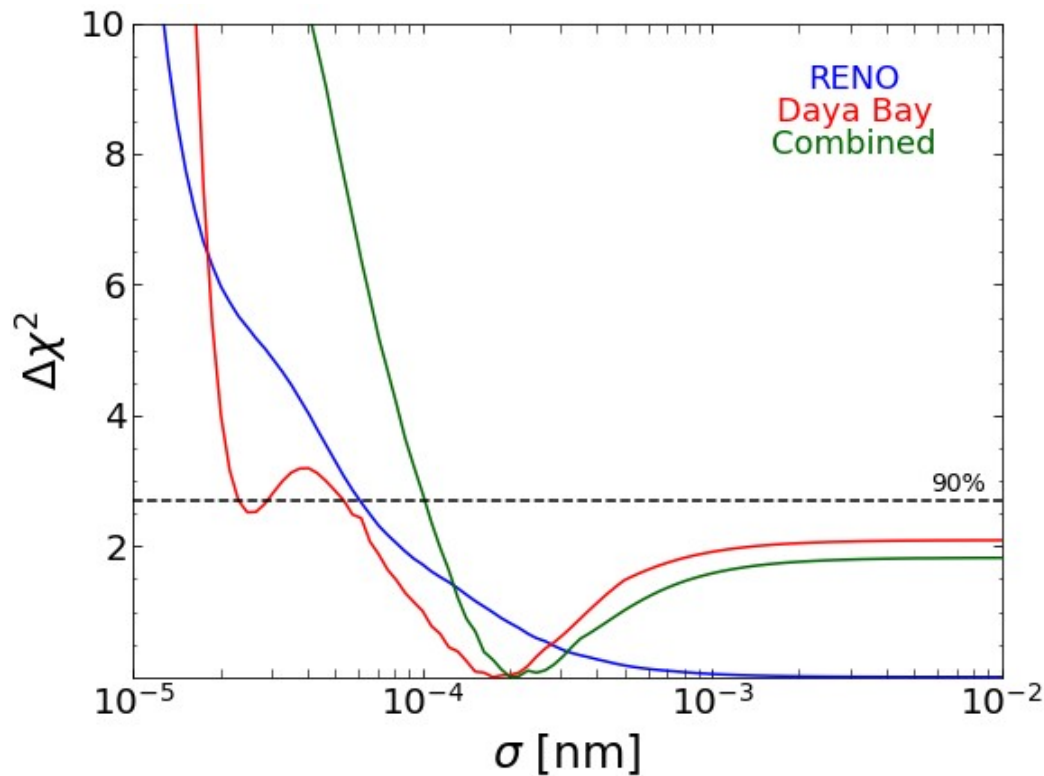


# Bounds from current experiments

From the combined analysis we obtain a lower bound:

$$\sigma > 1.0e-4 \text{ nm at 90\% CL}$$

For larger values of sigma the effect in the oscillation probability disappears. The lines extend to infinity.



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

# Sensitivity at JUNO

## JUNO:

12 reactors

1 far detector (+ 1 near detector under consideration)

JUNO will measure the solar parameters and the atmospheric mass splitting at below 1%

