# Fermilab SCD Post-Doc Talk

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#### A quick reminder about myself...



#### SBN Program



#### SBN Program



#### 1. Explore Sterile $\nu$ Oscillations





- Detailed sensitivity analysis includes cosmics backgrounds, correlated flux & cross section systematics, and beam backgrounds.
  - The SBN program will exclude the LSND 99% CL region at  $5\sigma$  in 3 years of taking data in the BNB

![](_page_4_Figure_5.jpeg)

#### 2. MiniBooNE Follow-up

![](_page_5_Figure_1.jpeg)

![](_page_5_Figure_2.jpeg)

#### Data Collected by MicroBooNE

- Surface-based, 85 ton active volume liquid argon
- Collecting cosmic and neutrino data since Fall 2015 with good uptime and purity
- Target is to make data collected until July 2019 available to be analyzed

![](_page_6_Figure_4.jpeg)

![](_page_7_Figure_0.jpeg)

![](_page_8_Figure_1.jpeg)

![](_page_9_Figure_1.jpeg)

![](_page_10_Figure_1.jpeg)

![](_page_11_Figure_1.jpeg)

![](_page_12_Figure_1.jpeg)

![](_page_13_Figure_1.jpeg)

### When electron drift through the argon...

![](_page_14_Figure_1.jpeg)

### When electron drift through the argon...

![](_page_15_Figure_1.jpeg)

A bit more about space charge...

![](_page_16_Figure_1.jpeg)

Two distinct features:

- A **squeezing** of the sides of the tracks in the transverse TPC directions that somewhat resemble a rotation ("A")
- A **bowing** of the track toward the cathode that is most pronounced in the middle of the TPC ("B").
- We simulate this effects and correct it out at the calibration stage.

#### dQdx Calibration

# Corrects for the misconfigured TPC channels, space charge effect, transverse diffusion

![](_page_17_Figure_2.jpeg)

$$(dQ/dx)_{Corrected}^{data} = (dQ/dx)_{Reconstructed} \cdot C_{YZ}$$

$$(dQ/dx)_{Corrected}^{MC} = (dQ/dx)_{Reconstructed} \cdot C_{YZ}$$

This YZ map is used to "smear" simulation to mimic data when we overlaid some cosmic data as part of our simulation also known as "overlay"

#### dQdx Calibration

- First calibration in LArTPC using the dedicated data-driven Space Charge Effects map provided by the University of Bern, CSU, and University of Minesota colleagues.
- Show that we can recover our "perfect calibration" sample to a good precision.
- Better detector simulation effects that better match our data as the input for the simulation

![](_page_18_Figure_4.jpeg)

## When electron drift through the argon...

![](_page_19_Figure_1.jpeg)

# Electron Lifetime

 Once the SCE are corrected, we can disentangle the electron lifetime effect from the convoluted : and t dependent detector effects

$$\frac{Q_A}{Q_C} = \exp(-t_{\rm drift}/\tau) \implies \tau_0 = \frac{(1/slope)}{v_d}$$

• Errors on the measurement is < 0.5%

![](_page_20_Figure_4.jpeg)

![](_page_20_Figure_5.jpeg)

Ensure that we are extracting the lifetime within 1% for smaller lifetime

The value extracted in data will be used as the input drift lifetime and will be calibrated later

![](_page_21_Figure_0.jpeg)

![](_page_22_Figure_1.jpeg)

# Flux and Cross Section Systematics

The flux and cross-section systematics are all calculated by weighing a central value according to different variations.

For example, axial mass for CC quasi-elastic,  $\pi$  absorption probability for neutrino interactions, and horn current,  $\pi^{\pm}$  production in flux.

The covariance matrix :

$$E_{ij}^{sys} = \frac{1}{N} \Sigma_k^N \left( N_i^{CV} - N_i^k \right) \left( N_j^{CV} - N_j^k \right)$$

Where N is the total number of variations,  $N_i^k$  is the value in the CV in the *i*-th bin of the k -th variation and  $N_i^{CV}$  is that of the central value.

![](_page_23_Figure_6.jpeg)

#### **Detector Systematics**

![](_page_24_Figure_1.jpeg)

- New method for detector systematics based on comparisons between data and MC, propagated to physics analyses by wire modification.
- Consider the detector's response to an energy deposition as a function of the relevant variables: x, (y, z), (θ, φ), and dE/dx
- Characterize the detector's response in terms of the charge and width of gaussian hits — using these as proxy for wire waveform properties

#### **Detector Systematics**

- Measure the values of Q and  $\sigma$  of hits vs. x
- Look at difference in data and simulation to bound detector differences in Q and σ, then perform a fit to those points to get the continuous functions RQ and Rσ
- Modify hits/wavefroms in simulation by that difference and rerun reconstruction, treating difference as a systematic
- The same ratio is also extracted in other variables and then propagated as alternate samples to the central value.

![](_page_25_Figure_5.jpeg)

#### Hit Charge Run 1 vs MC CV Ratio

![](_page_26_Figure_0.jpeg)

![](_page_27_Figure_1.jpeg)

## **Event Selection**

- Working with 'Pandora' pattern-recognition reconstruction framework (standard reconstruction tool in MicroBooNE) focusing on the excess related to electron channels.
- Muon Neutrino selection is used to constrain systematics

![](_page_28_Figure_3.jpeg)

## SBNfit

- Framework designed to perform simultaneous fits across data from multiple, correlated distributions.
- Developed by MicroBooNE/SBN collaborators at Columbia University (Georgia Karagiorgi, Mark Ross–Lonergan, Guanqun, Davio Cianci, et al.)

