Higgs signatures of MSSM Electroweak Baryogenesis

A. Menon
University of Michigan
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Based on work with:

David E. Morrissey, [arXiv:0903.3038]

Motivation: Baryon Asymmetry of Universe

Baryon density of the Universe:

$$\eta = \frac{n_B}{n_\gamma} = (6.5 \pm 0.3) \times 10^{-10}$$

where n_B is the excess number density of baryons and n_γ is the number density of photons. WMAP '06

 Baryon asymmetry of the Universe cannot be explained within the Standard Model.

Outline

- 1) The Mechanism of Electroweak Baryogenesis:
 - a) The Electroweak phase transition at finite temperature
 - b) Bubble Nucleation
 - c) Baryon Production
 - d) Sphalerons
- 2) Realizing Electroweak Baryogenesis
 - a) Electroweak Baryogenesis in the Standard Model
 - b) Electroweak Baryogenesis in the MSSM
 - c) The EWBG MSSM parameter space.
- 3) Collider implications of EWBG MSSM
 - a) Light stop searches at the Tevatron and LHC
 - b) Searches for Stoponium
 - c) Indirect Higgs Searches
- 4) Summary and Outlook

The Sakharov Conditions for Baryon Asymmetry

- Necessary conditions for generating baryon asymmetry (BAU) are:
 - 1) Baryon number B violation
 - 2) CP violation
 - 3) Violation of thermal equilibrium
- B violation clearly necessary for generating BAU

• B is odd under C and CP \Rightarrow BAU

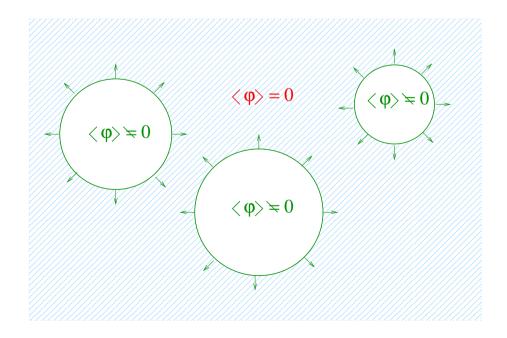
• B is also odd under $T \Rightarrow BAU$

The Electroweak Baryogenesis Scenario

Baryon production occurs at the electroweak phase transition.

[Kuzmin, Rubakov, Shaposhnikov '85]

- Electroweak symmetry breaking occurs as universe cools nucleating bubbles of broken phase.
- Baryon production occurs near the walls of the expanding bubbles.



The Electroweak Phase Transition

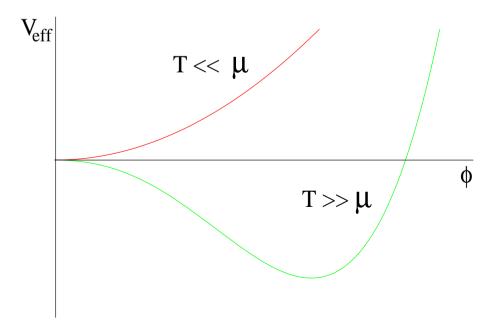
• The Higgs VEV $\langle \phi \rangle$ is the order parameter

$$\langle \phi \rangle = 0 \Rightarrow SU(2) \times U(1)_Y$$
 Unbroken

$$\langle \phi \rangle \neq 0 \Rightarrow SU(2) \times U(1)_Y \rightarrow U(1)_{em}$$

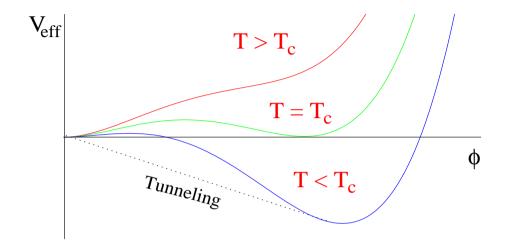
ullet Effective potential at finite temperature T

$$V_{eff} = (-\mu^2 + \alpha T^2)\phi^2 - \gamma T\phi^3 + \frac{\lambda}{4}\phi^4 + \dots$$