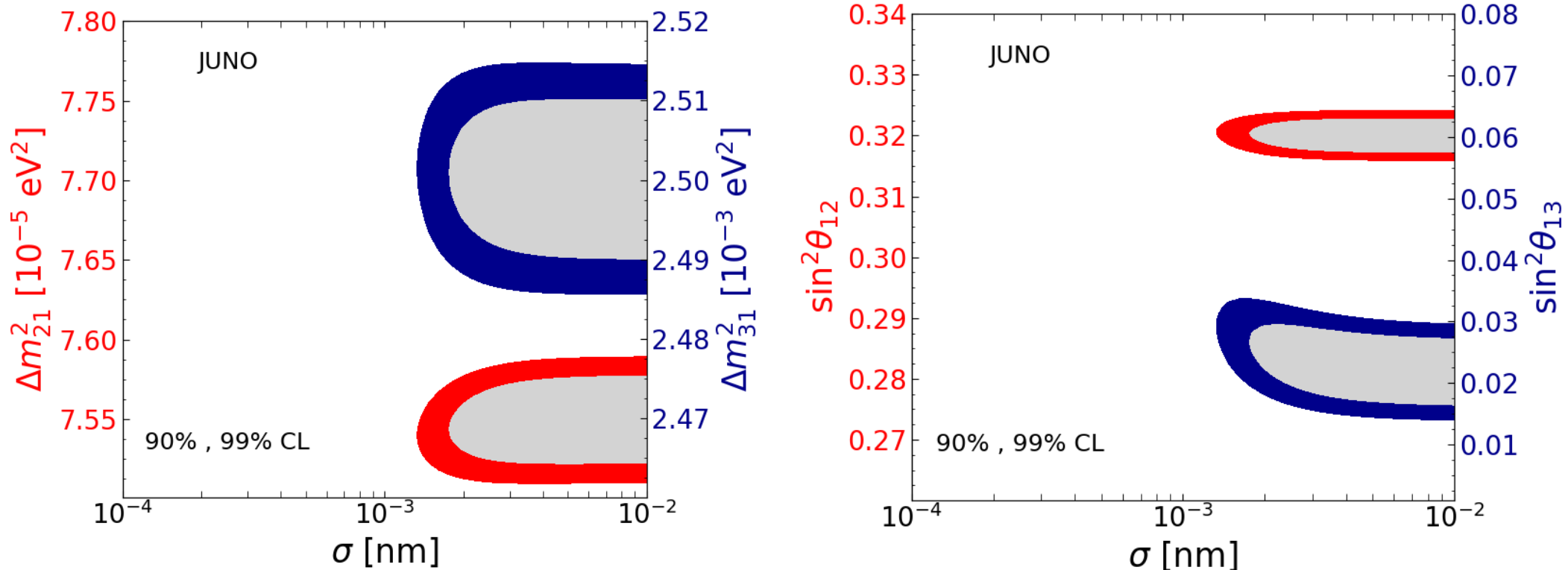
Sensitivity to both oscillation lengths should break the degeneracy observed in current experiments

JUNO, J.Phys. G, 1507.05613

JUNO+IceCube, PRD, 1911.06745

# Excluding decoherence

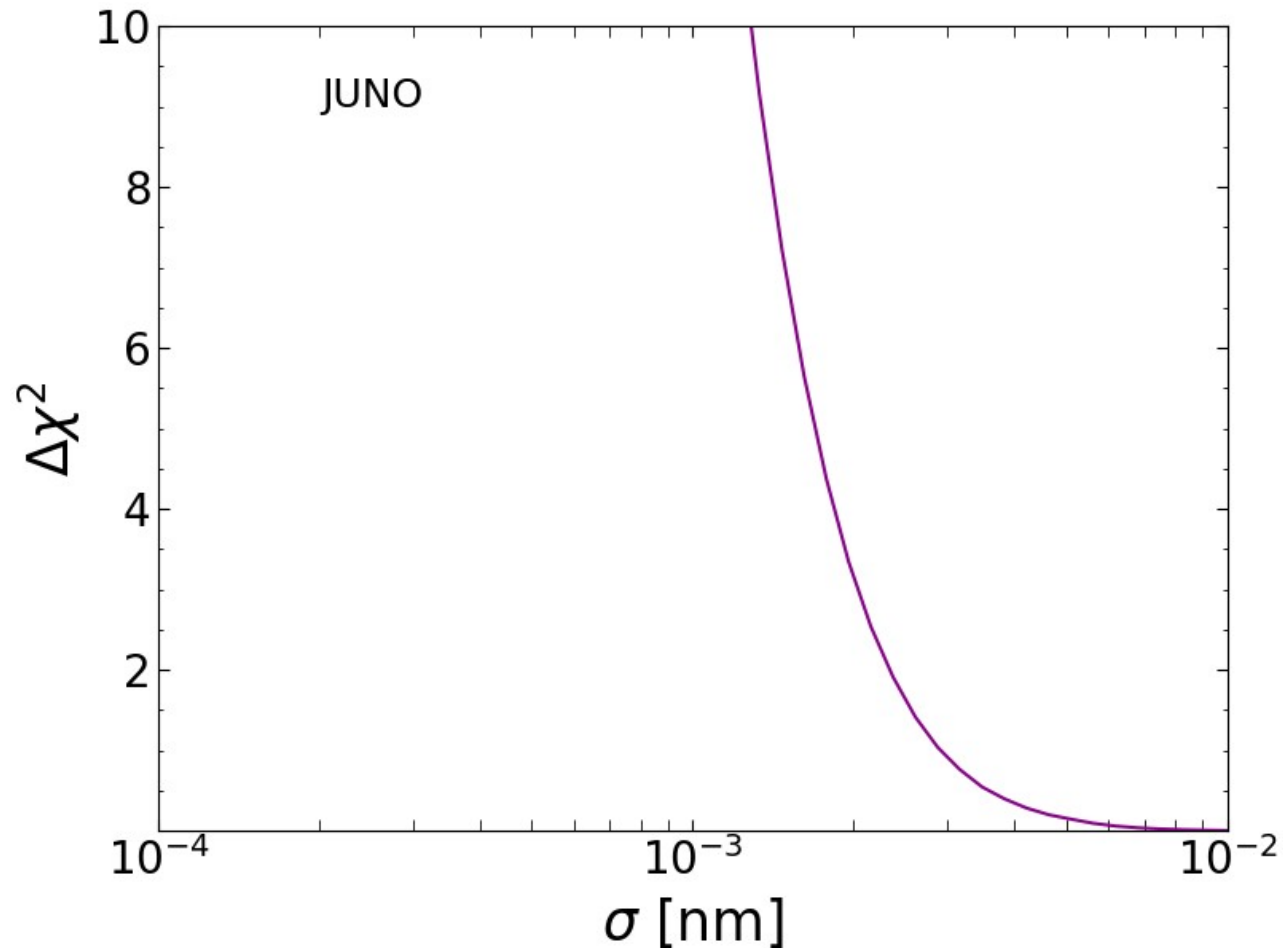
The measurement of the standard neutrino oscillation parameters remains mostly unaffected



