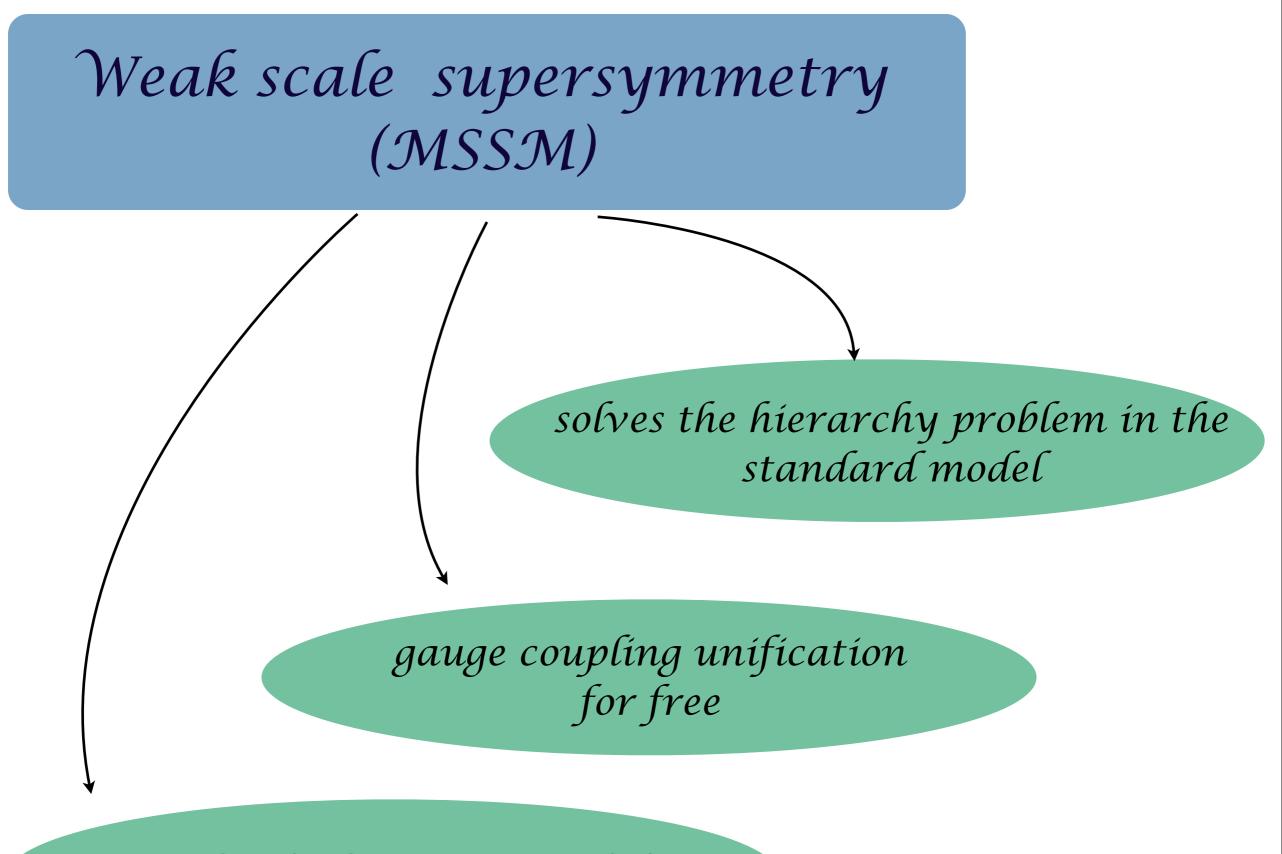


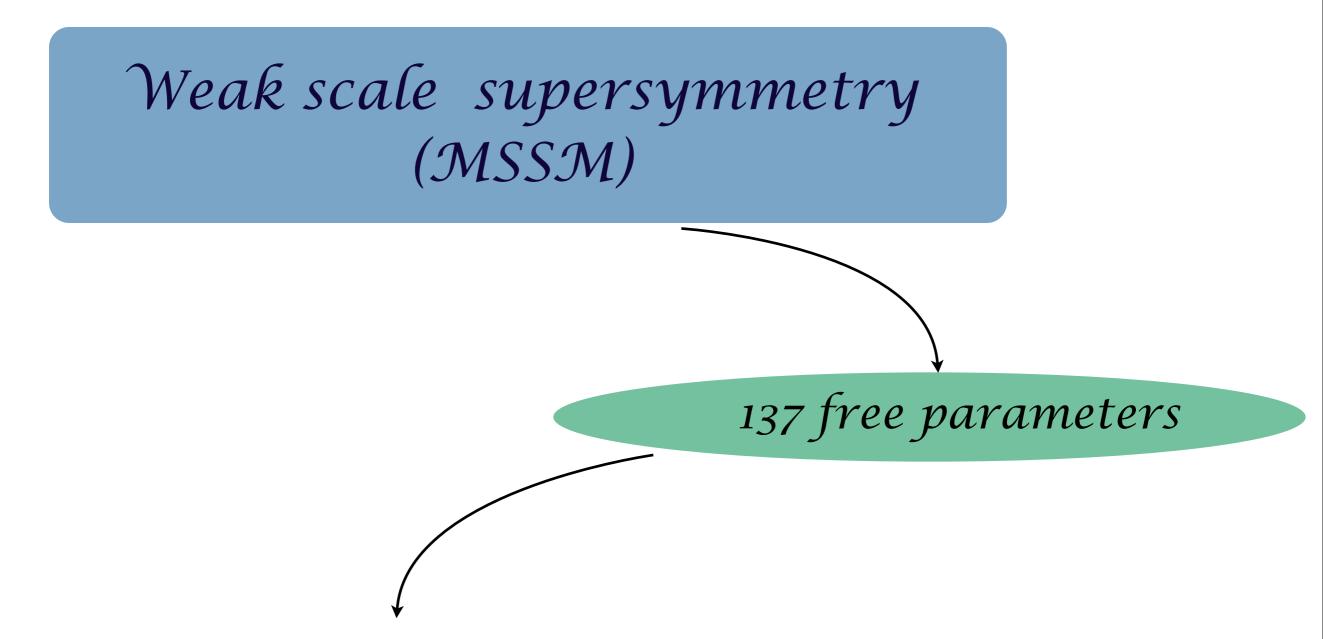
Tuhin Roy

University of Oregon

TR, Schmaltz (arXiv:0708.3593) Murayama, Nomura, Poland (arXiv:0709.0775) Perez, TR, Schmaltz (arXiv:0811.3206)



