

# Iceberg analysis with dataprep: Noise and SNR

## Iceberg analysis

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# Introduction

I continue look at Iceberg data

- Results plots posted at dedicated Iceberg link:
  - <https://internal.dunescience.org/people/dladams/protodune/iceberg/data/index.html>

Today look more at time dependence of noise

- Run-to-run variation of noise averaged over all wires
- Noise summaries and DFT spectra for low, typical and high noise

And estimates of SNR

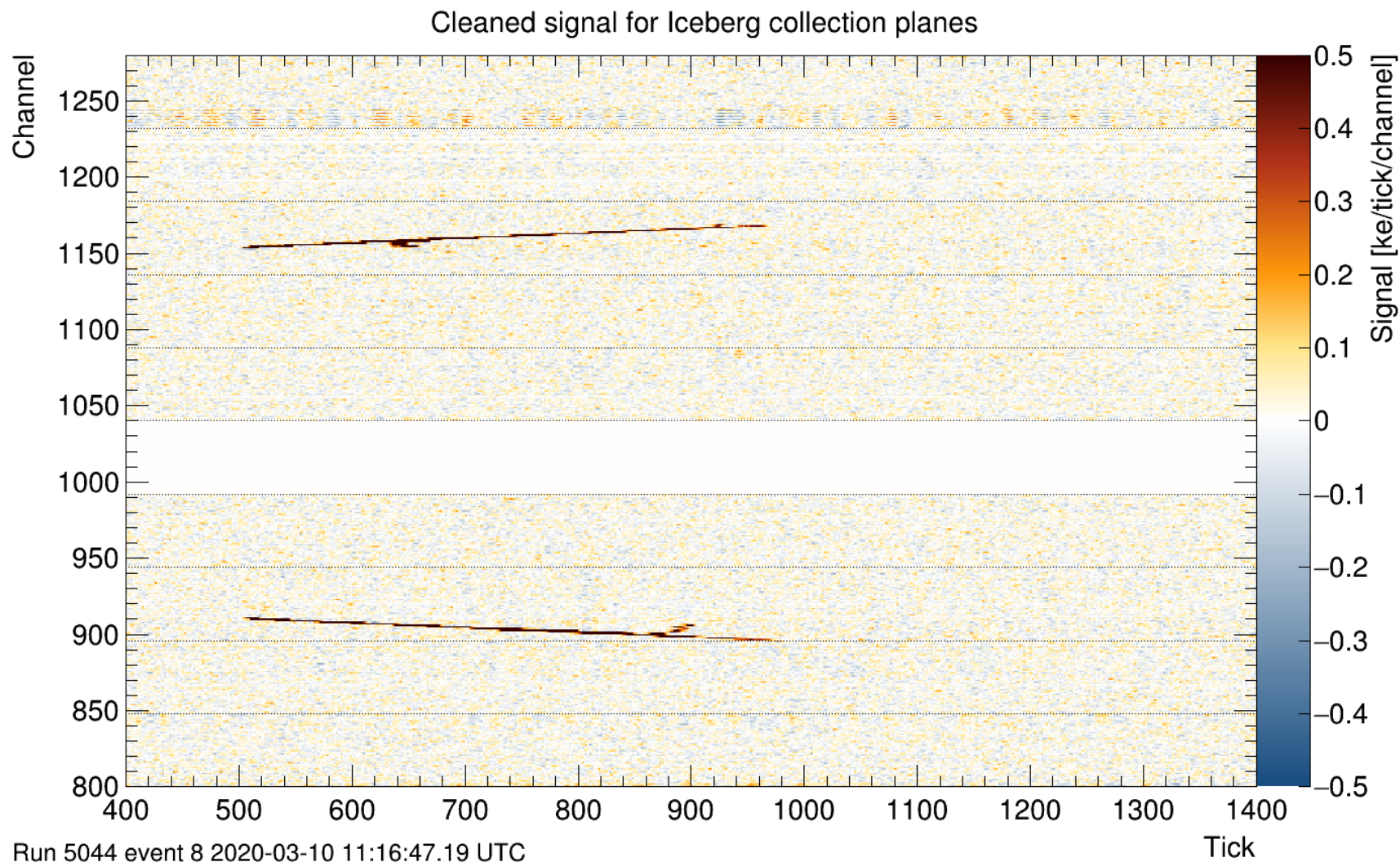
# Noise evaluation procedure

# Noise evaluation

## How noise is measured

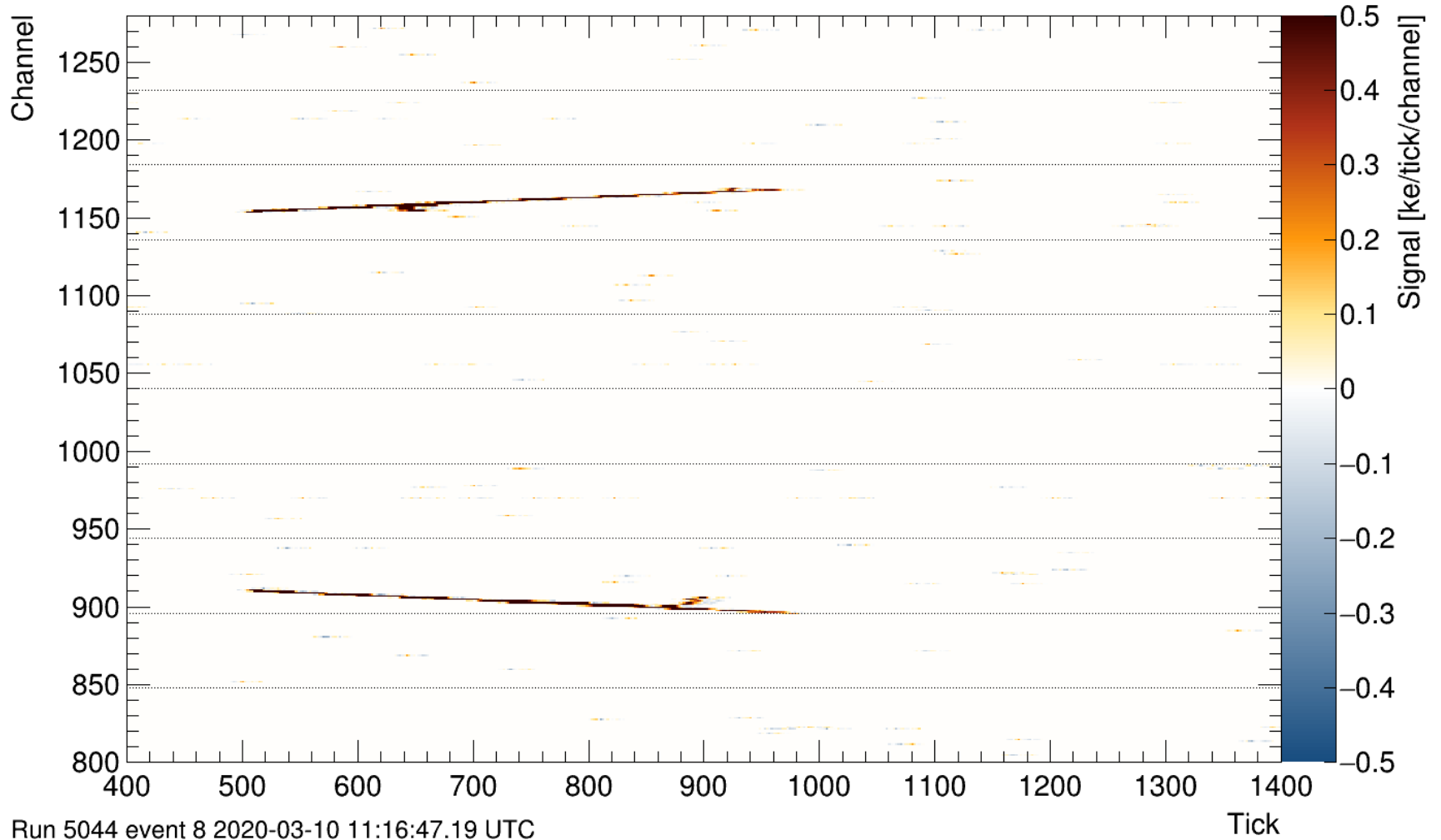
- Quick summary—see previous talks for more details
- Data preparation similar to protoDUNE
  - Identified sticky codes are flagged
  - Pedestal evaluated for each channel each event
  - Sticky codes mitigated (replace with interpolated values)
  - Tails (from amp-ADC AC coupling) are removed
  - Coherent noise removal (CNR) is applied
    - Subtracting median in each FEMB-plane (I think)
    - Noise results shown with and without CNR
- Signal-free regions selected
  - Dynamic threshold signal finder used to identify ROIs
    - Ticks inside ROIs are flagged as signal
  - Blocks of 1000 ticks w/o signal are identified
- Noise is defined as RMS over not-ROI ticks
- DFT power is evaluated us the signal-free 1000-tick blocks

# Example waveform after dataprep

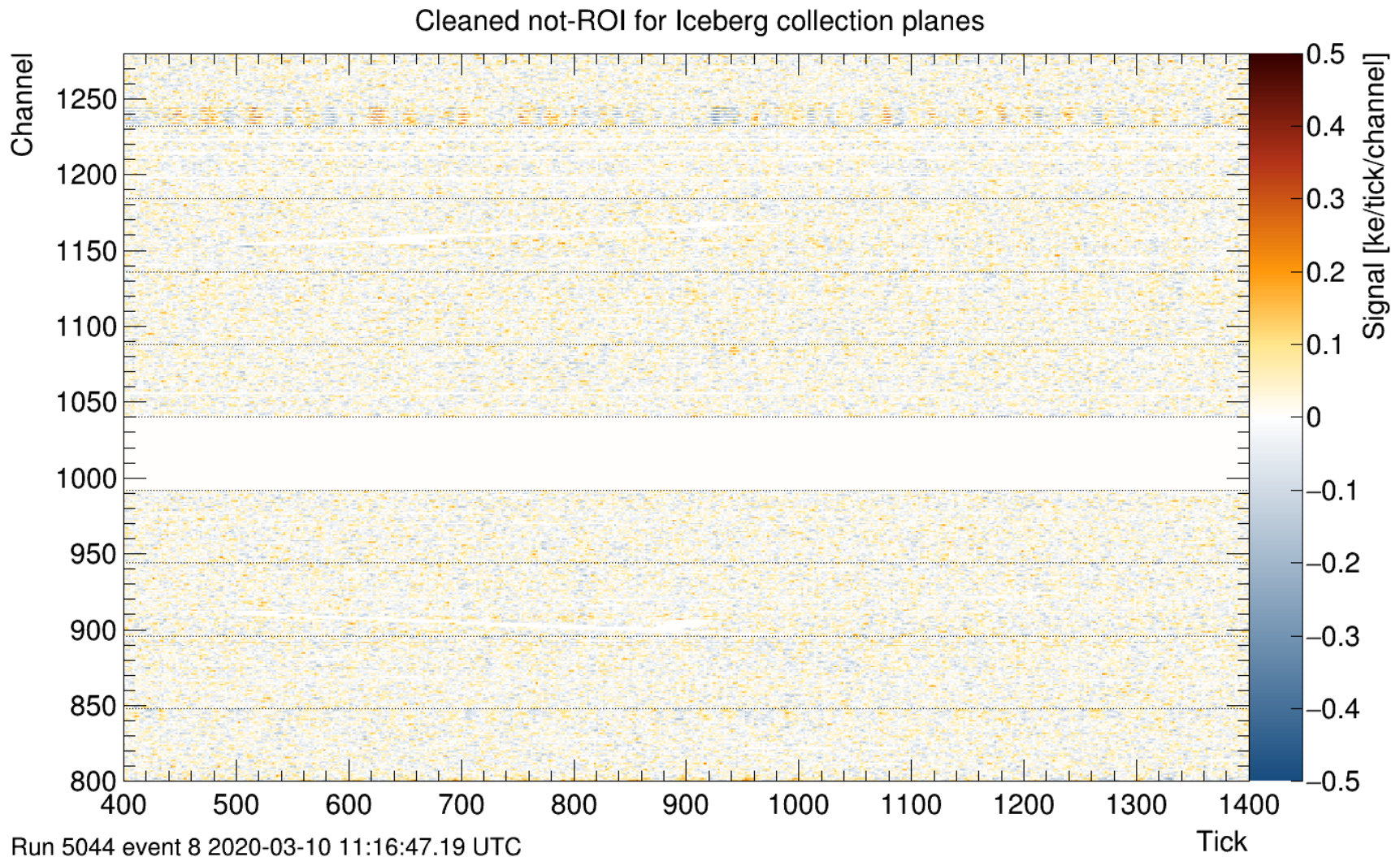


# After signal finding

Prepared signal for Iceberg collection planes



# Noise ticks (i.e. inverting signal finding)



# Noise summary

## Noise summary plots

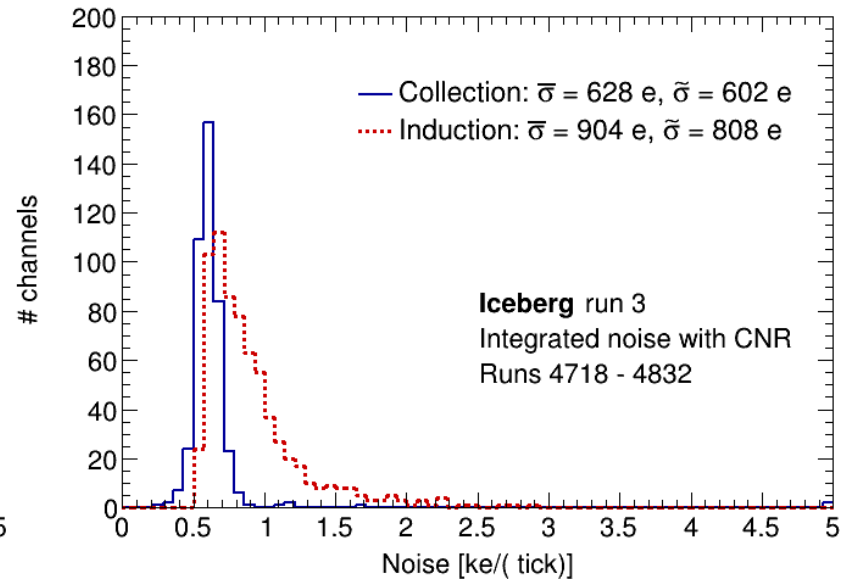
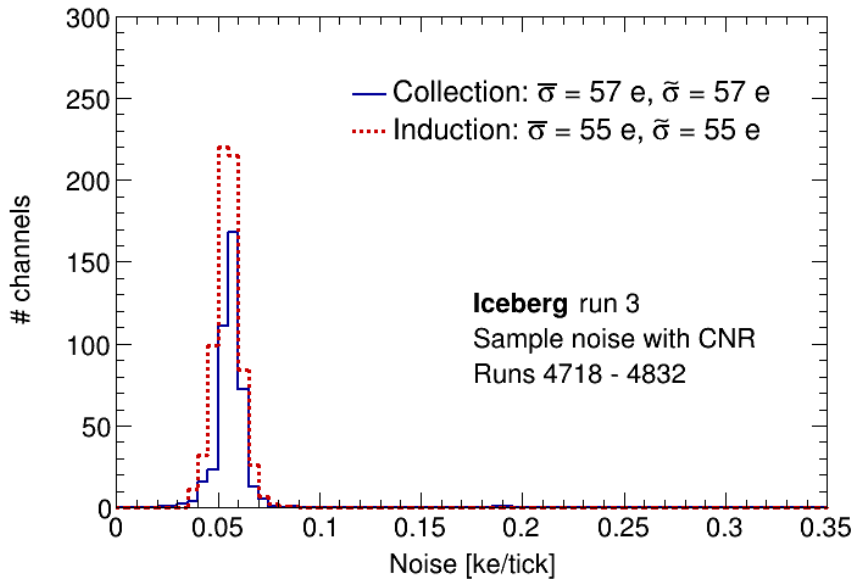
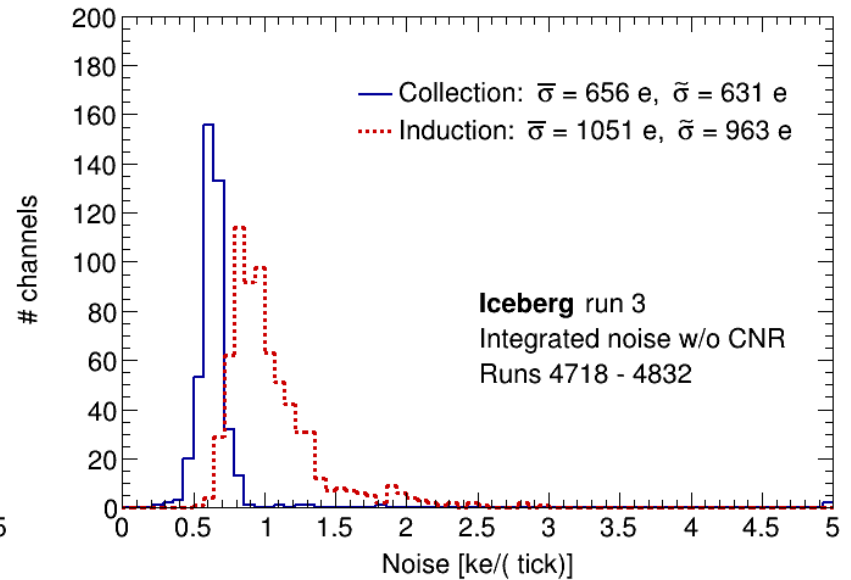
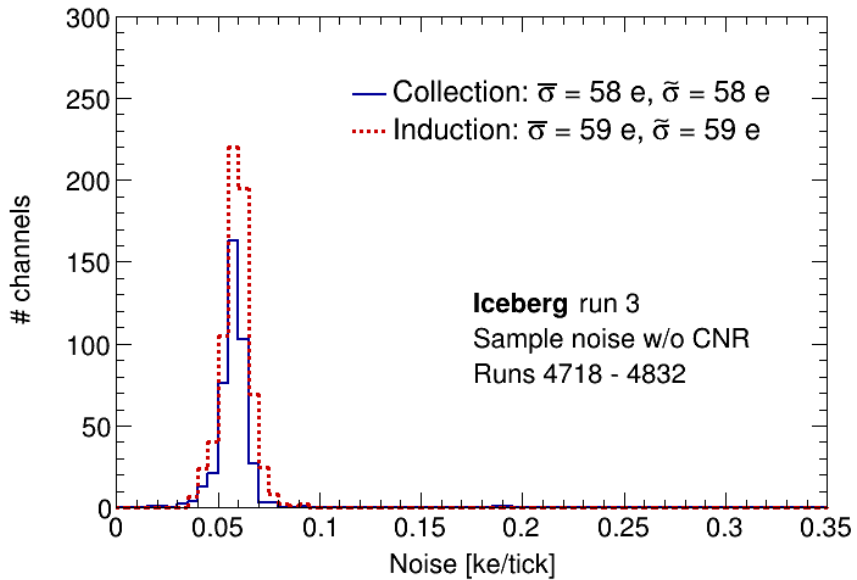
- Noise = RMS in signal-free regions
  - Sample noise = individual samples
  - Integrated noise = sums over 50 ticks
    - To approximate physics charge measurements
    - MIP deposit is 20-30 ke
- Eight curves with mean and median (new) are shown
  - Before and after CNR (top, bottom)
  - Sample and integrated (left, right)
  - Collection and induction (blue, red)
- One entry for each channel
  - Excluding channels flagged bad (134) or noisy (26)

## Summary for full run period on following page

- Sample noise looks good
- Integrated induction is very broad
- Integrated collection and induction have tails

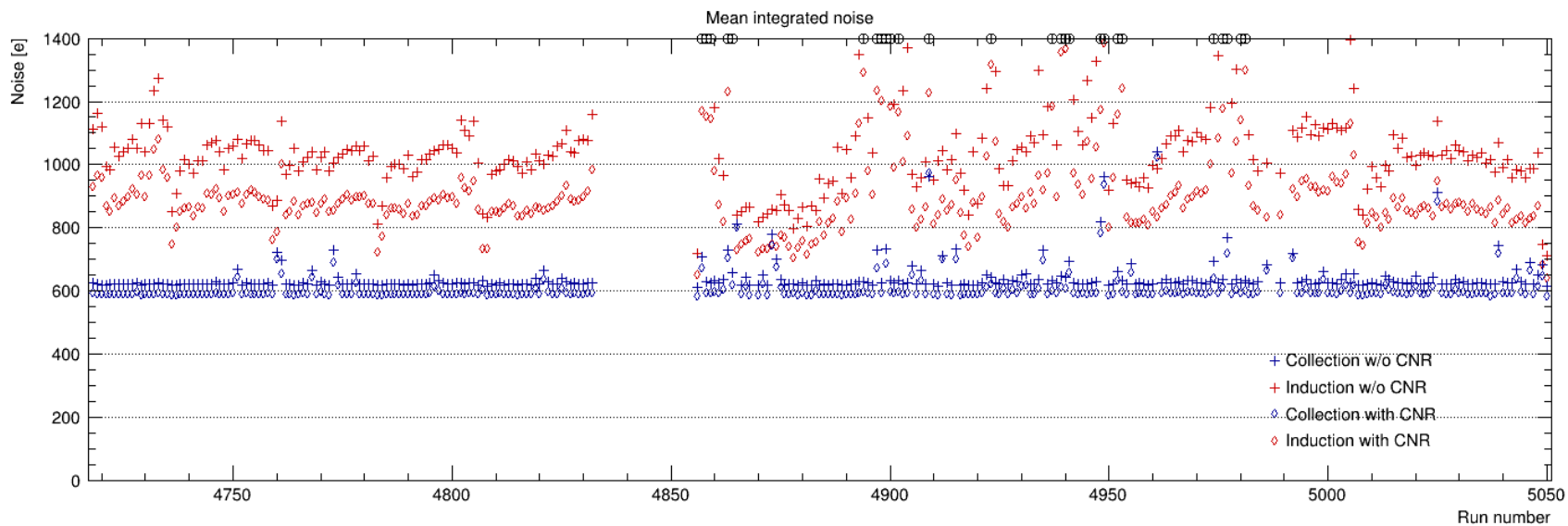
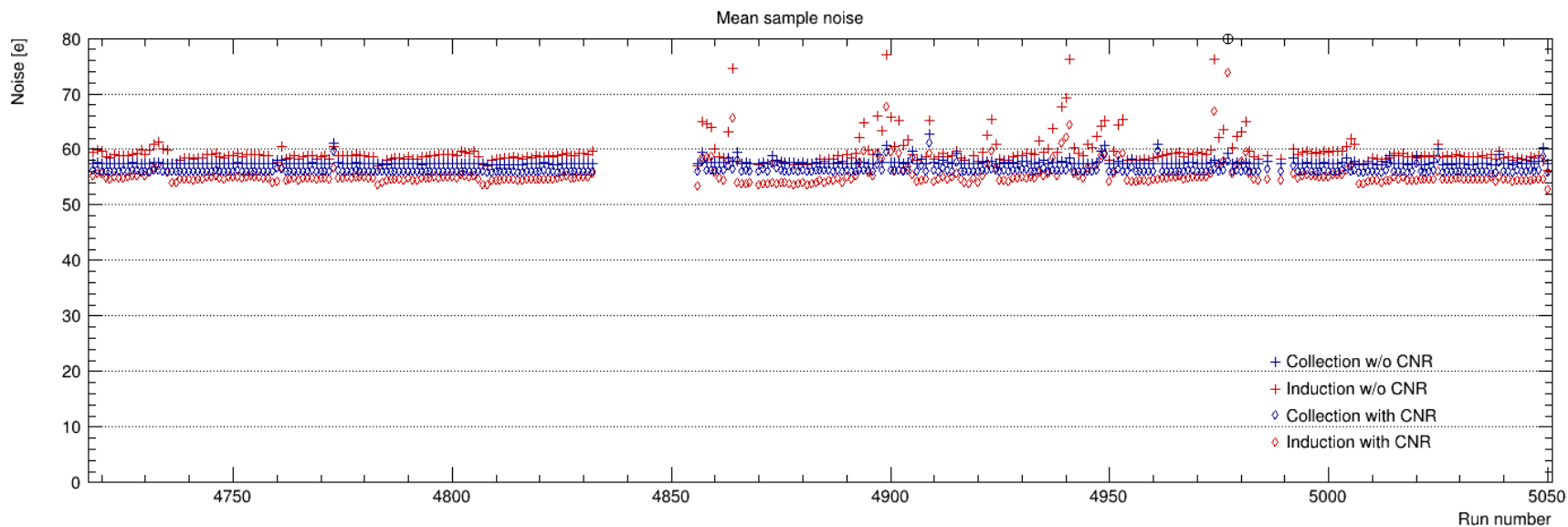


# Early runs (4718 – 4832)

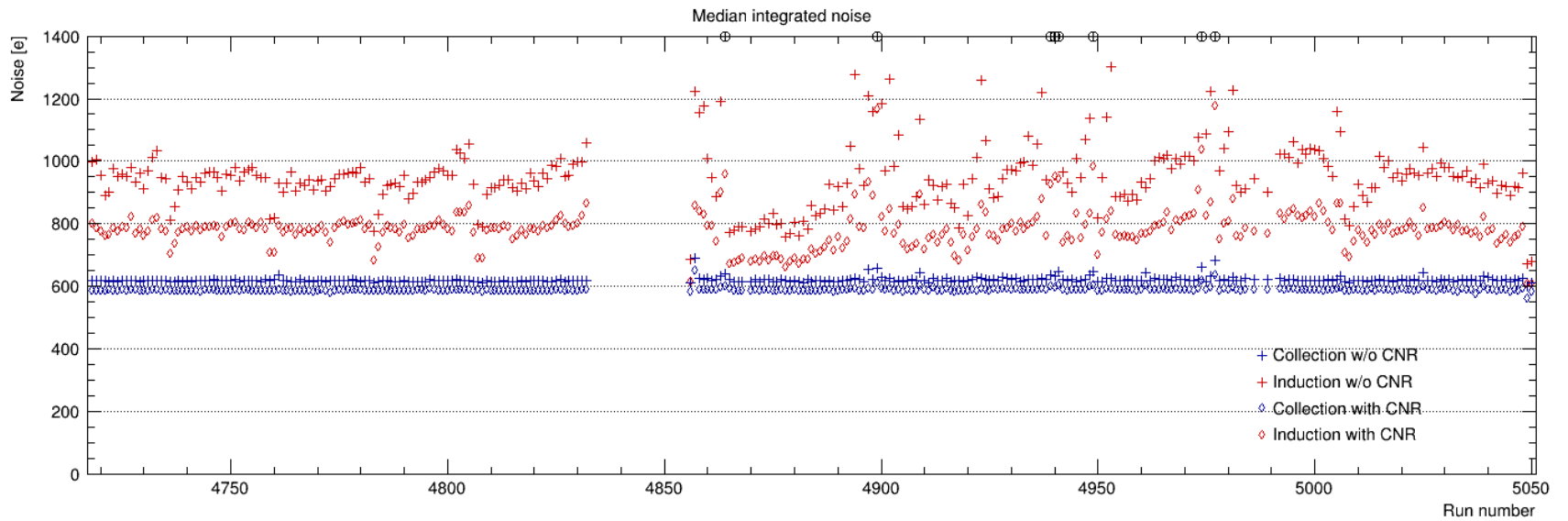
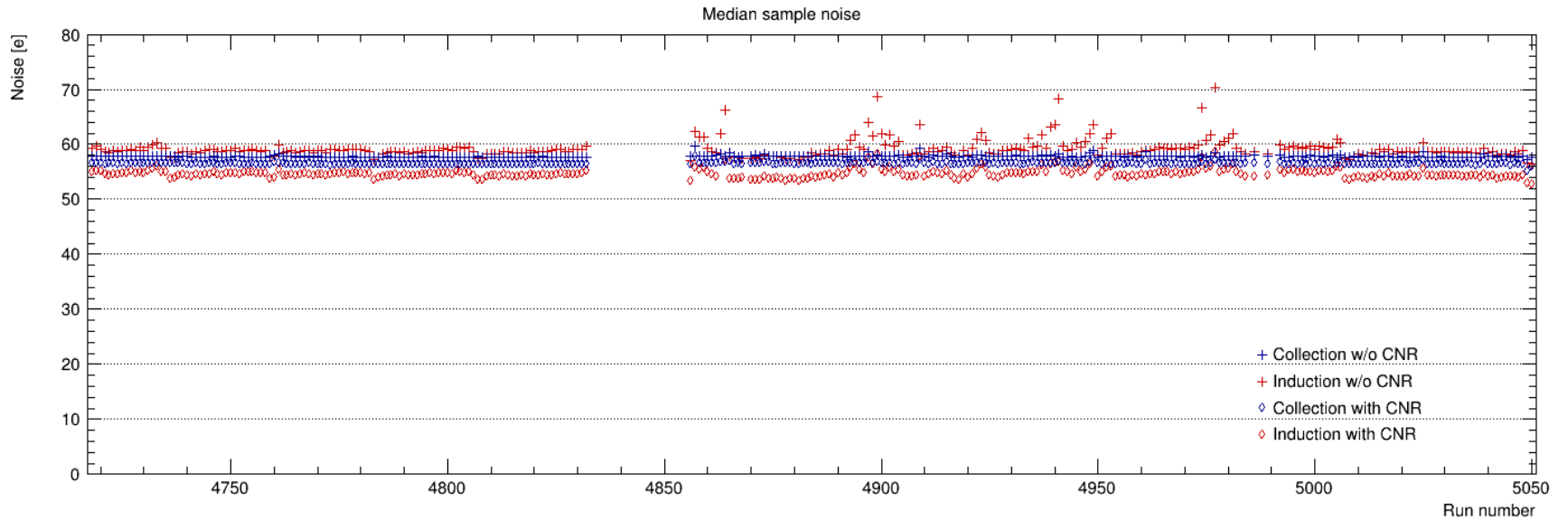


# Noise vs. run

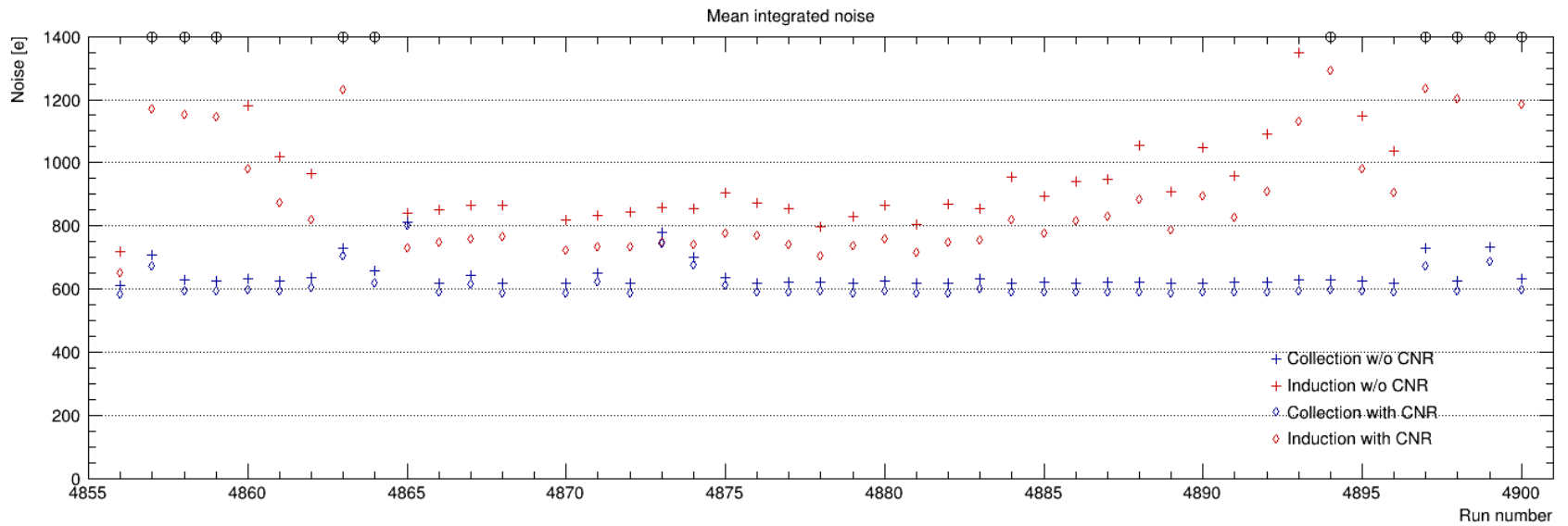
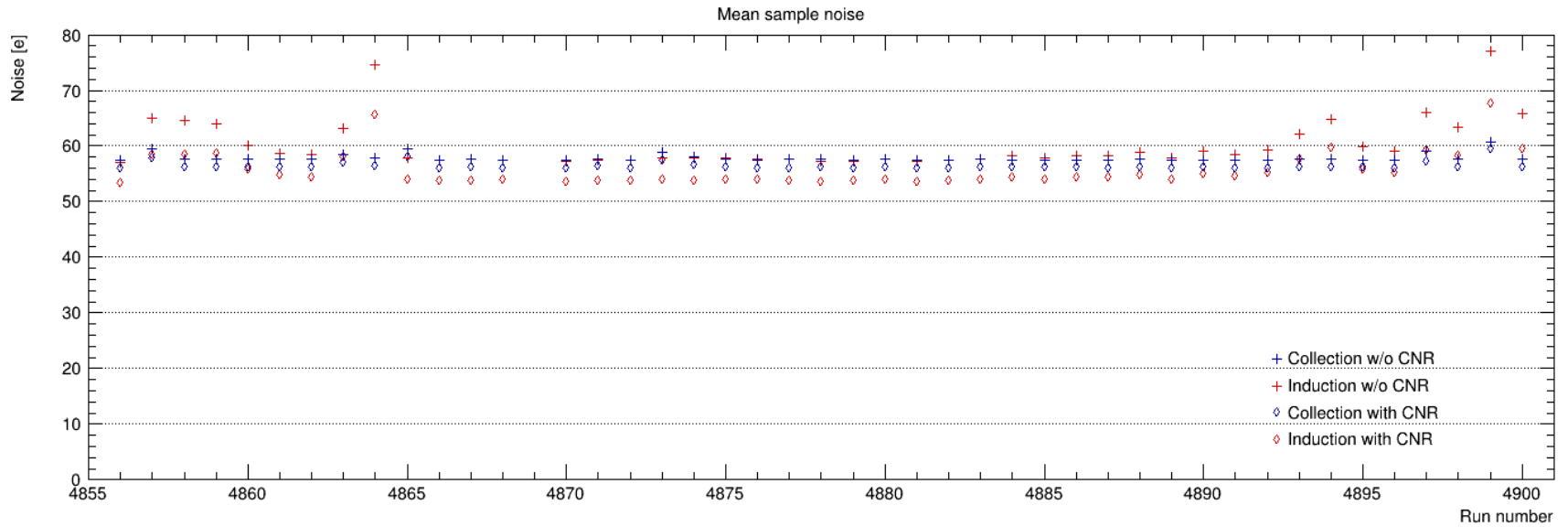
# Mean noise vs. run



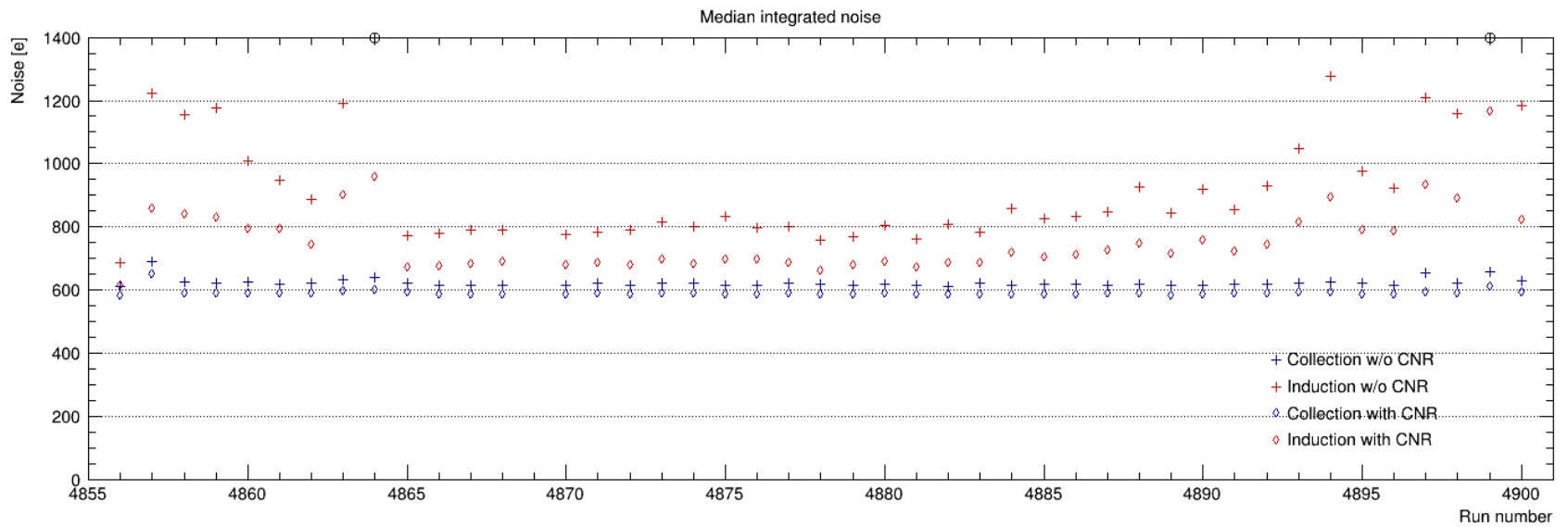
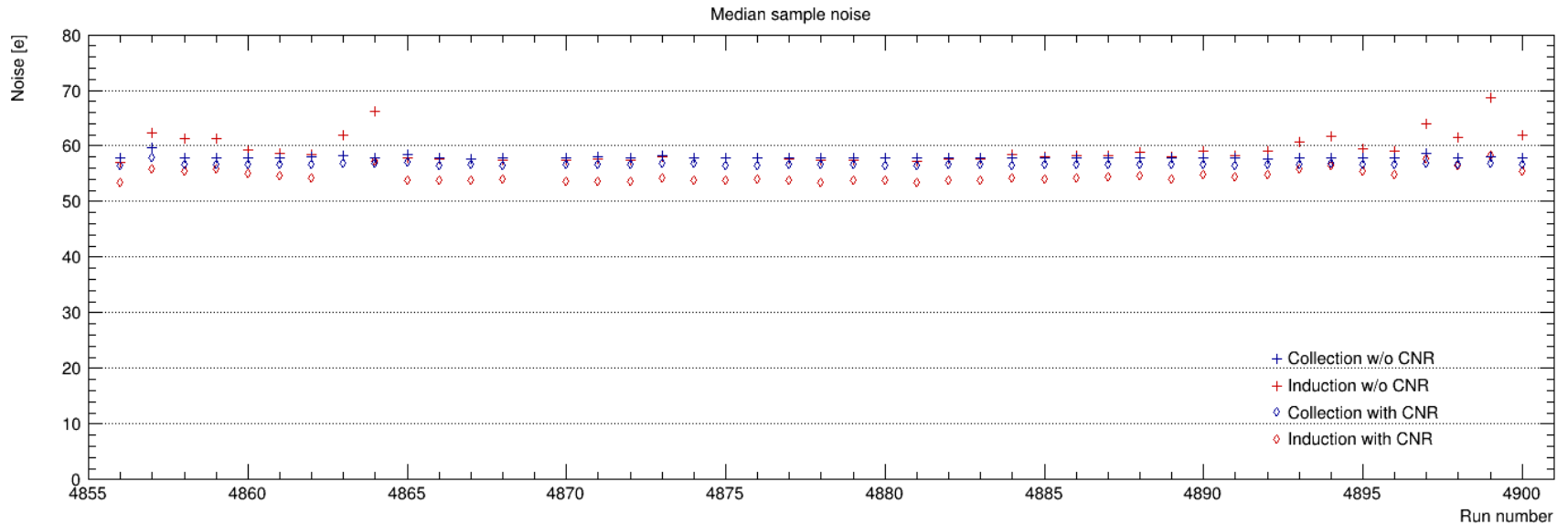
# Median noise vs. run



# Zoomed mean noise vs. run

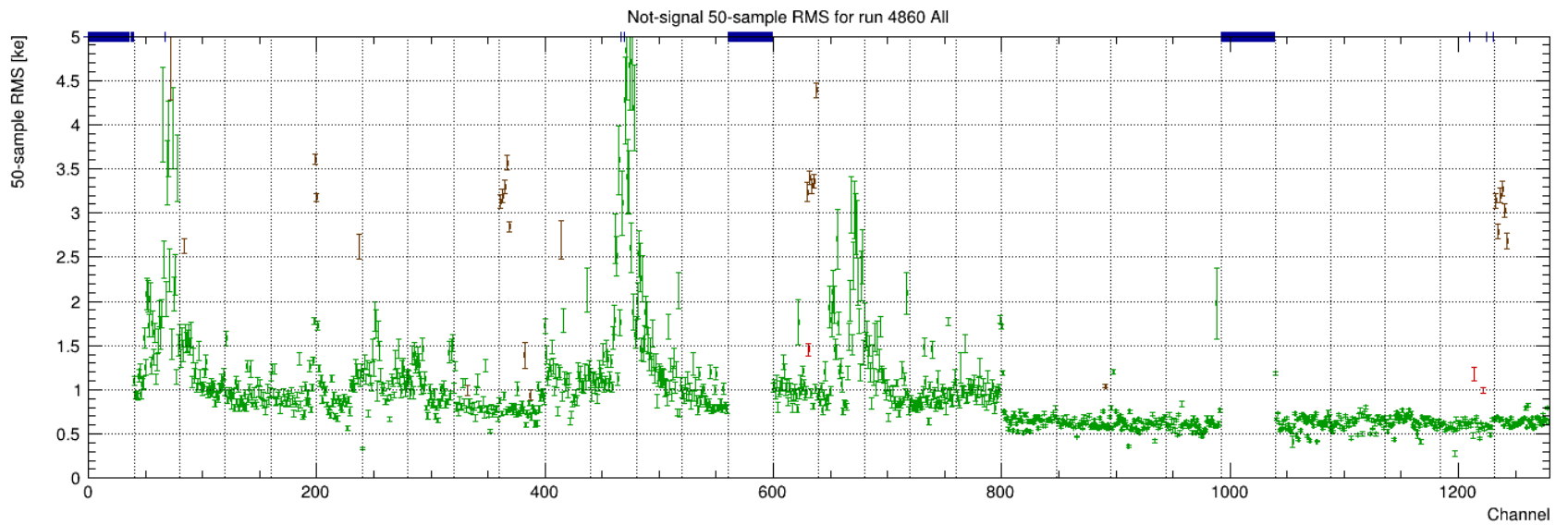
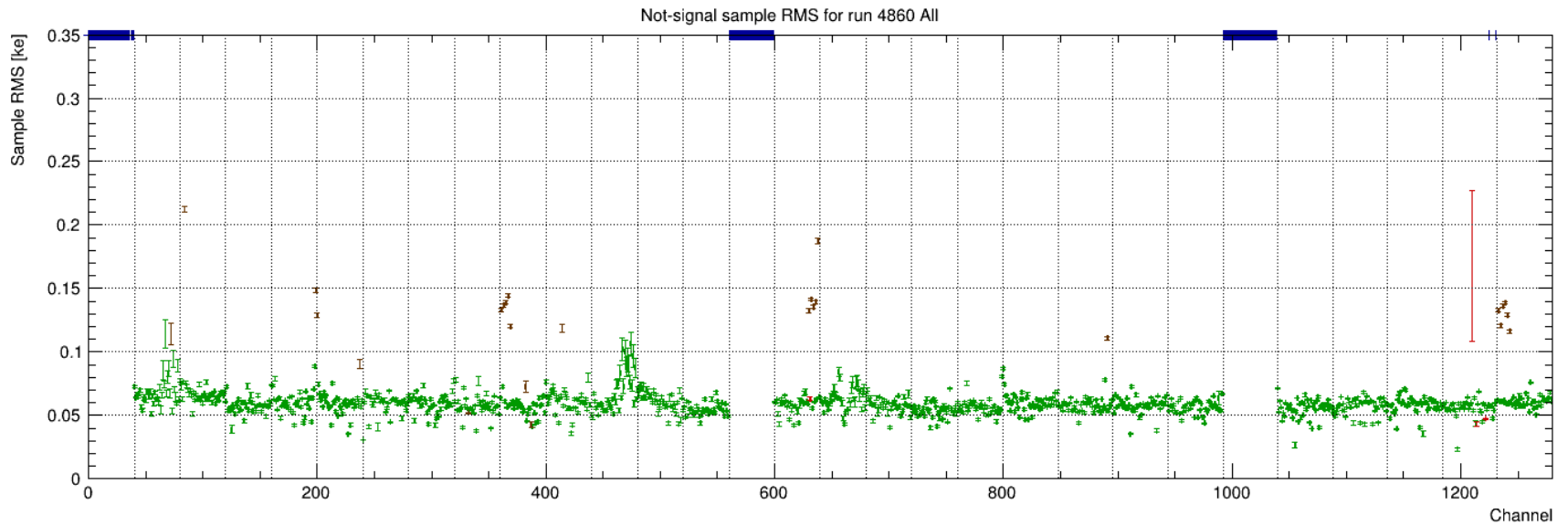


# Zoomed median noise vs. run



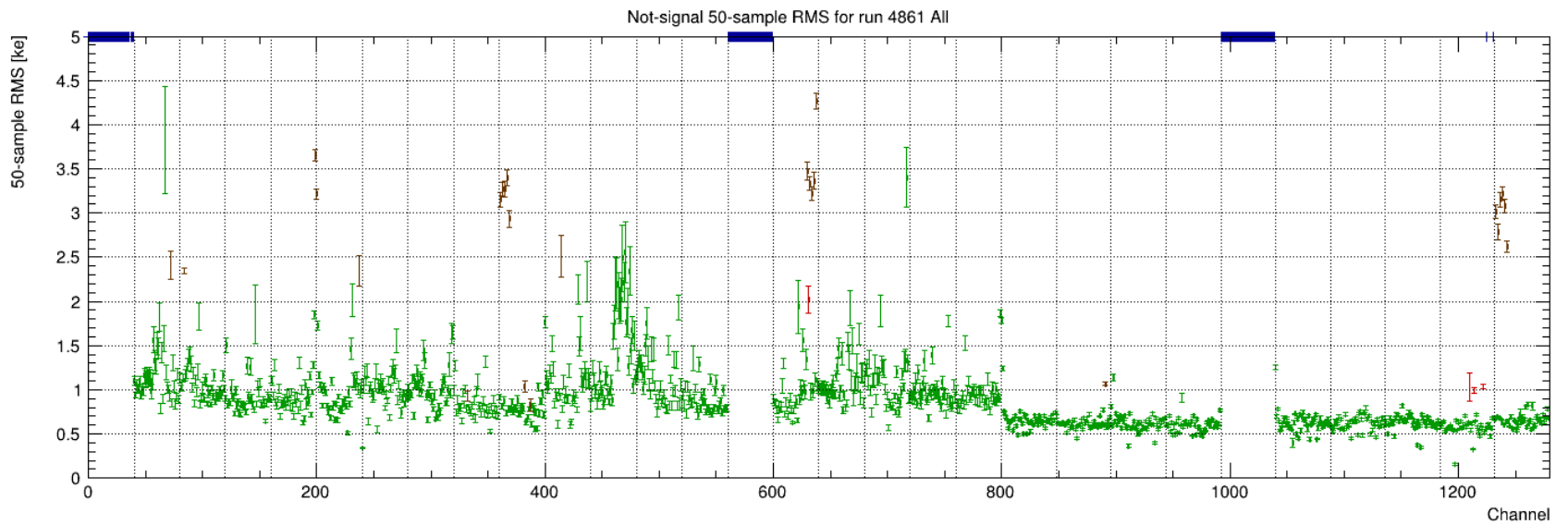
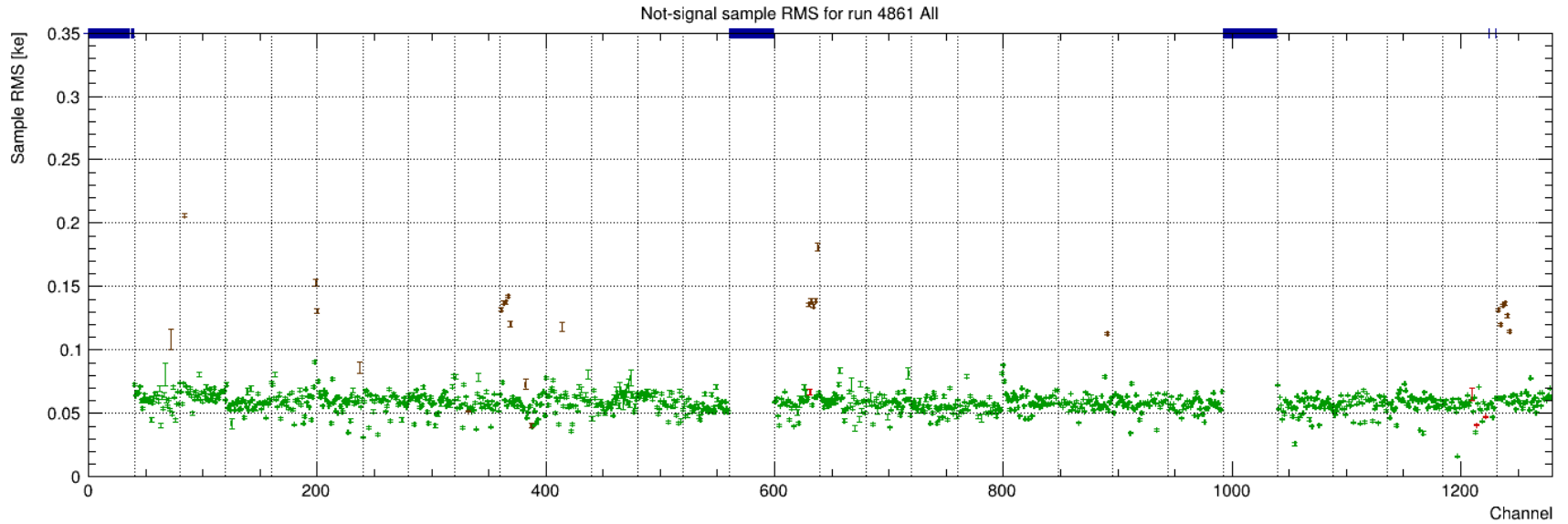
# Noise w/o CNR vs. channel by run

# 4860

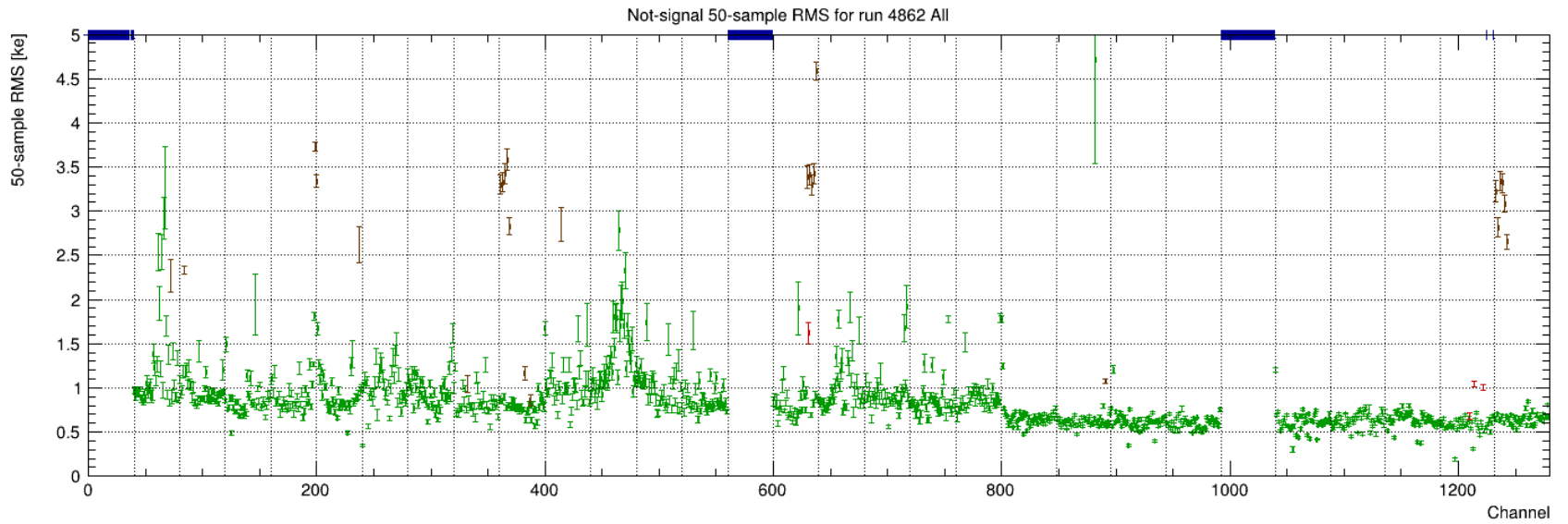
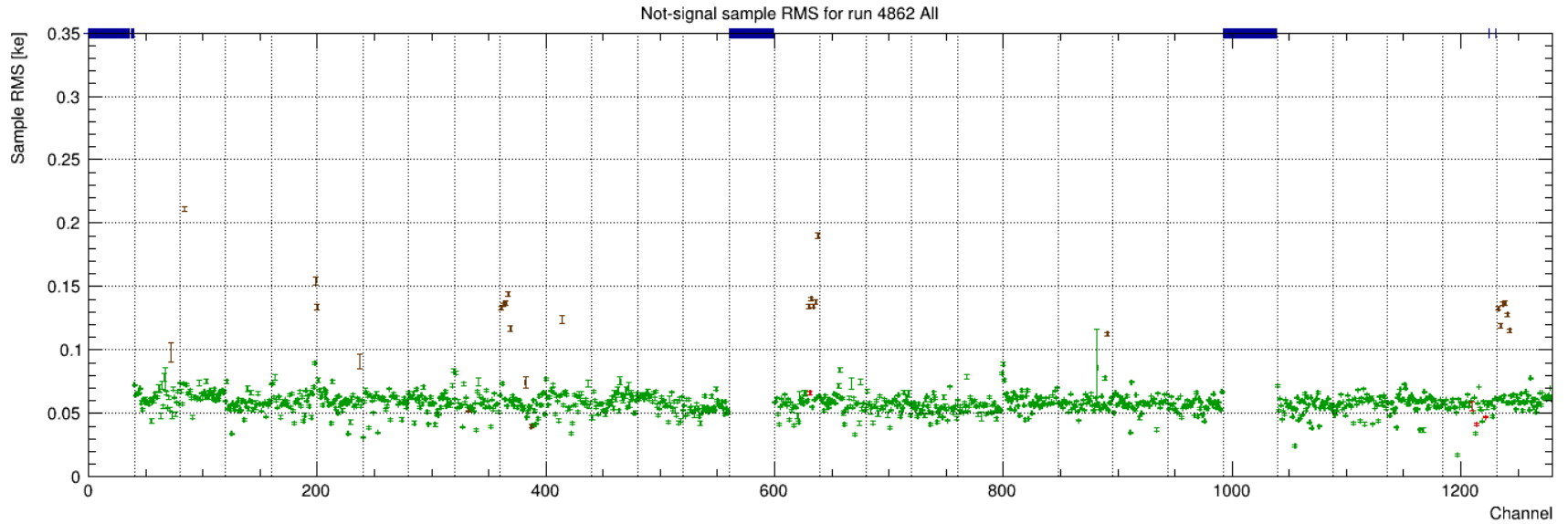




