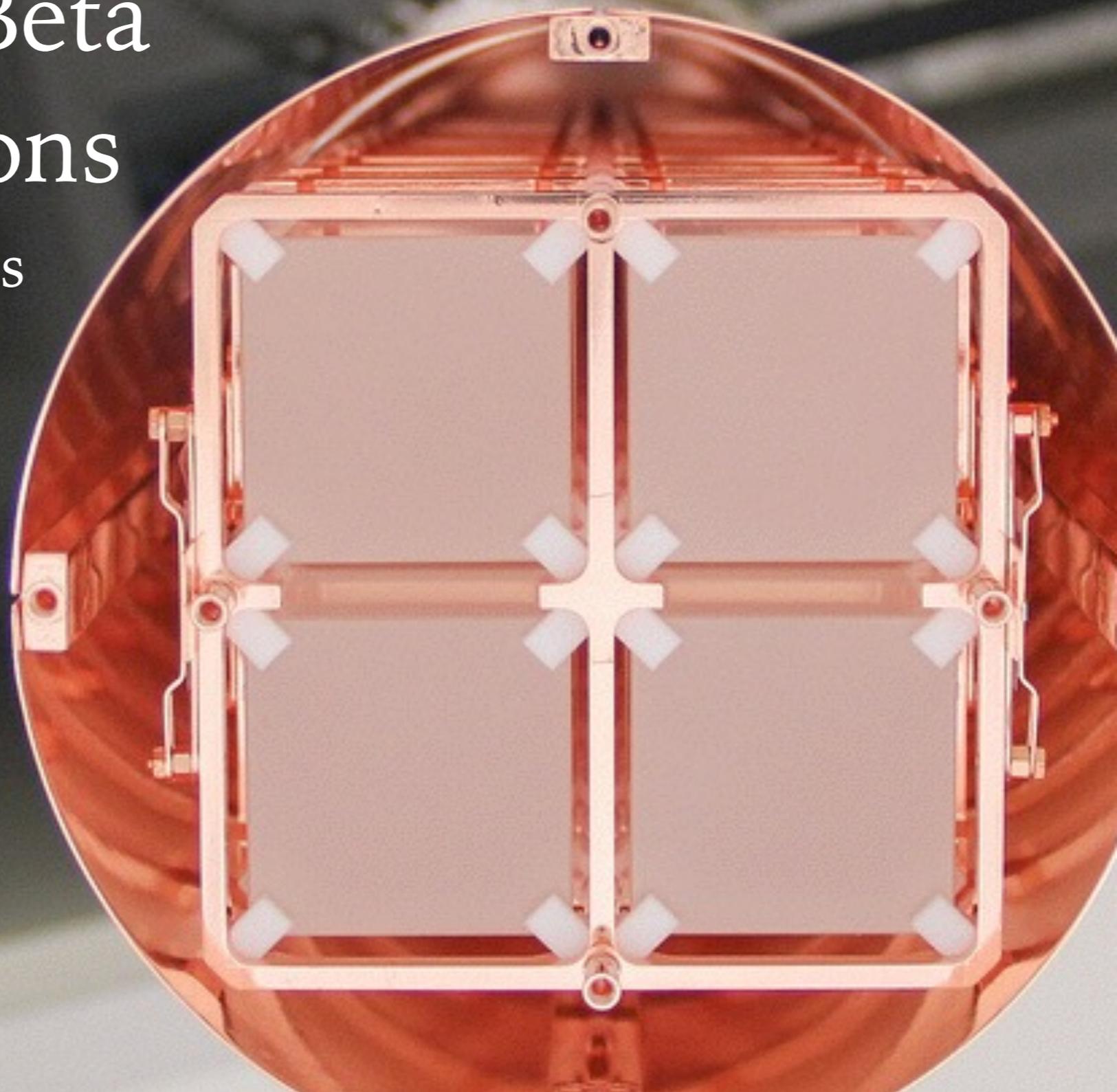


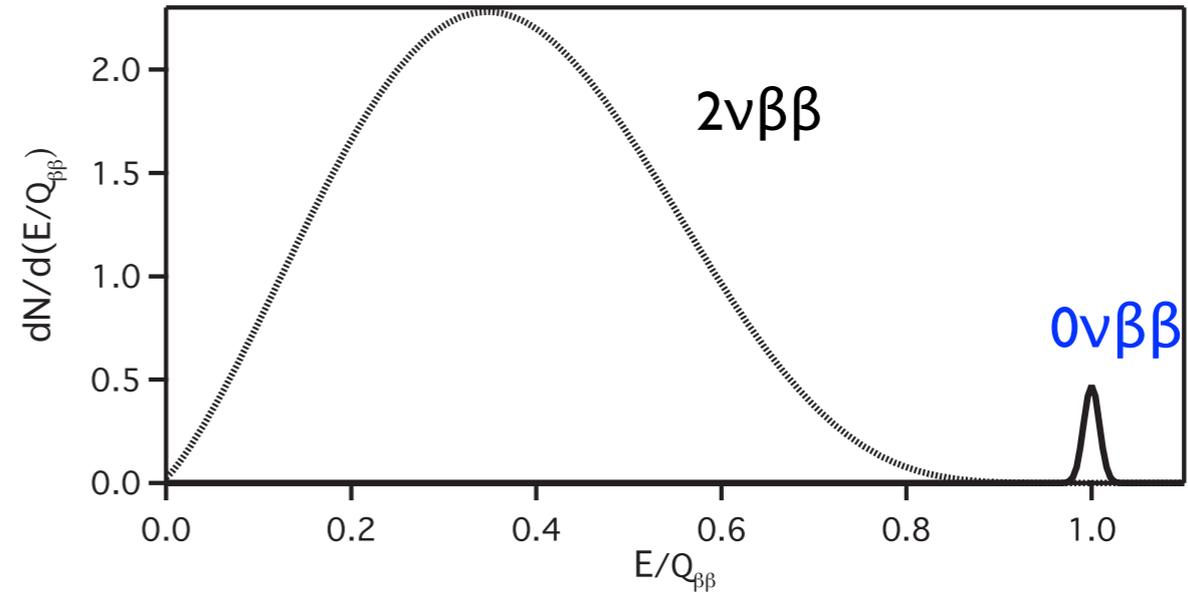
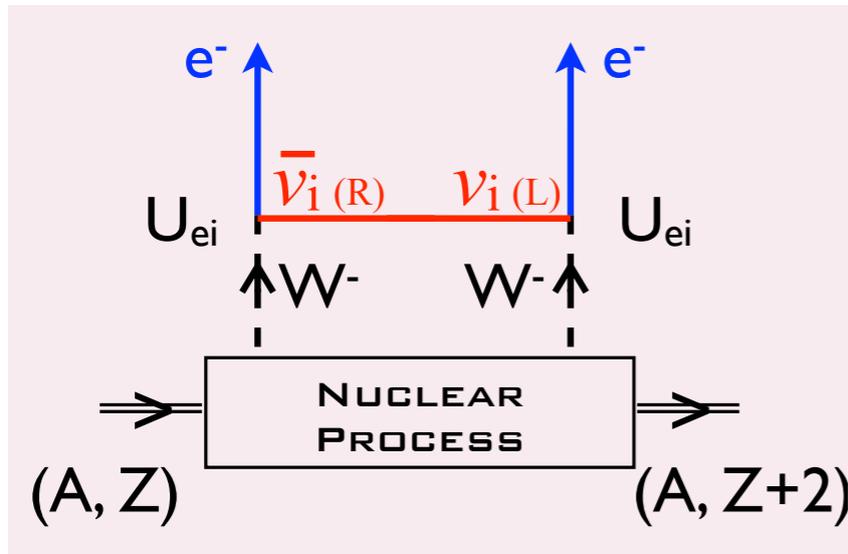
# TES for Double Beta Decay Applications

Raul Hennings Yeomans  
UC Berkeley



# Neutrino-less Double Beta Decay

- ▶ Hypothetical  $\beta\beta$  decay mode allowed if neutrinos are Majorana particles, i.e.  $\bar{\nu}_i \equiv \nu_i$



Phase space factor      Nuclear matrix element

$$\frac{1}{T_{1/2}^{0\nu}} = G^{0\nu} |M^{0\nu}|^2 |\langle m_{\beta\beta} \rangle|^2$$

Decay half-life      Effective Majorana  $\nu$  mass:

$$m_{\beta\beta} \equiv \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$$

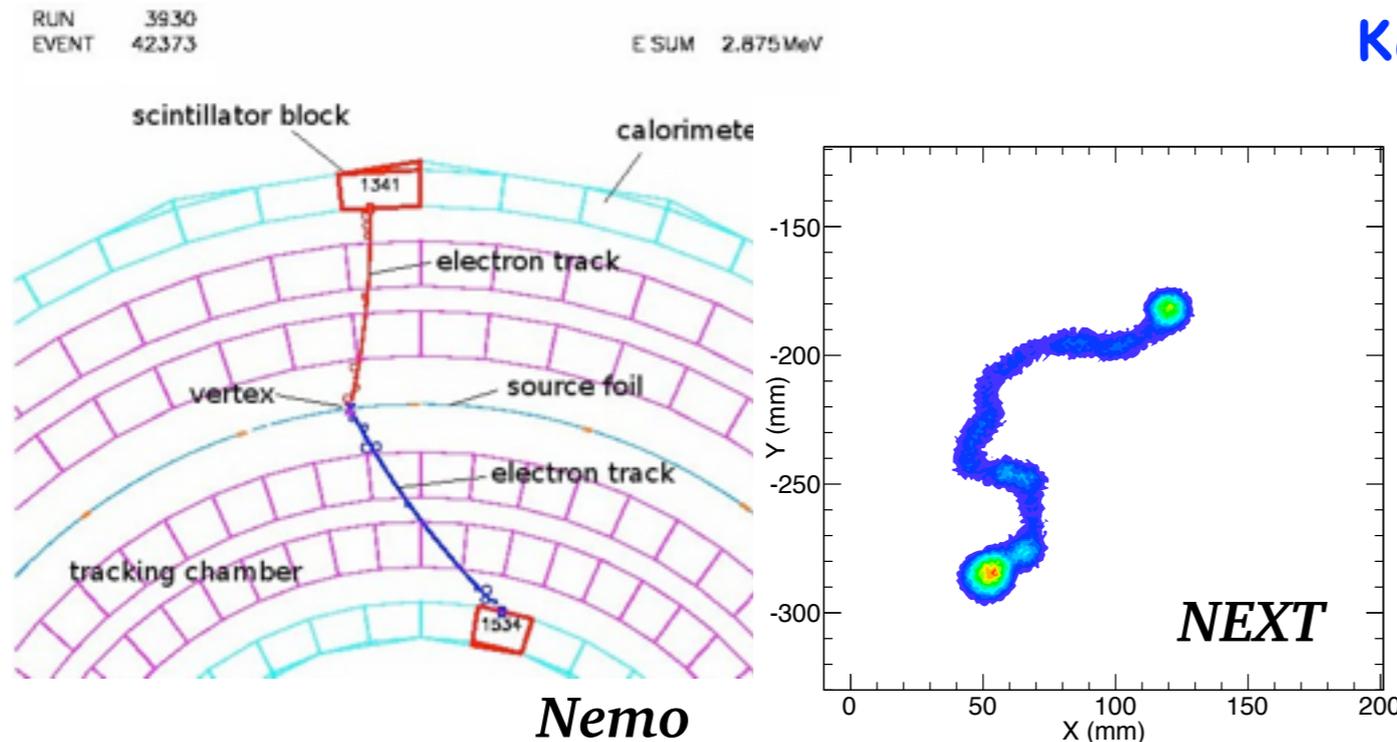
- ▶ For  $m_{\beta\beta} = 15$  meV estimated half lives  $10^{27} - 10^{28}$  years, depending on the nuclear system

- ▶ **Observation of  $0\nu\beta\beta$  would mean**
  - Lepton number violation
  - Neutrinos are Majorana particles
  - Rate measures (effective) electron neutrino mass

# Experimental approaches to $0\nu\beta\beta$

Source external to detector

Example: SuperNEMO



**Plus:** event topology, background rejection, multiple isotopes possible.

**Cons:** detector mass, resolution, acceptance.

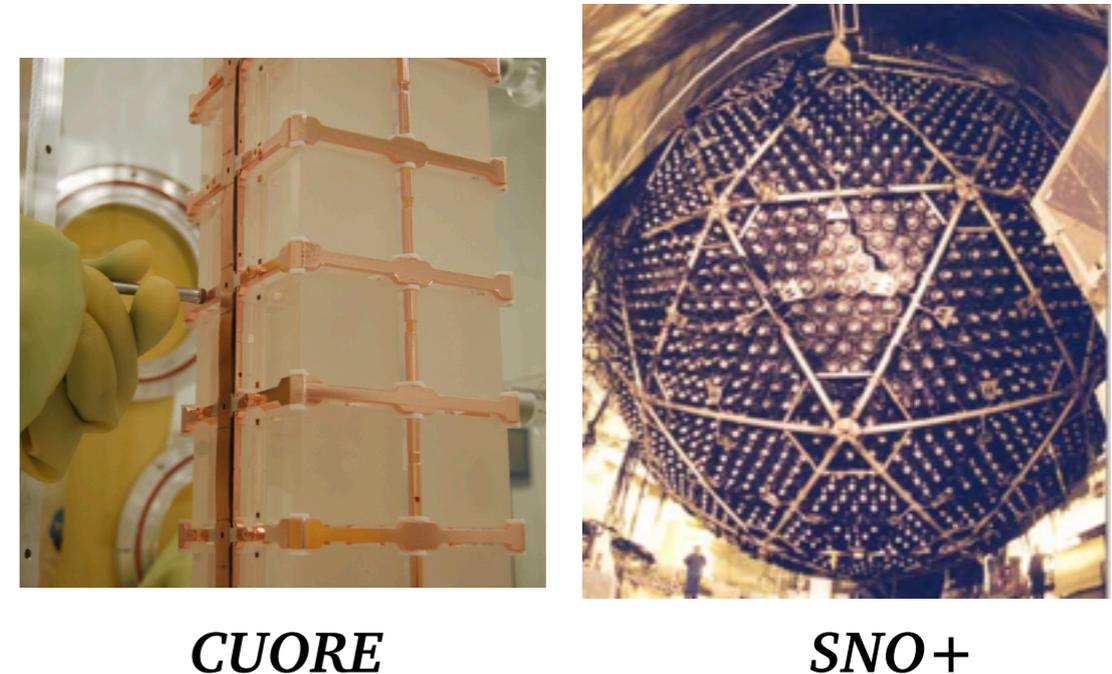
**Technology:** typically tracking detectors.

**Technology:**  
Pressurized TPC  
(10-16 atm)

May prove invaluable to test models  
once  $0\nu\beta\beta$  is discovered

Source internal to detector

Example: MAJORANA, EXO, CUORE, SNO+, Kamland-Zen, etc.



**Plus:** detector mass, energy resolution, acceptance

**Cons:** event topology, background rejection

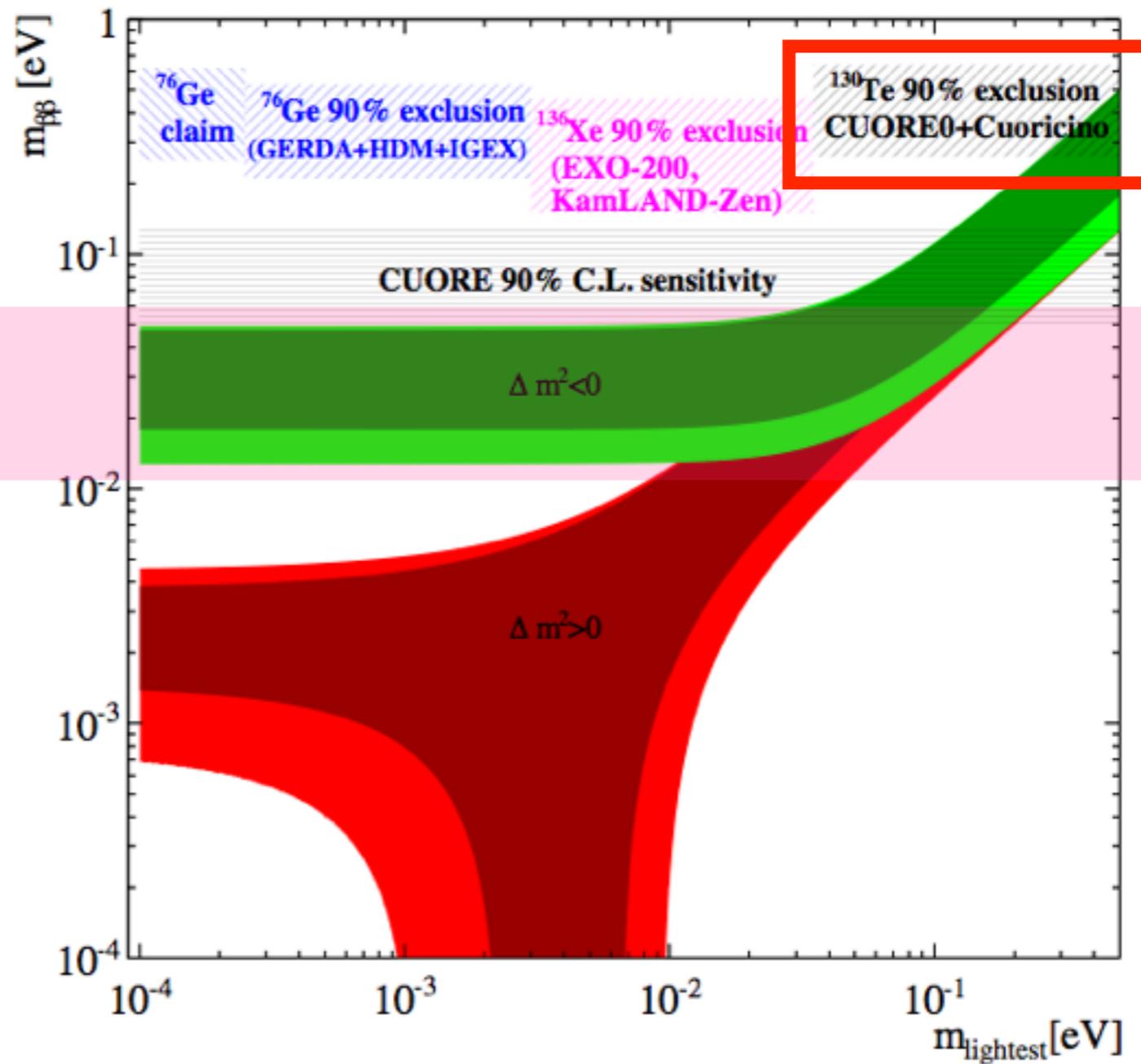
**Technology:** calorimeters (bolometers, ionization, scintillation), tracking

