

Probing Supersymmetry with Very Light Gravitino by Using Impact Parameter at the LHC

Takumi Ito
(U. Tokyo, Tohoku U.)

Collaborators
Masaki Asano (U. Tokyo),
Shigeki Matsumoto (IPMU, U. Tokyo),
Takeo Moroi (U. Tokyo)

To appear on arXiv [hep-ph].

Supersymmetry with light gravitino at the LHC

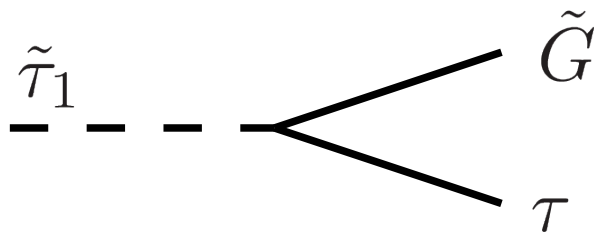
In this talk, I consider the supersymmetric model with

LSP : **Gravitino** mass $\sim O(10)$ eV

NLSP : **Stau** mass $\sim O(100)$ GeV

(Such a spectrum is realized, e.g., in low-scale gauge mediation models.)

$$c\tau_{\tilde{\tau}} \simeq 2 \text{ mm} \times \left(\frac{m_{3/2}}{10 \text{ eV}} \right)^2 \left(\frac{m_{\tilde{\tau}}}{100 \text{ GeV}} \right)^{-5}$$



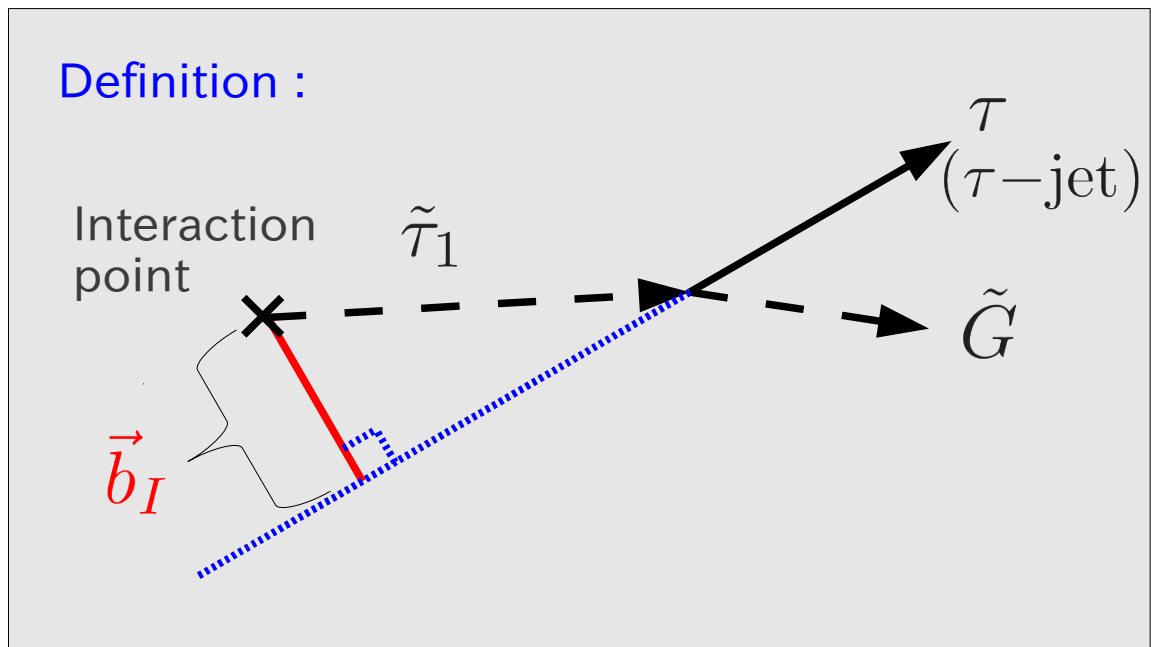
The stau decays before hitting inner detectors.

Impact parameter may be useful to probe such a signal.