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

# Excluding decoherence

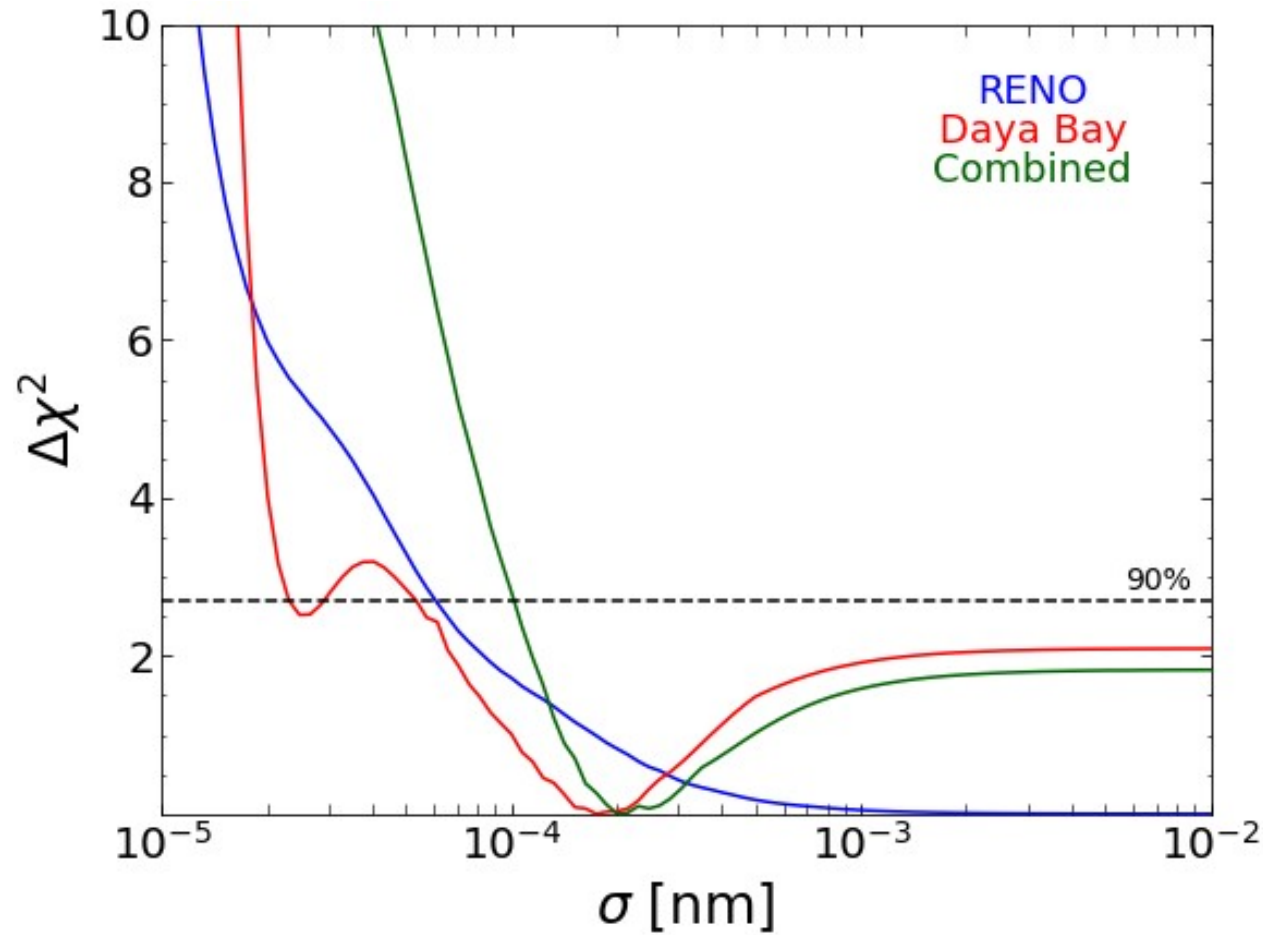
A bound more than a factor 20 stronger than the current one can be obtained after 6 years of data taking



