Fermilab: Present and Future

Young-Kee Kim Fermilab / Univ. of Chicago

January 6, 2010 Argonne National Accelerator Laboratory



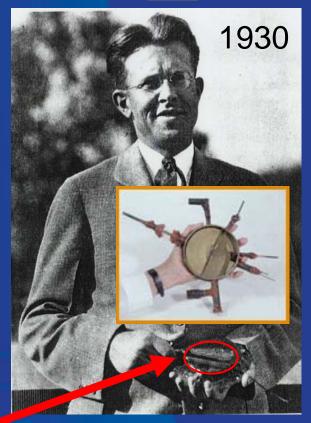
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What is the world made of? What holds the world together? Where did we come from?

the smallest things in the world interactions (forces) between them the Universe's past, present, and future

> Particle Physics: physics where small and big things meet, inner and outer space meet

Many generations of Accelerators created with higher and higher energy and intensity beams

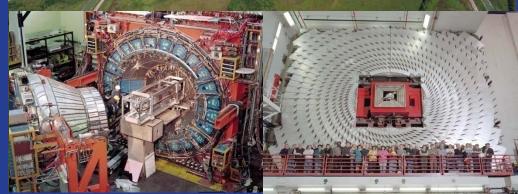


Ernest Lawrence (1901 - 1958)

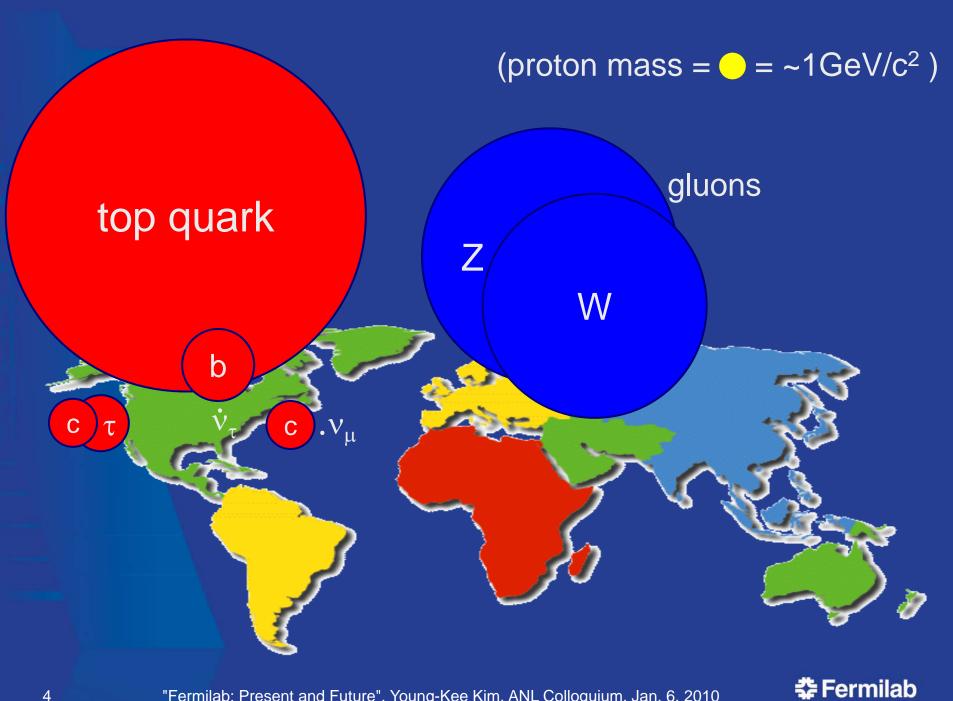
~2000 Scientists Fermilab experiments using accelerators > 2 publications every week ~2 Ph.D.s every week

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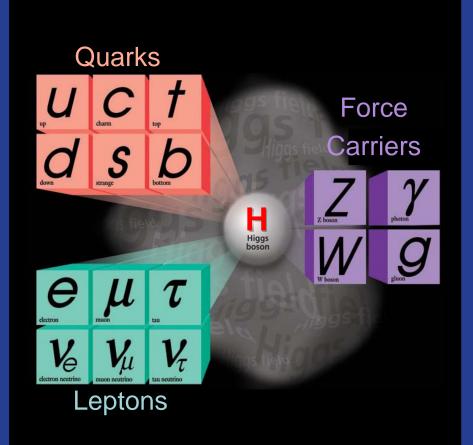
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Tevatron: x10⁴ bigger, x10⁶ higher energy Intense neutrino beams



- Present theory (Standard Model) is a remarkable intellectual construction
- Particle experiments ever done fits in the framework
- But huge questions remain unanswered.
- New physics is required to answer: e.g.
 - Supersym. extension of SM, extra dimensions,



Higgs yet to be discovered



1. What is the origin of mass for fundamental particles? 2/Why are there so many kinds of particles? 3. Do all the forces become one? 4. Are there extra dimensions of space? 5. What are neutrinos telling us? 6. What happened to the antimatter? 7. What is dark matter? 8. How can we solve the mystery of dark energy? 9. How did the universe come to be? 10. Are there undiscovered principles of natur new symmetries, new physical laws?

Evolved Thinker

Accelerators

Matter

History of the Universe

Create particles/antiparticles that existed ~0.001 ns after Big Bang.

- Direct: Tevatron, LHC,
- Indirect: v's, Rare Processes, ...

1 thousand million years

300 thousand years

3 minutes

10*34 seconds

10"10 seconds

10"43 seconds

Unification, New Symmetry Towards simple, elegant, complete theory

How did the universe come to be?

10³² degrees

Matter Antimatter

www.radiation

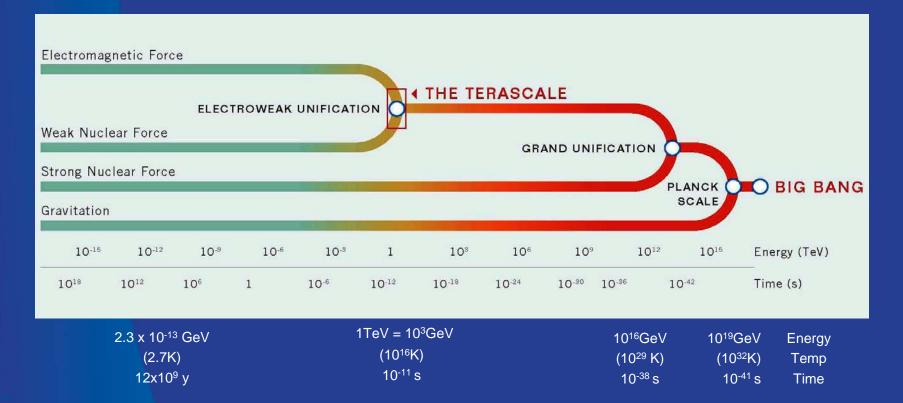
particles

heavy particles carrying the weak force positron (anti-electron)

neutron meson

meson

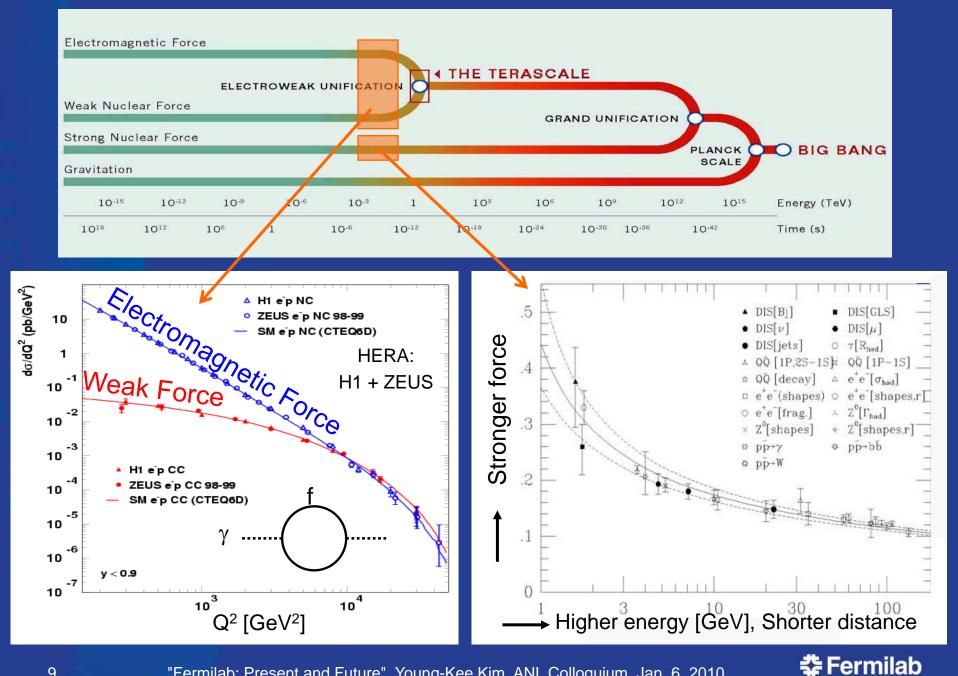
Unification

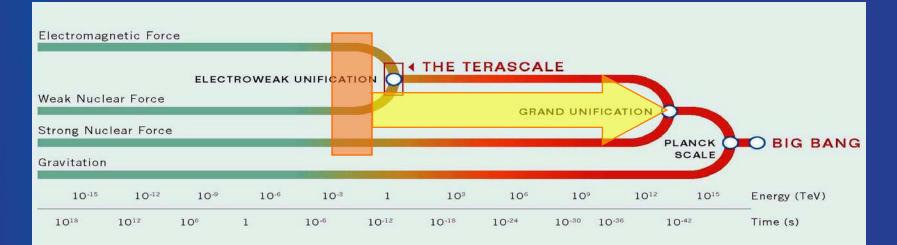


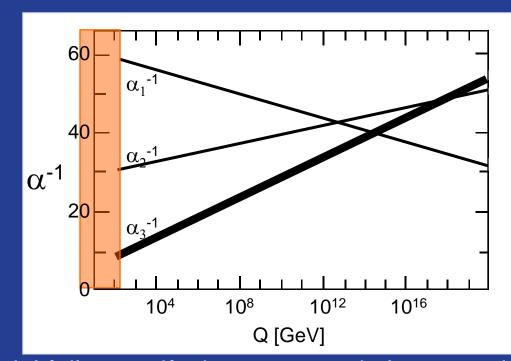
We want to believe that there was just one force after the Big Bang.