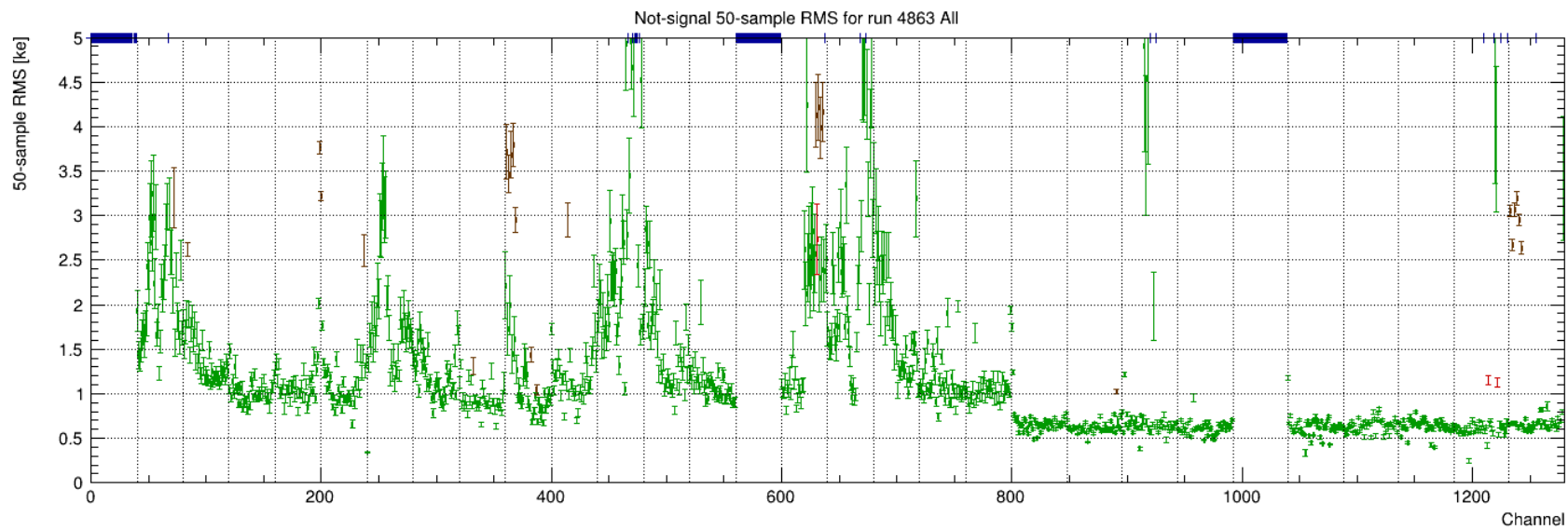
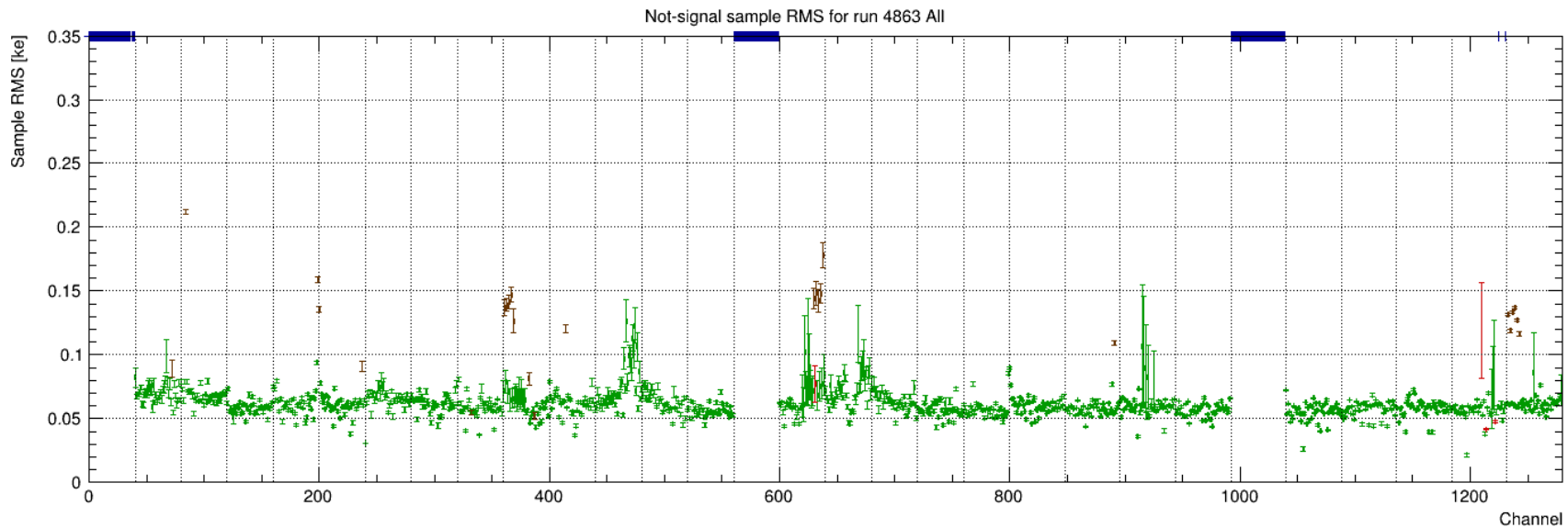
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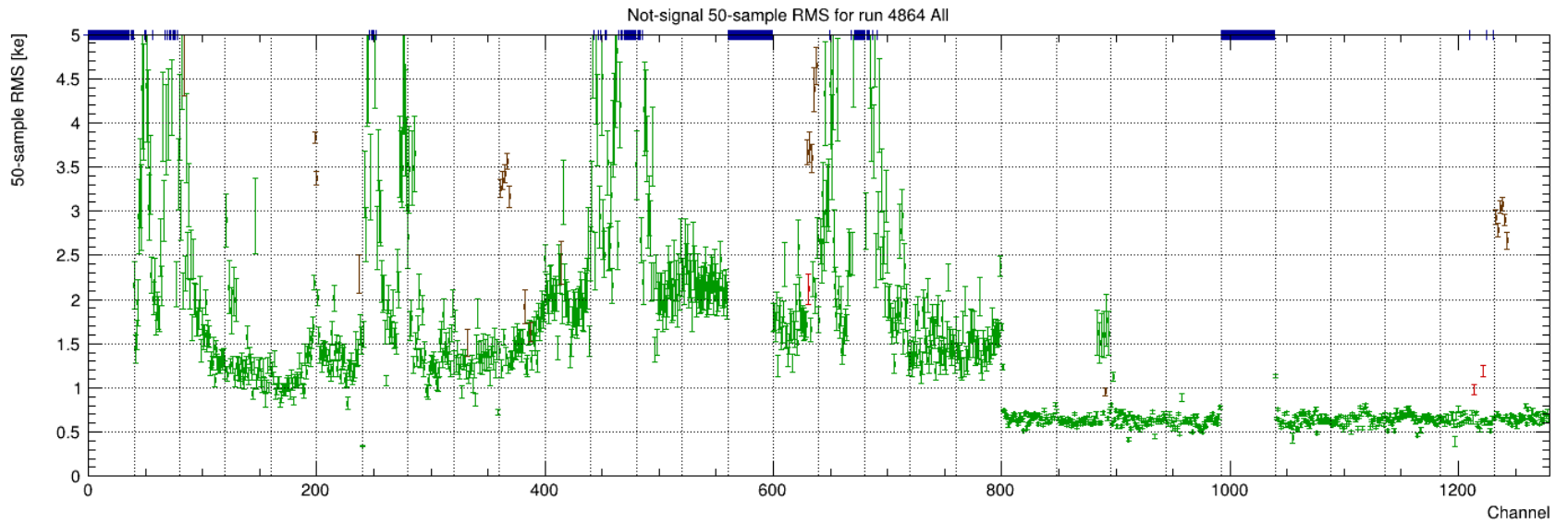
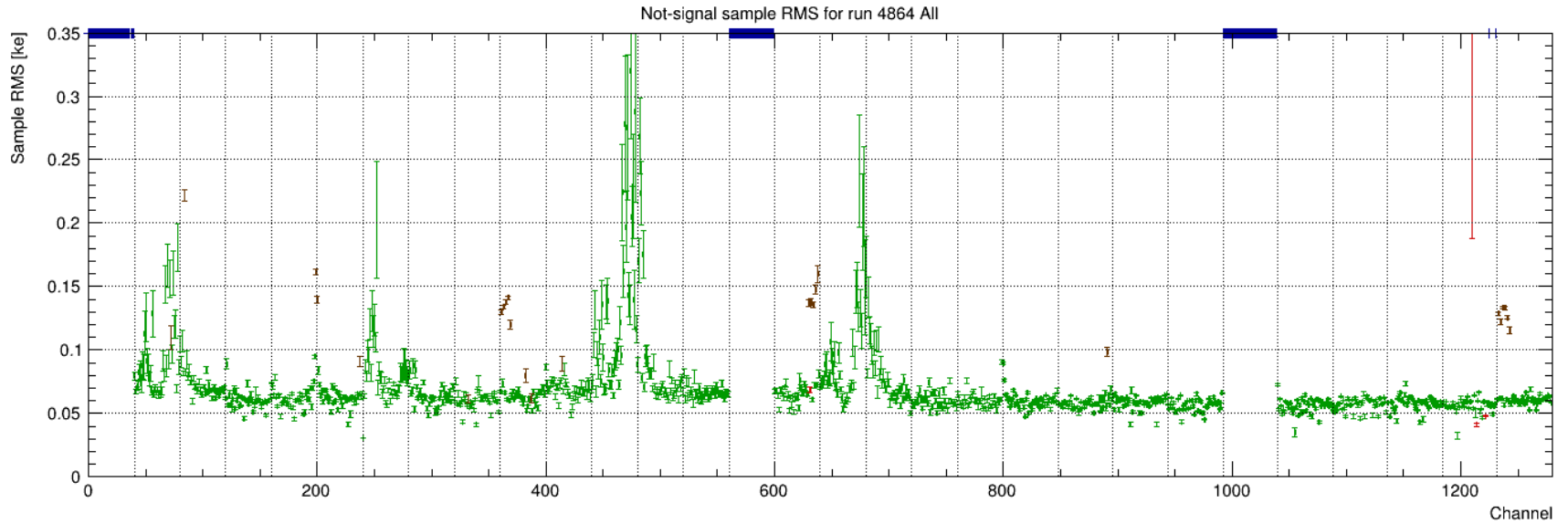
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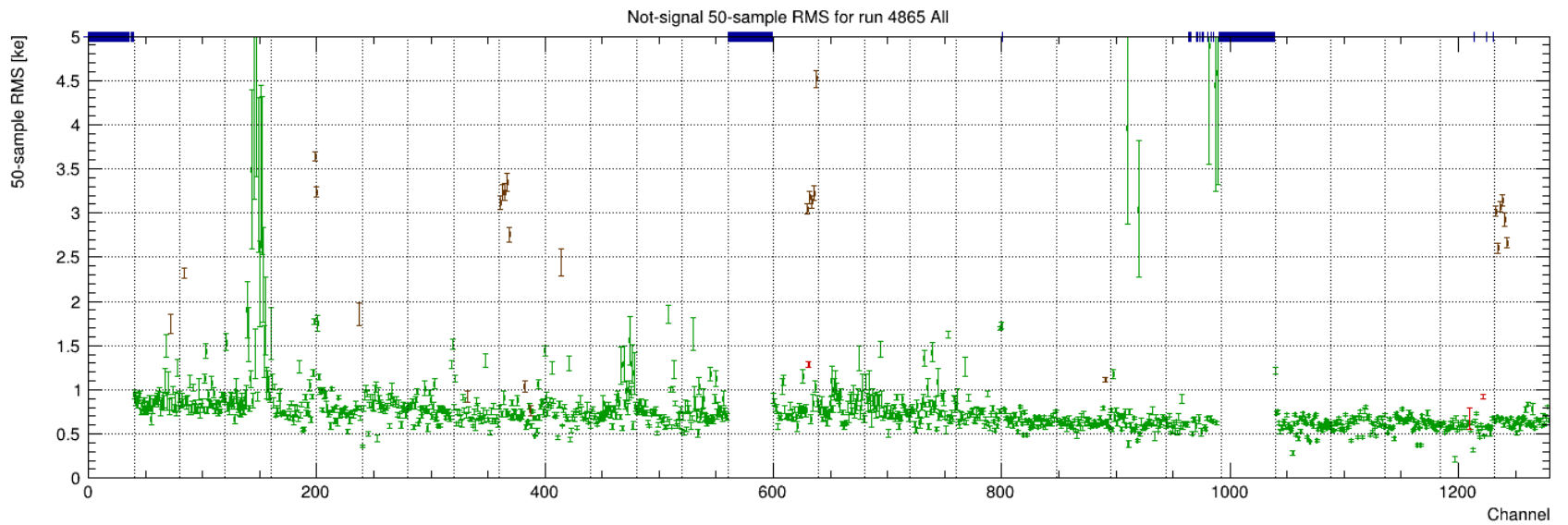
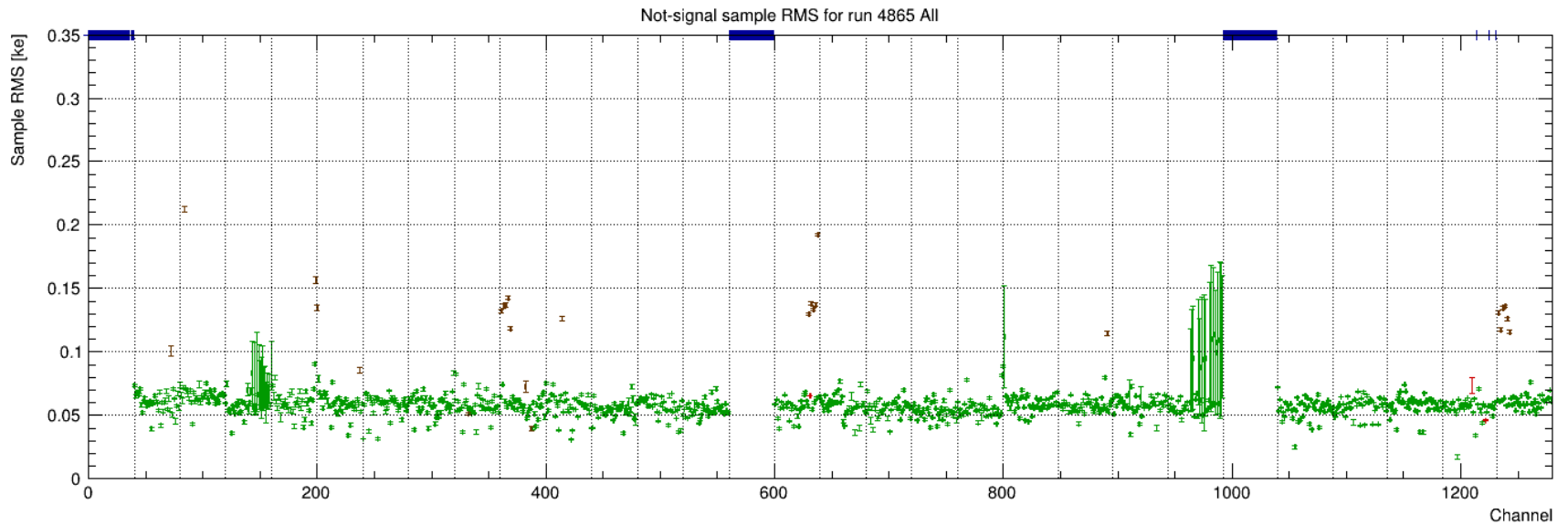
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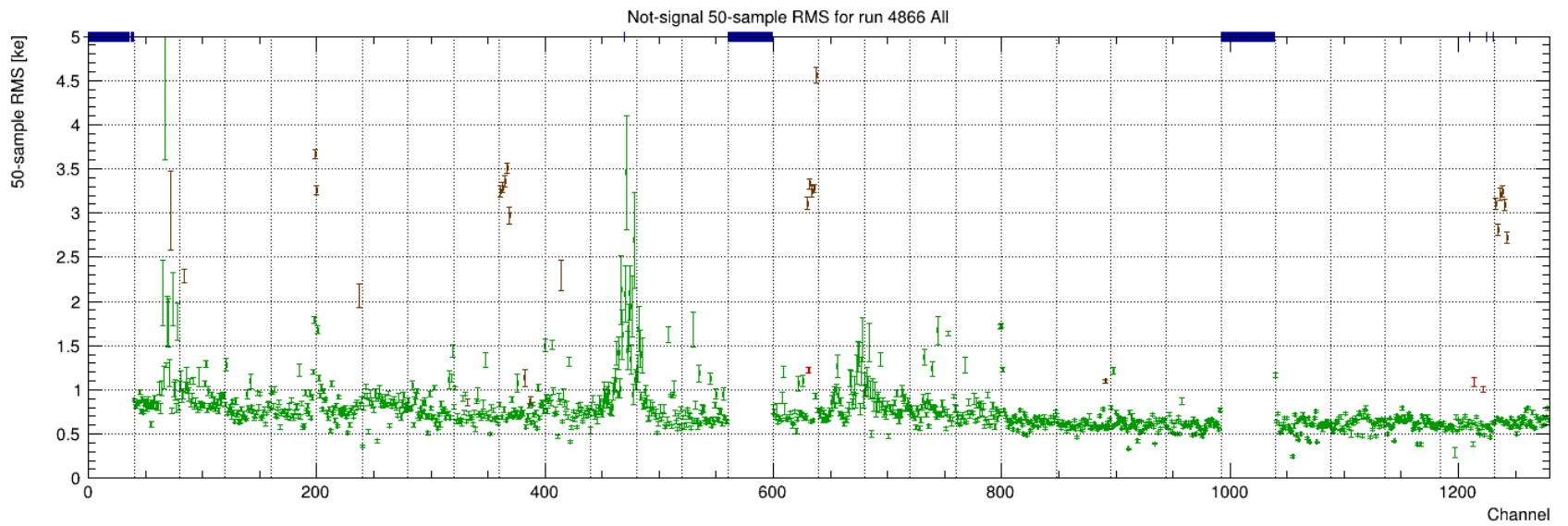
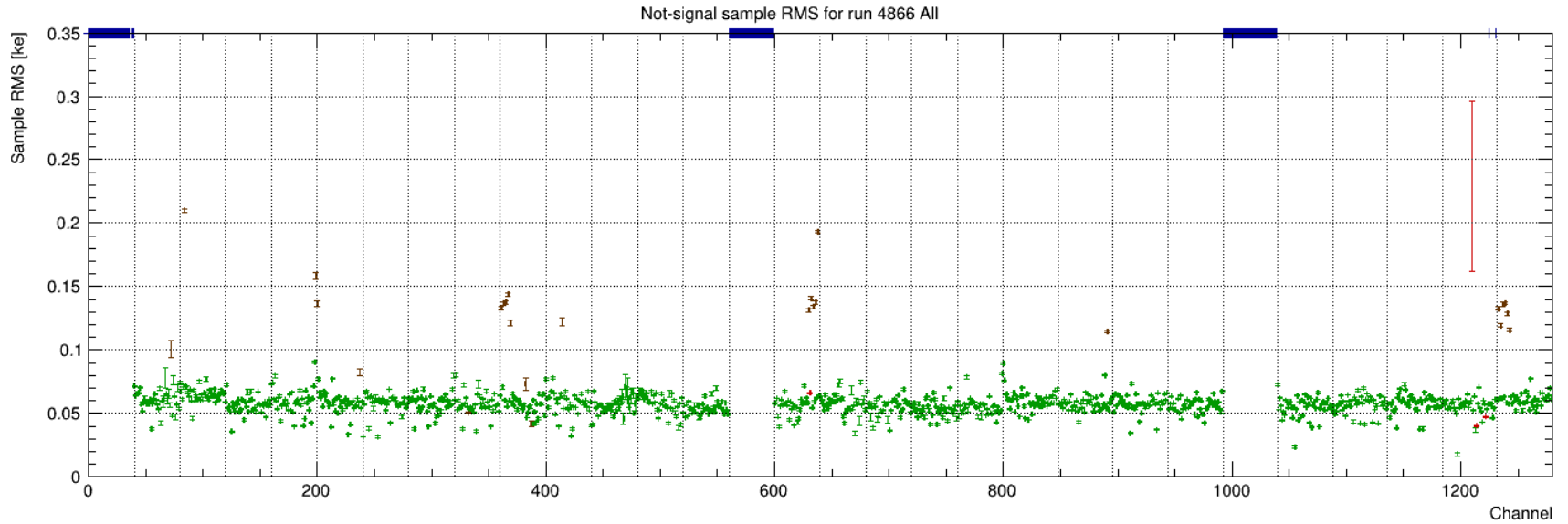
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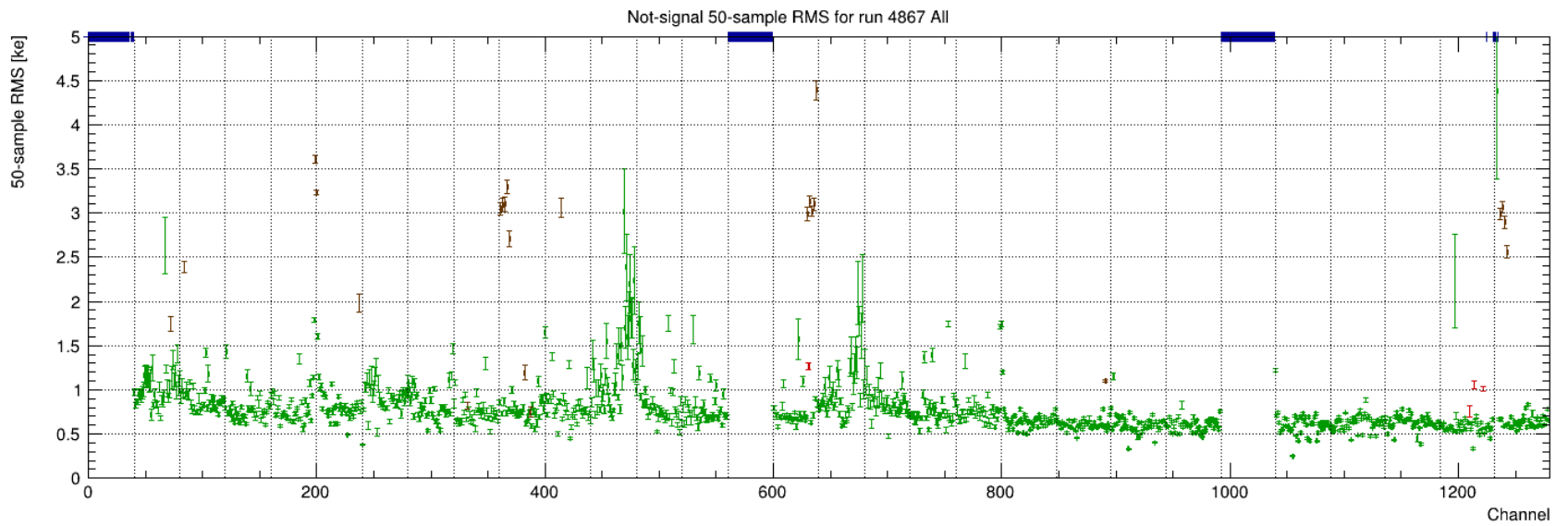
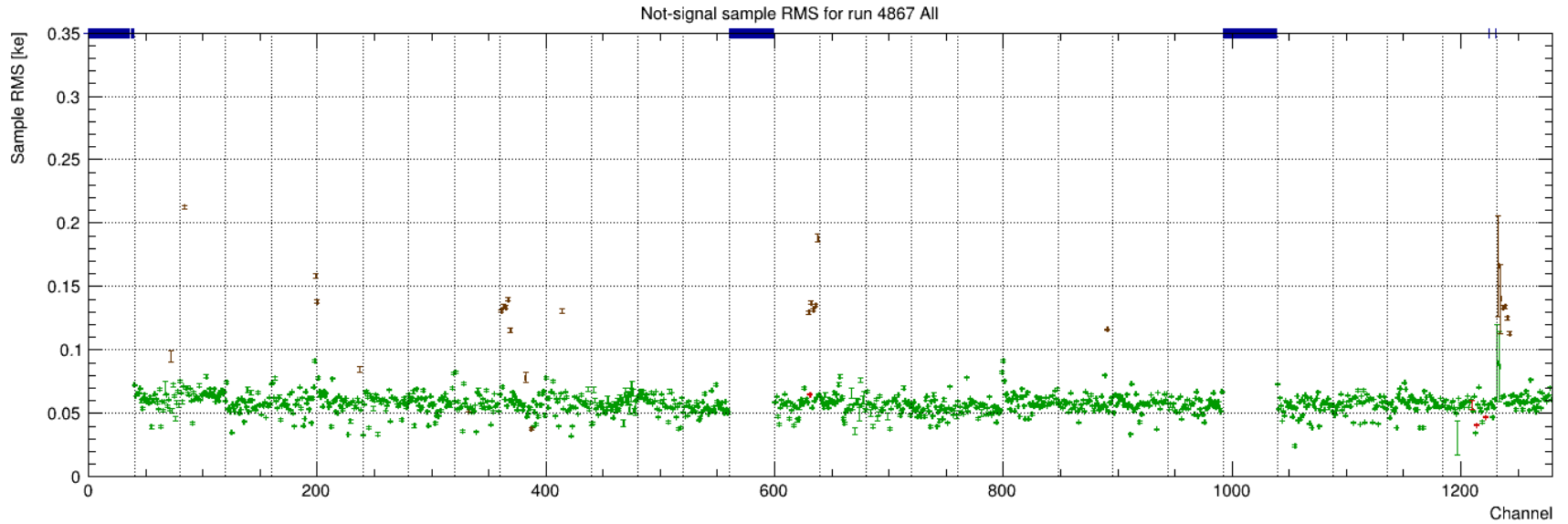
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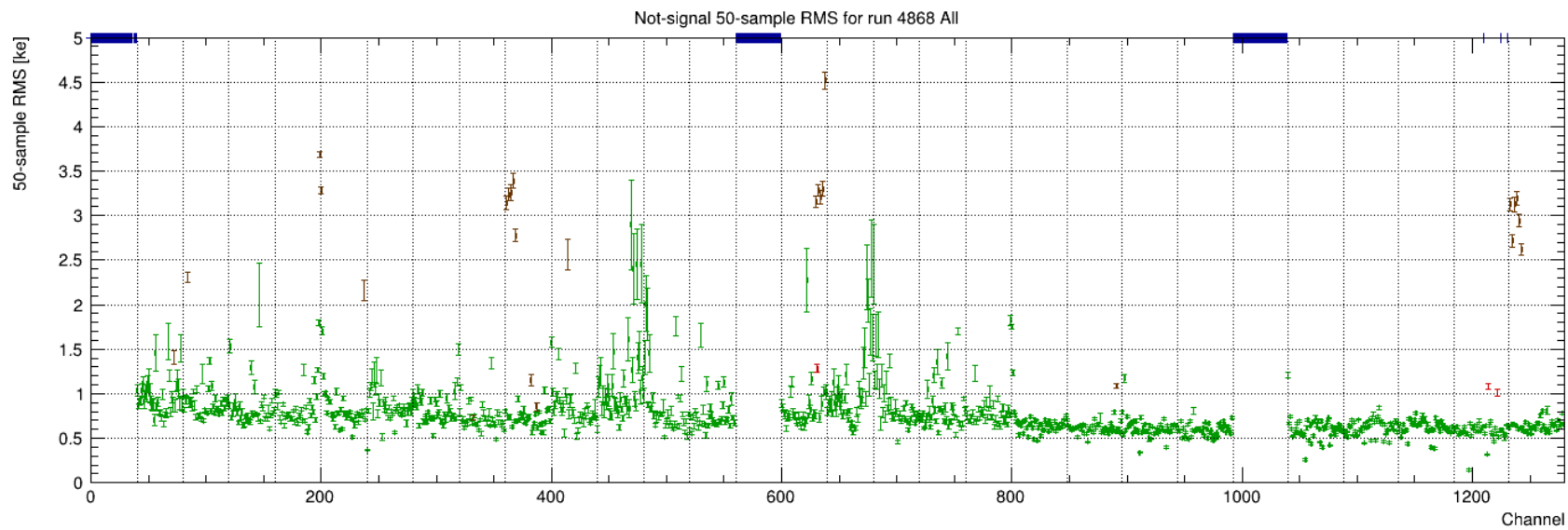
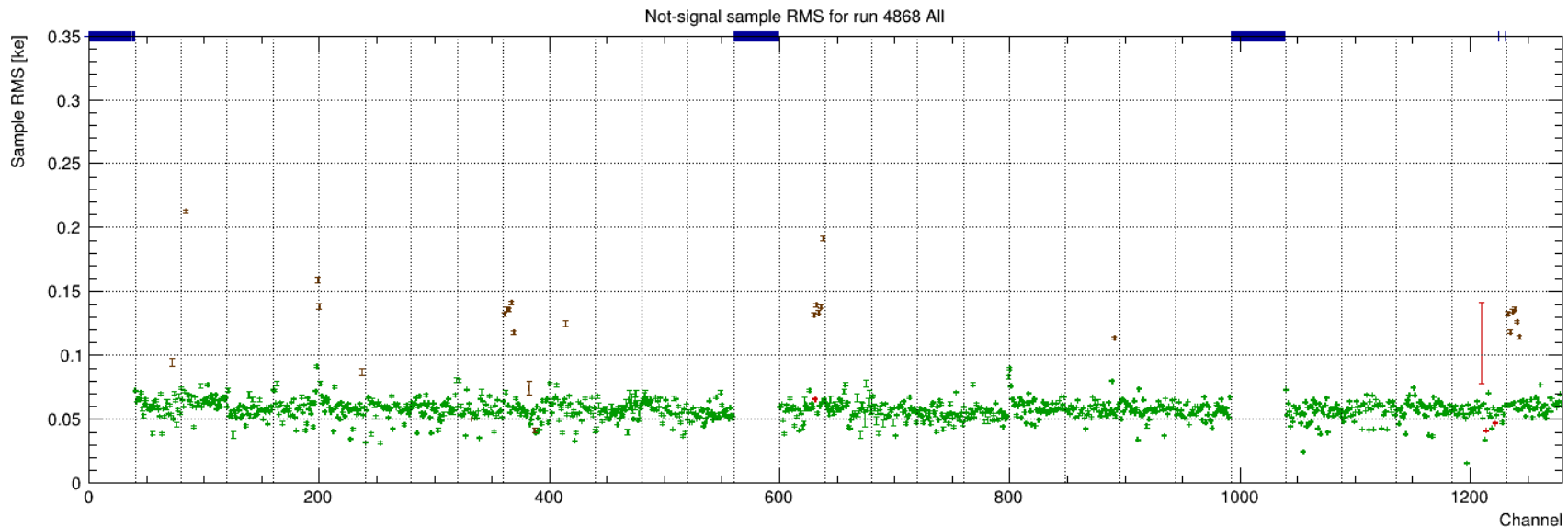
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# 4867

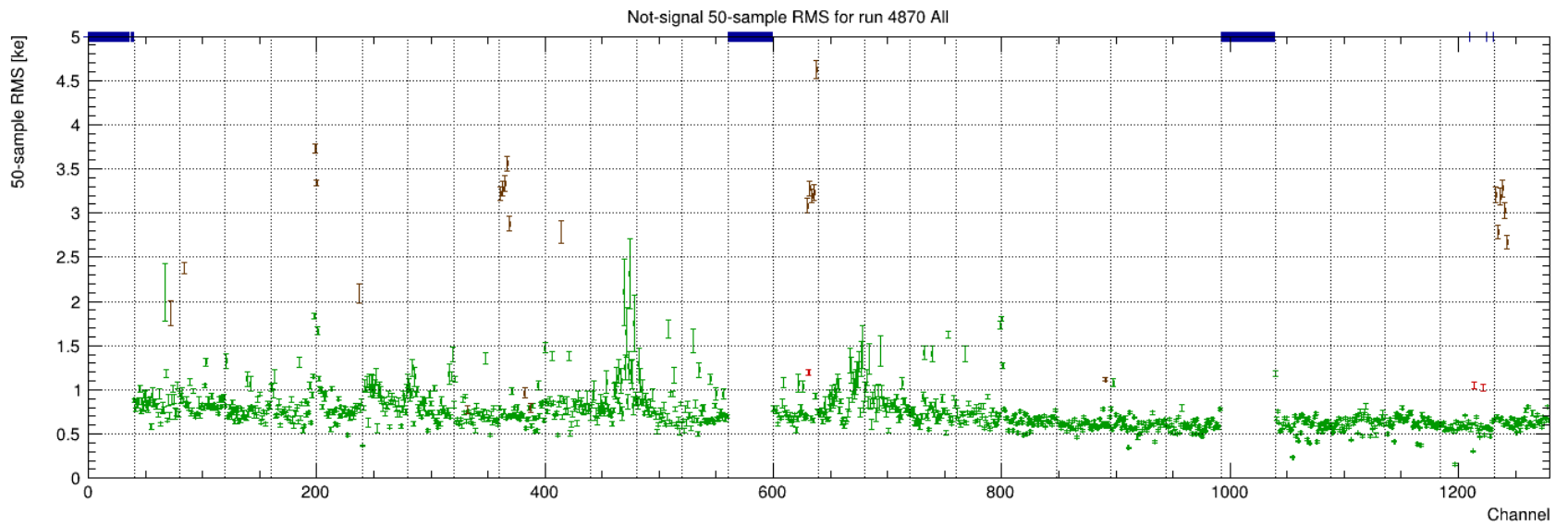
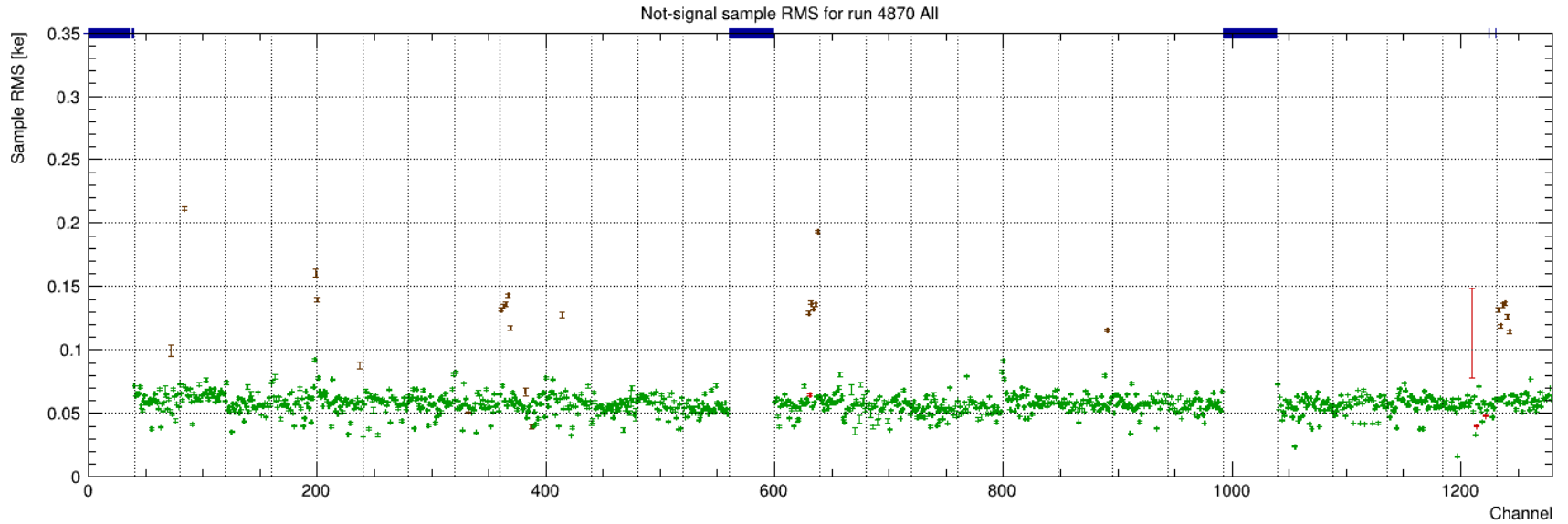


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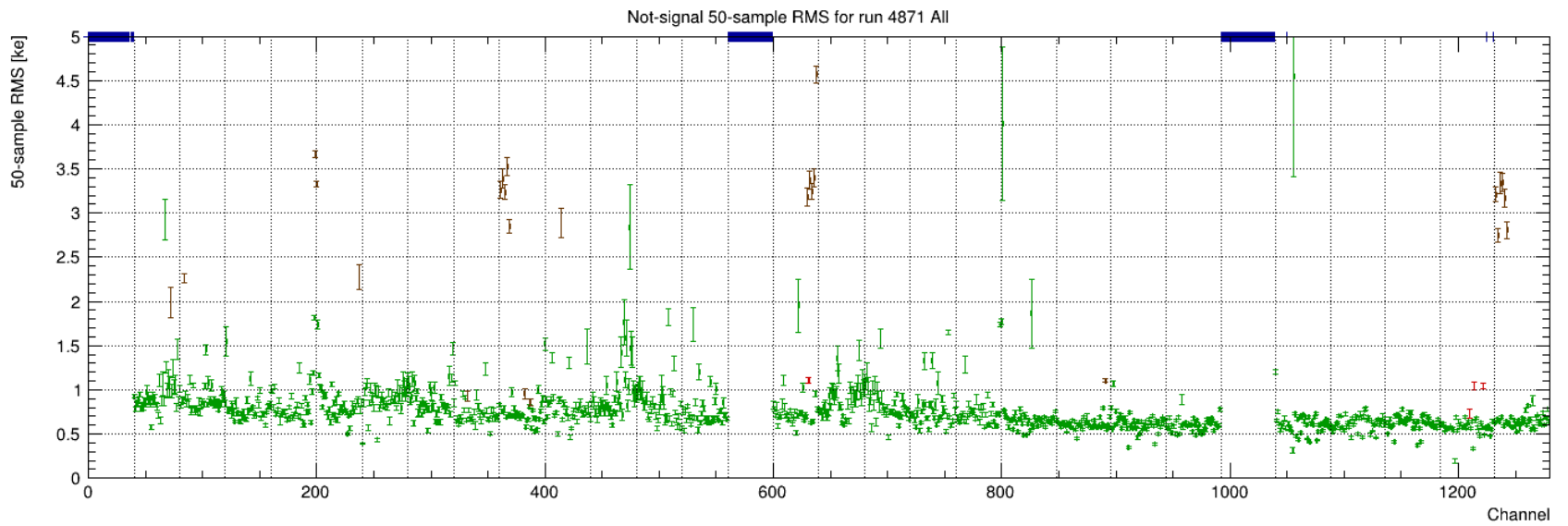
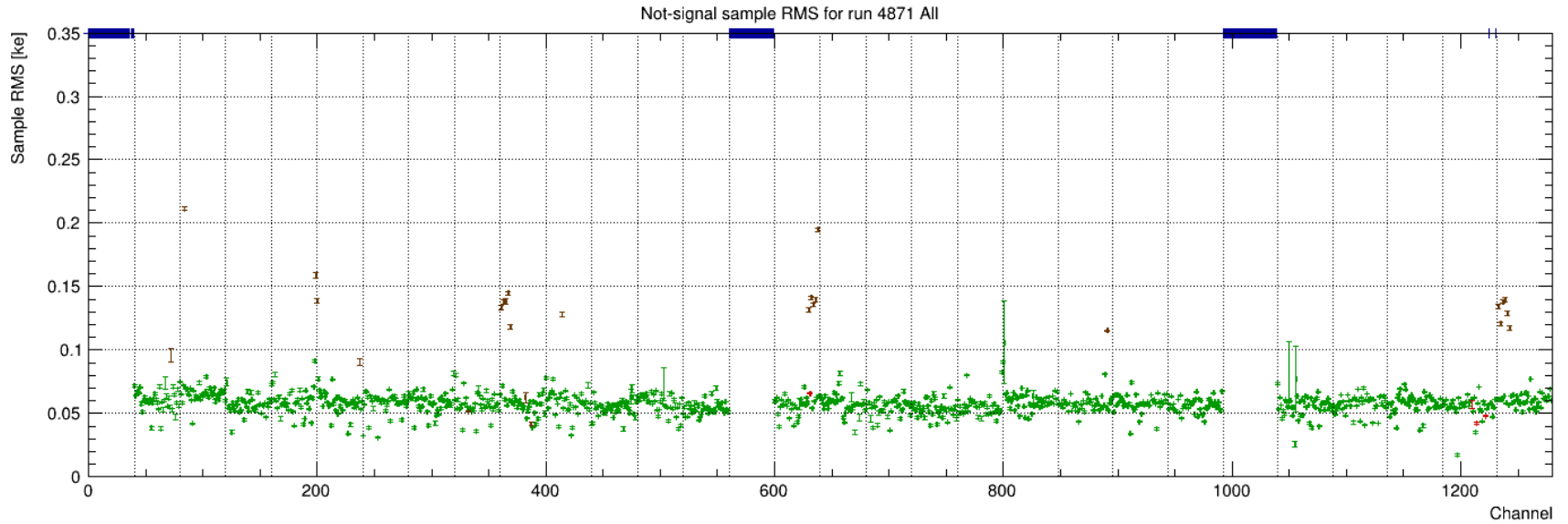




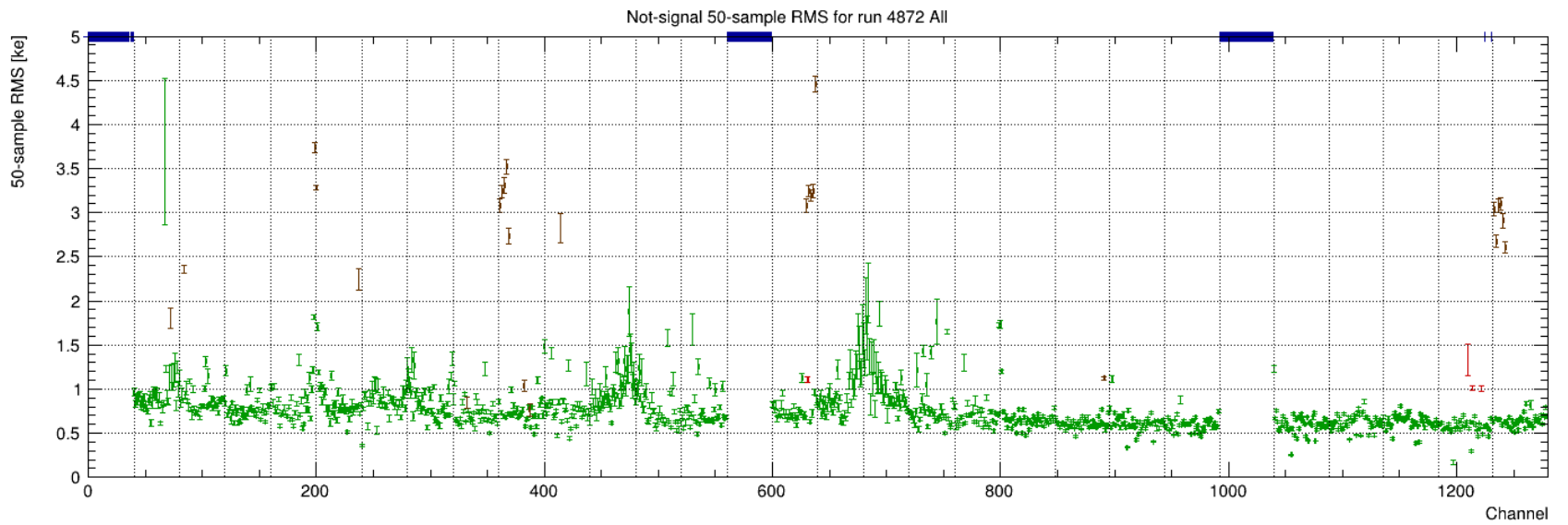
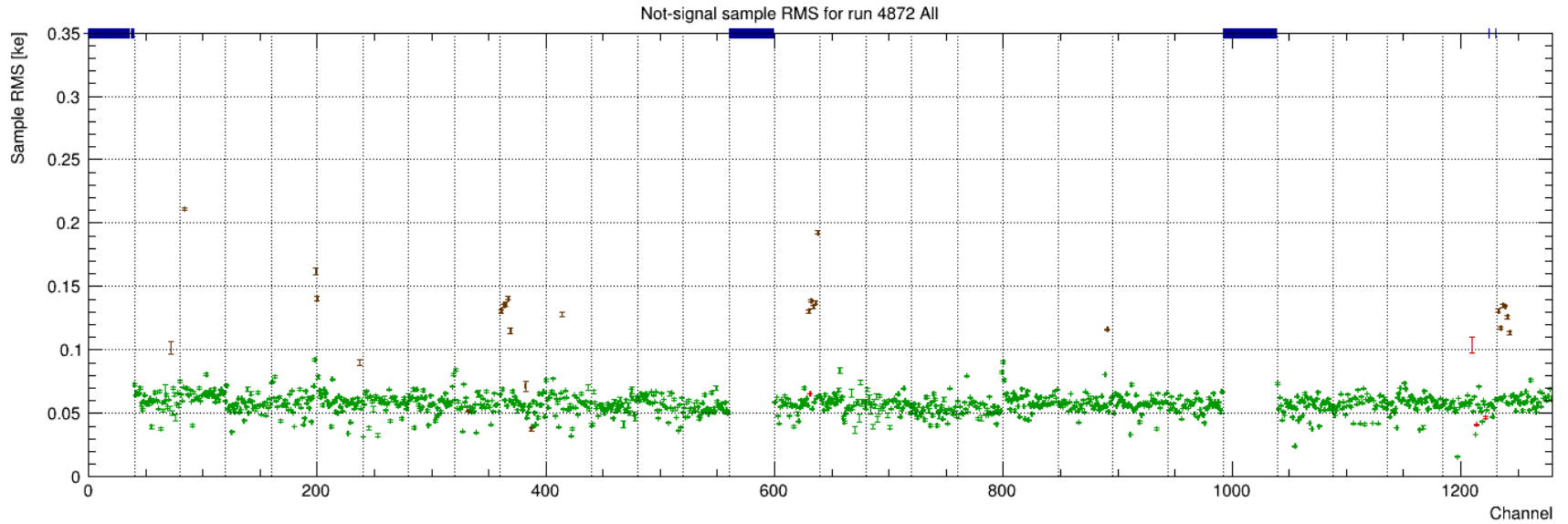
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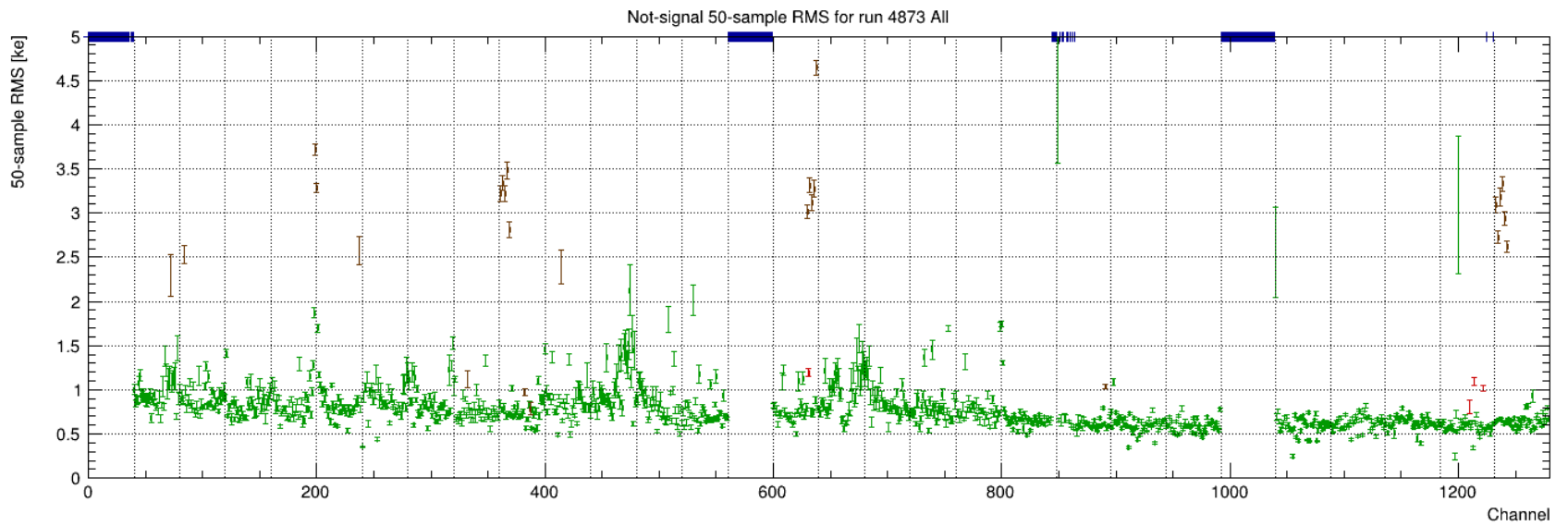
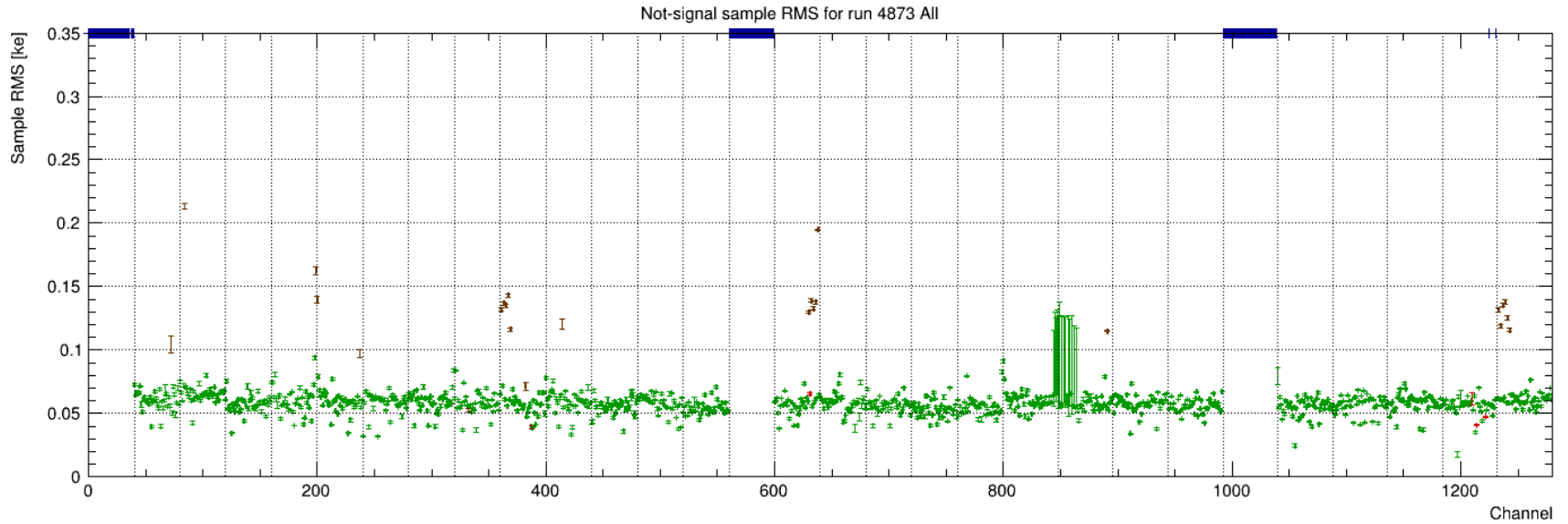
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# 4872

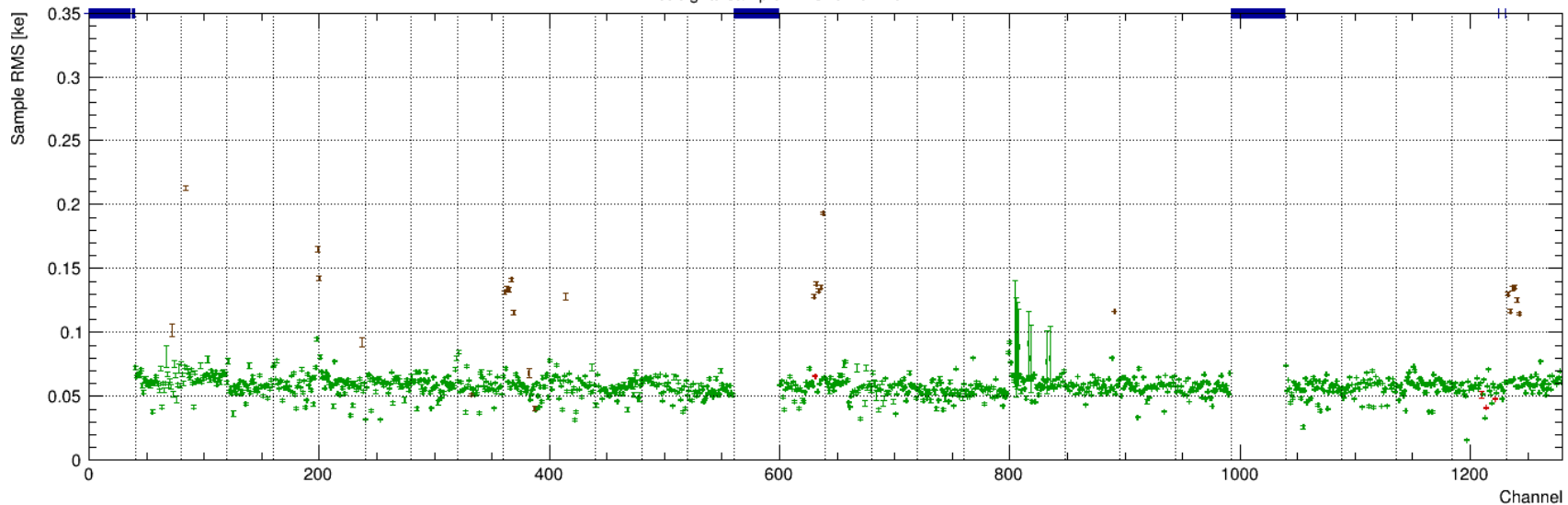


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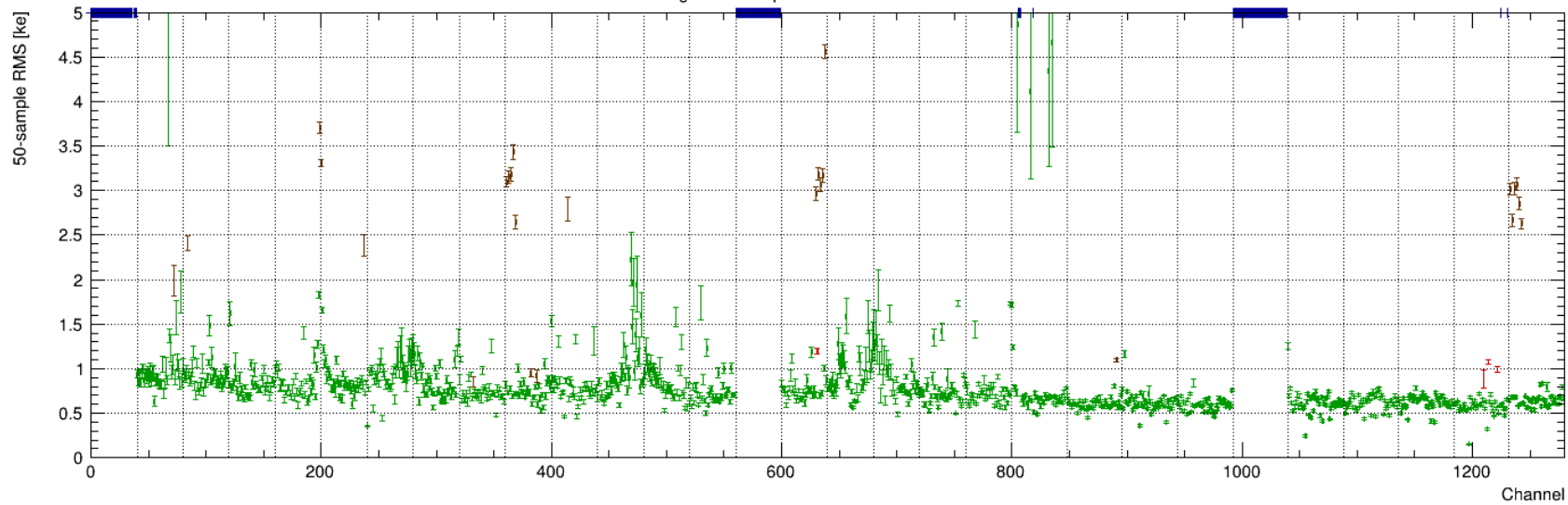


