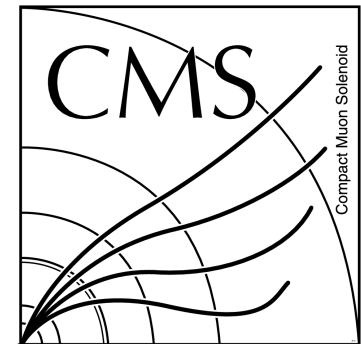


S-channel @ 100 TeV

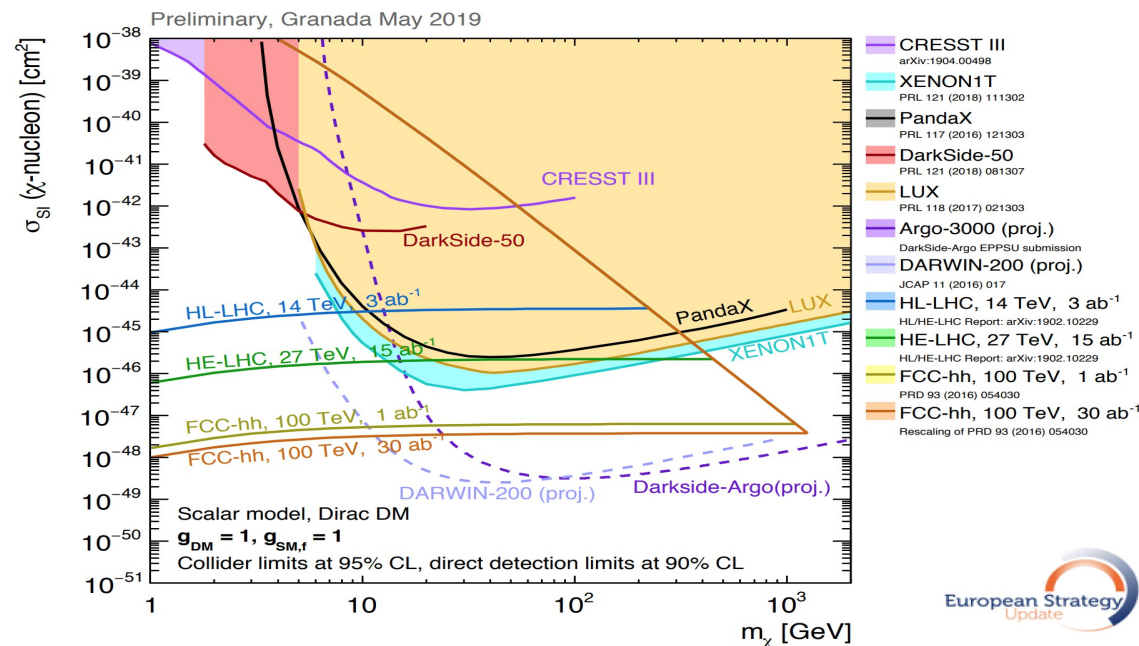


Phil Harris (MIT)



Point of this talk

- S-channel DM search discussion
 - An important consideration for future projections
 - Has been featured in various documents
 - Strategy has been to develop current analyses for :
 - 100 TeV pp collider with 30 ab⁻¹
 - 14 TeV pp collider with 3ab⁻¹