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

# Measuring decoherence

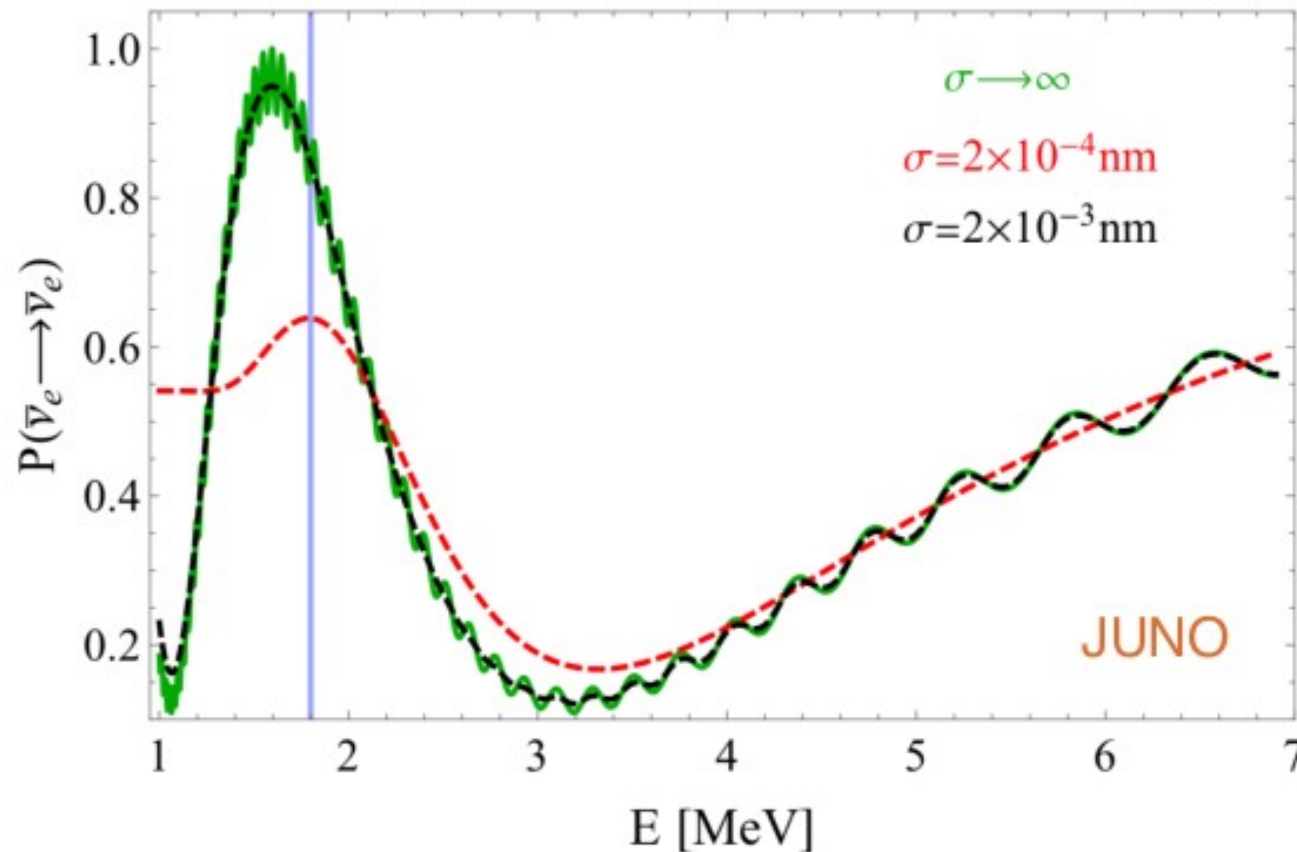
We create the fake data with the best fit value obtained in the analysis of current data