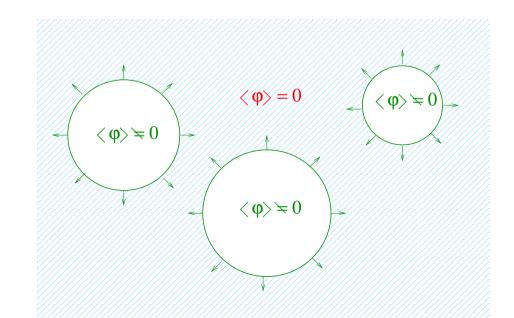


Bubble Nucleation

• First order phase transition

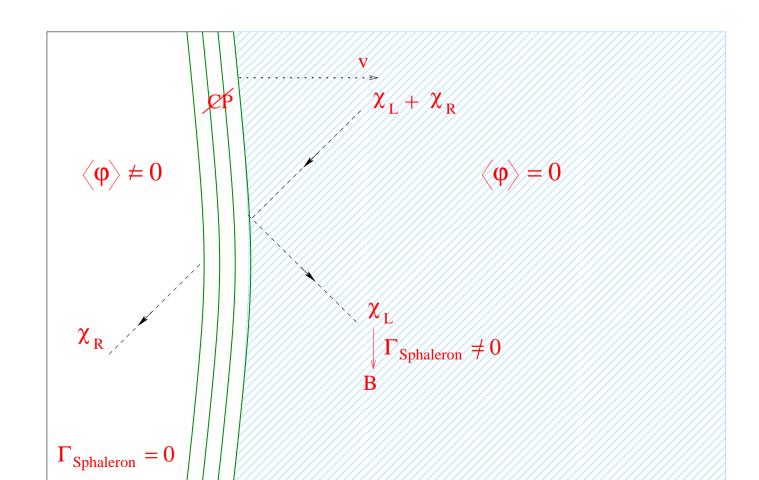


ullet Bubbles of broken phase nucleate for $T\lesssim T_c$



Baryon Production

- CP violation in bubble wall \Rightarrow chiral flux into the unbroken phase.
- Sphaleron transitions in the unbroken phase converts the chiral flux into baryons.
- Baryons produced at the bubble wall, enter the bubble as it expands.



Sphalerons

- $U(1)_{B+L}$ is anomalous in the SM and is violated by quantum effects.
- T=0 Instanton effects drive transitions

$$\Gamma \sim e^{-16\pi^2/g_2^2} \simeq 10^{-320}$$
.

• $T \neq 0$ Thermal fluctuations (sphaleron) drive transitions

$$\Gamma \sim \left\{ \begin{array}{ccc} T^4 e^{-4\pi\langle\phi\rangle/gT} & \langle\phi\rangle \neq 0 & \text{[Arnold,Mclerran '87]} \\ \alpha_w^4 T^4 & \langle\phi\rangle = 0. & \text{[Bodeker,Moore,Rummukainen '99]} \end{array} \right.$$

EWBG in the Standard Model

• Electroweak phase transition is first-order only if [Kajantie et. al.'98]

$$m_h \lesssim 70 \text{ GeV}$$

LEP experimental mass bound

$$m_h > 114.4 \text{ GeV}$$

• Not enough CP violation is SM [Gavela et. al. '94]

EWBG does not work in the Standard Model.

EWBG MSSM: First Order Phase Transition

• For a first order phase transition for

$$V_{eff} = (-\mu^2 + \alpha T^2)\phi^2 - \gamma T\phi^3 + \frac{\lambda}{4}\phi^4 + \dots$$

we need

$$rac{\langle \phi(T_c)
angle}{T_c} \simeq rac{\gamma}{\lambda} > 1$$

- \bullet γ is generated bosonic loops and the dominant contribution in the MSSM is from a light mostly right-handed stop. [Carena,Quiros,Wagner]
- Larger stop mass \Rightarrow large $\lambda \Rightarrow$ larger m_h .

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• The finite temperature effective Higgs potential in the MSSM has a term

$$V_{eff}^{MSSM} \subset -\frac{N_c T}{6\pi} (m_{\tilde{t}}^2 + \Pi_R(T)^2)^{3/2}$$

where $\Pi_R(T)^2$ is the positive self energy contribution and

$$m_{\tilde{t}}^2 \sim m_{U_3}^2 + y_t^2 \phi^2 \left(1 - \frac{|X_t|^2}{m_{Q_3}^2} \right)$$

• Generating a ϕ^3 term needed for a first order phase transition $\Rightarrow m_{U_3}^2$ cancels $\Pi_R(T)^2$ and $X_t \ll m_{Q_3}$.

EWBG MSSM: Light Stop Implications

- $\Delta \rho$ limits \Rightarrow light stop mass must be mostly right handed.
- For a strong first order phase transition and to avoid color breaking minima

$$-(100 \text{GeV})^2 \lesssim m_{U_3}^2 \lesssim 0$$
; $X_t/m_{Q_3} \lesssim 0.5$
 $\Rightarrow 120 \lesssim m_{\tilde{t}_1} \lesssim 170 \text{GeV}$.