This talk :
**Review analysis
 strategy**

An Analysis for all ages

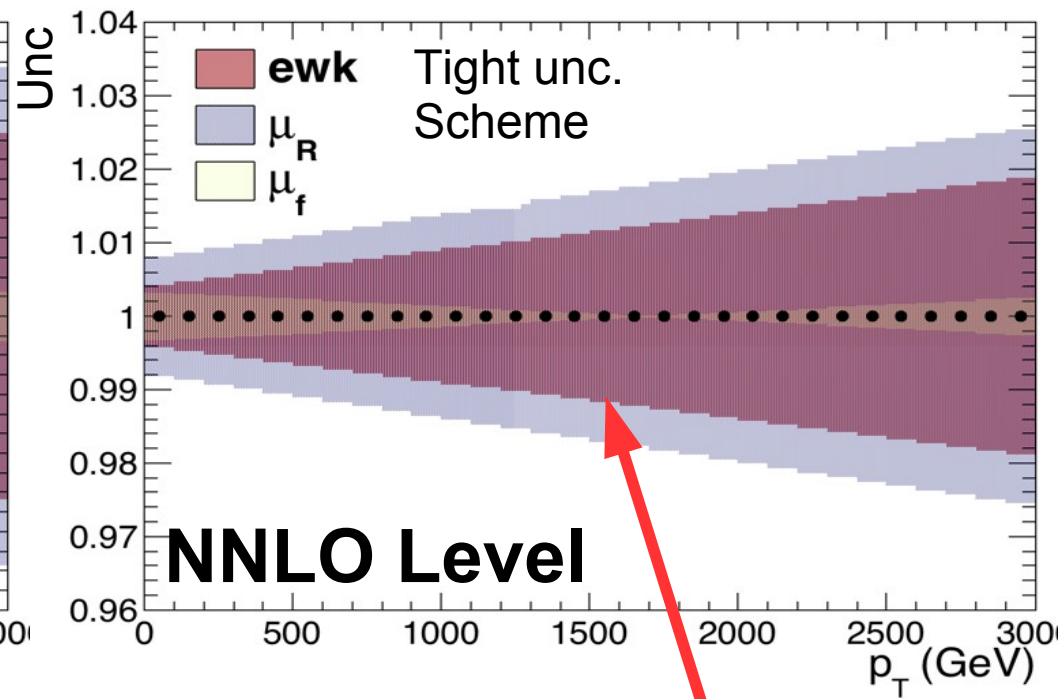
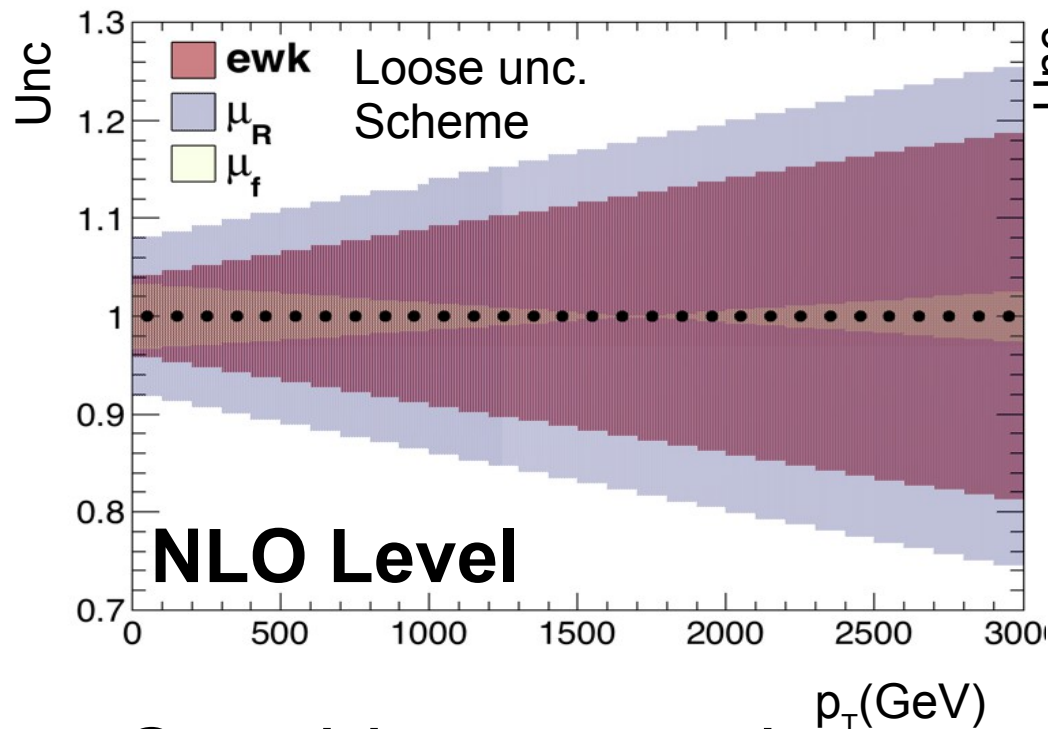
- In light of the LHC developments in monojet:
 - Big simultaneous fits
 - Improved theoretical constraints on EWK-corr
 - Fundamentally approach scales well with lumi
 - Goal should be to take advantage of this approach
- However 1st s-channel studies were done before
 - Studies were done before LHCDMWG recommends
 - Room for improvement
- Good news: mostly redone in Higgs to invisible

Experimental Approach in $H \rightarrow \text{Inv}$

- Use full simultaneous fit approach
- Delphes for simulation
 - In s-channel studies used toy smearing
- Weighted MC generation (makes things fast)
 - This was not done s-channel studies
- Same experimental setup otherwise as s-channel
 - Define control regions with leptons out to $|\eta| < 4.0$
 - Apply vetos based on this detector range
 - Approximate same lepton veto rates as LHC
 - Following CMS numbers (ATLAS is similar)
 - Skipped QCD background (its small in the end)