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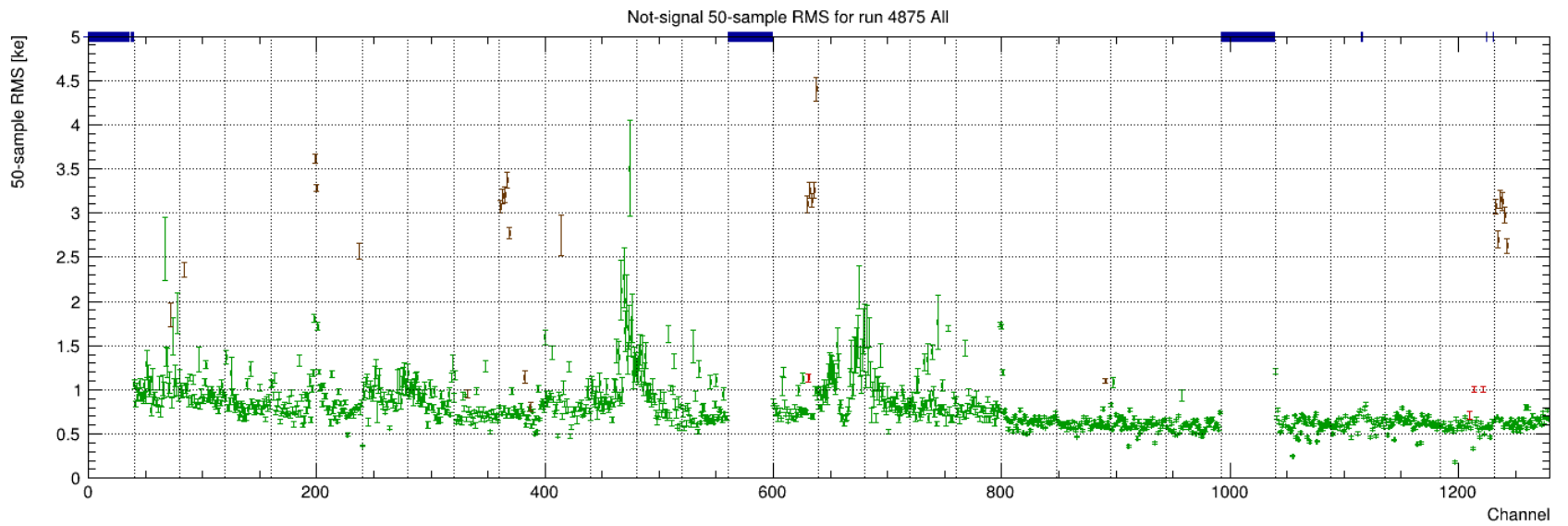
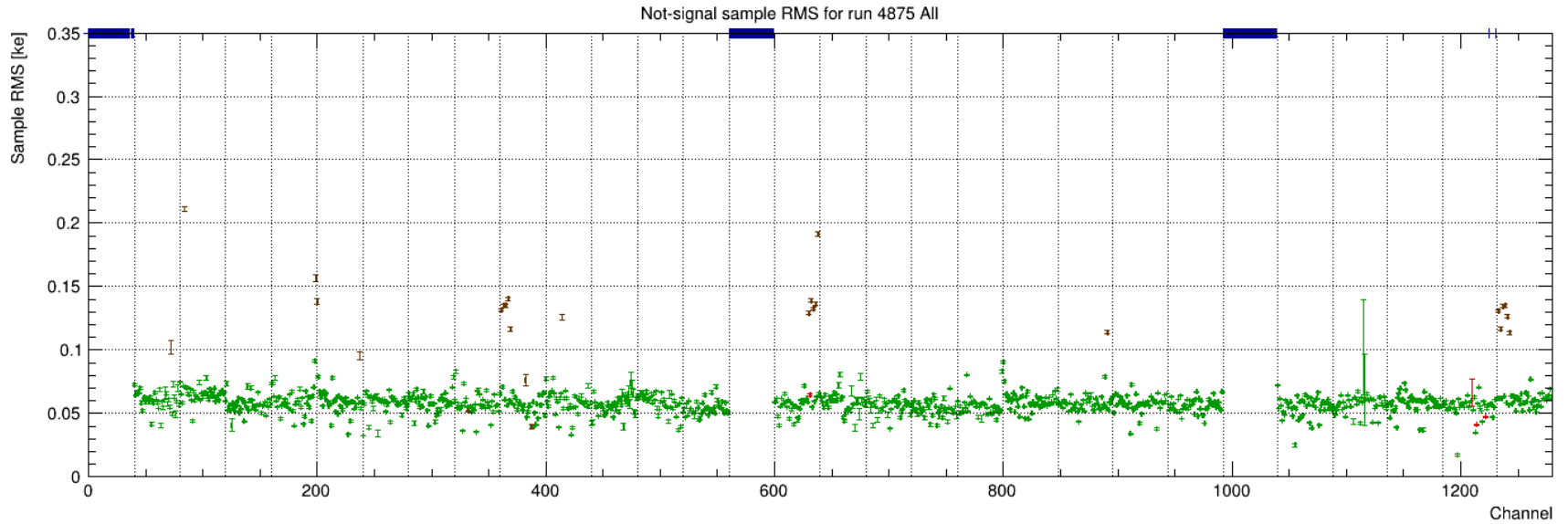
Not-signal sample RMS for run 4874 All



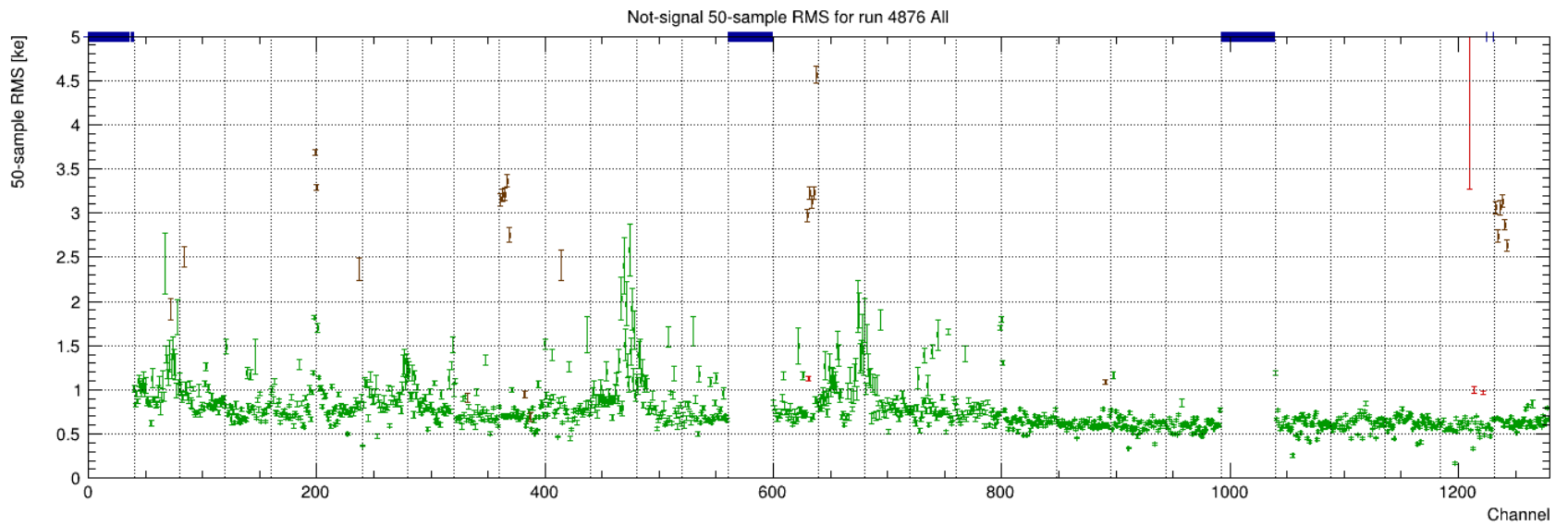
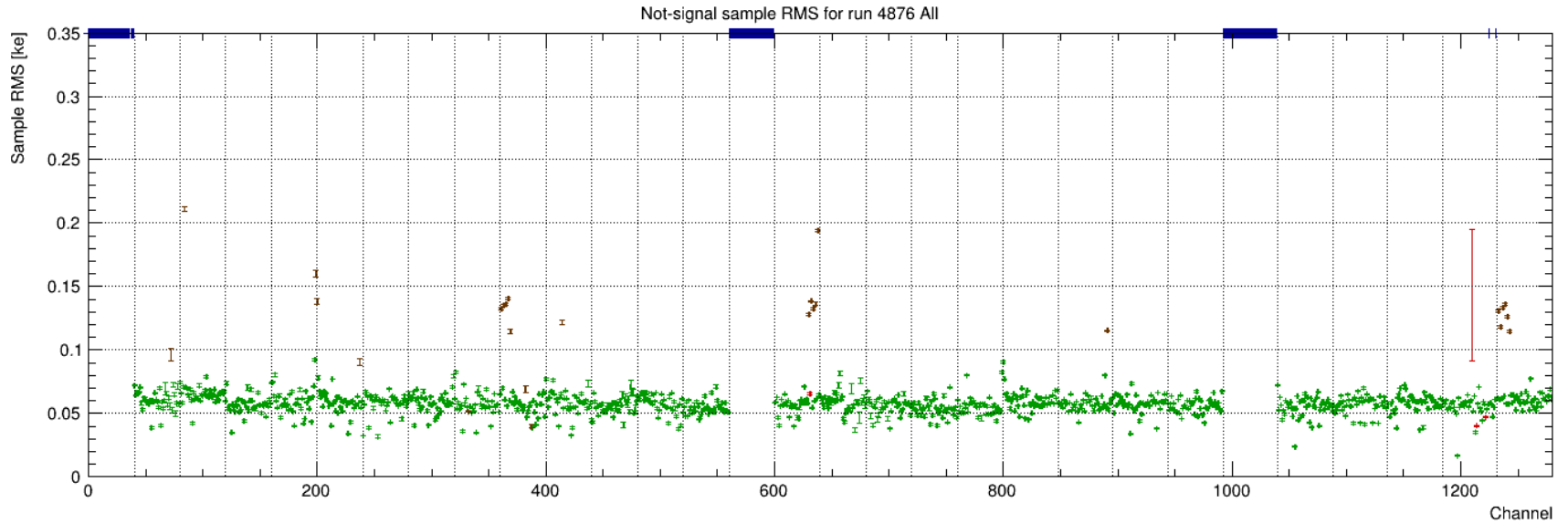
Not-signal 50-sample RMS for run 4874 All



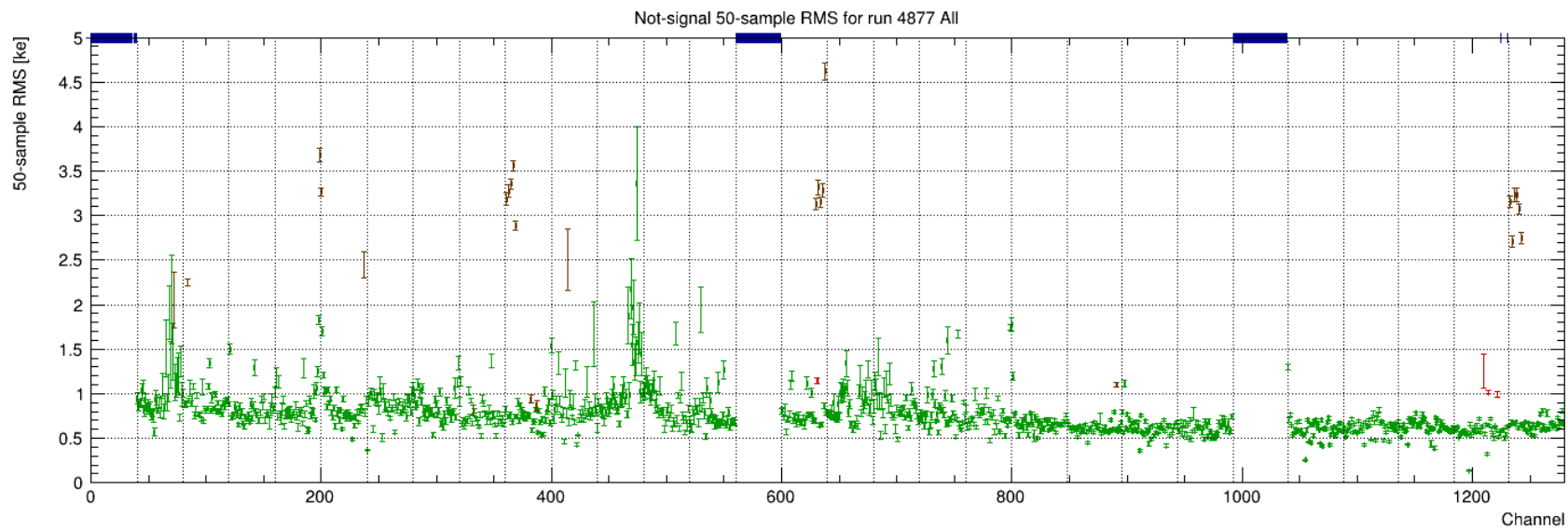
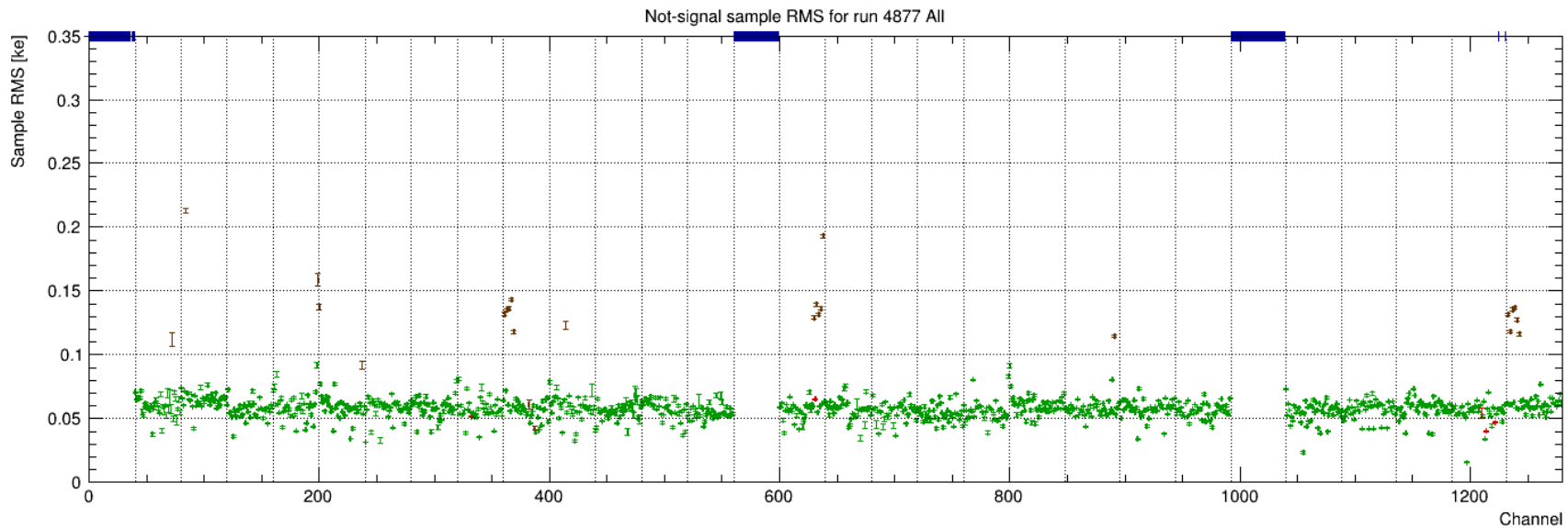
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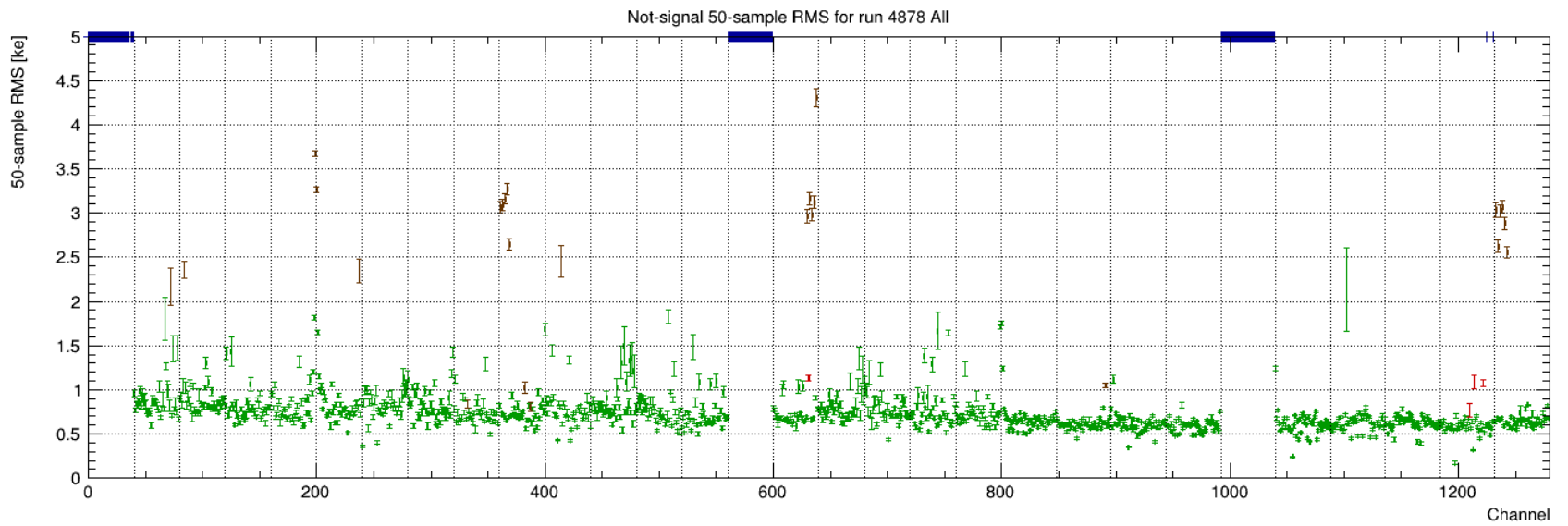
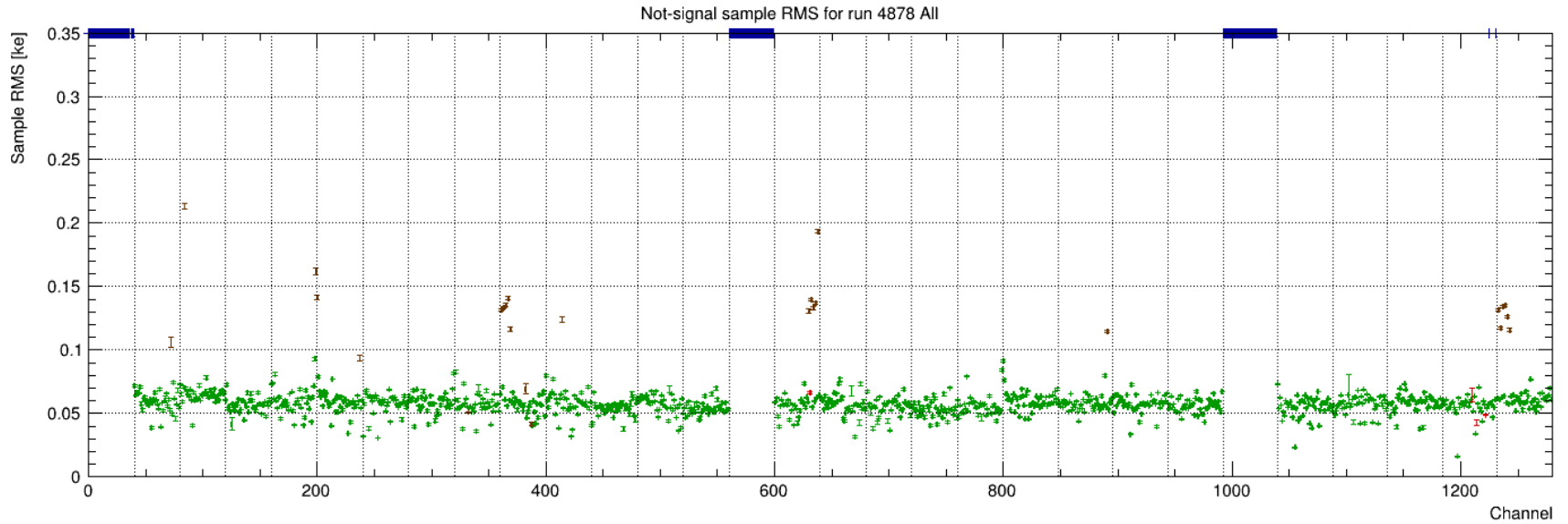


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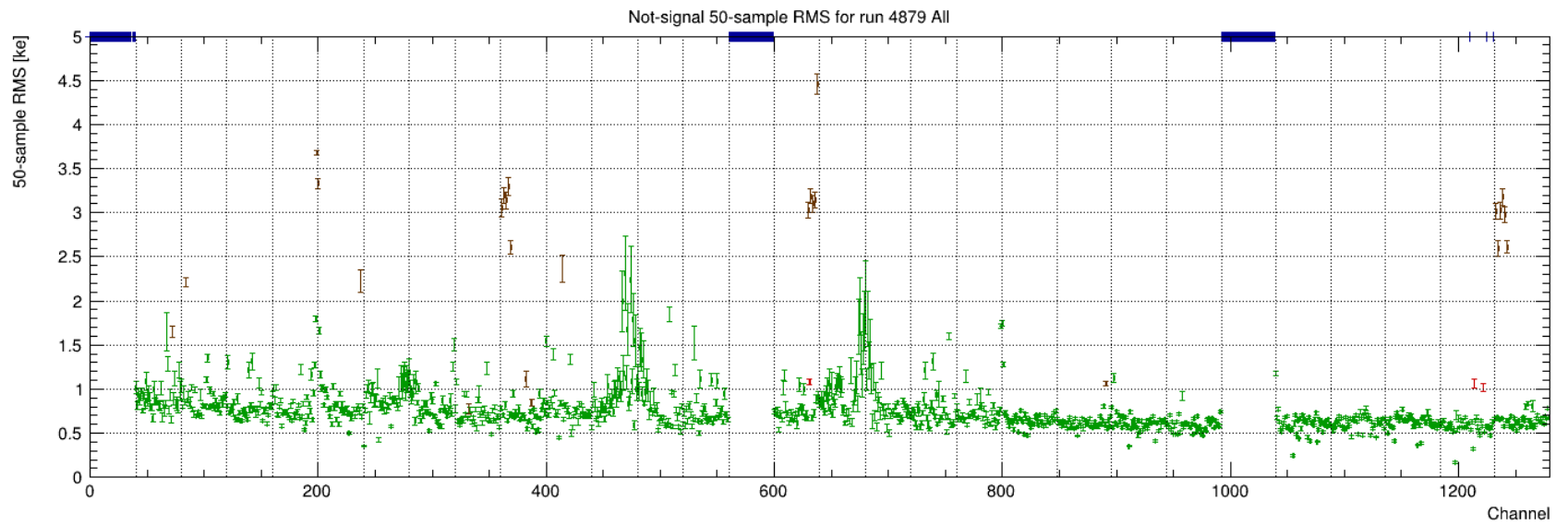
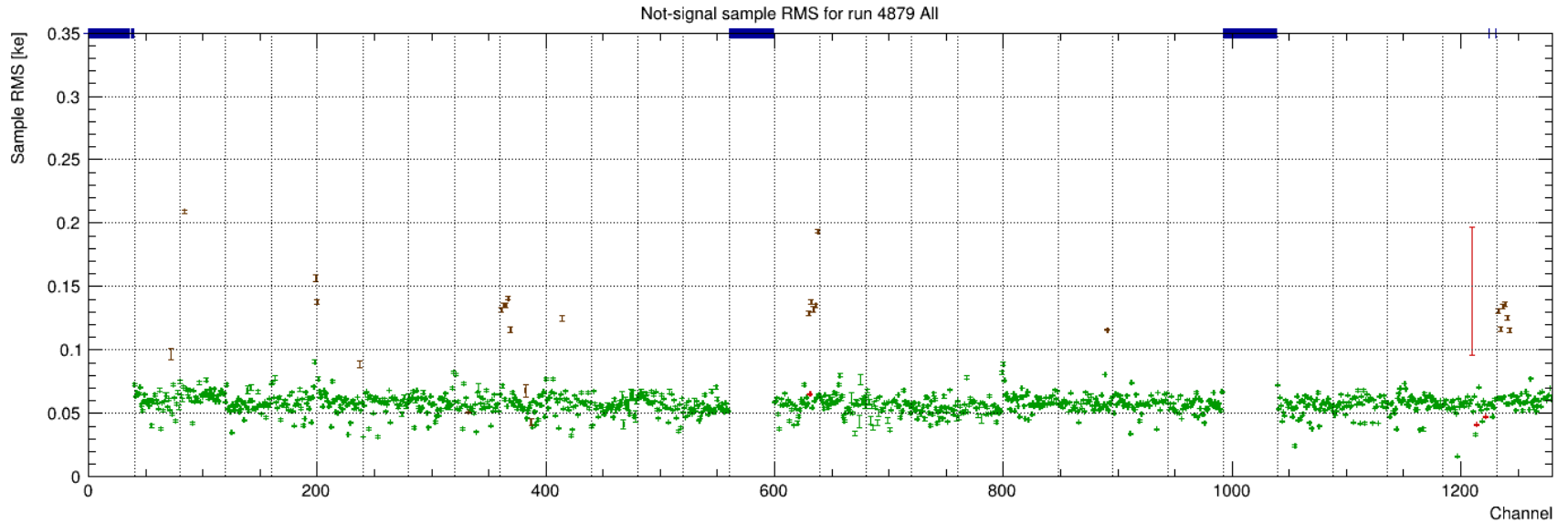




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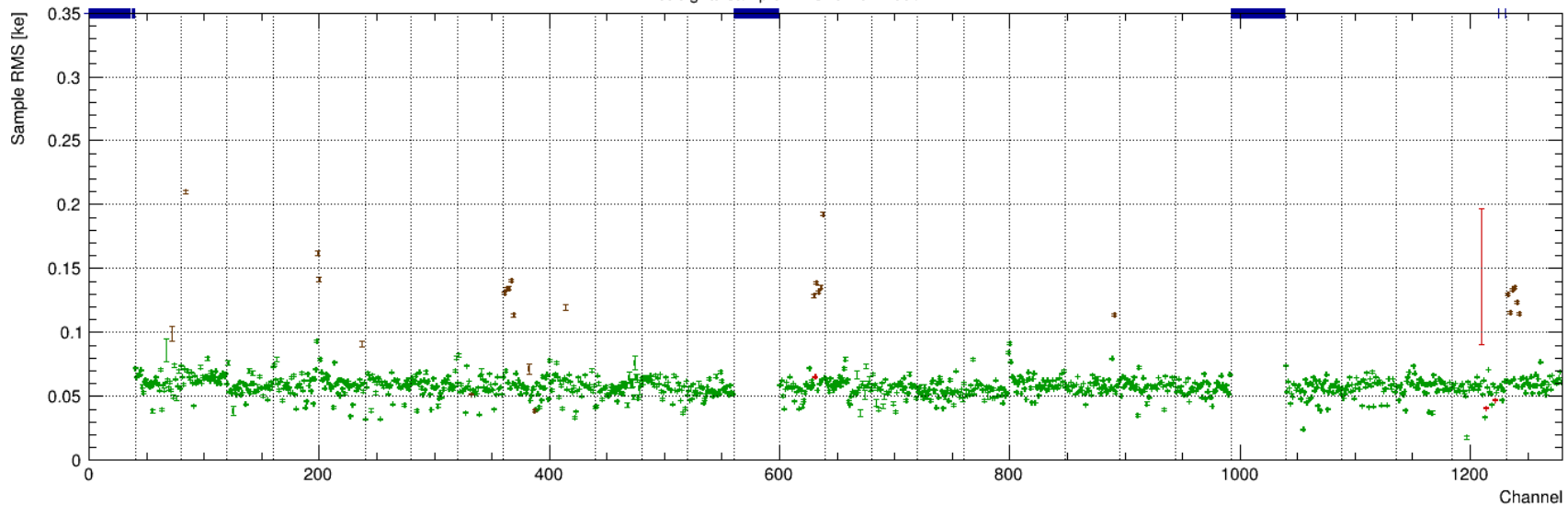


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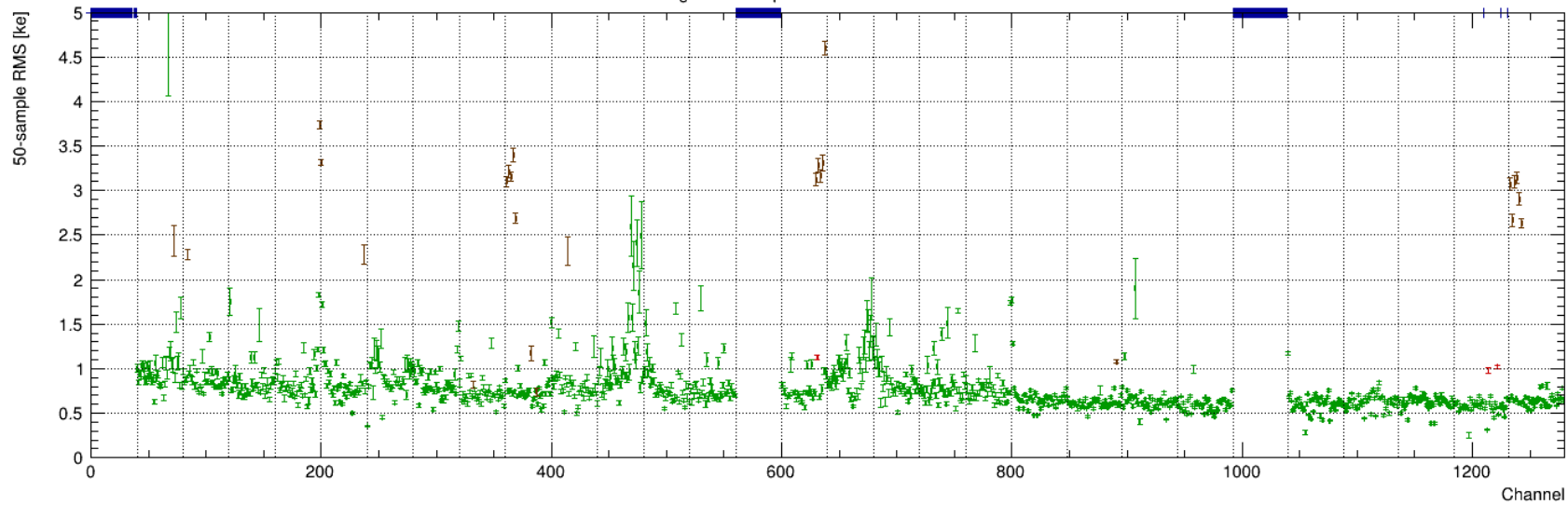


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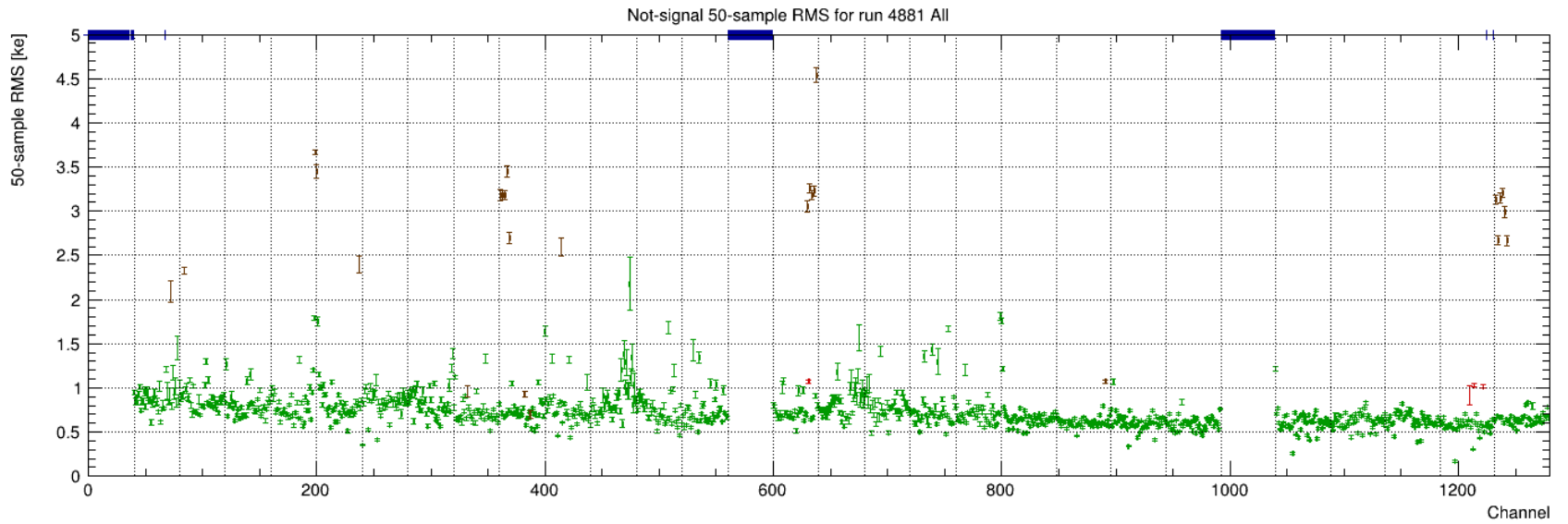
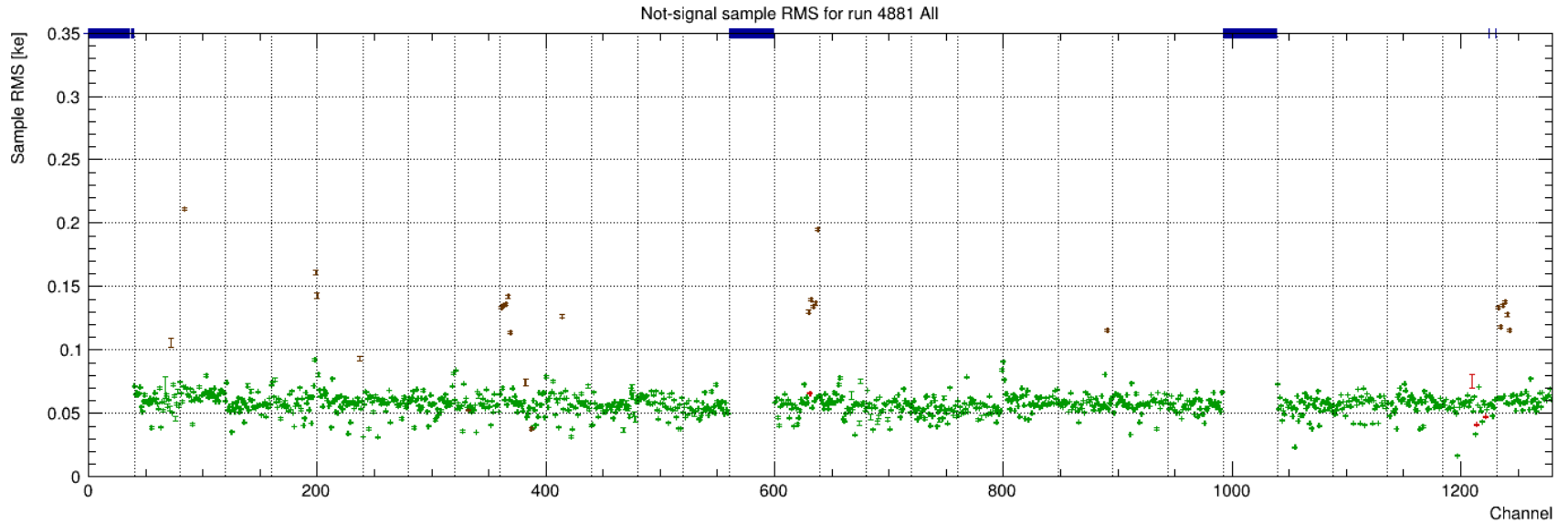
Not-signal sample RMS for run 4880 All



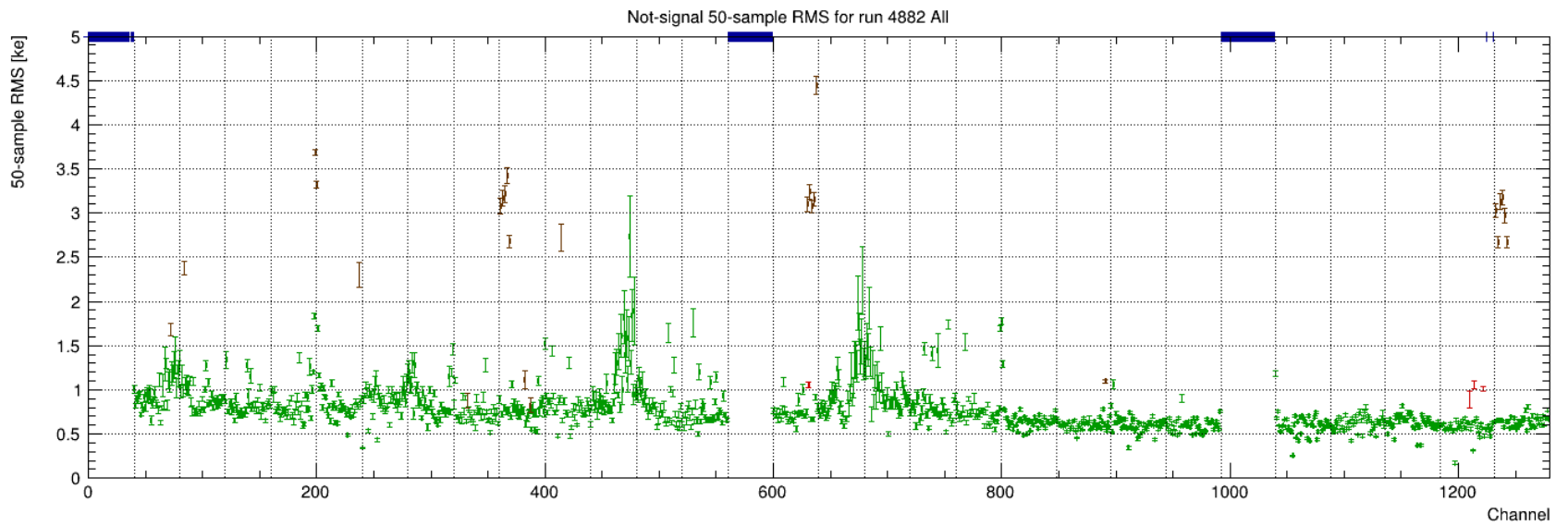
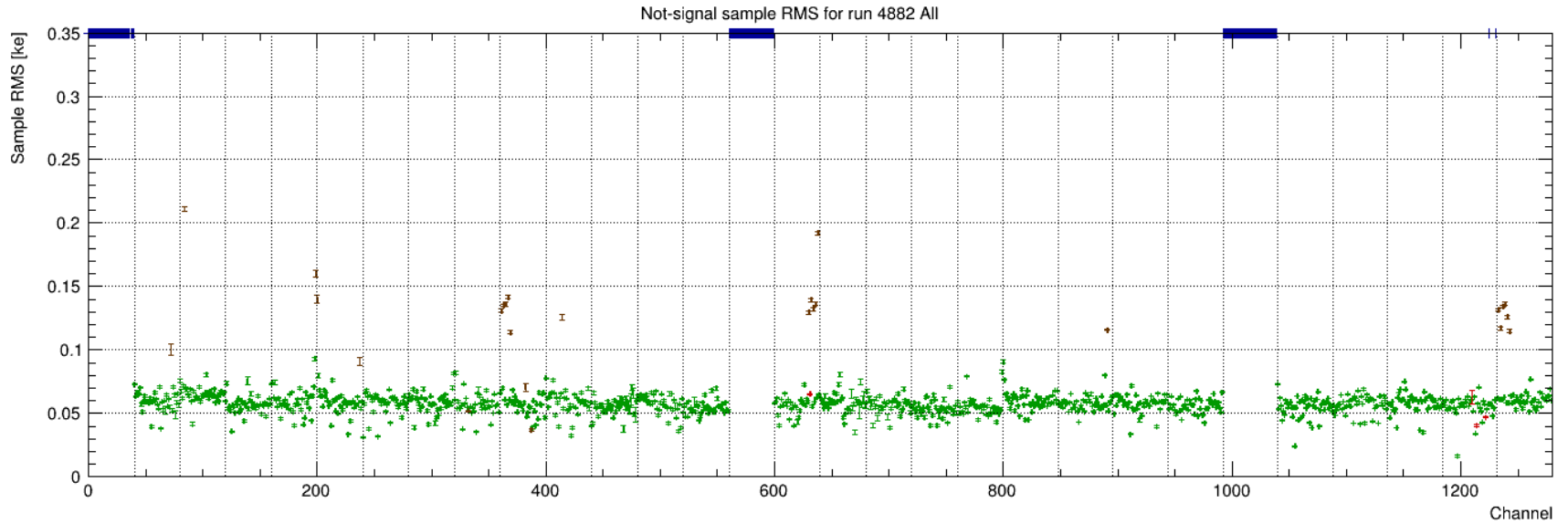
Not-signal 50-sample RMS for run 4880 All



# 4881

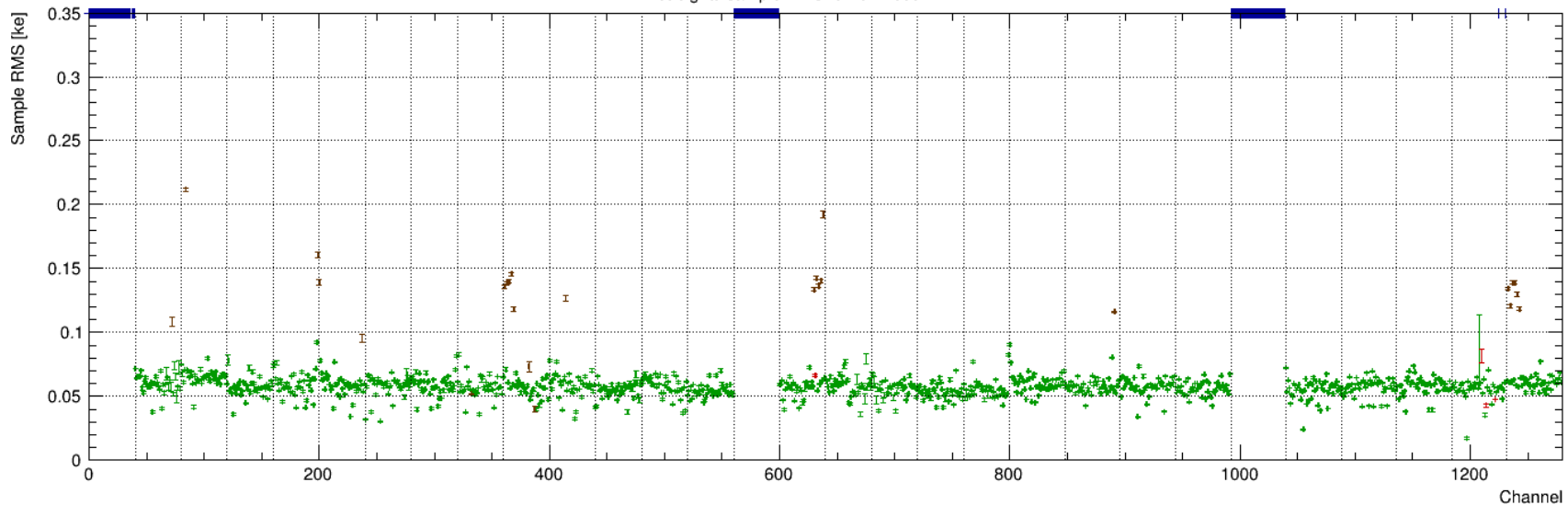


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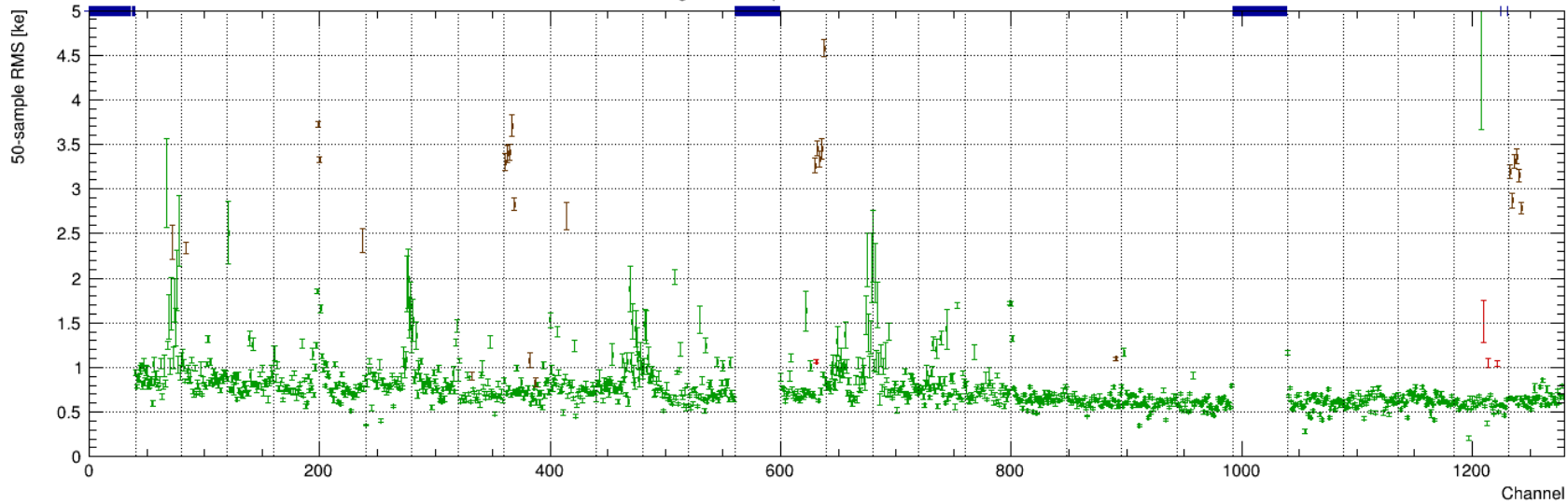


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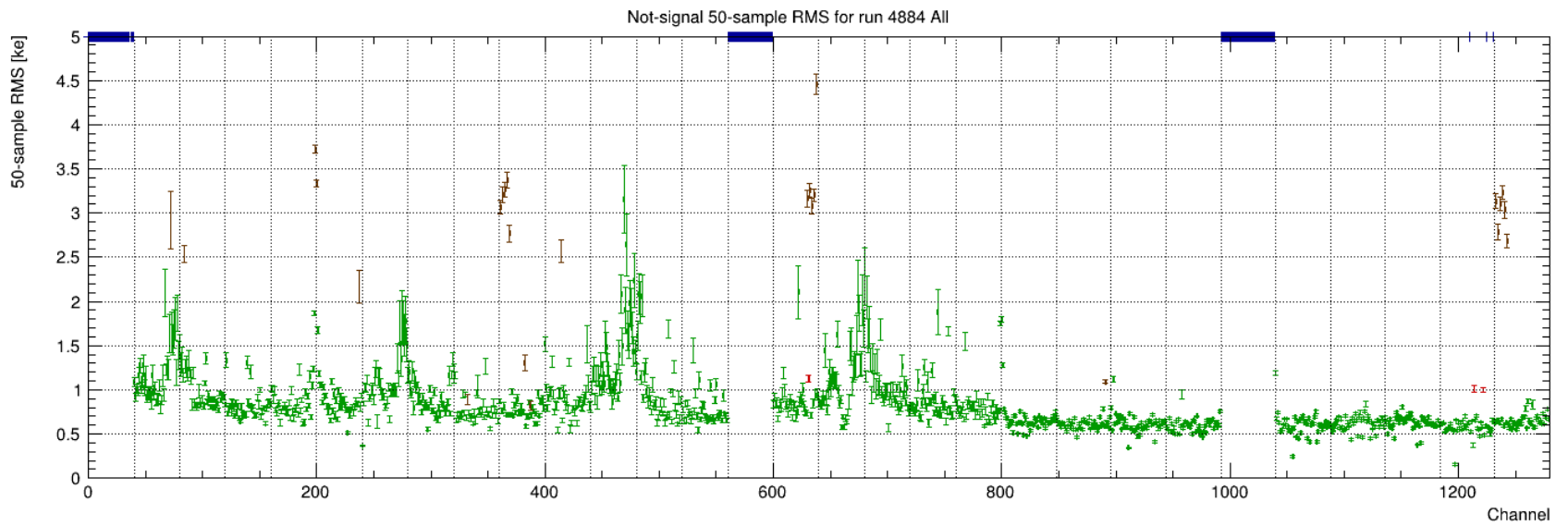
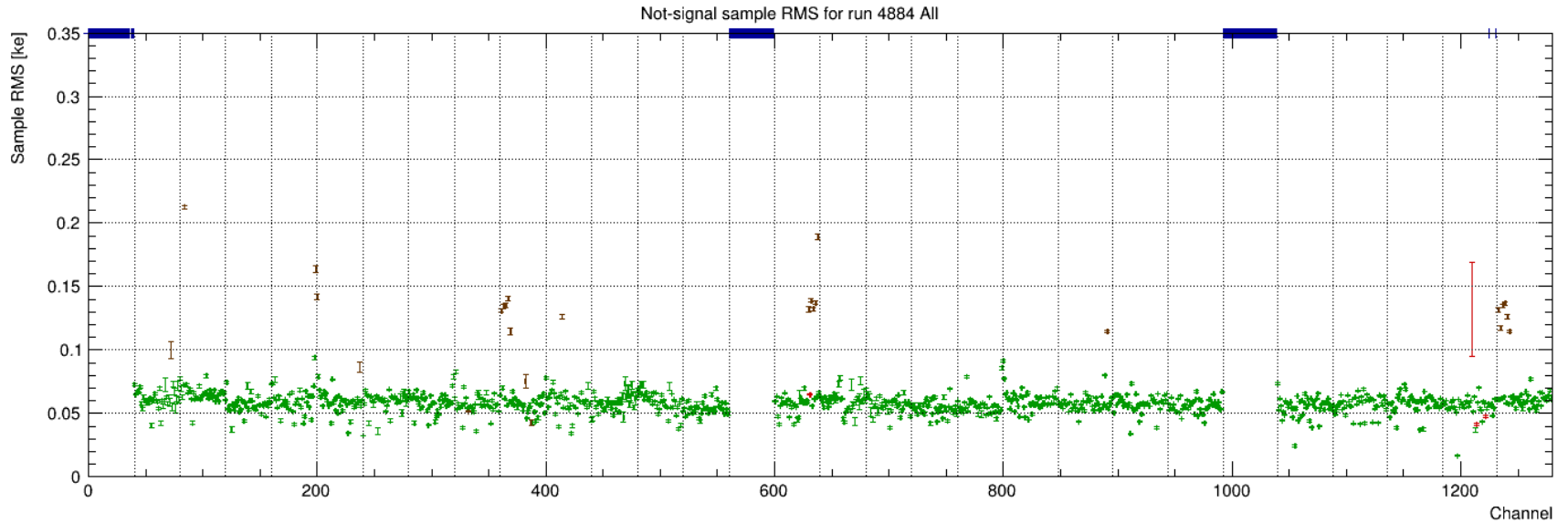
Not-signal sample RMS for run 4883 All



Not-signal 50-sample RMS for run 4883 All

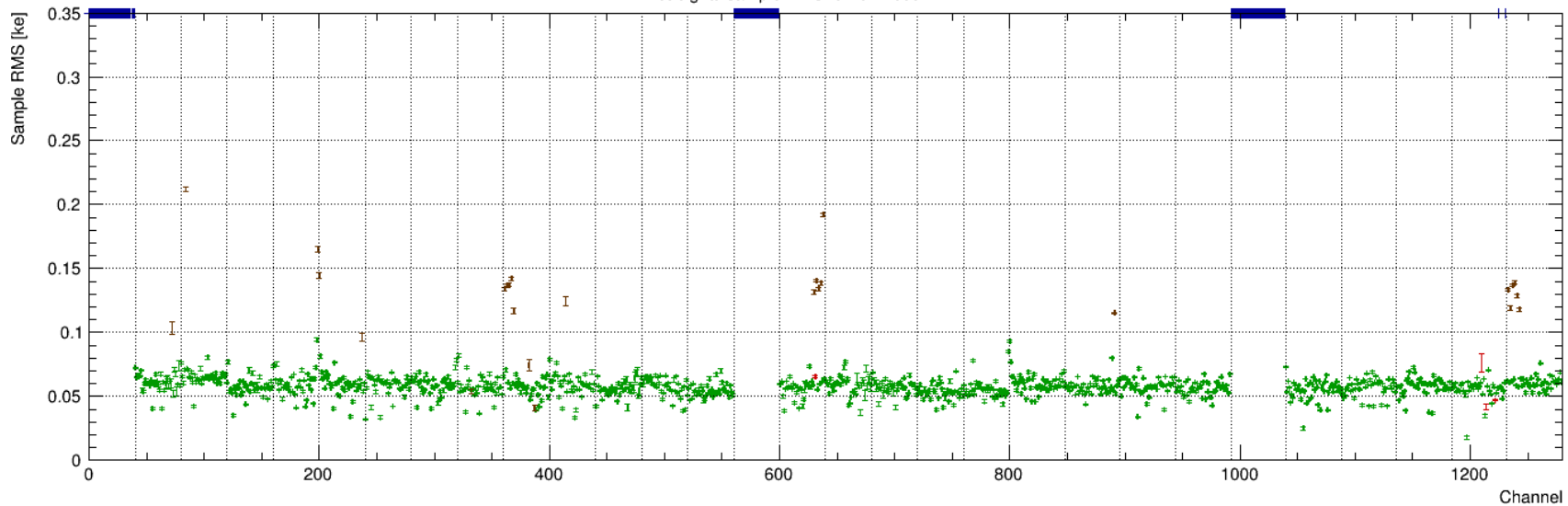


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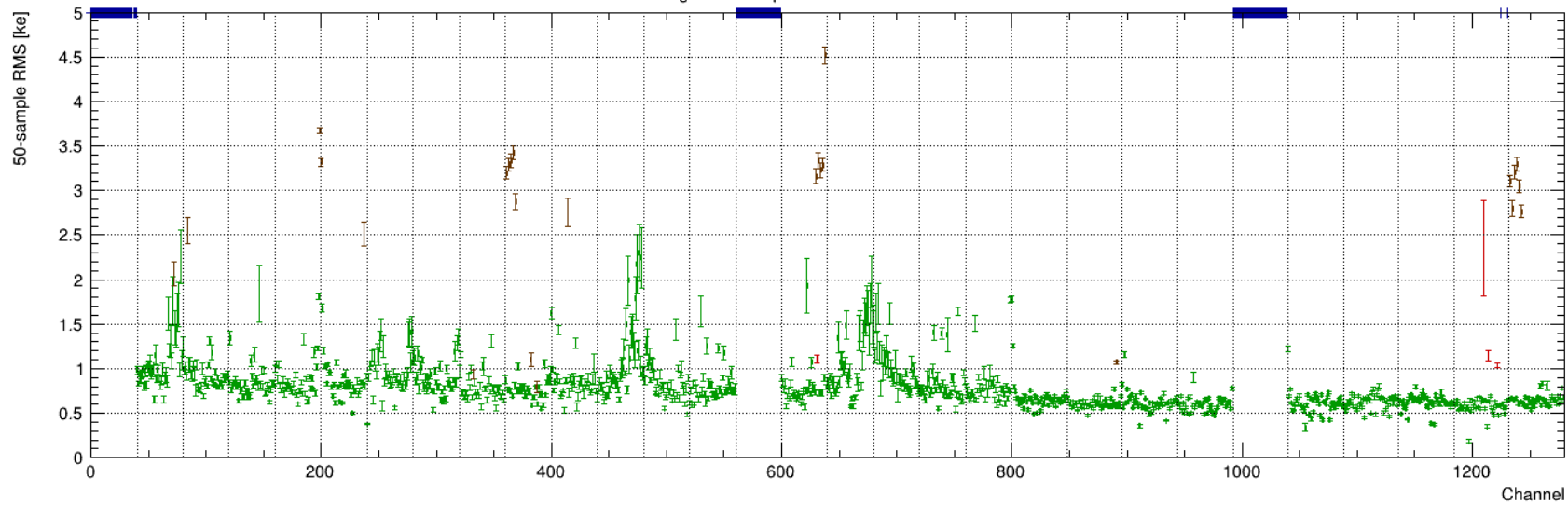