provídes dark matter candídates



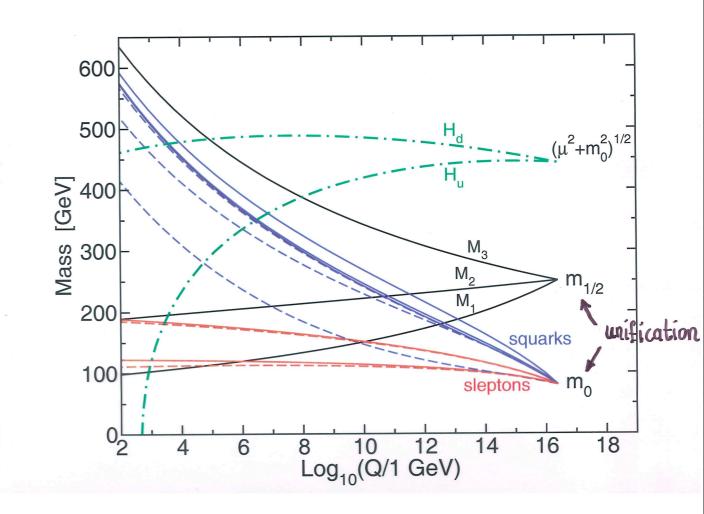
with further assumption of superpartner mass unification no. of parameters is reduced drastically

Superpartner Mass unification

e.g. SO(10)

State of the art MSSM runníng

- 2 loop running
- 1 loop matching at 1 TeV and at GUT
- automated
 - Isajet
 - Softsusy
 - Suspect
 - •
 - -



we missed a potentially very large effect and our scalar mass predictions may be totally wrong

Cohen, TR, Schmaltz (hep/ph 0612100)

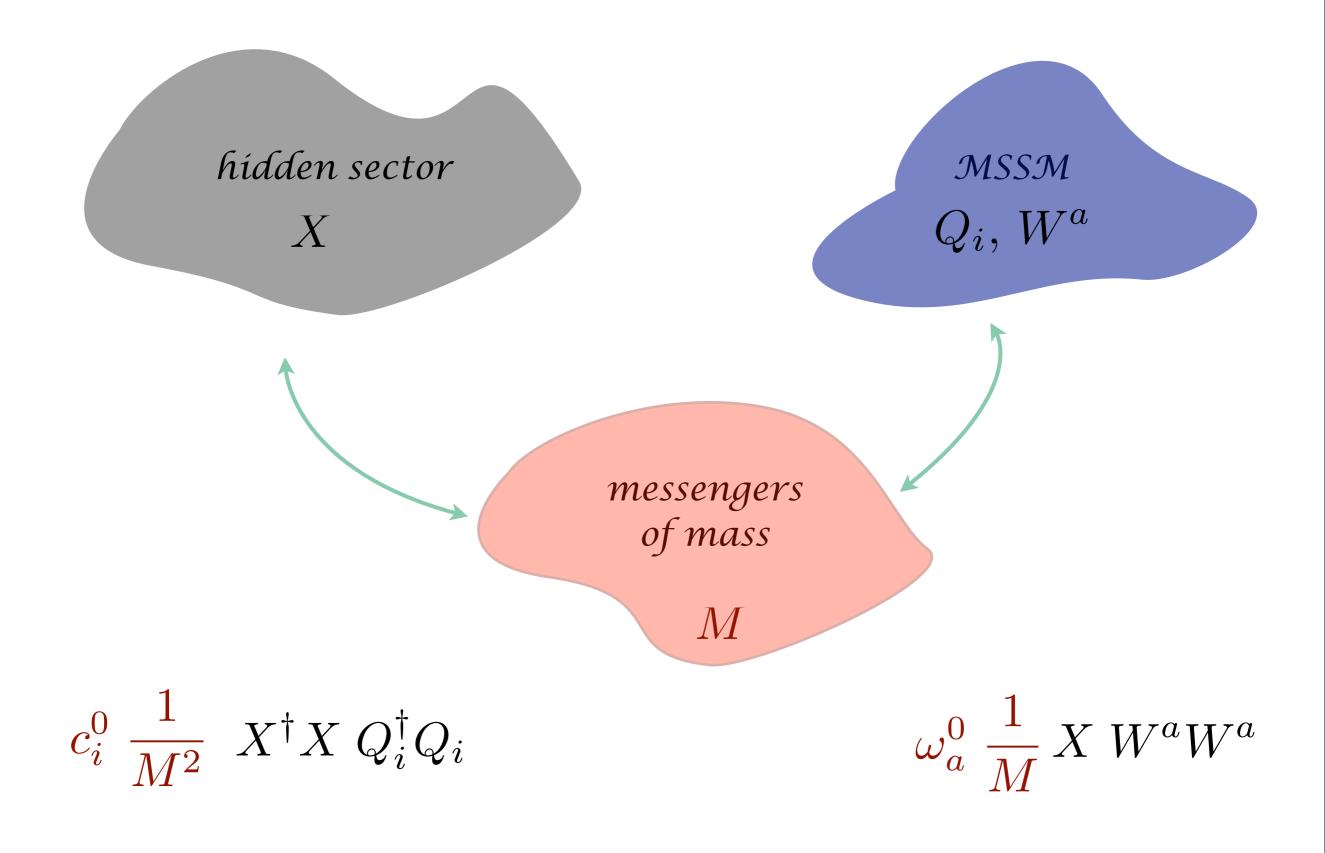
soft masses are also renormalized due to susy breaking interactions