Uncertainties

- What are reasonable uncertainty choices



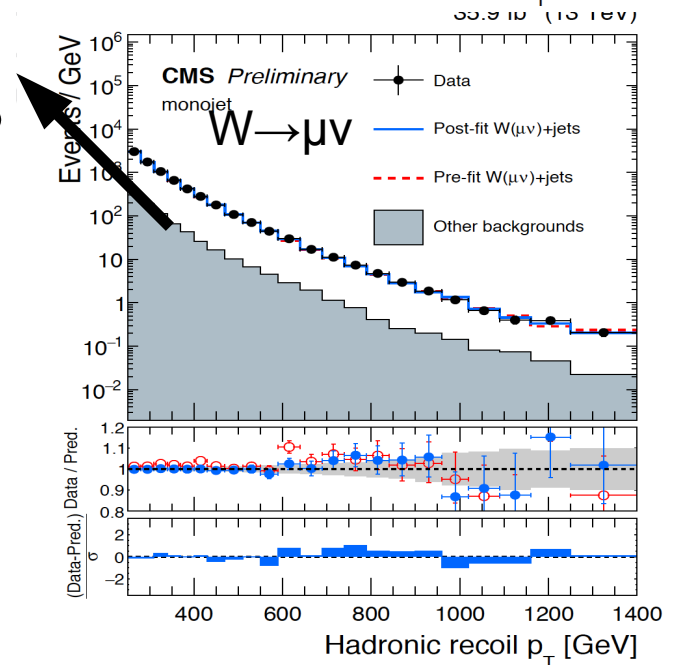
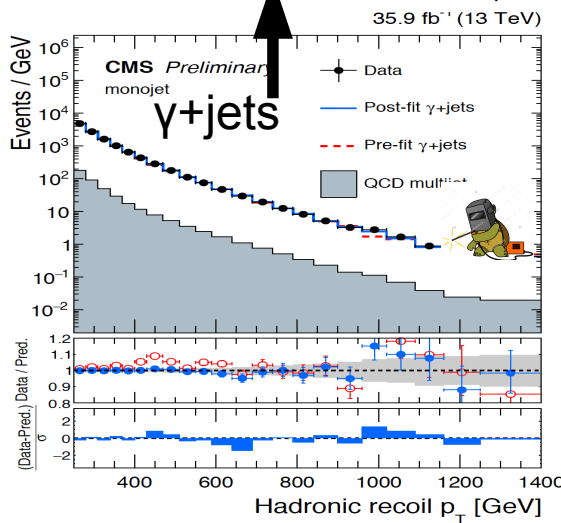
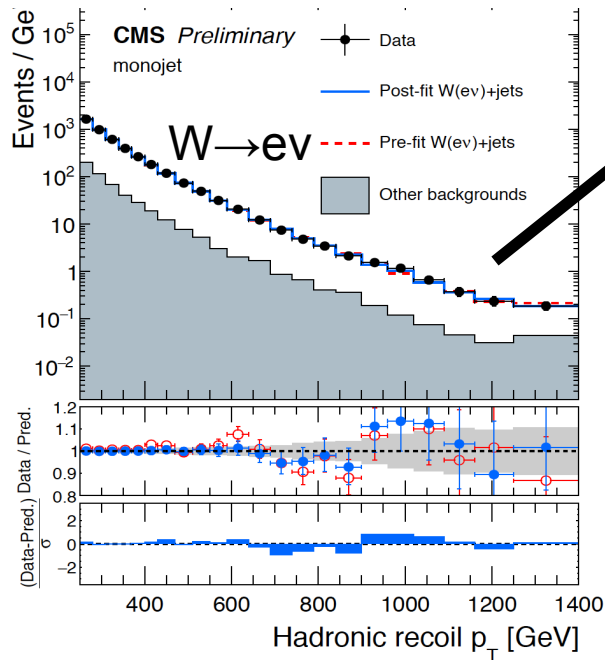
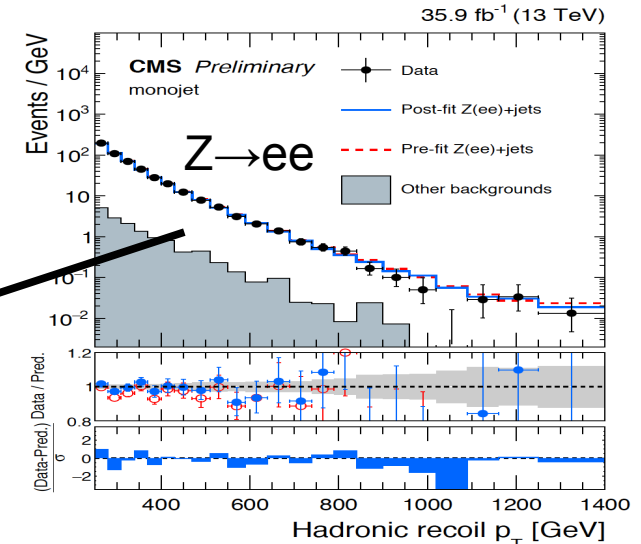
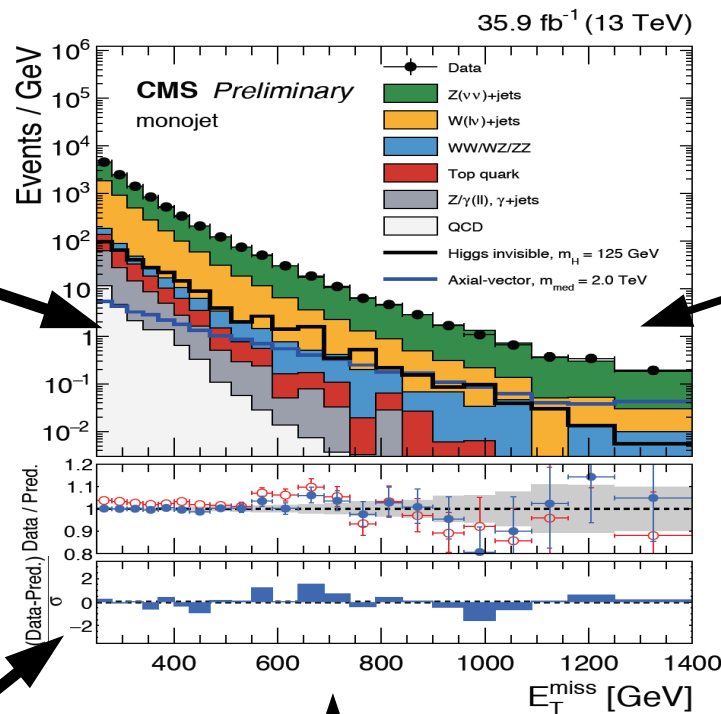
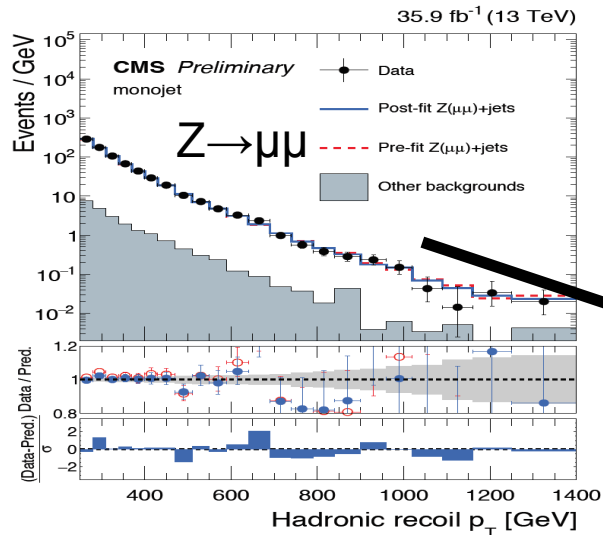
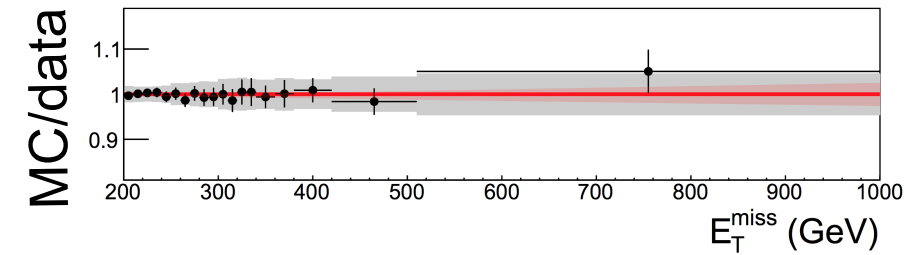
- Consider two options :