Typically aimed at  $0\nu\beta\beta$  discovery

# Beyond CUORE: towards covering the IHE

Goal of next generation experiments

$$m_{\beta\beta} \sim \frac{m_e}{\sqrt{F_N \cdot \varepsilon \cdot \eta} \sqrt{\frac{M \cdot t}{b \cdot \delta E}}}$$



For a zero background experiment

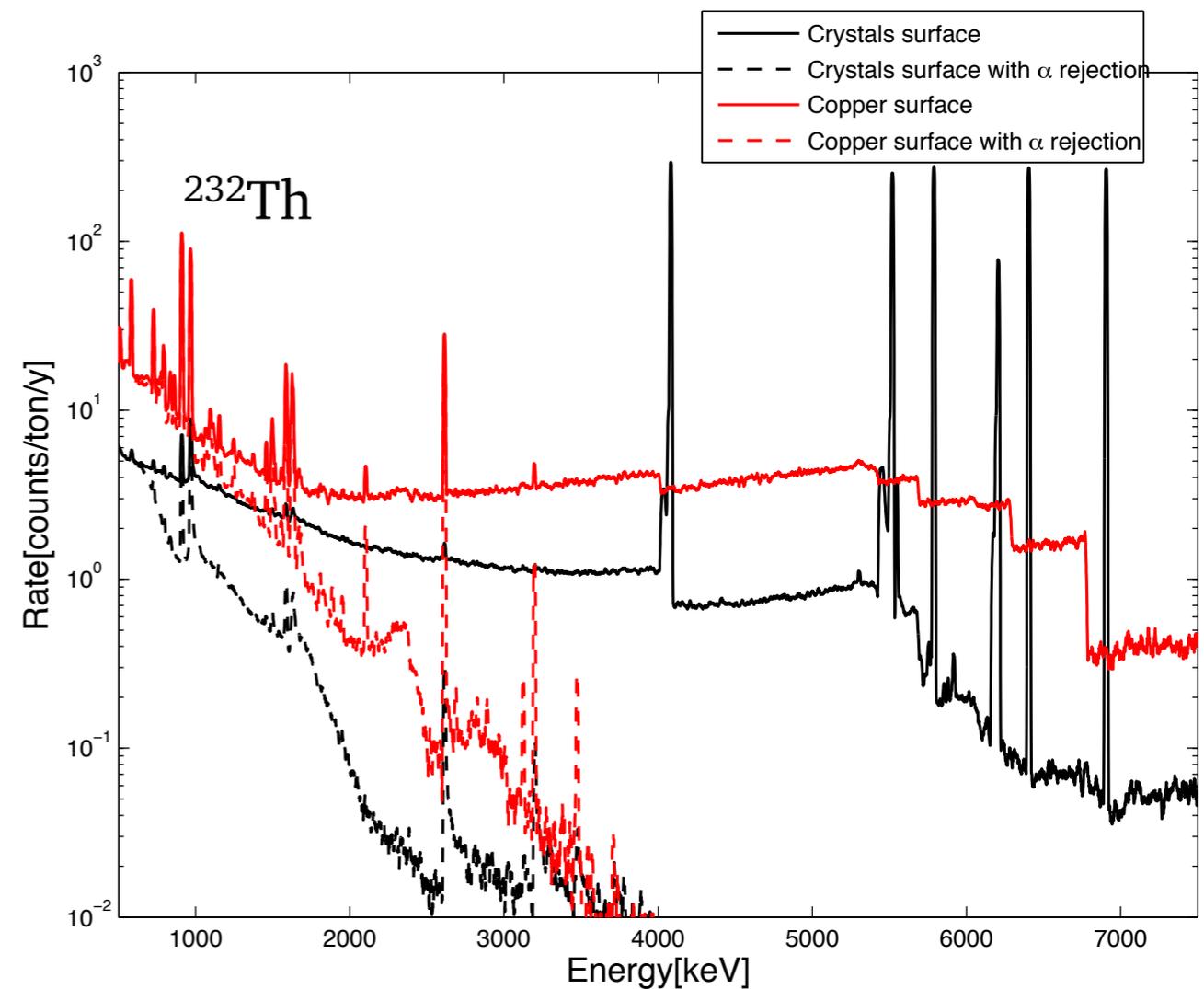
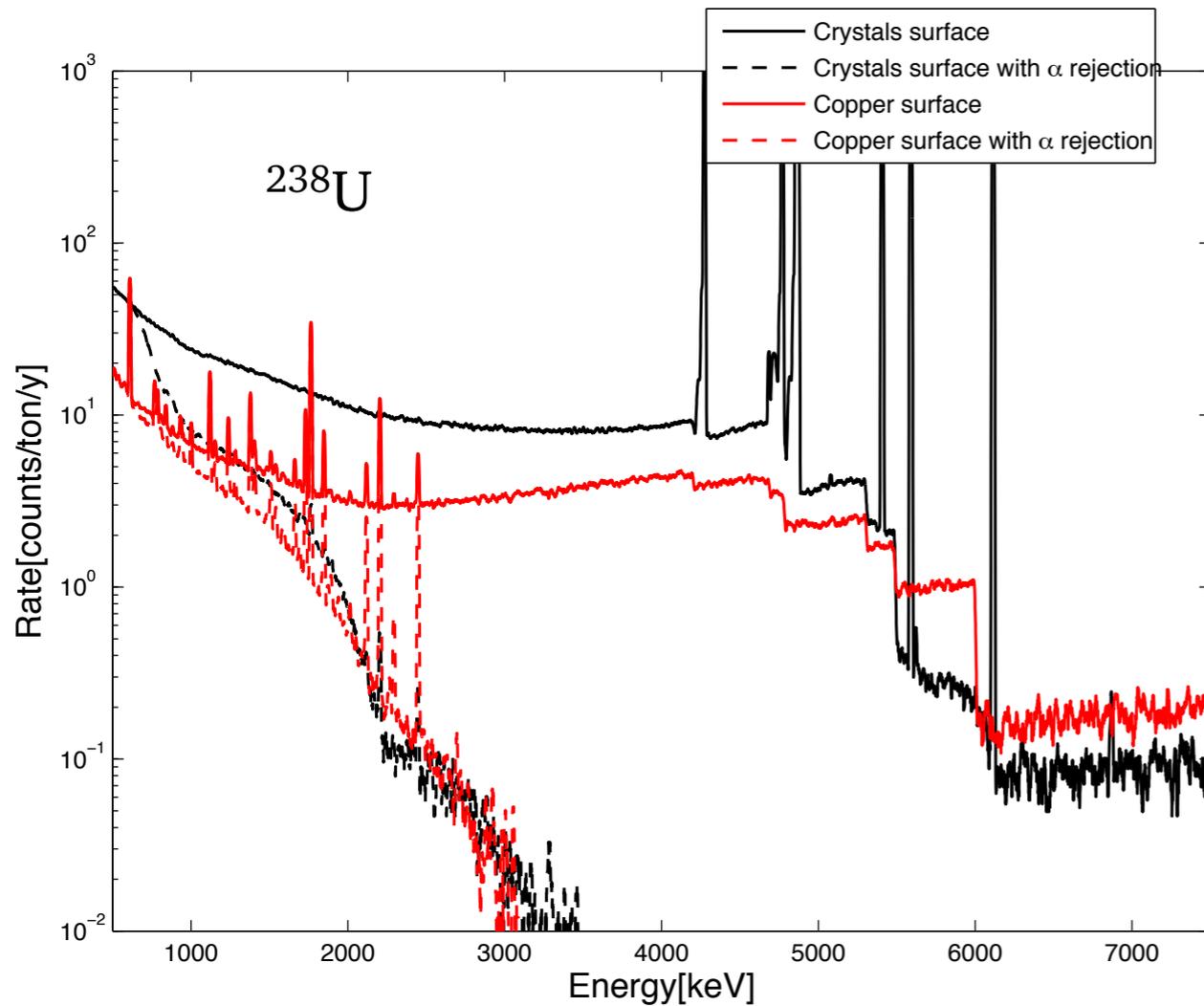
$$\langle m_\nu \rangle \propto 1 / \sqrt{T_{1/2}^{0\nu\beta\beta}} \propto 1 / \sqrt{Nt}$$

With background subtraction

$$\langle m_\nu \rangle \propto 1 / \sqrt{T_{1/2}^{0\nu\beta\beta}} \propto 1 / (Nt)^{1/4}$$

# Beyond CUORE: effect of alpha background rejection

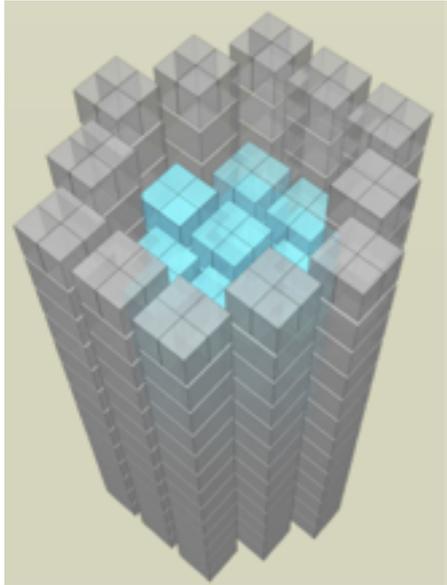
Simulation of surface contamination with an exponential depth profile and a mean depth of 5  $\mu\text{m}$ . Dashed histograms are without  $\alpha$ 's



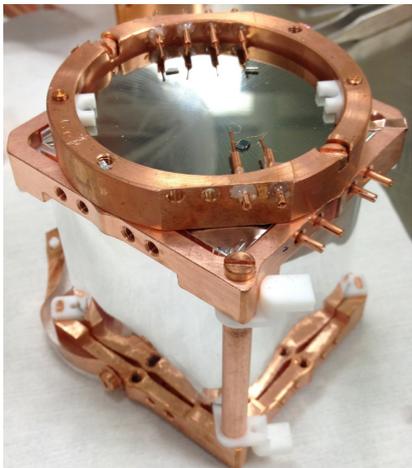
CUORE Collaboration Eur.Phys.J. C74 (2014) 10, 3096