Avoiding LEP limit on Higgs mass:

$$m_{Q_3} \gtrsim$$
 2TeV; $M_A \gtrsim$ 200 GeV ; 5 $<$ tan $\beta <$ 10 $\Rightarrow m_h \lesssim$ 125GeV

• Light stop search limits for $m_{\tilde{t}_1} \leq m_t$:

$$(m_{\tilde{t}_1} - m_{Wimp}) \lesssim 30 \text{GeV}.$$

as decay products are soft and difficult to find.

• Light stop coannihilations with Bino LSP \Rightarrow viable Ωh^2 for $m_{\tilde{t}_1}-m_{Wimp}\lesssim$ 30 GeV. [Balazs, Carena, A.M., Morrissey, Wagner '05].

EWBG MSSM: CP Violation

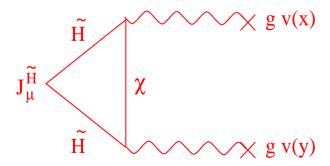
Higgsino is an important source of MSSM CP violation phase

$$M_{ ilde{\chi}^\pm} \sim \left(egin{array}{cc} |M_2| & g_2 v_u(z) \ g_2 v_d(z) & e^{i\phi} |\mu| \end{array}
ight)$$

where ϕ is the $Arg(\mu M_2)$.

• CP violation in bubble wall modifies the transmission and reflection coefficients leading the current

$$\langle J_{\mu}^{ ilde{H}}
angle \propto Im(\mu M_2)$$



ullet $\langle J_{\mu}^{H} \rangle$ in turn sources B violating process through sphalerons.

EWBG MSSM: Implications of CP Violation

• Generating baryon asymmetry:

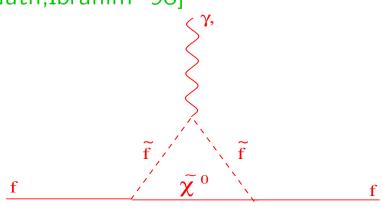
$$Arg(\mu M_{1,2}) \gtrsim 10^{-2}$$

 $\mu M_{1,2} \lesssim 400 \text{ GeV}$

[Carena, Quiros, Seco, Wagner '02, Lee, Cirigliano, Ramsey-Musolf '04]

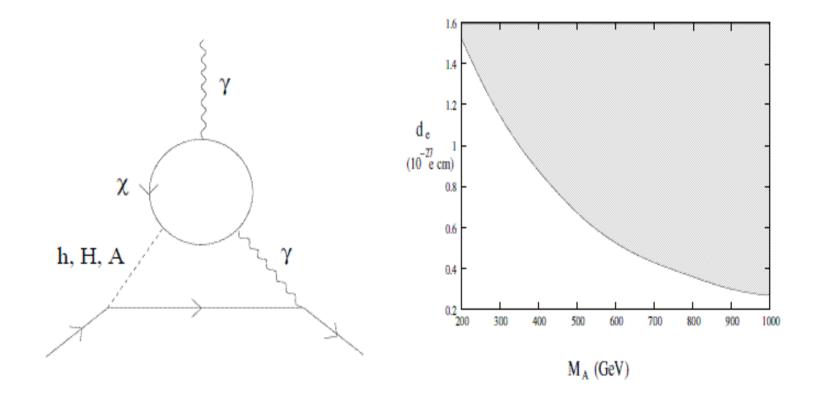
• EDM constraints:

• One loop EDM constraints $\Rightarrow m_{\tilde{f}}^2 > 2-5$ TeV or the first two generations are decoupled.[Nath,Ibrahim '98]



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ullet Two-loop contribution: $\propto Im(\mu M_2)$ [Chang, Chang, Kueng '02; Pilatfsis '02]



• EDM constraints weaker for $M_A \gtrsim 1$ TeV.

Split Stop Supersymmetry

■ EDM constraints ⇒ large sfermion and pseudo scalar masses.

$$m_{\widetilde{f}} \gtrsim$$
 5 TeV

ullet LEP Higgs mass limit \Rightarrow heavy mostly left-handed stop $m_{ ilde{t_2}} \gtrsim 2$ GeV.