Multi-mode	Multi-detector	Multi-channel
<ul> <li>Neutrino/anti- neutrino</li> <li>BNB/Numi Beam</li> </ul>	<ul> <li>SBN:</li> <li>SBND+MicroBooNE+ICA RUS</li> <li>MiniBooNE+MicroBooNE</li> <li>SBN+DUNE</li> </ul>	<ul> <li>1 electron only</li> <li>1 electron + N proton</li> <li>1 muon + N proton</li> </ul>

Allows for combined fitting of *arbitrarily large* number of modes, detectors and channels simultaneously, fully accounting for systematic correlations.

![](_page_30_Figure_0.jpeg)

![](_page_31_Figure_1.jpeg)

## **Constraining Systematics**

Each detector shares the same neutrino flux and argon cross-sections measurement is highly correlated.

Exploit the correlations between the  $\nu_{\mu} - \nu_{e}$  channels to reduce systematic uncertainties.

Combined analysis:

1) using multiple selections and observables

2) taking into account correlated systematic uncertainties through a covariance matrix

![](_page_32_Figure_6.jpeg)

10

15

20

Global Bin Number

nu uBooNE nue intrinsic

0

5

-0.2

-0.4

-0.6

# Not final selection

#### Constraining Systematics using $\nu_{\mu}$ selection

![](_page_33_Figure_2.jpeg)

![](_page_34_Figure_0.jpeg)

![](_page_35_Figure_1.jpeg)

#### Cartoon example

Sensitivity

[On-going work]

Have to simulate many experiments under the H0 hypothesis (no low energy excess), and H1 hypothesis .

For each toy experiment,  $\Delta \chi^2$  is calculated as  $\Delta \chi^2 = \sum_{i,j=1}^N \left( (n^i - \mu^i_{H_1}) E^{-1}_{i,j} (n^j - \mu^j_{H_1}) - (n^i - \mu^i_{H_0}) E^{-1}_{i,j} (n^j - \mu^j_{H_0}) \right)$ 

to obtain the  $\Delta \chi^2$  distribution.

The sensitivity quoted is the median sensitivity.

![](_page_36_Figure_7.jpeg)

## Looking Ahead

- Finalizing the systematics constraint, to include the  $1e0p0\pi$  and  $\nu_{\mu}$  channel in the constraint.
- Working on including detector systematics into the constraint and the final sensitivity.
- Finalizing the analysis internal note to be circulated to the group's Editorial Board
- Target: Neutrino 2020.

![](_page_38_Figure_0.jpeg)

## SBNfit $\chi^2$ Test Statistics for Oscillation Analysis

![](_page_39_Figure_1.jpeg)

# Feldman cousins method

![](_page_40_Figure_1.jpeg)

- 1. Given observed data spectrum D
- 2. Find the grid point with the minimum  $\chi^2$
- 3. Calculate  $\Delta \chi^2$  at each grid point
- 4. Calculate exactly the value of that 90% of experiments would be in by generating pseudo experiments.
- 5. Do this for  $3\sigma$  (10<sup>4</sup> pseudo experiments/"universes") and  $5\sigma$  (~10<sup>8</sup> pseudo experiments/"universes")

## Accelerating SBNFit Feldman Cousins on HPC

![](_page_41_Figure_1.jpeg)

#### Accelerating SBNFit Feldman Cousins on HPC

![](_page_42_Figure_1.jpeg)

![](_page_43_Figure_0.jpeg)

#### Scaling with N universes & N grid points

![](_page_44_Figure_1.jpeg)

![](_page_44_Figure_2.jpeg)

Scaling of the program run time with the number of universes, demonstrating a linear dependence on Nuniv Scaling of the program run time with the number of grid points, demonstrating a quadratic dependence on Nuniv.

#### Scaling on single node

![](_page_45_Figure_1.jpeg)

Measurement of the single Haswell node performance for a fixed problem size.

![](_page_45_Figure_3.jpeg)

Measurement of the single KNL node performance for a fixed problem size

### Scaling on Multi-node

![](_page_46_Figure_1.jpeg)

Strong scaling measurements on Haswell nodes.

Generally good scaling up to the point where the amount of work per rank becomes relatively small.

$\Delta (Y \cdot O)$	$N_{ m univ} = 10^4 \ (3\sigma)$	$N_{ m univ} = 10^8 \ (5\sigma)$
Cori phase 1 (Haswell)	$7.2  imes 10^5$	$7.2 \times 10^{3}$
Cori phase 2 (KNL)	$1.2  imes 10^6$	$1.2 \times 10^4$

Table 3: Upper boundaries on grid sizes that can be processed when running *a full day* on all of Cori phase 1/2.

# Looking Ahead

- Technical paper that highlights the improvement gained from the parallelization and its scalability at HPC titled "Grid-based minimization at scale: Feldman-Cousins corrections for SBN" has been submitted to arXiv: https://arxiv.org/abs/2002.07858
- Physics paper
  - A broader-scope SBN sensitivity/physics paper (in discussion)
    - Perform the non-FC method vs FC method. Compare the sensitivity outcomes for the two methods.
    - Testing the effects of different parameterizations of systematics, study biases and model dependencies of the current analysis
- Plan is to continue optimizing SBNfit fitting framework using GPU accelerators (Kokkos to replace Eigen3) and move away from the grid-based approach to allow us to probe higher dimension (3N+2 scenario).
- Plan to present the results on technical conferences (ICHEP).

## Backup