# CUPID: Cuore Upgrade with Particle ID

Te-130 enrichment

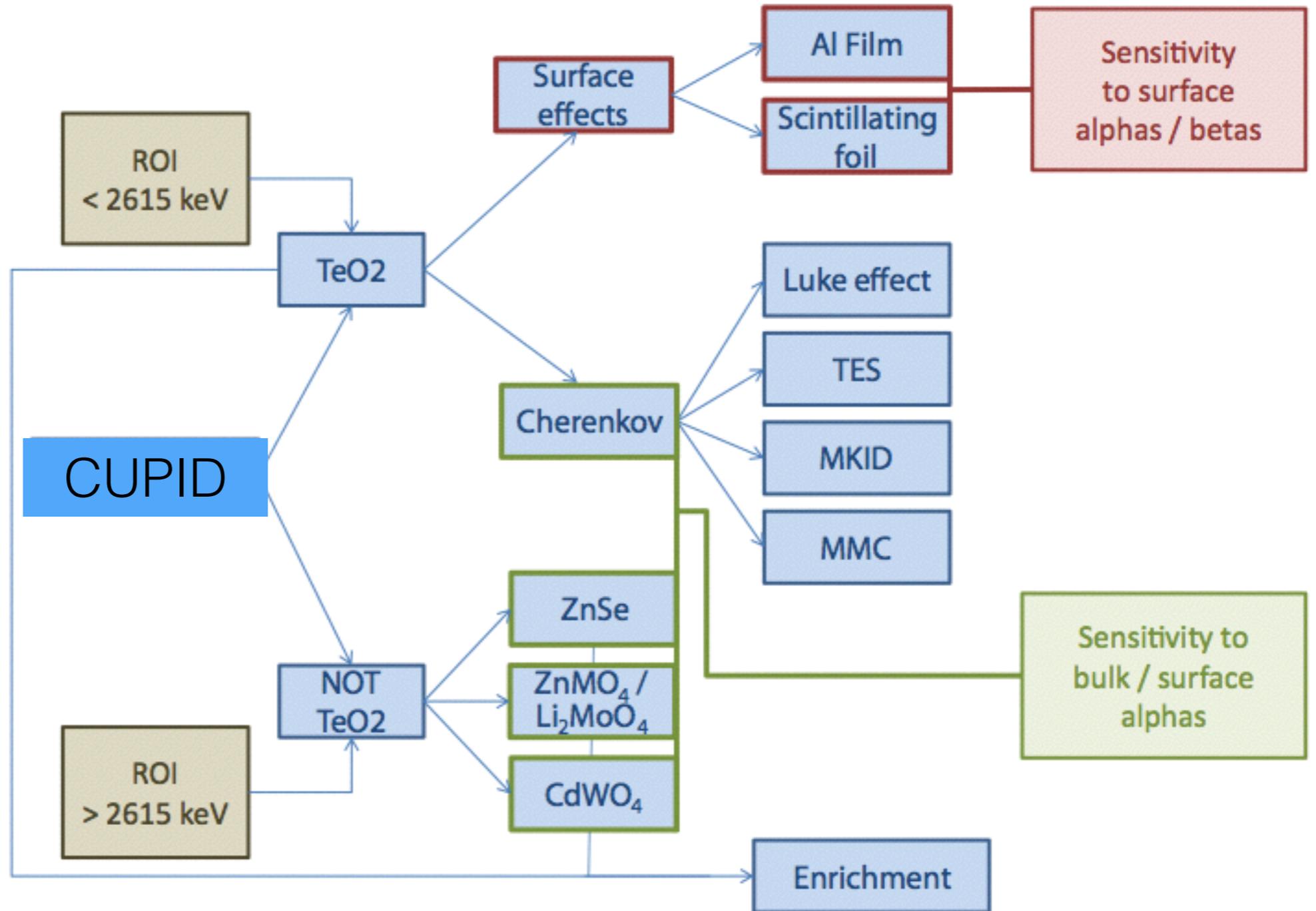


Particle ID



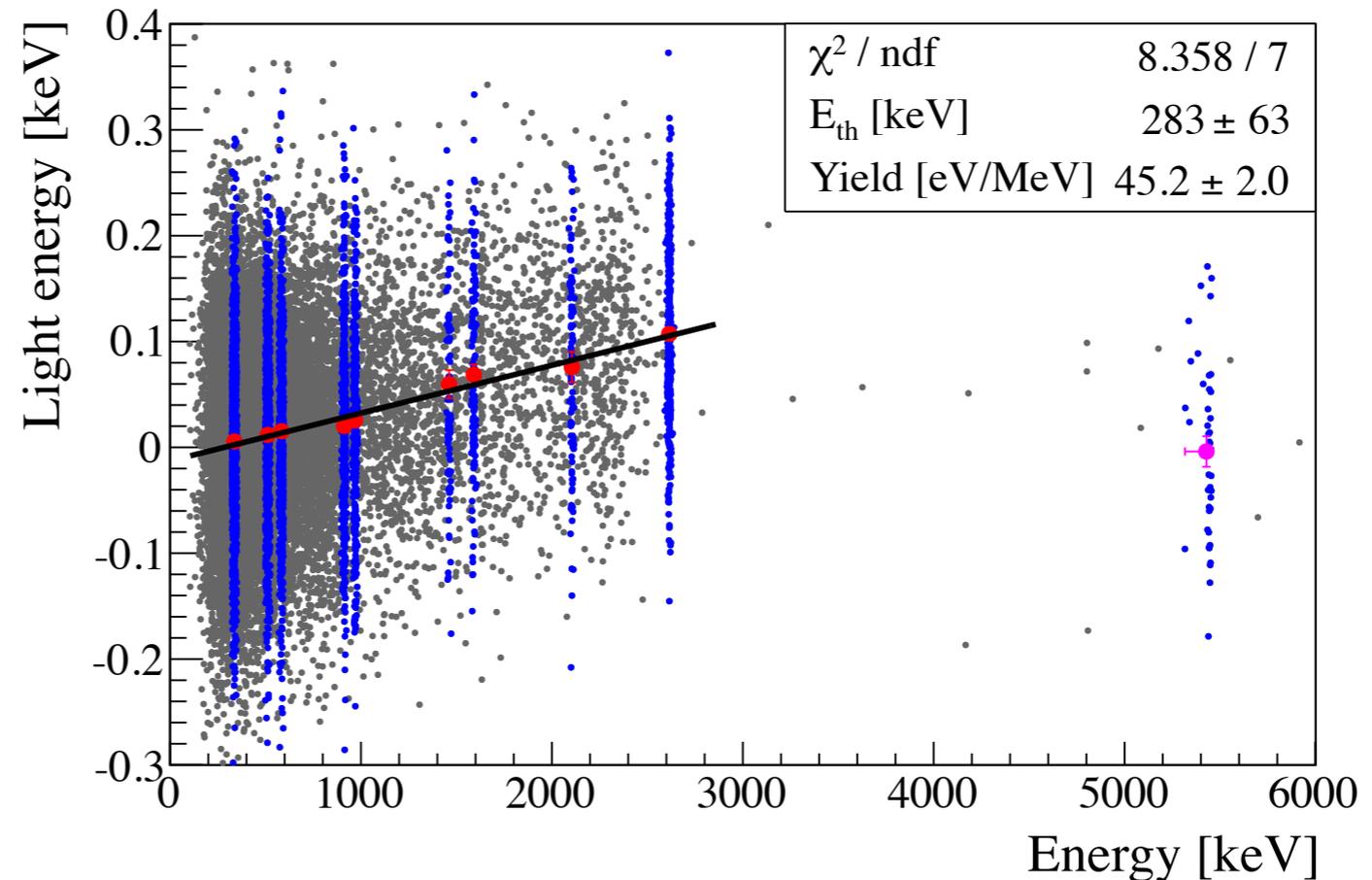
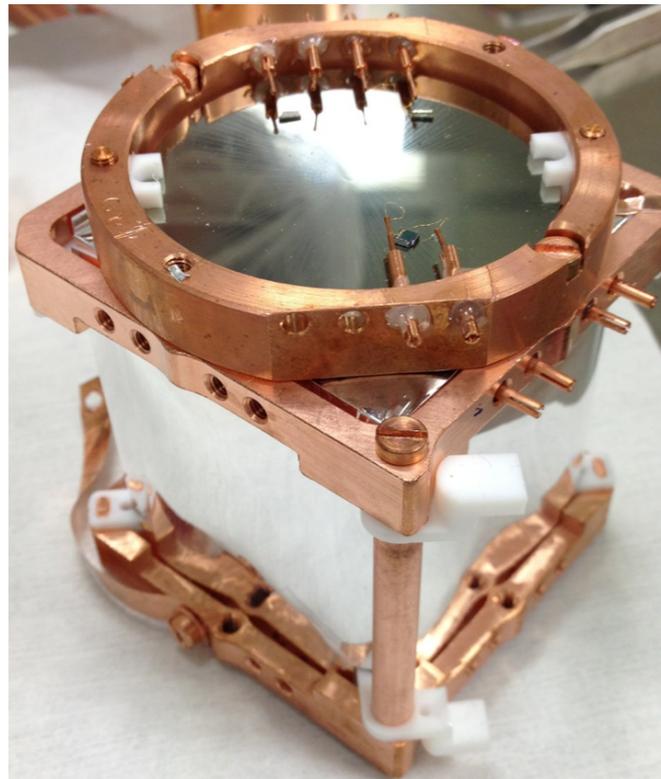
**Bolometer R&D:**

- CALDER
- Cherenkov/TeO<sub>2</sub>
- LUCIFER
- LUMINEU



# Particle ID via a the Light channel

Cherenkov light from a full 750g TeO<sub>2</sub> crystal



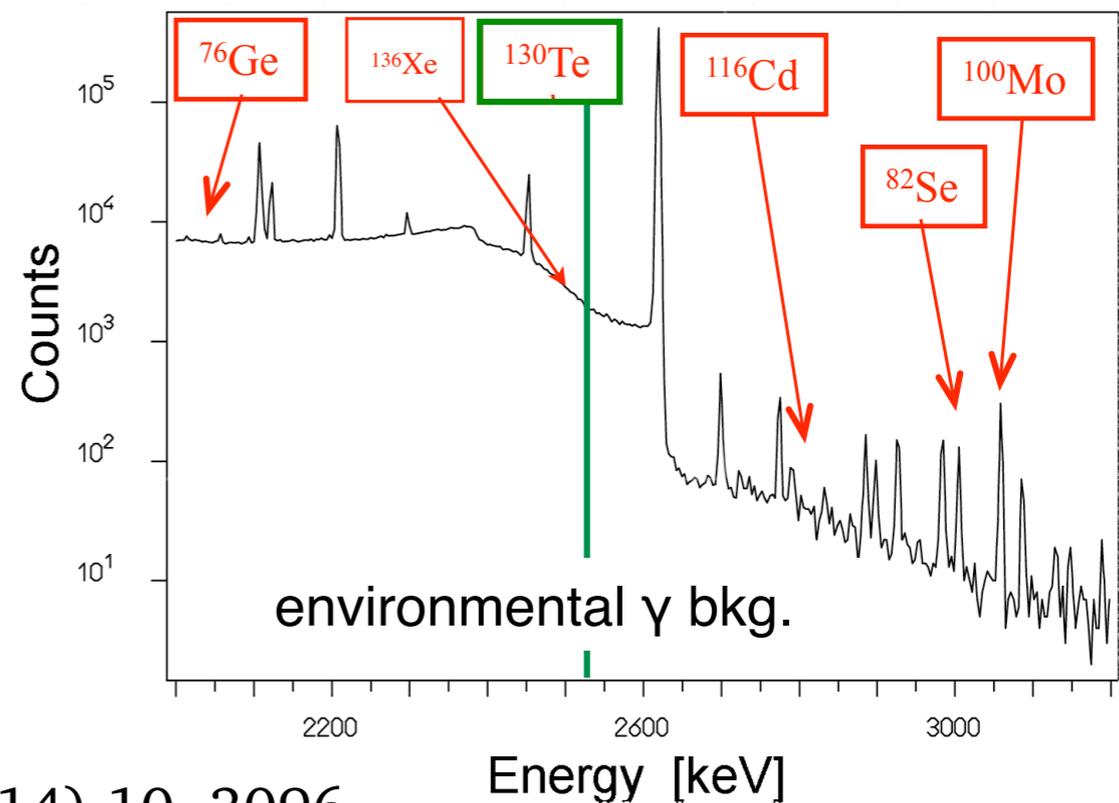
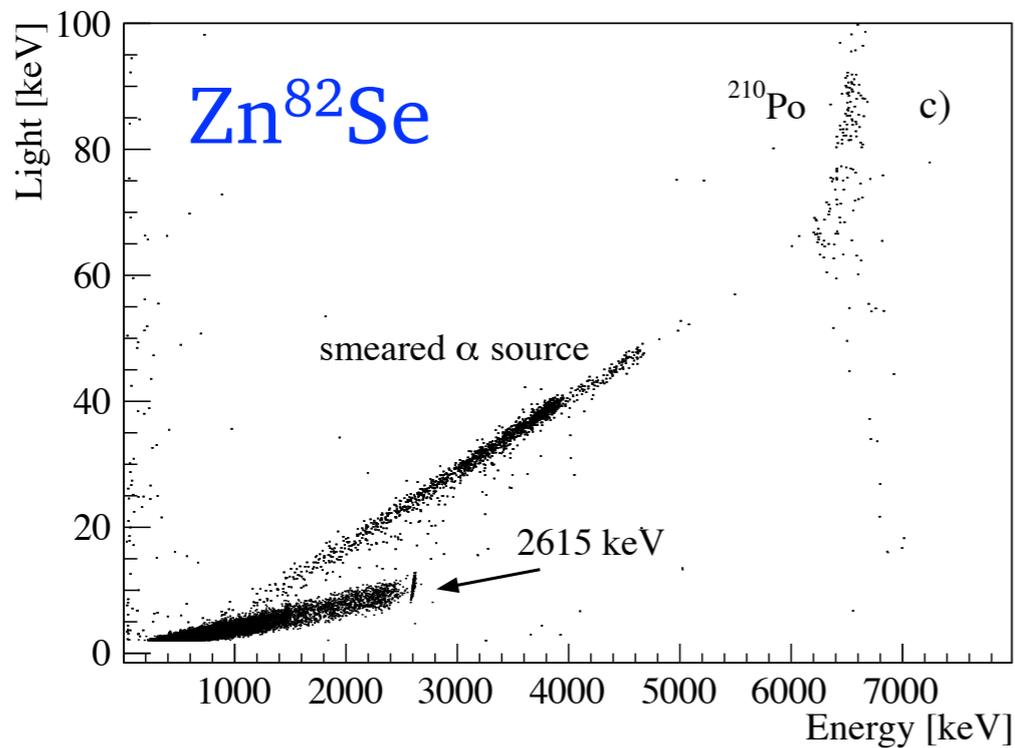
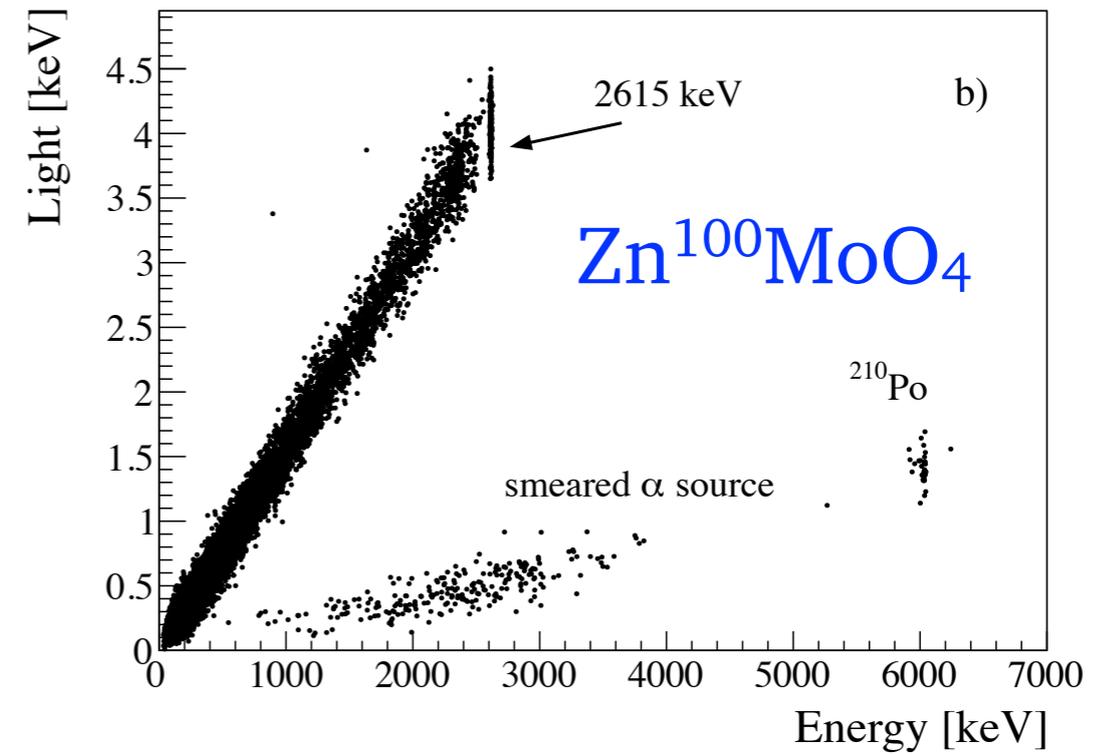
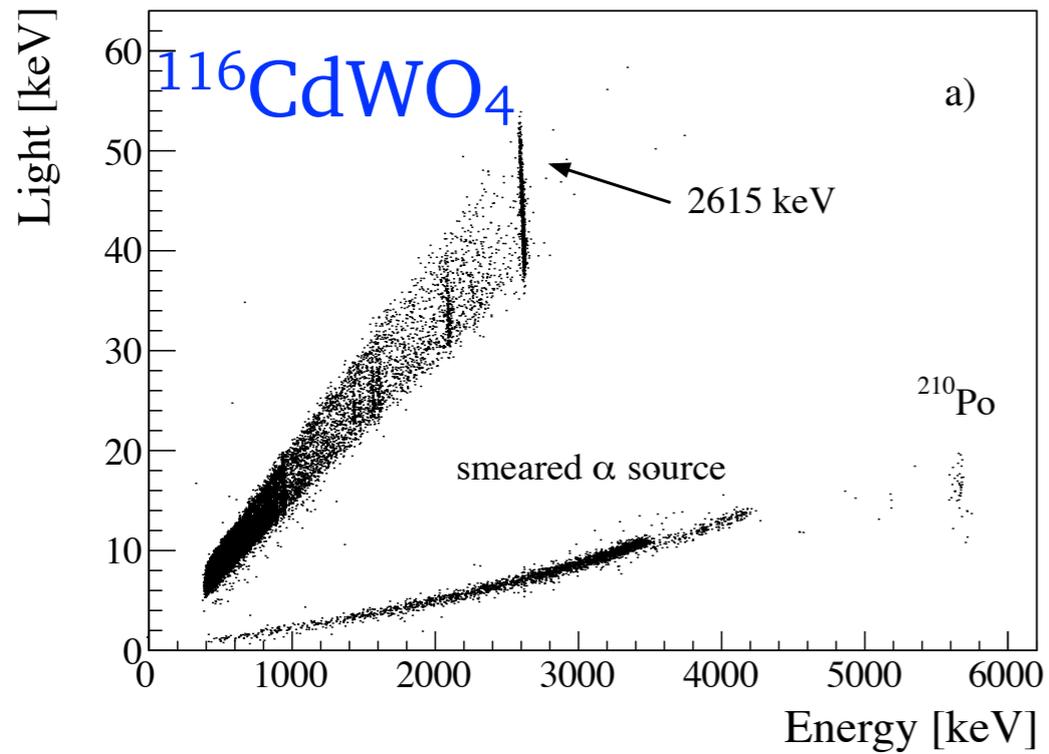
Casali et al. Eur. Phys. J. C75 (2015), 1, 12

101+/- 3.4 eV of light for a  $\beta/\gamma$  event with energy  $0\nu\beta\beta$  value

R&D on improved light detectors currently ongoing:  
MKIDs, TES, Neganov-Luke assisted, MMCs

We need to discriminate between  $\alpha$  and  $\beta/\gamma$  at  $5\sigma$ , ie 99.9% rejection  $\alpha$ 's with 90% efficiency.  
For a light yield of 100 eV implies resolution need of better than 20eV.

# Scintillating crystals



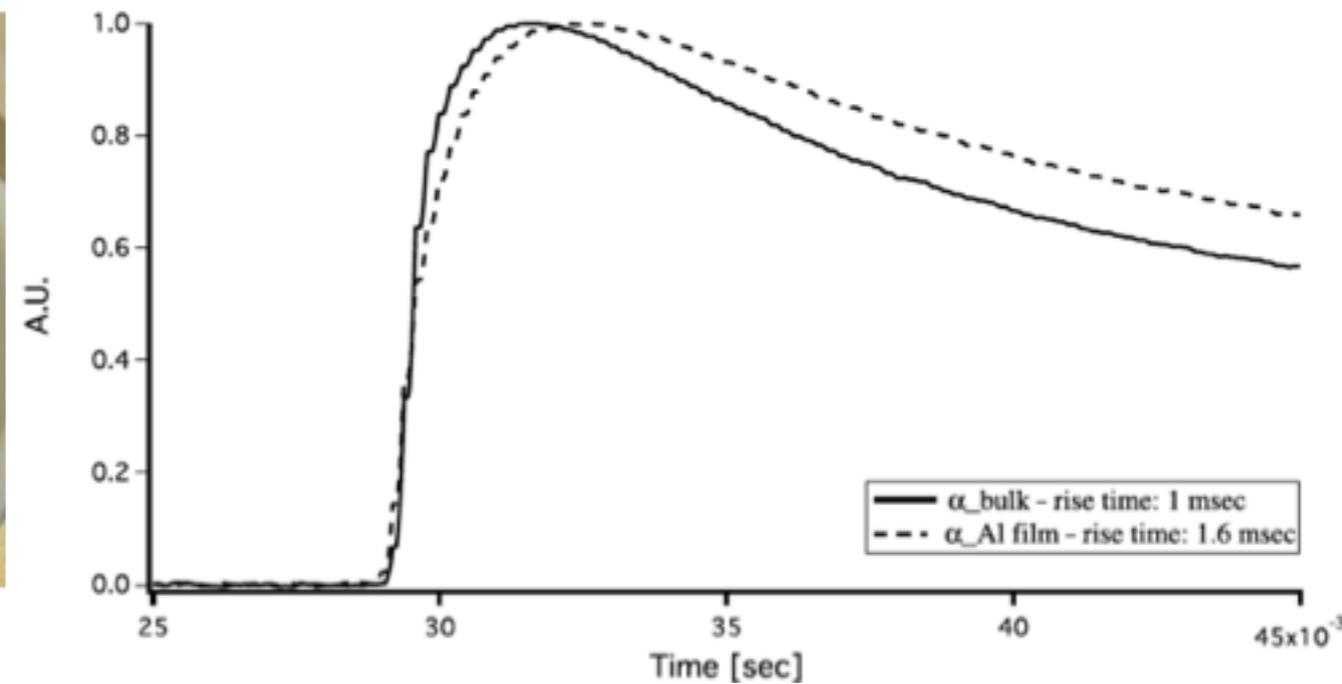
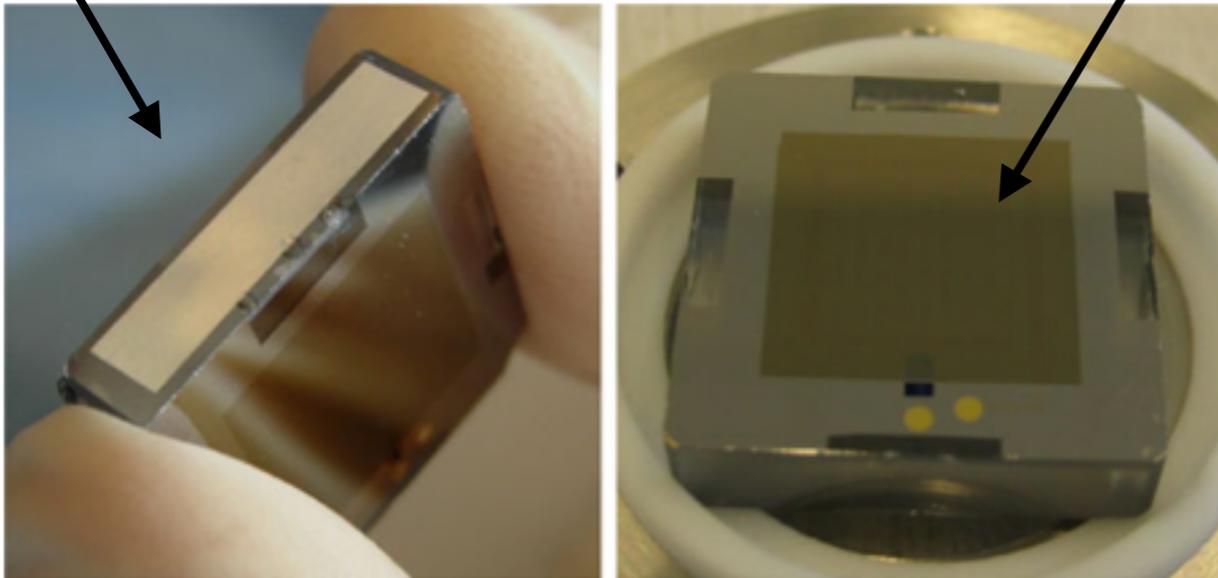
Cuore Collab. Eur. Phys. J. C74 (2014) 10, 3096