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

# Measuring decoherence

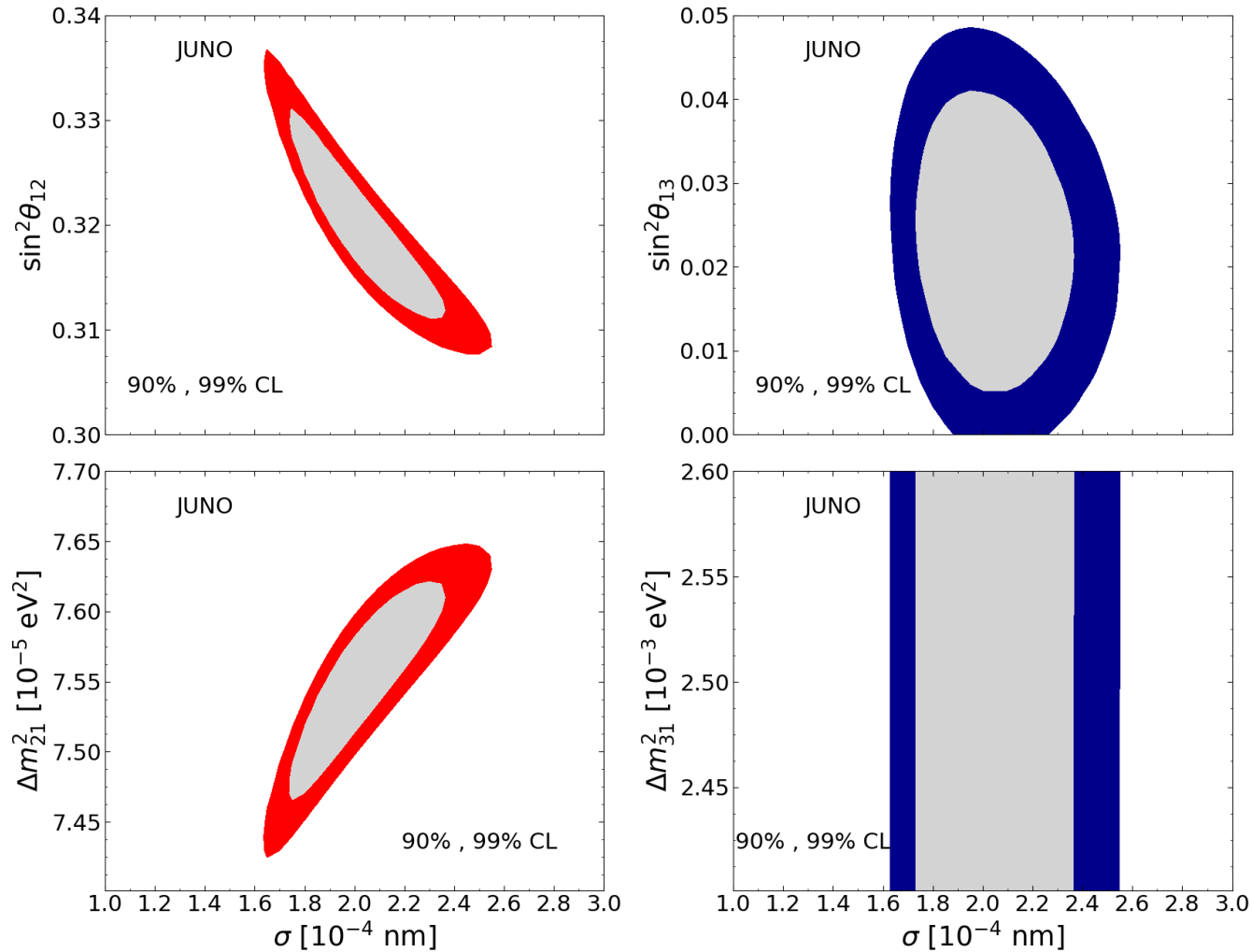
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# Measuring decoherence