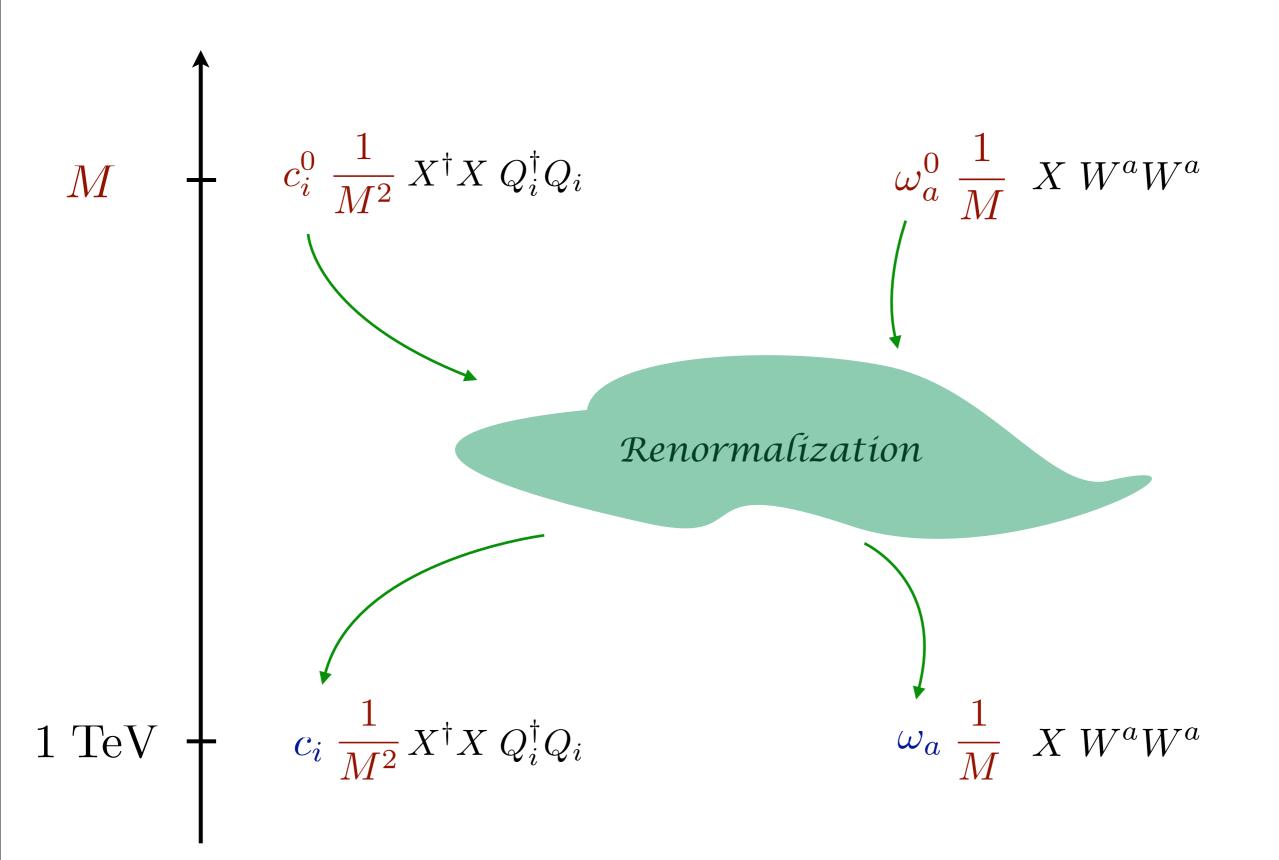
Cohen, TR, Schmaltz (hep/ph 0612100)