Advances in HEP, Vol 2013 (2013), Article ID 237973

# Particle ID via Pulse Shape with Aluminum coating

Aluminum film

NbSi sensor



J. Low Temp. Phys (2012) 167:1029-1034

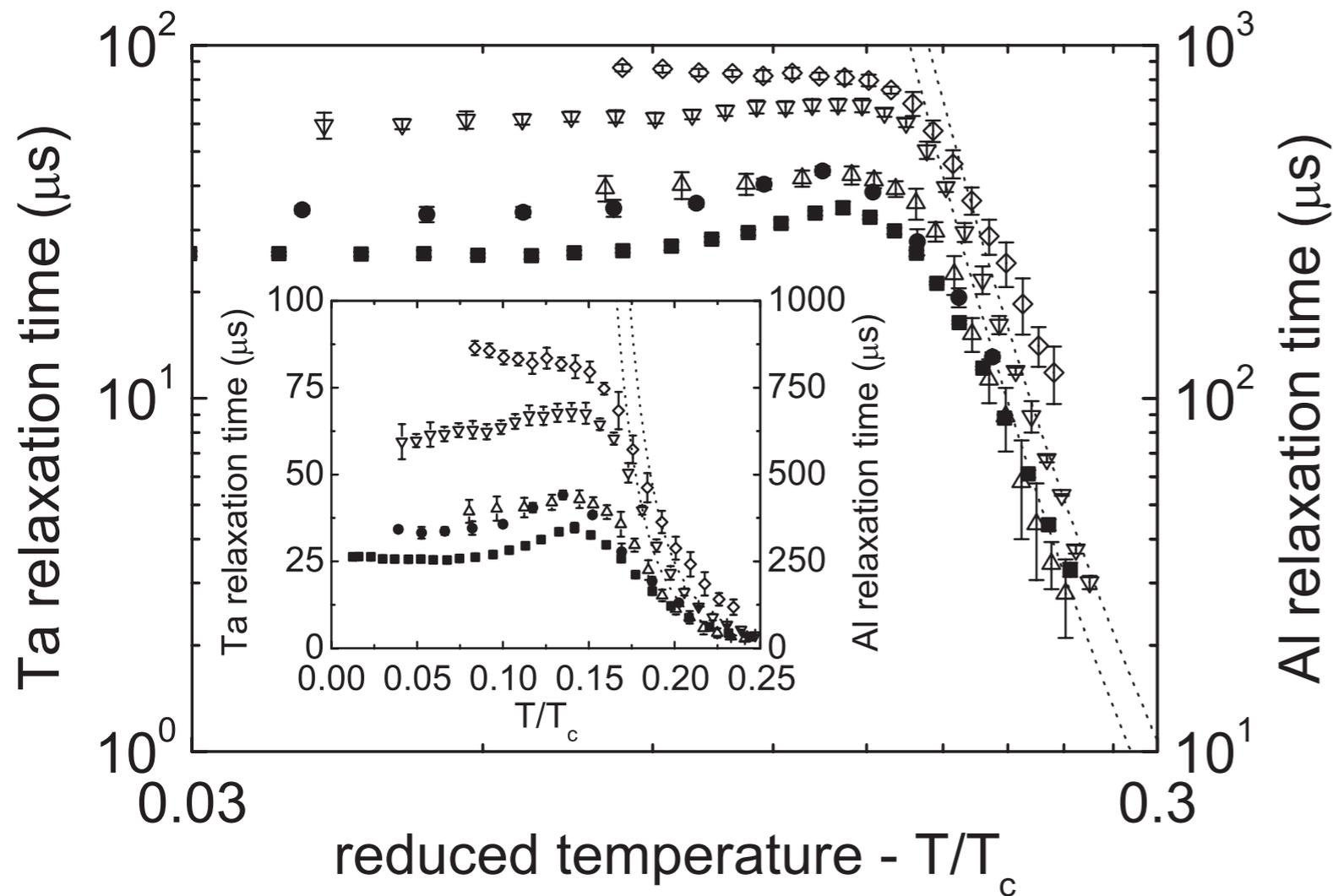
Quasiparticles “recombination” time:

$$\frac{1}{\tau_{\text{rec}}} = \frac{1}{\tau_0} \sqrt{\pi} \left( \frac{2\Delta}{kT_c} \right)^{5/2} \sqrt{\frac{T}{T_c}} e^{-\Delta/kT}$$

for example, for 1  $\mu\text{m}$  of Al on sapphire  $\tau_{\text{rec}} \sim 1.5 \mu\text{sec}$  from J. Shnagl NIMA 444 (2000)245-248

# Quasiparticles recombination time $\tau_R$

$$\frac{1}{\tau_{\text{rec}}} = \frac{1}{\tau_0} \sqrt{\pi} \left( \frac{2\Delta}{kT_c} \right)^{5/2} \sqrt{\frac{T}{T_c}} e^{-\Delta/kT}$$



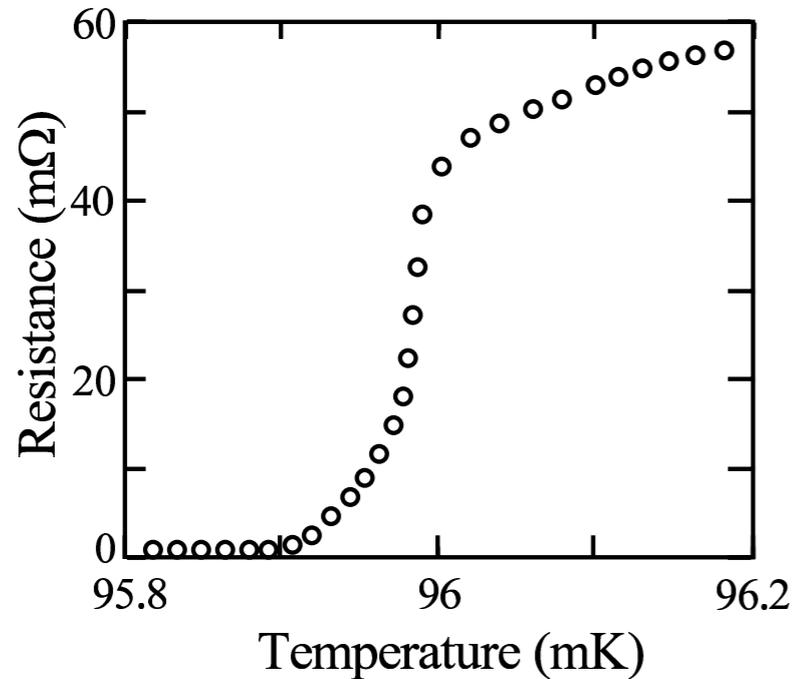
Film (substrate)	Thickness [nm]	$\tau_R$ [ $\mu\text{sec}$ ]
Al(Si)*	100	390
Al(Si)*	250	600
Al(Al <sub>2</sub> O <sub>3</sub> )*	250	860
Al(Al <sub>2</sub> O <sub>3</sub> **)	1000	1530 $\pm$ 350

$\tau_R \sim \sqrt{\text{thickness}}$   
 $\tau_R$  independent of Temp  
 for  $T/T_c < 0.1$   
 $\tau_R$  depends on substrate

\* R. Barends et al. PRL 100, 257002(2008)

\*\* J. Schnagl NIMA 444 (2000)245-248

# Transition Edge Sensors

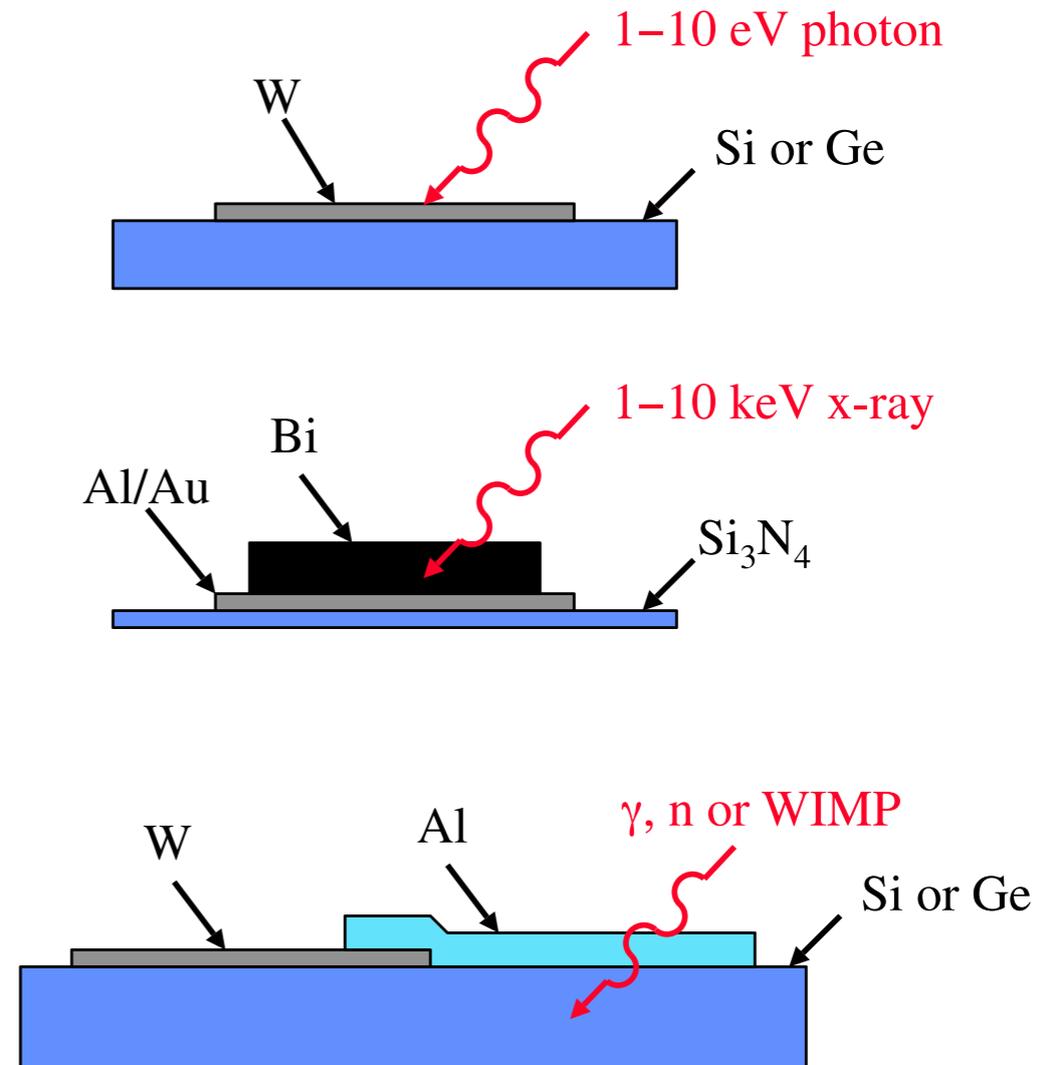


Topics Appl. Phys. 99, 63-149 (2005)

$$\Delta E_{\text{rms}} = \sigma_E \approx \sqrt{\frac{4k_B T^2 C_{\text{tot}}}{\alpha}} \sqrt{\frac{\beta + 1}{2}},$$

$$\alpha \equiv \frac{T}{R} \frac{dR}{dT}$$

$\beta$ : determined by thermal conductivity between the TES and absorber/heat-bath

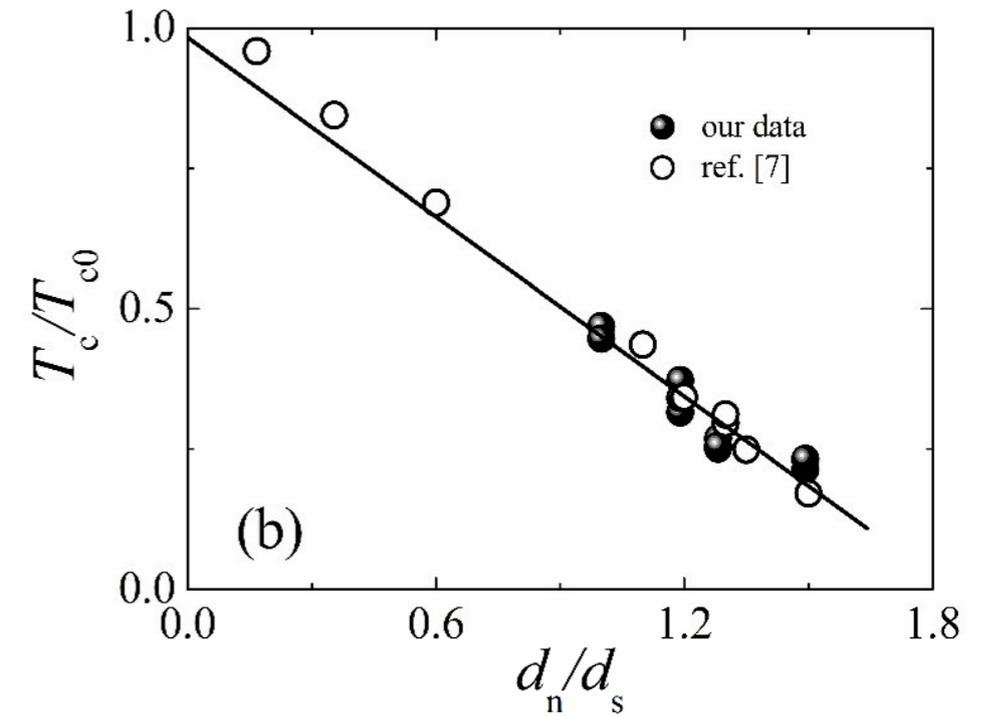
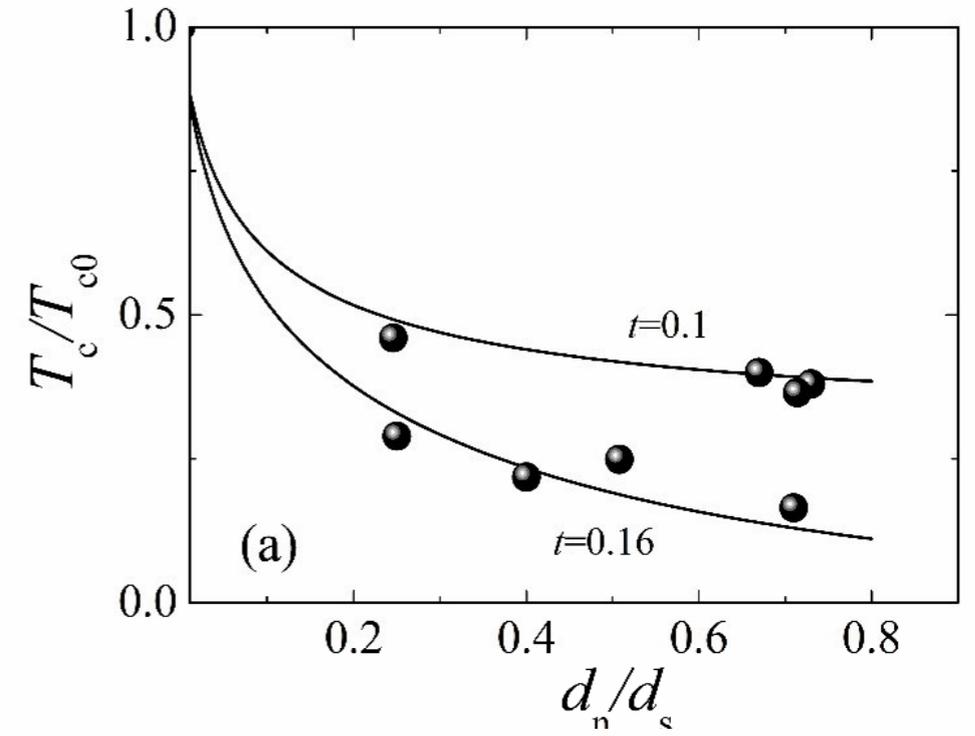
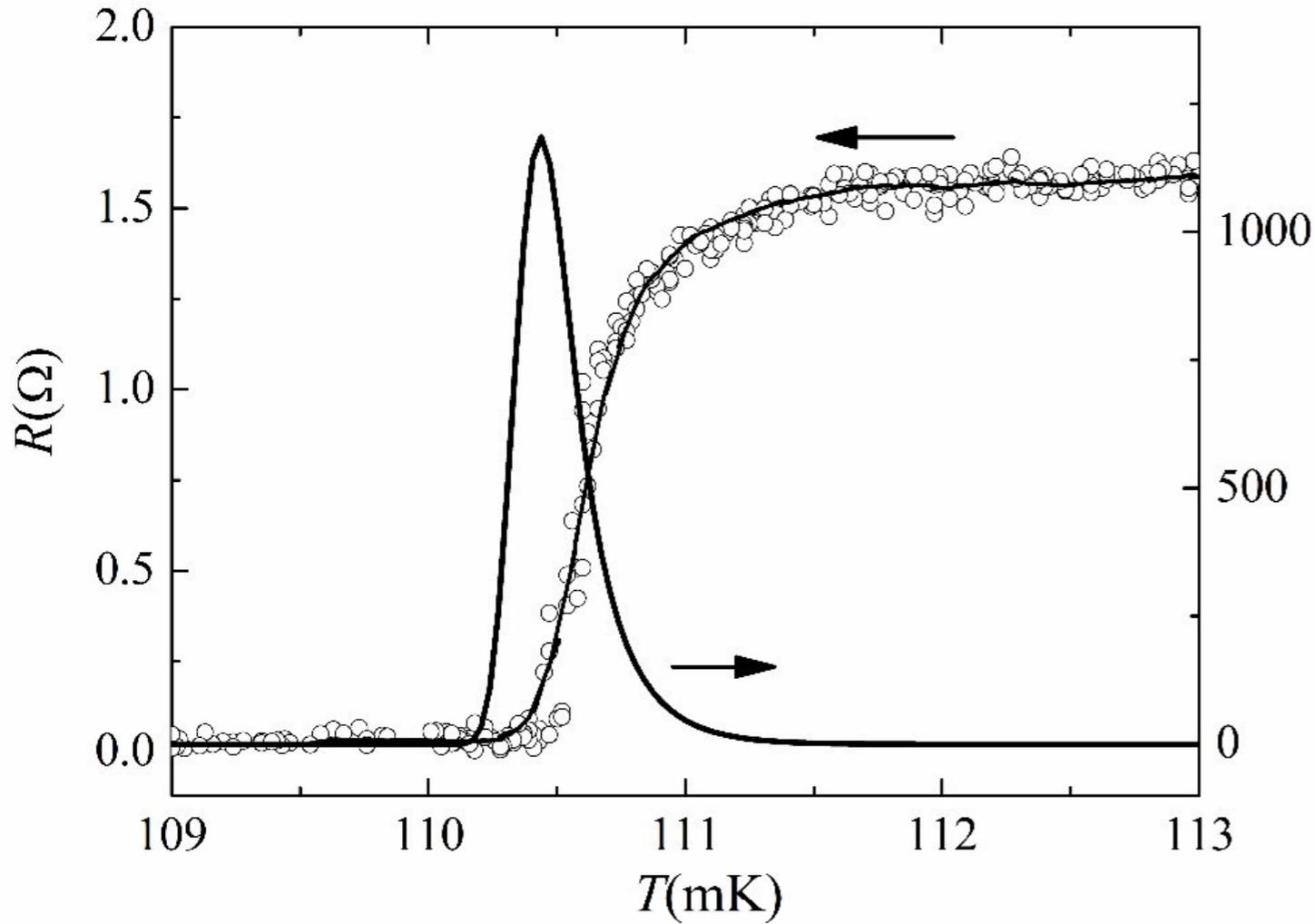