- A Loose uncertainty \rightarrow Comparable to NLO
- A Tight uncertainty \rightarrow Comparable to NNLO

definitively there

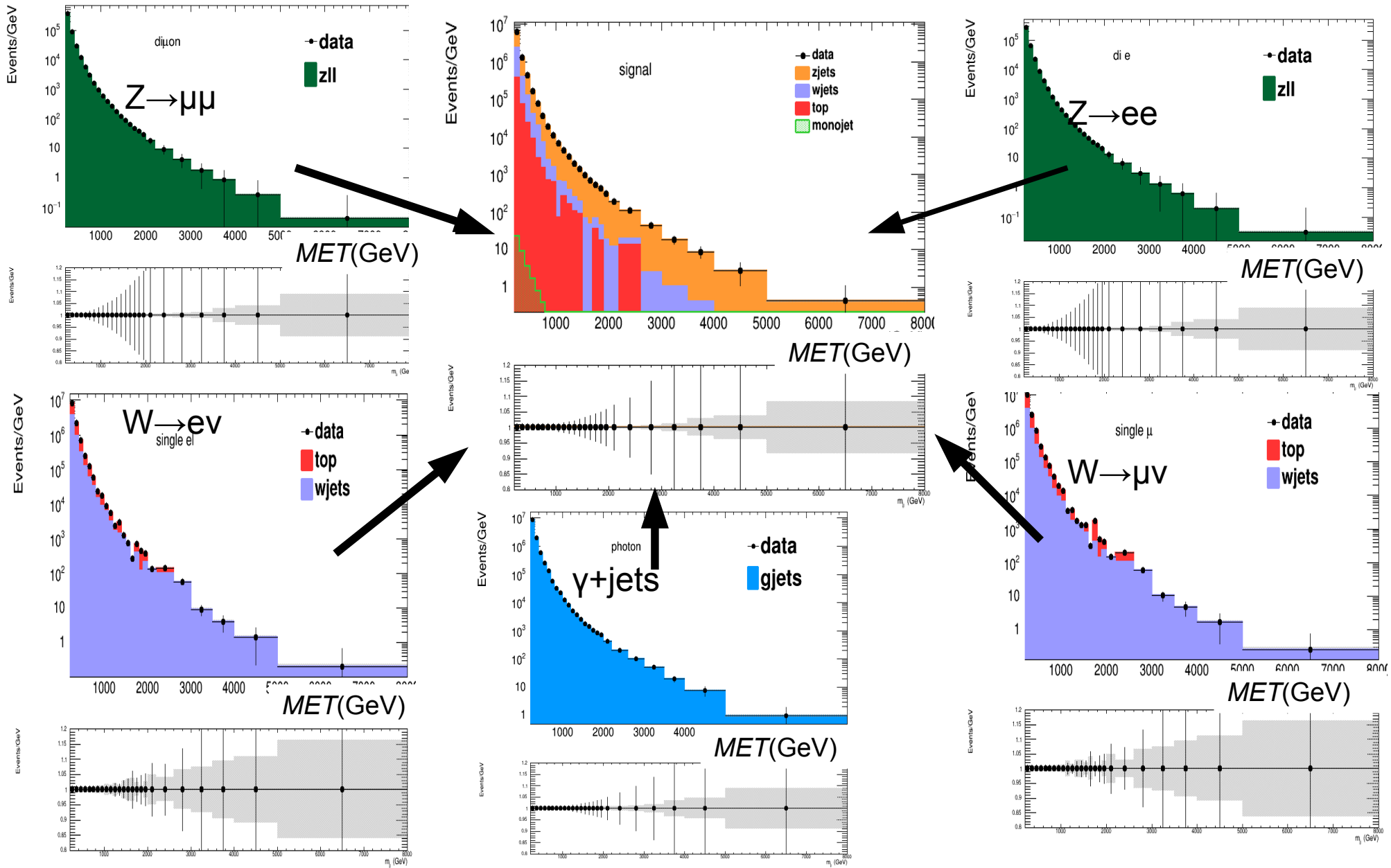
- Using : 0.5%/0.25%/5% e/ μ / τ efficiency & 1% lumi