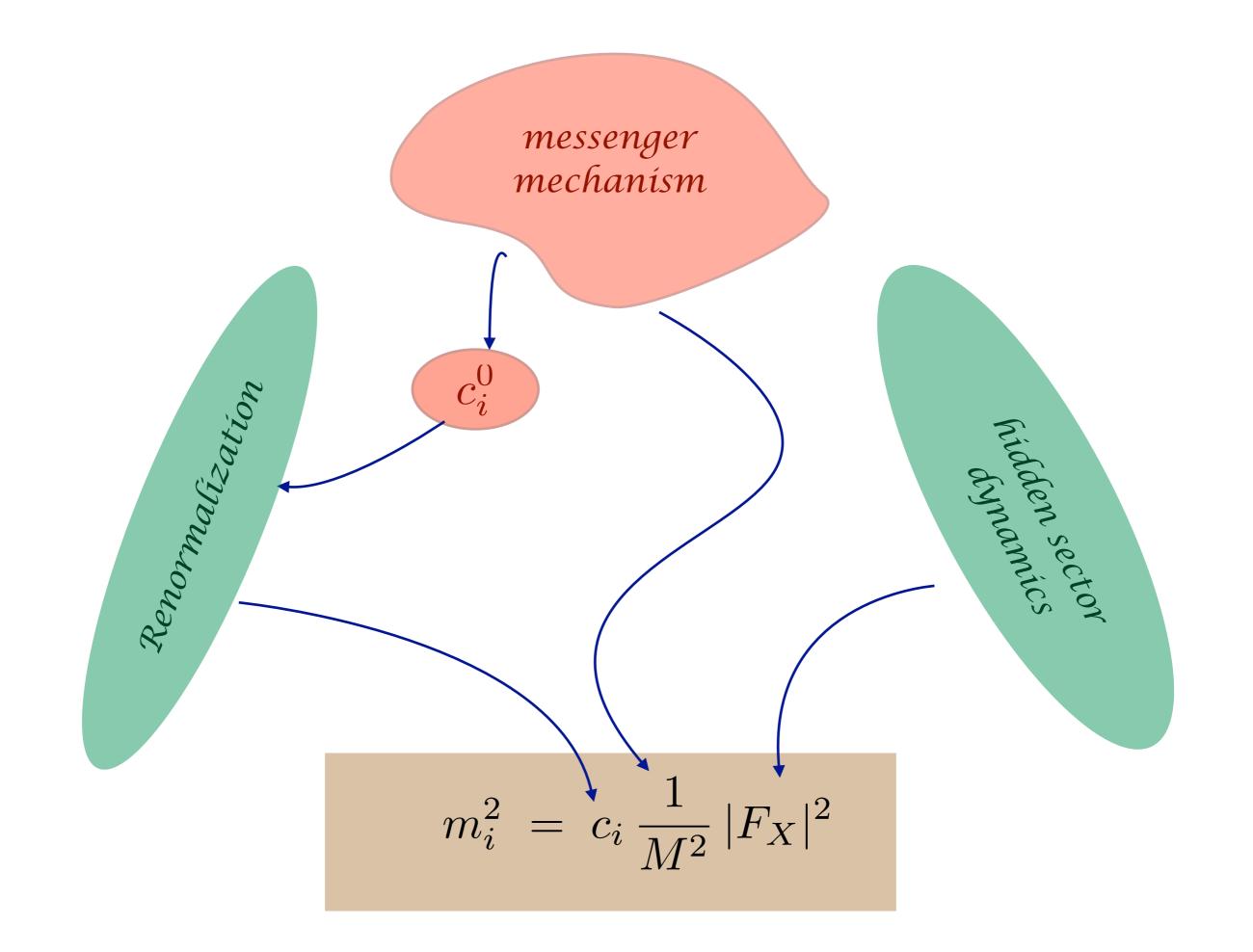
Outline

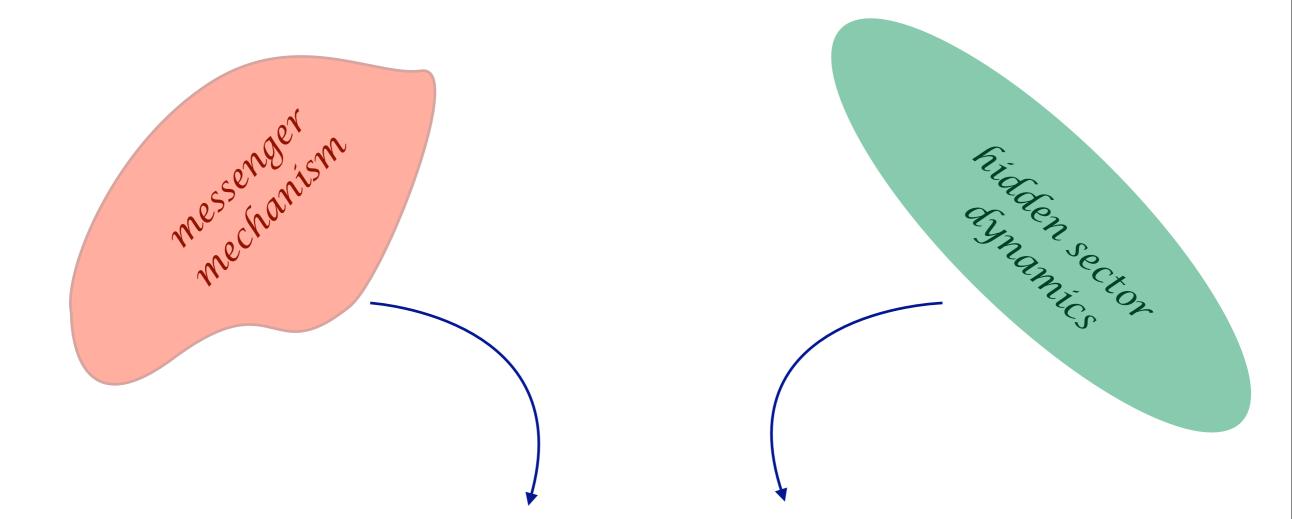
- 1. *Physics of soft masses and an intrinsic problem of all mediation mechanisms.*
- Corrections of soft masses due to hidden sector dynamics and a natural solution.
- 3. Scalar Sequestering
 - *definition*
 - *phenomenology and model independent predictions*

