



## The NOvA Experiment Status

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Fermilab Users' Meeting 2009



### The NOvA Collaboration





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## The NOvA Experiment



- NuMI Off-Axis  $v_e$  Appearance Experiment
- Uses a 15kT detector, NOvA
- Positioned off axis of the Fermilab NuMI beam
- Run for 6 years
- Goals:
  - Observation of the transition  $\nu_{\mu} \rightarrow \nu_{e}$
  - Measurement of  $\theta_{13}$
  - Measurement of  $\sin^2(2\theta_{23})$
  - Determination of mass hierarchy (sign of  $\Delta m_{23}^2$ )
  - $\ \, \text{Begin to localize } \delta_{CP} \\ \text{Leon Mualem } -\text{FNAL Users' Meeting 2009}$



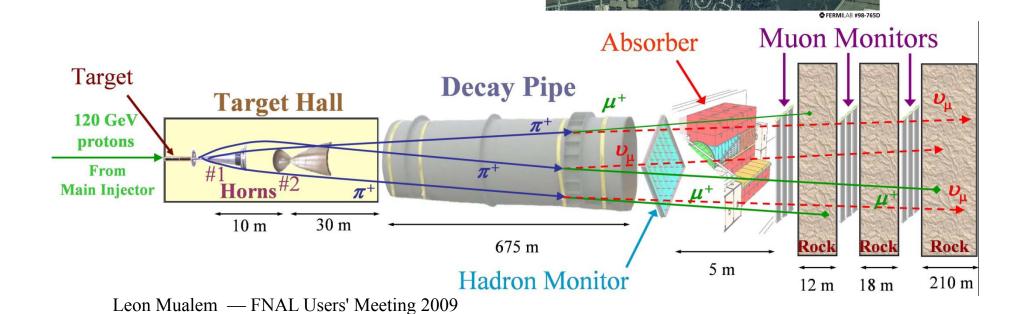
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## Fermilab NuMI Beam

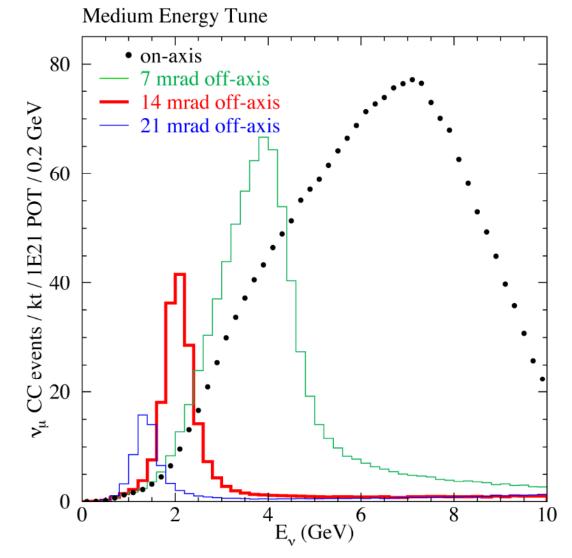




## Off-Axis beam features



- Narrow energy distribution
- Tuned Energy by selecting off-axis angle
  - Maximize v<sub>e</sub> appearance
  - Minimize  $\nu_{\mu}$
- Higher Intensity at a given energy
- Suppressed high energy tail
  - Reduces NC contamination





## Proton Plans



	Present Operating Conditions (May 2007)	Proton Plan Multi- batch Slip- stacking in MI	NO∨A Multi-batch Slip-stacking in Recycler
8 GeV Intensity (p/Batch)	4.3 - 4.5x10 <sup>12</sup>	4.3x10 <sup>12</sup>	4.3x10 <sup>12</sup>
Number of 8 GeV Batches to NuMI	7	11	12
MI Cycle Time (sec)	2.4	2.2	1.3
MI Intensity (protons per pulse or ppp)	3.3x10 <sup>13</sup>	4.5x10 <sup>13</sup>	4.9x10 <sup>13</sup>
MI to NuMI (ppp)	2.45x10 <sup>13</sup>	3.7x10 <sup>13</sup>	4.9x10 <sup>13</sup>
NuMI Beam Power (kW)	192	320	700
Protons/year to NuMI	2x10 <sup>20</sup>	3x10 <sup>20</sup>	6x10 <sup>20</sup>
MI Protons/hour	4.95x10 <sup>16</sup>	7.3x10 <sup>16</sup>	1.3x10 <sup>17</sup>