As the universe cooled down, the single force split into the four that we know today.







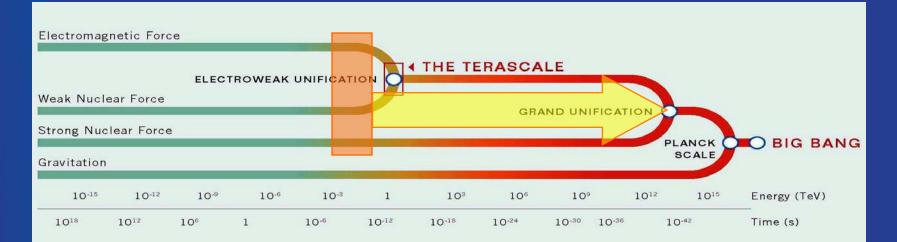


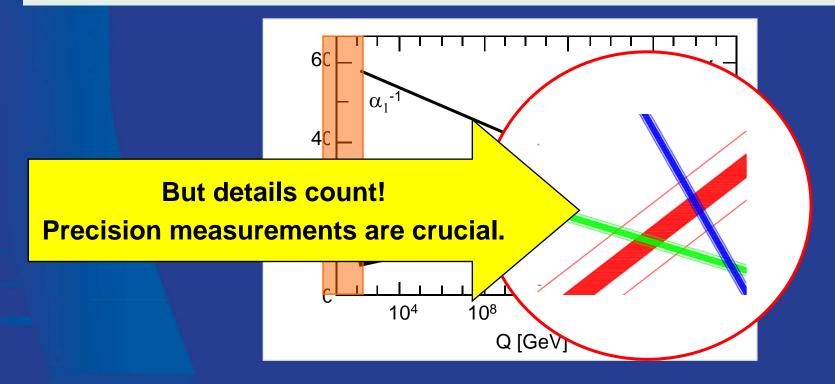
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The Standard Model fails to unify the strong and electroweak forces.

"Fermilab: Present and Future", Young-Kee Kim, ANL Colloquium, Jan. 6, 2010

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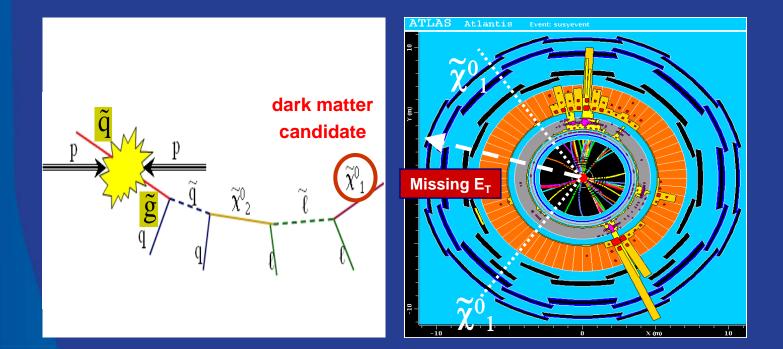






Early discovery at LHC

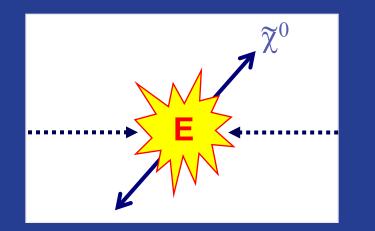
LHC could discover strongly coupled SUSY

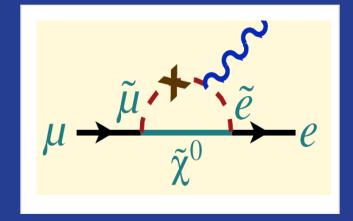


A host of new particles: fit roughly some masses, make assumption on couplings



Energy Frontier





The Intensity Frontier can probe new physics at a scale >> TeV.

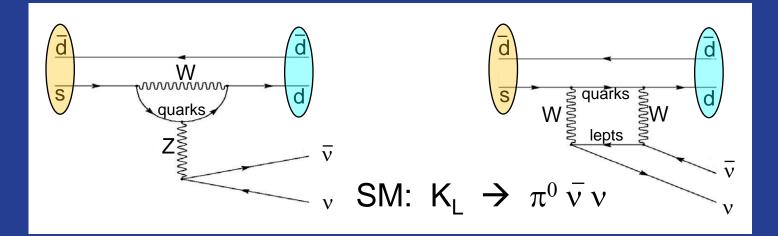
Muon to electron conversion: $\mu N \rightarrow eN$

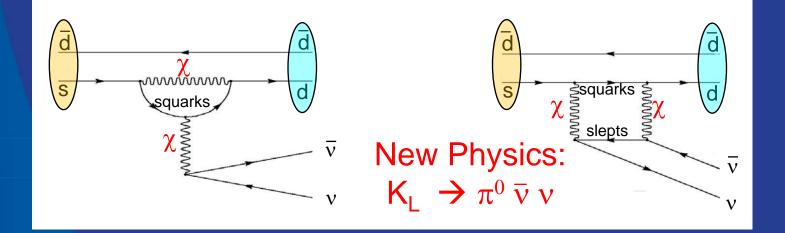
Neutrinos change from one kind to another. Do charged leptons do, too?





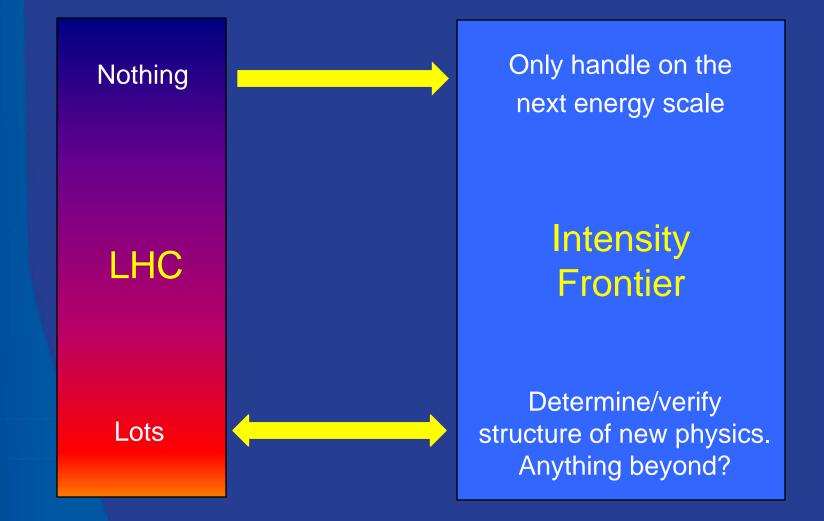
Intensity: Large effects in kaon decay rates







Interplay:Energy Frontier Intensity Frontier



‡Fermilab

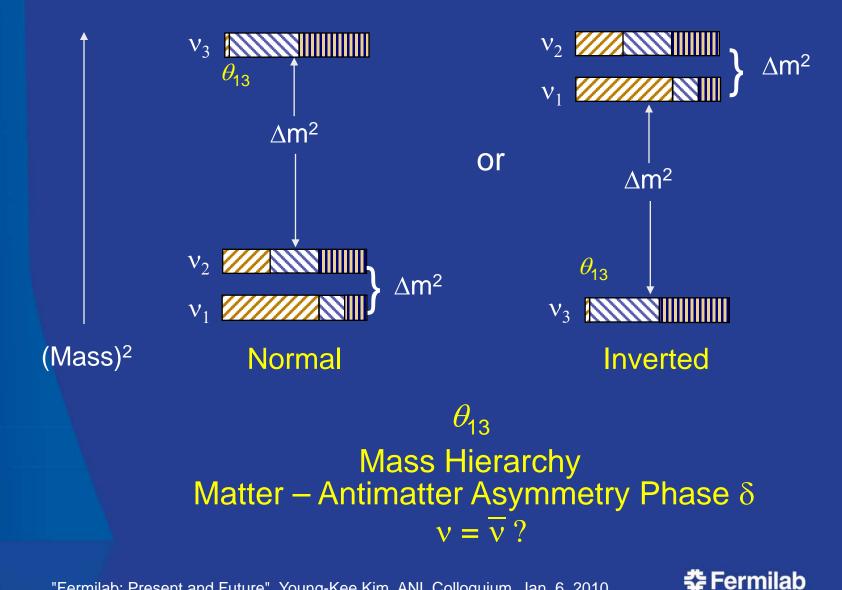
Intensity Frontier: Neutrinos

- Recent Discoveries
 - produced much excitement.
 - the only new physics seen so far in the lab.
- Behave so different from other particles
 - Mass, Oscillation pattern, $v = \overline{v}$ possibility
- A Matter-Dominate Universe
 - Require Matter-Antimatter Asymmetry (CP Violation)
 - Quarks can not explain. Maybe the leptons can.
- Unification
 - v mass, mixing point toward new symmetries (unification)
- Cosmic Connection
 - $\sim -10^9$ neutrinos / nucleon or electron in the Universe.
 - Neutrino mass affects large scale structure.