Physics of soft masses

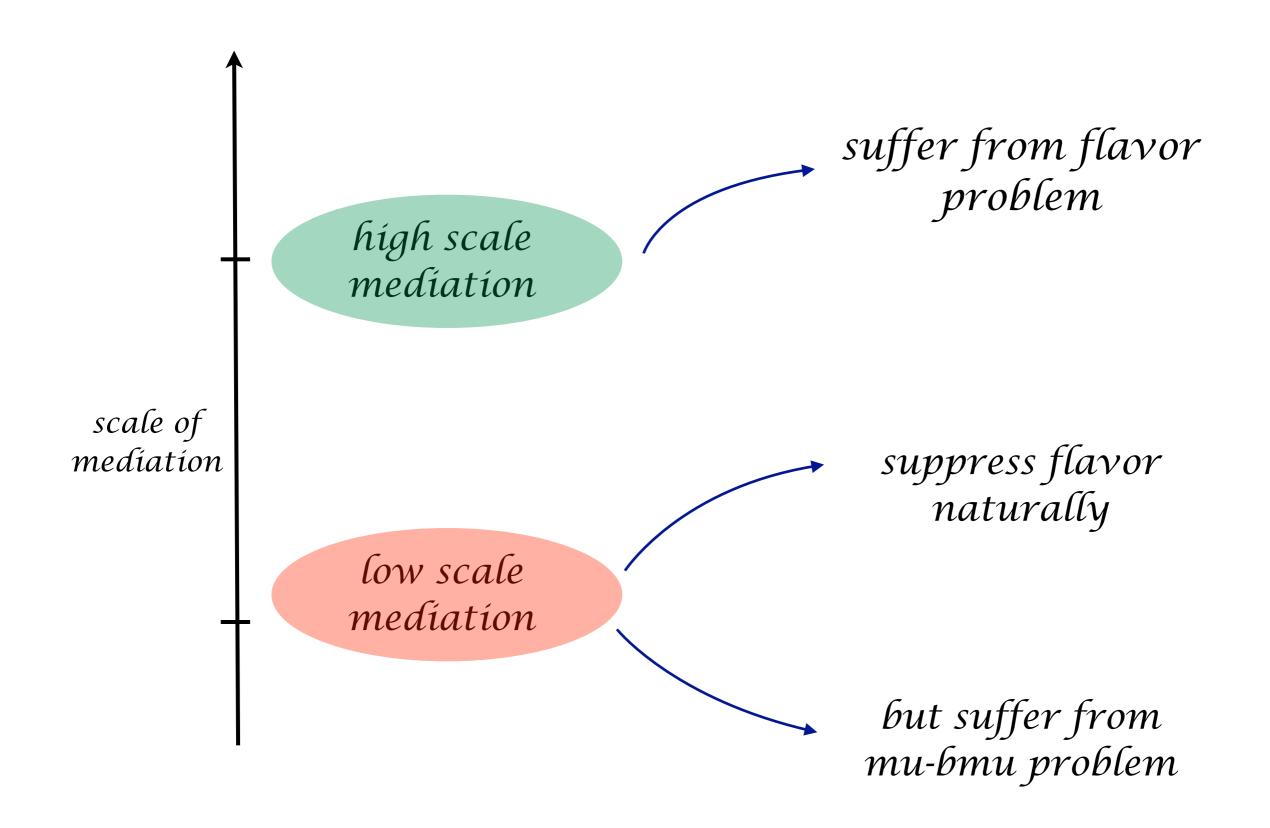




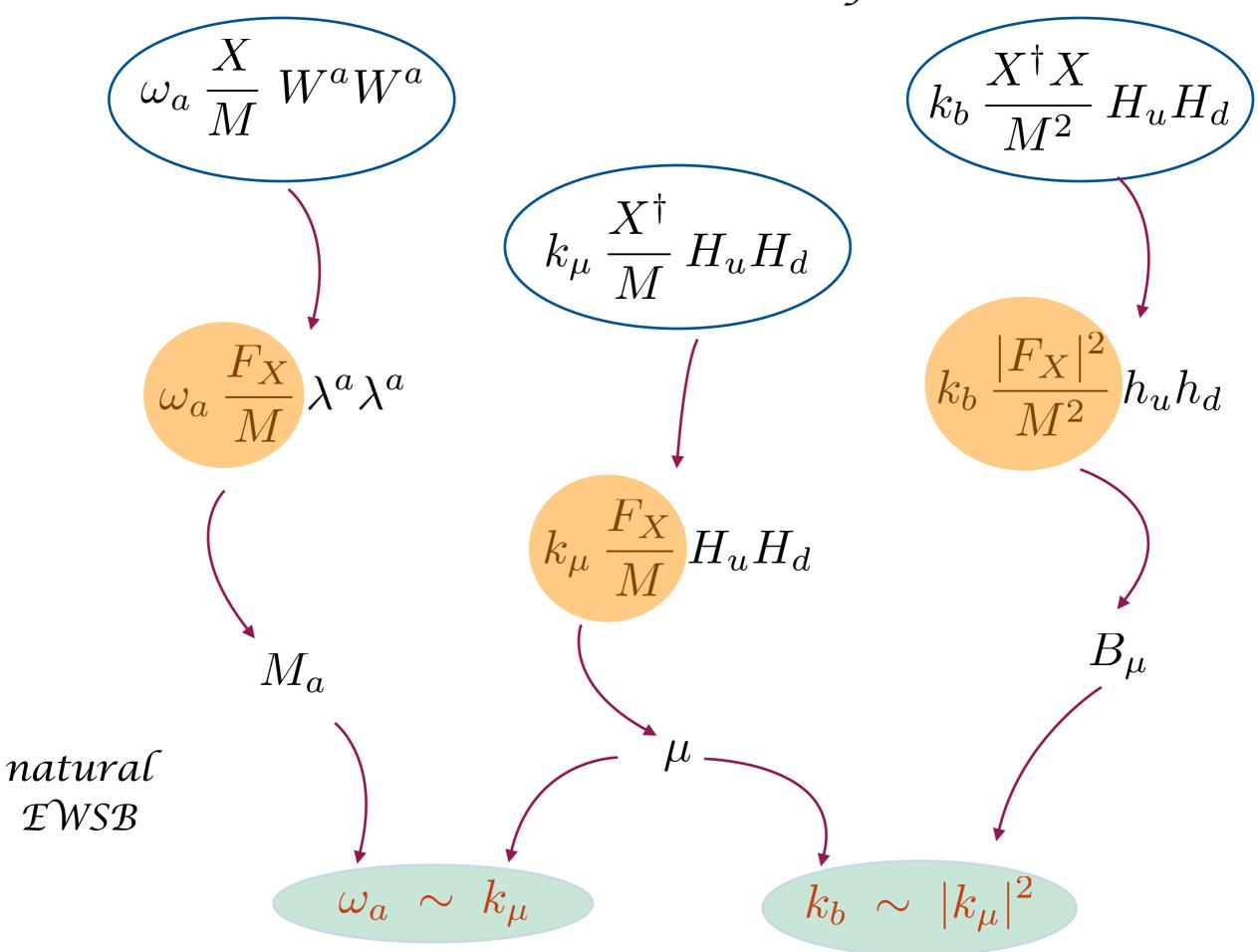




soft masses at EWSB scale summary of intrinsic problems of all messenger mechanism



more about the mu-bmu problem



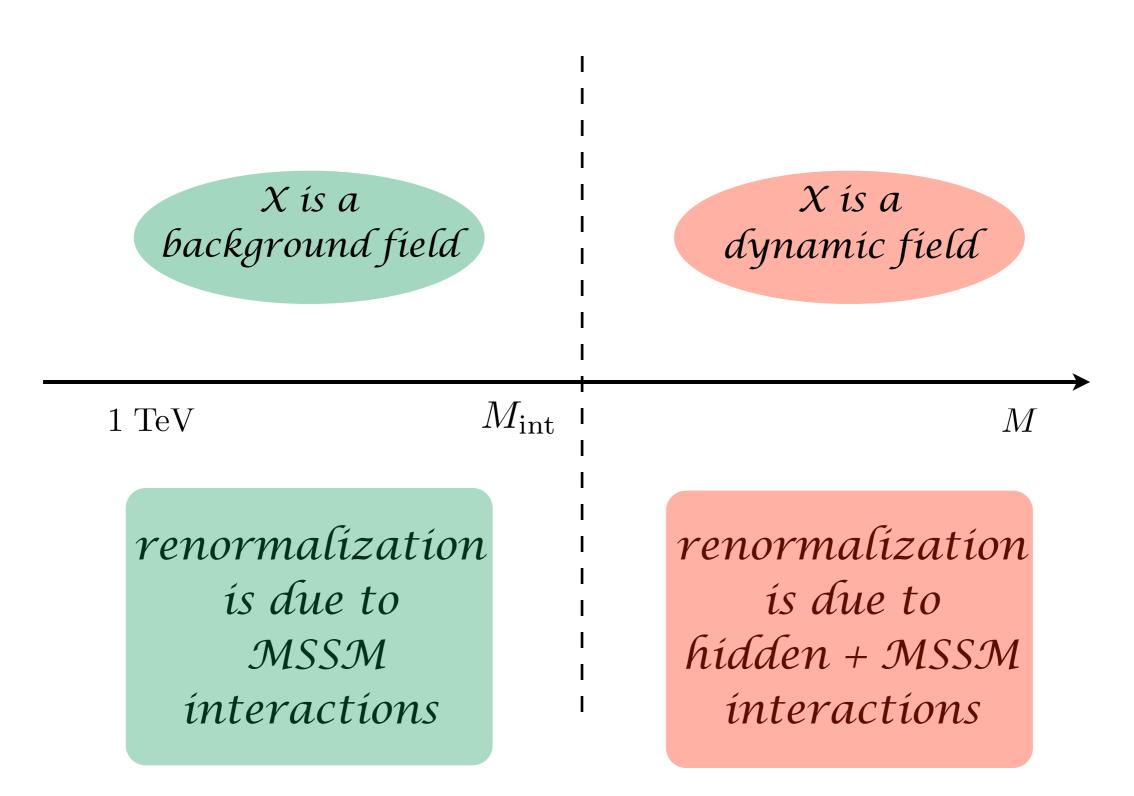
more about the mu-bmu problem

 $k_b \sim k_\mu \sim \delta$ $k_b \sim \frac{|k_\mu|^2}{\delta}$ $B_{\mu} \sim \frac{|\mu|^2}{s}$