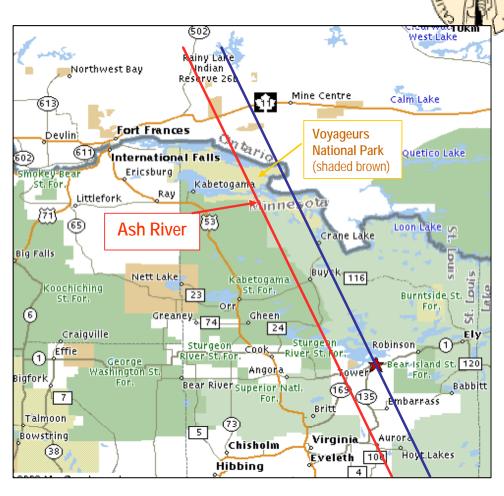


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## Site for Far Detector: Ash River, MN



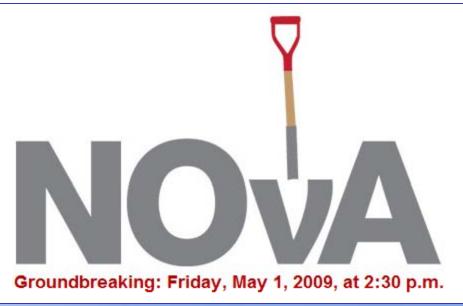


- This site is at 810 km from Fermilab, 11.77 km off-axis
- Farthest available site in the U.S. along the NuMI beam
  - US 53, St. Louis County Road 129 all weather access



## Far Site Status









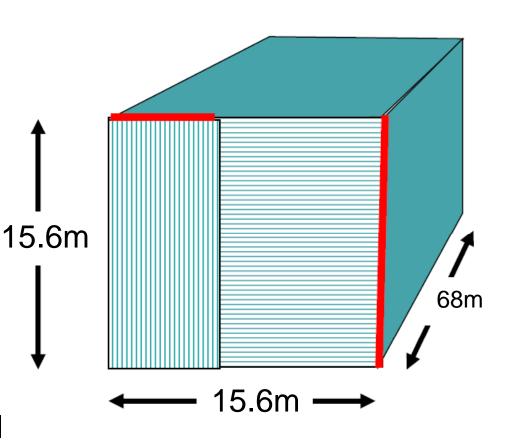




## NOvA Far Detector



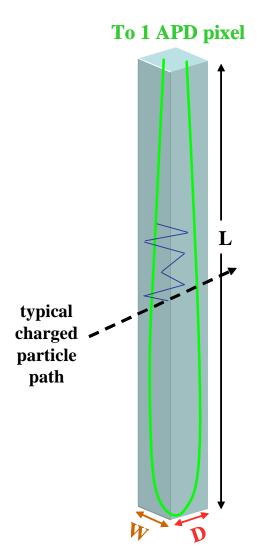
- > 15 ktons
- 15.6m x 15.6m x 68m
- 1003 liquid scintillator planes, (~73% active)
- Scintillator cells3.8 x 6.0 x 1540 cms
- Read out from one side per plane with APDs
- Expected average signal at far end of 30pe