Plan of talk

- Impact parameter of τ -lepton (τ -jet).
- Mass determination in SUSY cascade
 - squark, bino, stau
- Stau lifetime determination

Impact parameter



(3-dimensional) impact parameter

$$b_I \equiv \left| \vec{b}_I \right|$$

$$\delta(b_I) \simeq 100 \mu\text{m}$$

Transverse impact parameter

$$d_I \equiv \left| \left(\vec{b}_I \right)_T \right|$$

$$\delta(d_I) \simeq 10 \mu\text{m}$$

Better resolution. (c.f., arXiv:0901.0512[hep-ex])

Hence, we use d_I in the following.

Typically, $d_I \sim O(c\tau_{\text{NLSP}})$ for τ -jet from the stau decay.

Impact parameter is useful for the signal with $c\tau_{\text{NLSP}} \gtrsim 100 \mu\text{m}$
 (Notice that $c\tau_\tau = 80 \mu\text{m}$)

Monte Carlo study

* Sample spectrum (GMSB)

gluino ... 1096 GeV

squark_R ... 920 GeV

neutralino₁ ... 197 GeV

stau₁ ... 126 GeV

gravitino ... ~ 0 GeV

($\leftrightarrow c\tau_{\text{NLSP}} = 500 \mu\text{m}$)

* Total SUSY cross section ... 2.7 pb
($pp, \sqrt{s} = 14 \text{ TeV}$)

* Event generation ... HERWIG6.510

* Detector simulation ... PGS4

* Resolution of the impact parameter
is included in our simulation.

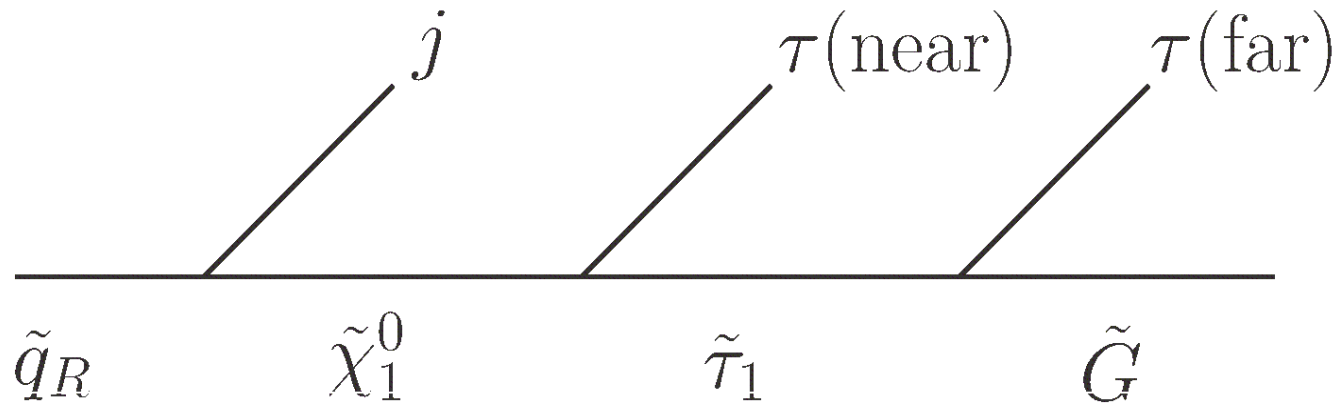
$$\delta(d_I) = 10 \mu\text{m}$$

Basic cut	SUSY	$t\bar{t}$
(Integrated luminosity of 100 fb^{-1})	273,600	49,610,000
$p_T > 150 \text{ GeV}$	190,504	1,644,160
The number of lepton-like objects ≥ 4	70,641	55,537
The number of b-jets = 0	56,228	41,673
The first jet with $p_T > 200 \text{ GeV}$	49,819	5,915
The second jet with $p_T > 100 \text{ GeV}$	41,007	1,984

After cut

Mass measurement of M_{squark} , M_{bino} and M_{stau}

SUSY cascade decay chain :



$M_{\tau\tau}$
$M_{j\tau(\text{near})}$
$M_{j\tau(\text{far})}$
$M_{j\tau\tau}$
$M_{T2}(\tilde{q})$
\vdots

* Hadronic decaying τ (“ τ -jet”) is used.

