

Detector R&D and Facilities at LBNL

Jim Siegrist 7 October 2010

Oct. 7, 2010 LBNL Detector R&D and Facilities





- Facilities used for Physics Division R&D
- Energy Frontier
- Cosmic Frontier
- Synergy with other LBNL R&D

LBNL Facilities/Capabilities for R&D

- R&D Facilities:
 - Composites Fabrication
 - IC Design
 - MicroSystems Lab
- Infrastructure:
 - Engineering: Mechanical & Electrical
 - Computing:
 - NERSC (National Energy Research Scientific Computing Center)

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- Accelerators/Testbeams/Radiation Testing
- Low background counting capability



More details follow

LBNL Composites Facility Central to Energy Frontier R&D





- Engineering and technical infrastructure for precision fabrication
- Expertise & collaboration with industry (SBIR)
- Used extensively in past for ATLAS construction and future R&D, now also STAR, PHENIX upgrades
- Experienced FT technical staff in-place
- Equipment includes large autoclave, CAD/CAM tools, fixturing, inspection and testing systems, cleaning, adjacent precision machine shop

IC Design Critical to all R&D



CCD control and readout

SOI Monolithic Active Pixels

- Experienced group with 5-7 FTE leveraging different funding lines
- Exchange and sharing of ideas, concepts, tools, design elements between different projects



Recent chips -- Pictures not to scale

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LBNL MicroSystems Laboratory Central to Cosmic Frontier R&D





- Class 10 clean room
- CMOS processing equipment: furnaces, lithography, dry & wet etching, sputtering, wet stations, etc.
- First CCD fabricated in 1996.
- MSL also includes CCD Testing Laboratory in nearby building:
 - Cryogenic light projection systems, QE measurements, charge diffusion, wafer cold probe station, packaging and wire bonding, etc.
- N. Roe MSL lead

R&D for Energy Frontier



- ATLAS Upgrades (G. Gilchriese, M. Garcia-Sciveres, C. Haber)
 - Pixel Integrated circuits
 - Mechanics and integration for silicon detectors
- New Concepts (M. Garcia-Sciveres, C. Haber, B. Heinemann)
 - Intelligent trackers example
- Multi-core Computing (P. Califiura)

Integrated circuits for hybrid pixels



- Critical to reduce the cost and power of pixel systems
- Push the limits of rate and radiation tolerance
- LBNL is leading the ATLAS collaboration in this area
- Brand new FE-I4 chip is largest in HEP to date: see poster
- Now working on generation after FE-I4 to reach highest possible data rates
- Power reduction includes on-chip DC-DC conversion: FE-I4 is first HEP chip to have one.





R&D on Hybrid Pixel Chips



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- FE-I4: Ambitious step towards system-on-chip
 - Full custom layout and synthesized logic tightly integrated down to pixel level
 - (see poster)





Hybrid Pixel Chips Beyond FE-I4



- Further reduce pixel size and at the same time increase rate capability
- Further reduce power
- LBNL leading effort to achieve goals using smaller feature size.
 - This is "The devil we know"
 - Working towards 65nm feature size CMOS submission end of yr.
 - Combined effort of 3 projects: Energy frontier (this one), Intensity Frontier (for Mu2e), and Cosmic frontier (fast ADCs for CCD).
- LBNL is also involved in European-led ATLAS parallel effort using 3D integrated circuits in Multi-Project Run organized by Fermilab.
 - This is a promising, but very challenging approach.
 - This is "The devil we don't know"

Integrated Macro-Assemblies (called staves and petals)



- GOALS: Faster construction and reduced material
- Expand the size of silicon detectors to the entire tracking volume
- Integrate electrical functions, mechanics, and power distribution
- LBNL led the ATLAS strip collaboration in this direction
- Now leading the application of these concepts to ATLAS pixels
- Exploit new carbon foam materials developed in collaboration with industry (SBIR)