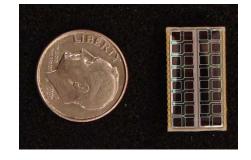


### The Basic Detector element





- Liquid Scintillator
  - 4.1% pseudocumene as scintillant
  - Mineral oil and waveshifters (PPO, bis-MSB)
- PVC cell for primary containment
  - Horizontals:3.87 cm x 6.0 cm x 15.4 m long
  - Verticals: 3.76 cm x 5.7 cm x 15.4 m long
  - Highly reflective, 15% titanium dioxide
- Looped wavelength shifting fiber to collect light
  - 0.7 mm diameter, double clad, K27 waveshifter
  - Almost perfect mirror, 3.6\*light in 1 fiber
- Avalanche Photodiode
  - 85% quantum efficiency
  - Gain of 100, operate at -15°C
- Low noise amplifier





### APD Photodetector

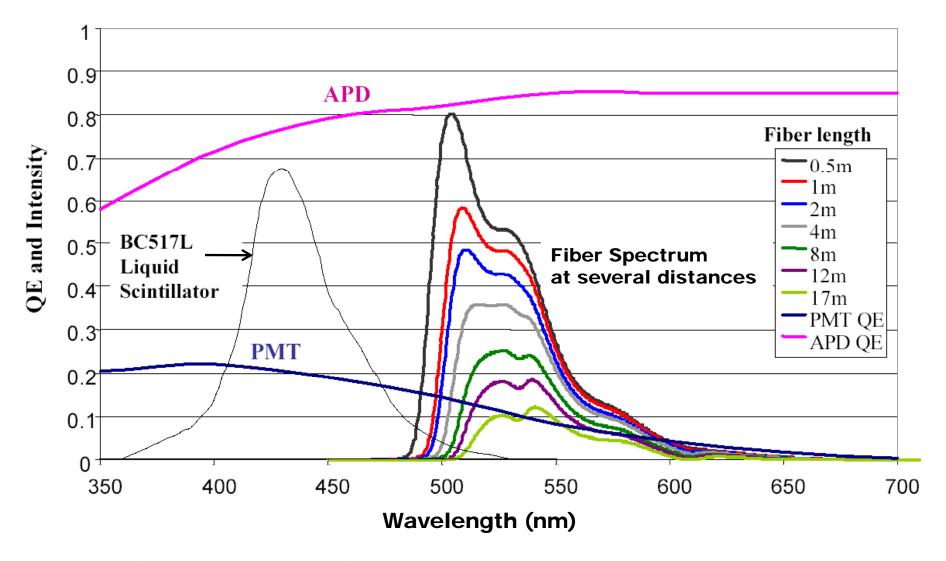
- Hamamatsu
   Si Avalanche
   Photodiode (APD)
  - S11211(X) CustomVariant of commercialS8550 SiAPD
  - Custom design to match to fiber aspect ratio
  - Bare die mounted toPCB via bump bonding