5 Control regions 15% uncertainty @ 1 TeV



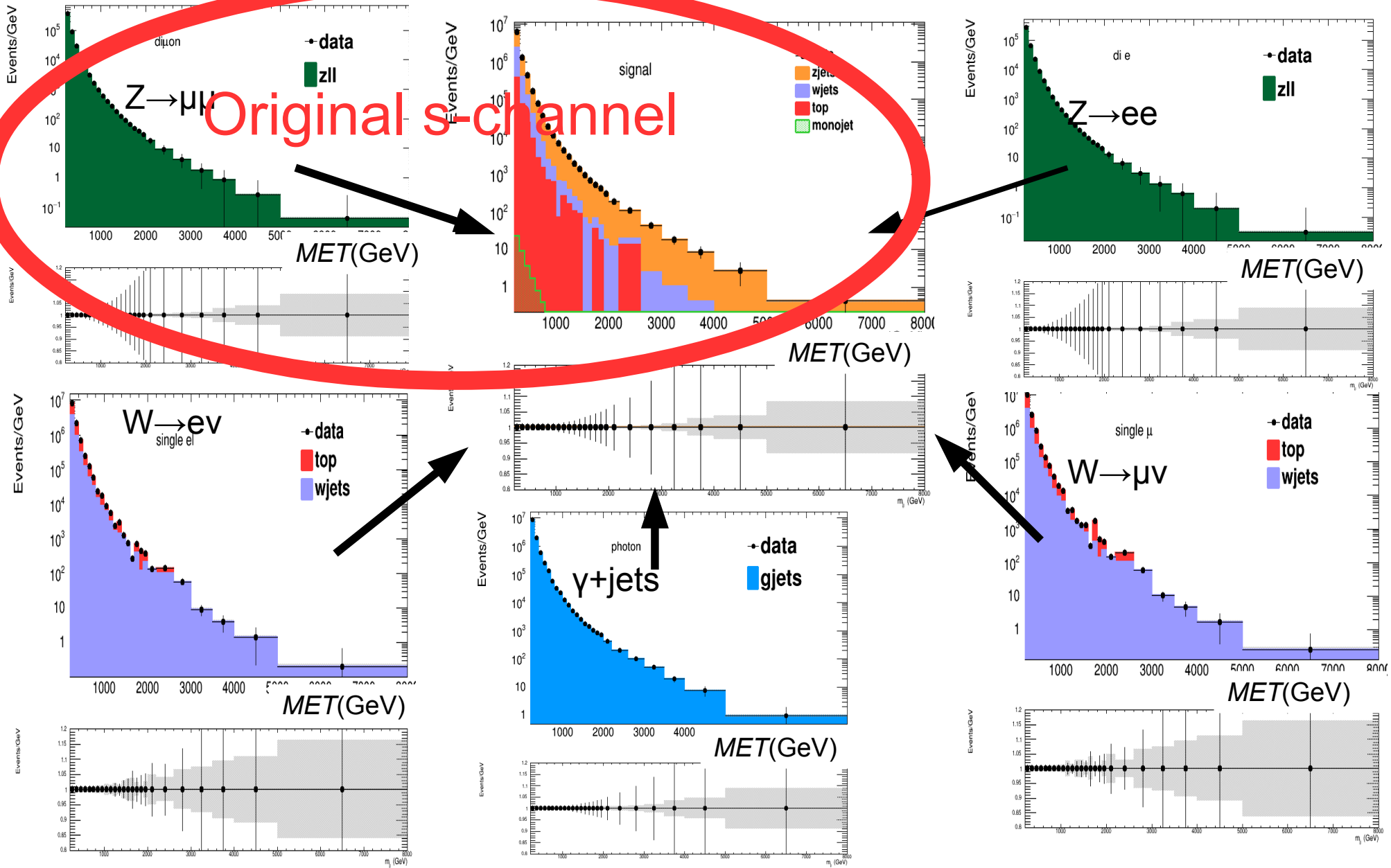
Monojet analysis @ CMS

The same fitting scheme applies to 100 TeV (fits 1ab^{-1})