The neutrino spectrum: unknowns

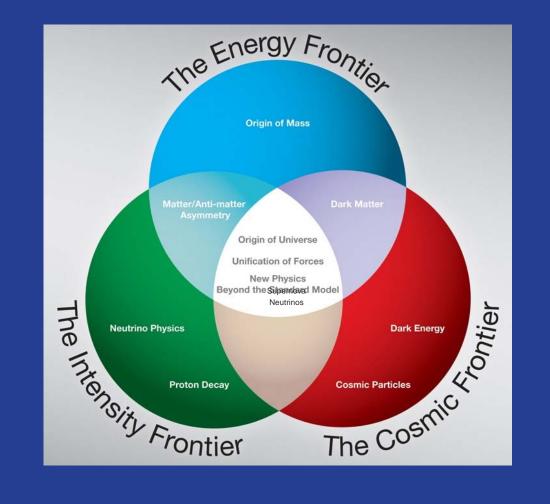


"Fermilab: Present and Future", Young-Kee Kim, ANL Colloquium, Jan. 6, 2010

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Particle Physics at the Three Frontiers

Endorsed by the US Particle Physics Community





Particle Physics

- Global enterprise
- Many laboratories have changed missions. A few principle particle physics laboratories in the world
- Important and healthy to maintain expertise, long term stability, and support in all three regions, and to engage the world wide community
- More coordination and collaboration



US Particle Physics Today

National Laboratories

- · Fermilab
 - Single mission particle physics
- Other laboratories: ANL, BNL, LANL, LBNL, SNAL, ...
 - . Multi missions including particle physics
 - . Particle physics is not the primary mission
- Universities
- We need to maintain expertise and uniqueness in laboratories and universities



Fermilab Programs at Three Frontiers (Now)

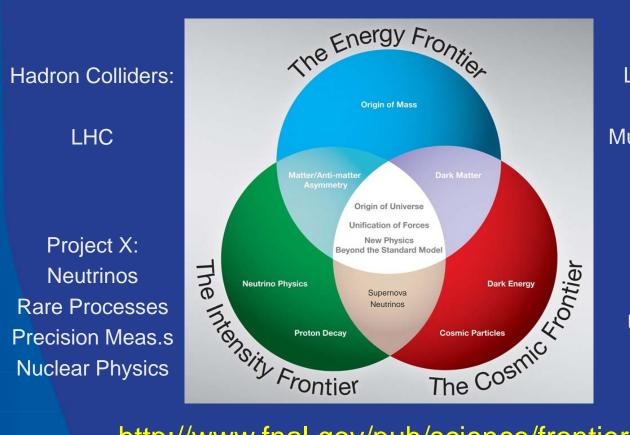


Dark Matter, Dark Energy, UHE Particles from Space

http://www.fnal.gov/pub/science/frontiers/



Fermilab Programs at Three Frontiers (Future)



Lepton Colliders: Sub-TeV: ILC Multi-TeV: µ Collider (CLIC)

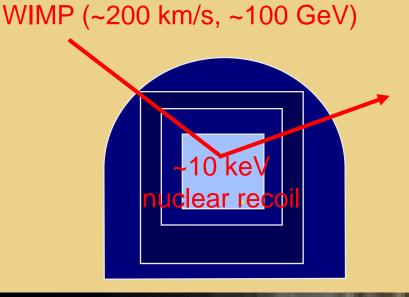
Dark Matter, Dark Energy, UHE Particles from Space

http://www.fnal.gov/pub/science/frontiers/



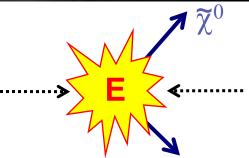
Cosmic Frontier: Dark Matter

Underground experiments may detect Dark Matter candidates.

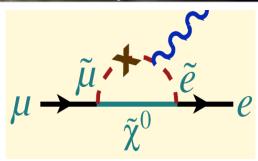


Cosmic Frontier

Energy Frontier

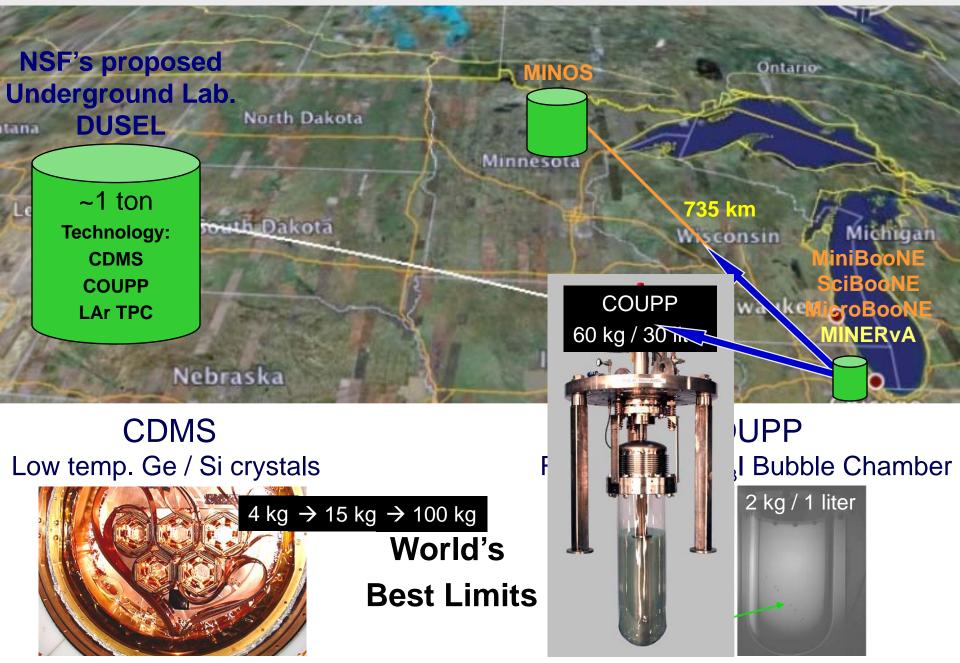


Intensity Frontier

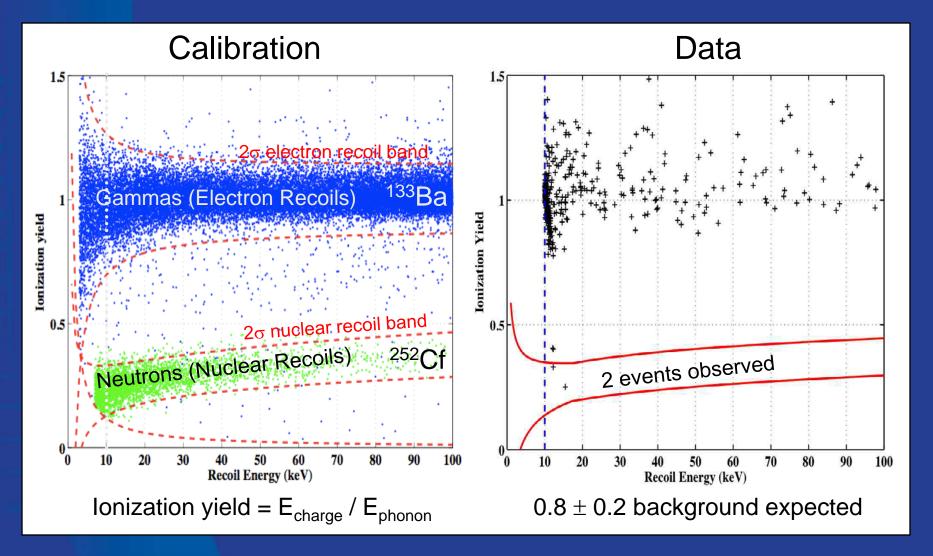


Accelerators can produce dark matter in the laboratory and understand exactly what it is.