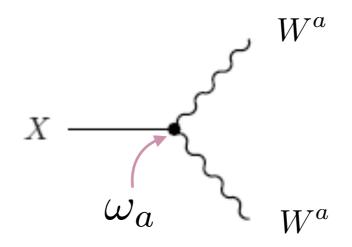


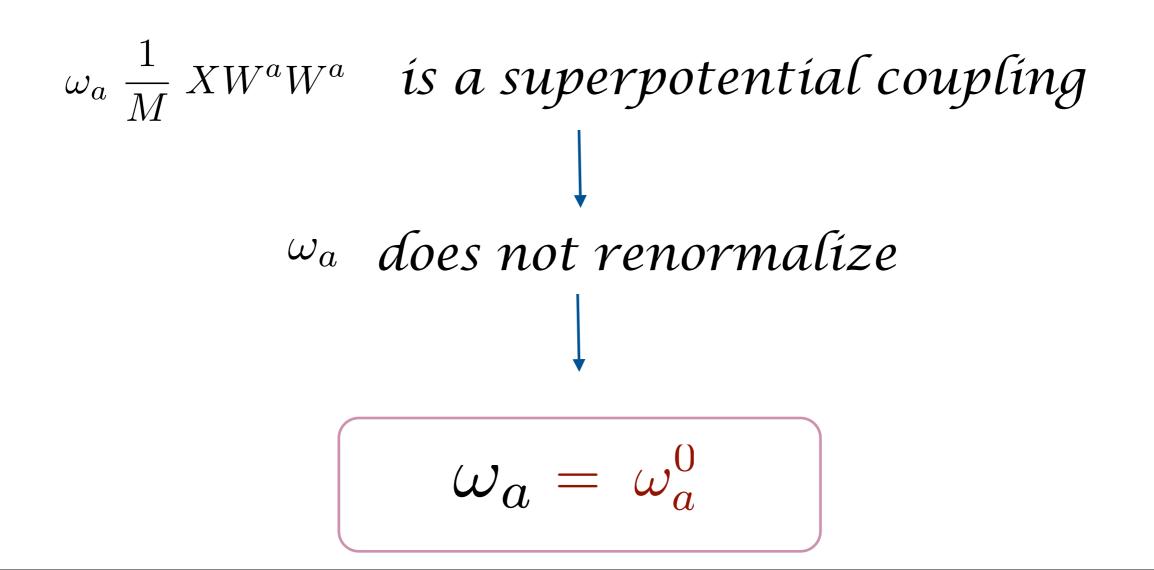
mu-bmu both can not have natural síze

Scales in renormalization

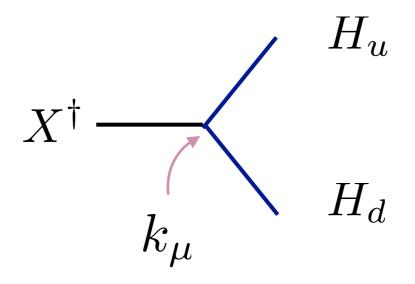


Renormalization

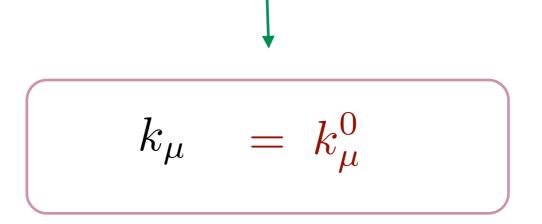


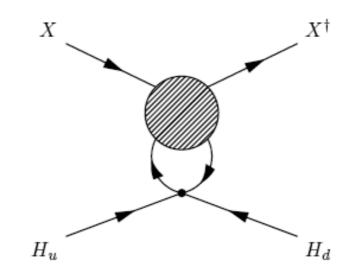


Renormalization



k_{μ} does not renormalize





$$\frac{dk_b}{dt} = \gamma k_b$$

$$k_b = e^{-\int \gamma} \quad k_b^0$$

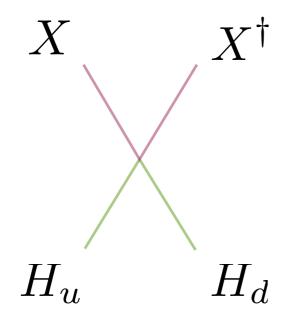
If γ is constant :

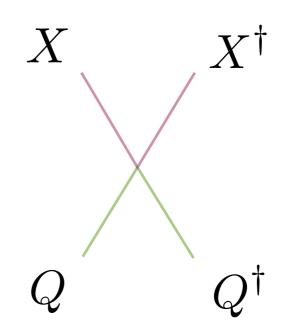
$$k_b(M_{\rm int}) = \left(\frac{M_{\rm int}}{M}\right)^{\gamma} k_b^0$$

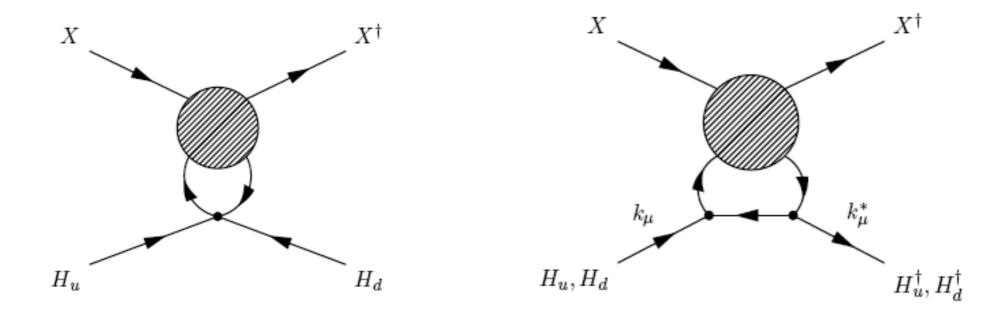
The hidden sector

- *1. supersymmetry is broken in the hidden sector.*
- 2. hídden sector dynamícs ís due to a superconformal field theory.
- 3. SCFT is such that $\gamma > 0$

hídden sector renormalization of scalar mass operators and the bmu operator is ídentical