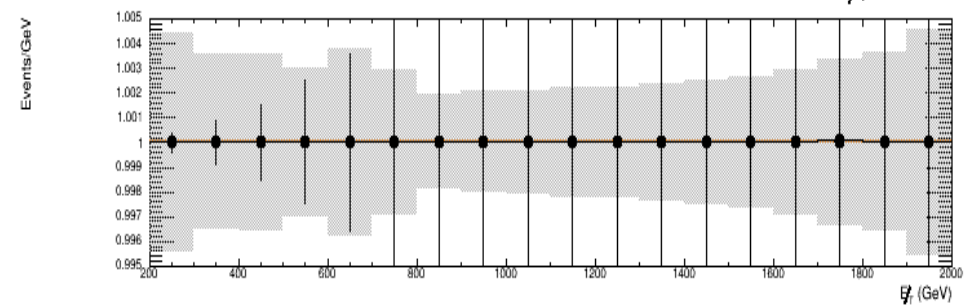
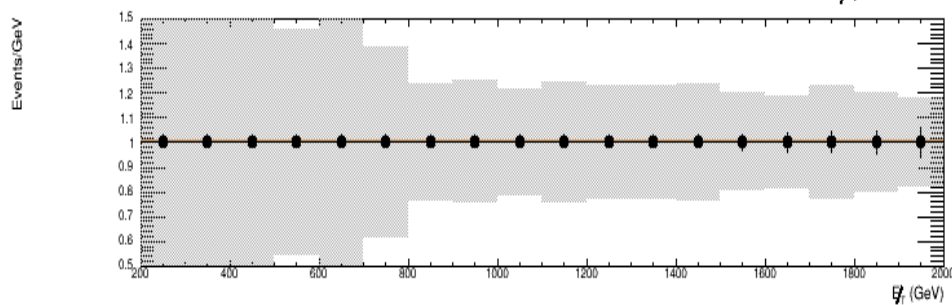
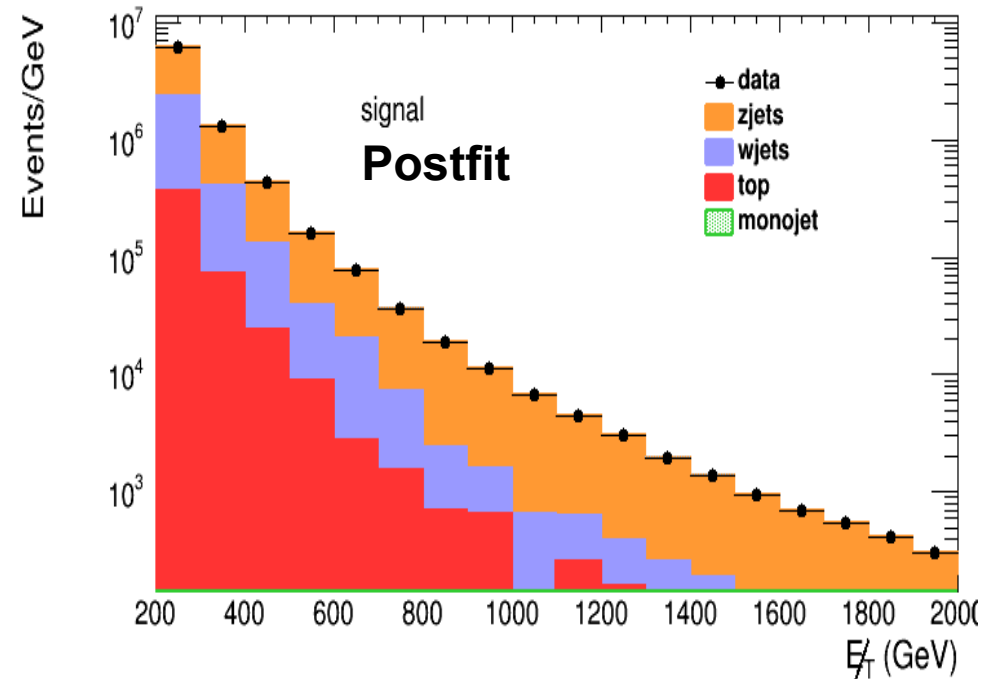
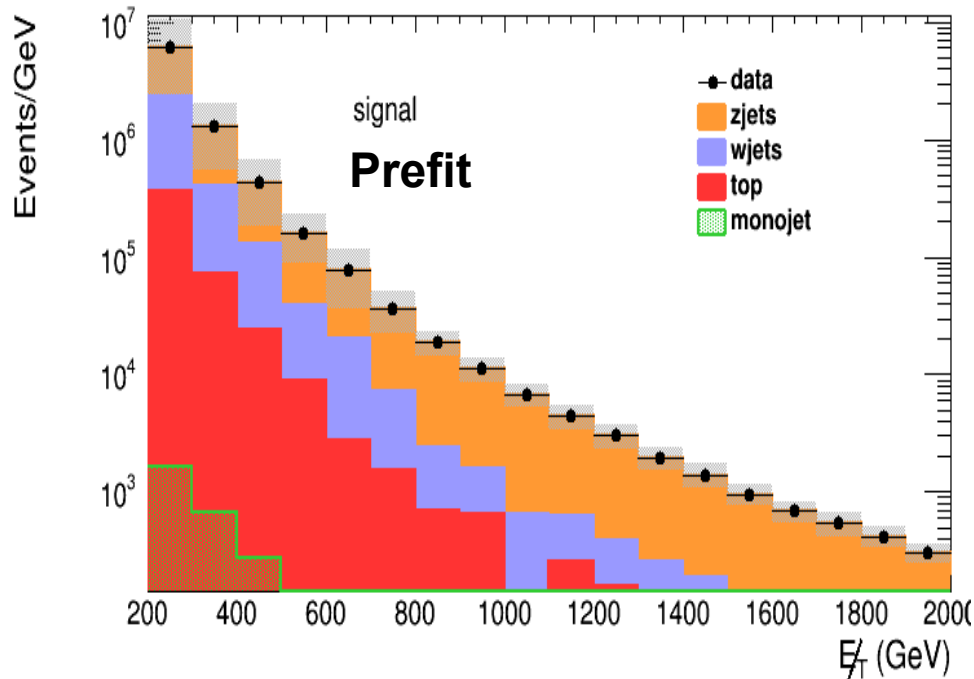
Monojet analysis @ CMS

The same fitting scheme applies to 100 TeV (fits 1ab^{-1})



What is the precision?

- Can probe a few % effects (NNLO precision)