* $\tau^{(\text{near})}$ and $\tau^{(\text{far})}$ will be distinguished by

$$d_I(\tau^{(\text{near})}) < d_I(\tau^{(\text{far})}).$$

* Fake τ -jets can be removed by the charge subtraction, i.e.,

$$(\text{Signal}) = \underbrace{(\tau^+\tau^- + \tau^-\tau^+)}_{\text{OS}} - \underbrace{(\tau^+\tau^+ + \tau^-\tau^-)}_{\text{SS}}$$

RESULT : Mass measurement of M_{squark} , M_{bino} and M_{stau}

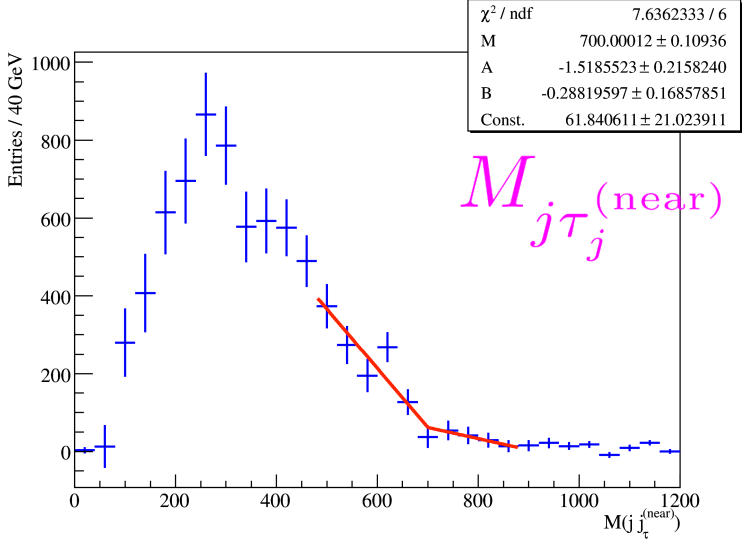
Thanks to dI information,
 $M_{j\tau(\text{near})}$ endpoint is available.

Assuming $m_{3/2} = 0$, we obtain

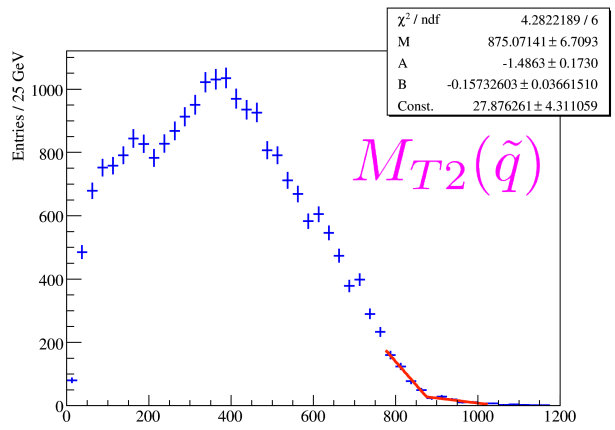
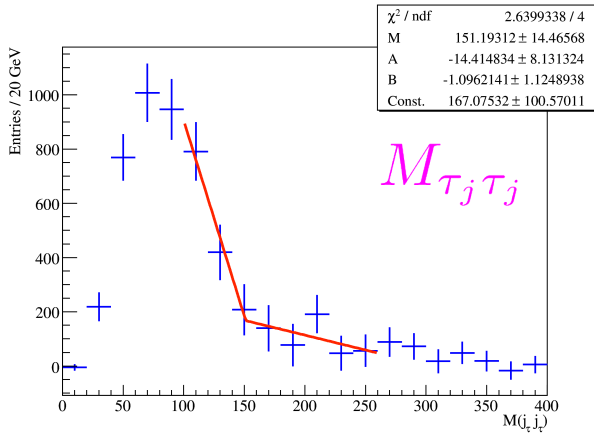
$$M_{\tilde{q}} = 916 \pm 6 \text{ GeV} \\ (m_{\tilde{q}} = 922 \text{ GeV})$$

$$M_{\tilde{B}} = 193 \pm 20 \text{ GeV} \\ (m_{\tilde{B}} = 197 \text{ GeV})$$

$$M_{\tilde{\tau}} = 121 \pm 18 \text{ GeV} \\ (m_{\tilde{\tau}} = 126 \text{ GeV})$$



In addition, two endpoints are studied.