## APD Advantage



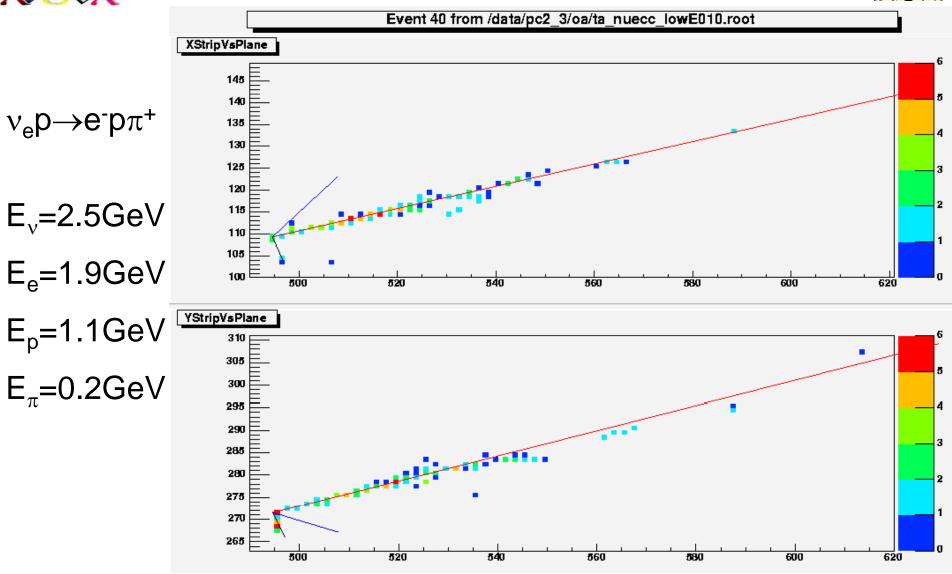


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## e CC event



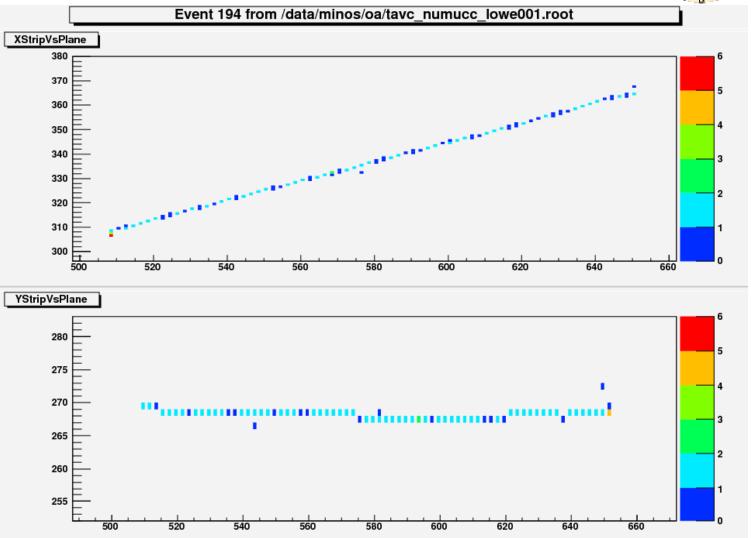




## μ CC event



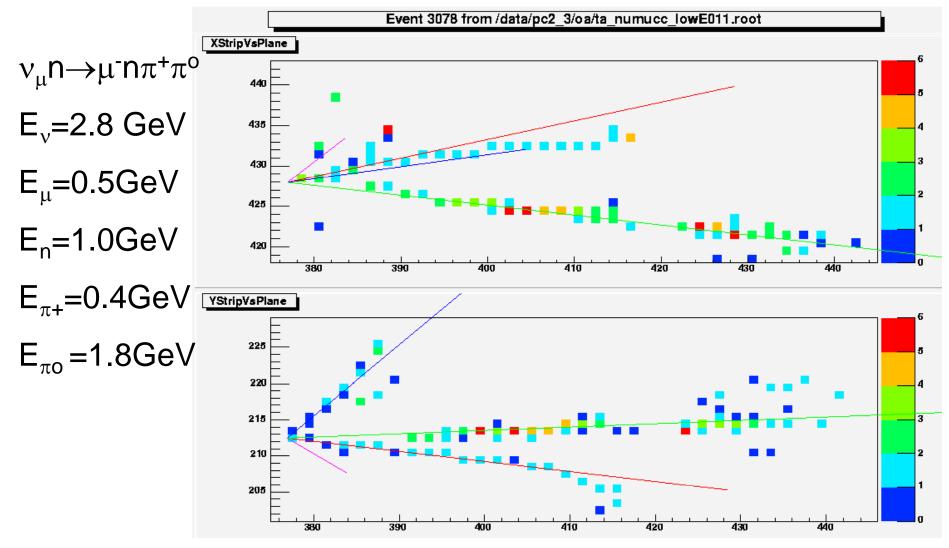