Dark Matter Searches – Underground Detectors

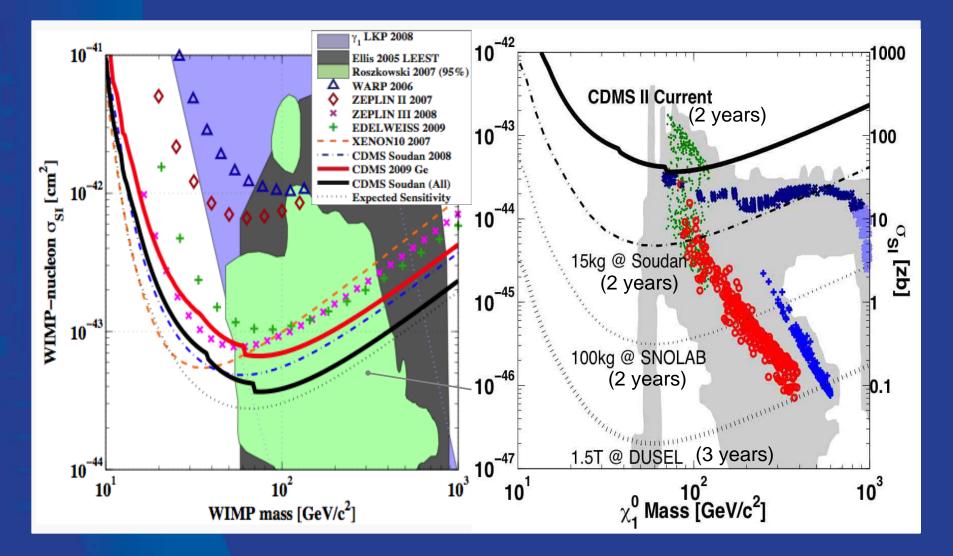


CDMS (4 kg) Results





90% CL Limit: Present and Future



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Cosmic Frontier: Probing Dark Energy

1. SDSS (Sloan Digital Sky Survey)

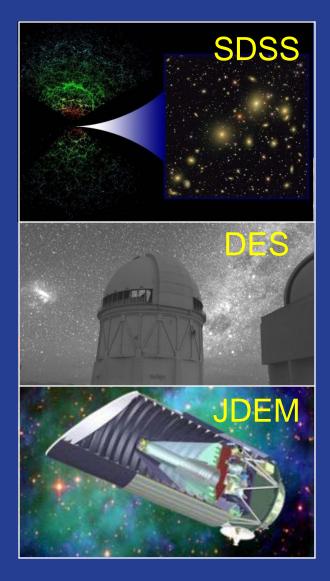
- 2.5 meter telescope in New Mexico
- Ranks as the facility with the highest impact in astronomy for the 3rd year in a row.
- Power spectrum of galaxies constrain dark energy density parameter.

2. DES (Dark Energy Survey)

- 4 meter telescope in Chile
- DES Camera under construction
- Operation: 2011 2016

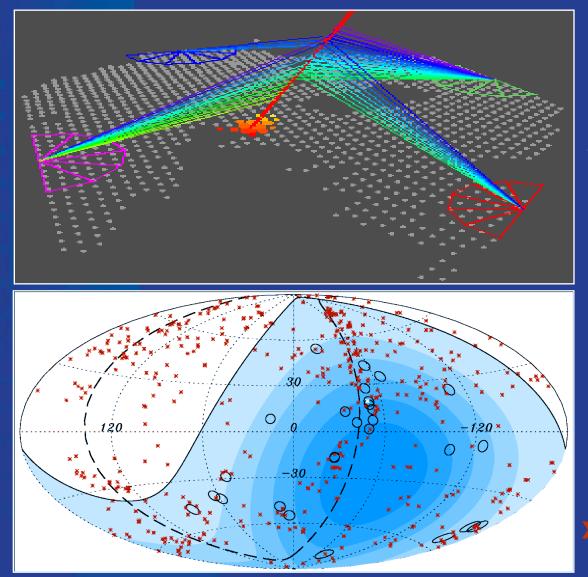
3. JDEM (Joint Dark Energy Mission)

- . Space telescope
- Fermilab Goal: Science Operation Center





Cosmic Frontier: High Energy Particles from Space



Auger Observatory studies ultra-high energy cosmic rays.

o – Cosmic rays with E > 57,000,000 TeV

Correlation

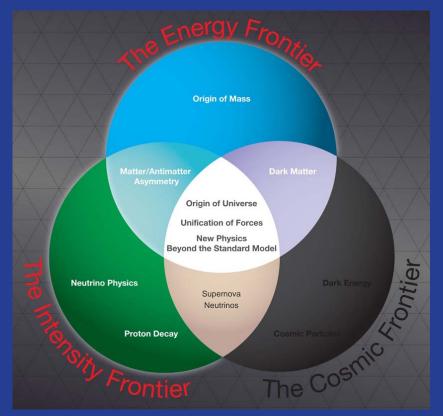
x – Active Galactic Nuclei

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Energy and Intensity Frontiers

Energy-Intensity Integrated Program

Today: operates Tevatron & highest power v beams



Future: continue to have the integrated plan

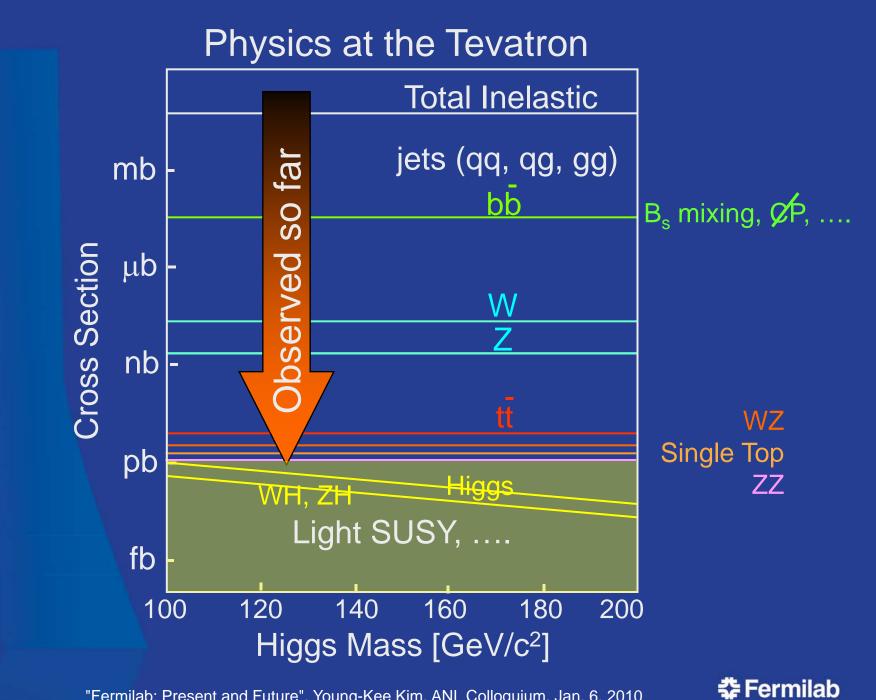


The Energy Frontier: The Tevatron (1985 – 2011)

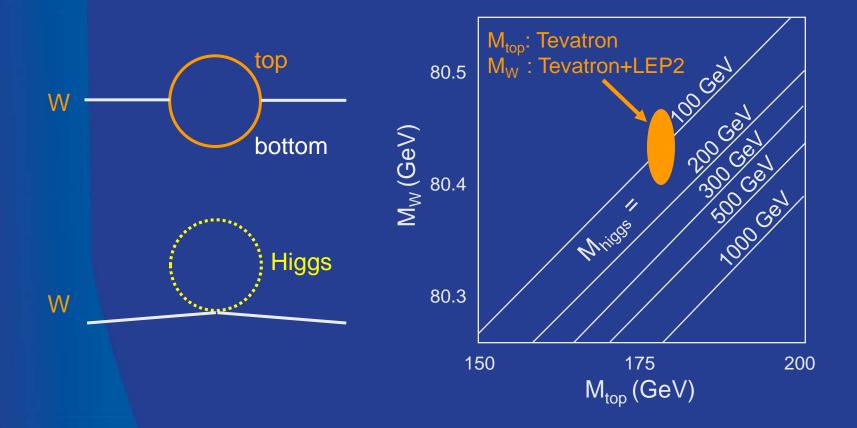
Progress: the Tevatron 7,500 FY2010 7,000 **Total Integrated Luminosity** 6,500 7.5 fb⁻¹ (a 6,000 FY2009 5,500 5,000 Luminosit 4,500 FY2008 4,000 3,500 ntegrated 3,000 FY2007 2,500 2,000 FY2006 1,500 FY2005 1,000 FY2002 FY2003 FY2004 500 2002 2005 2006 2007 2009 2010 2003 2004 2008 **Fiscal Year**

DZero

~100 publications / year, ~60 Ph.D.s / year Plan to run through FY2011: 7.5 fb⁻¹ (now) \rightarrow 12 fb⁻¹



The Tevatron Predicts Higgs Mass via Quantum Corrections



 $m_{H} = 87^{+35}_{-26} \text{ GeV} (m_{top} = 173.1 \pm 1.3 \text{ GeV})$

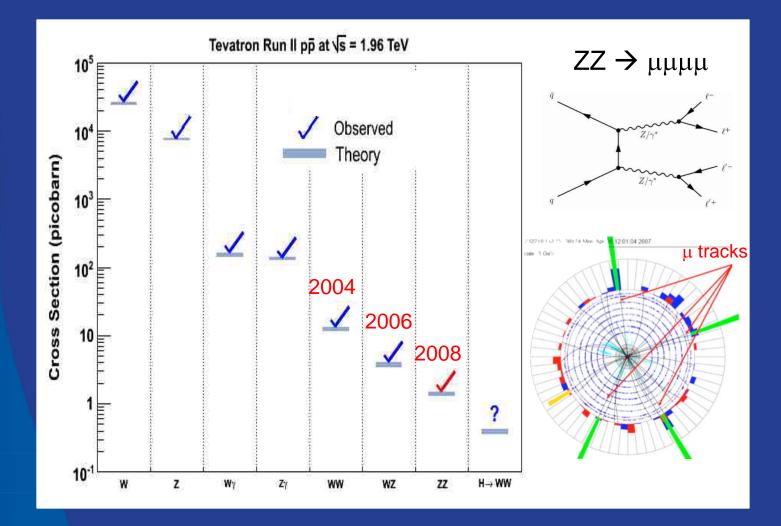
Favors "light" Higgs in the range where Tevatron has good potential.