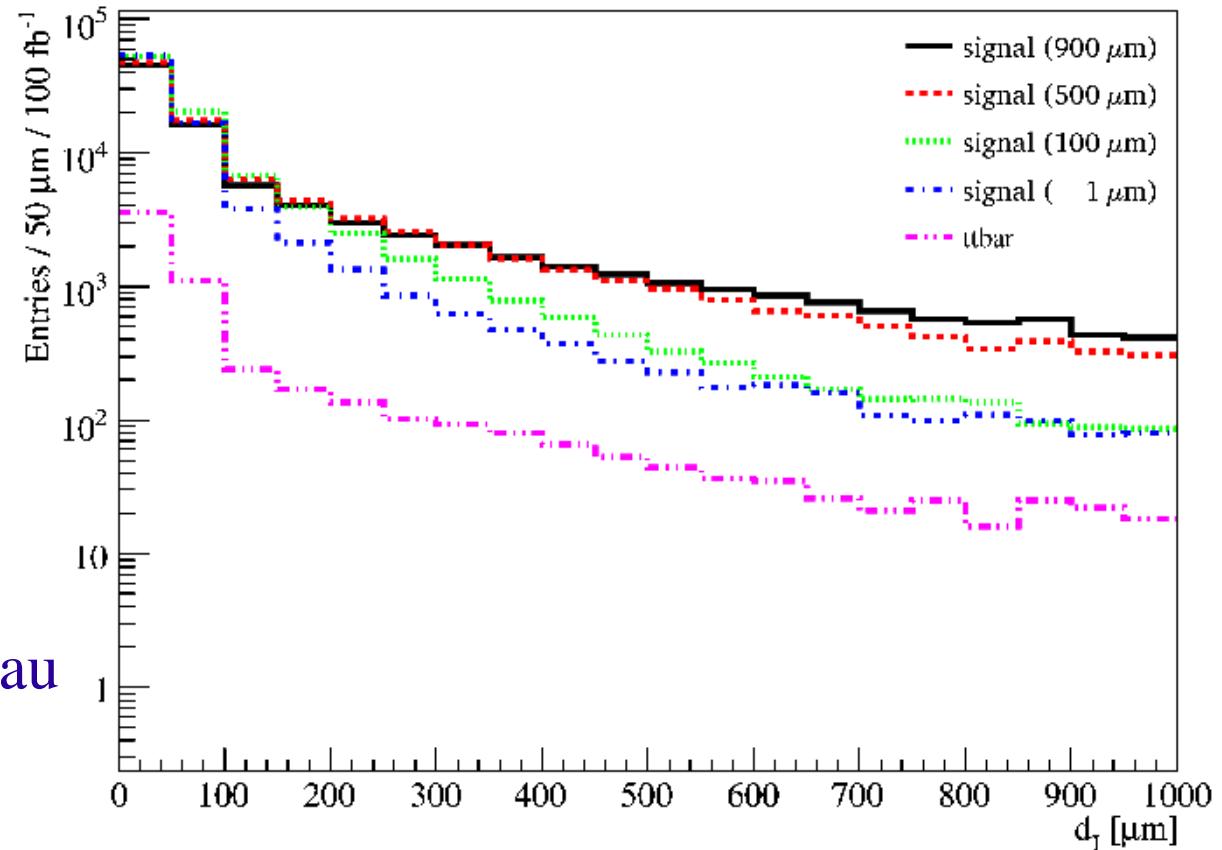
Distribution of the transverse impact parameter, d_I

Here, we treat the gravitino mass as a free parameter.

$\rightarrow c\tau_{\text{NLSP}} = 1, 100, 500, 900 \mu\text{m}$.

d_I depends on

- NLSP lifetime, and
- Lorentz boost factor of stau



After mass measurements, one would perform the MC simulation and obtain a template of d_I distribution for a *hypothetical* $c\tau_{\text{NLSP}}$.

NLSP lifetime is determined by comparing a real data with the template.

Lifetime determination

Here, we perform simple MC analysis as a demonstration.