2GeV muon CC event





## μ CC event

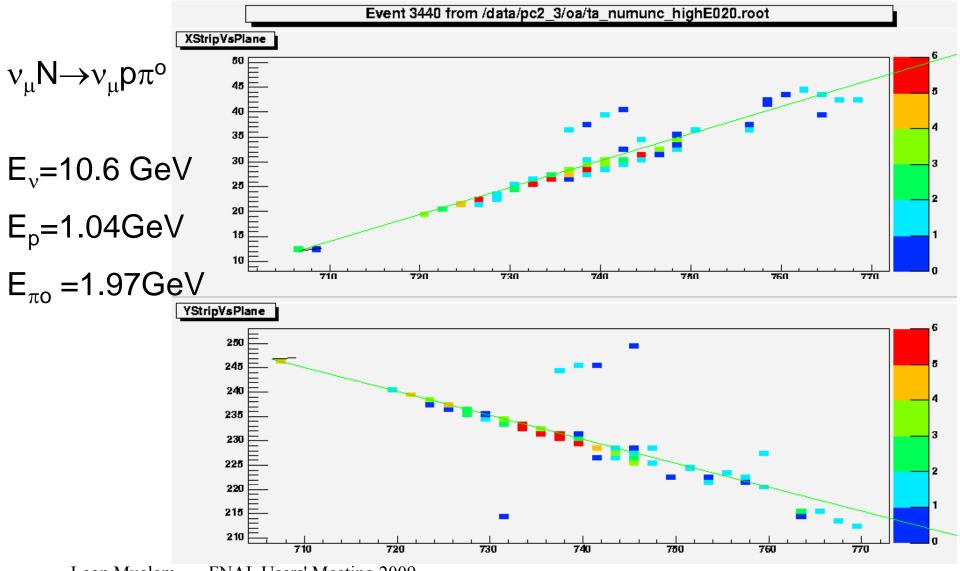






## NC event



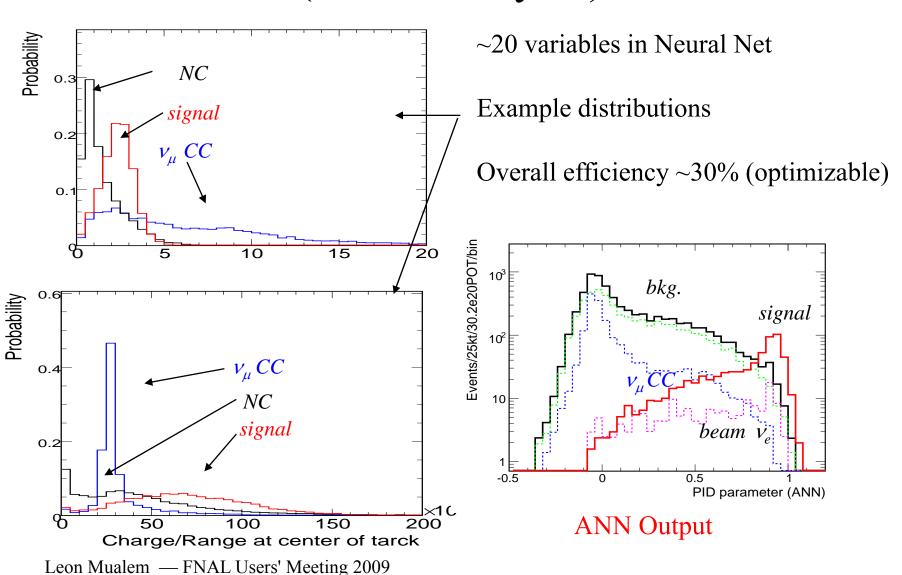


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# Electron Identification for NOvA (TDR Analysis)