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Observations: rare SM processes



Diboson production: more luminosity allows access to smaller cross sections



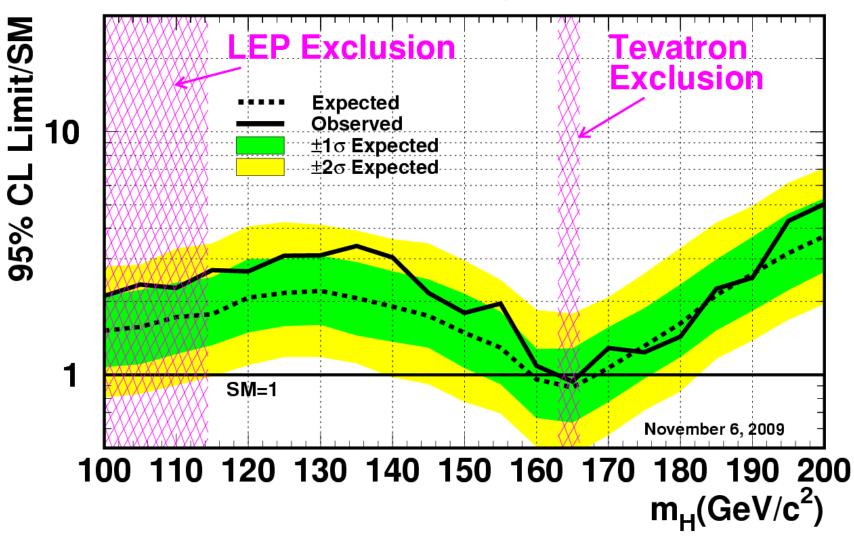
The Higgs Search

- The SM Higgs (if it exists) is being produced NOW at the Tevatron! We have enough energy
 - . Just not that often & it's buried in "backgrounds"
 - It's a story of luminosity, passion, persistence and luck
 - We know how to look for it and we are in fact closing in!
- Over the last years, there's been a dramatic infusion of people, effort and ideas, aimed at finding the Higgs





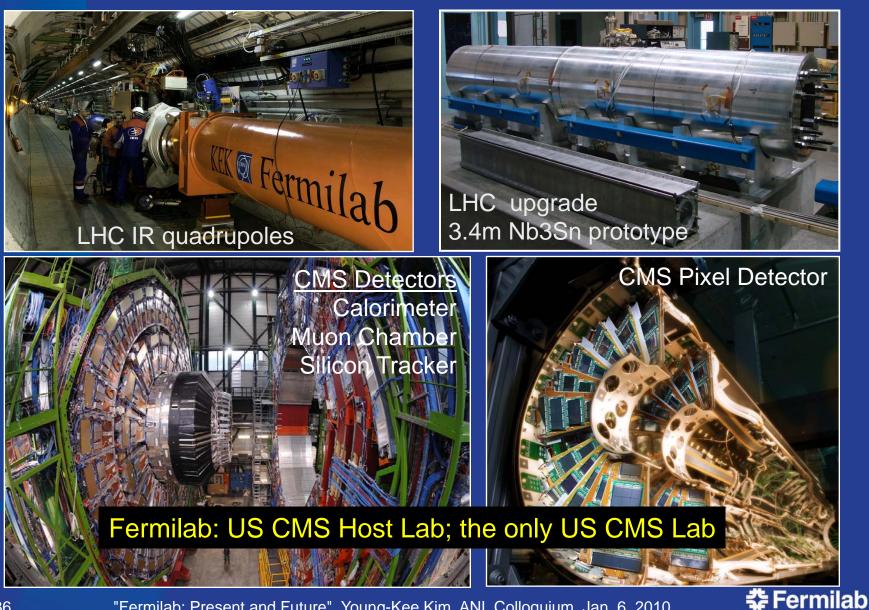
Tevatron Run II Preliminary, L=2.0-5.4 fb⁻¹





Fermilab and LHC:

Accelerator and Detector Design/Engineering/Construction and Upgrades





US CMS Host Lab; the only US CMS Lab

CMS Tier-1 Computing Center LHC Physics Center Support US CMS Community

Remote Operation Center (ROC): Detector Commissioning and Monitoring Accelerator Monitoring CERN Night = FNAL Day

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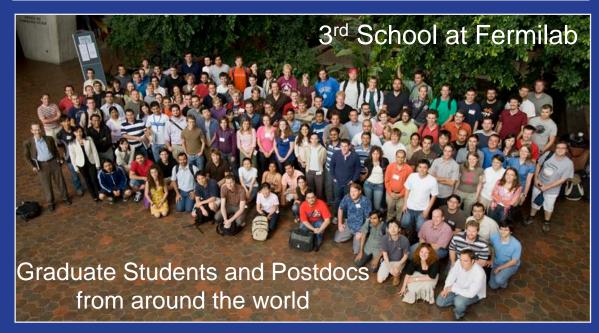
To make being at Fermilab as good as being at CERN. Requires critical mass (~100 Fermilab + University Scientists at Fermilab).

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Supporting the LHC Community

CERN-Fermilab Hadron Collider Physics Summer School

1 st	Fermilab	August 9-18, 2006
2 nd	CERN	June 6-15, 2007
3 rd	Fermilab	August 12-22, 2008
4 th	CERN	June 8-17, 2009
5 th	Fermilab	Summer 2010

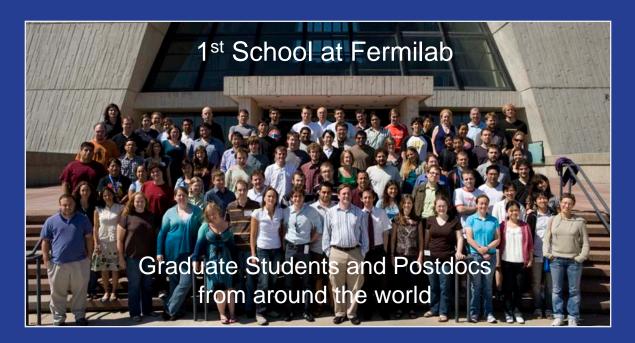




International Neutrino Summer School

Merging various neutrino schools into one coherent school Rotating in three regions

1 st	Fermilab	July 6-18, 2009
2 nd	KEK	2010
3 rd	Europe	2011





The Intensity Frontier: Neutrinos

260 kW 120 GeV MI protons & 8 kW 8 GeV Booster protons run simultaneously with the Tevatron