$$C_{\text{tot}} = C_{\text{bolo}} (\sim T^3) + C_{\text{TES}} (\sim T) + C_{\text{other}} \text{ (e.g. caused by impurities in the crystal)}$$

K.D. Irwin, Appl. Phys. Lett. Vol. 66, 1998 (1995)

# Example of Ti/Pd bilayer TES

Journal of Physics: Conference Series **150** (2009) 052168



$$\Delta E_{\text{rms}} = \sigma_E \approx \sqrt{\frac{4k_B T^2 C_{\text{tot}}}{\alpha}} \sqrt{\frac{\beta + 1}{2}}$$

K.D. Irwin, Appl. Phys. Lett. Vol. 66, 1998 (1995)

# TES scalability and low Tc



Tc tuning through  $^{56}\text{Fe}^+$  implantation

Low Tc TES fabrication:

- CDMS W-TES through ion implantation, cryogenic testing
- CRESST W-alpha phase TES
- Or we can utilize superconducting bilayers as TES (proximity effect)

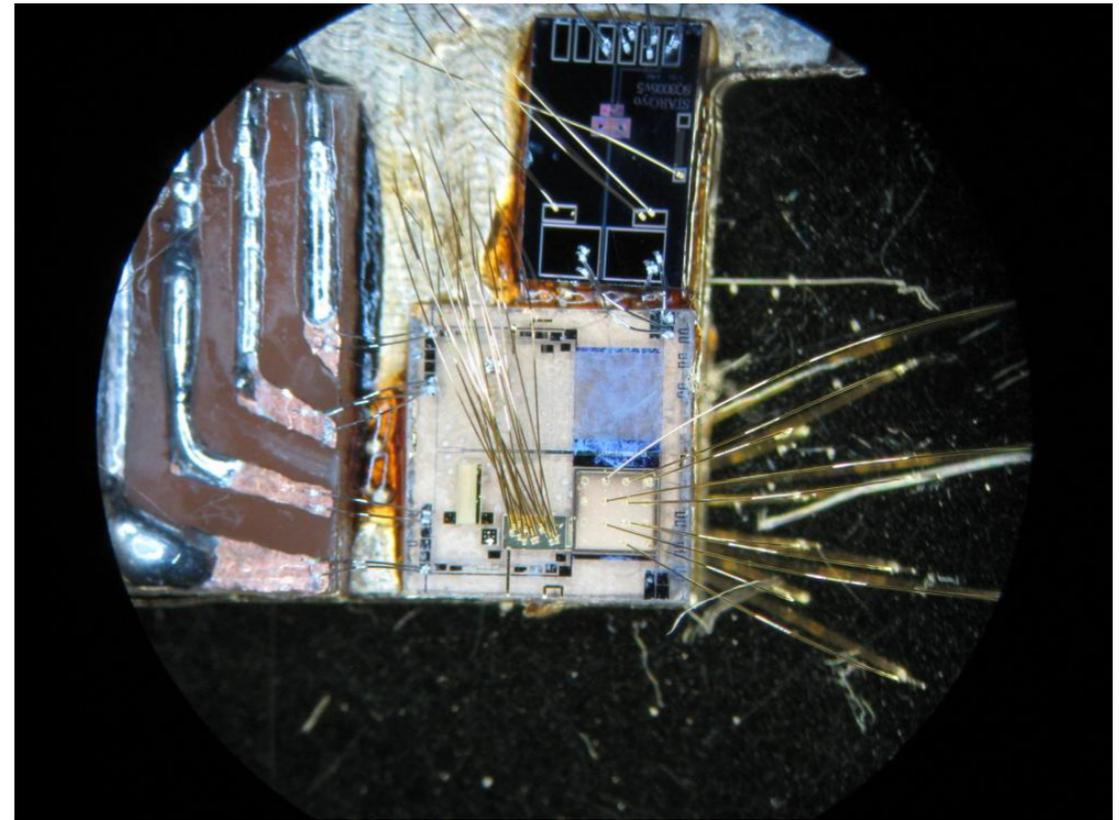
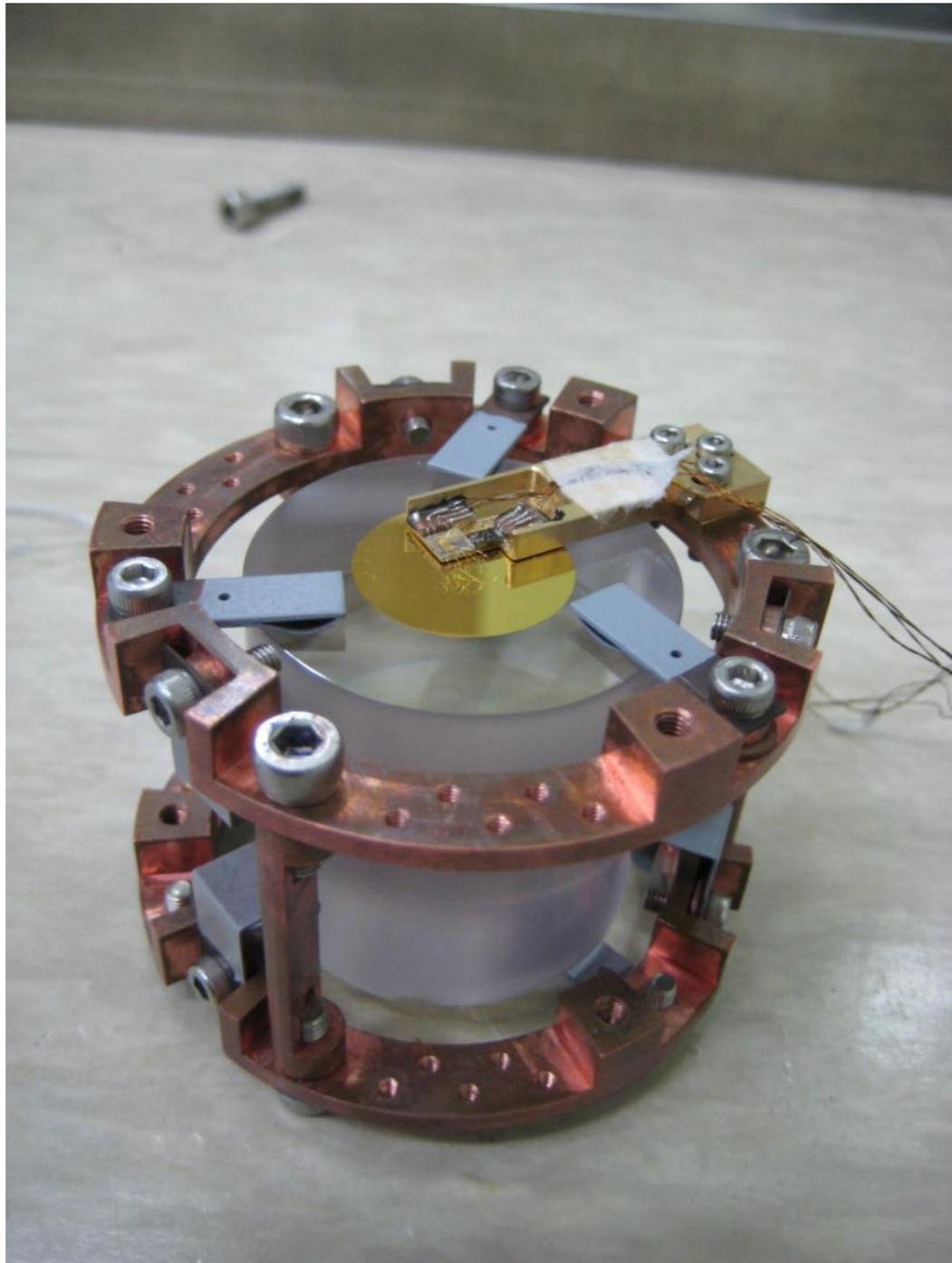
Large quantity (hundreds) production:

- Minimize # TES and SQUIDs on detector
- Simplify and tune thermal coupling TES/crystal-absorber
- Could the TES sit on the thermal bath?
- SQUIDs can be readout in arrays as large as up to 10,000

W-alpha phase for low Tc



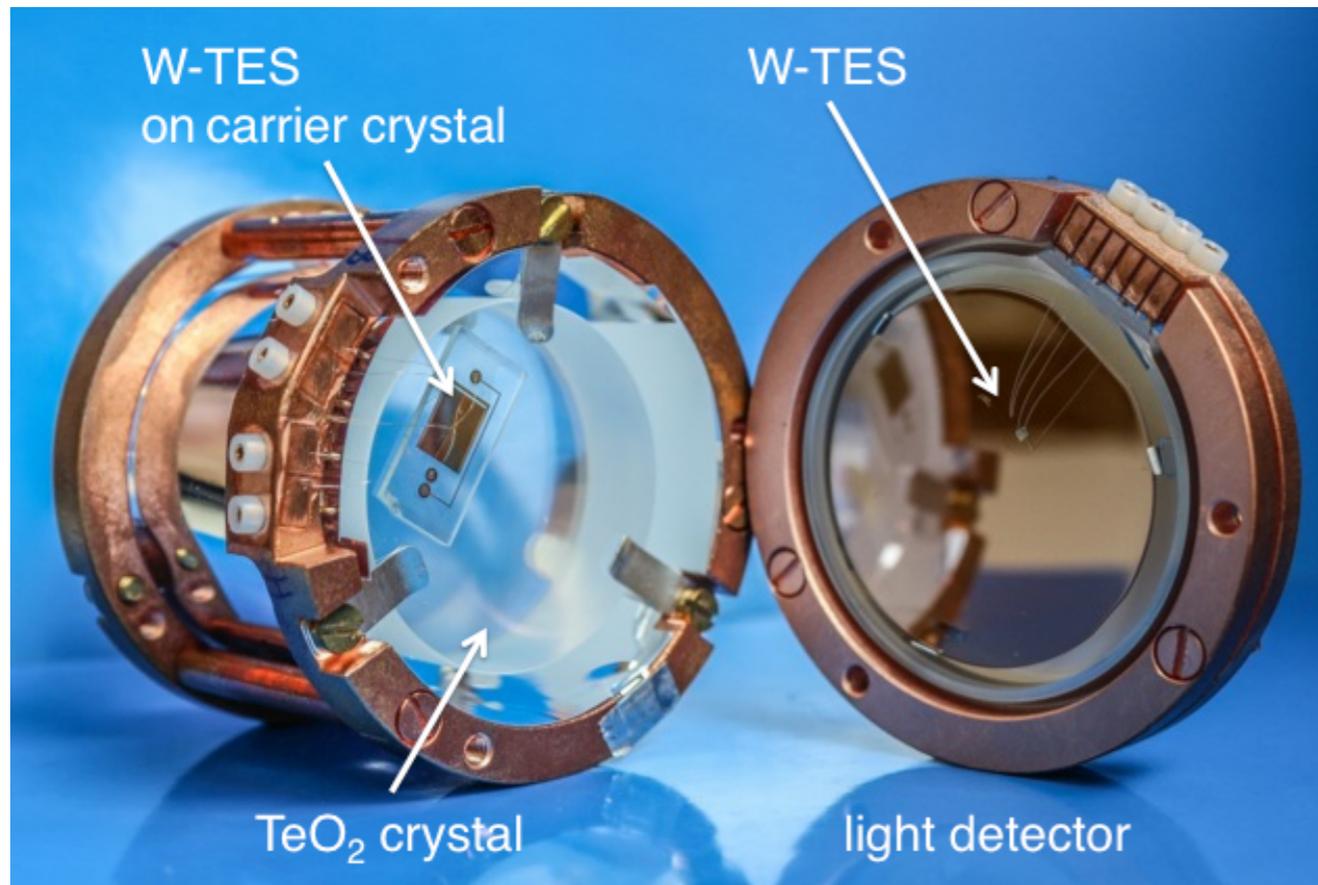
# AMORE prototype from 2012: Au patch coupled via gold wires to a floating MMC - *scalability*



- 2cm diameter, 200nm thickness gold film is evaporated on the center of the crystal top surface.
- Meander chip is placed on the brass support and it is connected to the gold film by ~10 gold wires of 25um diameter.