$$\frac{dC_H}{dt} = \gamma C_H + \gamma |k_\mu|^2$$

$$C_H^{\text{fixed}} = -|k_\mu|^2$$

$$m_H^2(M_{\rm int}) = -|\mu|^2$$

Spectrum at the intermediate scale

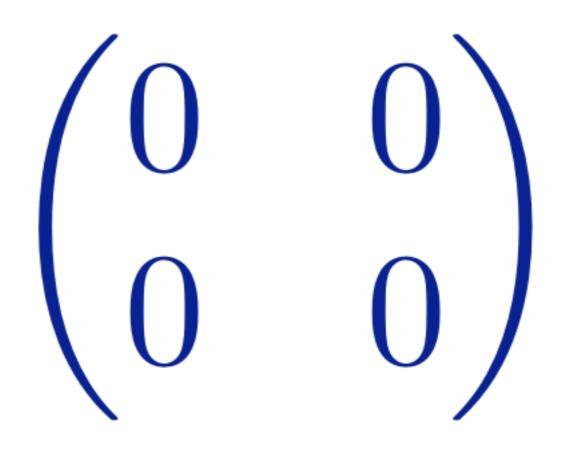
$$\mu \sim M_a$$

 $B_\mu = 0$
 $m_i^2 = 0$
 $m_{H_u}^2 = m_{H_d}^2 = -|\mu|^2$

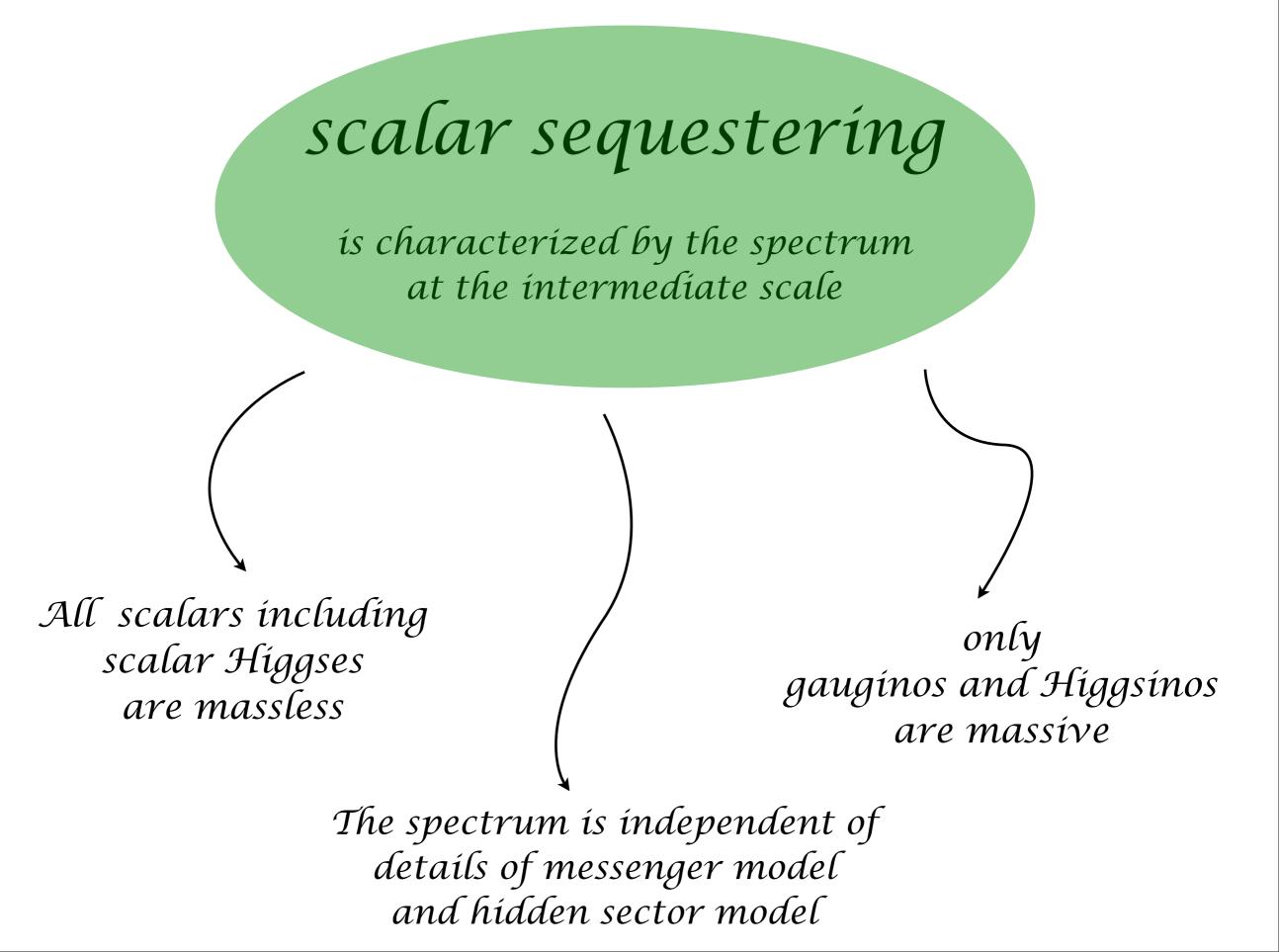
in particular: Higgs mass matrix

 $\begin{pmatrix} m_{H_u}^2 + |\mu|^2 & B_{\mu}^* \\ B_{\mu} & m_{H_d}^2 + |\mu|^2 \end{pmatrix}$

in particular: Higgs mass matrix



Scalar sequestering



Parameters of scalar sequestering

 μ

 $M_1 M_2 M_3$

 $a_t a_b a_{\tau}$

 $\log \frac{M_{\rm int}}{M_{\rm susy}}$

 $\tan\beta$

Parameters of scalar sequestering

 $\frac{\mu}{\frac{M_1}{g_1^2}} = \frac{M_2}{\frac{g_2^2}{g_2^2}} = \frac{M_3}{\frac{g_3^2}{g_3^2}} \equiv M_u$

 $a_t a_b a_{\tau}$

$$\log rac{M_{
m int}}{M_{
m susy}}$$

 $\tan\beta$

Parameters of scalar sequestering

 μ $\frac{M_1}{q_1^2} = \frac{M_2}{q_2^2} = \frac{M_3}{q_2^2} \equiv M_u$ $\frac{a_t}{\lambda_t} = \frac{a_b}{\lambda_b} = \frac{a_\tau}{\lambda_\tau} \equiv A_u$ $\log \frac{M_{\rm int}}{M_{\rm susy}}$