120 GeV Main Injector

ciBooNE

8 GeV

Booster

MiniBooNE Detector



Neutrino beam from 8 GeV Booster MiniBooNE: Excludes "4th gen." v

Tevatron

Low Eng Excess in v, Now running anti-v

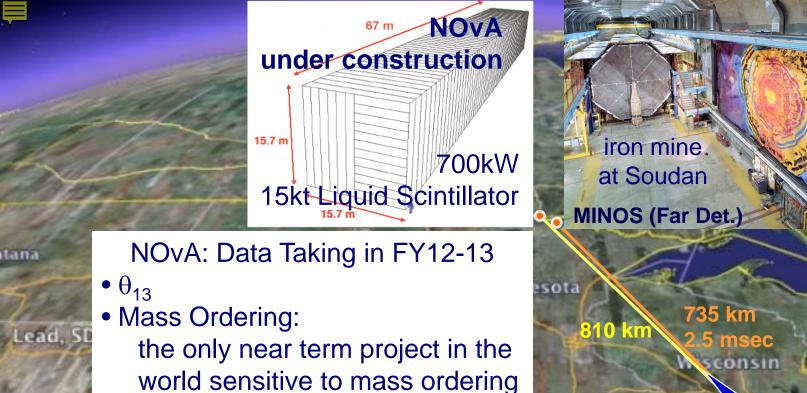
SciBooNE: v – Matter Interactions

MicroBooNE: 170 ton LAr TPC

Neutrino beam from 120 GeV MI

tana





improved precision: 2-3

Nebraska

Michigan

iniBoo ciBool OBOON MINOS (Near Det.) icago

ArgoNeuT

MNER

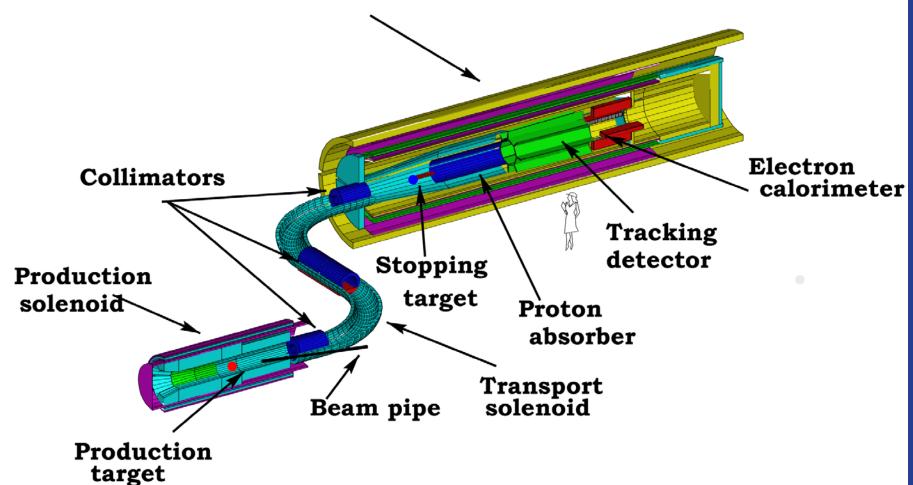
Milw

lowa

MINERvA: Ops. w/ Full Det.(2010) v - Matter Interactions

Muon to e Conversion ($\mu N \rightarrow eN$)

Detector solenoid



Muon g-2, K⁺ $\rightarrow \pi^+\nu\nu$ (1000 events) under consideration

Fermilab

Project X: intense proton accelerator

http://www.fnal.gov/pub/projectx/

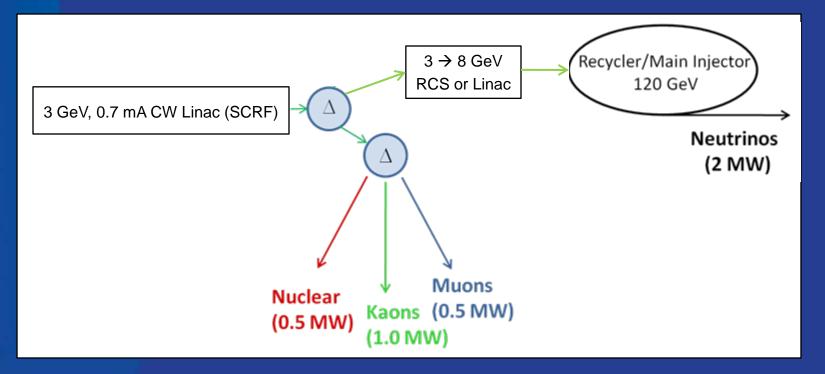
- The intensity frontier answers fundamental questions
- Project X is the key
- Project X can lead us back to the energy frontier





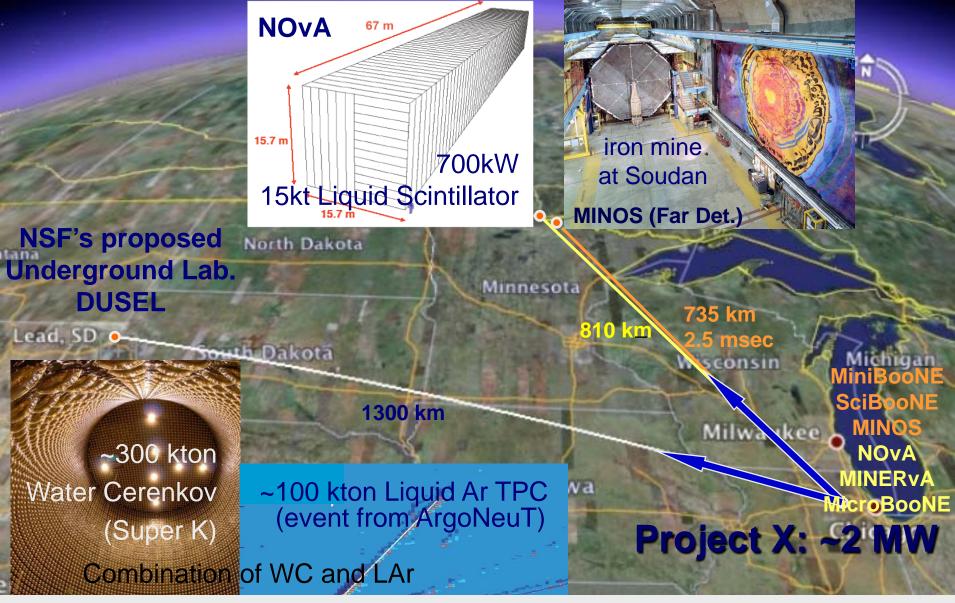
Evolution of Project X: 3 Simultaneous Beams

- 3 GeV protons, 2 MW CW (continuous pulses at 325 MHz)
 - Rare processes + precision measurements
 - flexible time patterns + pulse intensities
- 8 GeV protons, 20 200 kW: rare processes + precision meas.s
- 60 120 GeV protons, 2 MW (to Homestake) for neutrinos



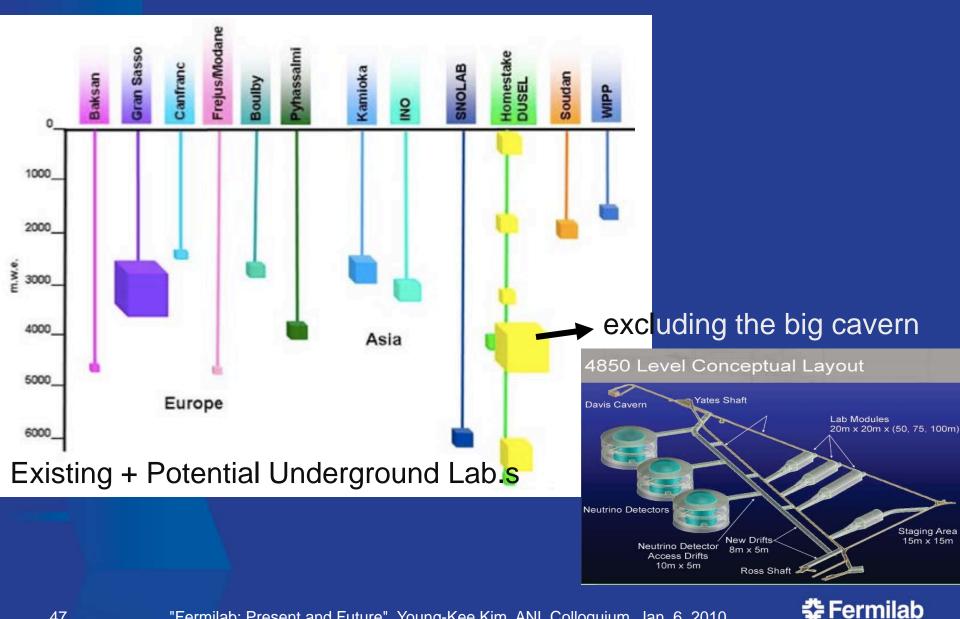


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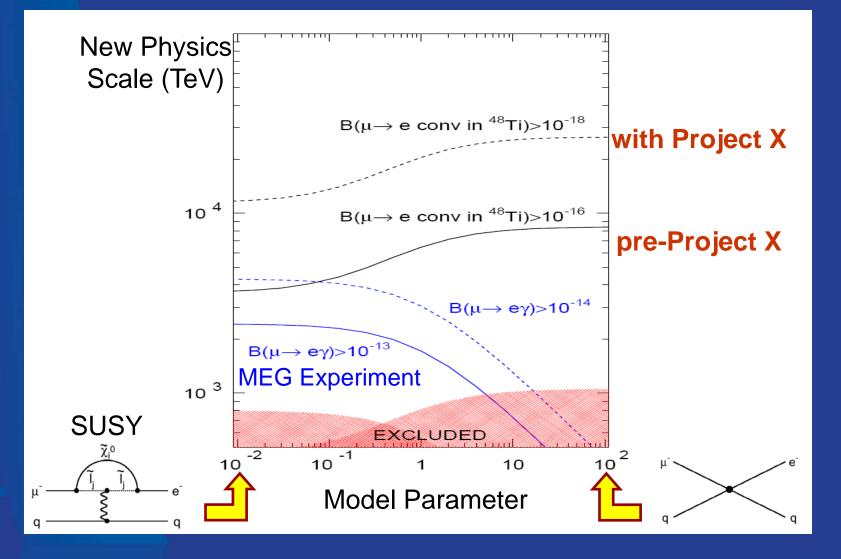
Matter – Antimatter Asymmetry with Neutrinos Proton Decay Supernovae Neutrinos