**For CUPID we need O(1000) channels**

# CRESST-II technology as Light Detector for TeO<sub>2</sub>



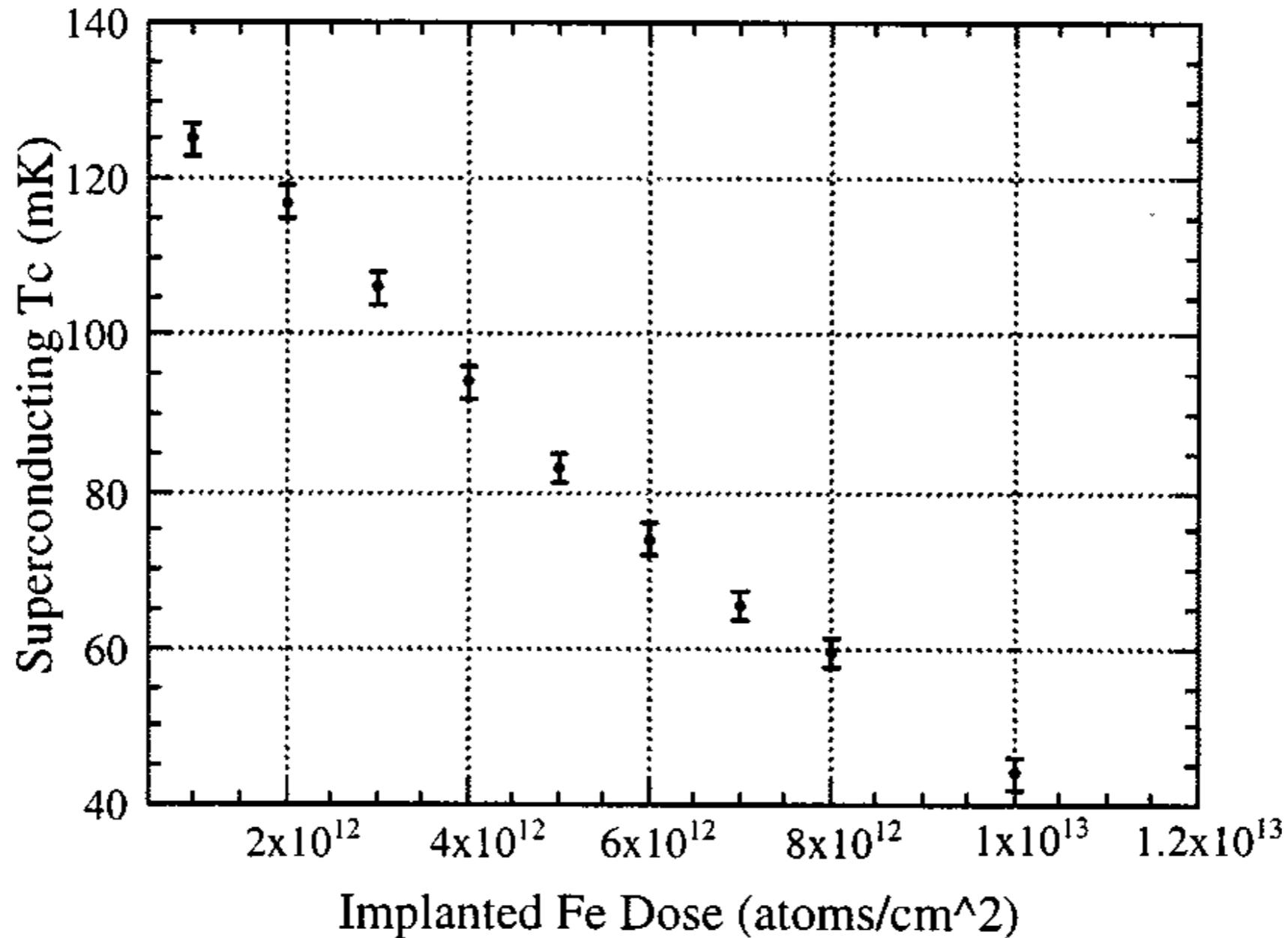
K. Schaffner et al. Astroparticle Physics (2015), pp. 30-36

- 280g TeO<sub>2</sub> crystal, 40mm diameter and height
- A small (20x10x2)mm<sup>3</sup> CdWO<sub>4</sub> carrier crystal, equipped with the W-TES was attached on one of the flat surfaces of the TeO<sub>2</sub> by vacuum grease
- Light Detector CRESST-II type, sapphire disc (Silicon on Sapphire) readout by W-TES
- Light Detector reached RMS baseline  $\sigma=24\text{eV}$ , although CRESST-II has obtained  $\sigma=5\text{eV}$

Absorber signal and sensor bandwidth mismatch minimization may improve energy resolution in this type of detectors as predicted by model in Pyle et al. arXiv:1503.01200

# Tc tuning in W for TES applications in CDMS

J. Appl. Phys., Vol. 86, No. 12, 15 December 1999



Measured superconducting transition temperature for 350 Å-thick W films implanted with Fe-56 ions at 50 keV kinetic energy

# Tc calculation for a superconducting bilayer using Usadel model

NIM A Vol. 444 (2000) 23-27

For thin films:

$$T_C = T_{C0} \left[ \frac{d_s}{d_0} \frac{1}{1.13(1 + 1/\alpha)} \frac{1}{t} \right]^\alpha$$

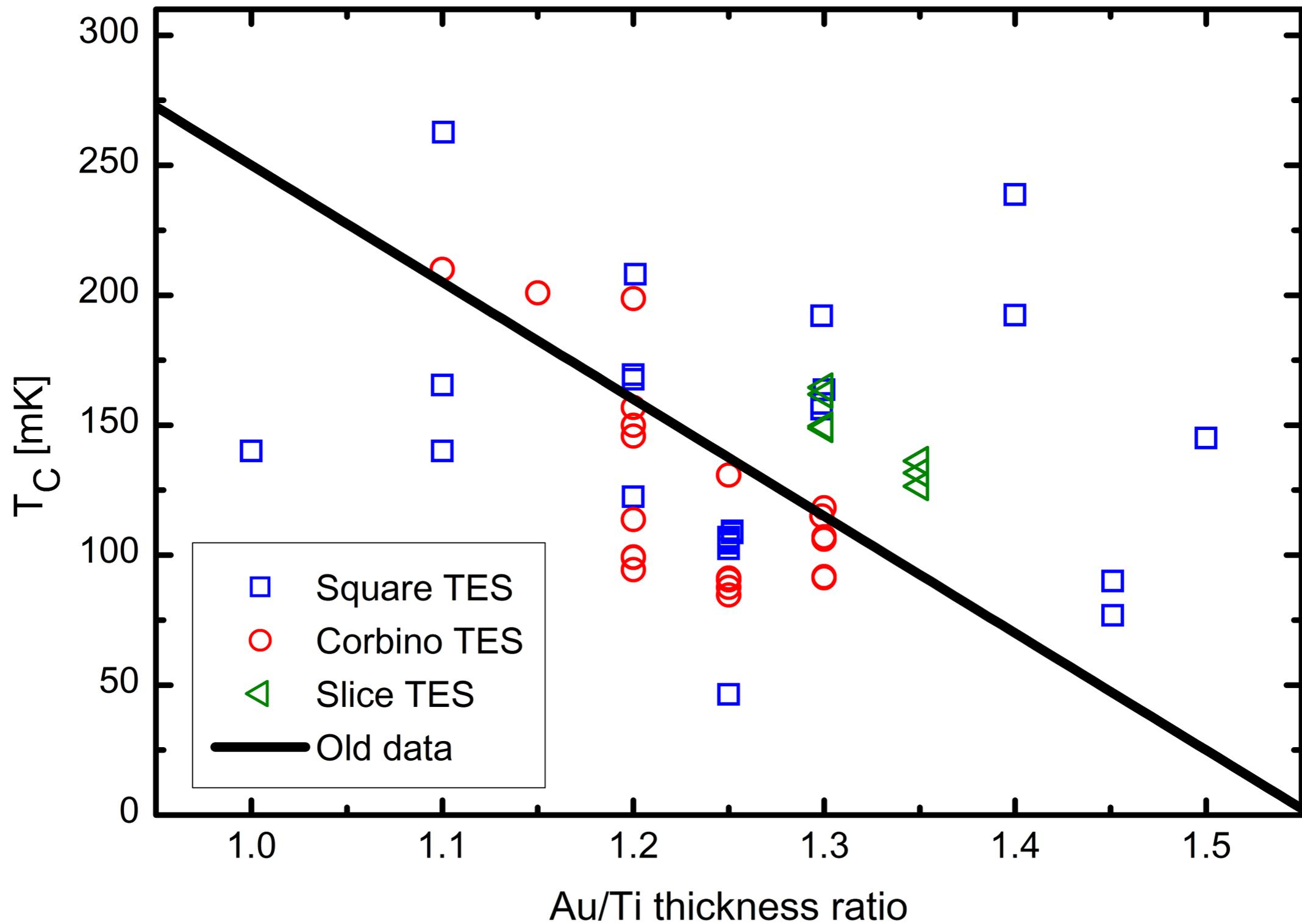
$$1/d_0 = (\pi/2)k_B T_{C0} \lambda_f^2 n_s$$

$$\alpha = d_n n_n / d_s n_s.$$

For a MoCu bilayer,  $d_0 = 1.18 \mu\text{m}$ ,  $n_n/n_s = 0.431$ ,  $\lambda_f = 0.464\text{nm}$ .  
From data they obtained  $t = 0.21$

# Although experimentally is a very challenging task

Kinnunen PhD Thesis

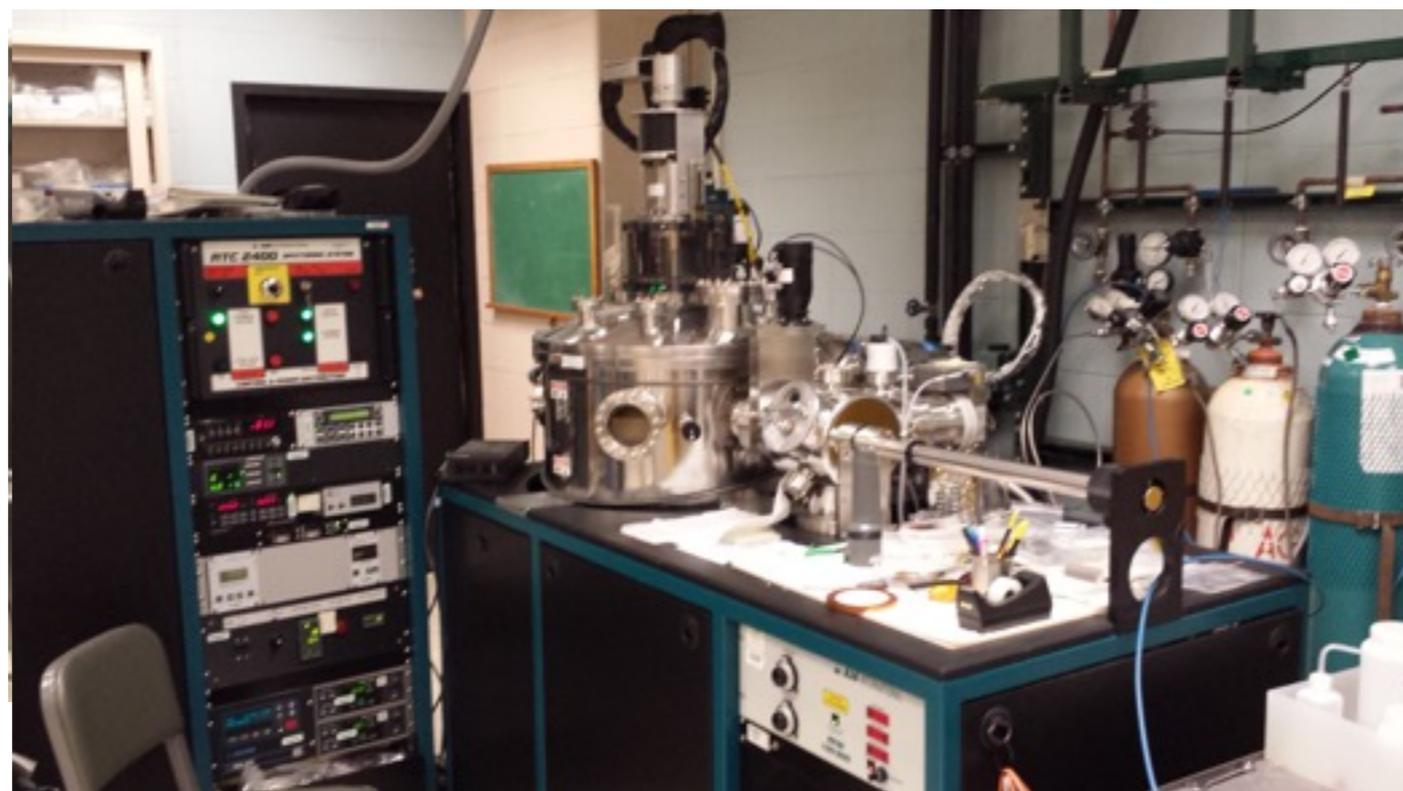


# Sample production at ANL and testing at Berkeley

Collaborating at ANL with G. Wang, V. Novosad, V. Yefremenko and C. Chang



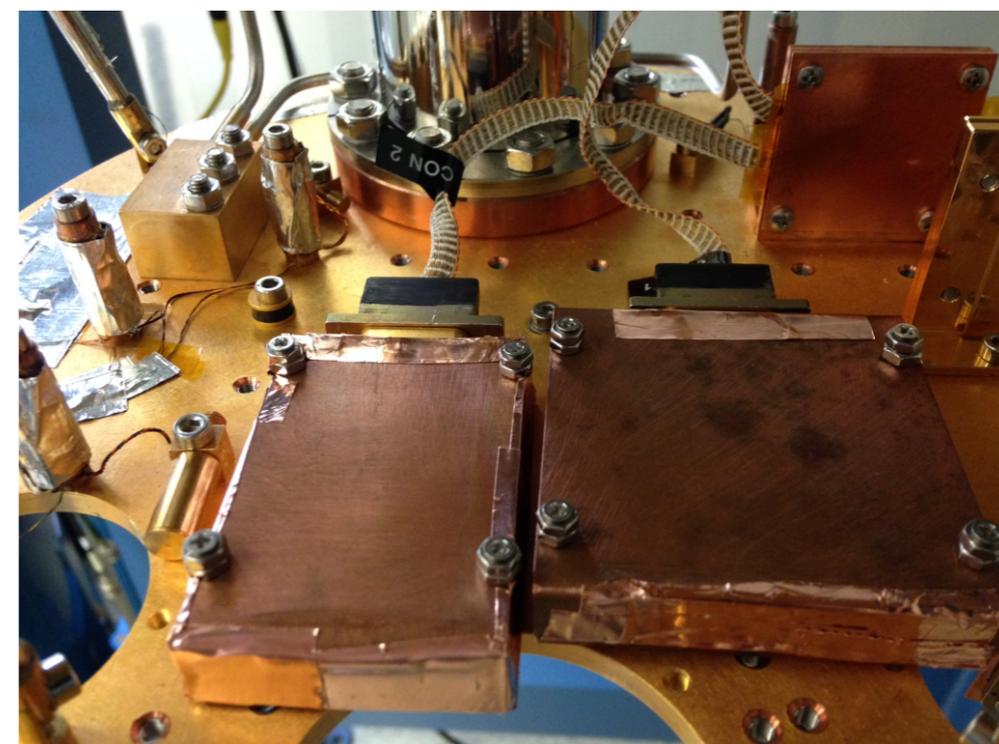
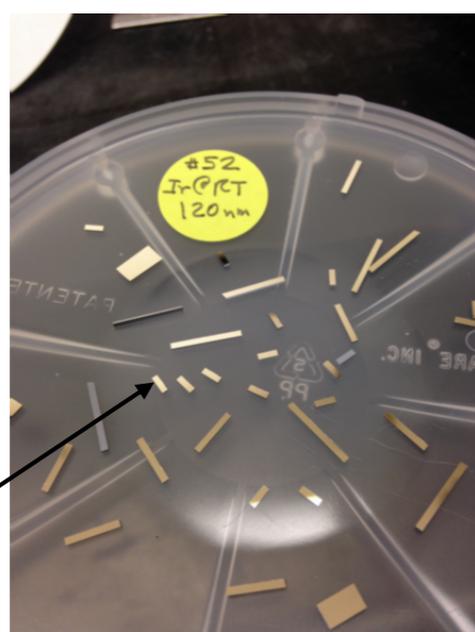
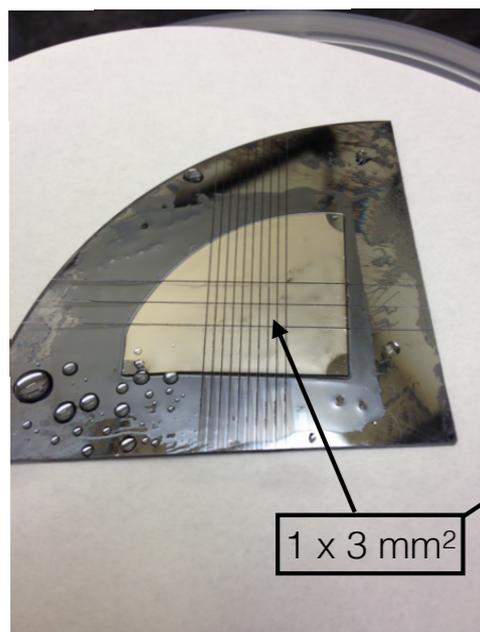
Cryogen free DR and Co-60 thermometry down to  $\sim 7\text{mK}$