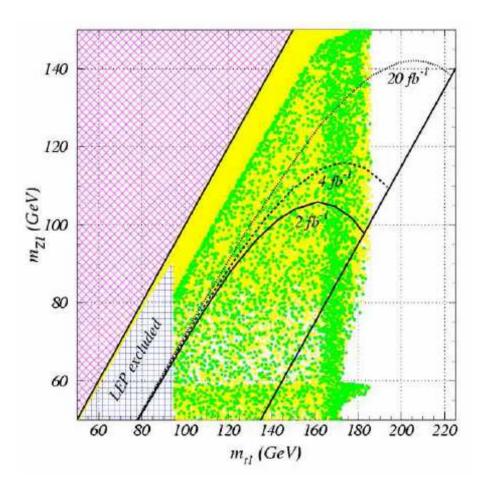
EWBG ⇒ light charginos, neutralinos and right handed stops.

$$m_{LSP} < m_{\tilde{t_1}} \lesssim m_t$$

⇒ Stop Split Supersymmetry. [Carena, Nardini, Quiros, Wagner '09]

Direct light stop searches: Tevatron

ullet Light stop searches the in $ilde t o c ilde \chi^0$ channel: [Balazs, Carena, Wagner '04]



Direct light stop searches: LHC

- \bullet \tilde{t}_1 produced copiously, but decays into soft jets so hard to find.
- Possible channels $\tilde{t}_1 \to c\chi_1^0$, $\tilde{t}_1 \to bW^*\chi_1^0$, $\tilde{t}_1 \to b\chi_1^{+*}$. [Balazs et. al. '04; Demina et. al. '99; Hiller, Nir'08]
- \bullet Can be seen in gluino pair production and decay into same sign leptons, jets and MET with 30 fb $^{-1}$ for $m_{\widetilde{g}}\lesssim 1$ TeV. [Kramel, Raklev '05,06]
- Parameter space determination is generally challenging.

Direct light stop searches: Stoponium

- $\eta_{\tilde{t}_1}=\tilde{t}_1\tilde{t}_1^*$ can form as $\Gamma_{\tilde{t}_1\to c\chi_1^0}\ll$ binding energy of $\eta\tilde{t}_1$. [Nojiri, Dress '95]
- $\eta_{\tilde t_1} \to \gamma \gamma$ may be observable with 100 fb⁻¹ for $m_{\eta \tilde t_1} \lesssim$ 250 GeV. [Martin '08]

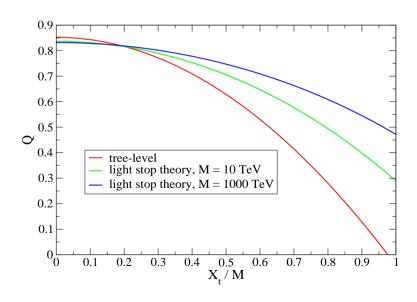
Precise measurement of stoponium mass.

Indirect light stop effects on Higgs Searches

- Light stops can modify Higgs production cross-sections. [Kane et. al. '95; Dawson et. al. '96; Djouadi '98; Dermisek, Low '07]
- Effective light stop Higgs coupling in EWBG

$$g_{h\tilde{t}_1\tilde{t}_1} \sim m_t^2 \left(1 - \frac{|X_t|^2}{m_{Q_3}^2} \right)$$

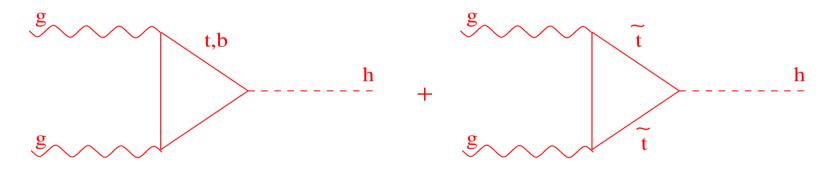
where $Q = \sqrt{2}vg_{h\tilde{t}_1\tilde{t}_1}$



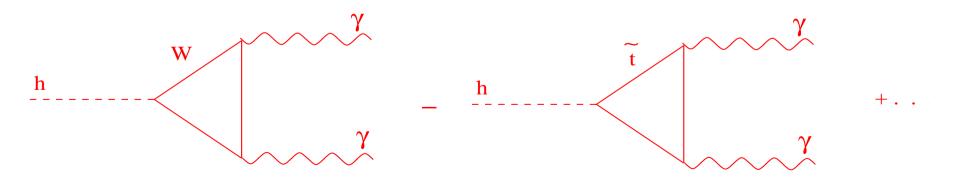
• Stronger first order EW phase transition \Rightarrow larger $g_{h\tilde{t}_1\tilde{t}_1}$ coupling, if $X_t \ll m_{O_2}$.