# 4885

Not-signal sample RMS for run 4885 All

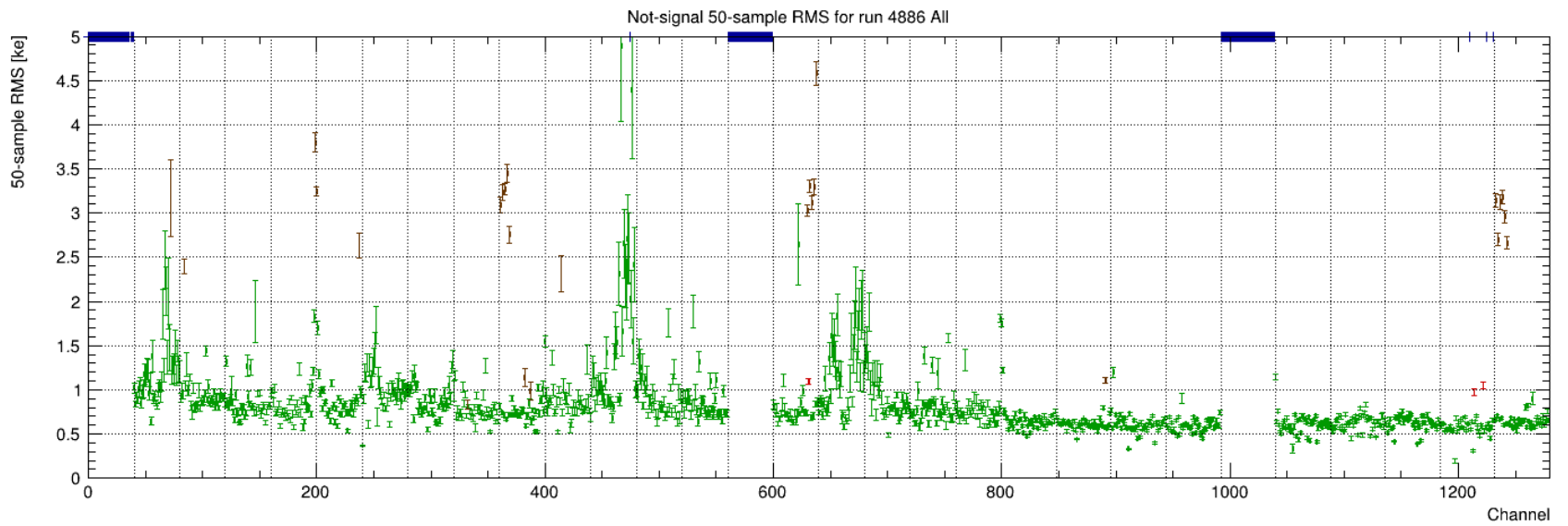
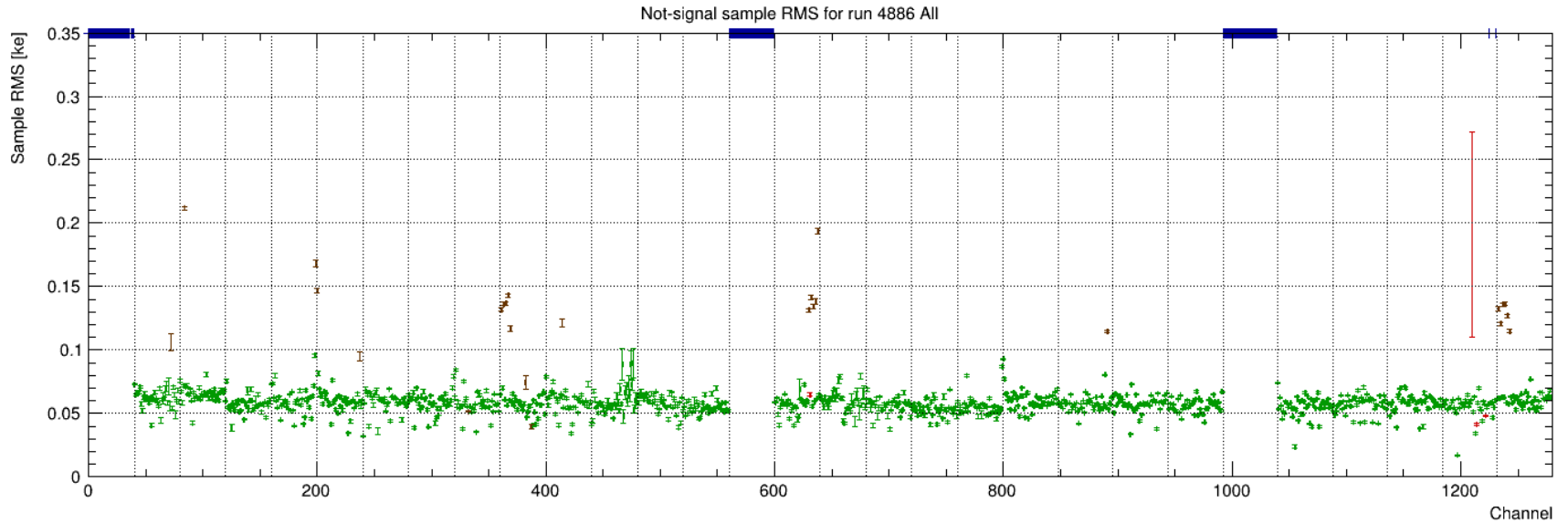


Not-signal 50-sample RMS for run 4885 All



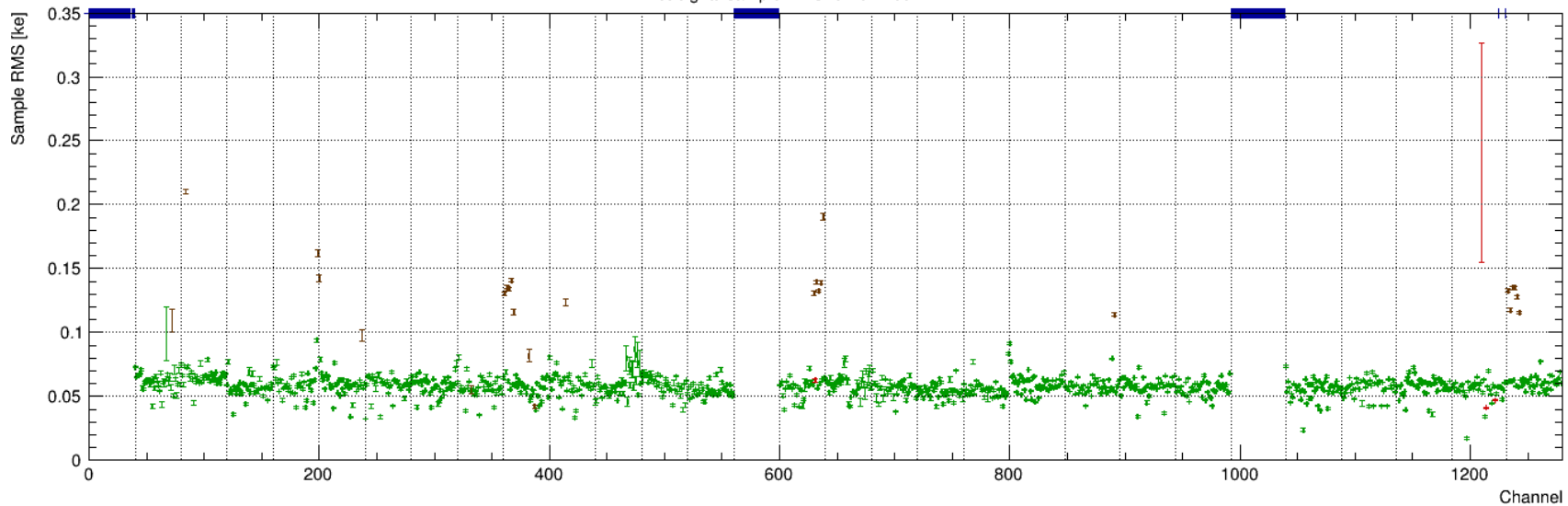


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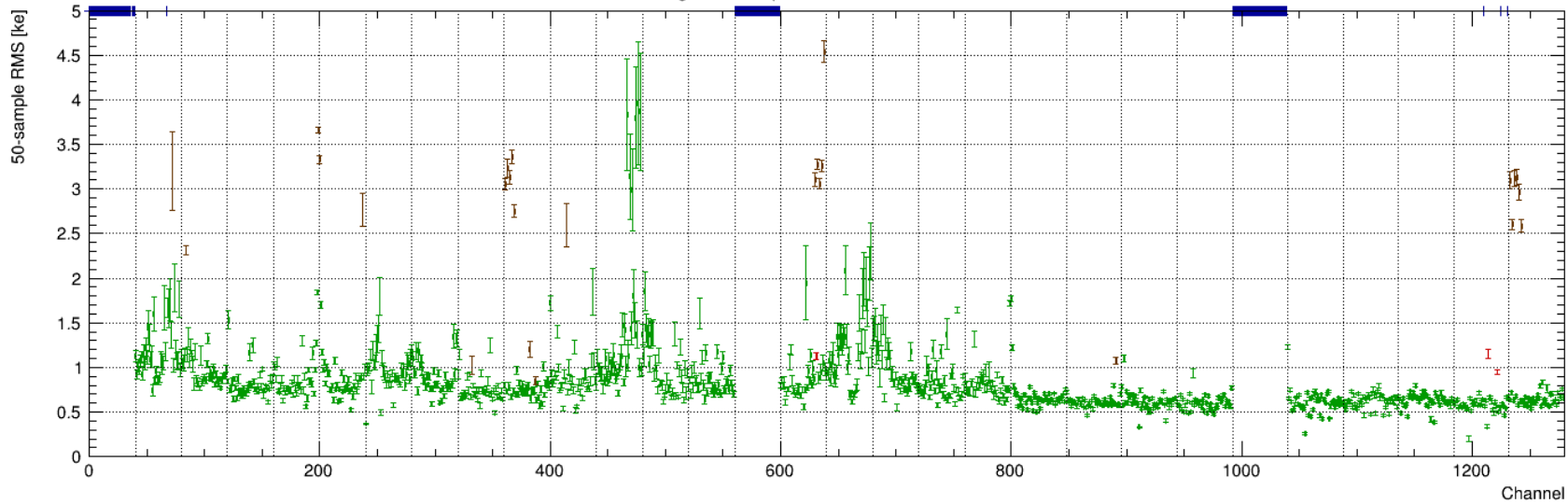


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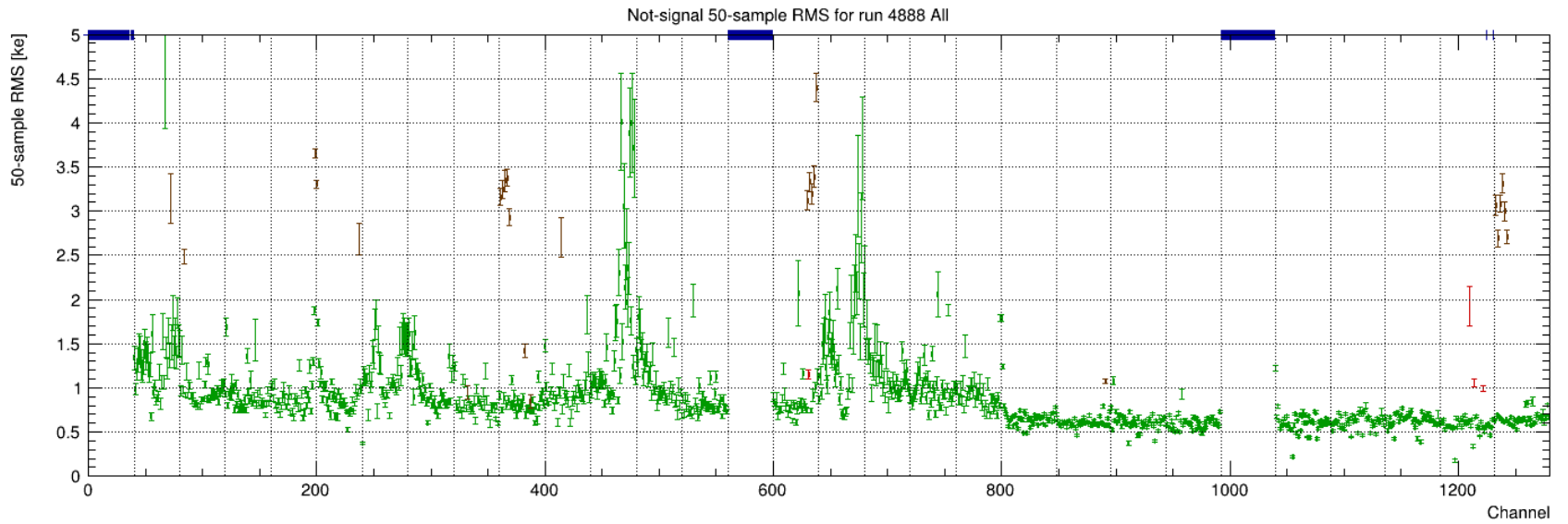
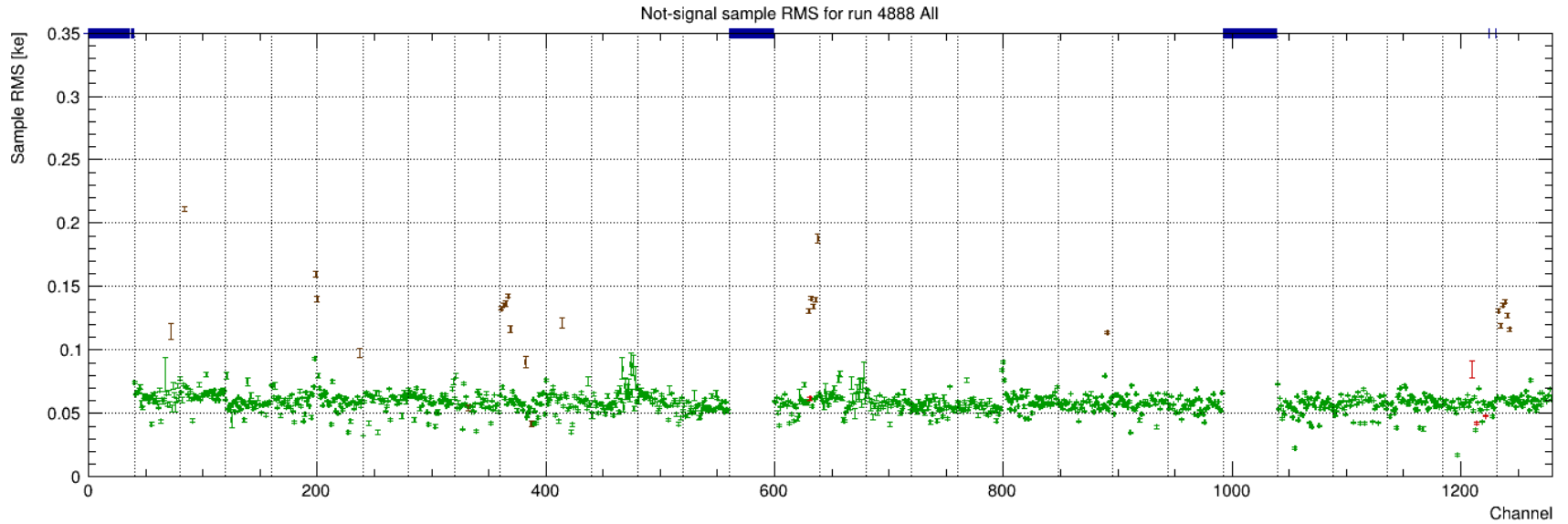
Not-signal sample RMS for run 4887 All



Not-signal 50-sample RMS for run 4887 All

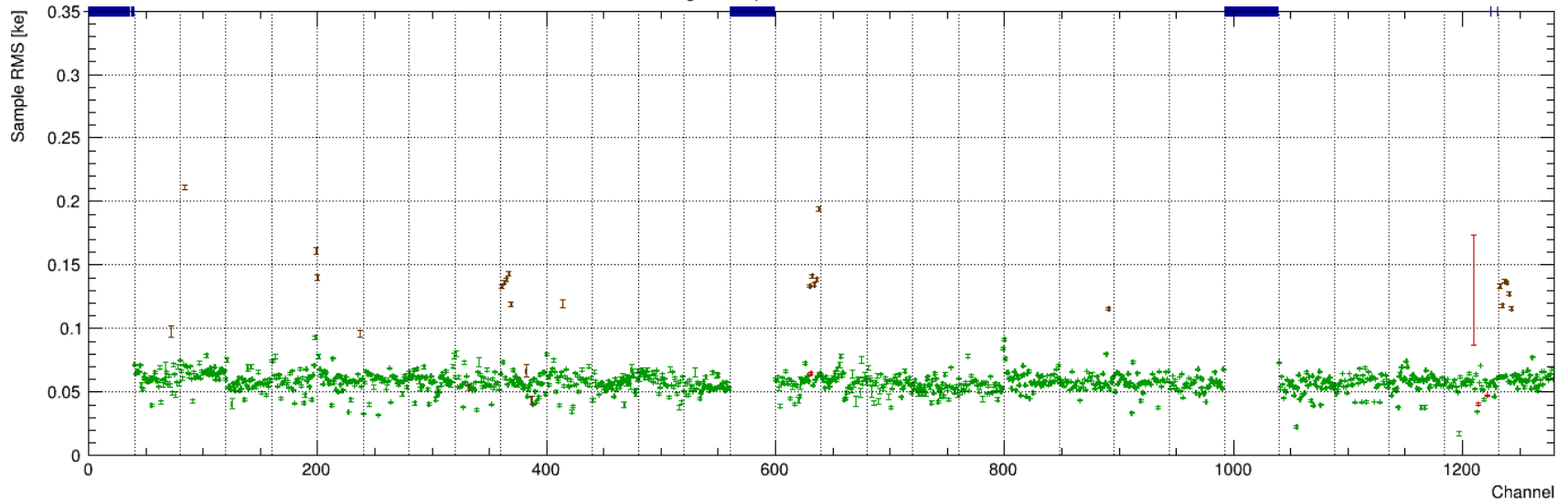


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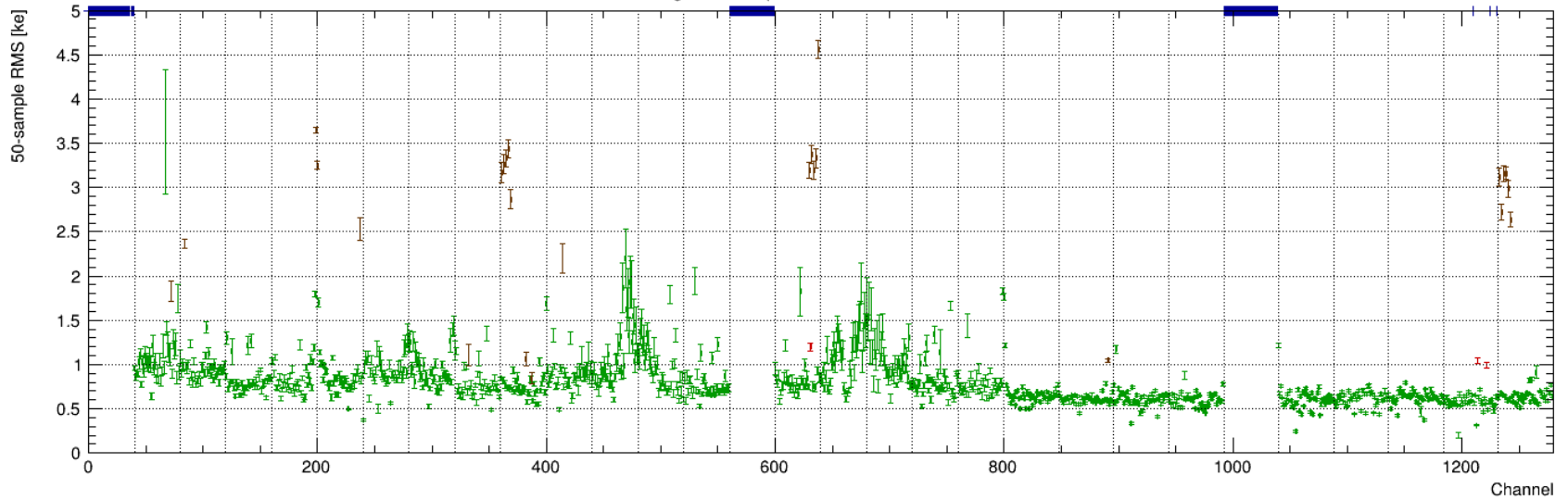


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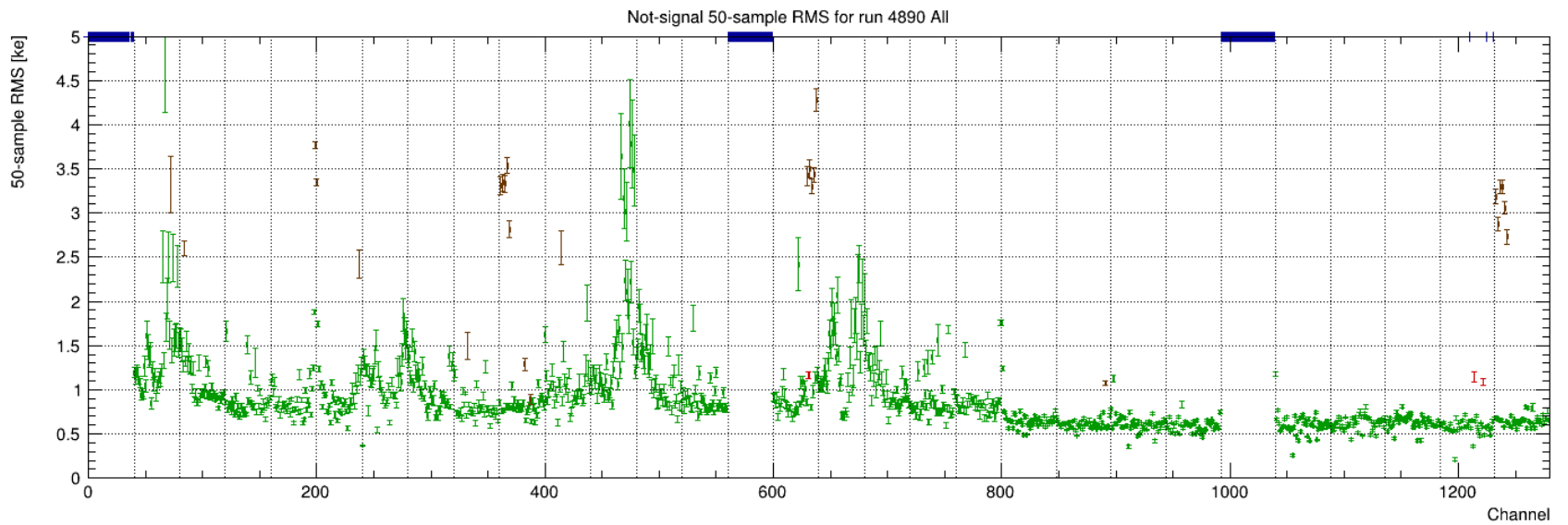
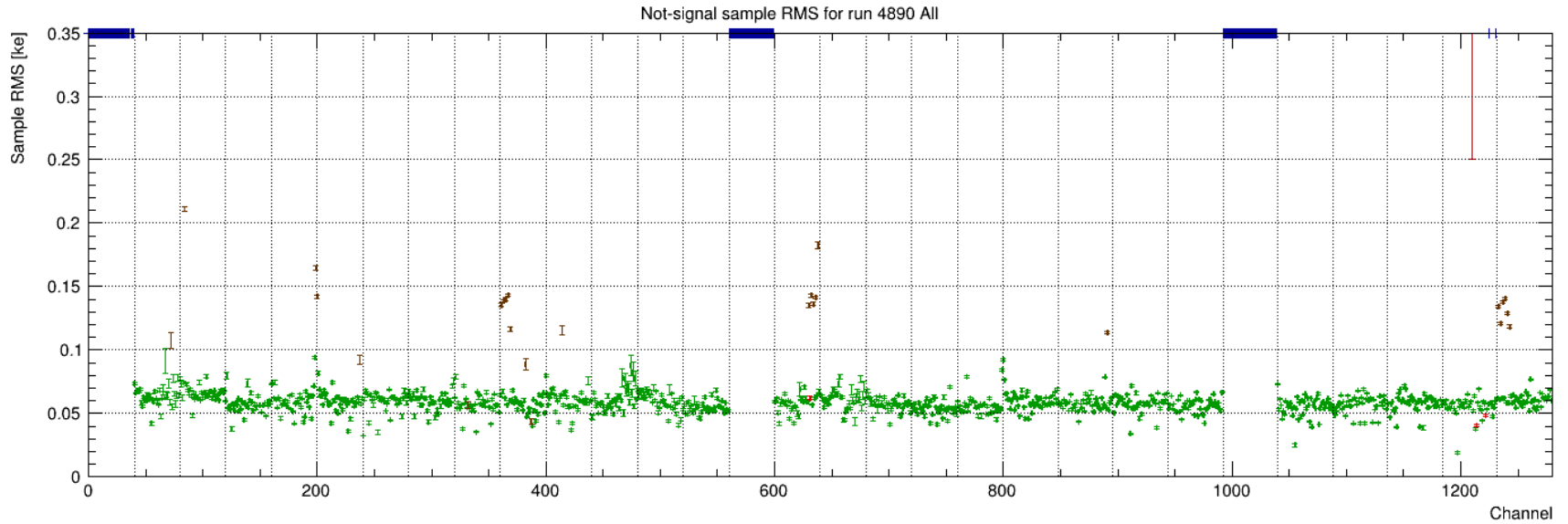
Not-signal sample RMS for run 4889 All



Not-signal 50-sample RMS for run 4889 All



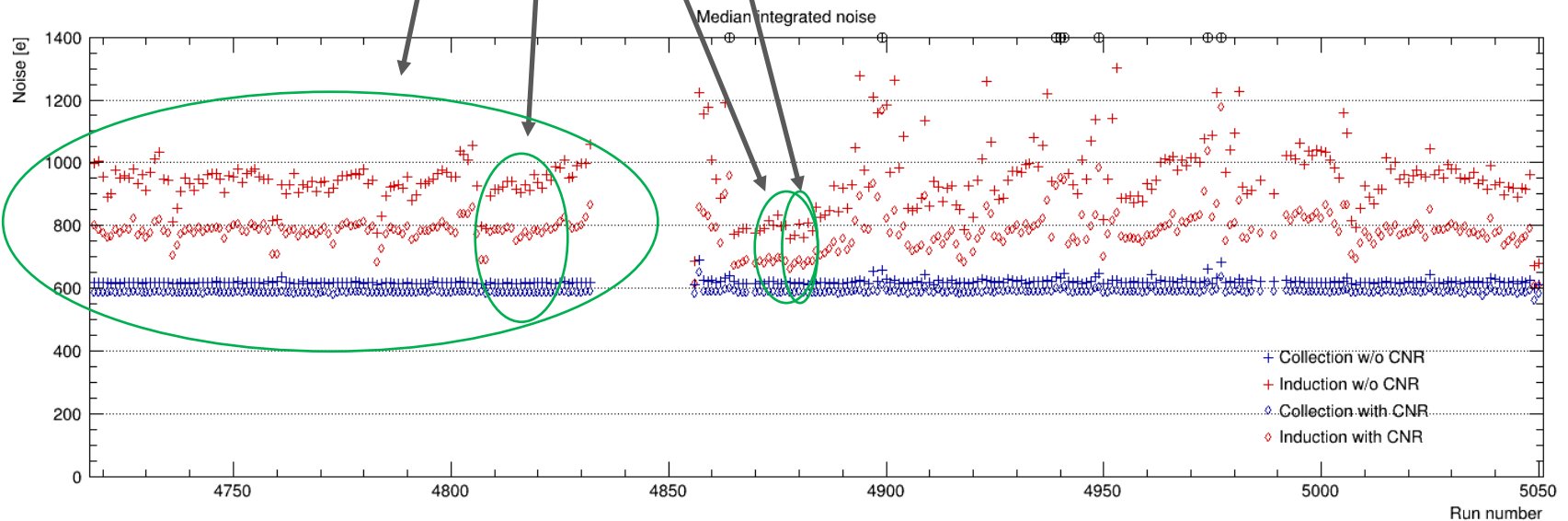
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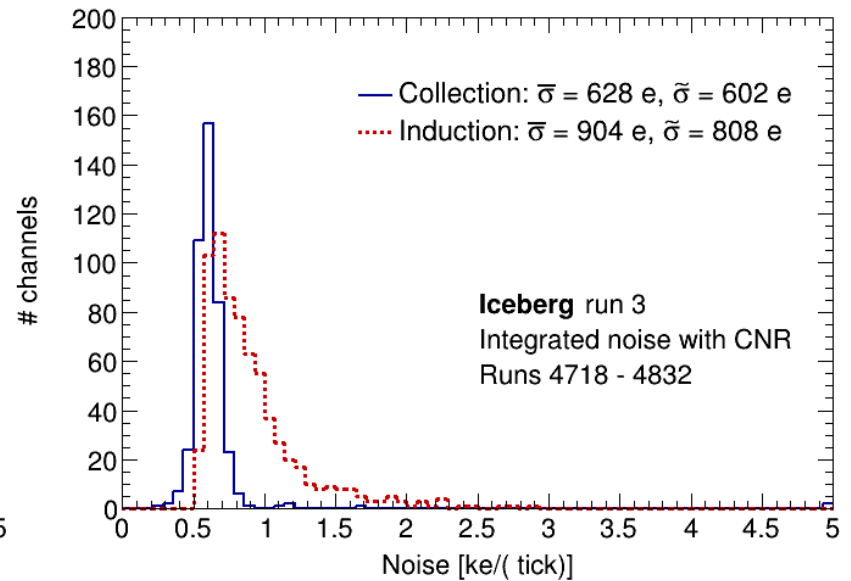
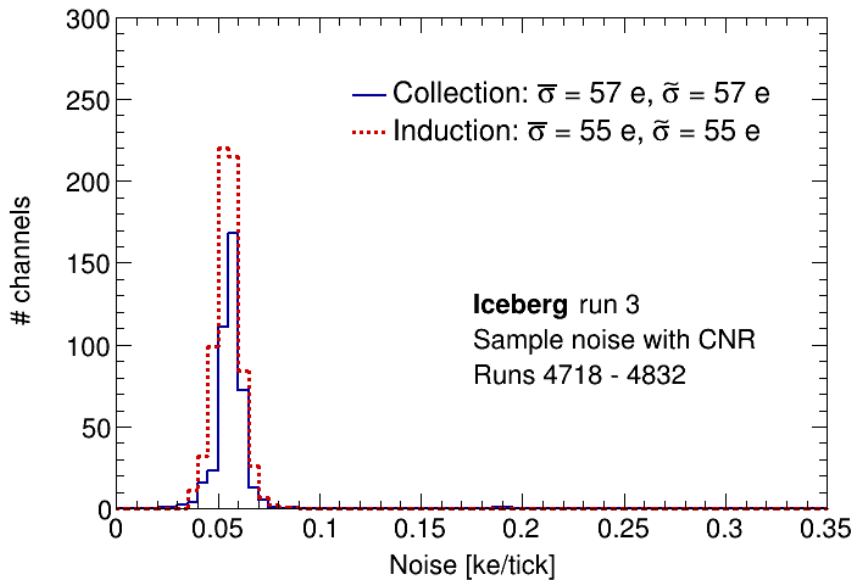
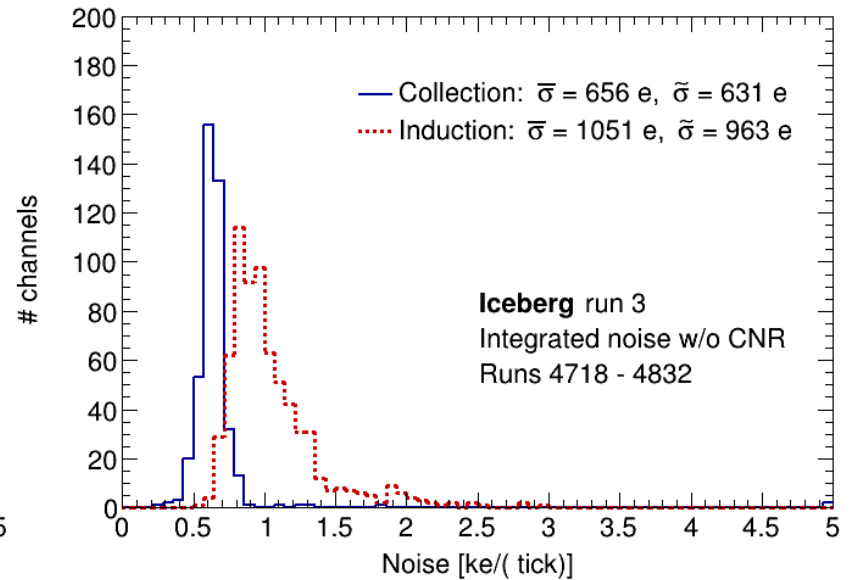
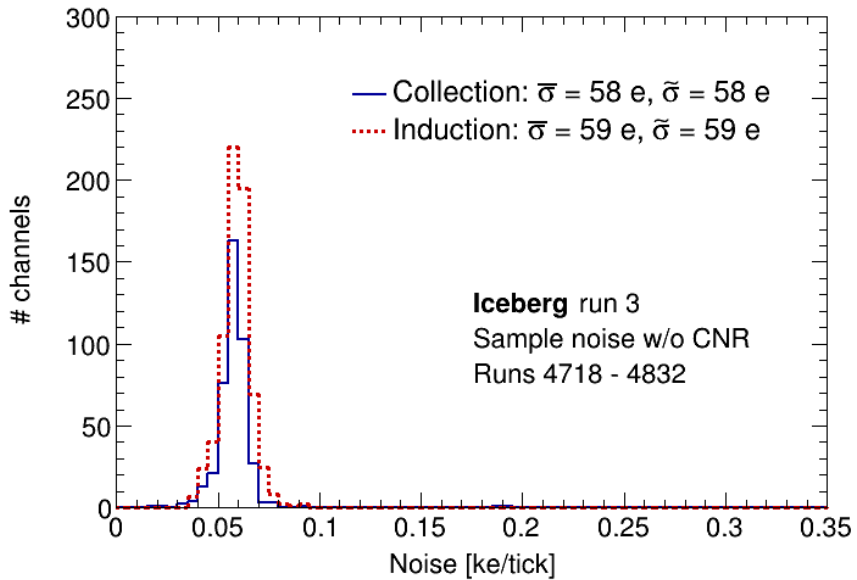
# Noise summary selections

Want noise summary values

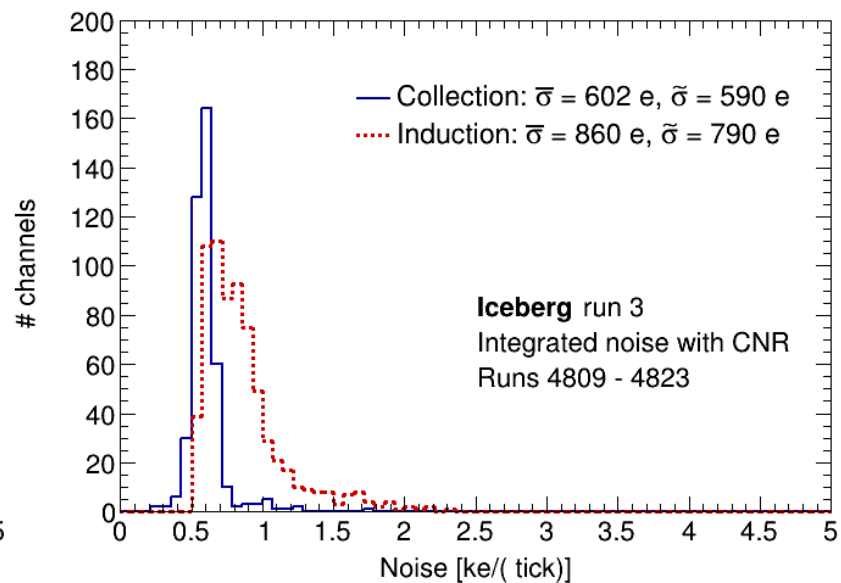
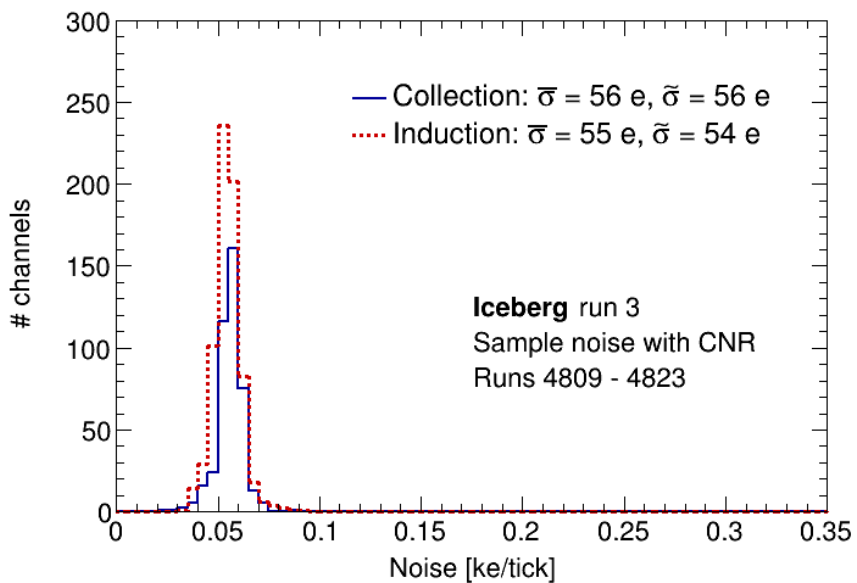
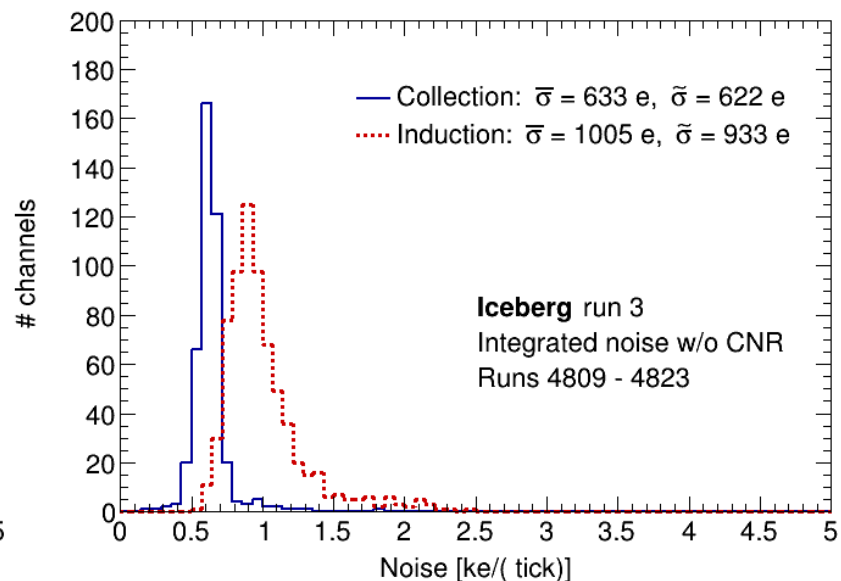
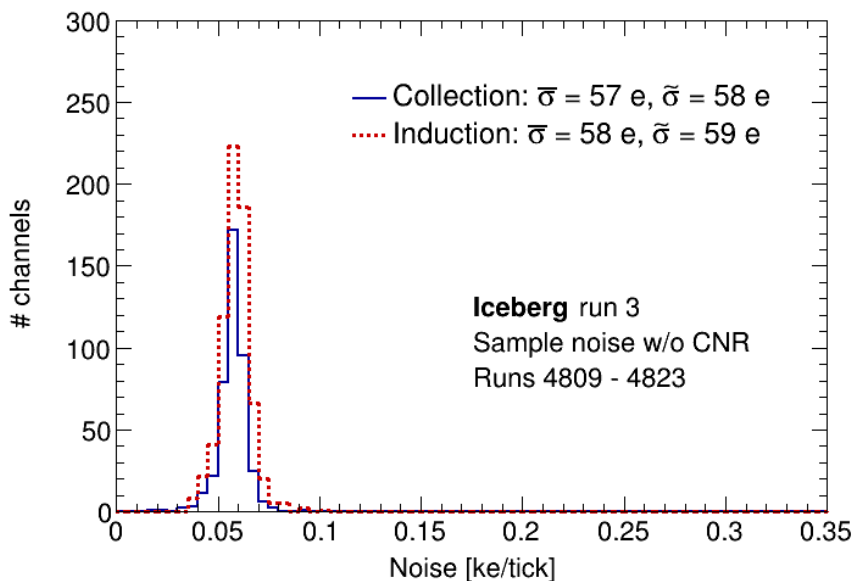
- For collection, induction with and w/o CNR
- Look at some run periods
  - First period (4718 -4832)
  - Typical: runs 4809 – 4823 (much of the data looks like this)
  - Quiet: runs 4865 – 4885
  - Very quiet: runs 4877 - 4882



# First period (4718 – 4832)

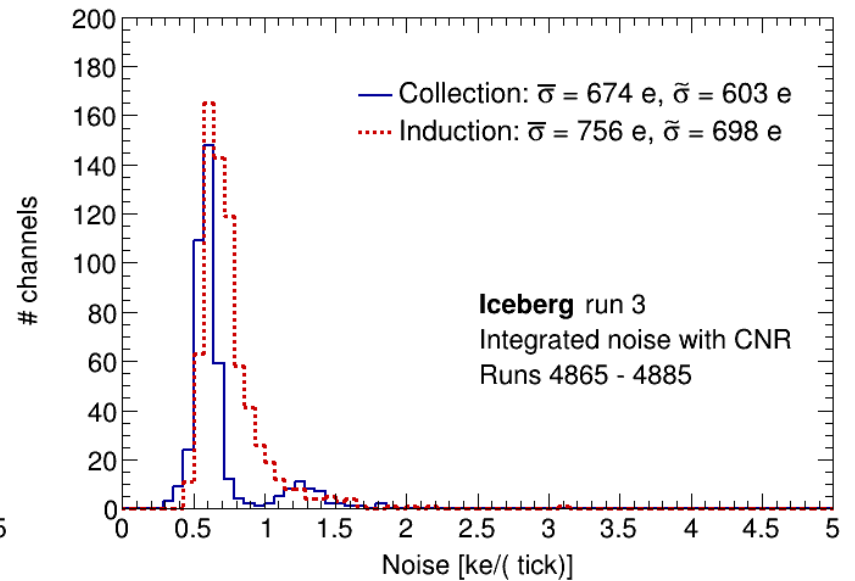
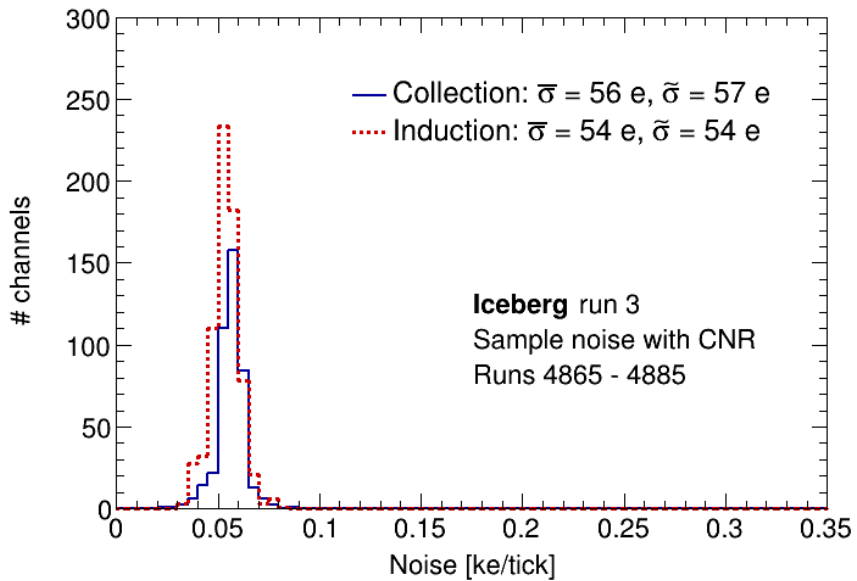
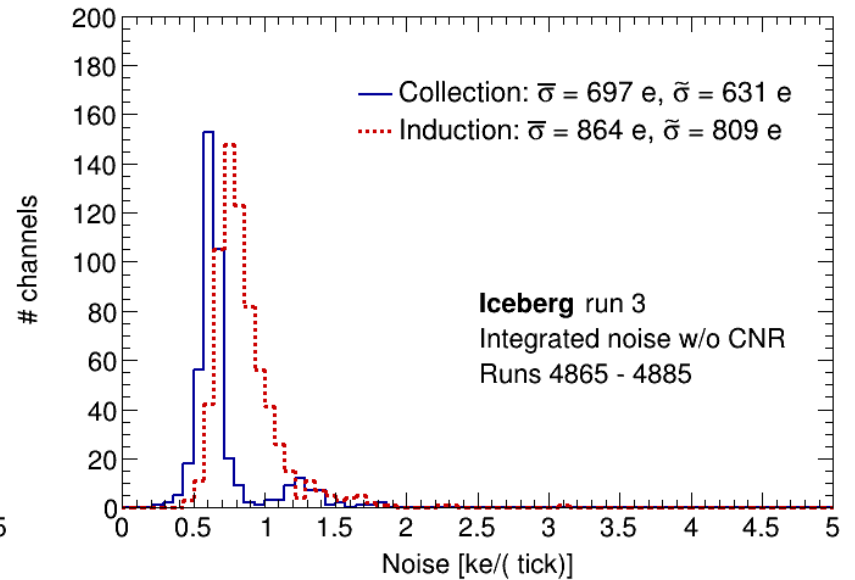
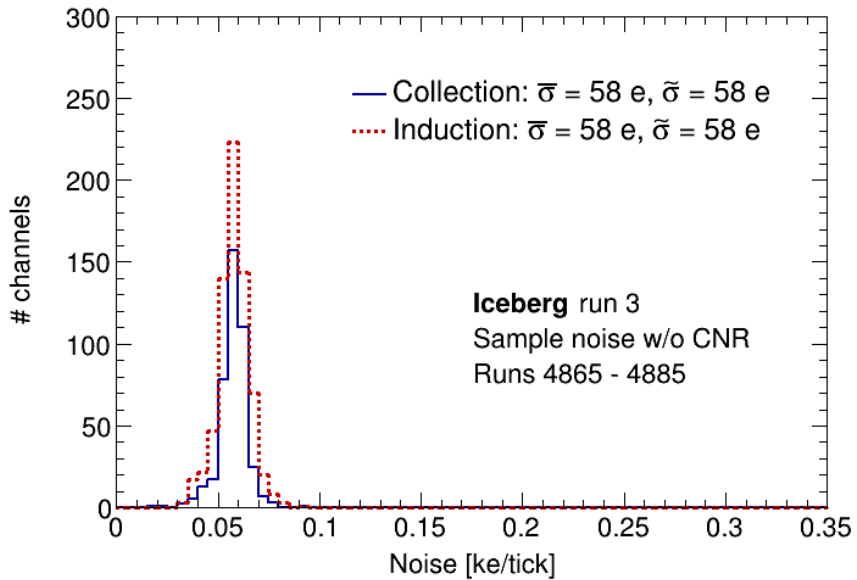


# Typical (4809 – 4823)

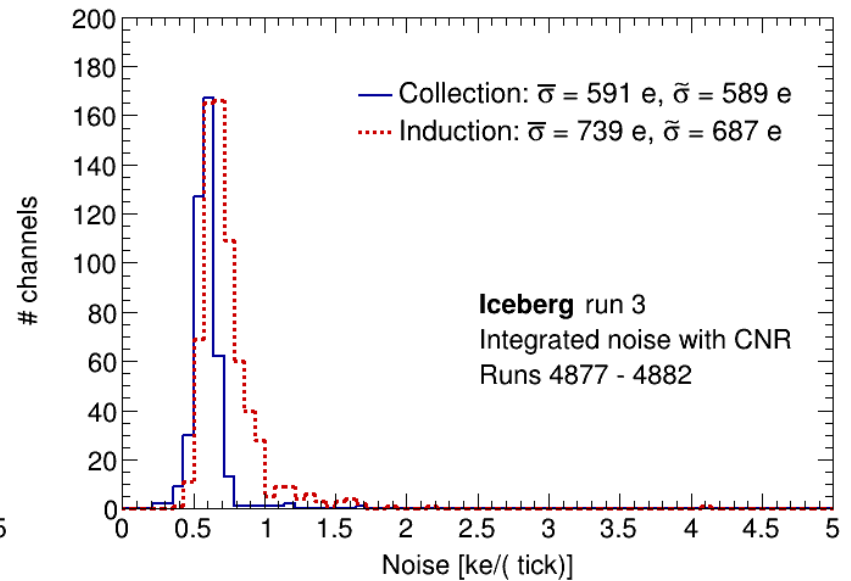
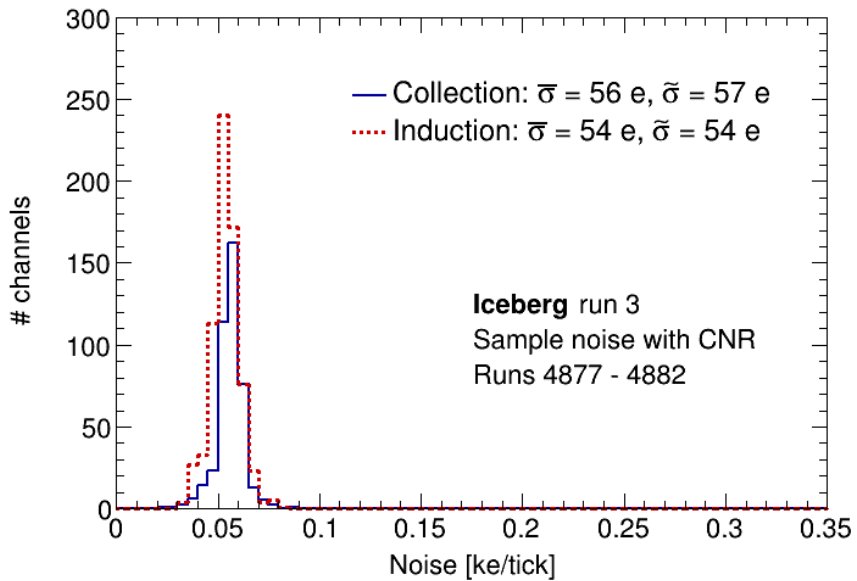
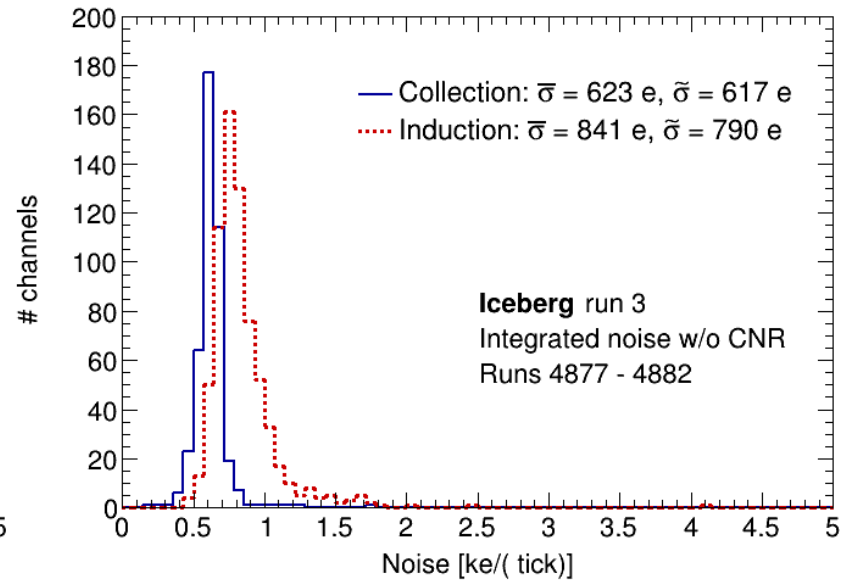
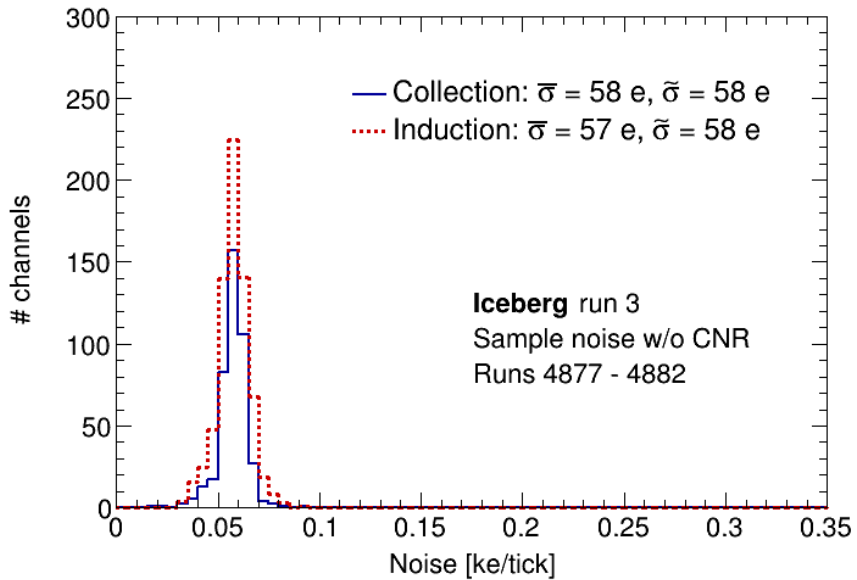