At ANL with sputtering chamber on the back

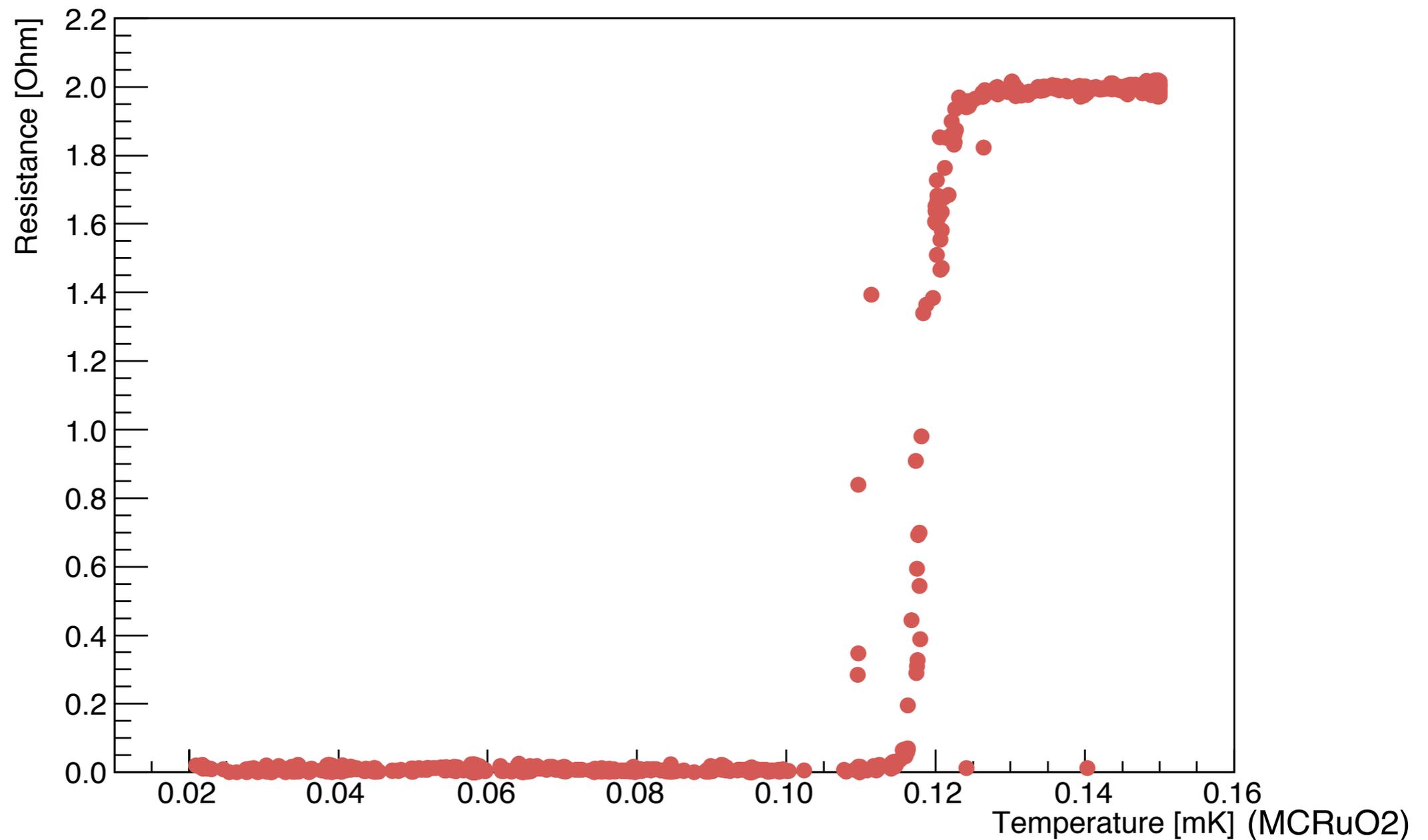


Iridium bilayers



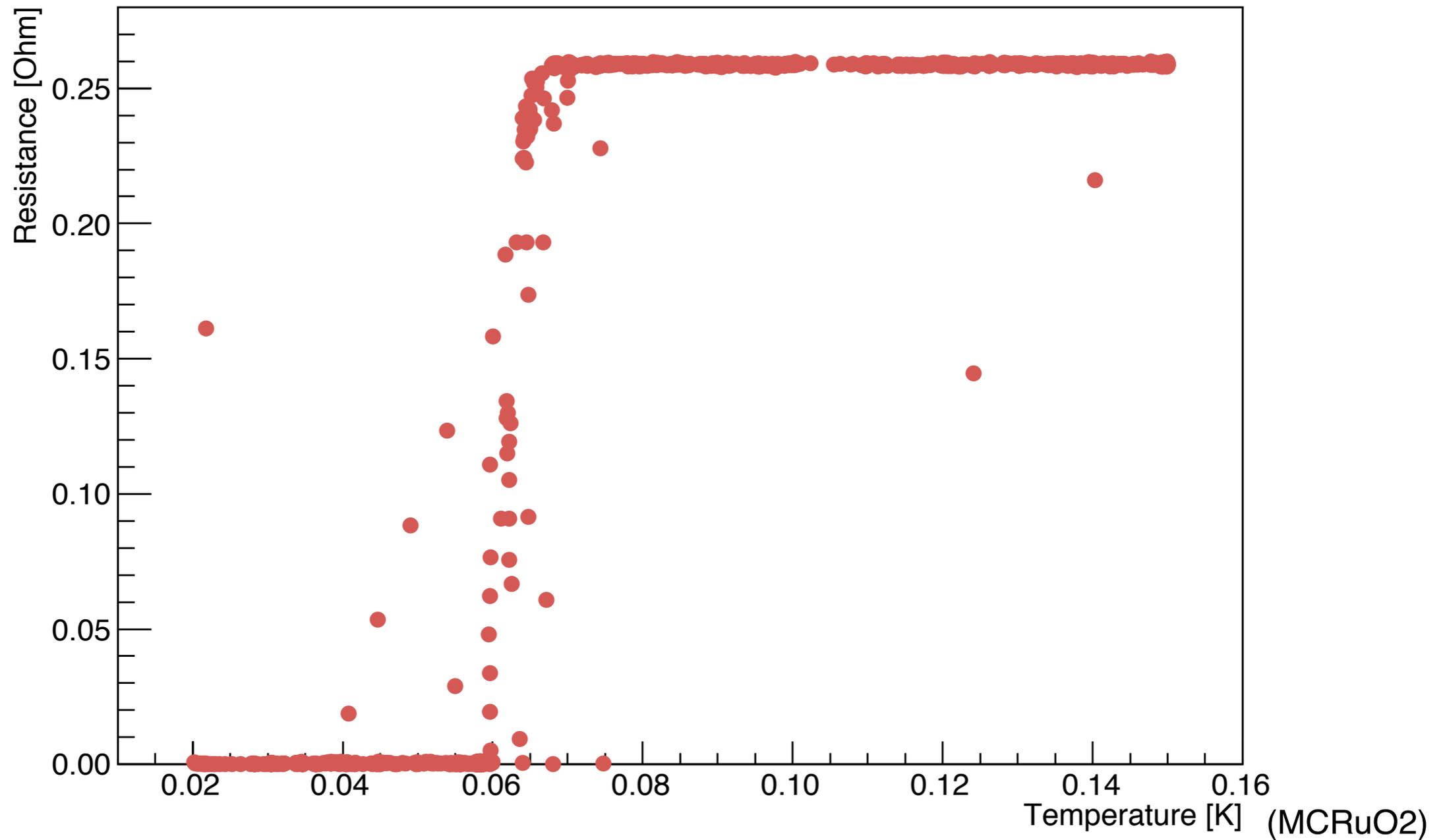
Ir at RT / Au  
Ir = 80 nm  
Au = 120 nm

$T_c = 120$  mK  
Channel 07



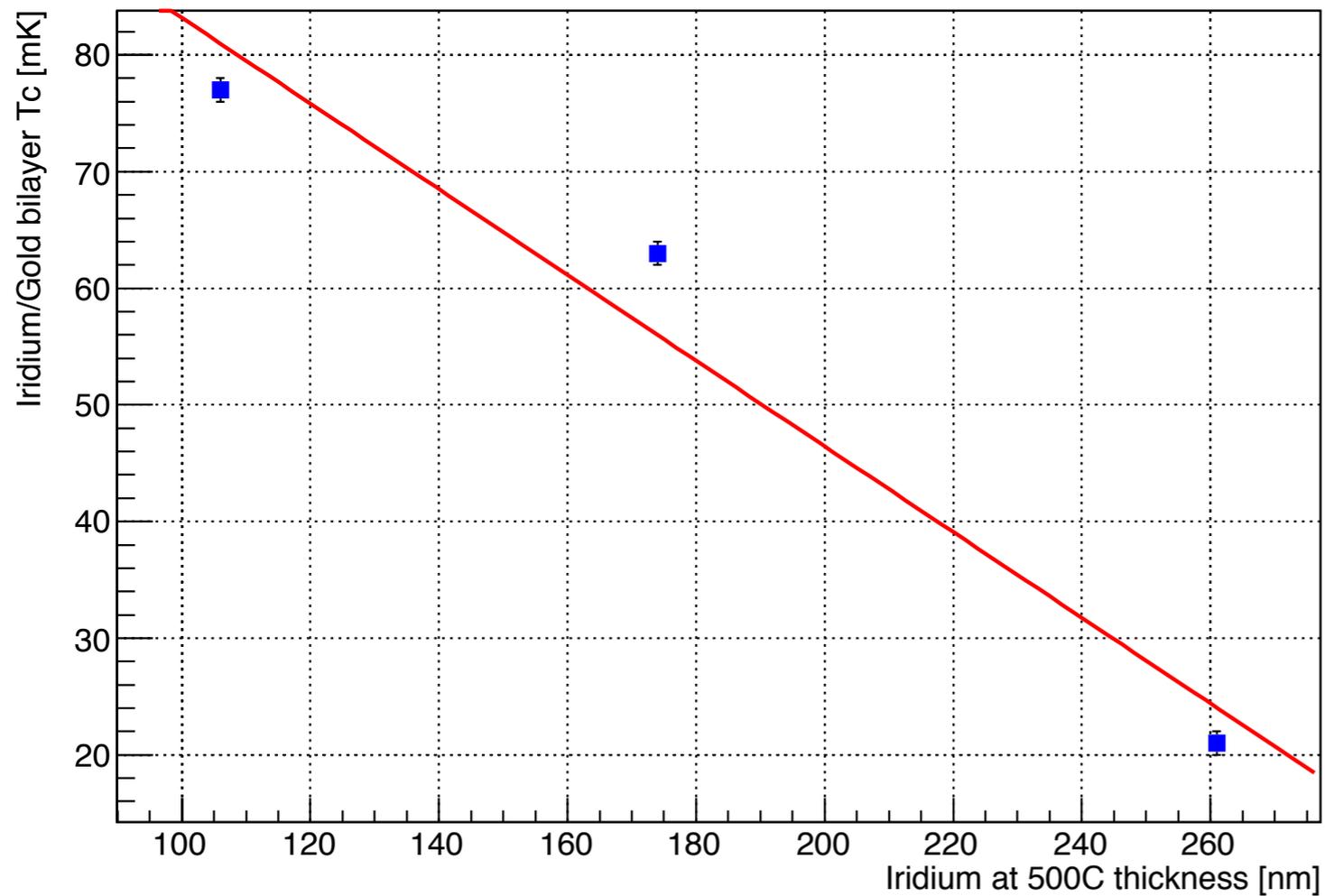
Ir at 600 °C / Au  
Ir = 80 nm  
Au = 120 nm

$T_c = 60$  mK  
Channel 06

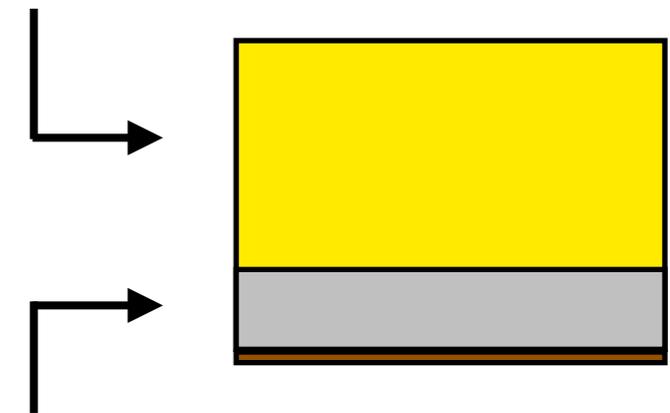


# Ir@500C/Au bilayer $T_c$ as a function of Au thickness

Ir = 101nm at 500C / Au = x-axis thickness



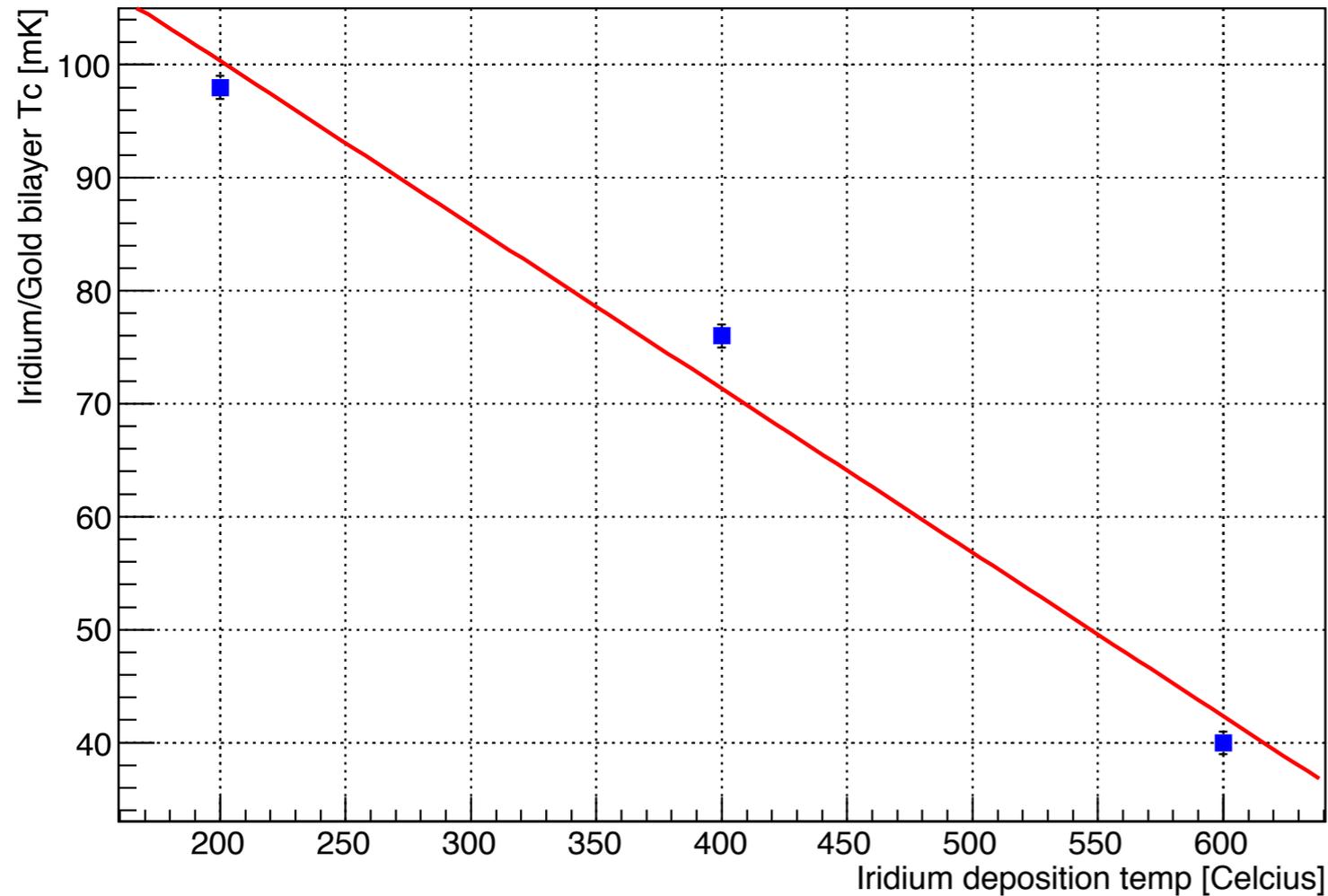
**Au layer at room temp**



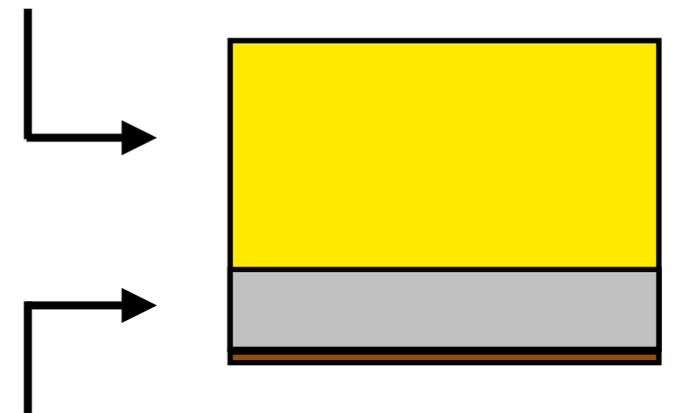
**100nm Ir deposited at 500C**

# Ir/Au bilayer $T_c$ as a function of Ir temp

Ir = 80nm at x-axis temp / Au = 160nm



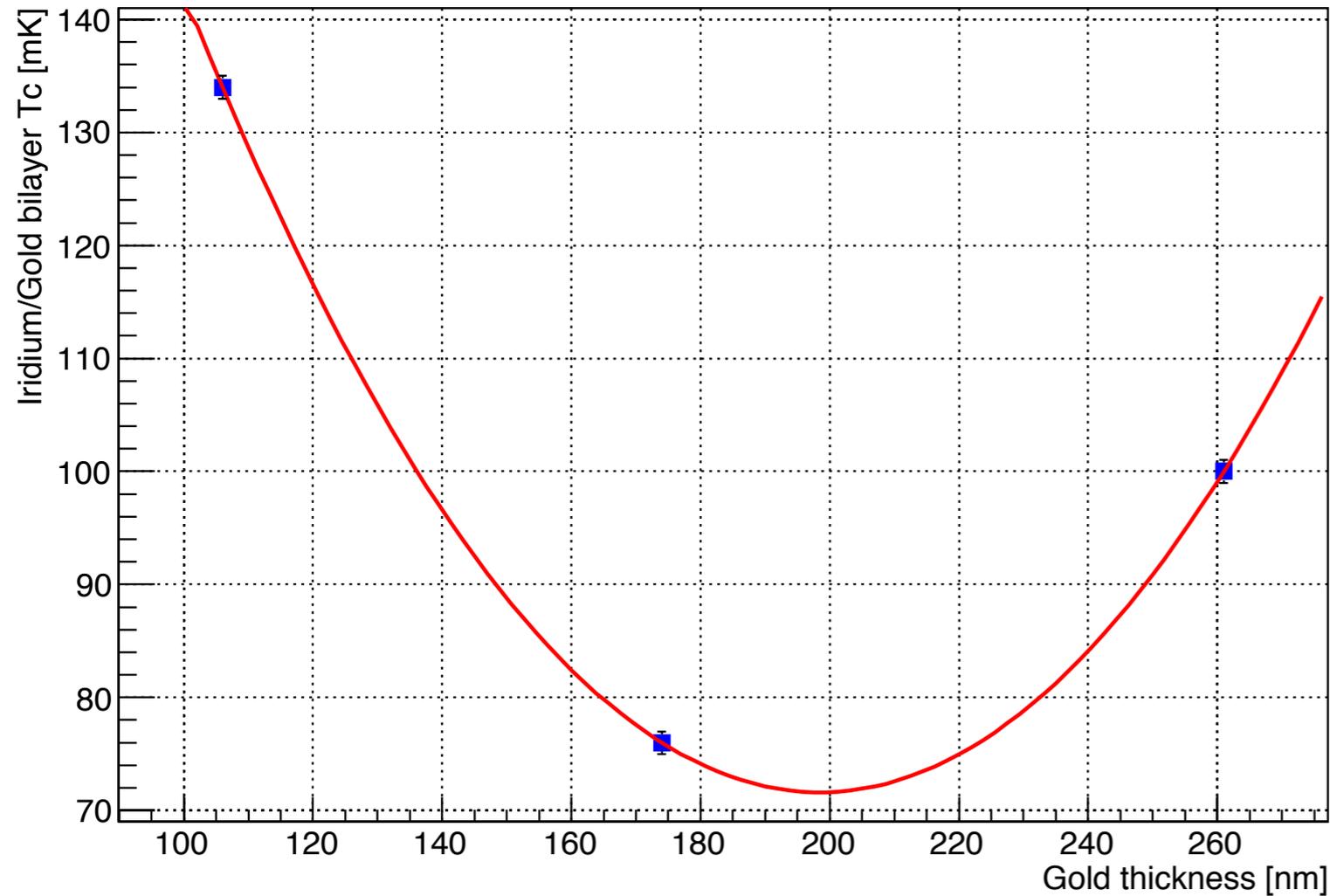
**260nm Au at room temp**



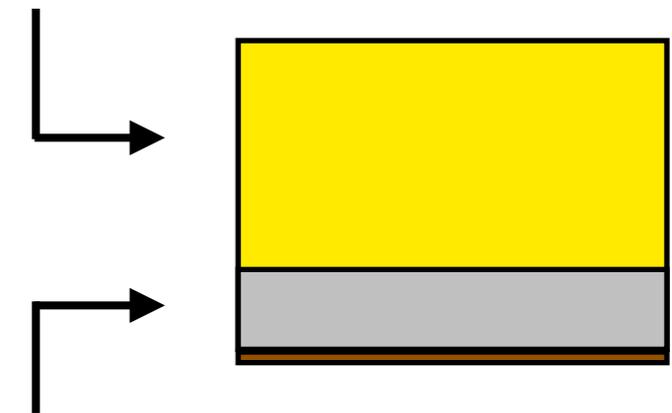
**100nm Ir annealed at different temperatures**