The Intensity Frontier: Fermilab \rightarrow DUSEL Option





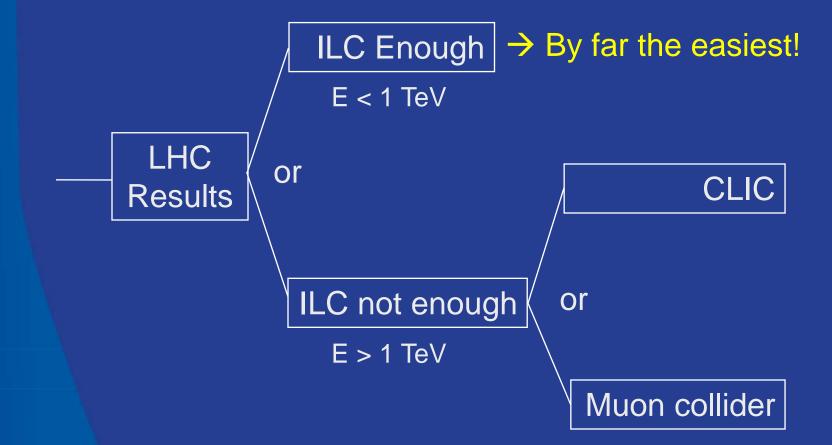
Mu2e can probe ~10⁴ TeV



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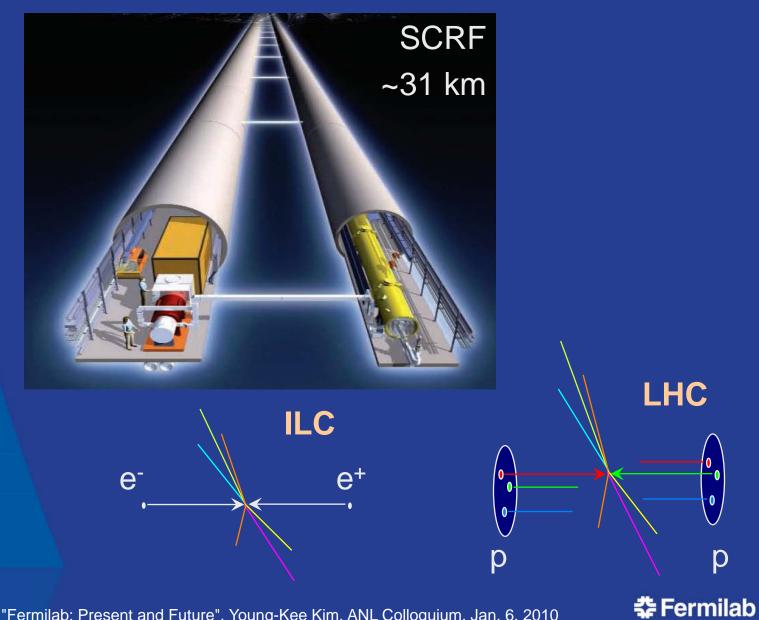
A path back to the energy frontier?

Lepton colliders beyond LHC





International Linear Collider (ILC)



Muon Collider Conceptual Layout

Project X

Accelerate hydrogen ions to 8 GeV using SRF technology.

Compressor Ring Reduce size of beam.

Target

Collisions lead to muons with energy of about 200 MeV.

Muon Cooling

Reduce the transverse motion of the muons and create a tight beam.

Initial Acceleration

In a dozen turns, accelerate muons to 20 GeV.

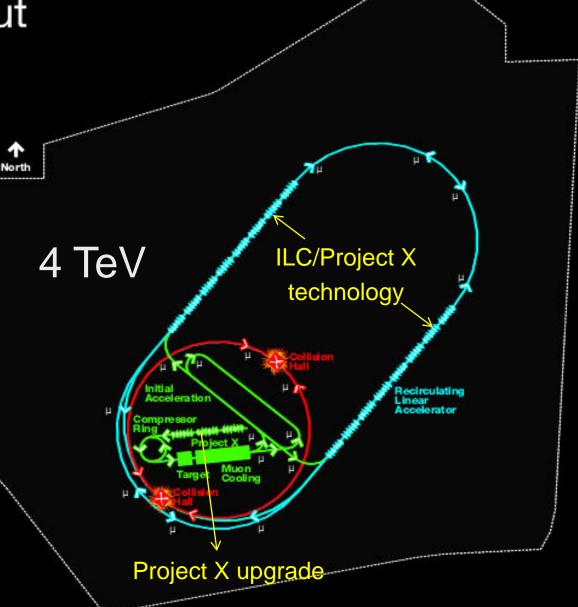
Recirculating Linear Accelerator

In a number of turns, accelerate muons up to 2 TeV using SRF technology.

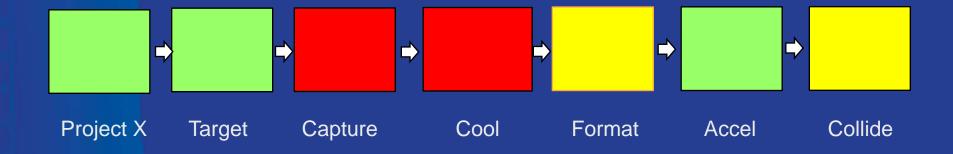
Collider Ring

Located 100 meters underground. Muons live long enough to make about 1000 turns.





Muon collider functional layout



Color indicates degree of needed R&D (difficulty) and demonstration



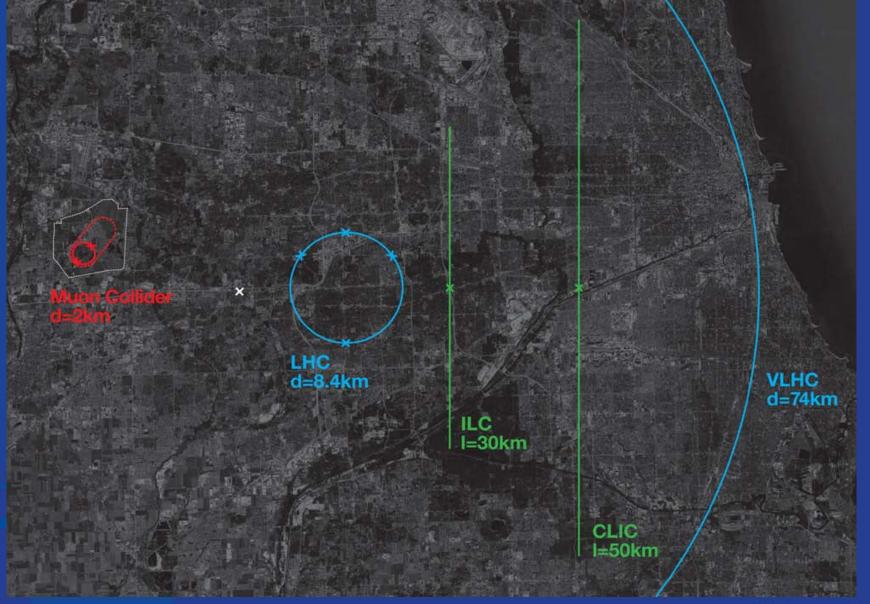
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ILC / Project X / Muon Collider technology at Fermilab



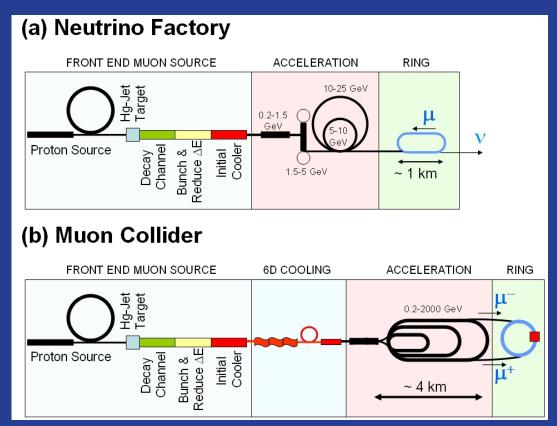
Comparison of Particle Colliders To reach higher and higher collision energies, scientists have built and proposed larger and larger machines.