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Strip Stave 09



- Prototyping effort being carried by large collaboration:
- LBNL, BNL, UCSC, Yale, NYU, Duke, RAL, Liverpool, Oxford, Cambridge, Sheffield, UCL, Valencia, CERN.
- 1 Stave has up to 120k channels- compare to 46k of original CDF SVX



Pixel Staves 2010





Stave2010 prototype with embedded cable, LBNL Sept. 2010



Model of electrical stave with functional modules using FE-I4 chip. Goal for April 2011





Lots of measurements to tune simulation parameters

All based on new material – low density thermally conducting carbon foam – developed with industry (SBIR)

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Intelligent Trackers



- Community wide interest in increased functionality tracking layers for improving triggers in very dense high rate events.
- "Workshop on Intelligent Trackers" held at LBNL in Feb. 2010.
- 50 participants from LHC experiments and beyond
- Proceedings published in JINST
- <u>LBNL Plans</u>: continue work on coupled layer design. Add collaborators: (Uppsala, US groups...)



Continue development of Si + C-foam interposer wafers



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Multi-core computing



- Computers are changing
- The traditional CPU is a thing of the past
- From now on it's all Multi-Core, rapidly increasing 8, 32, 64.... (supercomputers have been this way for a long time, but now ALL computers are this way)
- This is a challenge, but also a big opportunity for physics software
- LBNL R&D aims at the next generation physics software, taking full advantage of parallel processing



Multi-core development for ATLAS





- AthenaMP code (2-8 cores)
 - Lightweight, process-based, event parallelism
 - Use Linux fork() to share memory automatically
- Many-core (>32) code challenges
 - Reduce memory footprint, maximize data and code locality
 - Group modules according to their code and data "closeness"



Representation of module grouping algorithm

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R&D for Cosmic Frontier



- CCDs for Space and Ground (S. Holland, N.Roe)
 - MicroSystems Lab (MSL) is devoted to this
 - MSL is both an R&D and a fabrication facility
- IR Detectors (C. Bebek)
- WFIRST (M. Levi)
- CMB detectors (A. Lee, H. Spieler)
- Direct Detection of Dark Matter (D. Nygren, S. Loken)
 - Low-Threshold/High-Resolution/Imaging/Directionality
 - Germanium cost reduction strategies
- DUSEL planning (G. Gilchriese)

LBNL CCDs



Poly gate

electrodes

n--(10 kΩ-cm)

CCD structure

photo-

- Thick, fully-depleted CCDs were invented and first fabricated at the LBNL MicroSystems Lab (MSL) in the mid 1990's
- More sensitive at red wavelengths (=high redshifts)
- Very radiation tolerant for space
- HV-compatible design has very good point resolution
- sensitive Partnership with volume 100 10000.00 DALSA semi. $(300 \, \mu m)$ 90 1000.00 enables large 80 BNI 86135.7.7: Quantum Efficiency [%] 250 µm thick (-140C) 70 scale production [m] 100.00 60 Length Commercial CCD: ~ 20 µm thick for projects 50 10.00 such as Dark Absorption 40 1.00 Absorption Length at -120/-100C 30 Energy Survey 20 (over 160 wafers 0.10 10 processed) 0.01 300 500 700 900 1100 Wavelength (nm)

CCDs produced and deployed





Keck 10m telescope LRIS spectrograph



Lick Observatory/Mt Hamilton Hamilton Echelle spectrograph



Kitt Peak/Mayall 4m MARS and RC spectrographs



Palomar Hale 200"telescope SWIFT spectrograph



MMT 6.5m/Mt. Hopkins Red Channel spectrograph



2.5m SDSS telescope BOSS spectrographs

Future CCDs





Dark Energy Survey 2011 CTIO Blanco 4m Telescope





BigBOSS 2015 - proposed KPNO Mayall 4m Telescope



Mayall 4-Meter Telescope



Joint Dark Energy Mission/ WFIRST ?