Light stop loop effects on Higgs Branching Ratios

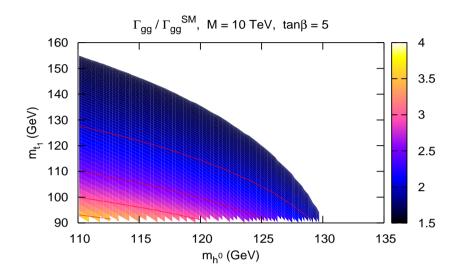
Gluon Fusion

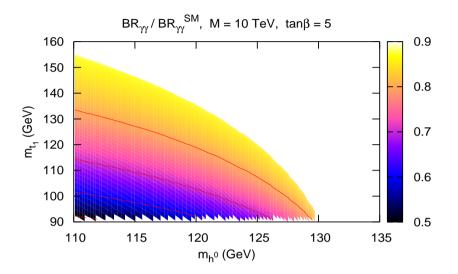


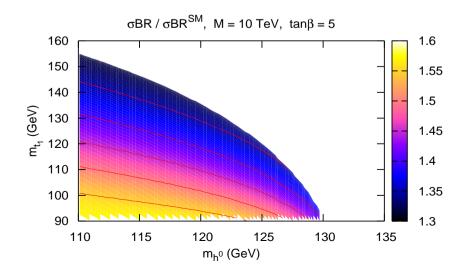
Higgs Photon coupling



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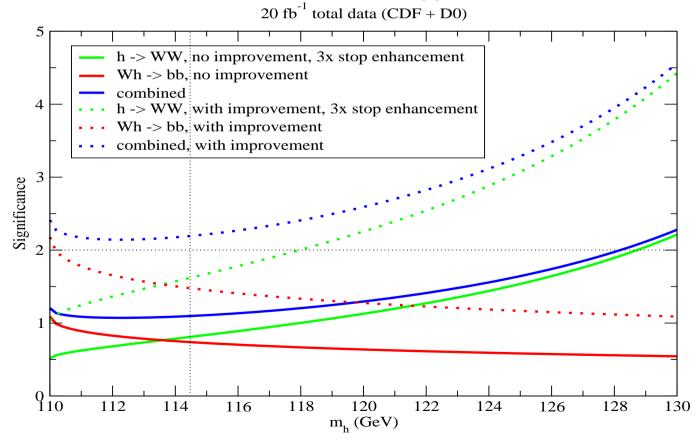




Light stop effects on Higgs searches at the Tevatron

- \bullet For $m_h \lesssim 135$ GeV: Associated Higgs production $Wh \to Wb\bar{b}$ is dominant
- For $m_h \gtrsim 135$ GeV: $gg \to h \to WW^*$ dominant.

Rescaled Tevatron Higgs Results



Improvements in analysis $\Rightarrow 2 - 3\sigma$ observation of the Higgs?

Light stop and Higgs searches at LHC

- 5 σ Higgs discovery in $gg \to h \to \gamma \gamma$ with 30fb⁻¹ and error in Higgs mass 0.2 GeV. [ATLAS Collaboration'09; CMS Collaboration '07]
- 30 % error in Γ_{hgg} with 200fb⁻¹ of data. [Zeppenfeld et. al. '00,02; Belyaev, Reina '02; Duhrssen et. al. '04]
- 20 % error in $\Gamma_{h\gamma\gamma}$ with 200fb⁻¹ of data. [Zeppenfeld et. al. '00,02; Belyaev, Reina '02; Duhrssen et. al. '04]

Large enhancements of Γ_{hgg} in EWBG \Rightarrow deviations in $\Gamma_{ggh}/(\Gamma_{ggh})_{SM}$ should be observable with 200fb⁻¹ at 3σ .

• Finding deviations in $\Gamma_{h\gamma\gamma}$ is more challenging.

Summary

- EWBG MSSM is a very predictive scenario \Rightarrow stop split SUSY light stop, charginos and neutralinos with the remaining scalars being heavy.
- Observing this scenario at the LHC is very challenging.
- Higgs branching fractions are a potential indirect probe of this scenario and could lead to an enhancement of the signal significance at the Tevatron.

Future avenues

- Stoponium decays into leptons are velocity suppressed but lower backgrounds in this channel may make it a viable channel at the LHC.
- Light charginos and neutralino decays involving intermediate stops could be an additional way of probing this scenario at the LHC.