# Quiet (4865 – 4885)



# Very quiet (4877 – 4882)



# Noise summary medians

		<b>Sample noise [e]</b>			
	<b>CNR</b>	<b>All</b>	<b>Typical</b>	<b>Quiet</b>	<b>V. quiet</b>
<b>Collection</b>	no	58	58	58	58
	yes	57	56	57	57
<b>Induction</b>	no	59	59	58	58
	yes	55	54	54	54

		<b>Integrated noise [e]</b>			
	<b>CNR</b>	<b>All</b>	<b>Typical</b>	<b>Quiet</b>	<b>V. quiet</b>
<b>Collection</b>	no	630	620	630	620
	yes	600	590	600	590
<b>Induction</b>	no	930	930	810	790
	yes	790	790	700	690

# Noise comments

Preceding is direct measurement of collection noise

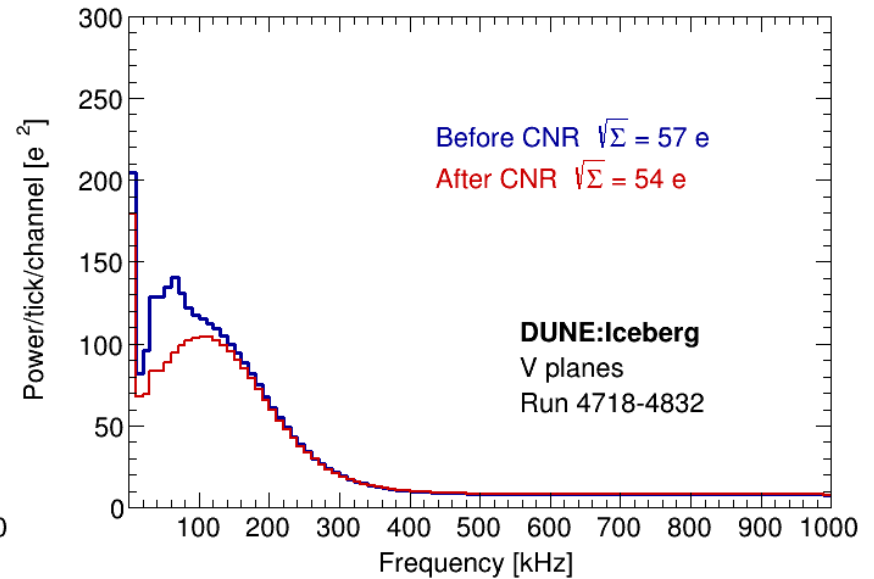
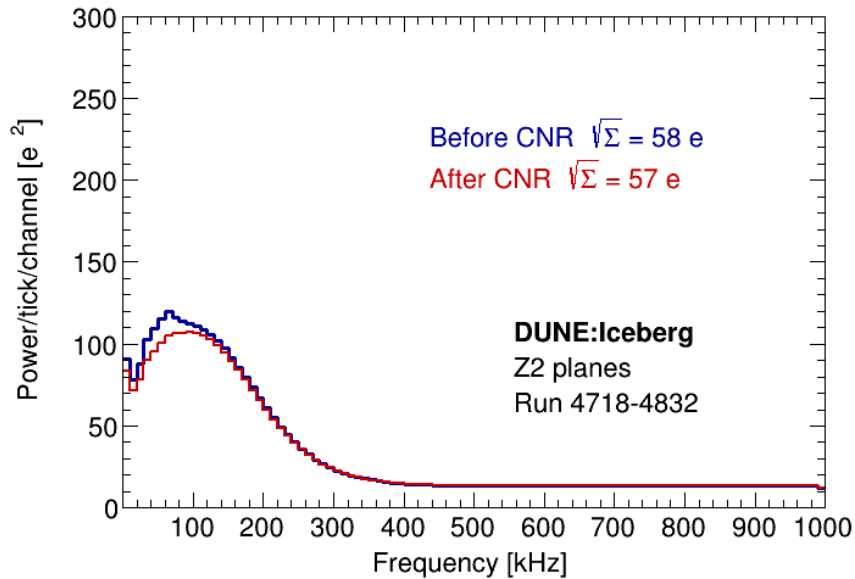
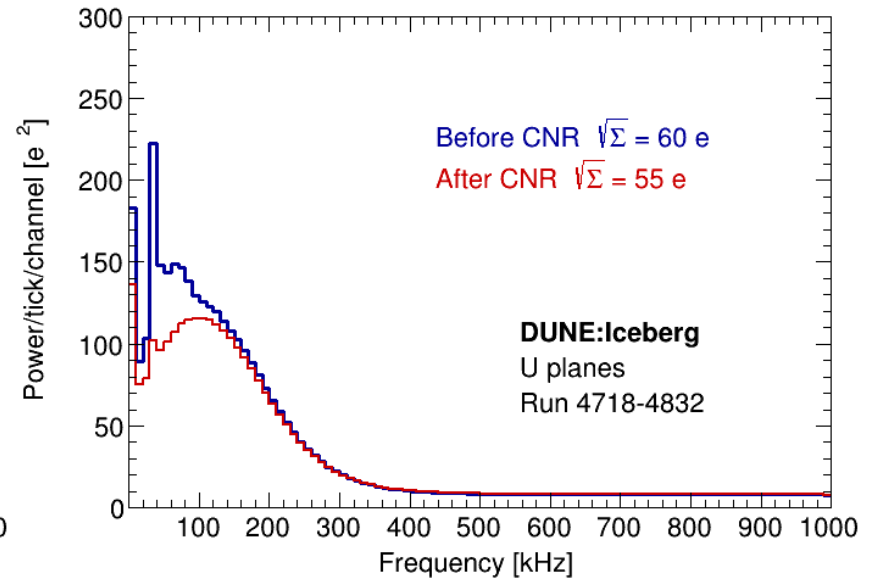
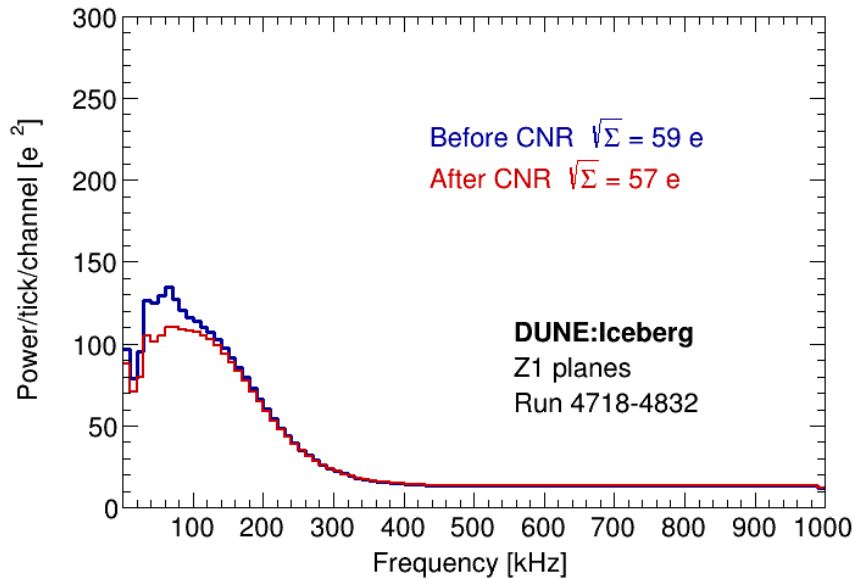
- Sample noise: 58  $\rightarrow$  57 e (before  $\rightarrow$  after CNR)
  - CNR = coherent noise removal
  - ENC =  $5.58 \times$  (sample noise): 320  $\rightarrow$  320 e
    - Dune doc 15523
- Integrated noise: 620  $\rightarrow$  590 e
  - Expect this uncertainty for charge measurements
    - Deconvolution should not change it
  - protoDUNE values are 1250, 920 e

Corresponding results for induction planes

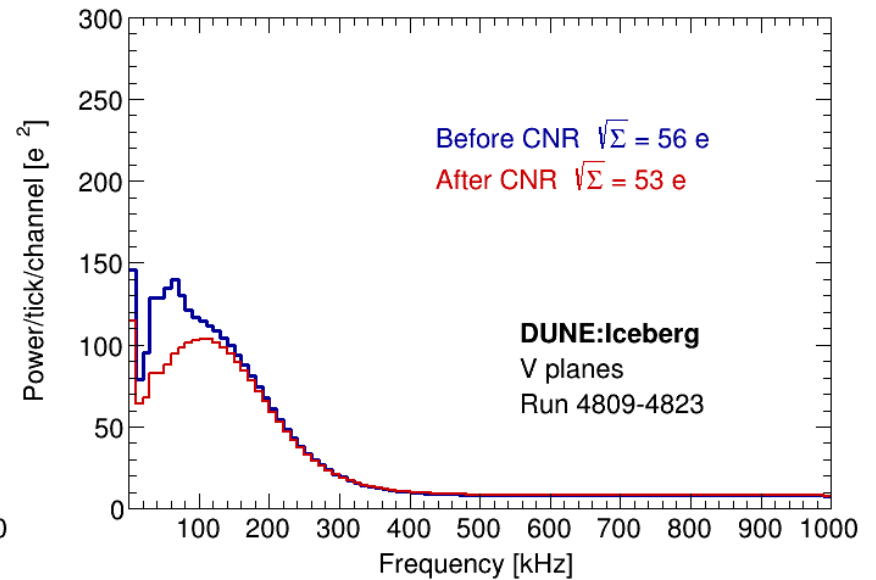
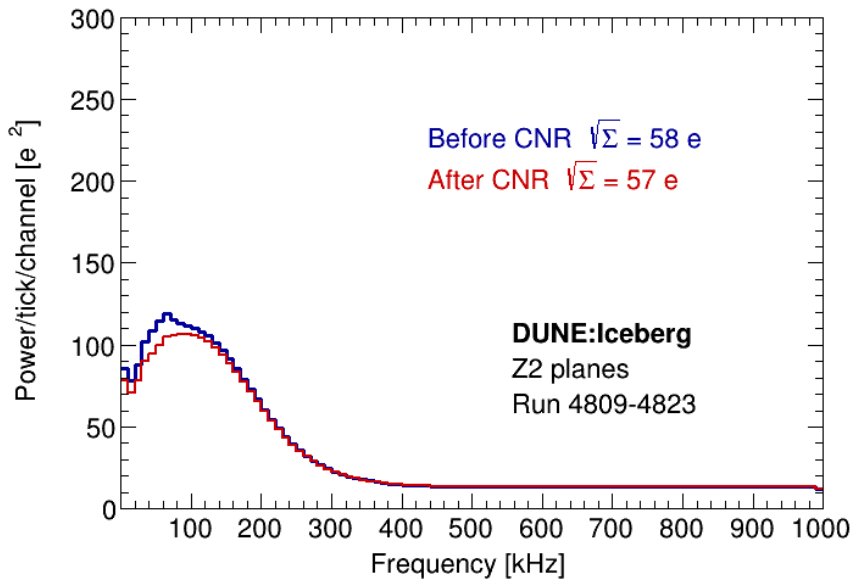
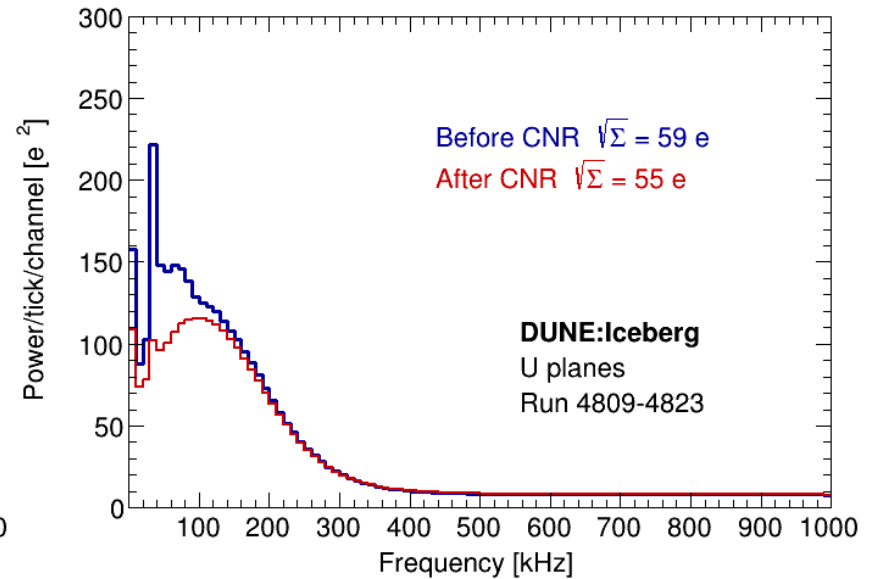
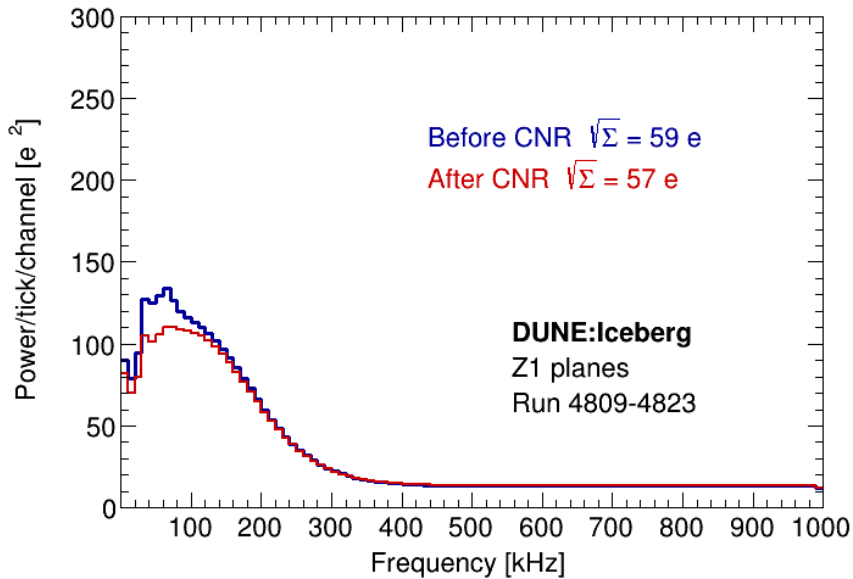
- Sample: 58  $\rightarrow$  54 e
  - ENC: 320  $\rightarrow$  300 e
- Integrated: 790  $\rightarrow$  690 e
  - Deconvolution  $\rightarrow$  much larger uncertainty on charge measurements
    - To be determined

# DFT power

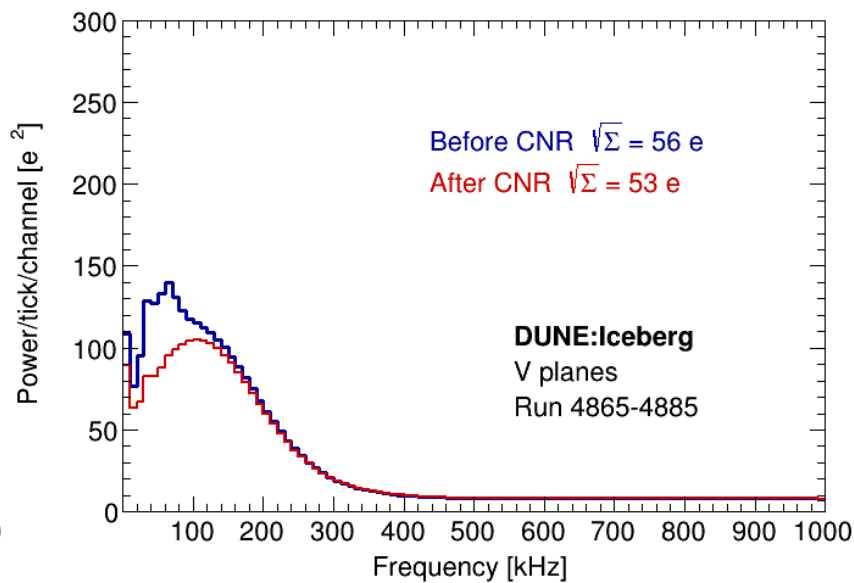
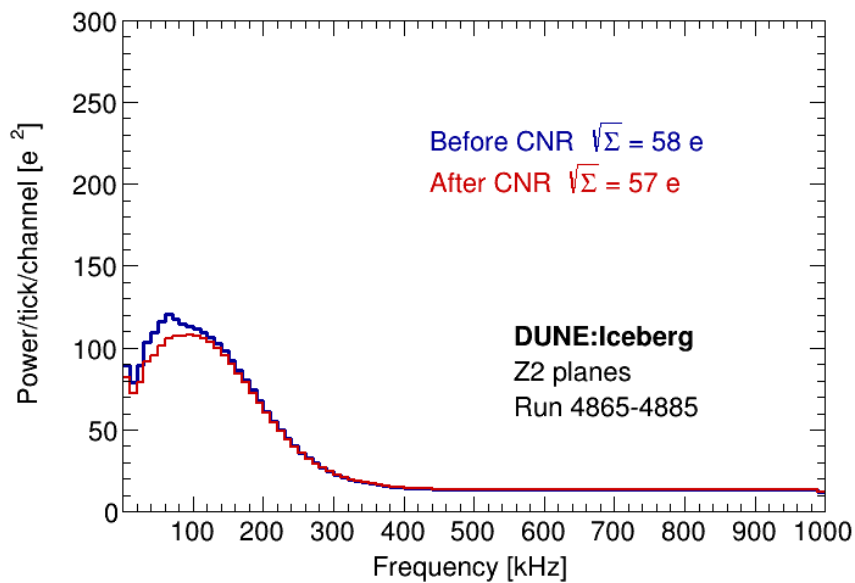
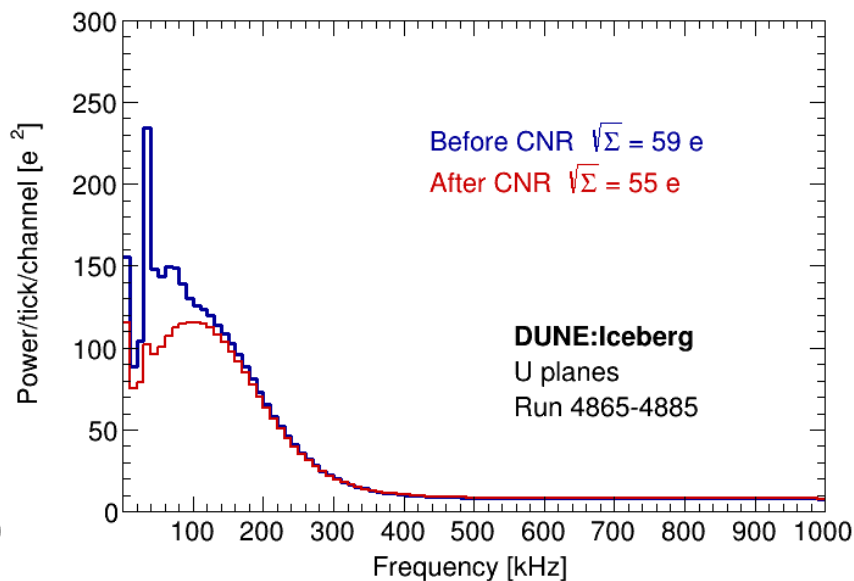
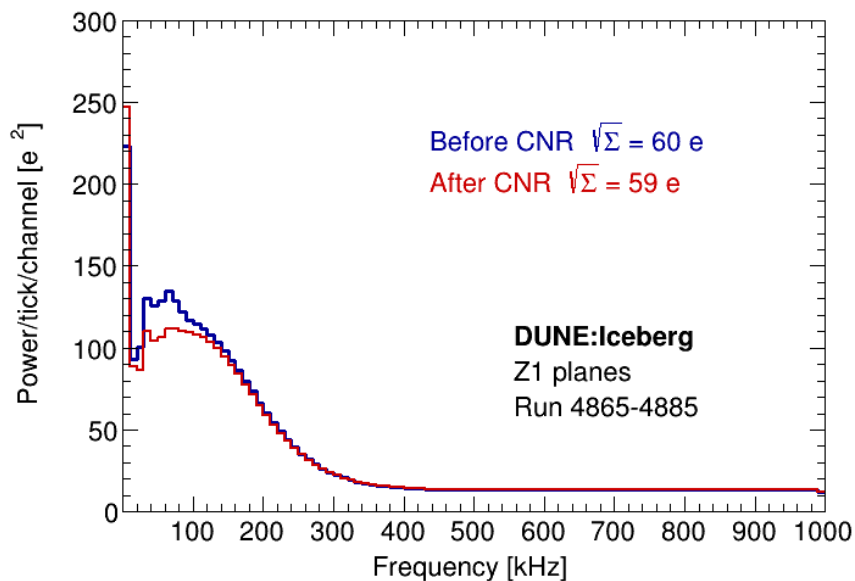
# First period (4718-4832)



# Typical (4809-4823)

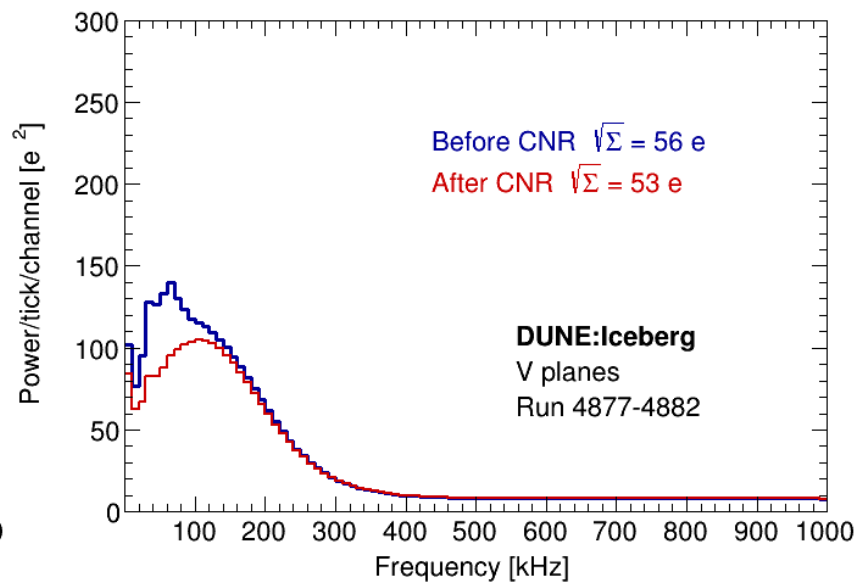
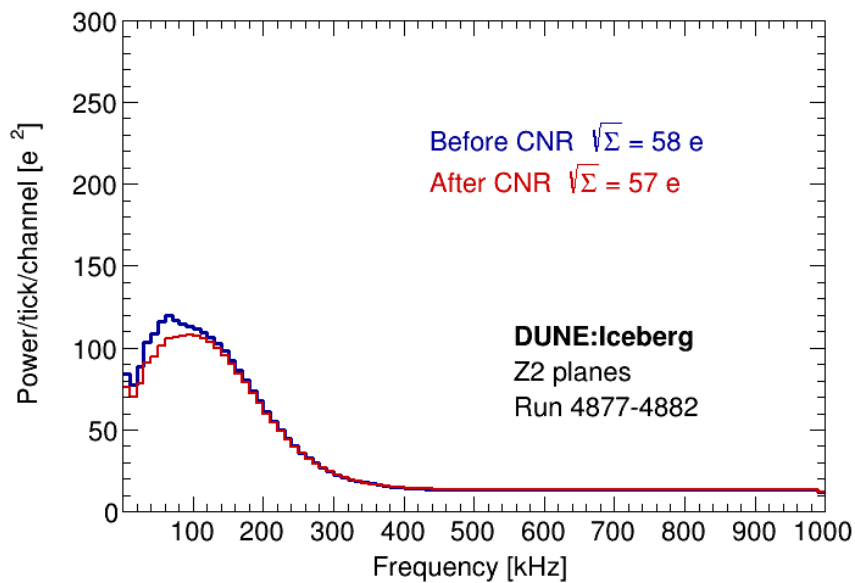
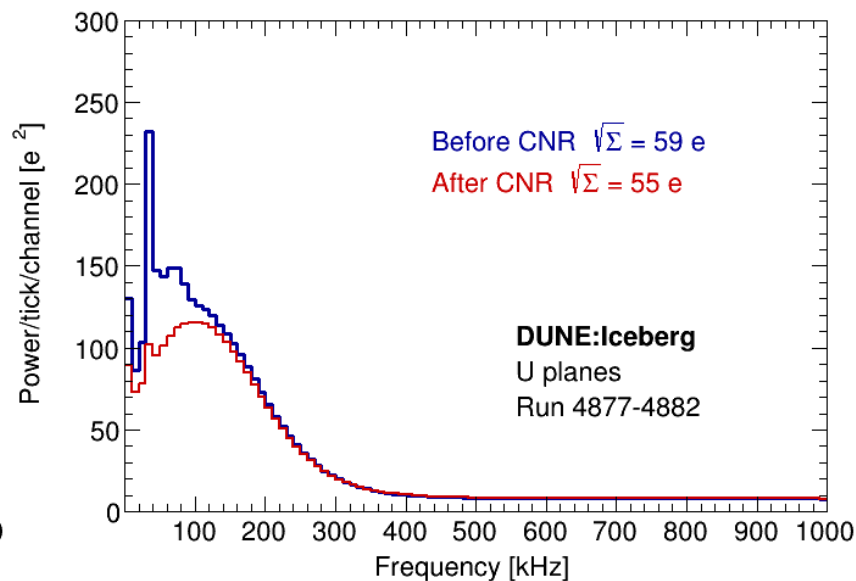
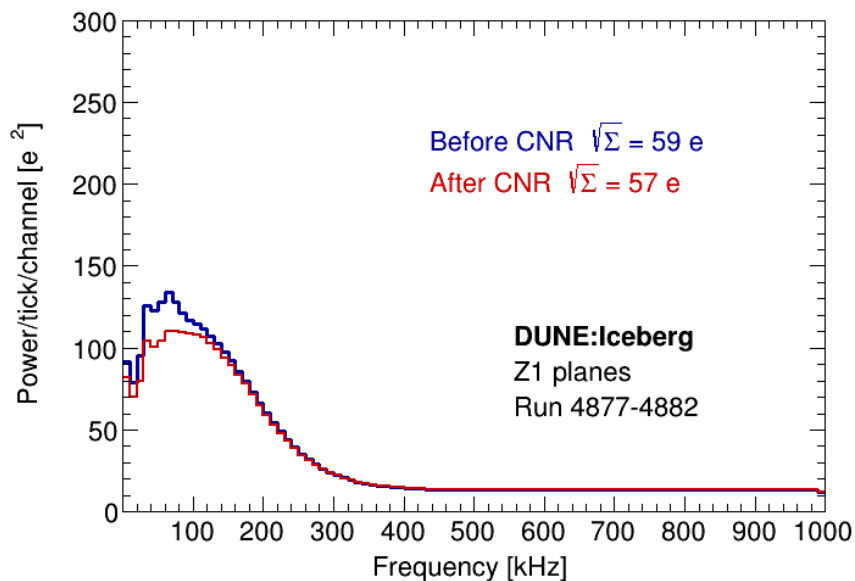


# Quiet (4865-4885)

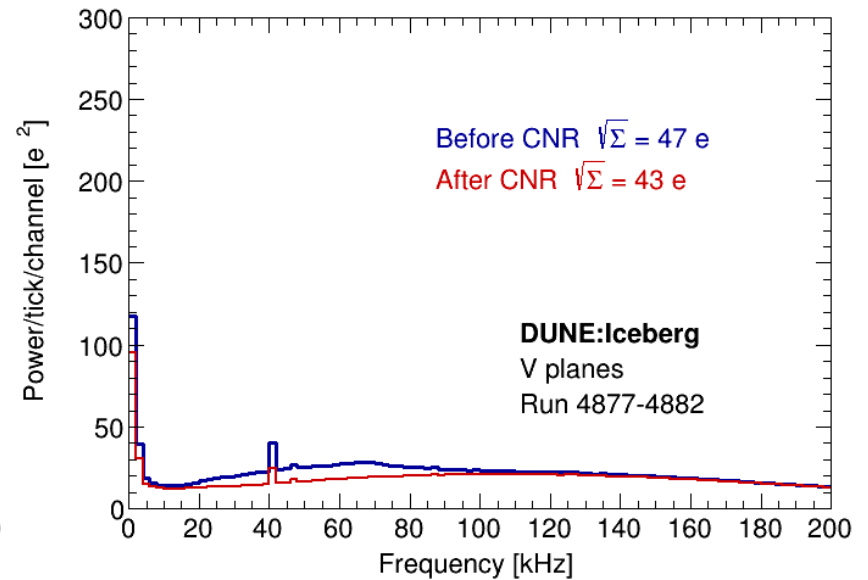
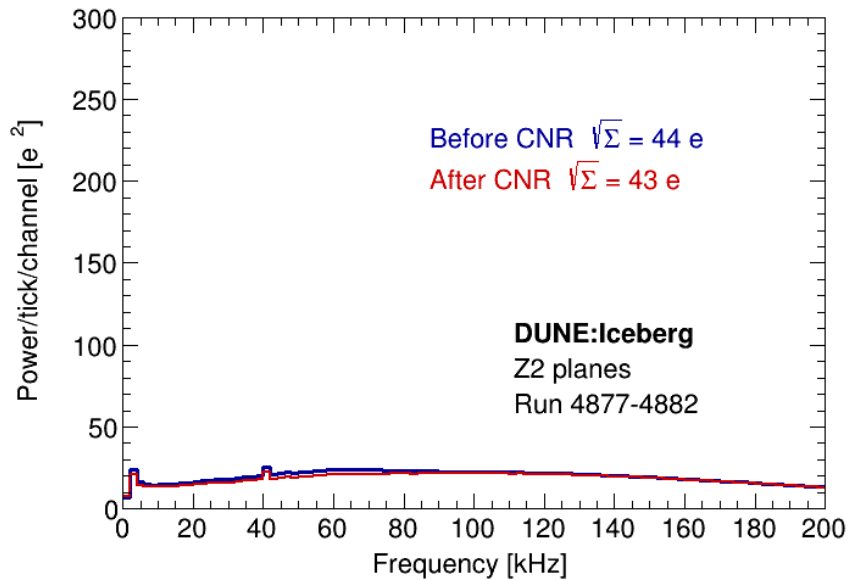
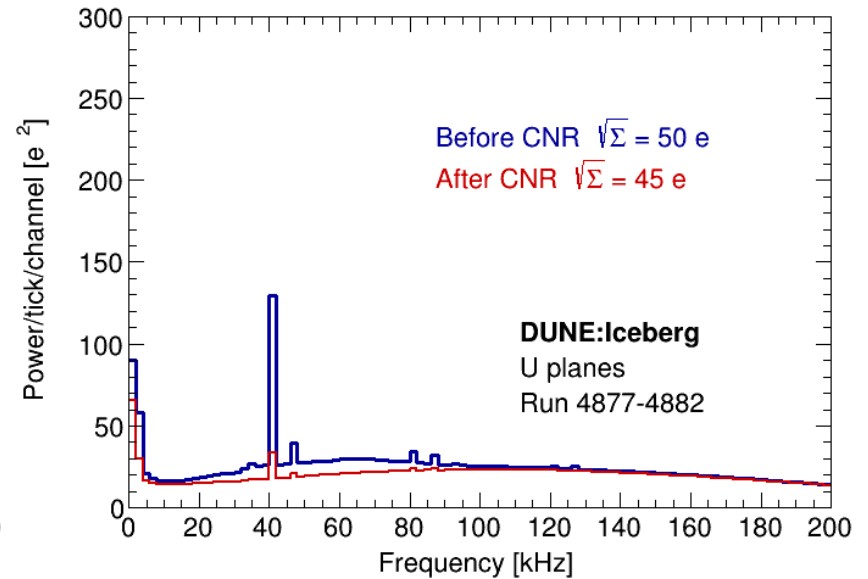
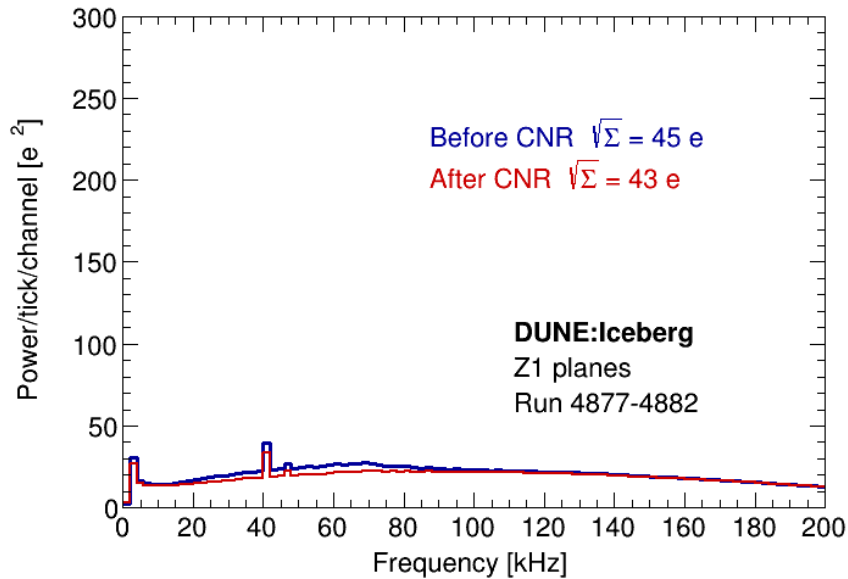




# Very quiet (4877-4882)



# Very quiet (4877-4882) zoomed



# Signal to noise (SNR)

# Signal to noise

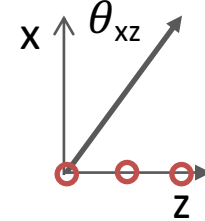
## Reported difficult to repeat protoDUNE SNR study

- Wanwei, Tingjun talk May 21
- Because tracks typically deposit charge on few wires
  - I.e.  $\theta_{xz} \approx 0$  deg
- But we expect signal height similar to that for  $\theta_{xz} \approx 90$ 
  - See following slides
  - Need a different analysis to pull this out
- And there are quite a few tracks (1 in 10 events?) where track is not quite perpendicular
  - Charge is distributed over 20+ wires
  - Amenable to traditional analysis?
  - Example plot follows

# Perpendicular tracks

# Tracks perpendicular to wire plane

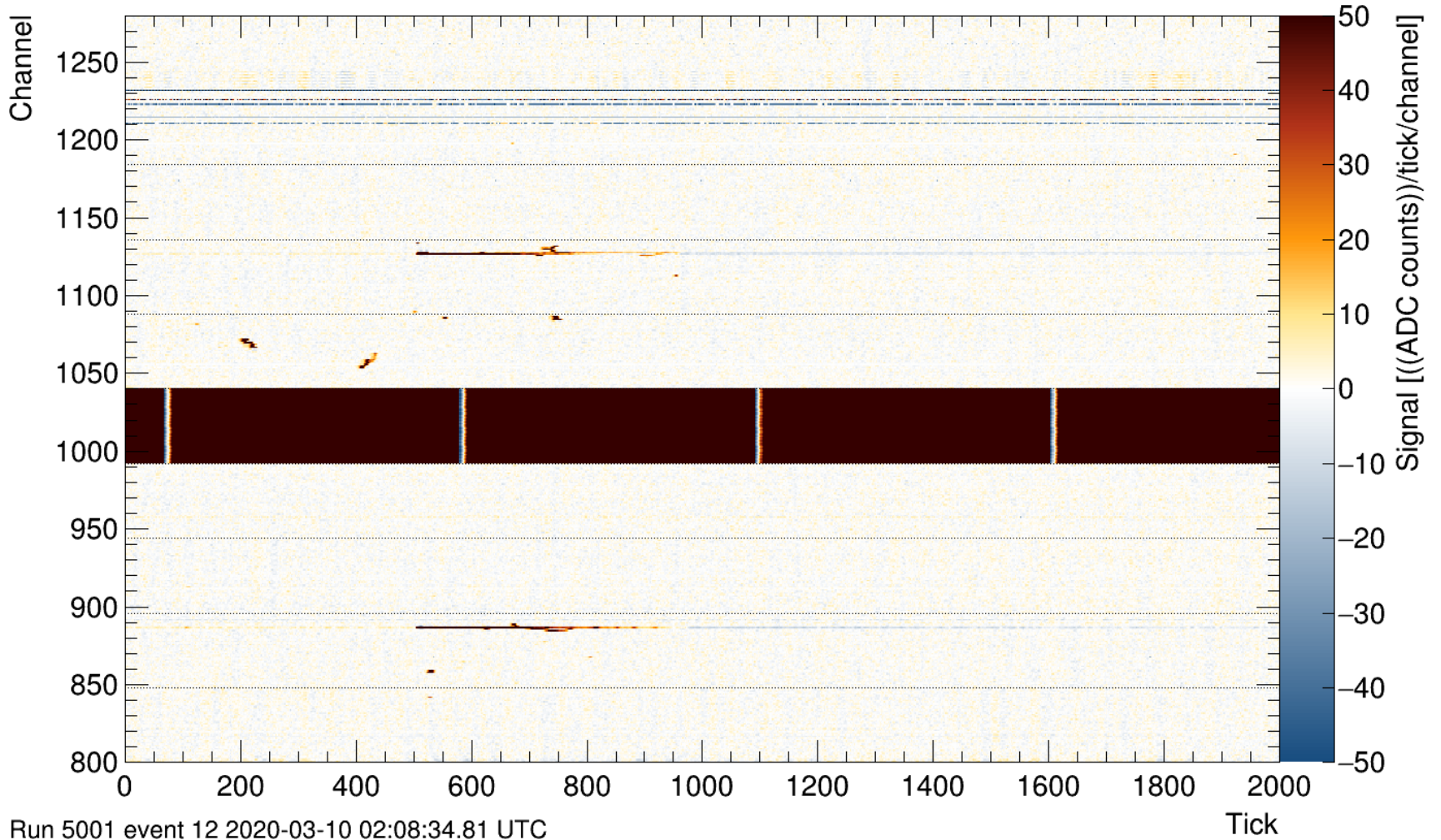
Expected signal for  $\theta_{xz} \approx 0$



- I.e. track perpendicular to wire plane
- MIP produces 62 ke/cm after recombination
  - Smearing from shaping, diffusion have little effect on these tracks
    - As long as signal does not migrate to neighboring wire
  - Attenuated by electron lifetime
  - Boosted by  $1/\cos(\theta_{zy})$  (Small effect with out trigger)
- Drift speed  $1.6 \text{ mm}/\mu\text{s} = 0.8 \text{ mm}/\text{tick} \rightarrow 5.0 \text{ ke}/\text{tick}$ 
  - Sample noise of 0.06 ke  $\rightarrow$  SNR = 85
- CE response 40 ADC/ke  $\rightarrow$  200 ADC/tick
  - Examples follows
- Observed waveforms are not flat
  - Exponential decrease with drift time due to electron lifetime
  - Bumps in spectra from delta rays--expect 20% signal in deltas
  - Signal is doubled where track passes by collection wire
  - Many example follow