LSST has also baselined fully depleted CCDs

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CCD R&D directions



- Thinner CCDs using TiN metal for reduced diffusion, improved yields
- Charge multiplying CCD for single photon sensitivity
- Lower noise output transistors
- Higher speed output transistors
- Combination of the above in a single device, e.g. an LSST prototype with multiple readouts, low noise and thin substrate using TiN.
- Fast CCD with multiple outputs for soft x-ray detection at ALS

R&D done for JDEM (now for WFIRST)



- HV-compatible architecture to reduce Point Spread Function
 - Achieved 4 μ m rms with 100V bias voltage
- Irradiation studies at LBNL 88" cyclotron
 - Demonstrated robustness for 6 yr space mission
- Buttable package with space qualified components and compact readout module



JDEM / WFIRST heterogeneous focal plane prototype





- 32-detector silicon carbide focal plane prototype
- Space qualified Near Infrared (NIR) detectors
- Space qualified fully-depleted CCDs from MSL
- Instrument now assembled and undergoing electrical tests.

NIR modules for focal plane



- Commercial components
 - Teledyne SIDECAR ASIC
 - Hg(x)Cd(1-x) 2kx2k photo diode arrays
- LBNL electro-mechanical packaging
 - Supports direct connect to detector
 - Supports 36-channel mode
 - LVDS digital interface
 - Four science data links
 - Temperature monitoring
 - Two delivered; four more in assembly





CMB detectors (LBNL & UCB)



- Superconducting Transition Edge Sensor bolometer arrays
- New technology yields unprecedented sensitivity and systematic error control for mm wavelength radiation
- To be used in several detectors:
 - Galaxy Clusters
 - <u>APEX-SZ</u>
 - First science from new generation TES arrays
 - South Pole Telescope (SPT-SZ)
 - Discovery of first clusters via SZ effect
 - CMB Polarization
 - POLARBEAR
 - First science data is 2010/2011 with 1274 bolometer array
- Use LBNL-developed frequency-domain mux
- Pixel design in cooperation with UCB
- Future: Array integration and high-speed low-power readout

POLARBEAR: CMB Polarization



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High Pressure Xenon for Dark Matter and Double Beta Decay

- Excellent intrinsic energy resolution:
 - (only 3 times worse than Ge)
- TPC + Electroluminescence = 3D tracking (background suppression by event topology) + excellent E resolution
- Low background
- 1 kg prototype nearly completed at LBNL, to demonstrate ~0.5% FWHM
 @ 511 keV gammas
- Optimizing design towards a 100 kg enriched ¹³⁶Xe for 0ν ββ search
 @ Canfranc underground lab

(D. Nygren)

A. Bolotnikov, B. Ramsey / Nucl. Instr. and Meth. in Phys. Res. A 396 (1997) 360-370



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Fig. 5. Density dependencies of the intrinsic energy resolution (%FWHM) measured for 662 keV gamma-ray

High Pressure Xe prototyping





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TPC with GEMs & Pixels: Collaboration with a university group

- New collaboration with U. of Hawaii
- Spinoff of pixel chip technology
- Full 3D tracking in large gas volumes
- Multiple applications
 - Directional Dark Matter Detection
 - Directional Neutron Detection (Homeland security)
- Current UH Work funded by NSF-ARI
- Only possible with LBNL help:
 - Initial proof-of-concept (for ILC) part of detector R&D at LBNL
 - UH PI (Vahsen) is former LBNL postdoc
 - Employs ATLAS pixel electronics developed at LBNL
 - Gas setup expertise provided by LBNL