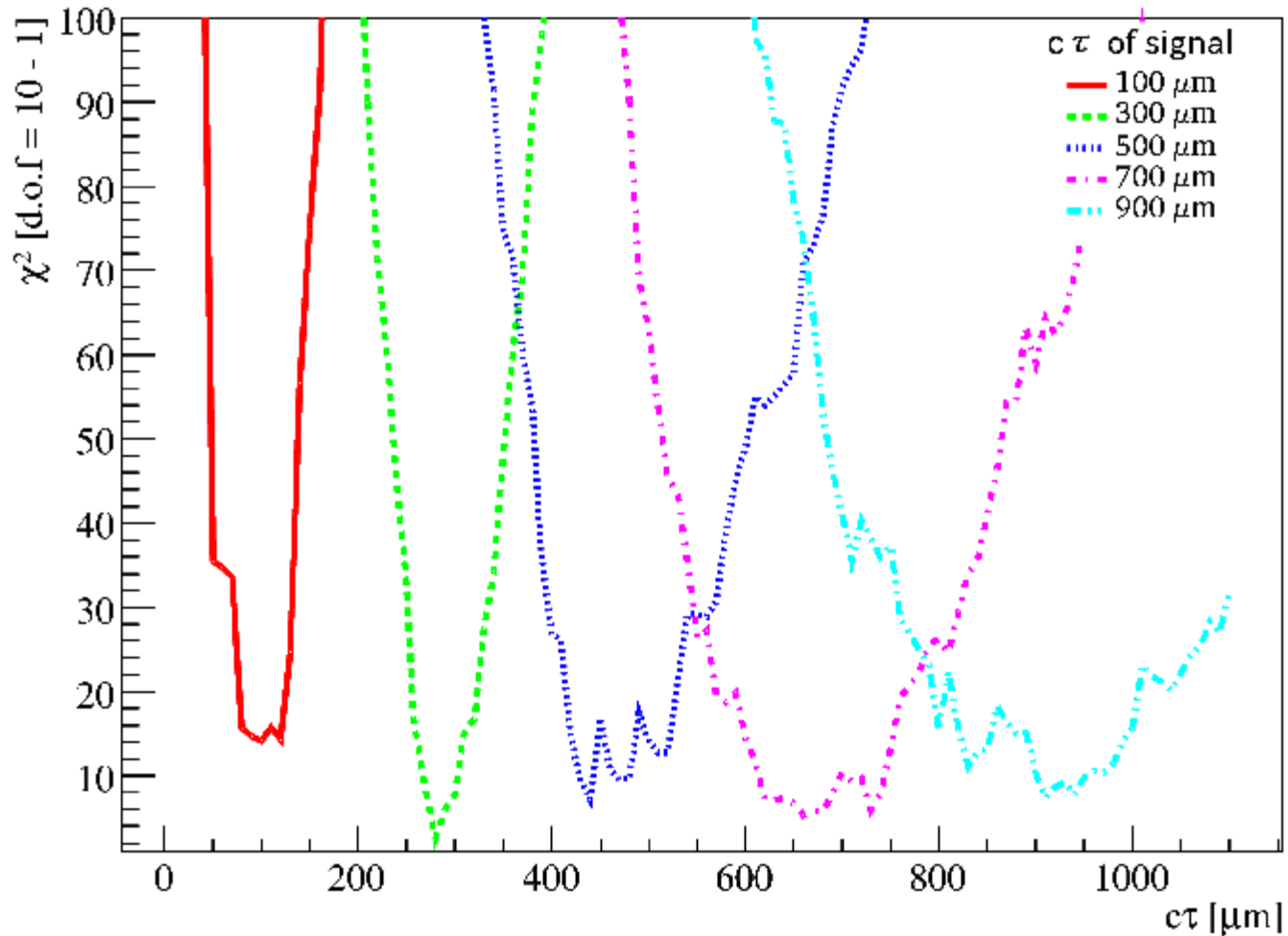
Making templates of the d_I distribution

- Taking some hypothetical lifetime, $c\tau_{\text{NLSP}}$ ($= 10, 20, \dots, \mu\text{m}$).
- The stau is produced at the origin with some fixed β .
- Then, the stau decay into tau & gravitino is simulated.
- We obtain the template of the d_I distribution.

We perform χ^2 -test.

The d_I distribution of the “signal” is compared to the templates.

RESULT : Lifetime determination



* “Signal” (all SUSY events) : $c\tau_{\text{NLSP}} = 100, 300, 500, 700, 900$ μm .

* In order to enhance the signal, χ^2 -test is performed in the range $0.5 \times c\tau < d_I < 2 \times c\tau$, for each $c\tau$.

Summary

- * I have discussed LHC phenomenology on SUSY with light gravitino-LSP and stau-NLSP.
- * For $m_{3/2} = \mathcal{O}(10)$ eV case, impact parameter information will be useful to detect a NLSP decay.
 - Mass measurements (squark, bino, stau)
 - Lifetime determination (stau)

BACKUP

Mass measurements (including gravitino)