Evolution of v Program: Neutrino Factory

 Muon Colliders & Neutrino Factories require similar, & potentially identical, muon sources:

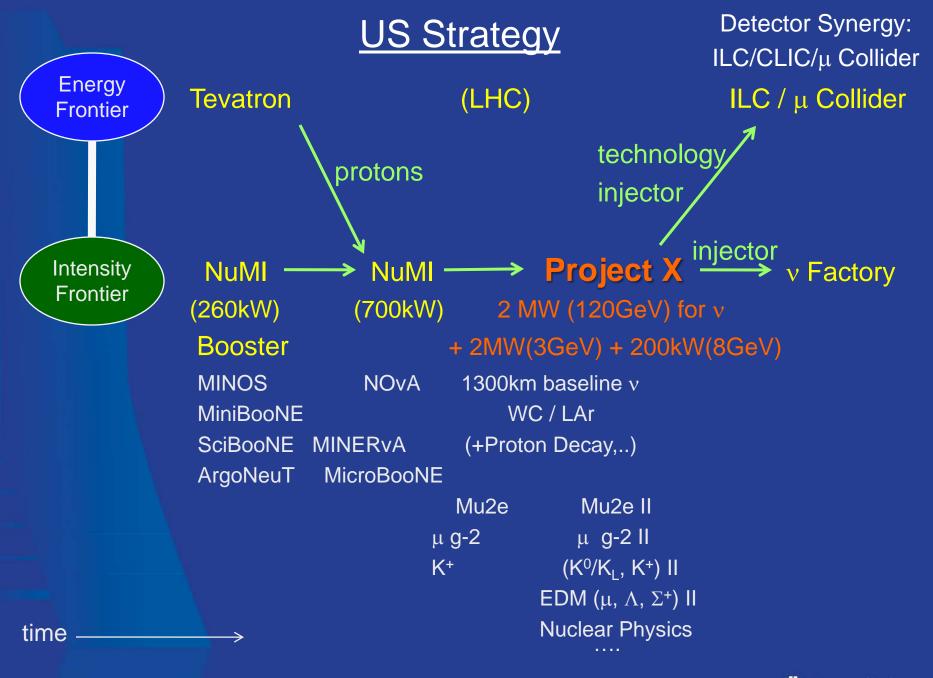




Project X

- Would be a fantastic machine at the intensity frontier for neutrino physics, kaon physics, muon physics, and nuclear physics
- Would develop and exercise the technologies to position the US to host a global facility at the energy frontier (or contribute to one elsewhere)
 - . ILC and muon collider







Closing Remarks

- Compelling Questions in Particle Physics
 - Require three interrelated frontiers
 - . The Energy Frontier
 - . The Intensity Frontier
 - . The Cosmic Frontier



• Fermilab: Current and Future

- . A balanced program at 3 interrelated frontiers
- Project X (intense proton source)
 - . Intensity Frontier Facility (broad physics program)
 - . A path back to the Energy Frontier
 - ILC technology
 - Front end of a muon collider (and/or v factory),
 - Acceleration technology for a muon collider



END

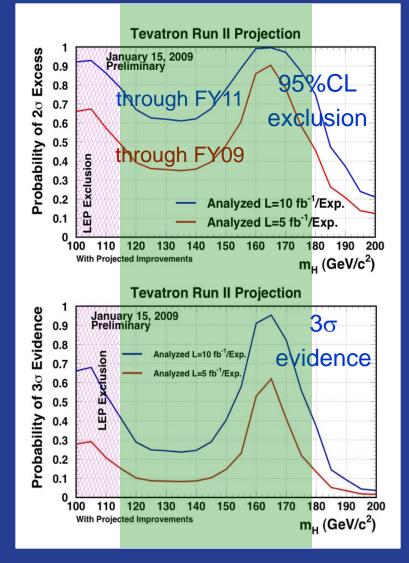


and some luck

Favored mass region

Higgs reach with continued analysis improvement

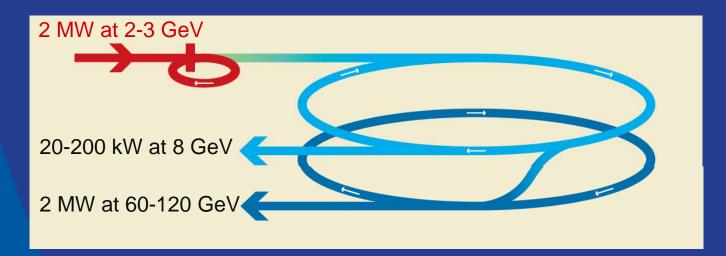
running through FY09 (red) FY11(blue)



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Evolution of Project X: 3 Simultaneous Beams

- 2 MW CW (continuous pulses at 325 MHz) 2-3 GeV protons rare processes and precision measurements flexible time patterns and pulse intensities
 - 20 200 kW 8 GeV protons rare processes and precision measurements
- 2 MW 60 120 GeV protons (to Homestake) for neutrinos



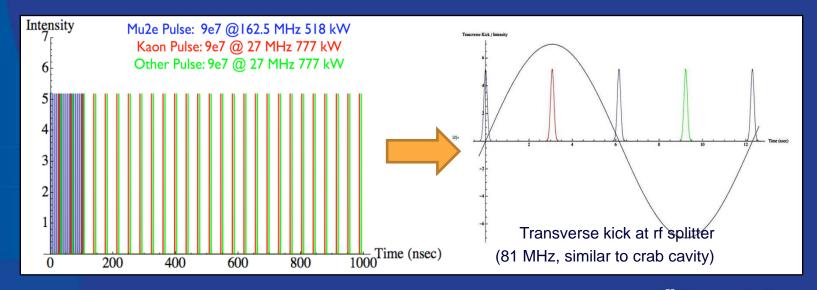
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Flexible bunch format

- Variable H⁻ ion source provides current 1 to 10 mA DC
- Variable bunch formats:
 - Ion source at 1 mA, no beam chopping: 1.9x10⁷ protons per bunch at 325 MHz rate
 - Ion source at 10 mA, 90% beam chopping: 1.9x10⁸ protons per bunch at 32.5 MHz rate (1 mA ave current)
 - Bunch-by-bunch chopping example (ion source at 4.7 mA), chopping and rf splitting for 3 experiments



Fermilab

Other applications

- Nuclear Physics
 - Can drive an ISOL target for Nuclear Physics applications. Totally complementary program for nuclear EDMs and fundamental experiments on atomic traps just with ISOL target
- Muon Spin Rotation
 - Currently done in Rikken, PSI and TRIUMF
 - Would produce the most intense muon beams available, including, polarization and monochromatization



Muon experiments

- Next generation μ→e conversion experiment, new techniques for higher sensitivity and/or other nuclei.
- μ→3e
- Next generation (g-2) if motivated by theory, next round, LHC
- Other:

 μ edm. $\mu^+e^- \rightarrow \mu^-e^+$ $\mu^-A \rightarrow \mu^+A'$

Systematic study of radiative μ capture on nuclei.

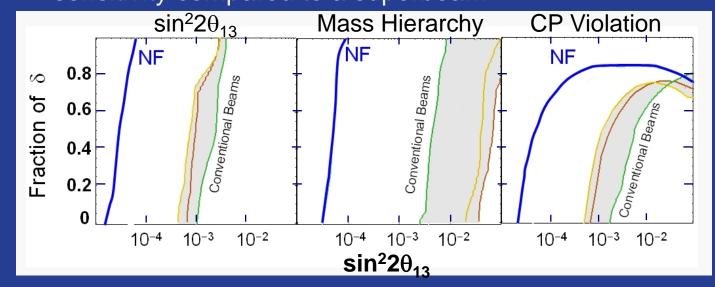
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Evolution of v Program: Neutrino Factory

International Design Study

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- If $\sin^2 2\theta_{13}$ is small
 - Choose a NF energy of 25 GeV & a very long baseline (e.g. ~3000km) – up to ~ x100 improvement in sensitivity compared to a superbeam



- If θ_{13} is large (>.005)
 - A 4 GeV NF aimed at Homestake gives clean reach into CP violation, mass hierarchy and any unusual features

Multi-TeV Lepton Colliders

- Muon Collider Approach: Fermilab's Focus
 - Based on a secondary beam: we have experience basing colliders on antiprotons. For μ 's we must do it in 20 msec.
 - Advantages: narrow energy spread (no beamstrahlung) and small physical footprint (no synchrotron radiation)
 - No new methods of acceleration, but new method of deceleration!: muon cooling
- CLIC Approach: CERN's Focus
 - Advantages: stable particles, polarization
 - Two-beam accelerator scheme
- Physics/detector
 - ILC-CLIC-Muon Collider Synergy



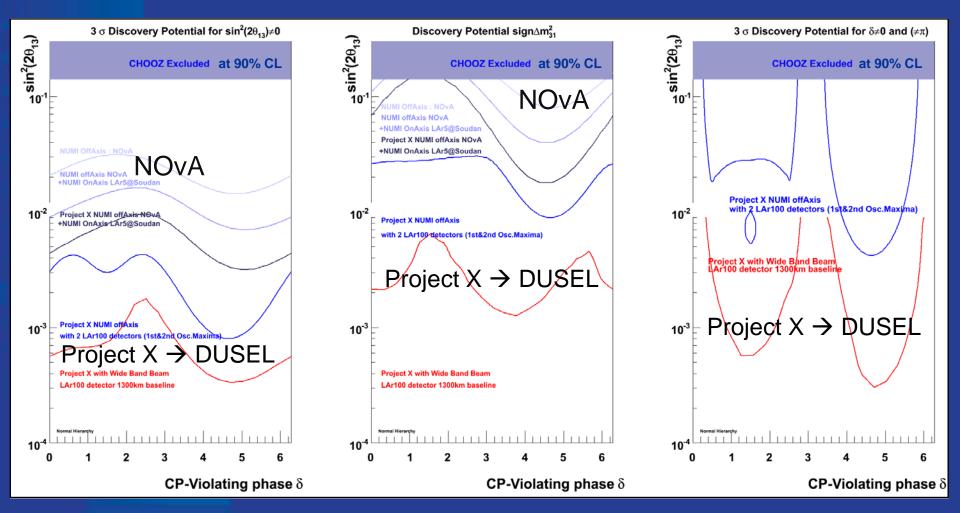


The 3σ reach (2 MW, 100 kton LAr TPC)

$sin^2 2\theta_{13}$

Mass Hierarchy

CP Violation



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