## Detector Performance



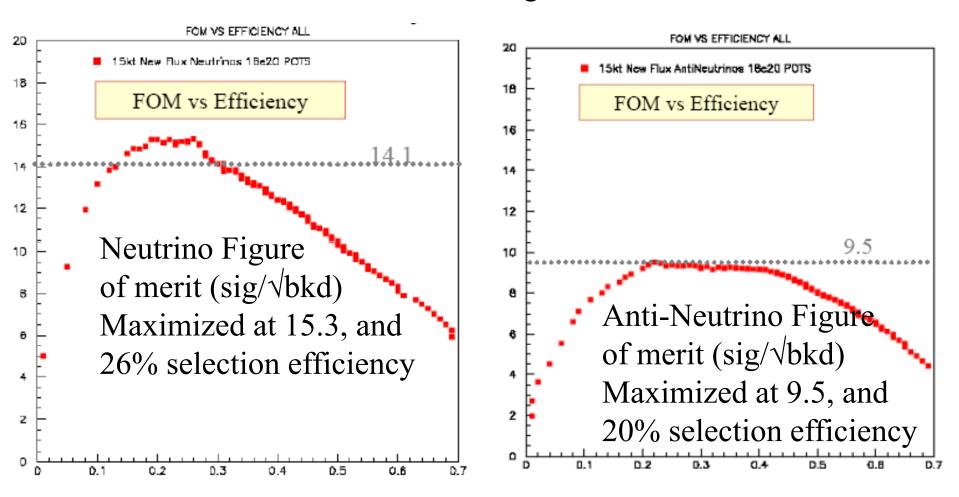
- Optimize Figure of Merit (FOM=sig/√bkd) performance separately for v and anti-v
- Use 18E20 POT,  $\sin^2(2\theta_{13}) = 0.1$ ,  $\Delta m^2 = 2.4E-3$ , no matter effects
- Upgrades to neutrino event generator
- Updates to predicted neutrino flux
- Updated detector mass and configuration
  - 15kT, slightly shorter modules, photodetector readout



## Performance optimization



### Combined Figure of merit 18.0

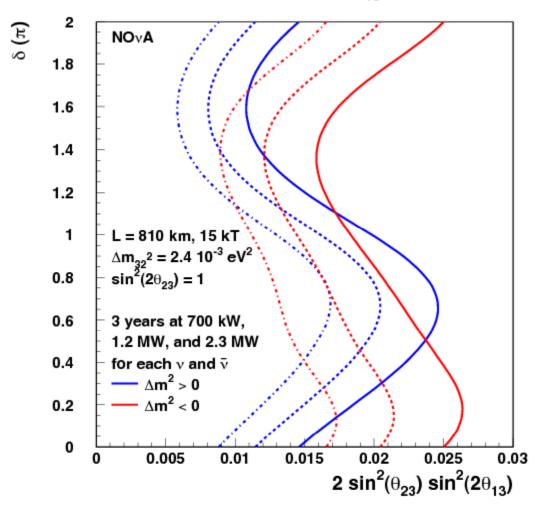




## 23 $\sigma$ Sensitivity to $\sin^2(2\theta_{13}) \neq 0$



### 3 $\sigma$ Sensitivity to $\sin^2(2\theta_{13}) \neq 0$



Sensitivity vs  $\delta$  for non-zero  $\sin^2(2\theta_{13})$  at 700kW, 1.2MW, and 2.3MW

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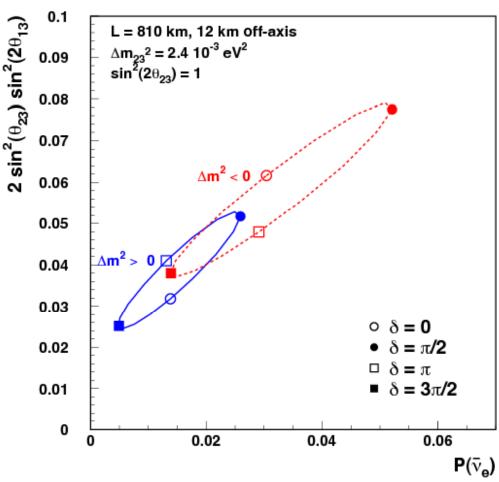
## Mass Hierarchy Resolution



If you assume a measurement of  $P(v_e)=0.02$ , what does that imply about  $\theta_{13}$  and the mass hierarchy, and how can you distinguish them.