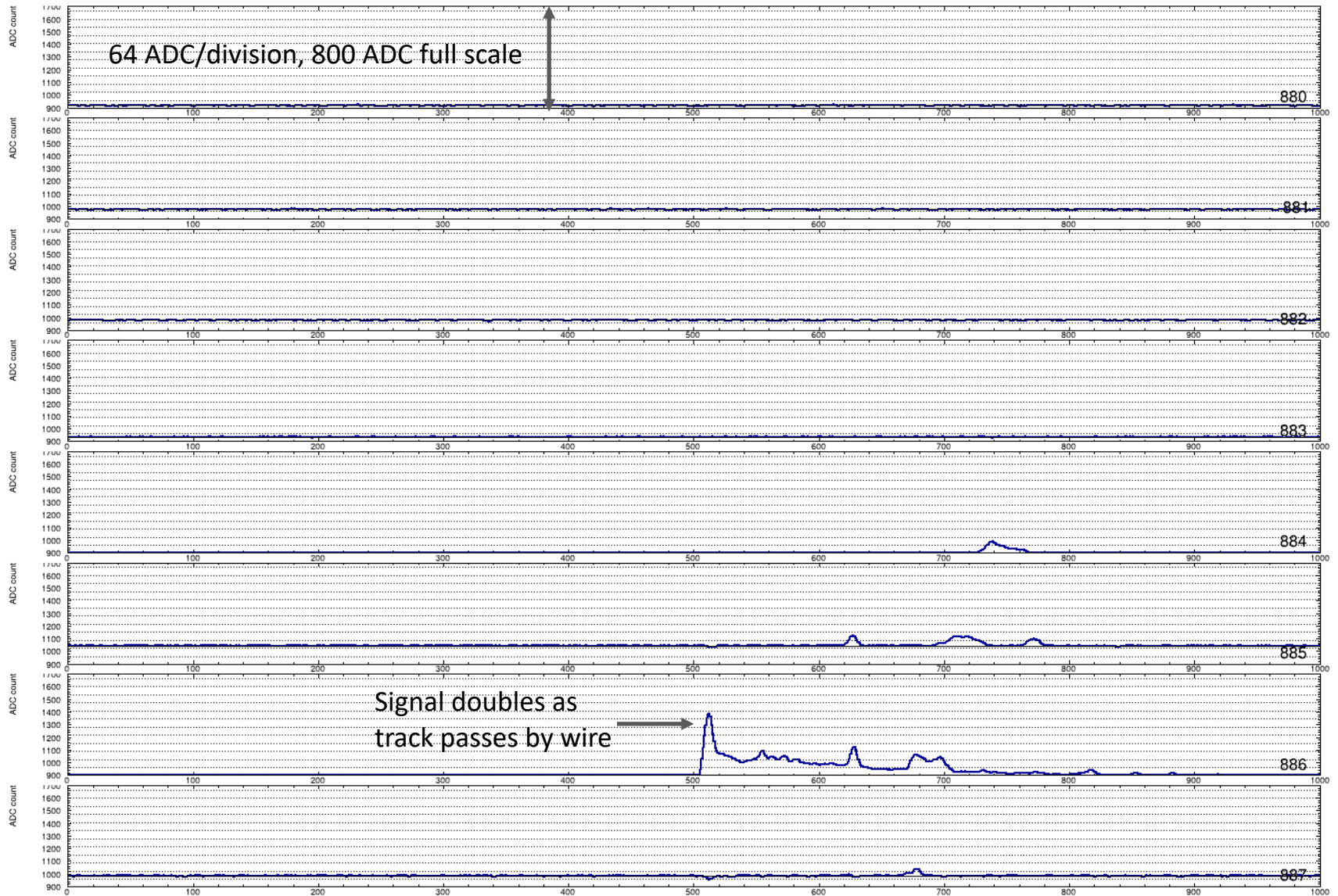
# 5001-12

Raw ADC for Iceberg collection planes



# 5001-12 z1

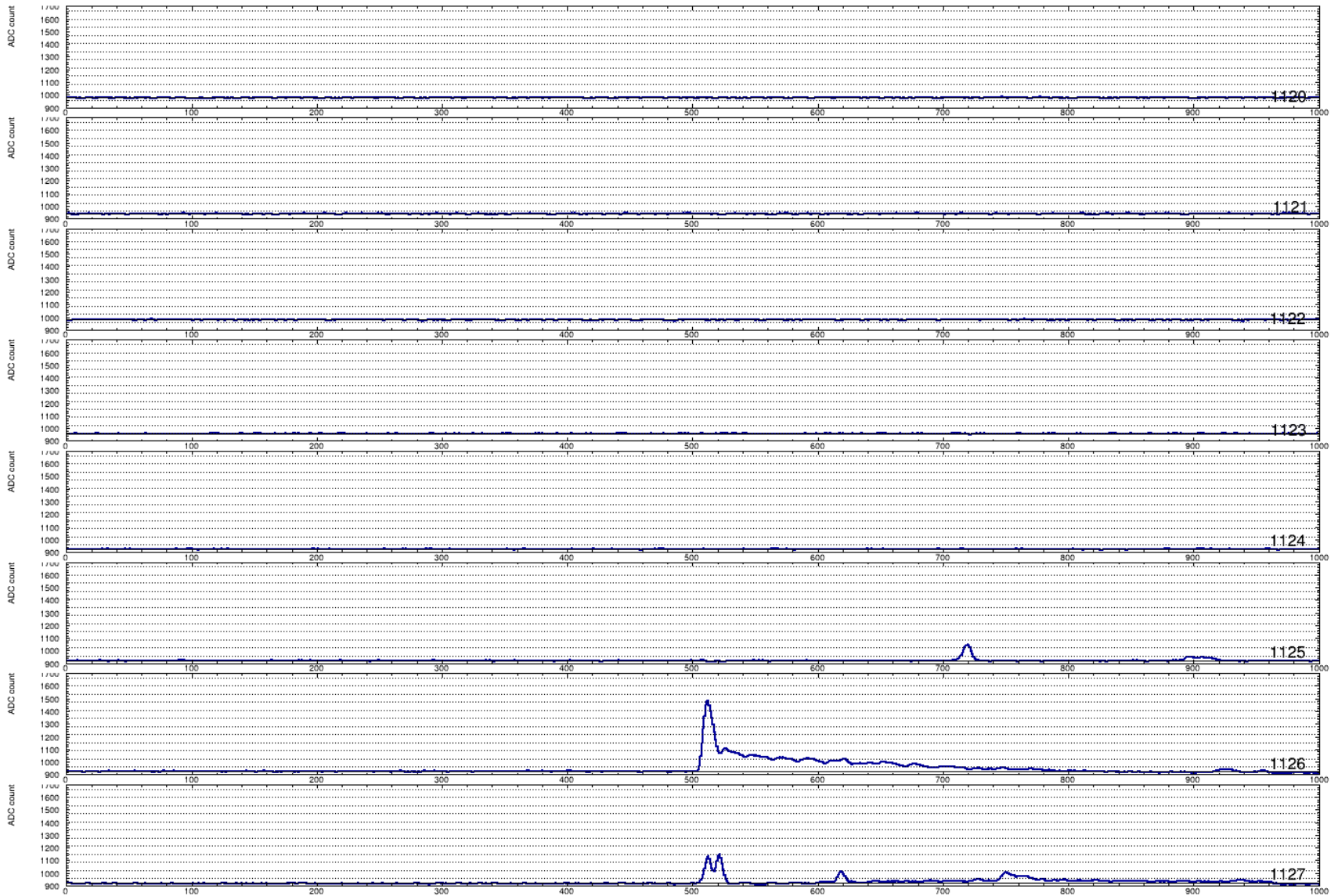
ADC raw run 5001 event 12





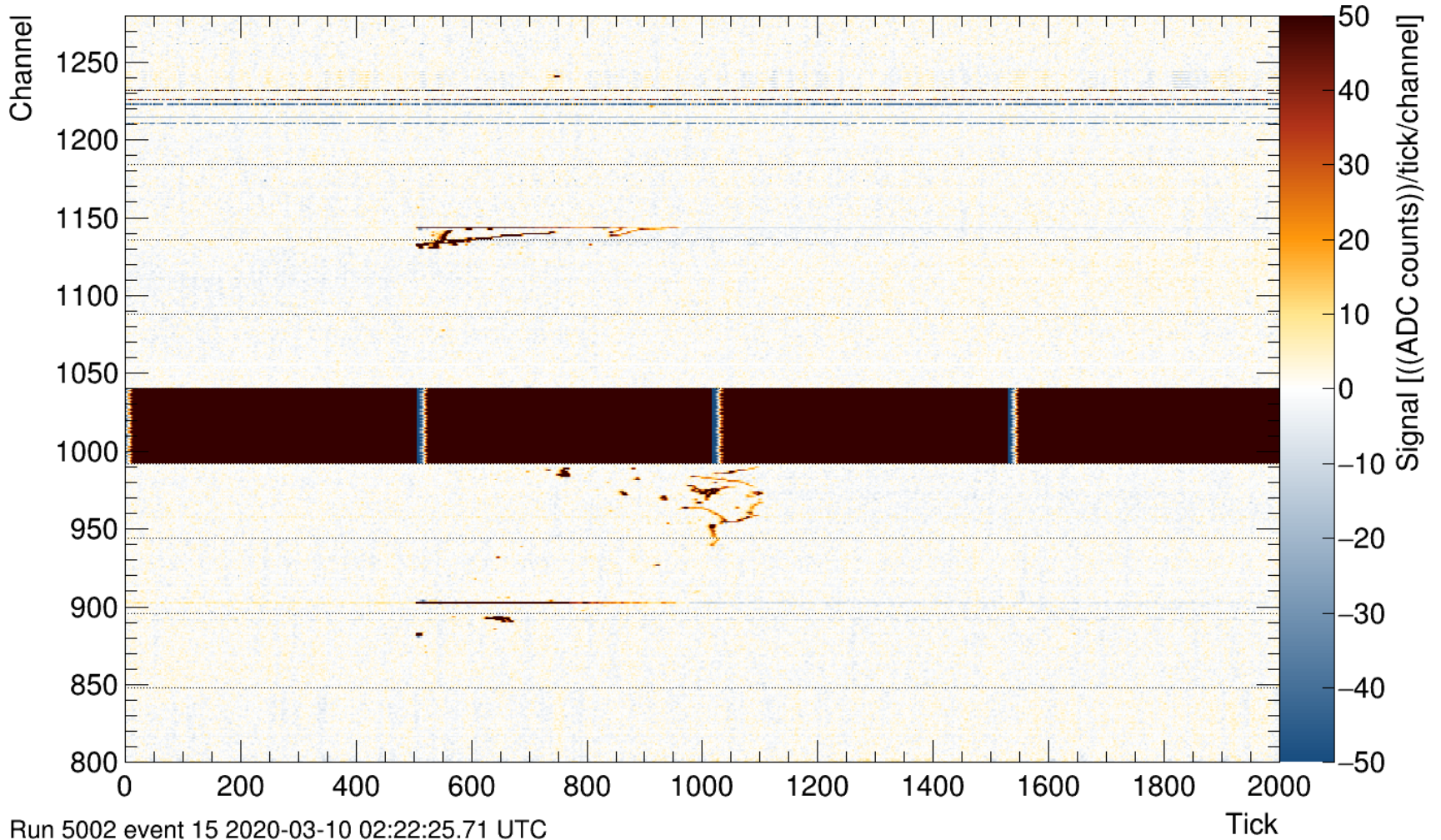
# 5001-12 z2

ADC raw run 5001 event 12



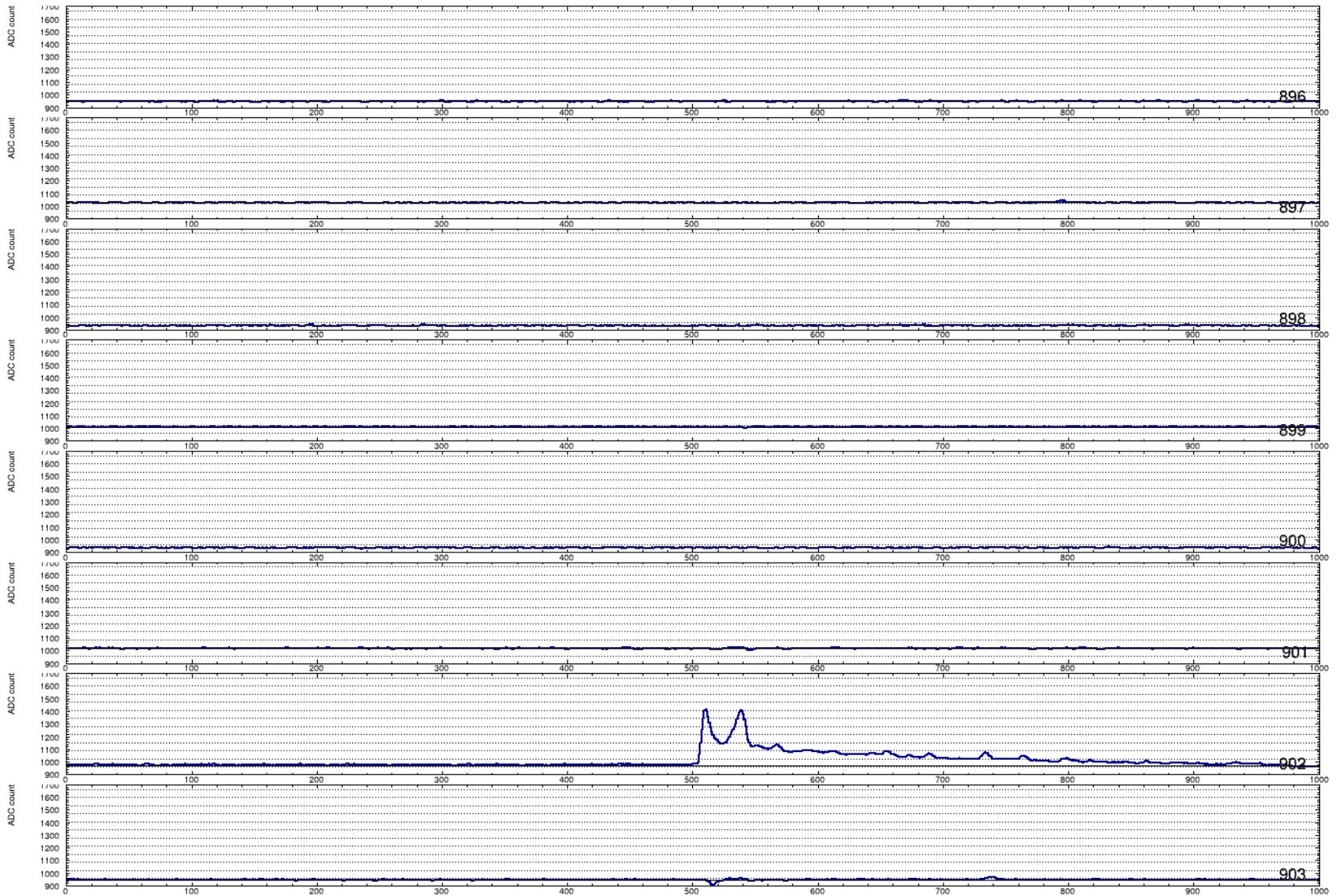
# 5001-15

Raw ADC for Iceberg collection planes



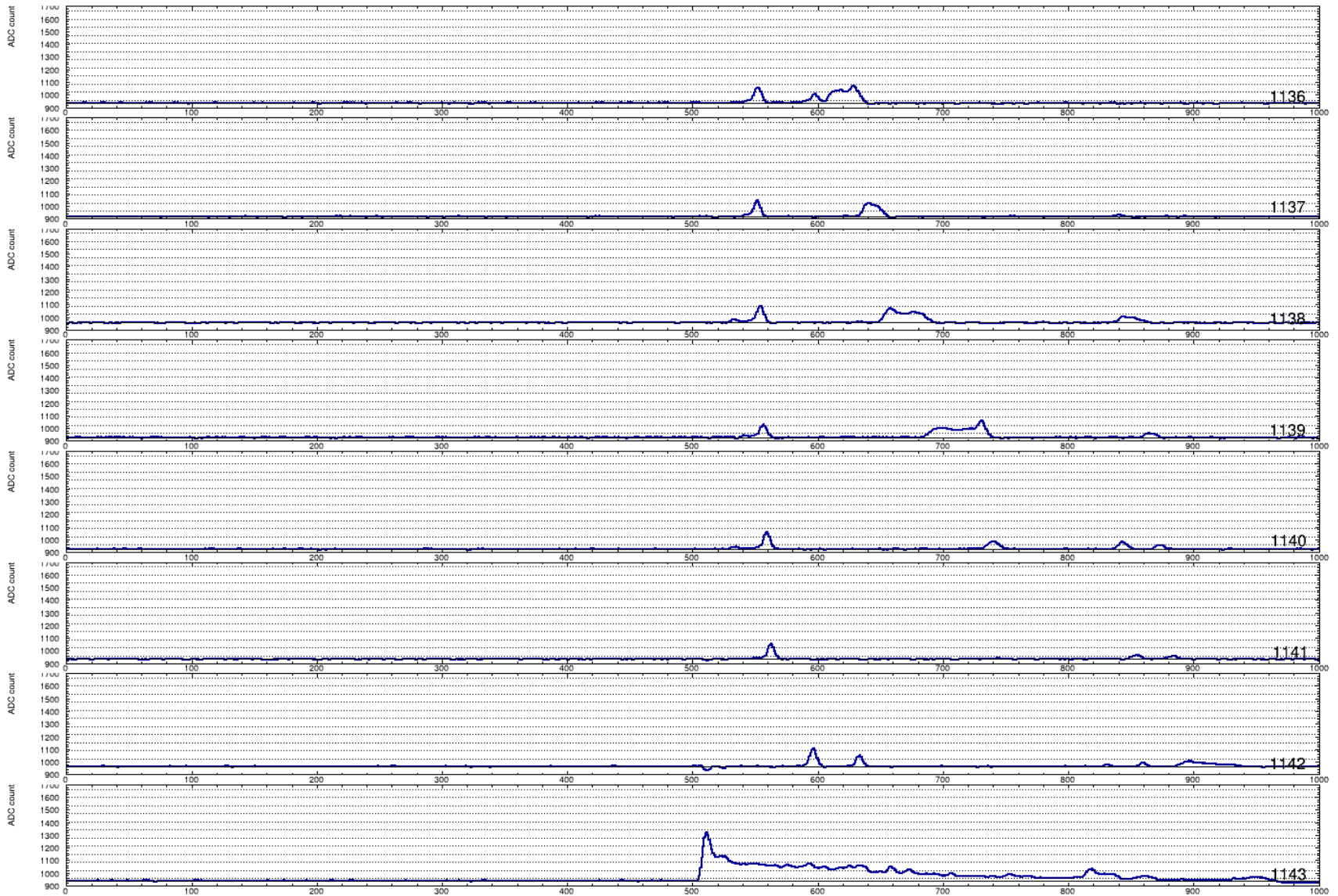
# 5001-15 z1

ADC raw run 5002 event 15



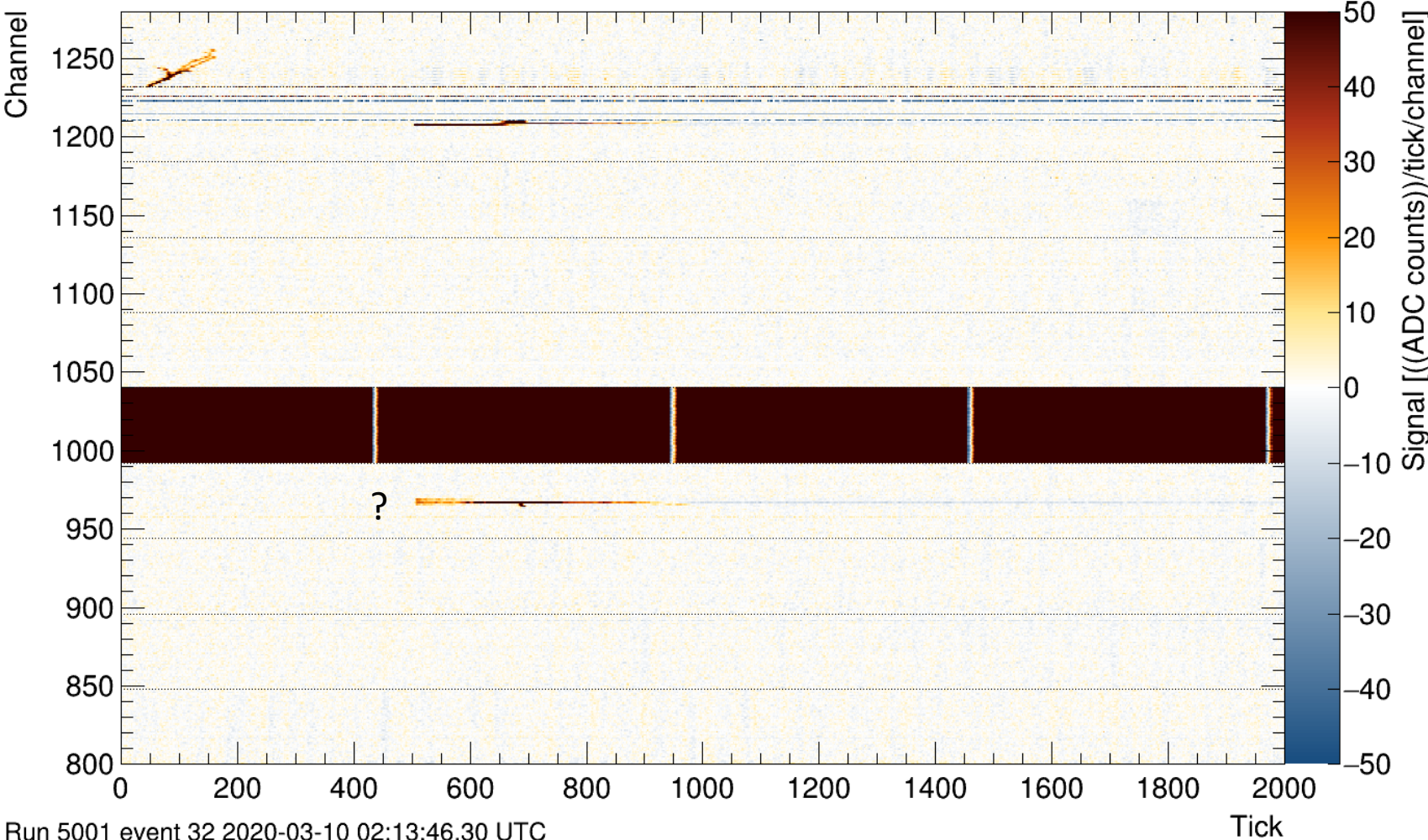
# 5001-15 z2

ADC raw run 5002 event 15



# 5001-32

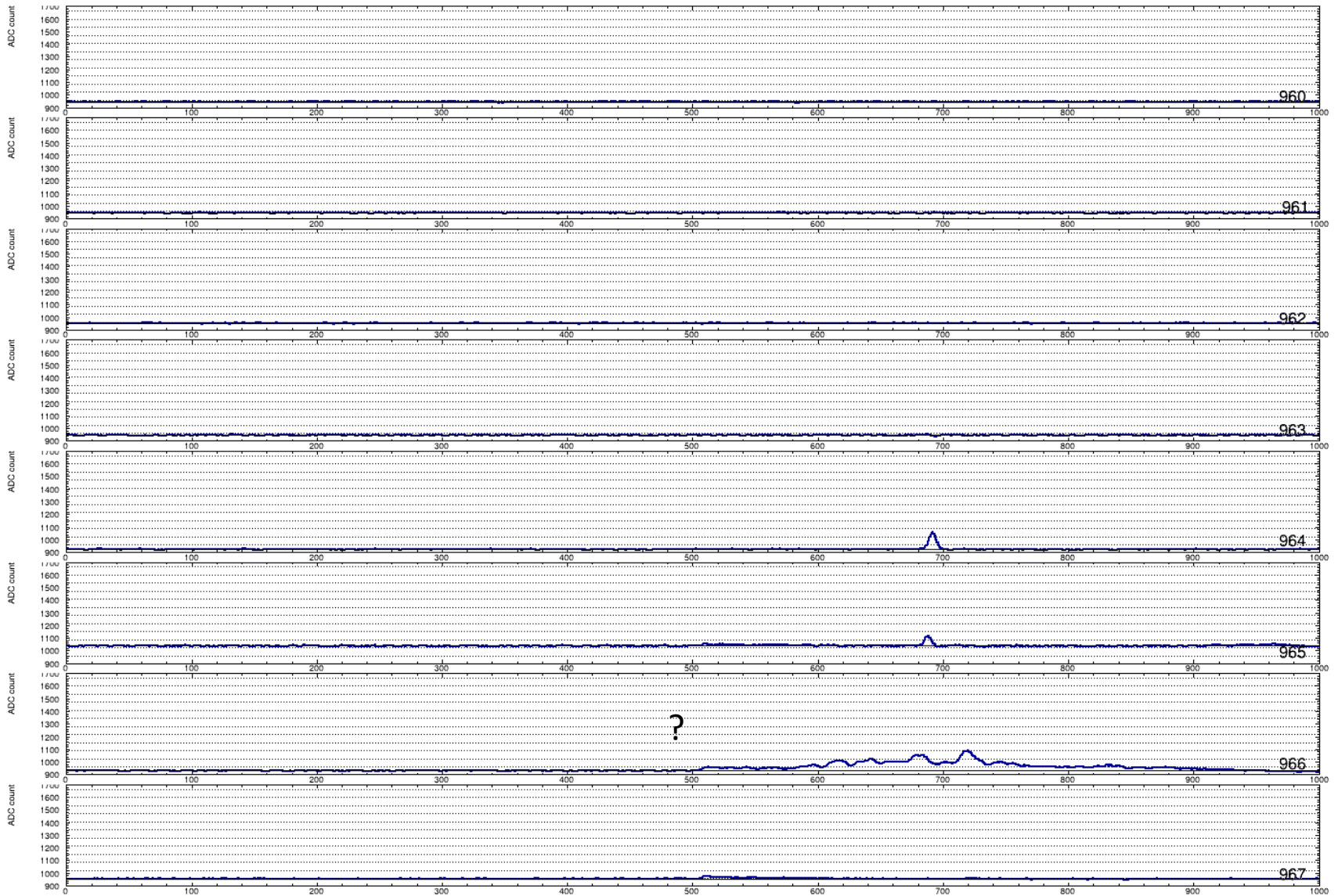
Raw ADC for Iceberg collection planes



Run 5001 event 32 2020-03-10 02:13:46.30 UTC

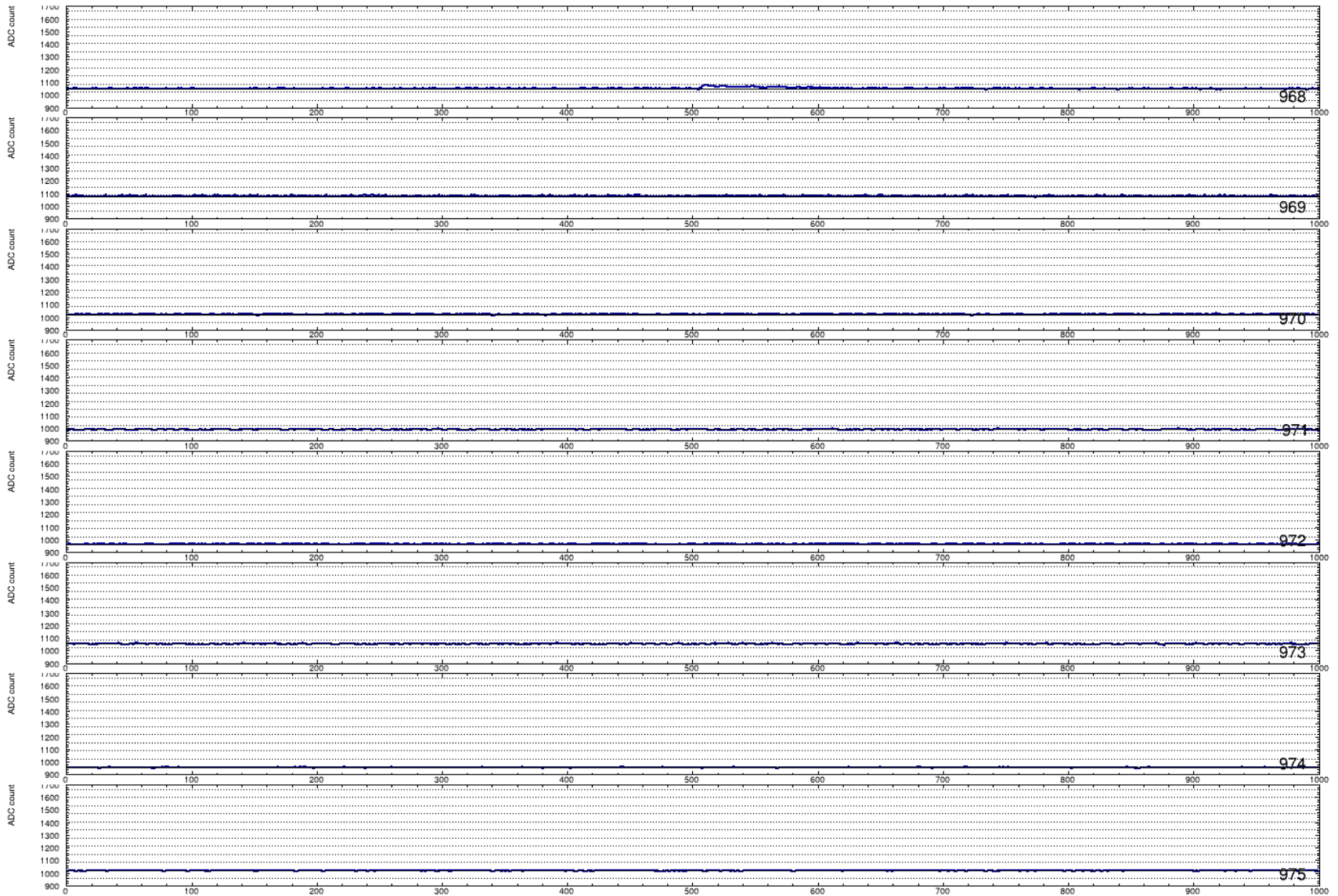
# 5001-32 z1

ADC raw run 5001 event 32



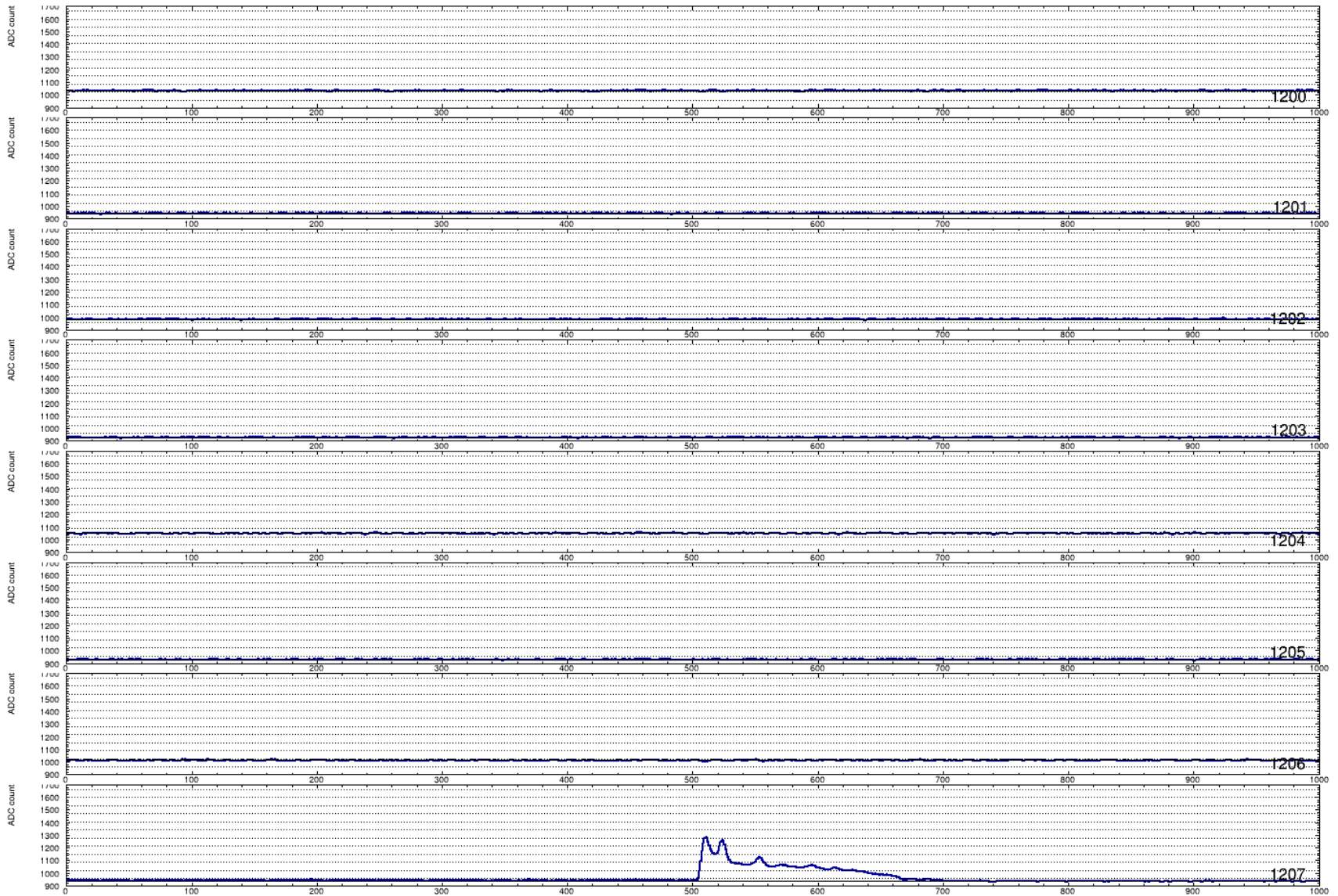
# 5001-32 z1

ADC raw run 5001 event 32



# 5001-32 z2

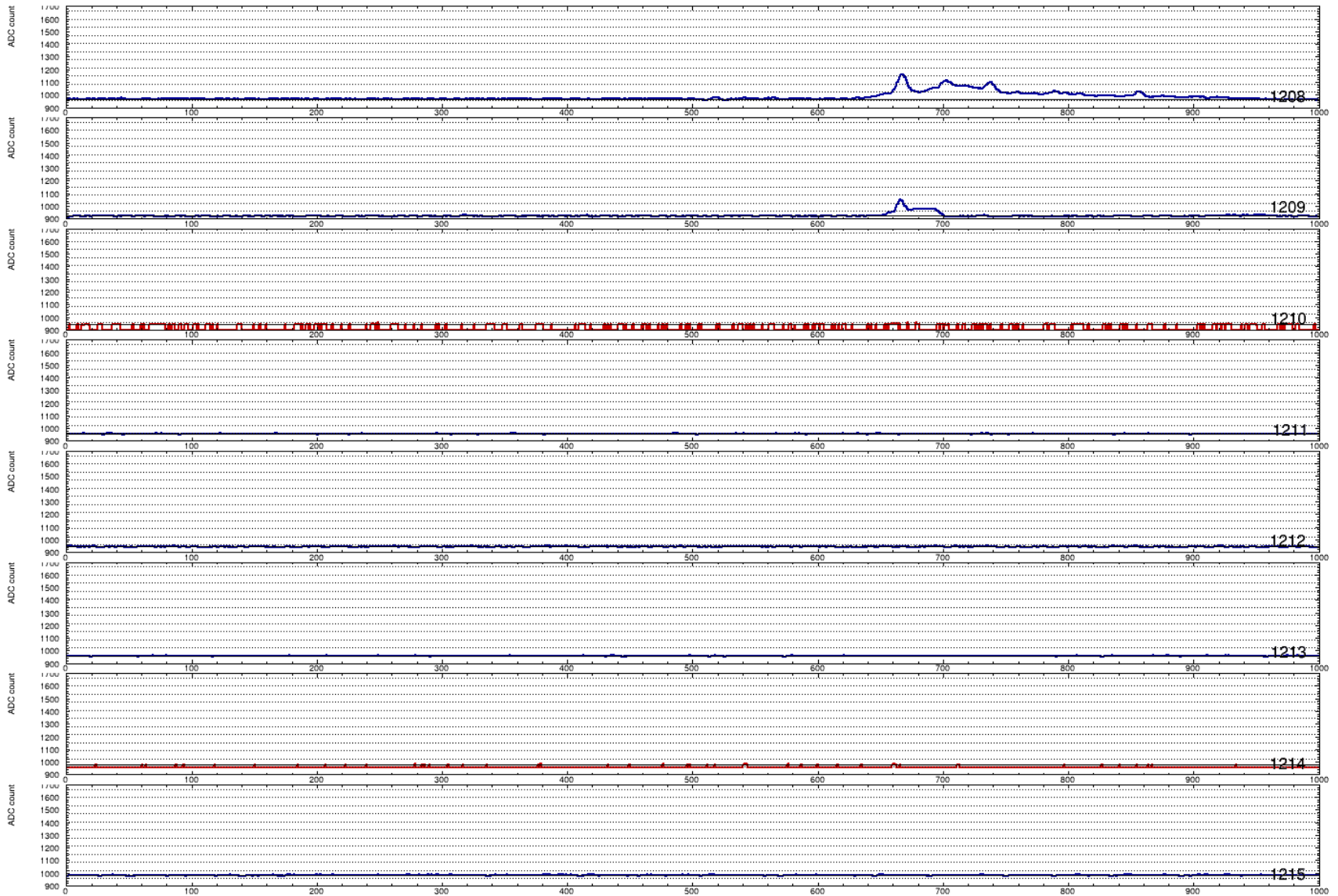
ADC raw run 5001 event 32





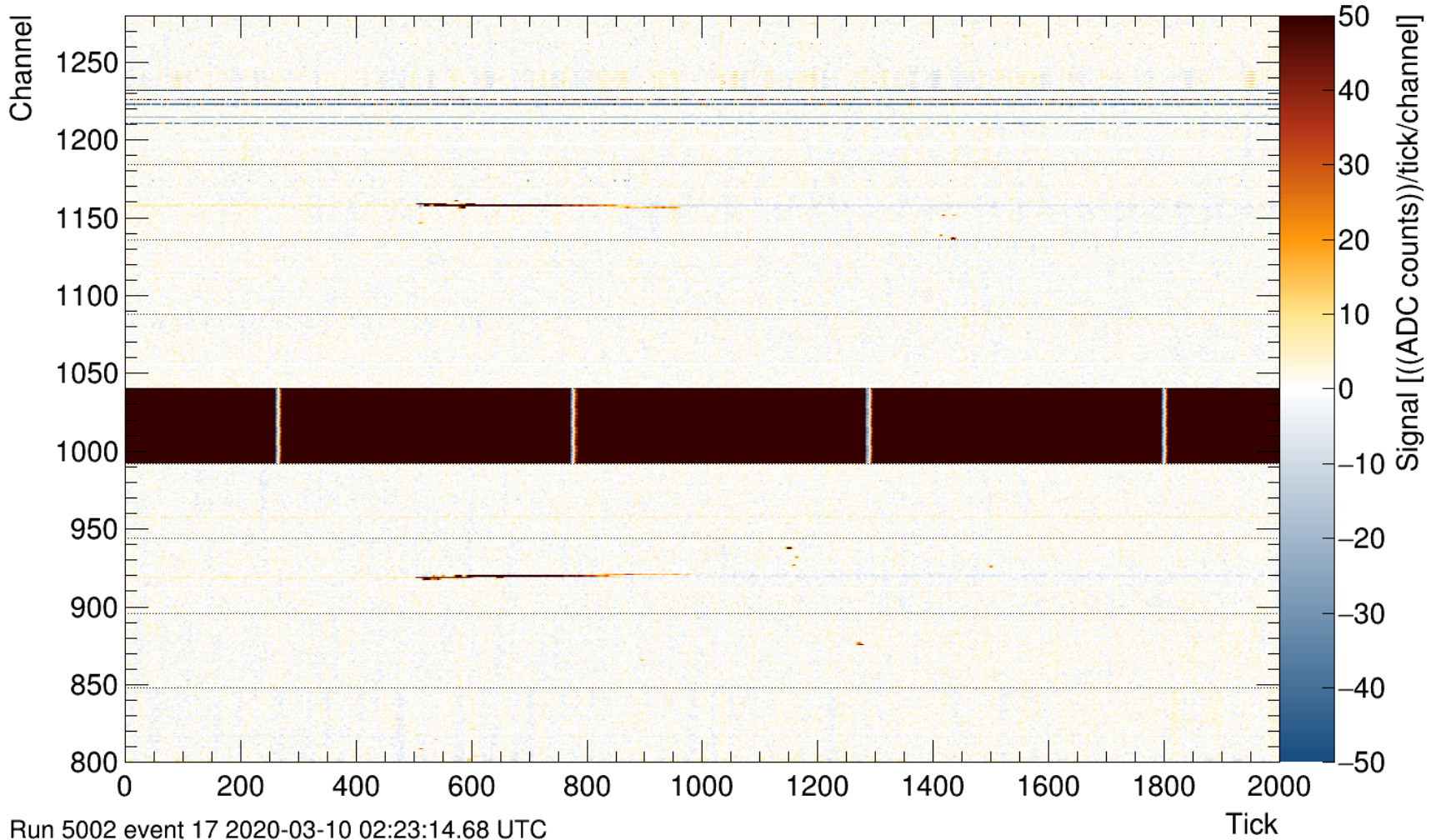
# 5001-32 z2

ADC raw run 5001 event 32



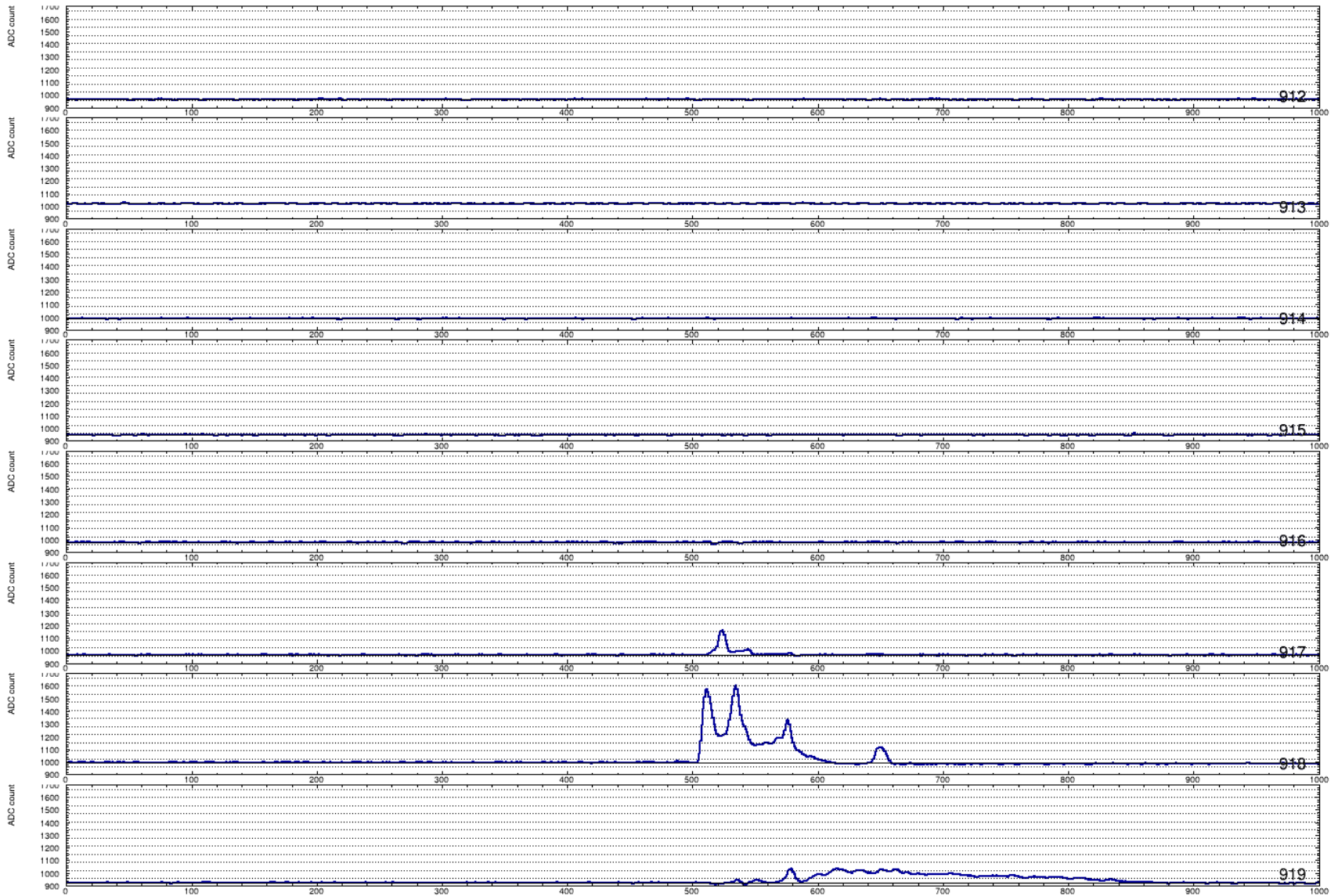
# 5002-17

Raw ADC for Iceberg collection planes



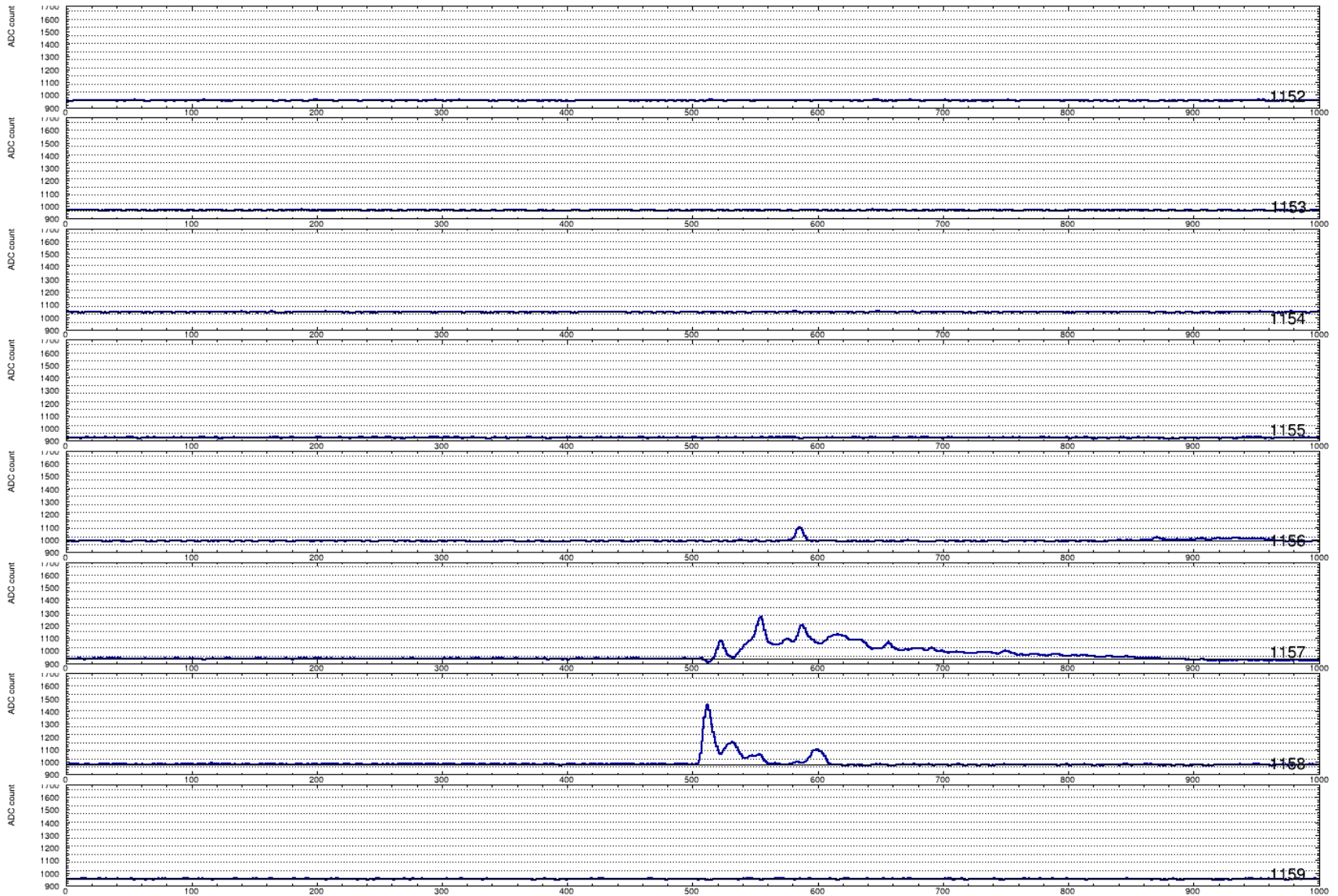
# 5002-17 z1

ADC raw run 5002 event 17



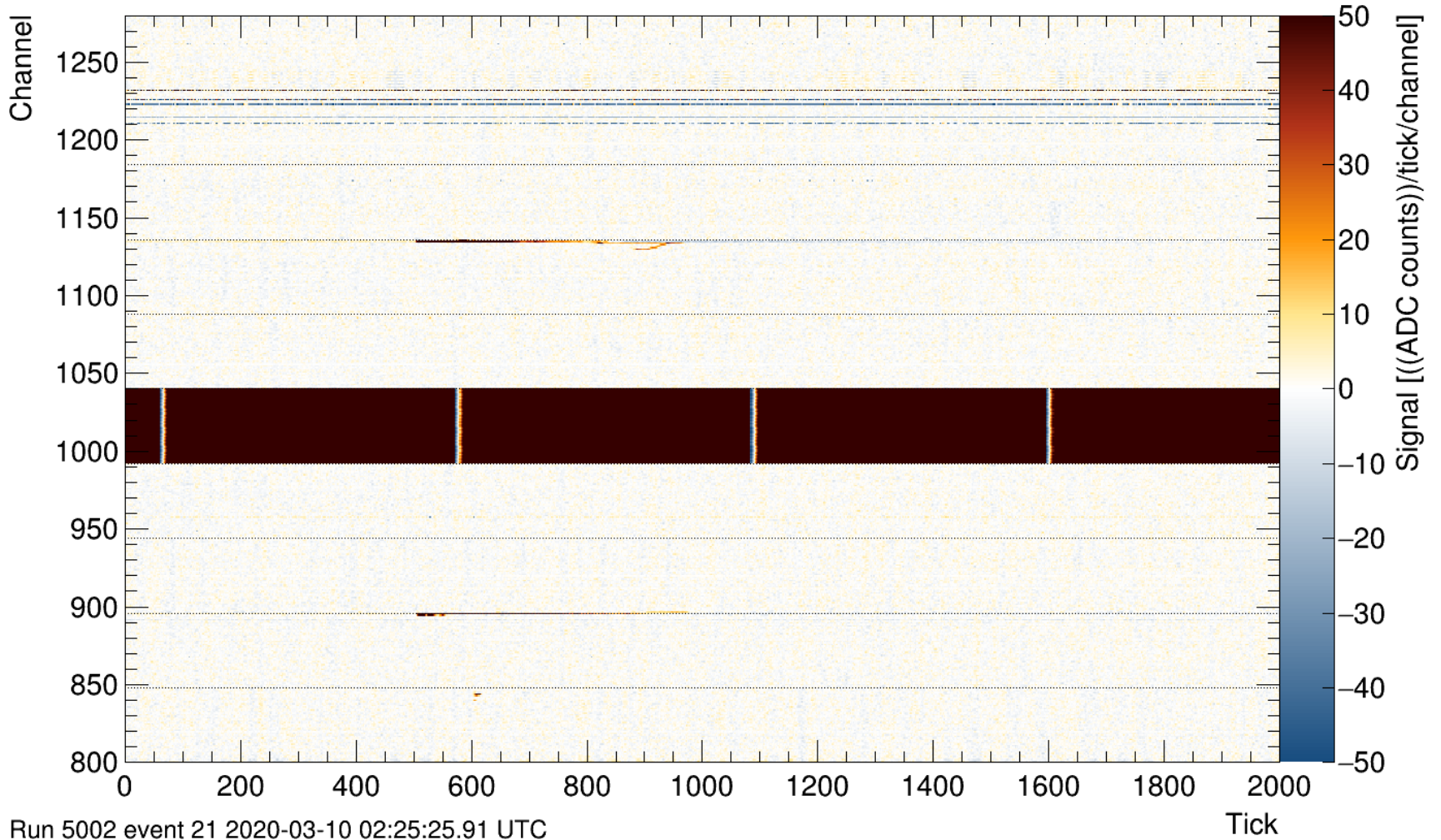
# 5002-17 z2

ADC raw run 5002 event 17



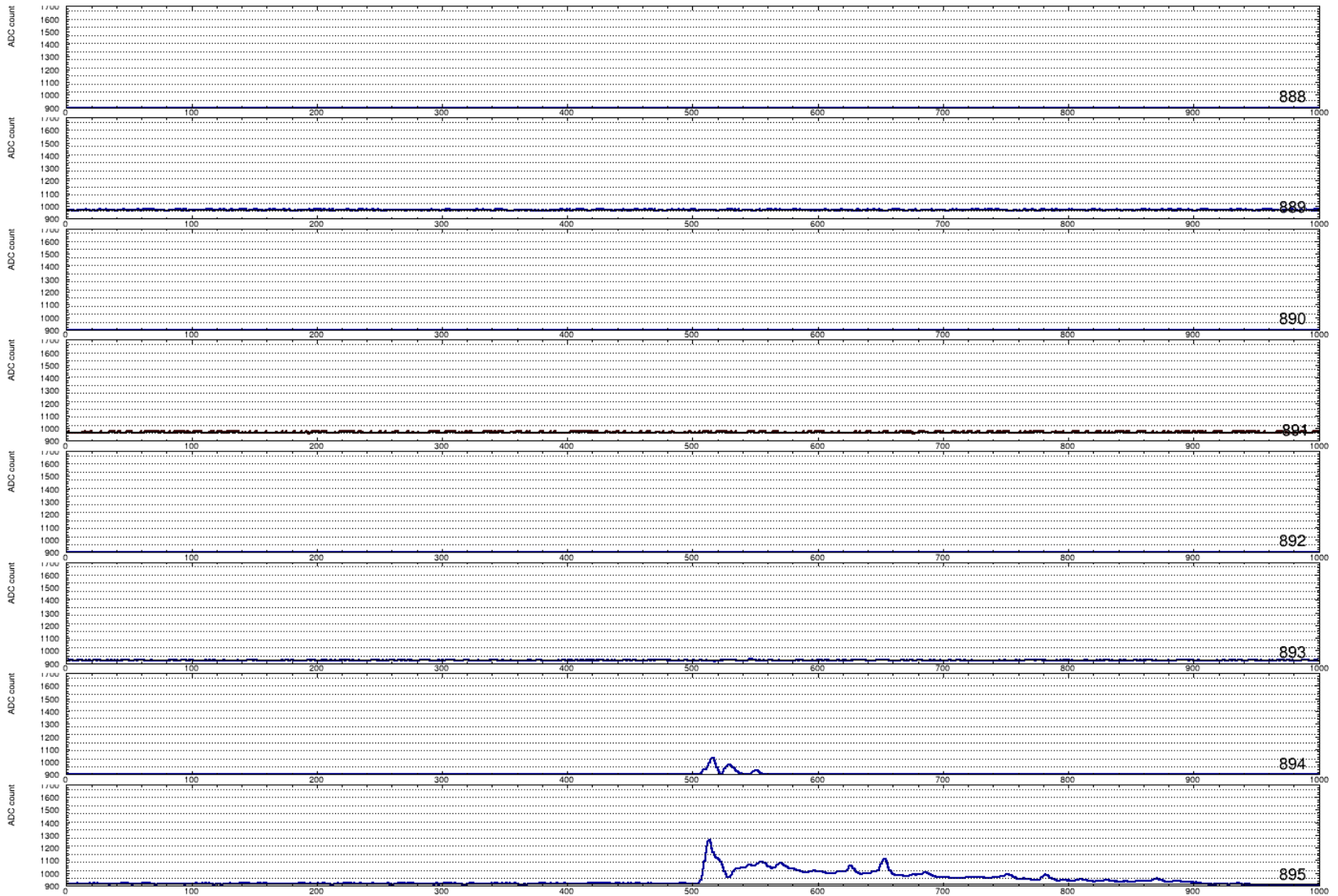
# 5002-21

Raw ADC for Iceberg collection planes



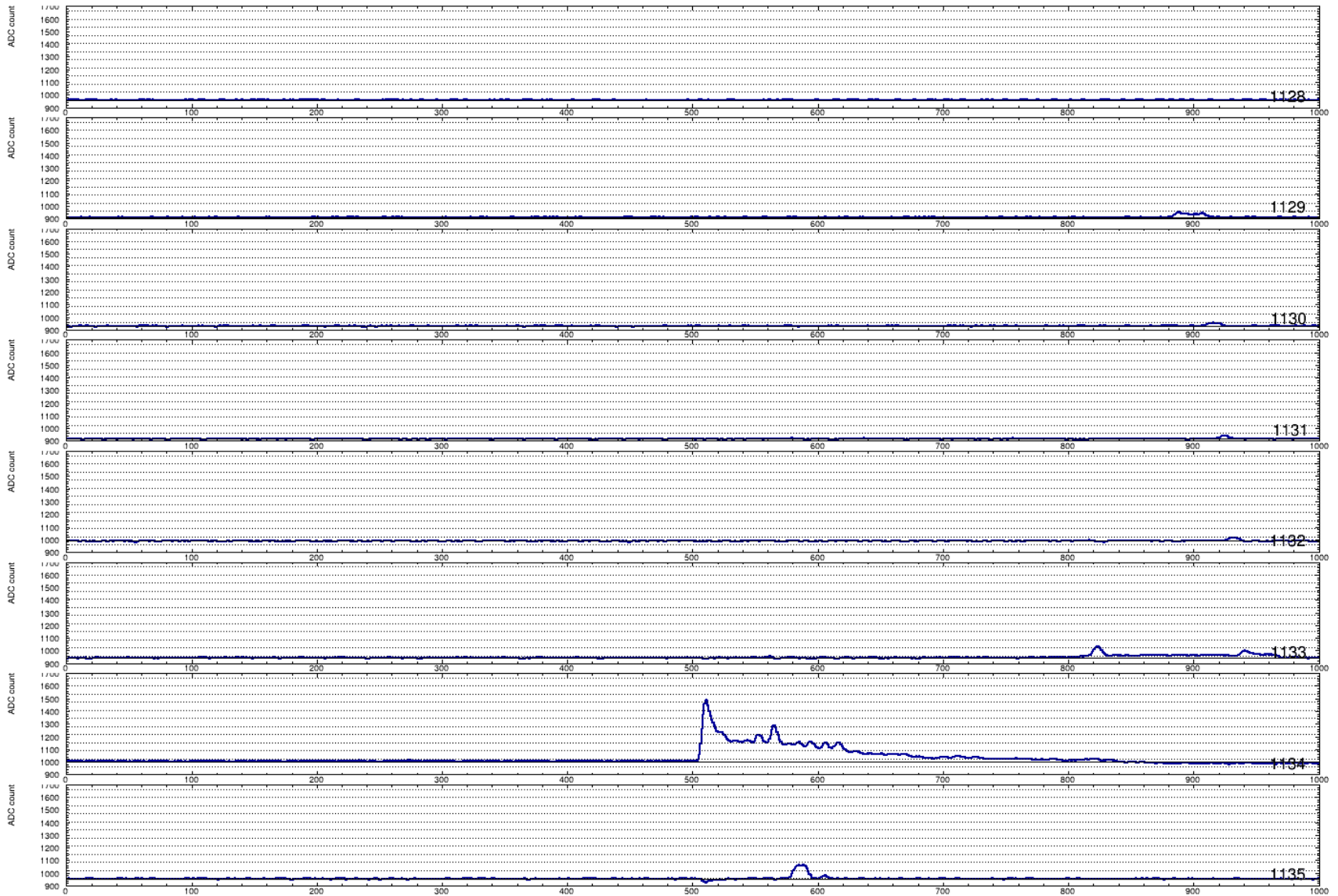
# 5002-21 z1

ADC raw run 5002 event 21



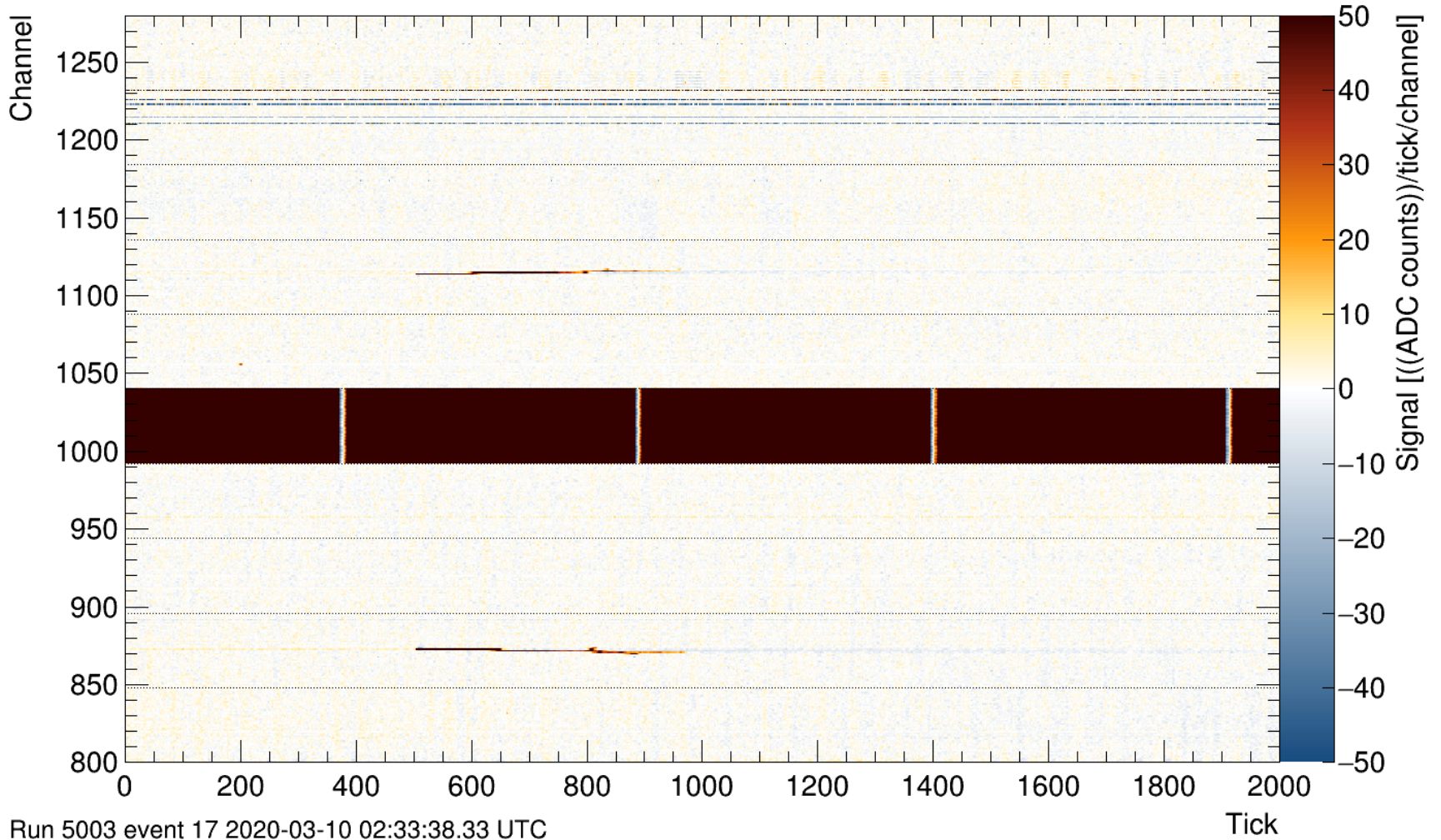
# 5002-21 z2

ADC raw run 5002 event 21



# 5003-17

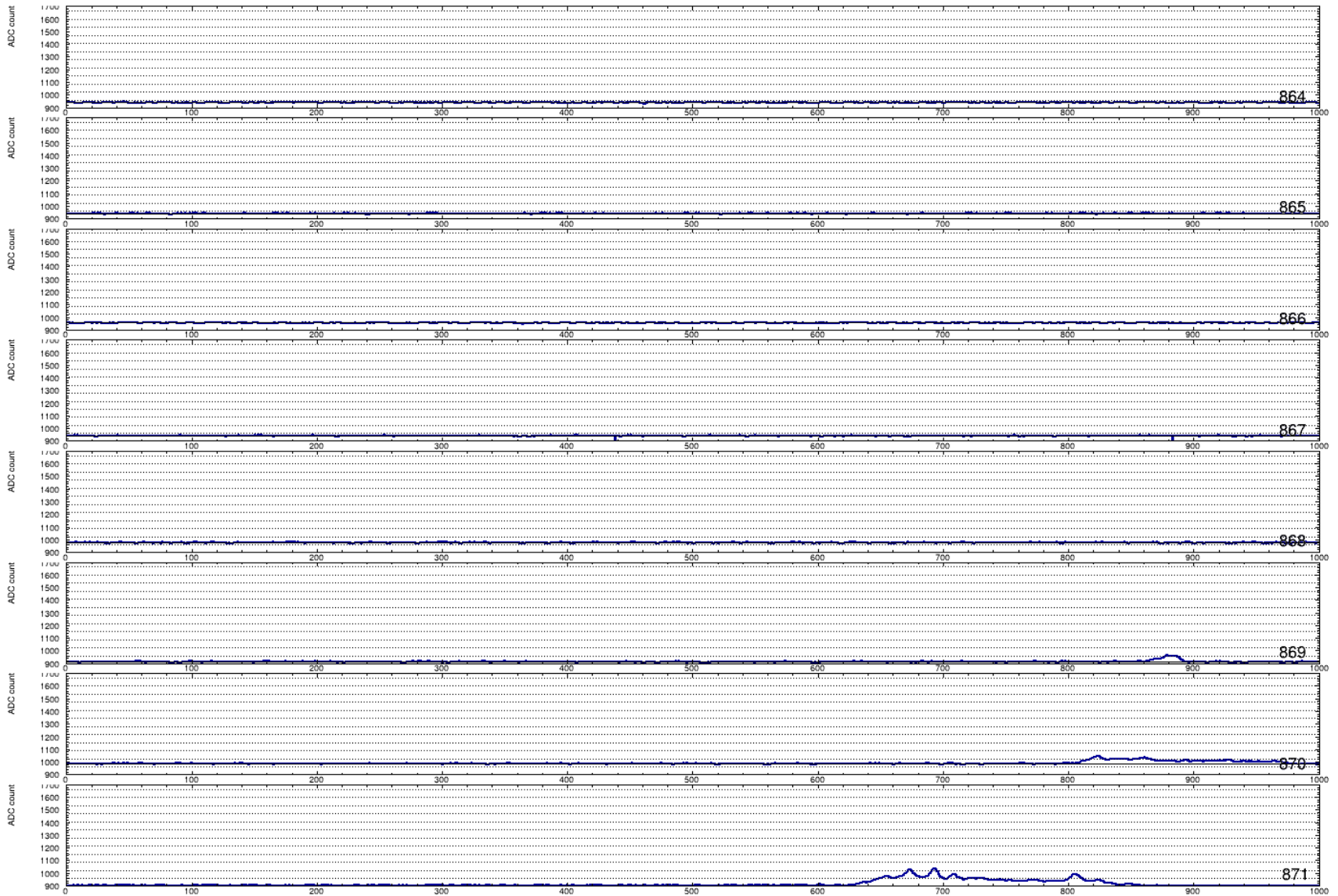
Raw ADC for Iceberg collection planes





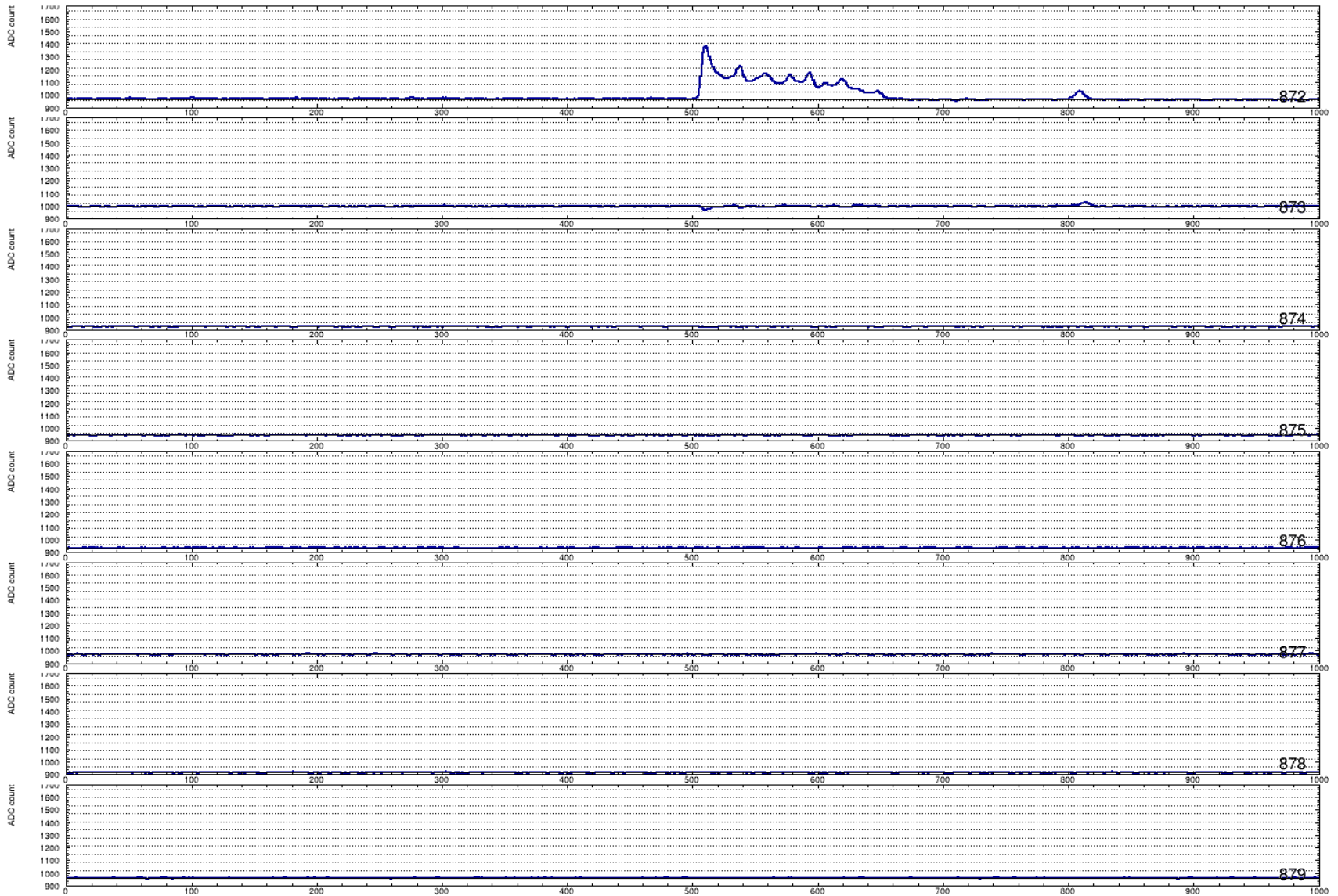
# 5003-17 z1

ADC raw run 5003 event 17



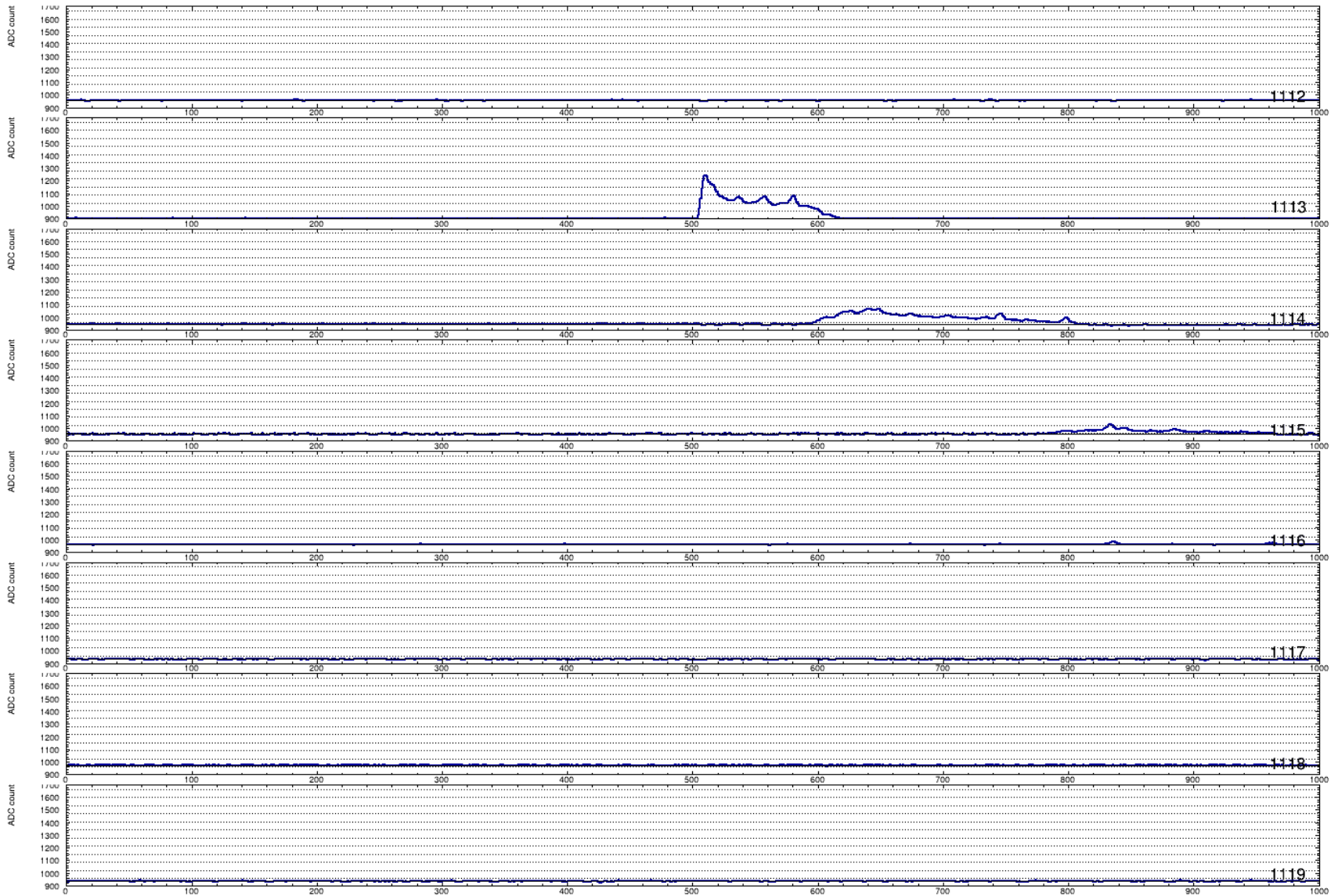
# 5003-17 z1

ADC raw run 5003 event 17



# 5003-17 z2

ADC raw run 5003 event 17



# Comments on perpendicular tracks

We do observe large signals

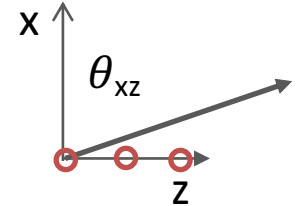
- Height around 200 – 250 ADC counts
  - After correcting (by eye) for electron lifetime and doubling of signal near collection wire
- Prediction was 200 ADC counts without accounting for vertical angle
  - I.e. signal is enhanced by  $1/\cos(\theta_{xy})$
  - Trigger counters are at 25 deg (near horizontal)
  - → enhancement is 1.09, i.e. expect 220 ADC
  - Consistent