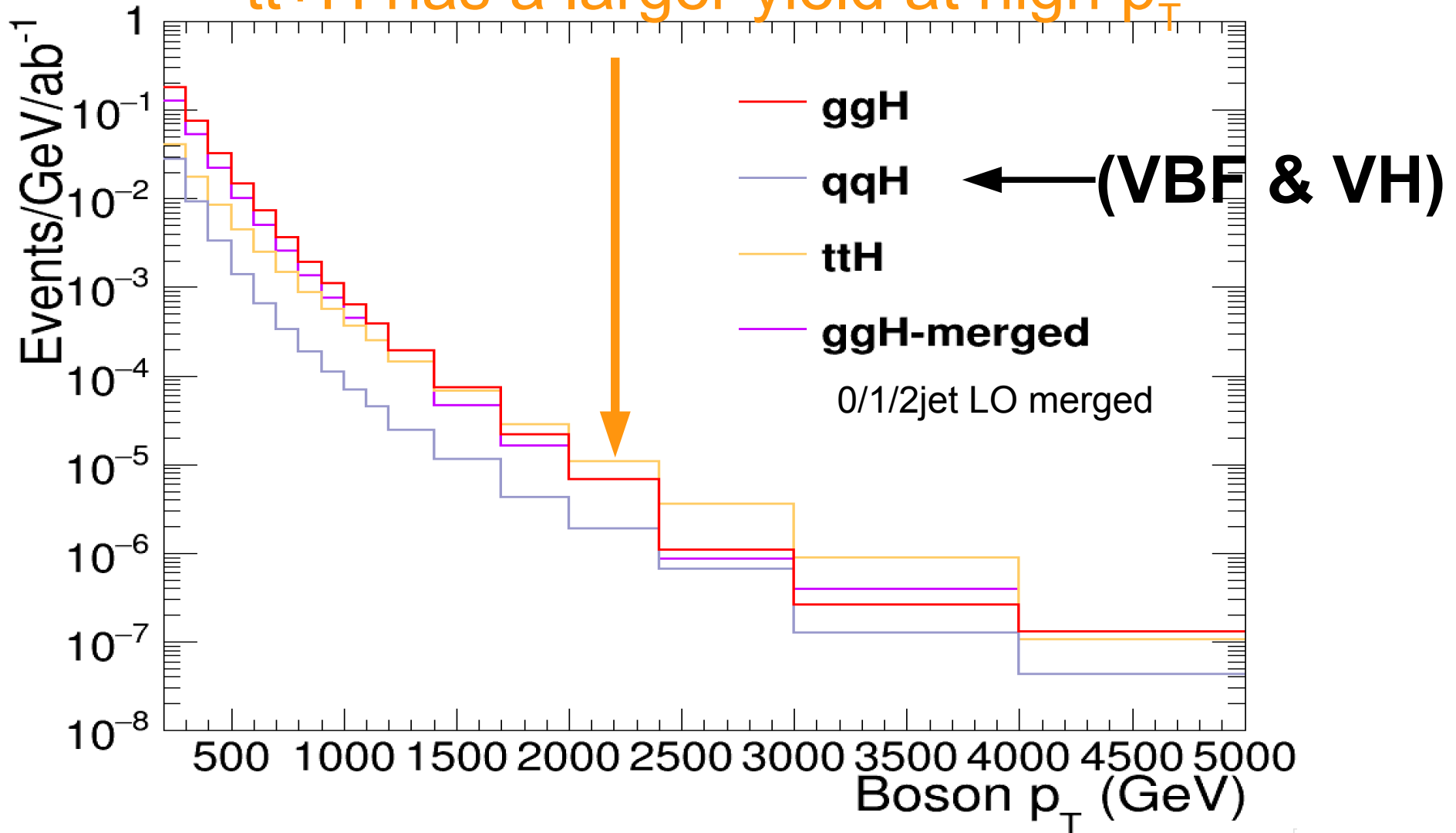


Through this scheme we can probe boson pT to 10^{-4} level

Higgs Invisible search

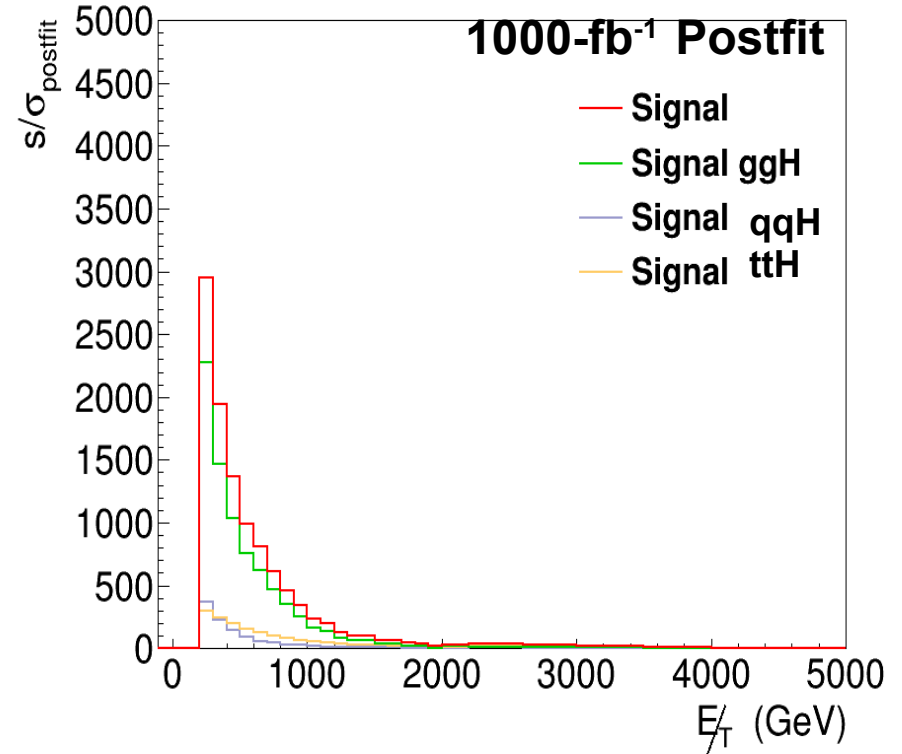
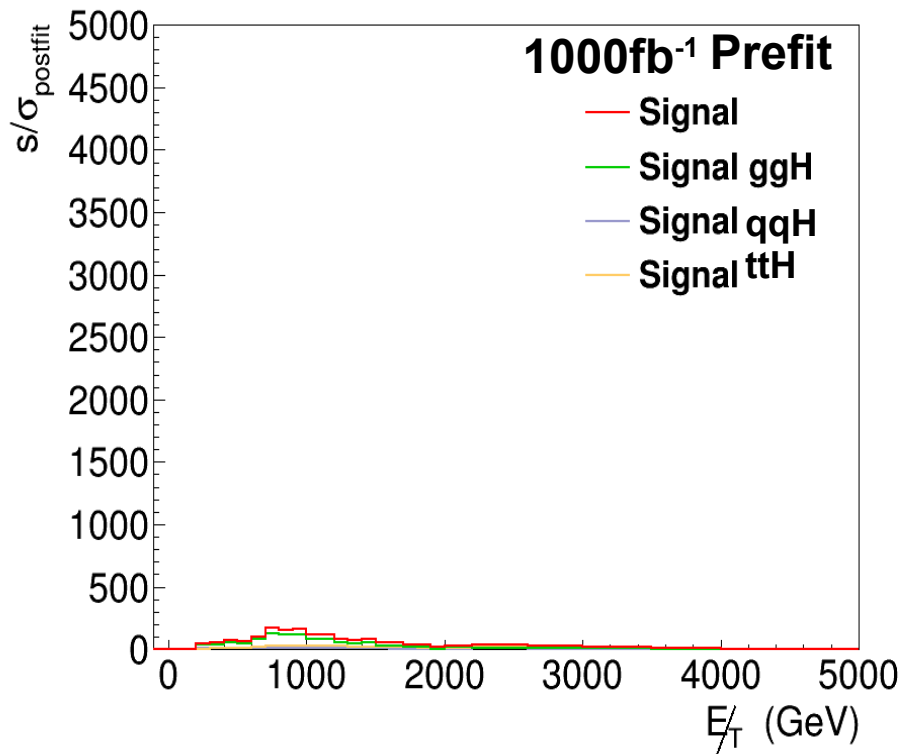
- A key feature at high p_T