Determine P using antineutrinos and/or measure  $\theta_{13}$  in reactor experiment.

### $\sin^2(2\theta_{13})$ vs. $P(\bar{v}_e)$ for $P(v_e) = 0.02$

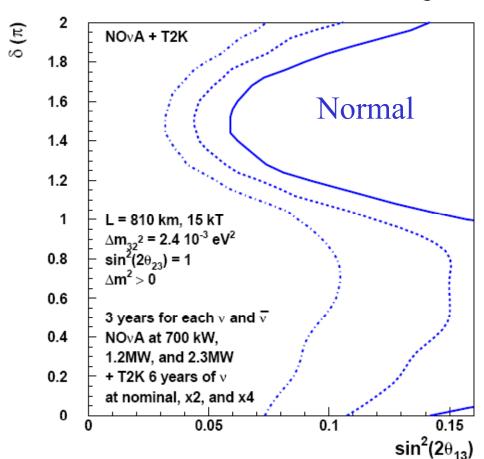




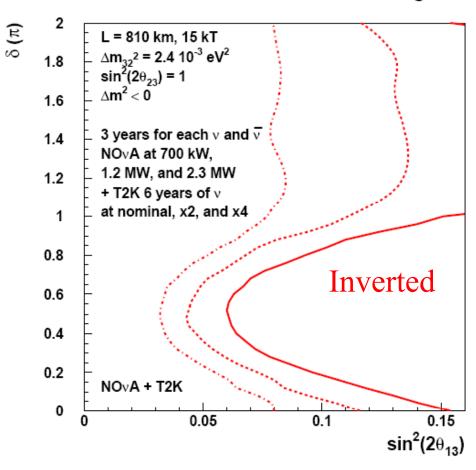
# 95% CL Sensitivity to the Mass Ordering



### 95% CL Resolution of the Mass Ordering



#### 95% CL Resolution of the Mass Ordering

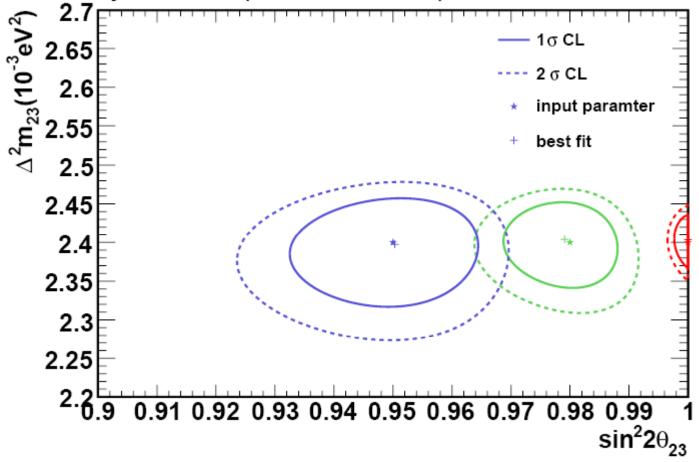




## Measurement of $\sin^2(2\theta_{23})$







If  $sin^2(2\theta_{23}) = 1$ , then it can be measured to 0.004.

Otherwise, it can be measured to ~0.02.

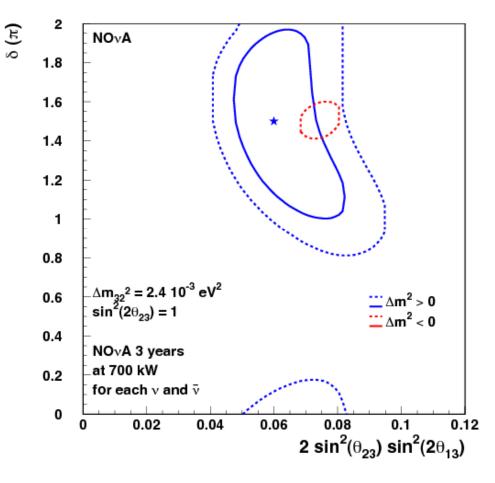


## CP phase



 Assuming a normal hierarchy, and oscillation at the starred point, NOvA measurement resolves the mass hierarchy, and constrains  $\delta_{CP}$  to the top half of the plane.