Tracking of cosmic ray at LBNL



Lower Cost Germanium



Needed to meet growing R&D demand

Application	Detector Requirements	Total mass desired (next ~10 years)
Nuclear Physics (GRETA, AGATA)	n-type tapered hexagonal segmented detectors	~500 kg
0νββ Decay, Dark Matter (MAJORANA, GERDA, 1TGe, CDEX)	p-type point-contact detectors Large volume desired	Tonnes
Space Applications (GRIPS, NCT)	Planar n-type Large area desired	~100 kg
Medical Imaging	Large area desired	Hundreds+ kg/yr
Nuclear Safeguards, Homeland Security, Environmental Monitoring, Emergency Response	Depends on success of current- generation systems	Hundreds+ kg/yr

 Must work with industry to increase yield and reduce cost of detector grade material (HPGe)

Development of a low cost production process for HPGe





Organization:

Analysis and design: LBNL lead with selected experts and in collaboration with HPGe producer Equipment fabrication: subcontract Equipment testing: LBNL in collaboration with HPGe producer

Synergy and collaboration



- Physics Division is one of several at LBNL
- There is a great deal of instrumentation R&D outside of physics
 - Many ideas, resources, and advances are shared
 - Will give a few examples to illustrate.
- Much of the R&D described involves collaborations of many labs and universities in the US and abroad
- In many cases LBNL has initiated the R&D and then expanded to involve collaborators.
 - Example: FE-I4 chip, 11 designers from 5 institutes outside LBNL and many more for testing
 - Example: Stave R&D, now 13 outside institutions
 - Example: WFIRST, now 5 outside institutions

Synergy example: STAR and Phenix upgrades



- Composites facility producing mechanics for BOTH detectors.
- STAR HFT upgrade is a major LBNL project using engineering resources and composites.
- Know-how from ATLAS construction being applied, and new concepts from STAR HFT seeding development for ATLAS upgrades



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Synergy example: SOI & bulk active pixels (silicon on insulator)

BERKELEY LAB

Started as ILC R&D. Now fully supported outside HEP for Electron Microscopy and Photon science (*P. Denes*)



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Parallel readout CCDs can greatly increase acquisition speed of light source experiments.

- This is a different solution than SOI active pixels. Each solution has different advantages.
- Next step is to use 65nm technology to fit one ADC per CCD column
- pixel readout for Energy Frontier and digitizer for Intensity frontier

Combined 65nm submission with

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Synergy example: parallel CCDs for photon science



Parallel readout CCD detail



Outlook



R&D challenges driven by physics

- Energy Frontier detectors:
 - LHC upgrades and Lepton Colliders detectors: Low mass, low power, high rates, rad hard, new ICs, etc.
- Cosmic Frontier detectors:
 - Branching out from HEP detectors to photon sensors: CCDs, IR, CMB
 - Future work: single photon counting, high rates, wide wavelength coverage
 - Dark matter, $0\nu \beta\beta$, neutron detection
 - Future work: new applications of TPCs, lower the cost of GE to enable large scale detectors
- Continued support for R&D facilities and infrastructure critical for development of advanced detectors
- Continued collaboration with university groups is an essential part of the LBNL program



Backup

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Turning staves into track trigger staves

- Redundant doublet of axial strips is sensitive to track momentum
- no stereo angle offset=2 offset=0 Data reduction vs cut offset. Simulations show data reduction factor of 100 is possible • Pile-up (x1-x400) 3.00% (T. Mueller, B. Heinemann, M. Garcia-Sciveres) has little effect -X1 clusters 2.509 ---X 100 Number of Hits -X 400 2.009 NAR STREET, ST # 20GeV 1.50% π⁺ 1GeV coincidences 20GeV 1.009 π 1Ge 10² 0.509 0.00% 0.5 1 1.5 2 2.5 з 3.5 corrected phi difference cut in units of strip pitch 10 20 Corrected Phi Difference in Units of Strip Pitch offset Added flexible circuit showing path 'Stave" assembly Implementation concept : from top to bottom (existing) with silicon strips on (C. Haber, M.Garcia-Sciveres) both sides. Bottom O O not visible

high pT

or secondary

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low pT