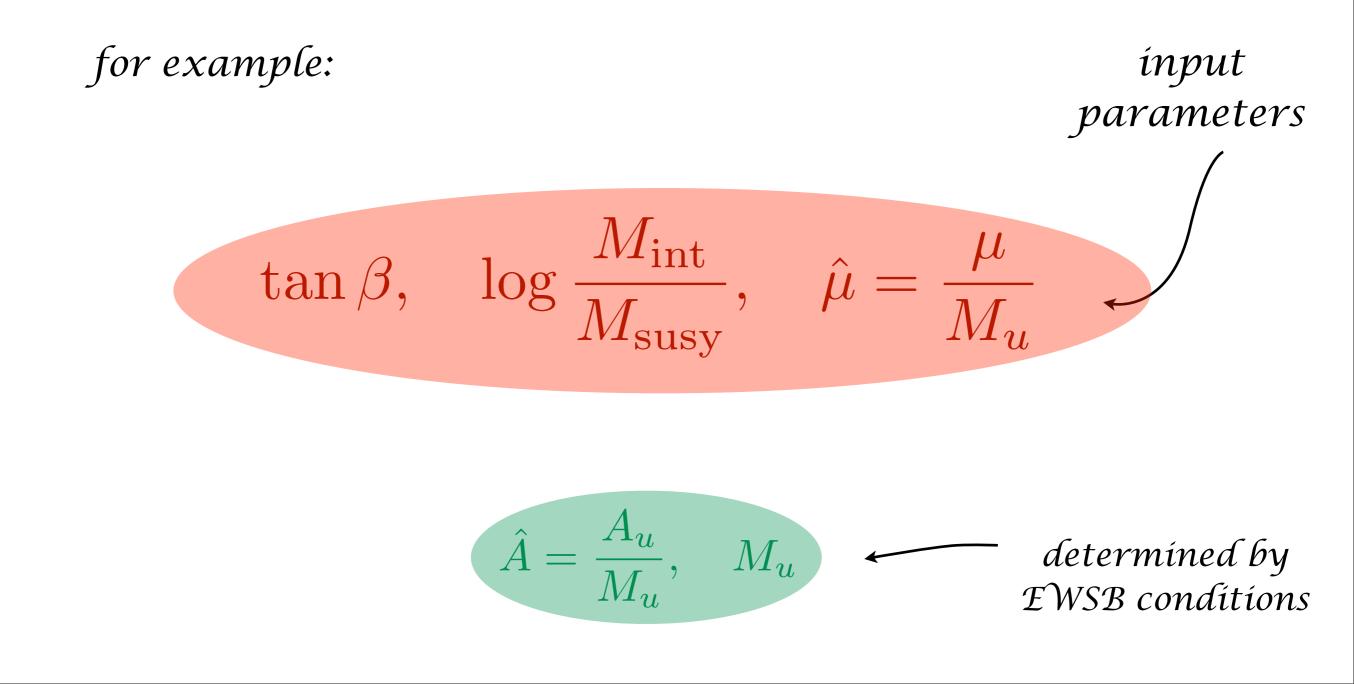
 $\tan\beta$

two equations of constraints at the EWSB scale

$$\frac{|m_{H_u}^2 - m_{H_d}^2|}{\sqrt{1 - \sin^2 \beta}} - \left(m_{H_u}^2 + m_{H_d}^2 + 2|\mu|^2\right) = m_Z^2$$
$$B_\mu - \frac{\tan \beta}{1 + \tan^2 \beta} \left(m_{H_u}^2 + m_{H_d}^2 + 2|\mu|^2\right) = 0$$

we can eliminate two more parameters

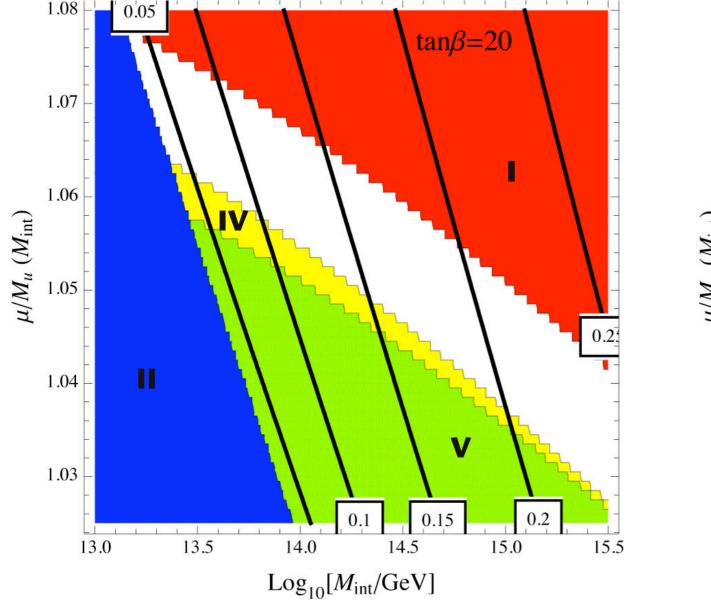
Scalar sequestering + unification spectrum is determined only in terms of 3 parameters

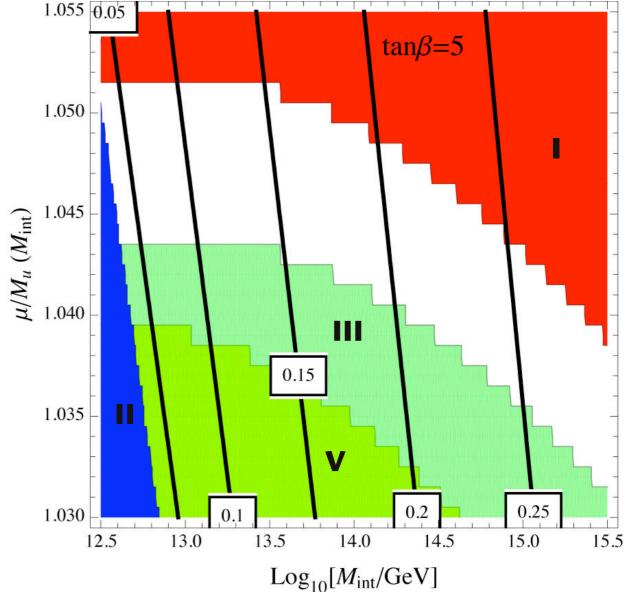


the parameter space in scalar sequestering + unification is the most minimal

it is more minimal than even MSUGRA and actually have theoretical foundation

there are regions of parameter space open for scalar sequestering + unification





features of scalar sequestering at the EWSB scale

- all scalar Higgses are much lighter that the Higgsinos $m_A^2 \ll |\mu|^2$
- third generation scalars receive additional positive contribution due to large and negative Higgs soft masses.

We predict:

 $\left(m_{\tau_1}^2 + m_{\tau_2}^2\right) > \left(m_{L_i}^2 + m_{E_i}^2\right) \qquad i \in \{1, 2\}$

other features of the spectrum are shared with gaugino mediation

Conclusion

- Soft masses are sensitive to hidden sector physics.
- Scalar sequestering is an extreme example where the effect due to the hidden sector dominates those due to the visible sector.
 - Scalar Sequestering solves non-trivial problems in susy model building using hidden sector dynamics
 - the spectrum is independent of details of messengers or hidden sector interactions
 - with further assumption of unification, the entire spectrum is determined in terms of only three parameters.
 - unique predictions
 - Higgsinos are far heavier than all scalar Higgses
 - *staus are generically heavier than smuons and selectrons*