#### 1 and 2 or Contours for Starred Point for NOvA





## Progress



- Advances on all fronts
  - Accelerator and NuMI Upgrades (ANU) extensive shutdown plans beginning upgrades necessary for 700kW operation.
  - Module components prepared, or contracts made for prototype detector. (Integration Prototype Near Detector (IPND), and more, where approved)
  - Assembly prototypes under construction
    - Static leak tests, full height, in C0, full pressure at UMN
    - FSAP, Full Size Assembly Prototype (See collaboration photo "Dance Floor")
    - FHEP, Full Height Engineering Prototype, 31 planes 2 modules wide



## Module Assembly and Testing







**Bubbler Construction** 

Testing Bubblers

Extrusion production



Rolling Lifting Fixture



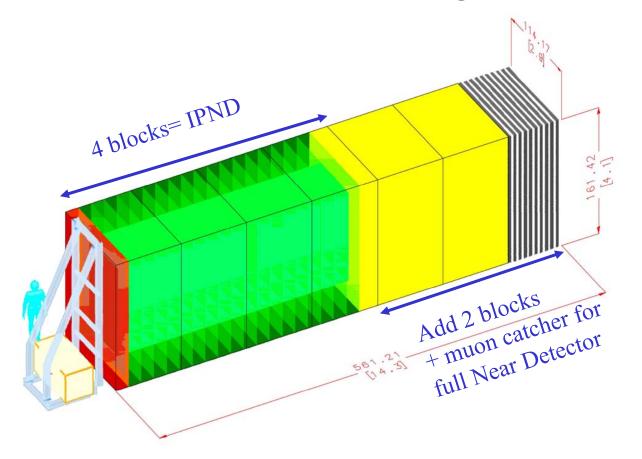
Adhesive Dispenser



## Integration Prototype Near Detector



- Constructing prototype detector to run on the surface near MINOS Surface Building
- 4 Blocks (6+)
- 2.9m wide
- 4.2m tall
- 8.4m long (15.3m)





## Approximate Schedule



- May 2007: US DOE CD-1 Approved
- Sep 2007: Cooperative Agreement signed
- Nov 2007: US DOE CD-2/3A review
  - Baseline, advanced procurement
- Dec 17, 2007 R&D funding for FY08 cut to zero.
- July 1, 2008 Supplementary funding restarts progress
- Sept 15, 2008, Oct 24, 2008: CD-2, CD-3A Approved
- April 2009: Reinvestment Act, and FY09 funding approved
- May 1, 2009 Groundbreaking at Ash River
- Fall 2009 Begin Building IPND modules
- Winter 2009/10 Begin building IPND Blocks
- Summer 2010 Begin operation of IPND (prototype Near Det.)
- Summer 2010: Beneficial Occupancy of the Far Detector building
- Winter 2011: Begin to Install FD, Data taking after a few kT installed
- 2013 Installation Complete



## Conclusions



- The NOvA experiment exhibits significant sensitivity to the current and future issues of neutrino oscillations
  - Electron neutrino appearance --  $\sin^2(2\theta_{13})$
  - Muon neutrino disappearance --  $\sin^2(2\theta_{23})$
  - mass hierarchy resolution
  - CP violation measurement
- Upgraded beam intensity needed to reach all of these goals
- Final thanks to NOvA Collaborators for their contributions



## Extras





## Far Detector Neutrino Spectrum



(L=810 km, theta=14 mrad) 10<sup>2</sup> Signal at CHOOZ limit