# Ir/Au at Room Temperature was discarded

Ir = 101nm at room temp / Au = x-axis nm

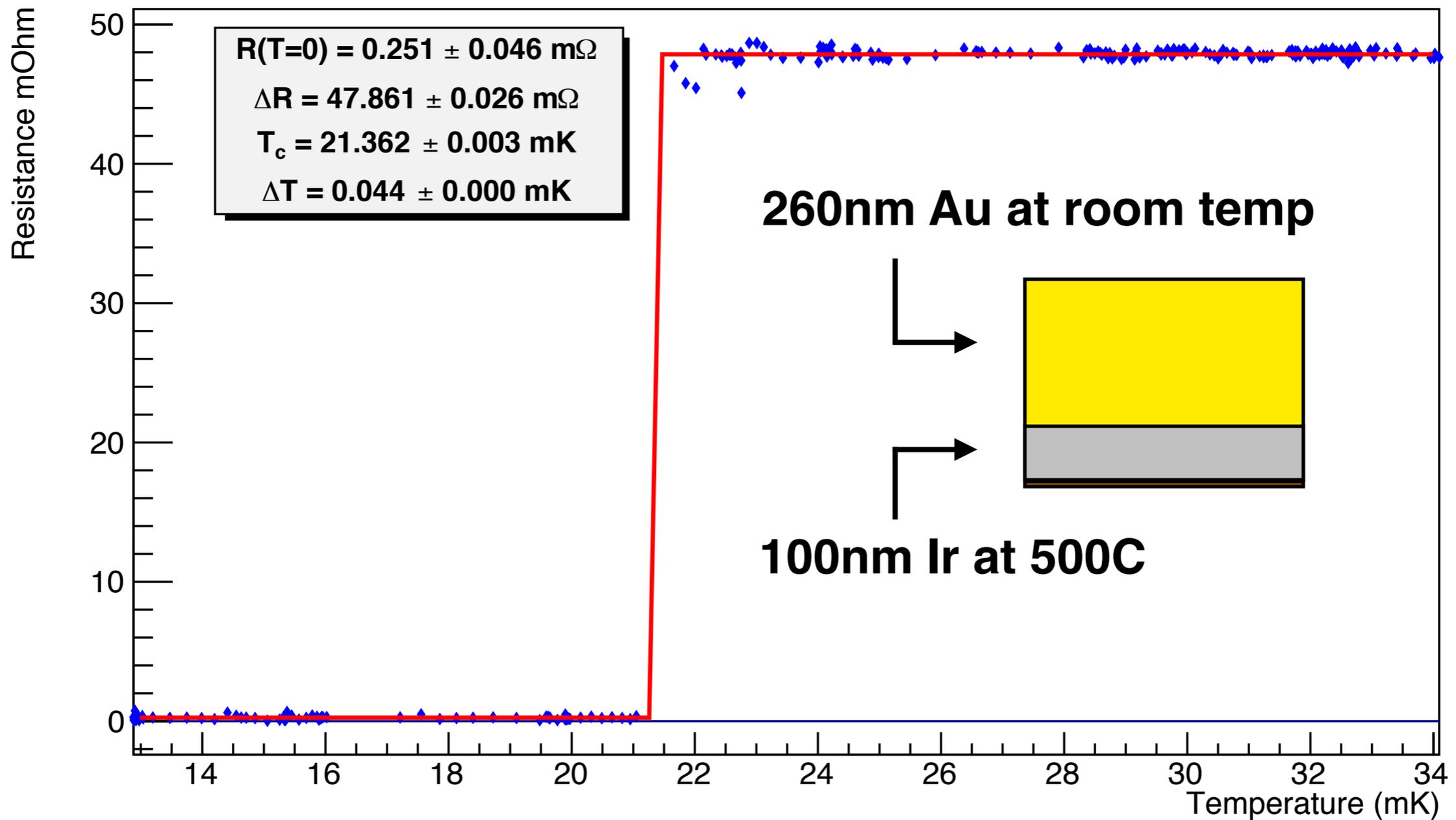


**100-260nm Gold**



**100nm Ir at room temp**

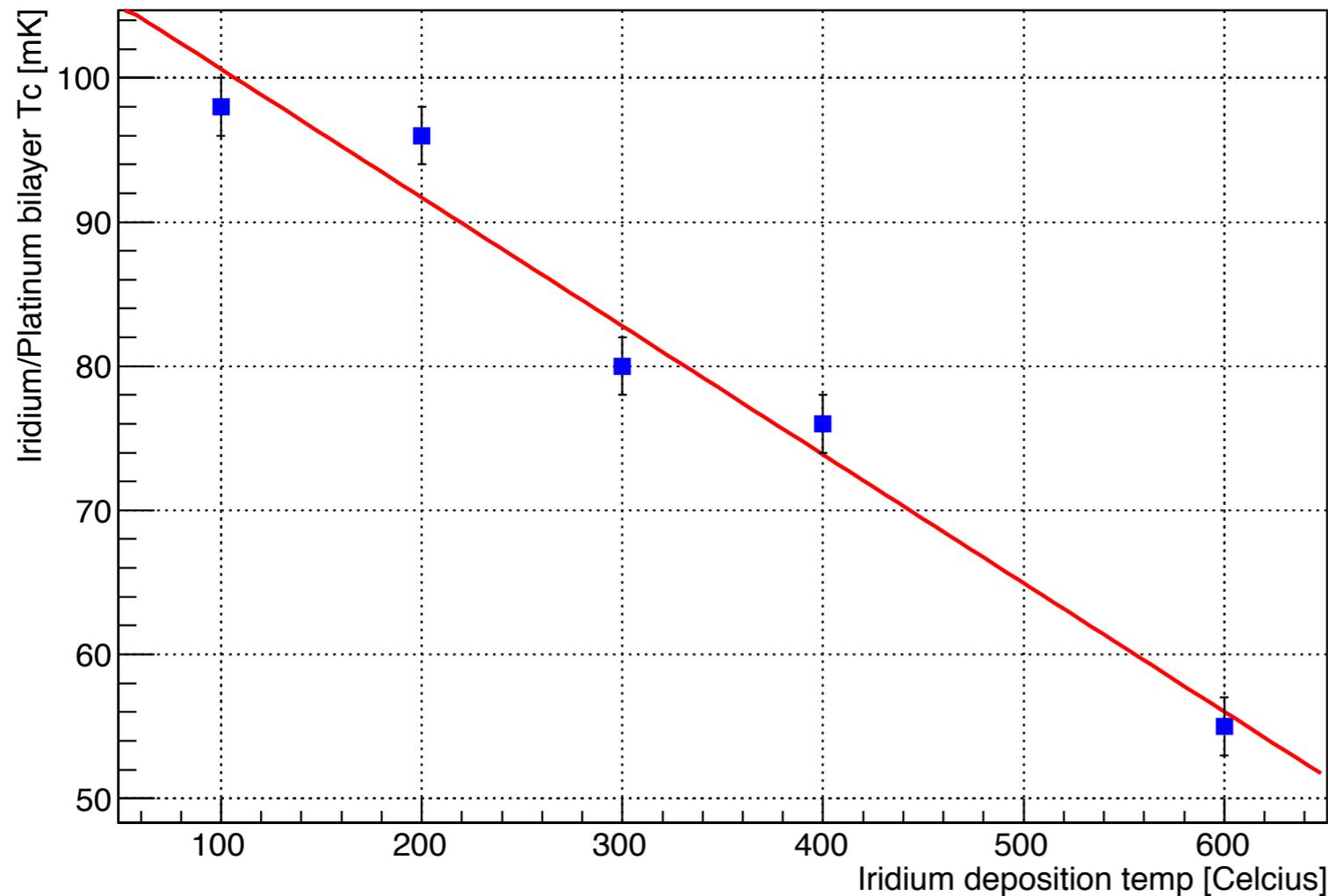
# Our lowest Ir/Au bilayer $T_c=21.4$ mK



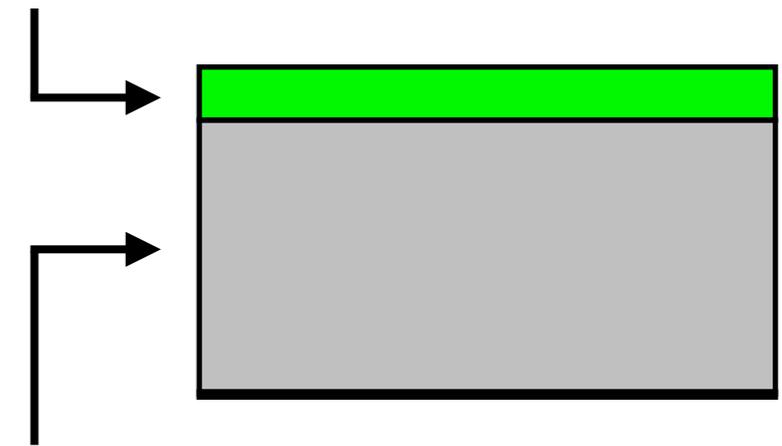
# A promising new bilayer for Low Tc TES: Ir/Pt

Tc dependence on Iridium deposition temperature for an Ir/Pt

Ir = 80nm at x-axis temp / Pt = 20



20nm Pt at room temp



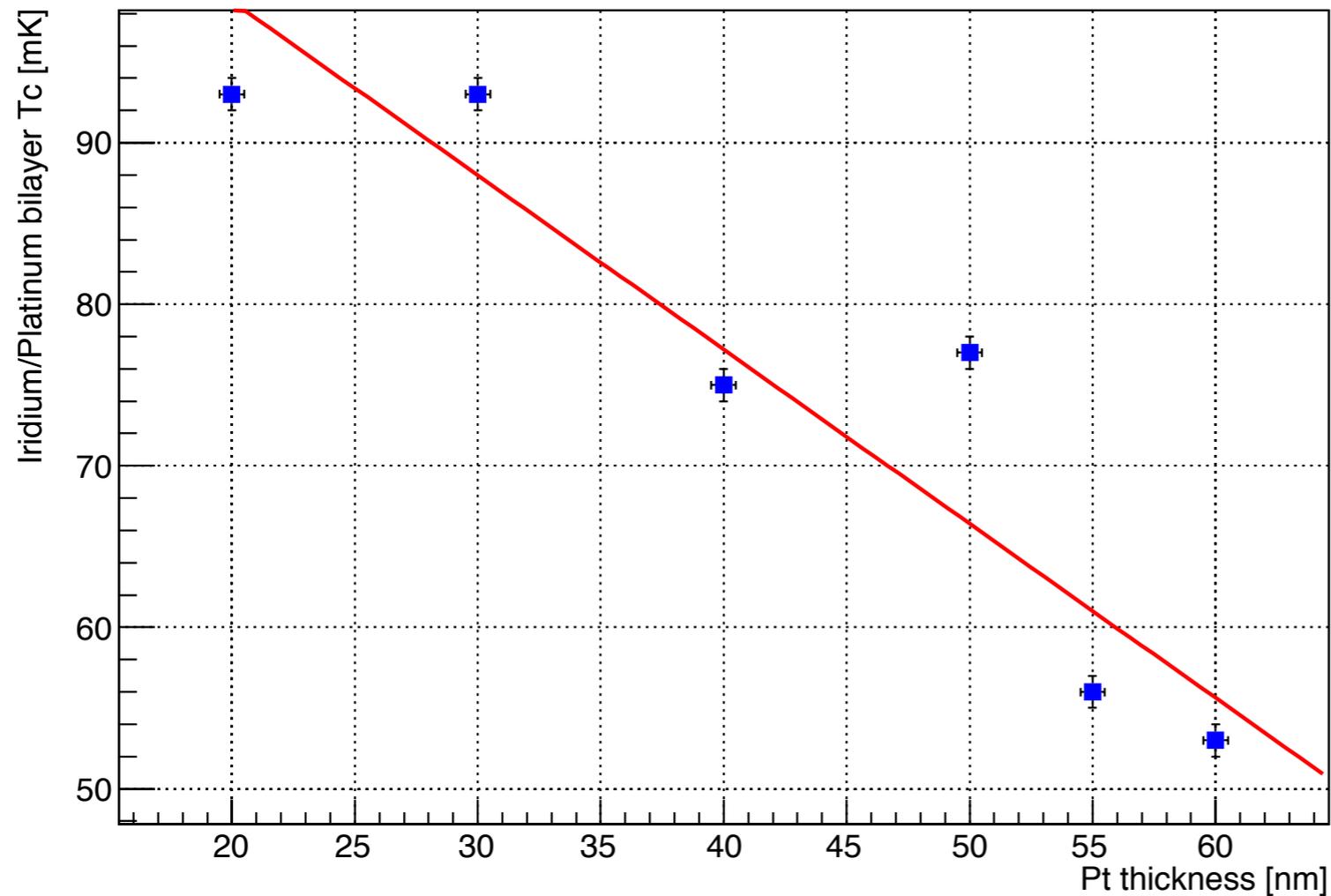
80nm Ir at varying temp

Normal resistance  $\sim 10x$  of IrAu for same Tc

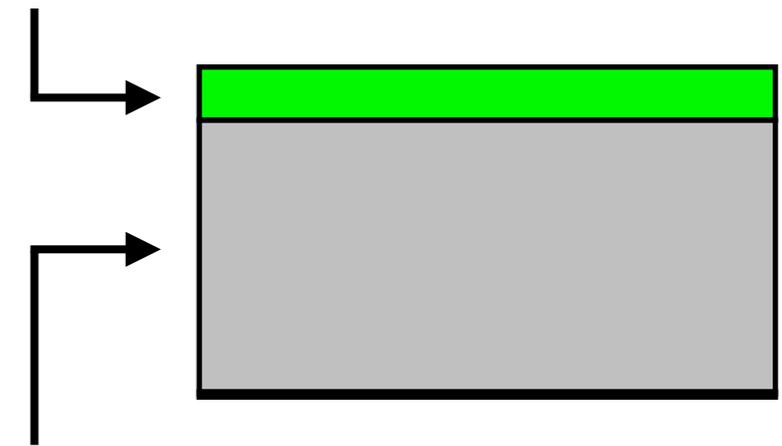
# Ir/Pt at room temperature

## Tc dependence on Platinum thickness

Ir = 100nm at Room Temp / Pt = x-axis



20nm Pt at room temp



80nm Ir at varying temp

Room temperature deposition of TES would make application to large crystals possible

# Summary

- TES technology for Double Beta Decay may be applicable in order to improve energy resolution and reduce backgrounds
- TES as light detectors for macro-bolometers in large quantities  $O(1000)$  remains to be demonstrated for the future of double beta decay with bolometers
- Low  $T_c$  material for TES with fast fabrication time may be achievable utilizing Iridium/Gold and Iridium/Platinum
- Ir/Au only works if we heat Ir and let it anneal to room temperature before depositing Au at room temperature
- We have found a new promising bilayer, Ir/Pt that works both with Ir annealing and at room temperature. It also shows large  $R(\text{normal})$  compared to Ir/Au ( $\sim 10x$ ).
- Large  $T_c$  suppression (reproducibility) remains an issue without dedicated film deposition chambers
- SQUID multiplexing may be required, although the channel number has been already accomplished by other experiments in CMB community