# Non-perpendicular tracks

# Tracks not perpendicular to wire plane

Standard MIP signal:  $\theta_{xz} \approx 90$  deg

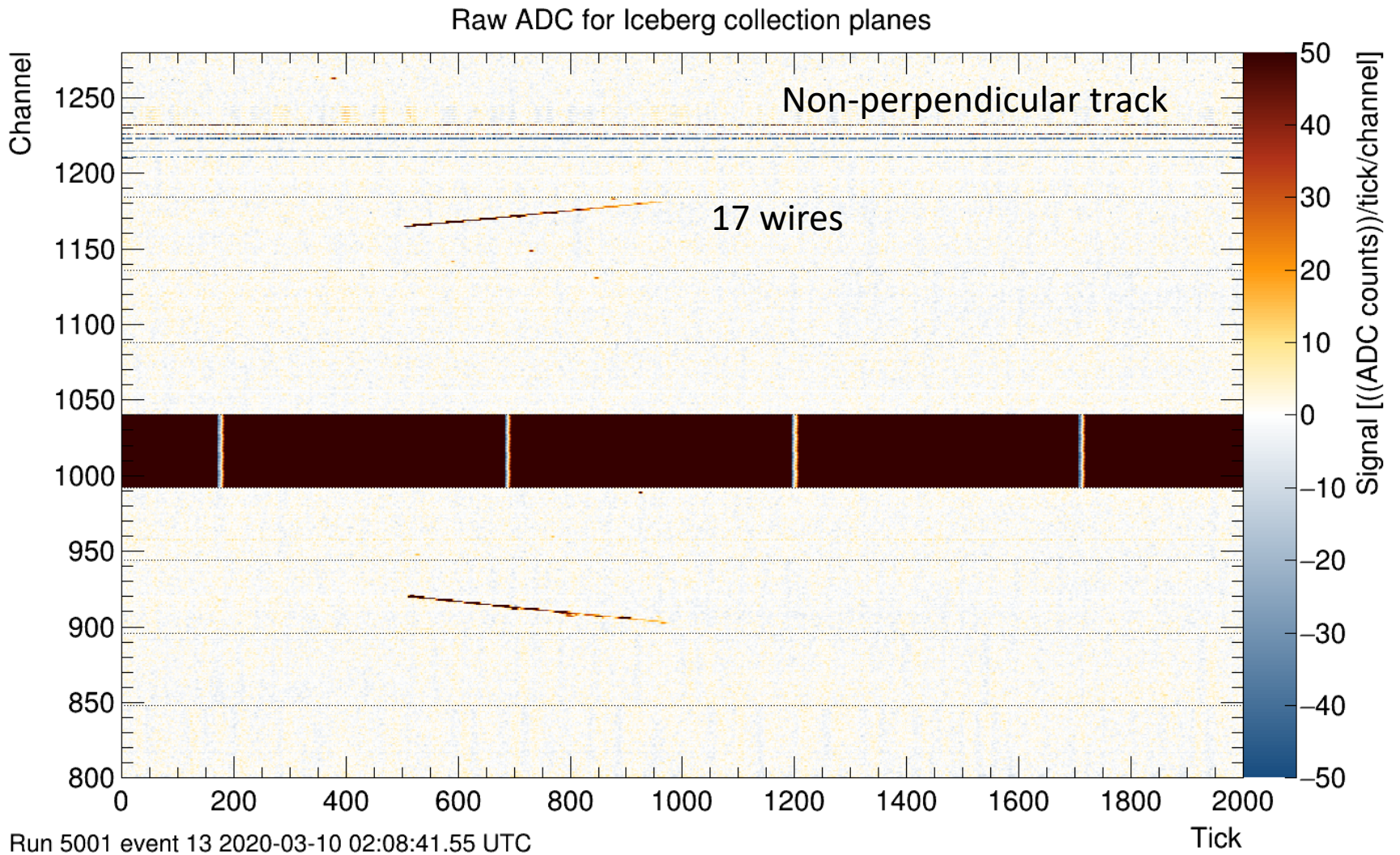
- I.e. nearly parallel to wire plane
- Integrated charge/wire is (4.8 mm) (62 ke/cm) = 30 ke
  - For track in horizontal plane
- Peak height is area/5.58 from shaping
  - plus smearing due to path lengths and diffusion
  - → Peak height  $\lesssim 5$  ke = 200 ADC
- I.e. height is about the same as perpendicular tracks



Trigger counters accept  $|\theta_{xz}| \lesssim 25$  deg

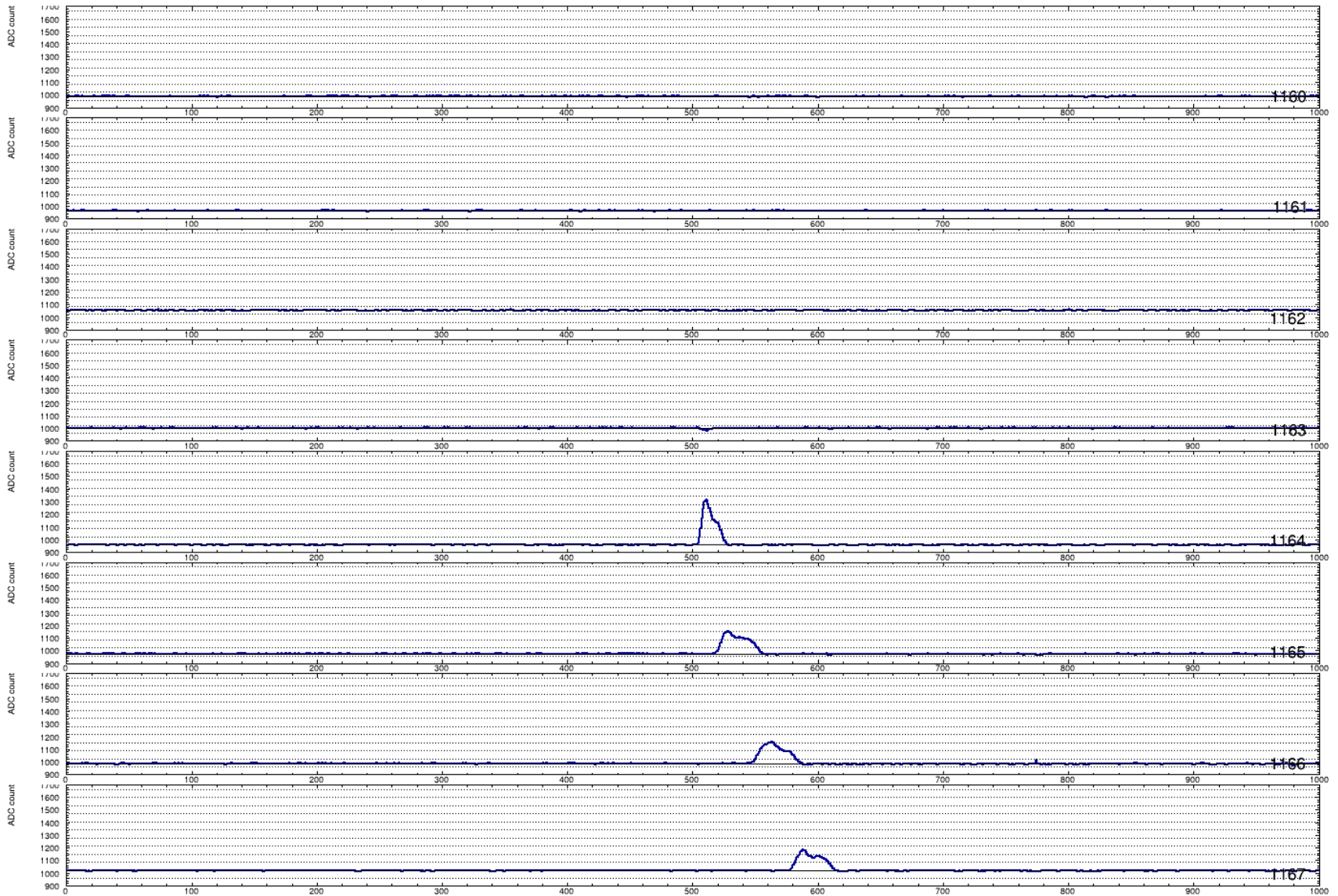
- Far from parallel but there are some tracks significantly away from the perpendicular
- Tracks cross fewer than 30 wires on either side of APA
  - Maybe 20-30 wires are enough for protoDUNE-like SNR study?
- Following pages show a few tracks away from perpendicular