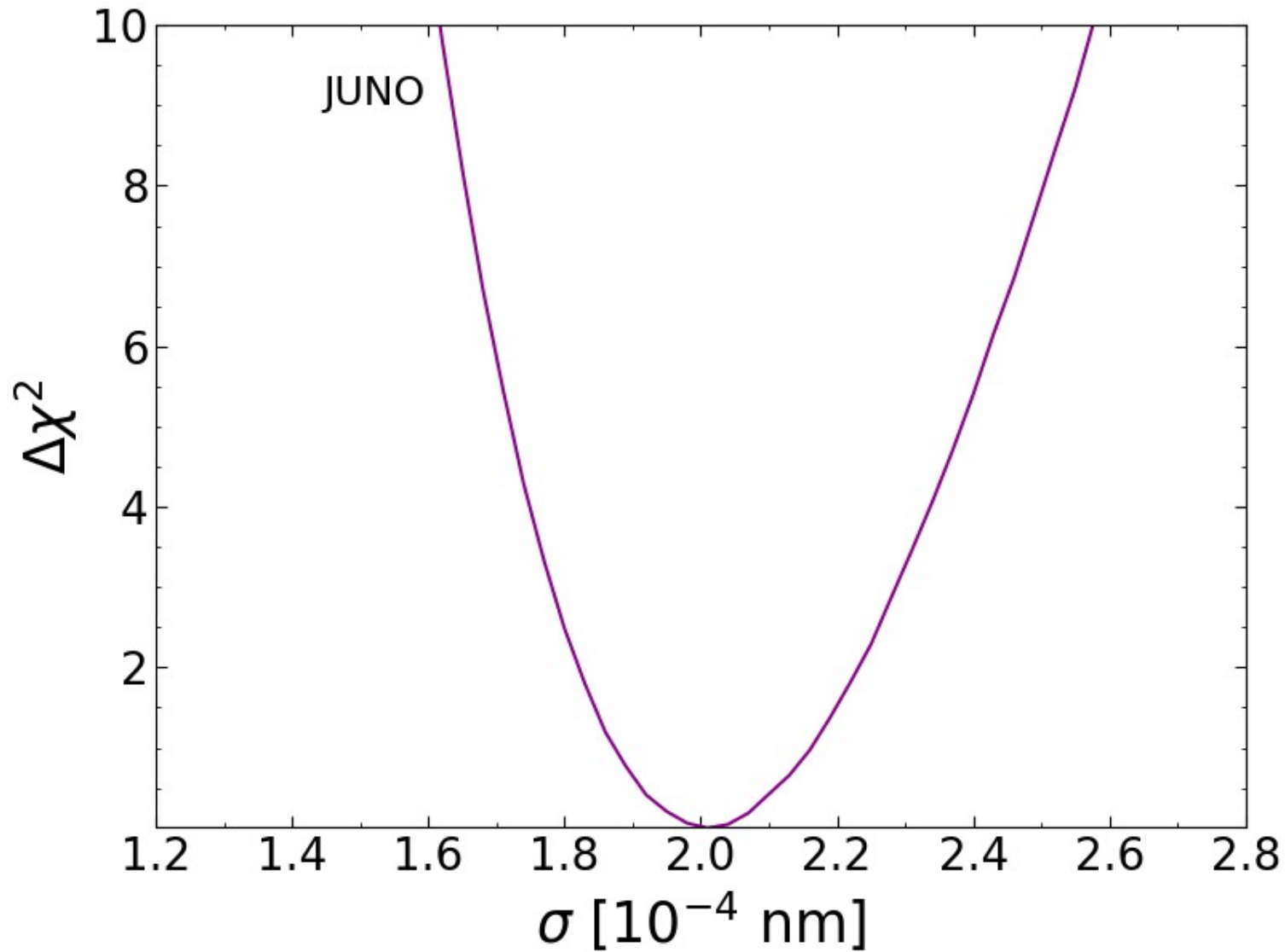
The correlations re-appear because now a measurement of the atmospheric angle is not possible anymore



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes

# Measuring decoherence

The coherence scenario would be excluded at more than 10 sigma



arXiv:2005.03022, de Gouvêa, De Romeri, Ternes



# Conclusions

We analyzed data from RENO and Daya Bay to obtain lower bounds in the neutrino wave packet width

We find that  $\sigma > 1.0e-4$  nm at 90% CL

Assuming 6 years of running time for JUNO this bound could be improved by a factor of 20

If the true value for  $\sigma$  lies within the sensitivity of JUNO, a clear measurement would be possible