# 5001-13



# 5001-13 z1

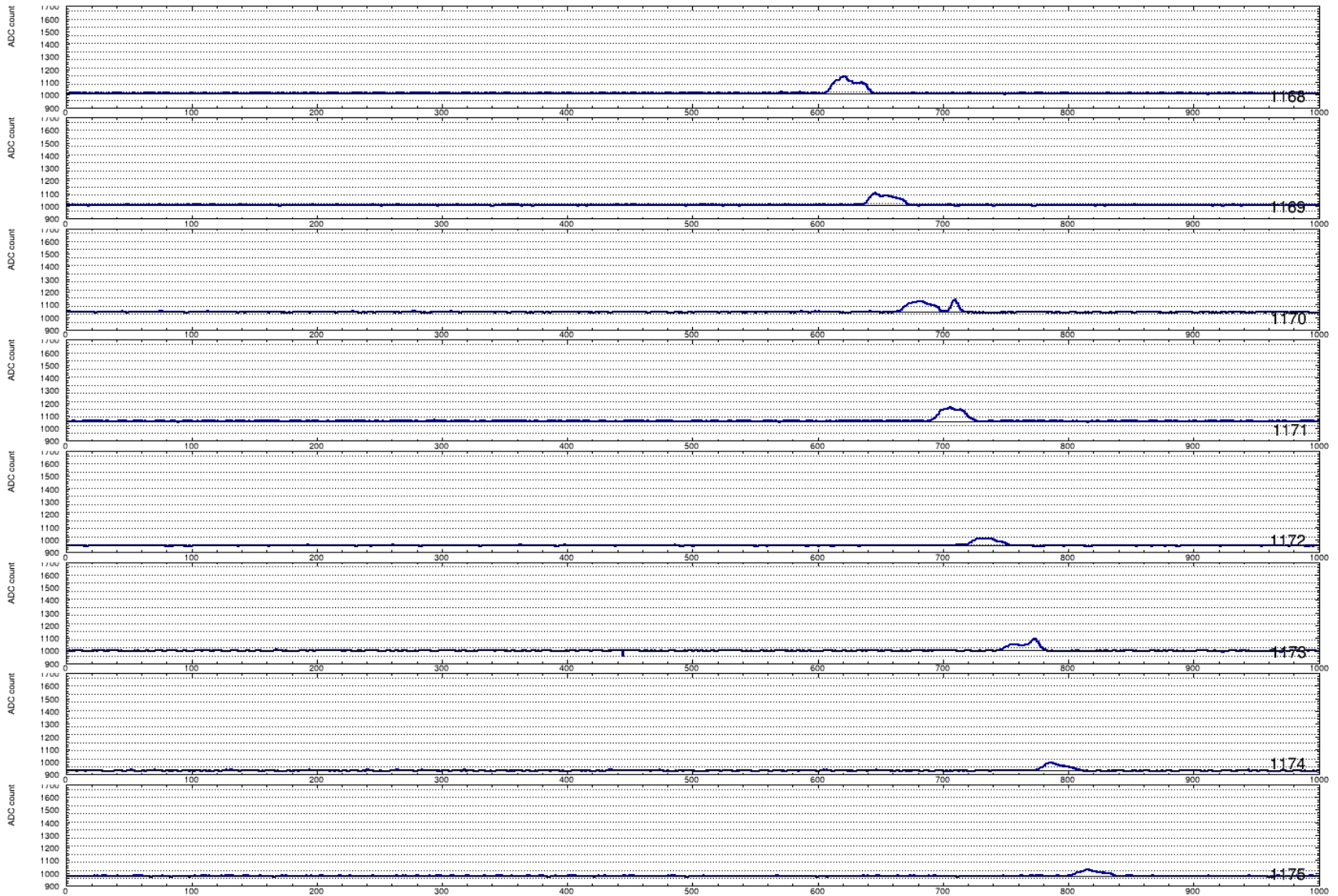
ADC raw run 5001 event 13





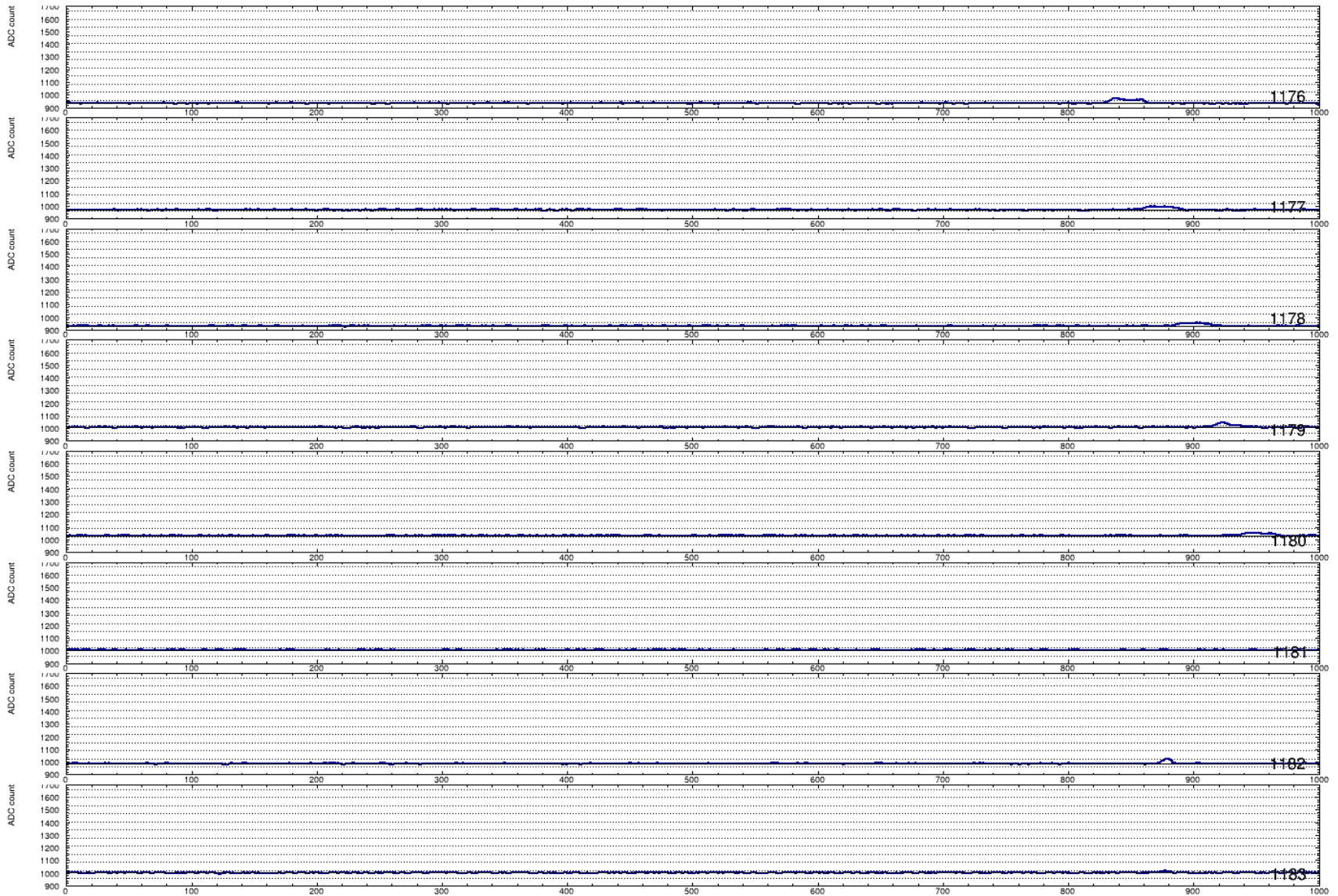
# 5001-13 z1

ADC raw run 5001 event 13



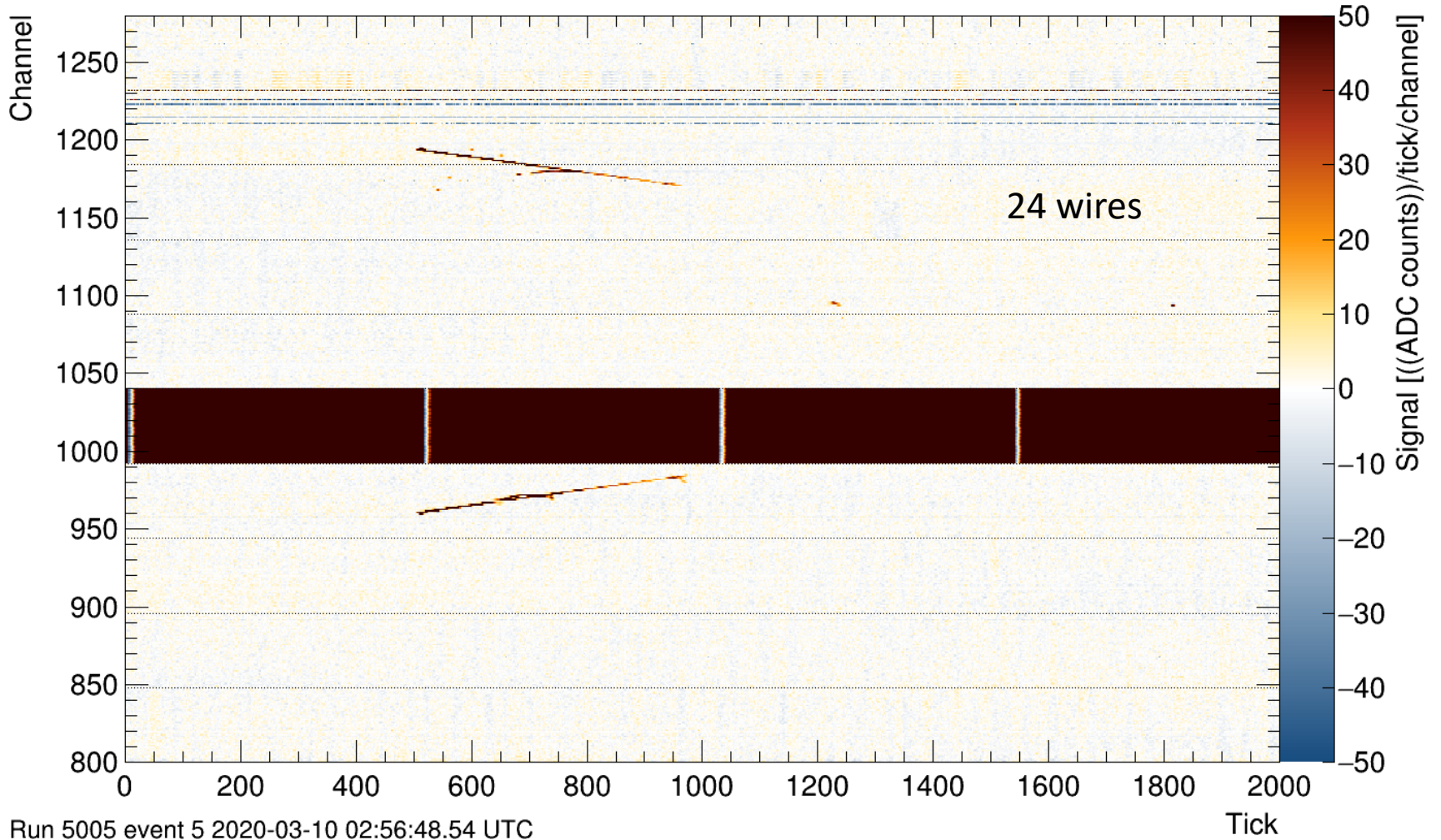
# 5001-13 z1

ADC raw run 5001 event 13



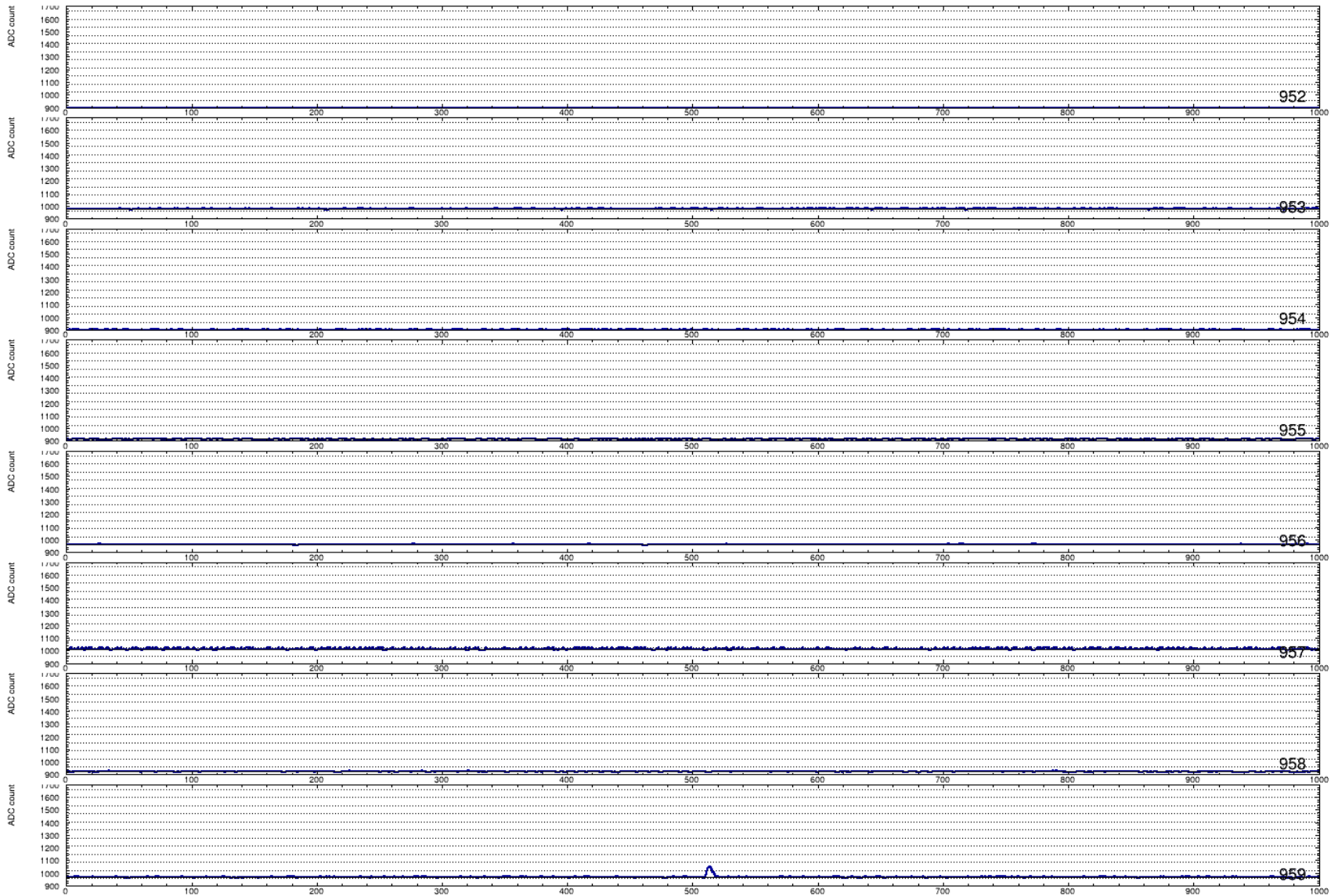
# 5005-5

Raw ADC for Iceberg collection planes



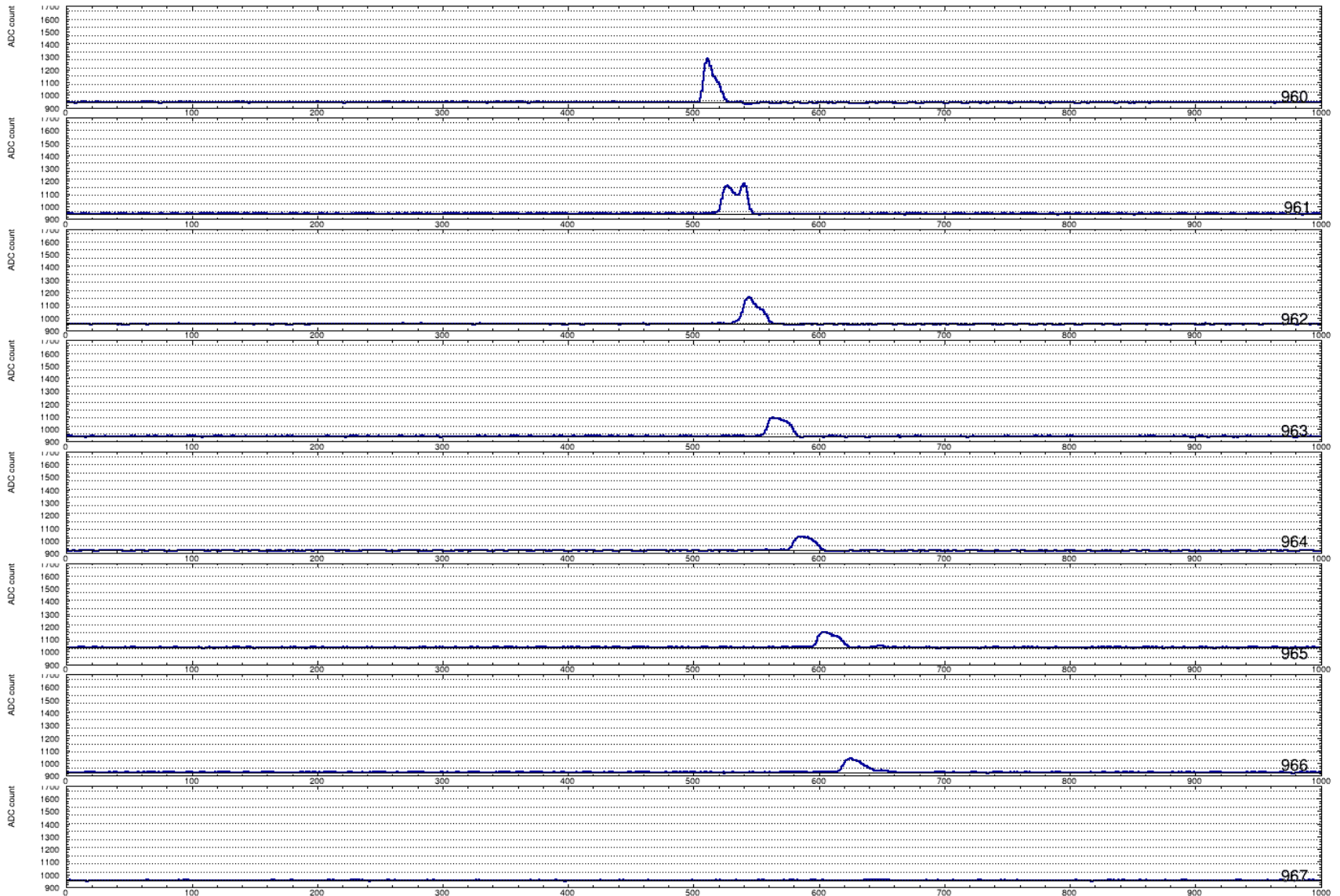
# 5005-5 z1

ADC raw run 5005 event 5



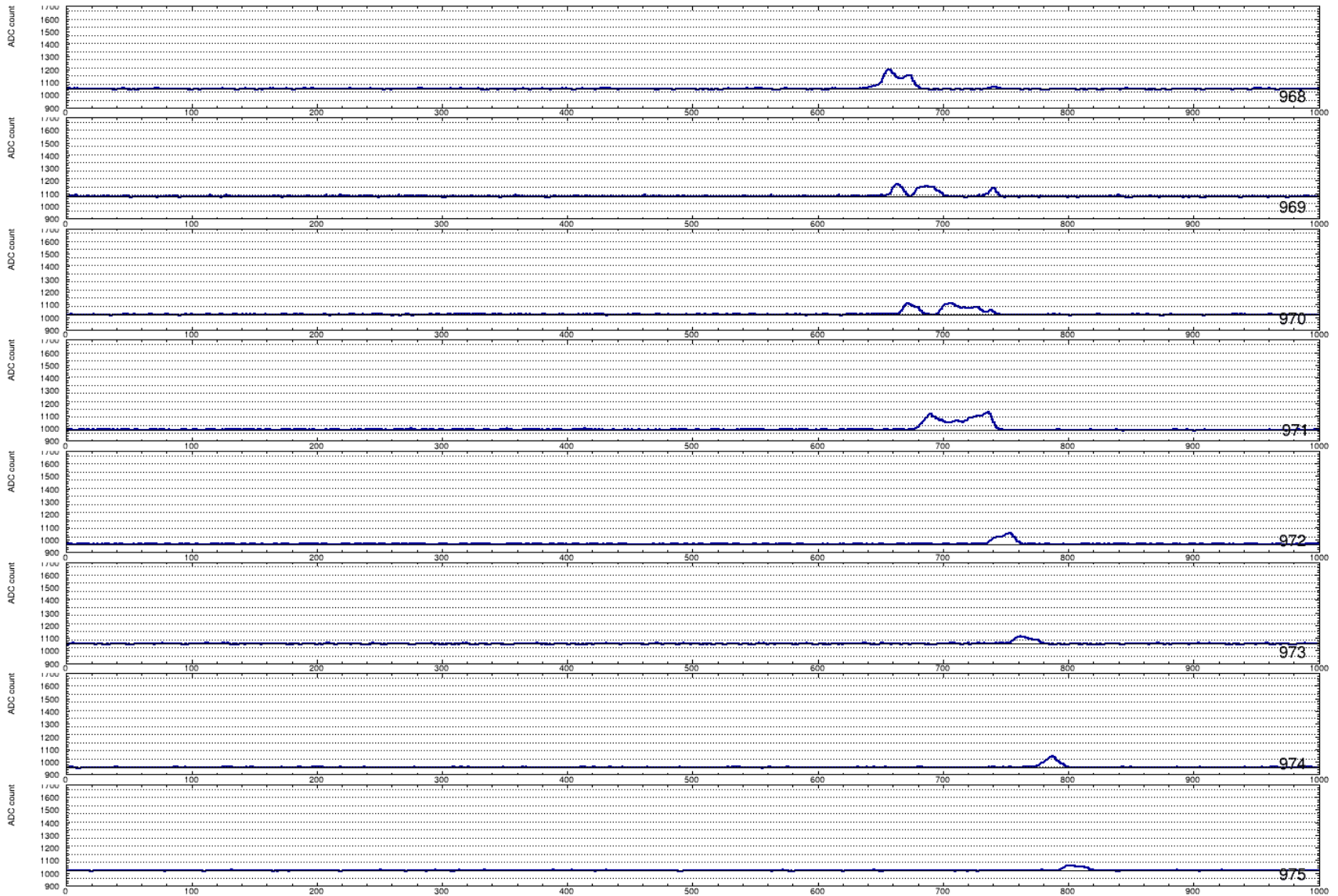
# 5005-5 z1

ADC raw run 5005 event 5



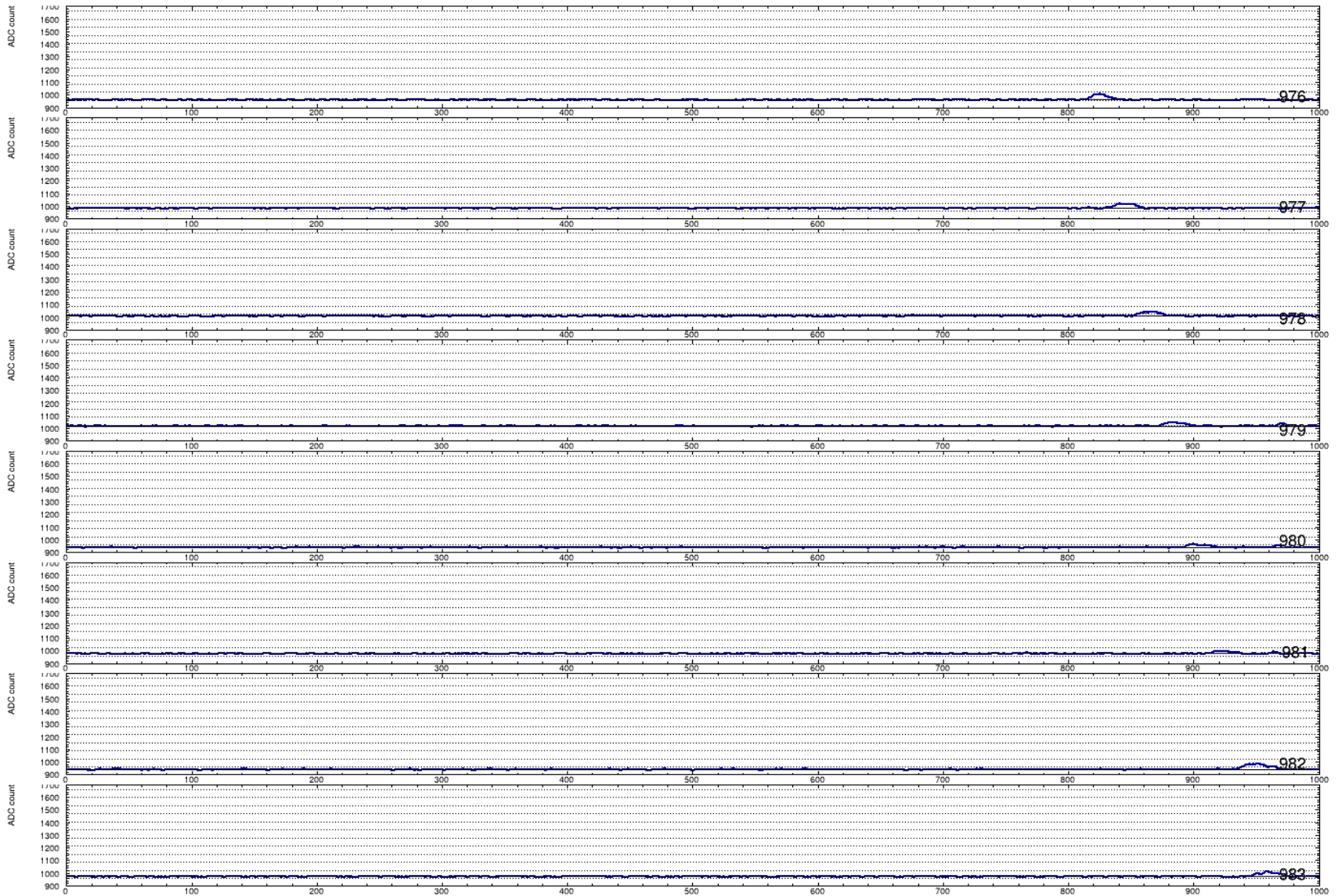
# 5005-5 z1

ADC raw run 5005 event 5



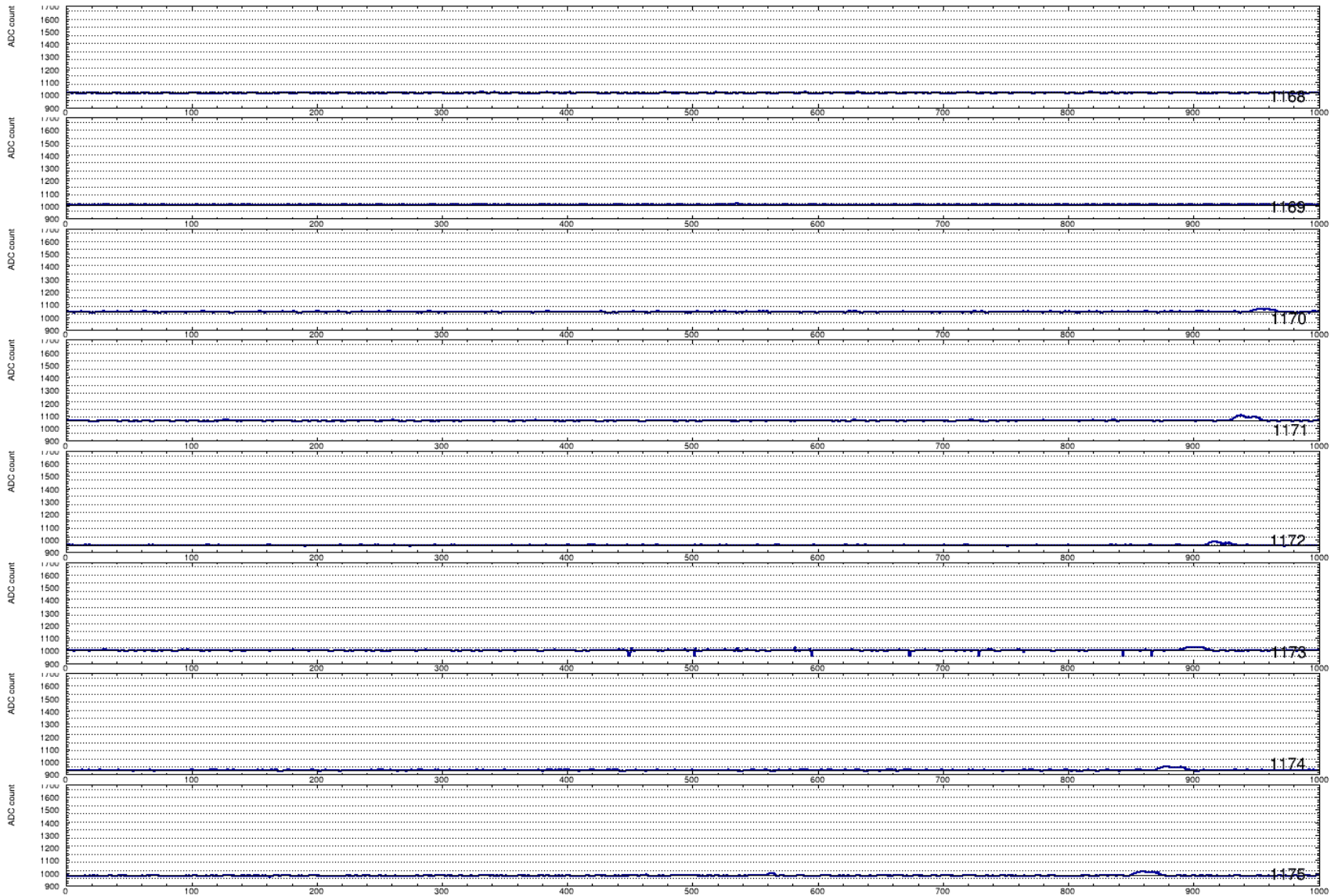
# 5005-5 z1

ADC raw run 5005 event 5



# 5005-5 z2

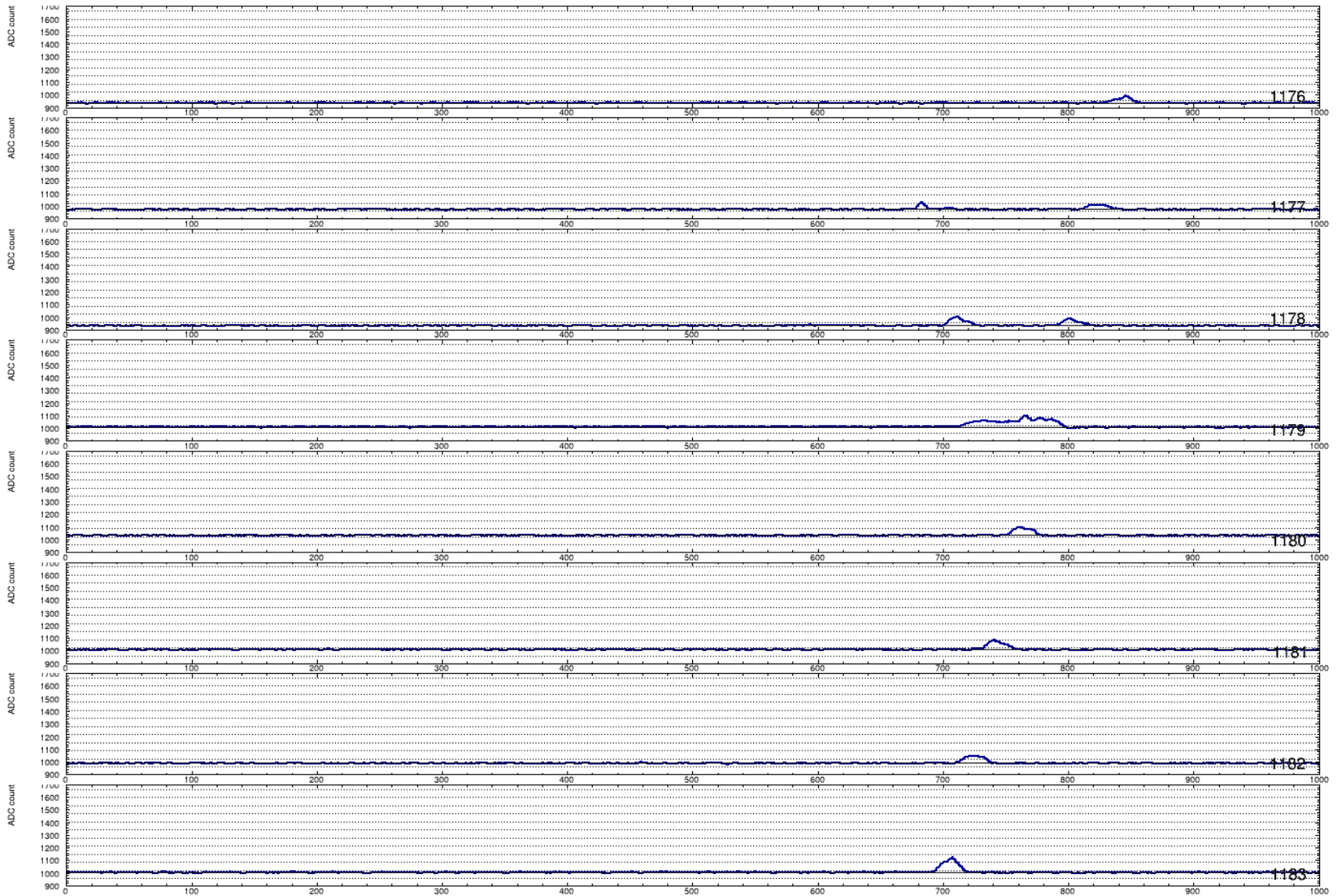
ADC raw run 5005 event 5





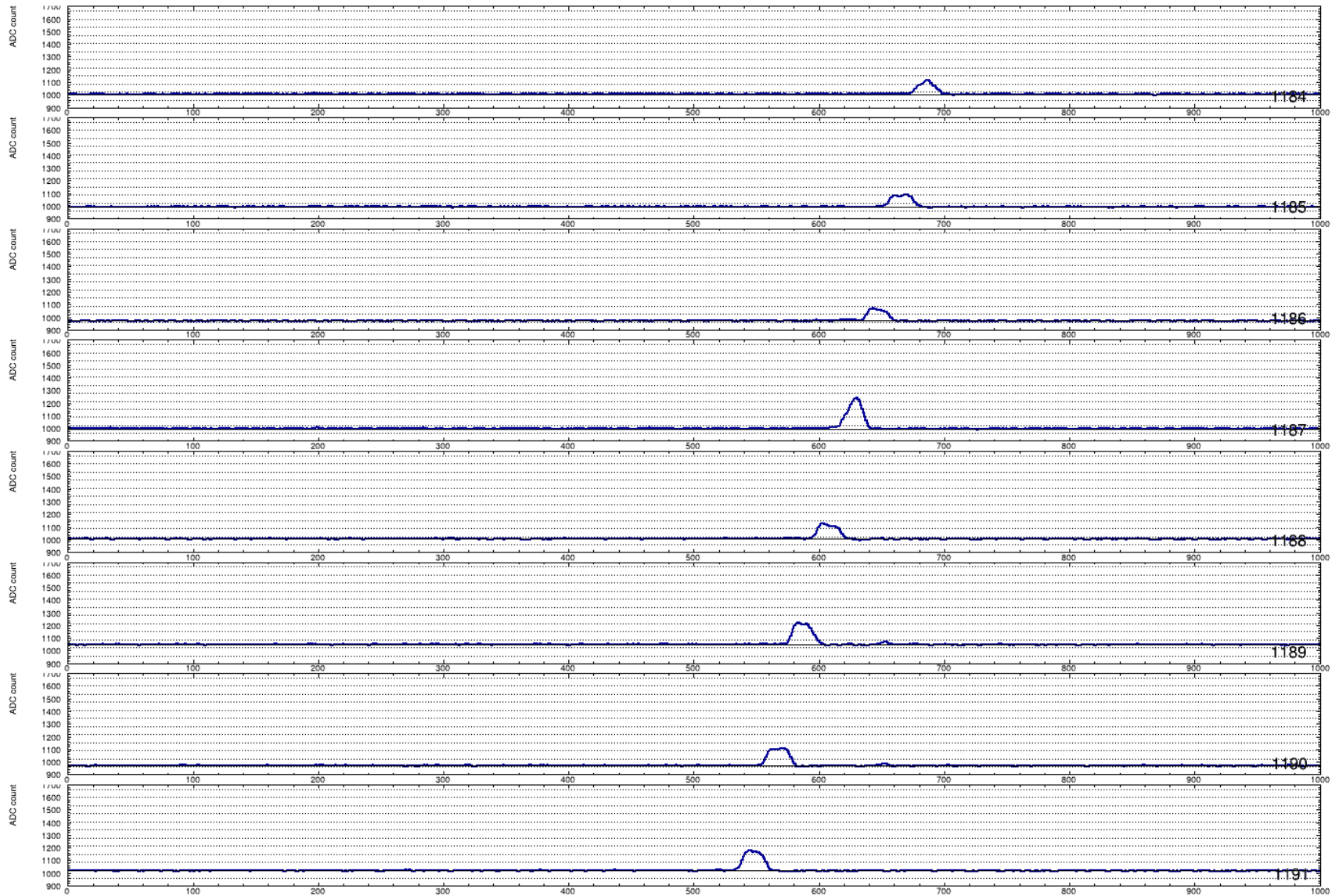
# 5005-5 z2

ADC raw run 5005 event 5



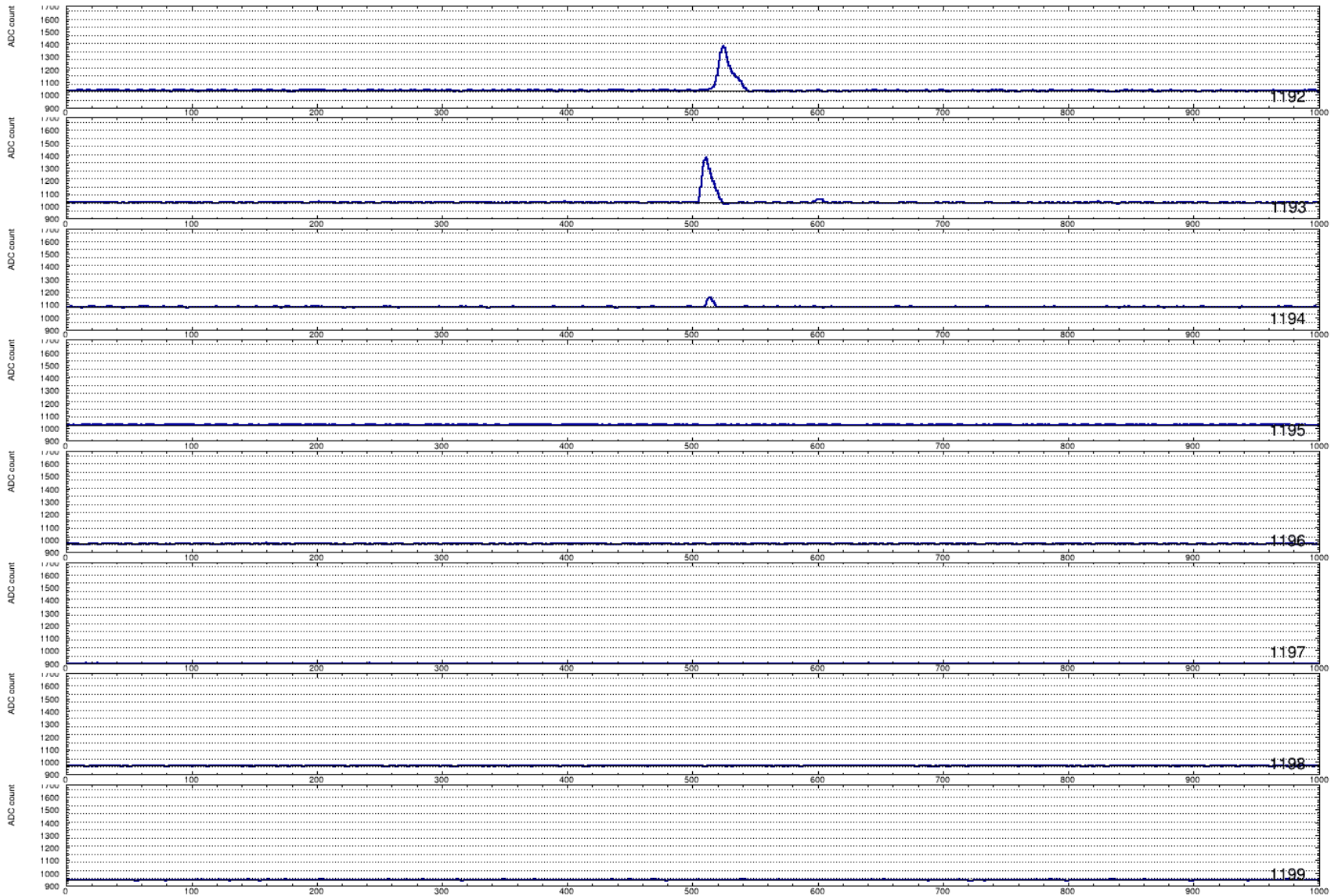
# 5005-5 z2

ADC raw run 5005 event 5



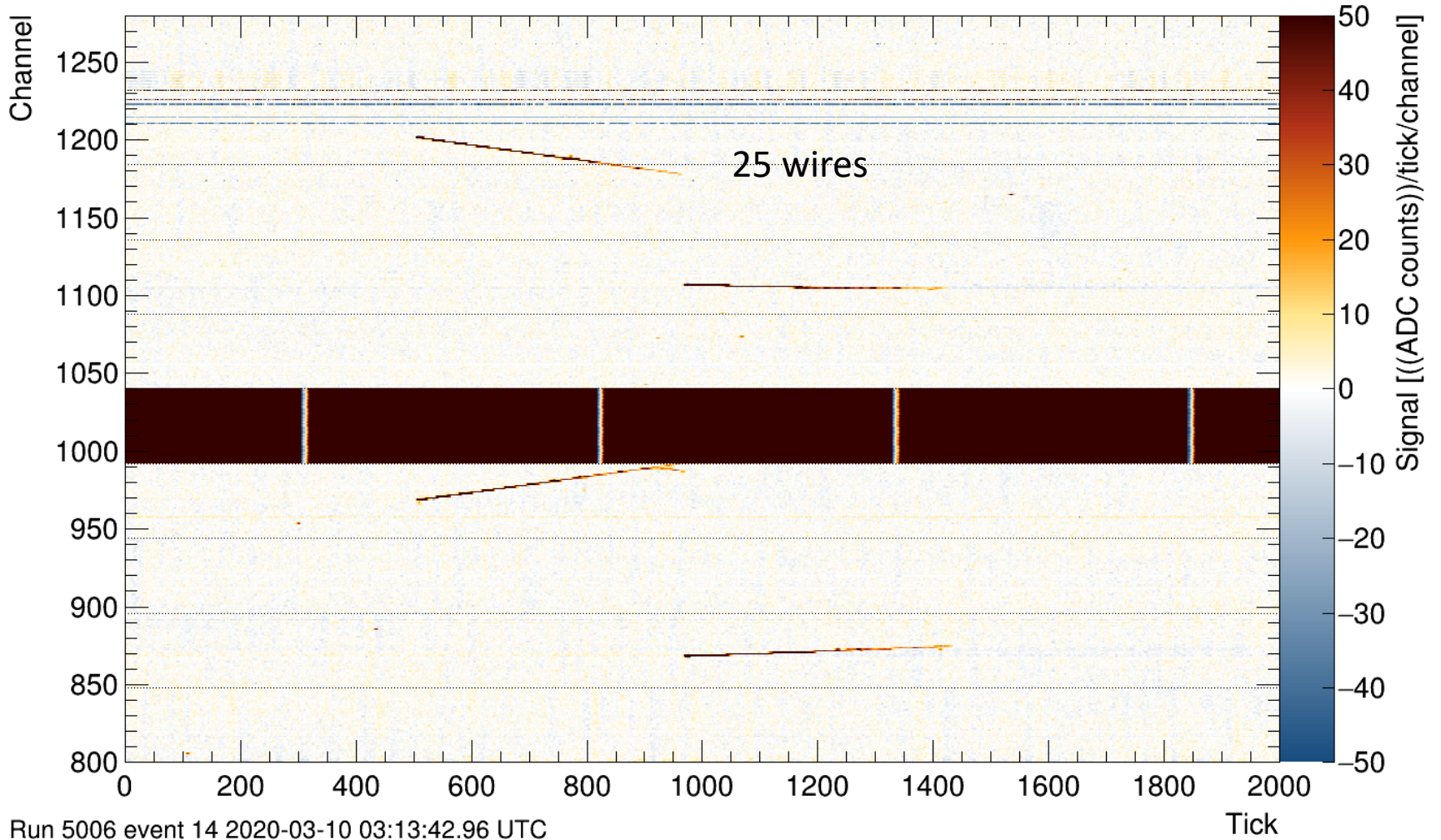
# 5005-5 z2

ADC raw run 5005 event 5



# 5006-14

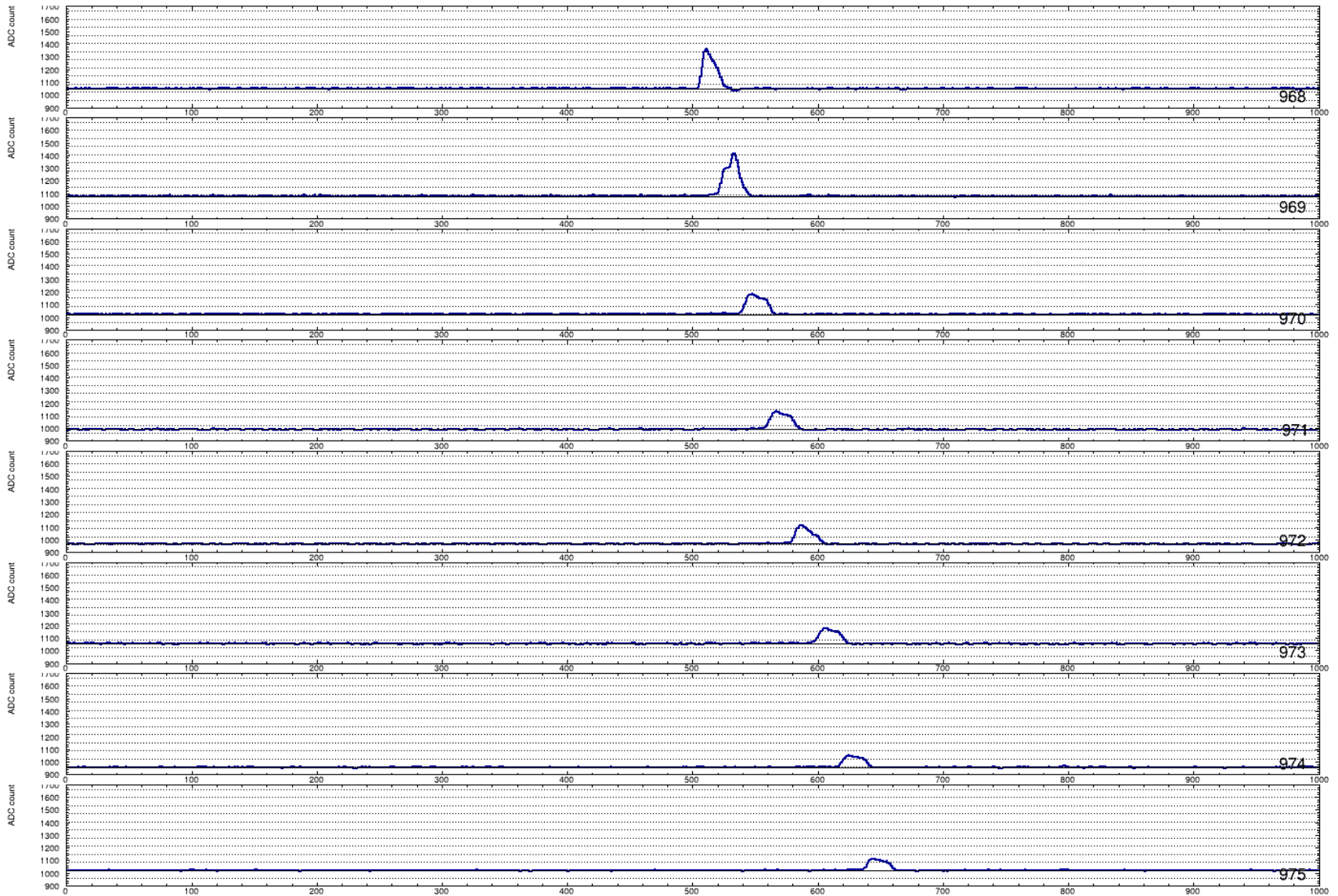
Raw ADC for Iceberg collection planes



Run 5006 event 14 2020-03-10 03:13:42.96 UTC

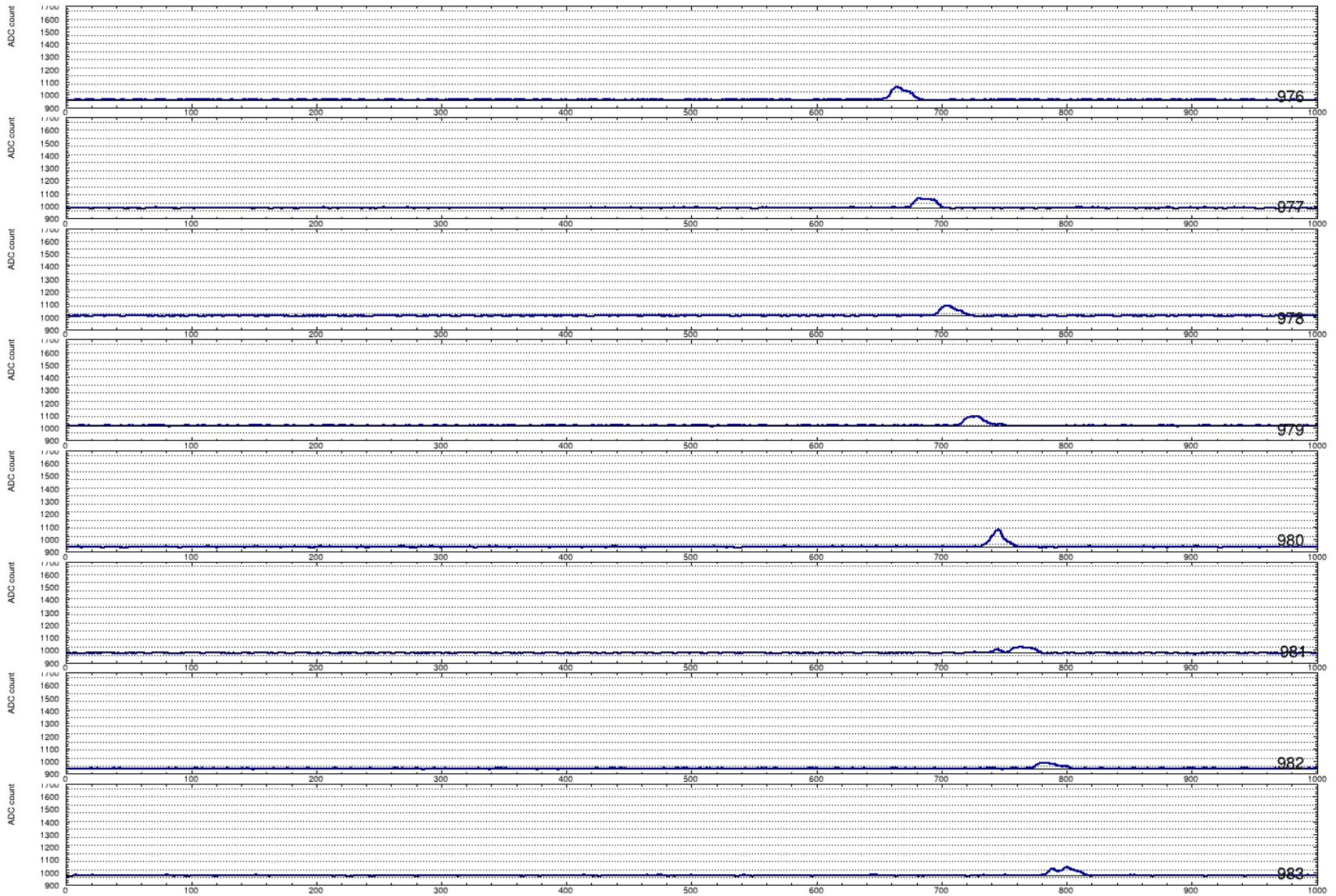
# 5006-14 z1

ADC raw run 5006 event 14



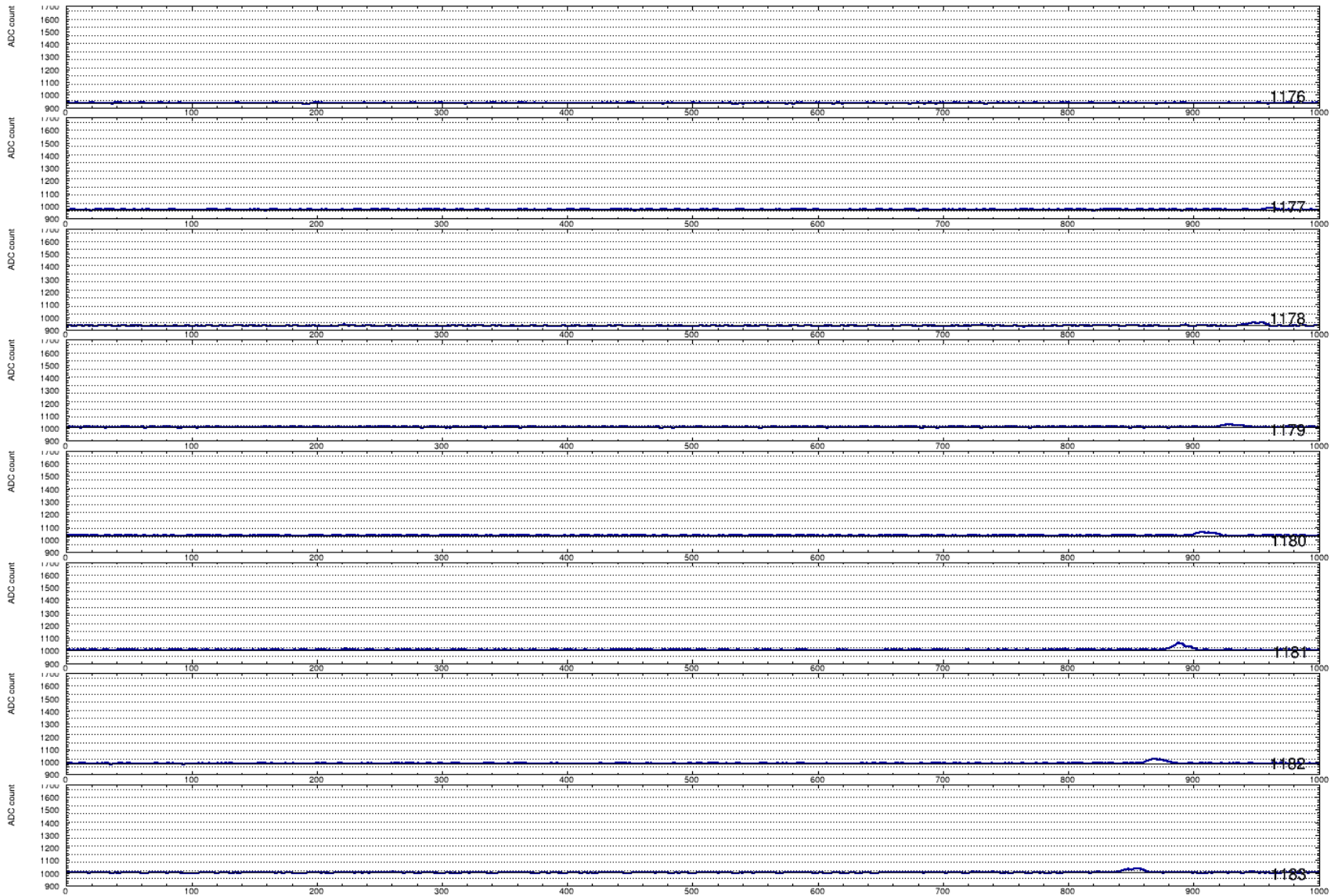
# 5006-14 z1

ADC raw run 5006 event 14



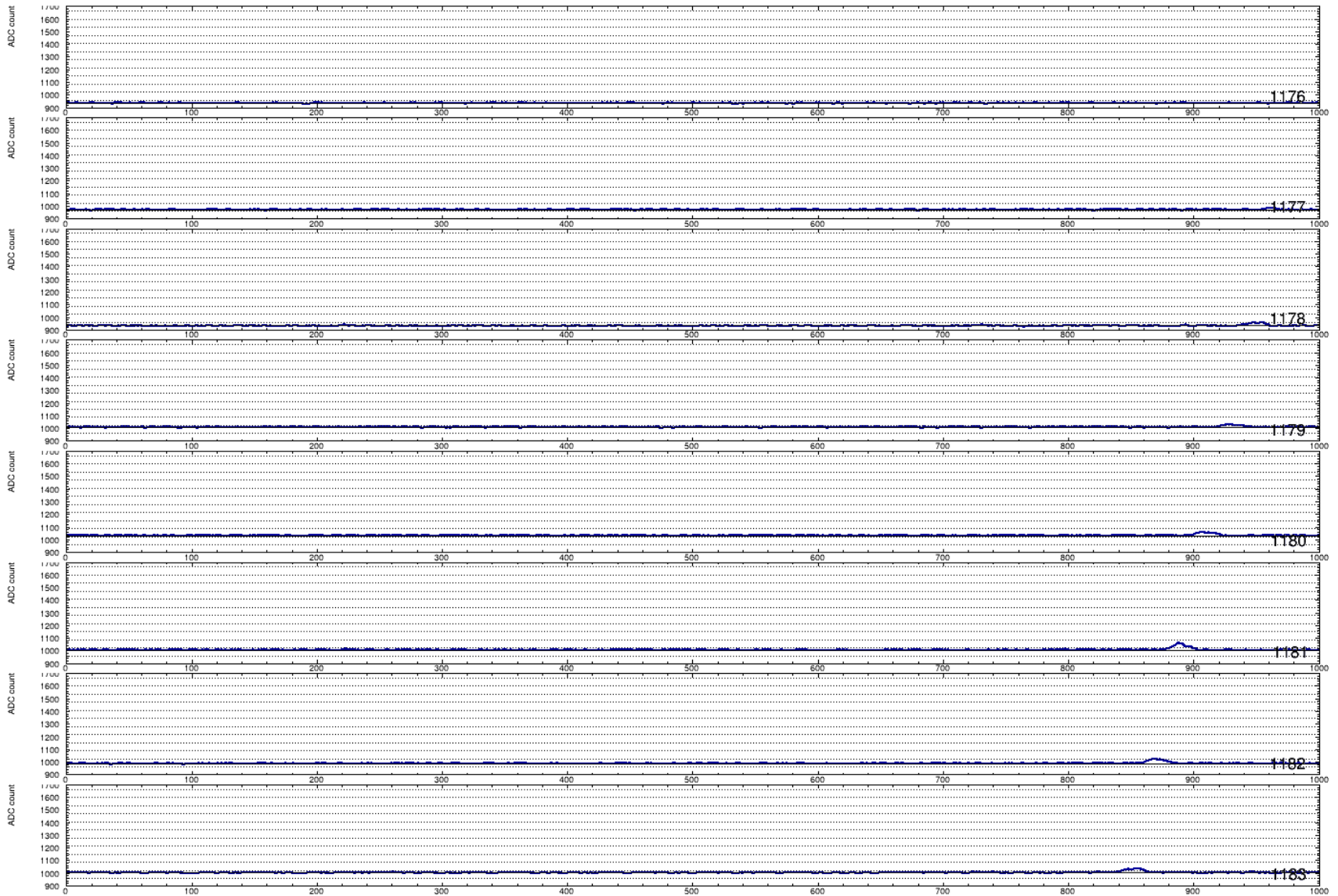
# 5006-14 z1

ADC raw run 5006 event 14



# 5006-14 z2

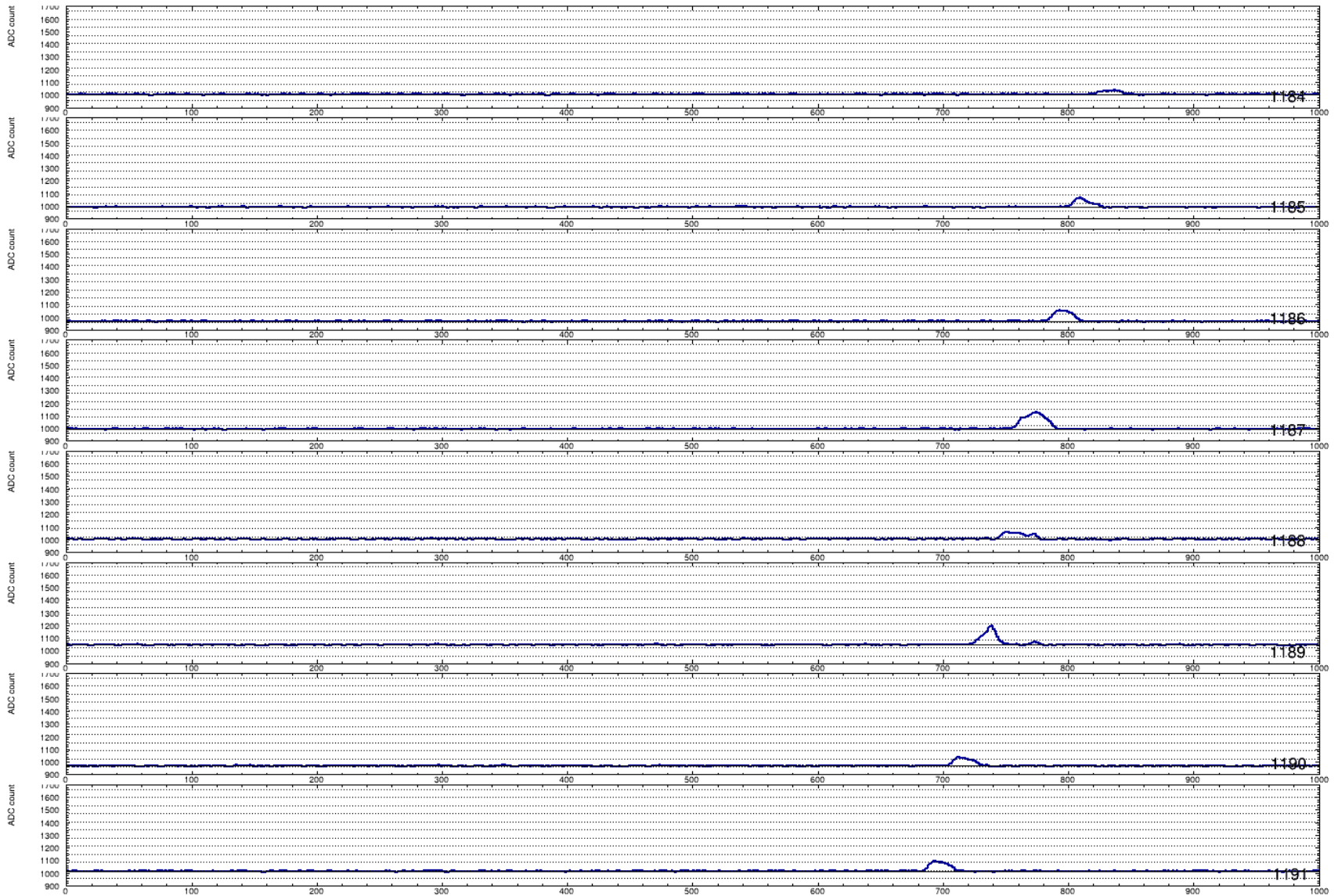
ADC raw run 5006 event 14





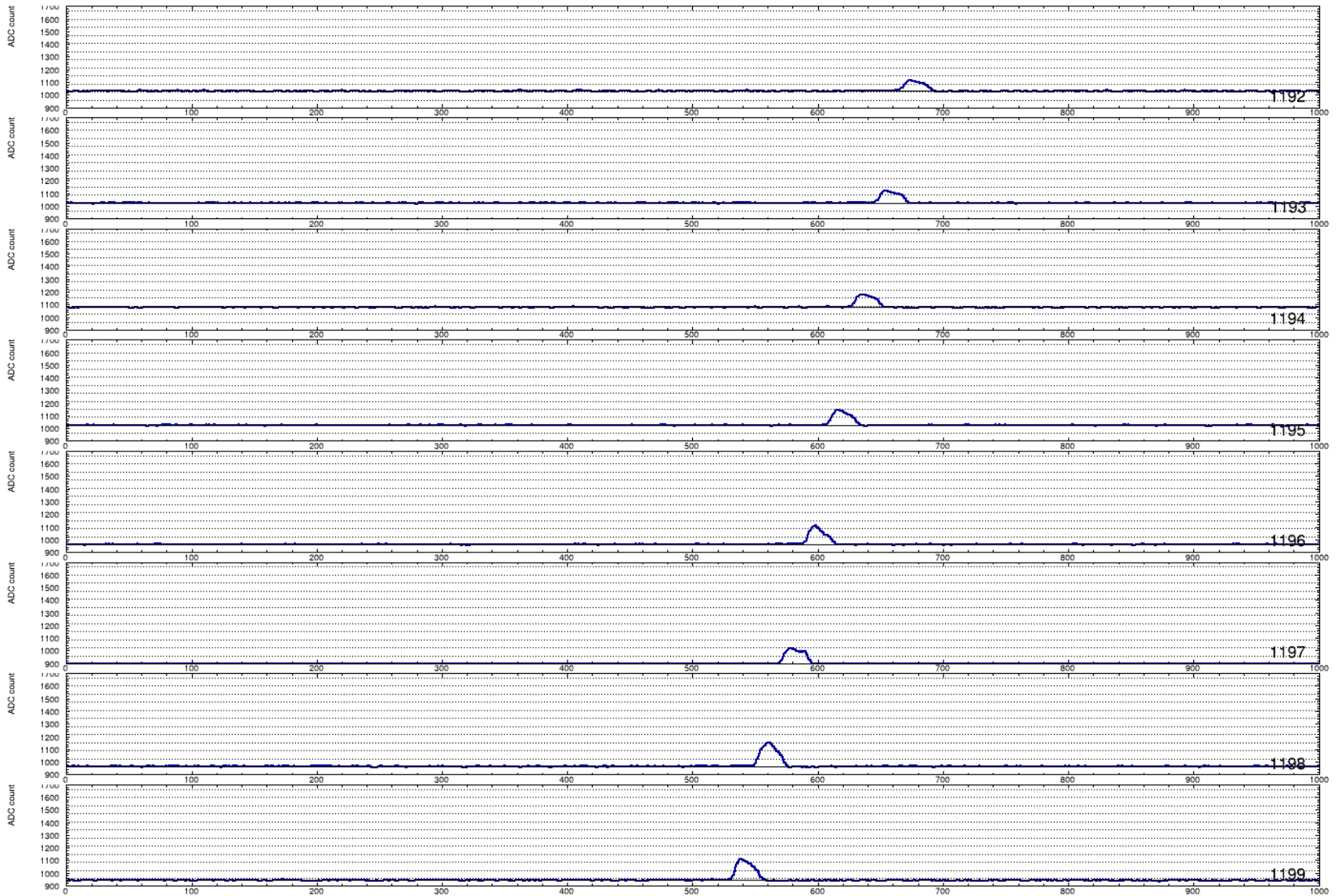
# 5006-14 z2

ADC raw run 5006 event 14



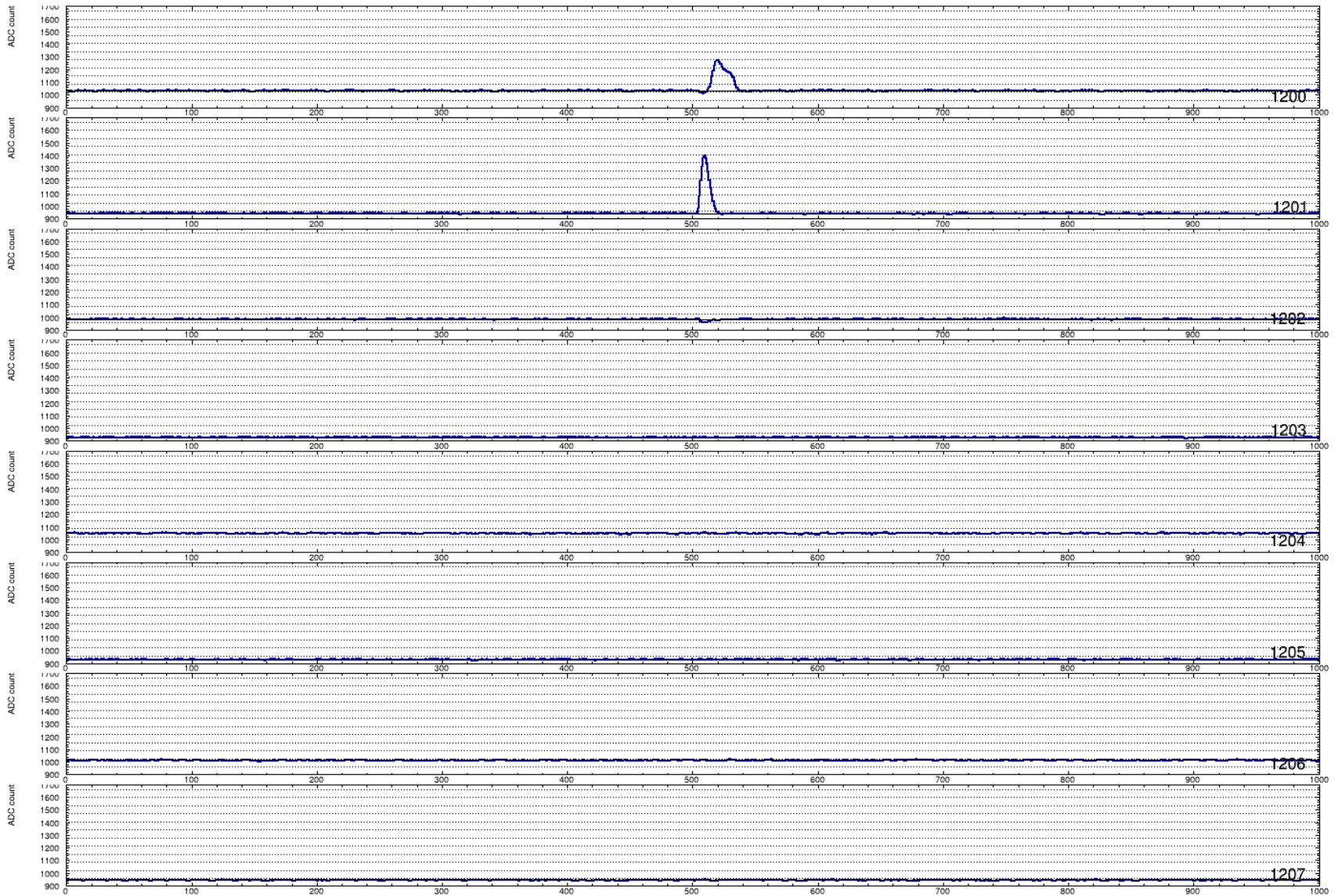
# 5006-14 z2

ADC raw run 5006 event 14



# 5006-14 z2

ADC raw run 5006 event 14



# Conclusions

## Summary noise levels presented

- Sample and integrated; before/after CNR, collection/induction
- And for various run periods
- Sample noise  $\lesssim 58$  e (ENC 320 e)
- Integrated noise before  $\rightarrow$  after CNR
  - Collection: 620  $\rightarrow$  590 e
  - Induction 790  $\rightarrow$  690 e
  - Both are medians for “very quiet” run period

## Calibration scale check

- Signals from perpendicular tracks have the expected ADC signal height  $\rightarrow$  our calibration scale is roughly correct
- Height SNR for perpendicular tracks is around 85
  - Expect about the same for parallel tracks
- Area SNR for parallel tracks is  $(30 \text{ ke}) / (600 \text{ e}) = 50$
- All before degradation from lifetime

# Next

## What might I/we do next?

- ~~Charge calibration~~
- ~~Identify sticky codes~~
- Tune reconstruction sequence
  - ◉ ~~Use the above (calibrate, mitigate SC)~~
  - ◉ ~~Check tail removal (fix if needed)~~
    - Timing mitigation?
  - ◉ ~~Check noise removal~~
    - Deconvolution
- Noise studies
  - ◉ ~~Sample and integrated estimates before deconvolution~~
  - ◉ ~~DFTs for each view~~
  - ◉ ~~Same combining many runs~~
  - ◉ ~~Same after removing noisy channels~~
    - Same after deconvolution
    - Remove event channels with overflows

# Extras

Track angles from Wanwei

Trigger counter geometry (Shekhar)

Analysis code packages

Still more noise summary groups

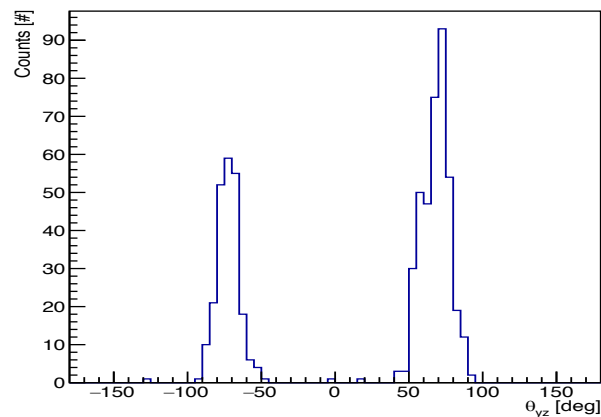
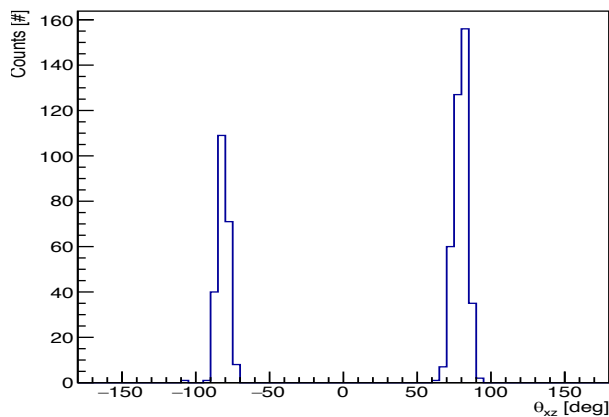
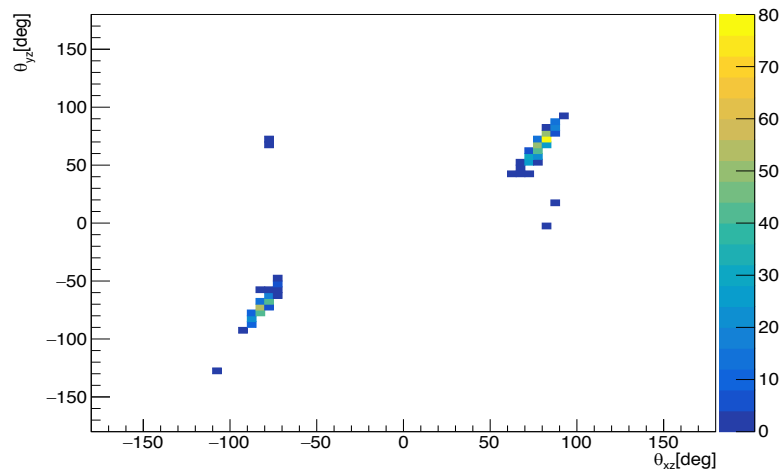
Analysis code

Run groups

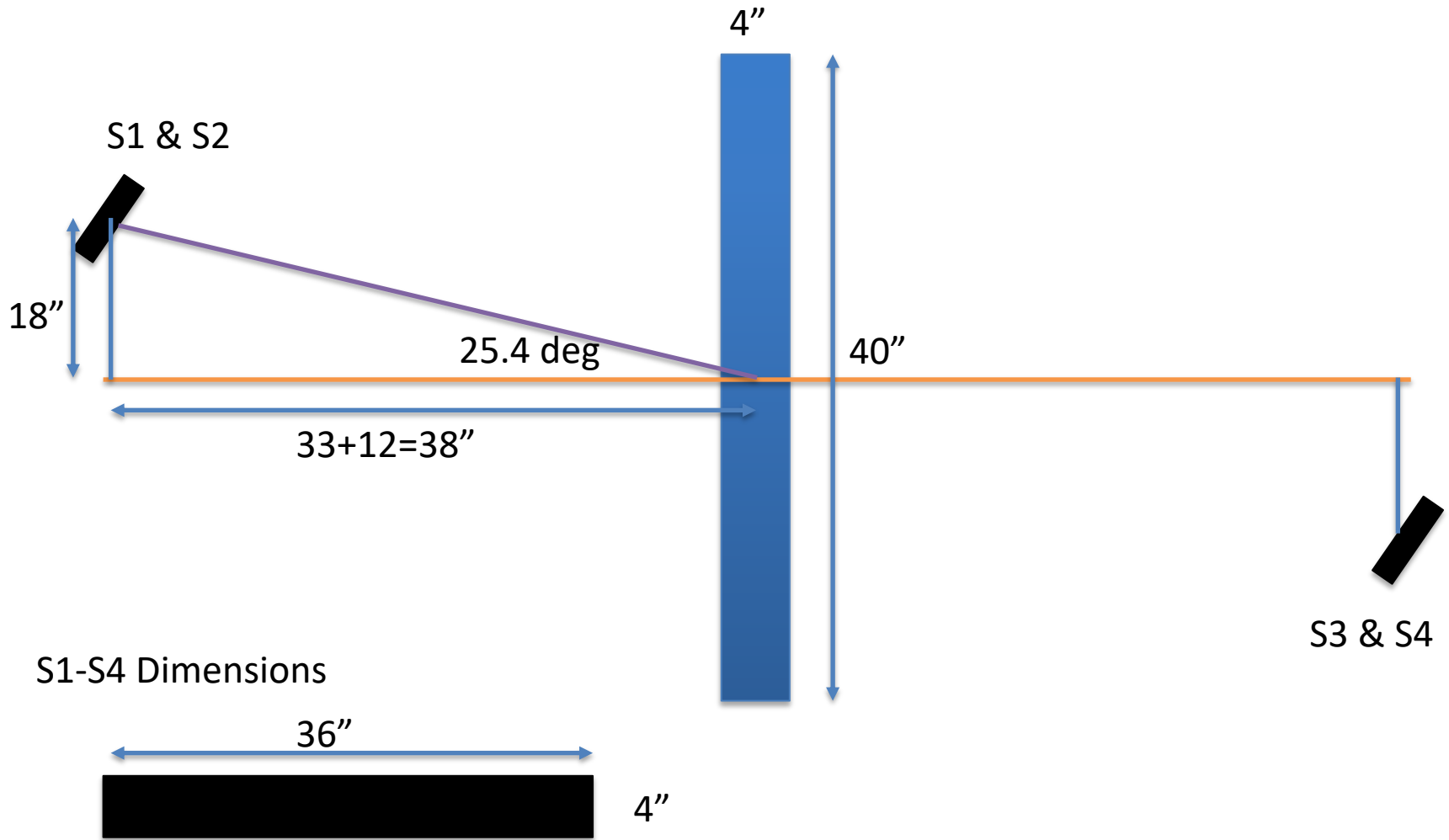
# Track angles from Wanwei

## Today's Topics—Track Angle

- Most of tracks are close to the drift direction
- Many of tracks are also close to wire direction

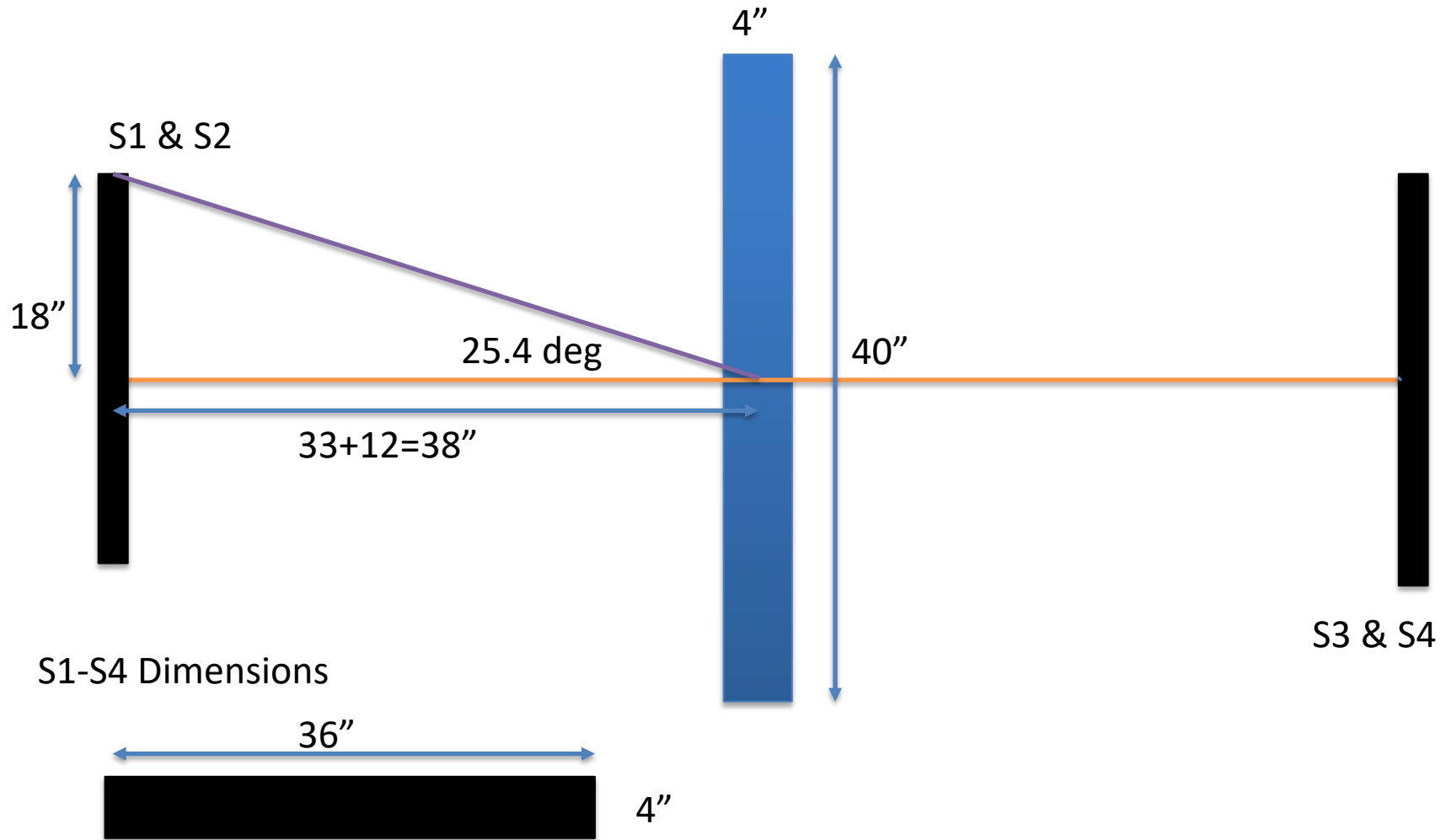


# CRT Vertical Angle

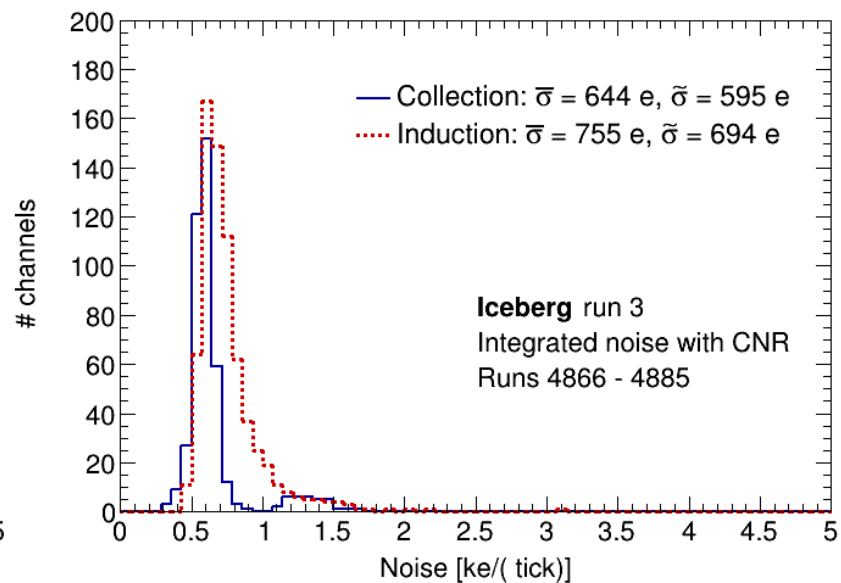
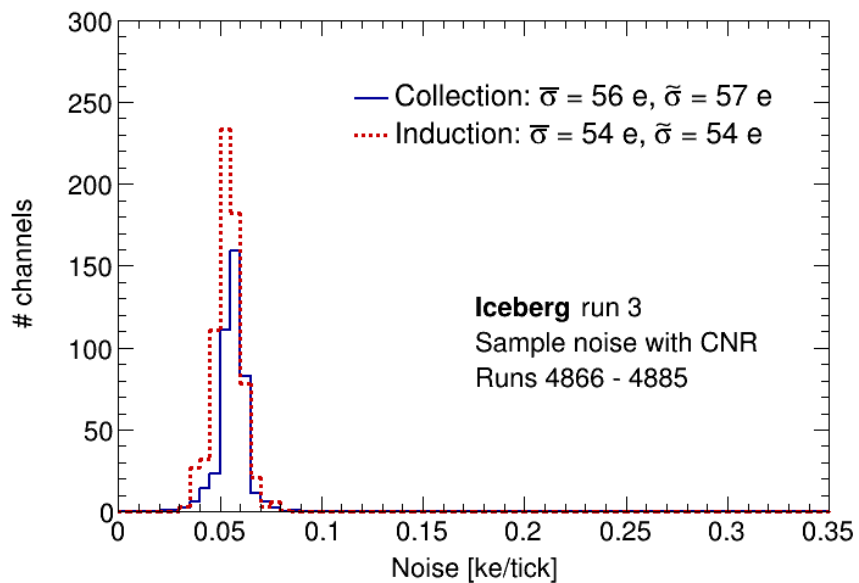
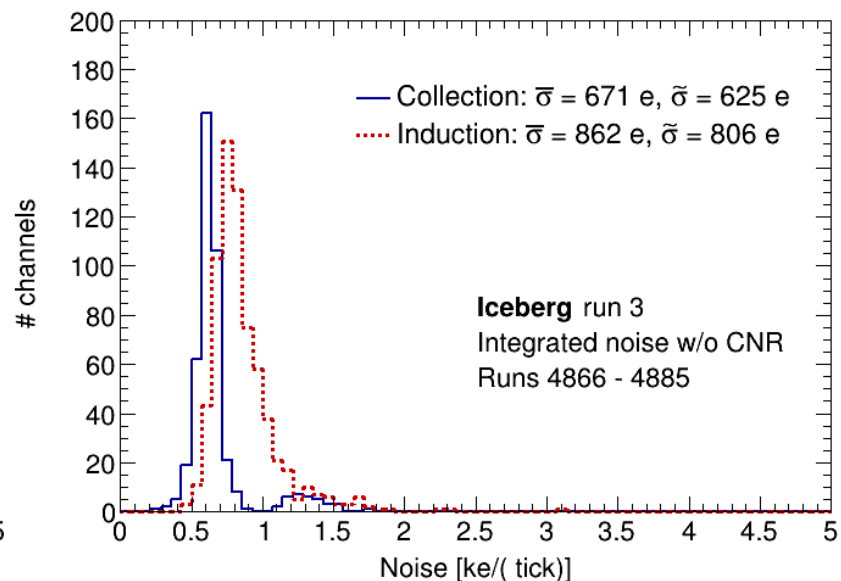
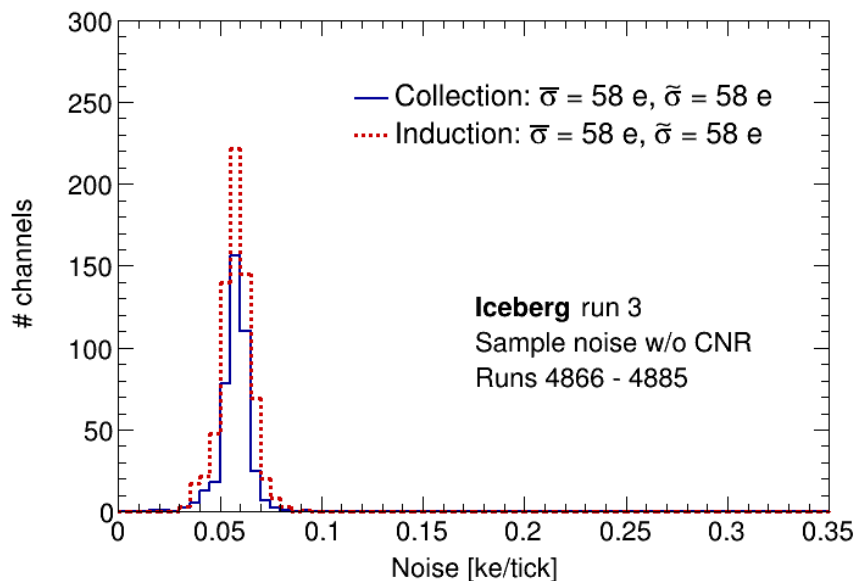




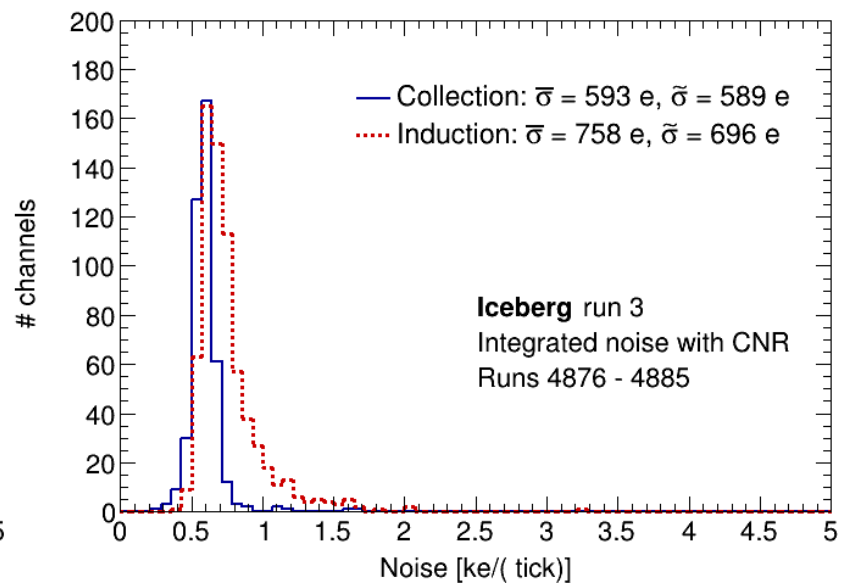
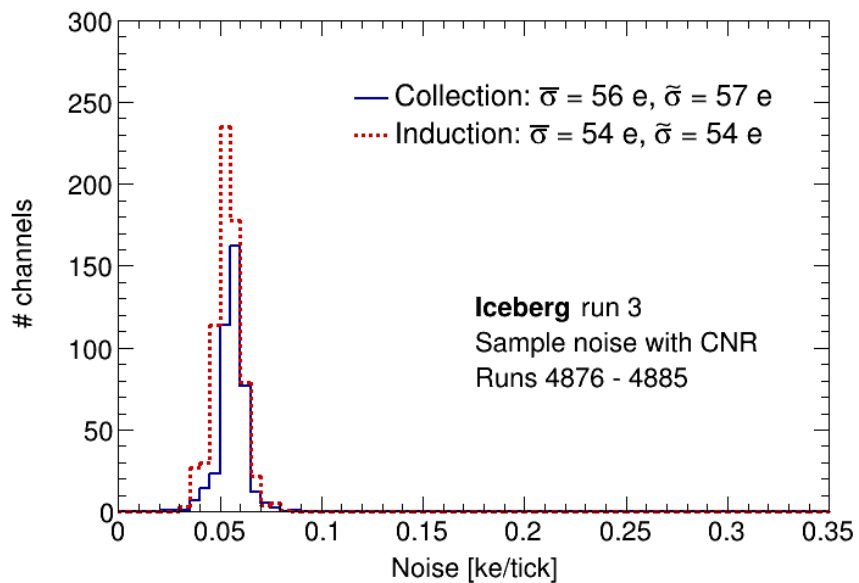
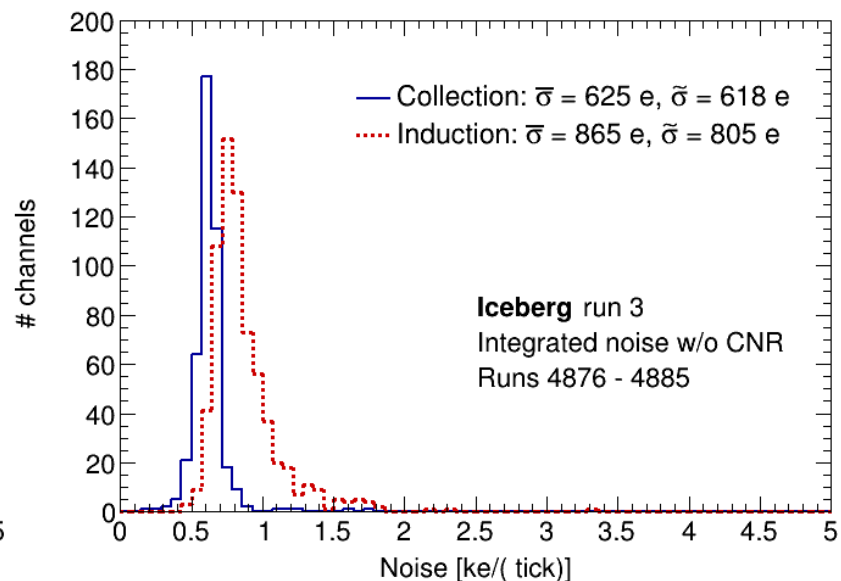
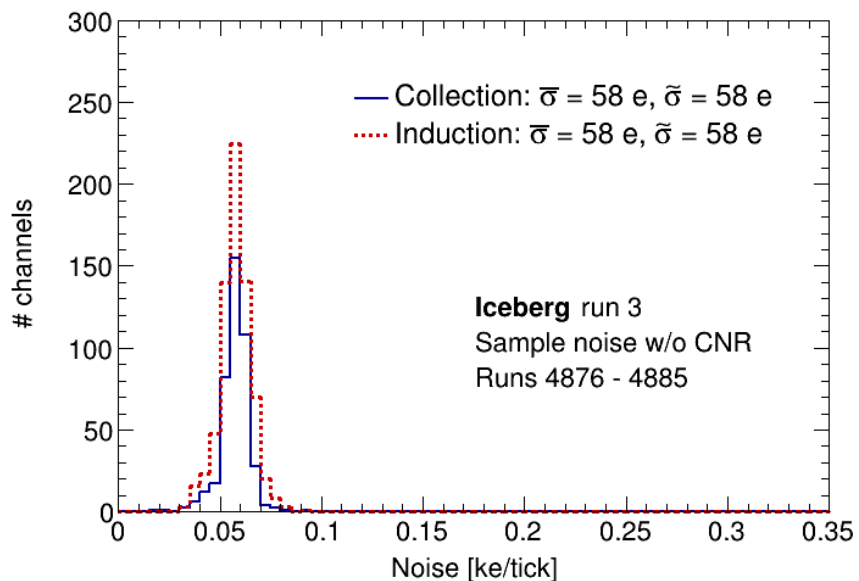
# CRT Horizontal Angle



# Quiet 2 (4866 – 4885)



# Quiet 3 (4876 – 4885)



# Analysis code

I have been migrating my analysis scripts into github

- A series of small (and so quick to build) cmake projects
- For my use but available to all
- Happy to get feedback and have others create similar

## Current projects

- myproj is a toy example showing the conventions:
  - Script “build” to build and install the project
  - Cmake examples to build utility libraries, art services and root dictionaries, and to install headers, fcl and python
  - <https://github.com/dladams/myproj/blob/master/README.md>
- duneproc is my base level project
  - Script duneproc is a wrapper for running lar jobs
  - Fcl file for protoDUNE and (new) Iceberg
  - <https://github.com/dladams/duneproc/blob/master/README.md>
    - Note link to Iceberg tutorial
- dunececalib has charge calibration code
- dunenoise has code for noise studies and DFT

# Run groups

Iceberg Run 3 run grouping

David Adams

May 5, 2020

Run range Configuration

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3366 3382 Initial CRT (cosmic trigger) configuration. 15 deg?  
3384 3466 Final CRT. 15 deg?  
3467 3611 Cosmics with CRT at 15 deg  
3616 3626 TPC off  
3627 3695 Cosmics with CRT at 15 deg?  
3696 3839 Cosmics with CRT at 15 deg  
3870 3874 TPC off  
3875 4430 Cosmics with CRT at 15 deg?  
4431 4480 Cosmics with CRT at 15 deg  
4481 4507 Pulser runs  
4508 4530 Cosmics with CRT at 15 deg  
4531 4658 TPC off  
4679 4698 Pulser runs  
4699 4699 Pulser DAC=0 (cosmic0, random trigger?)  
4700 4703 Pulser runs (pulser)  
4704 4705 Pulser DAC=0 (pulser0)  
4710 4711 Pulser runs  
4714 4717 Pulser DAC=0 (pulser0)

Begin stable running

4718 4860 New 60 deg CRT (cosmic trigger) with S4  
4861 5050 New 60 deg CRT w/o S4  
5060 5108 Investigation of timing fragment  
5109 5112 Random trigger with  $g=4.7$ .  
5113 5132 Pulser with  $g=4.7$   
5133 5153 Pulser with  $g=7.8$   
5154 5174 Pulser with  $g=25$   
5175 5195 Pulser with  $s=0.5$   
5196 5222 Pulser with  $s=1.0$   
5223 5243 Pulser with  $s=3.0$   
5244 5264 Pulser with  $b=200$  mV  
5265 5285 Pulser with  $b=200/900$  mV  
5286 5306 Pulser with  $s=0.5$   
5307 5327 Pulser with  $I=1000$  pA  
5328-5348 Pulser with AC coupling  
5349-5359 Cosmics vary settings run to run

4504, 4699, 5109 taken without cosmic trigger and TPC on.