$tt+H$ has a larger yield at high p_T



At 100 TeV ttH is more important, ggH still leads

Understanding sensitivity



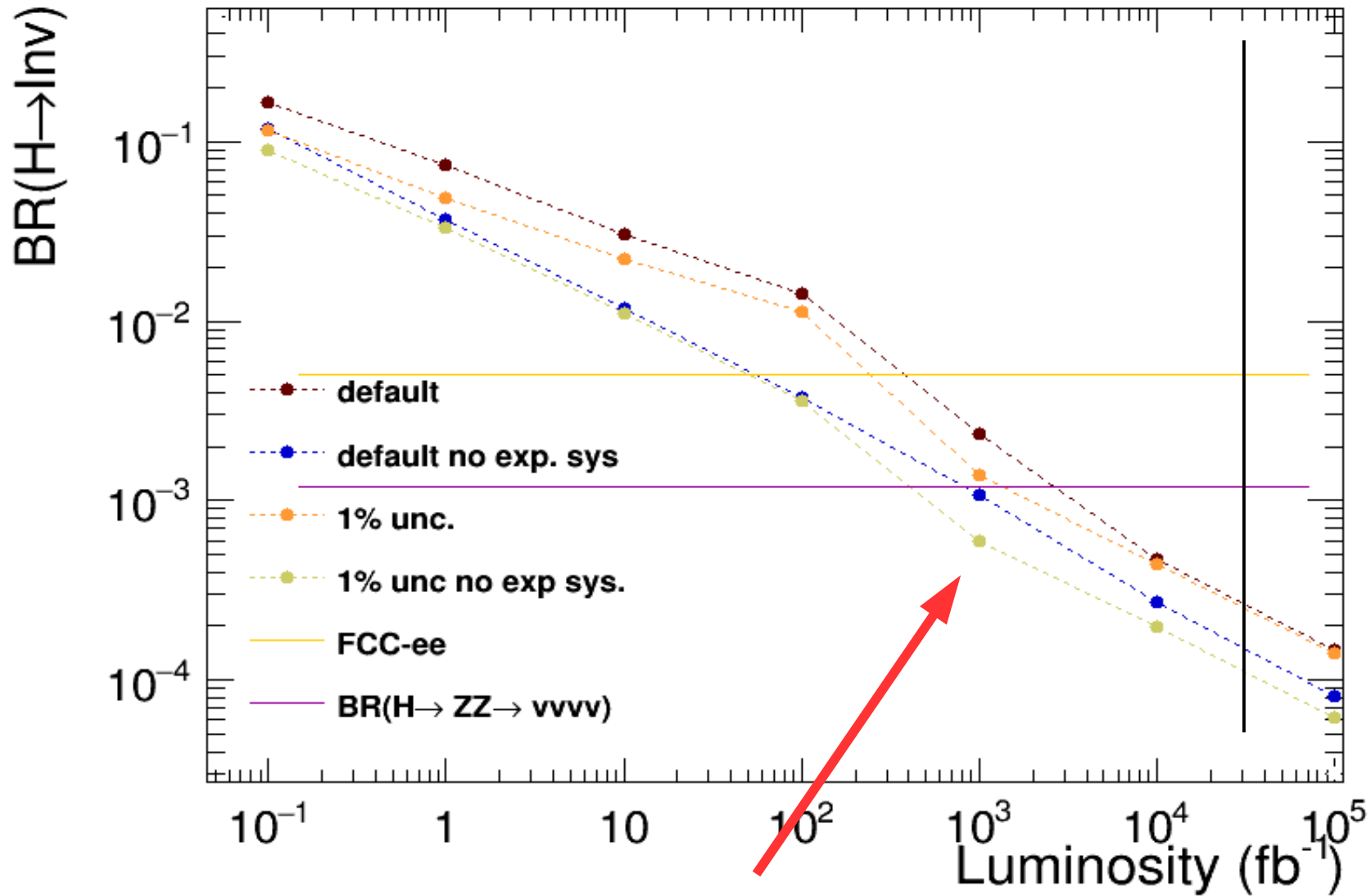
In both cases monojet dominates tt+H signal for sensitivity

Transition to ttH happens at 1-2 TeV (note no top selection)

Postfit brings an improvement in sensitivity

Especially at low MET : still critical for search

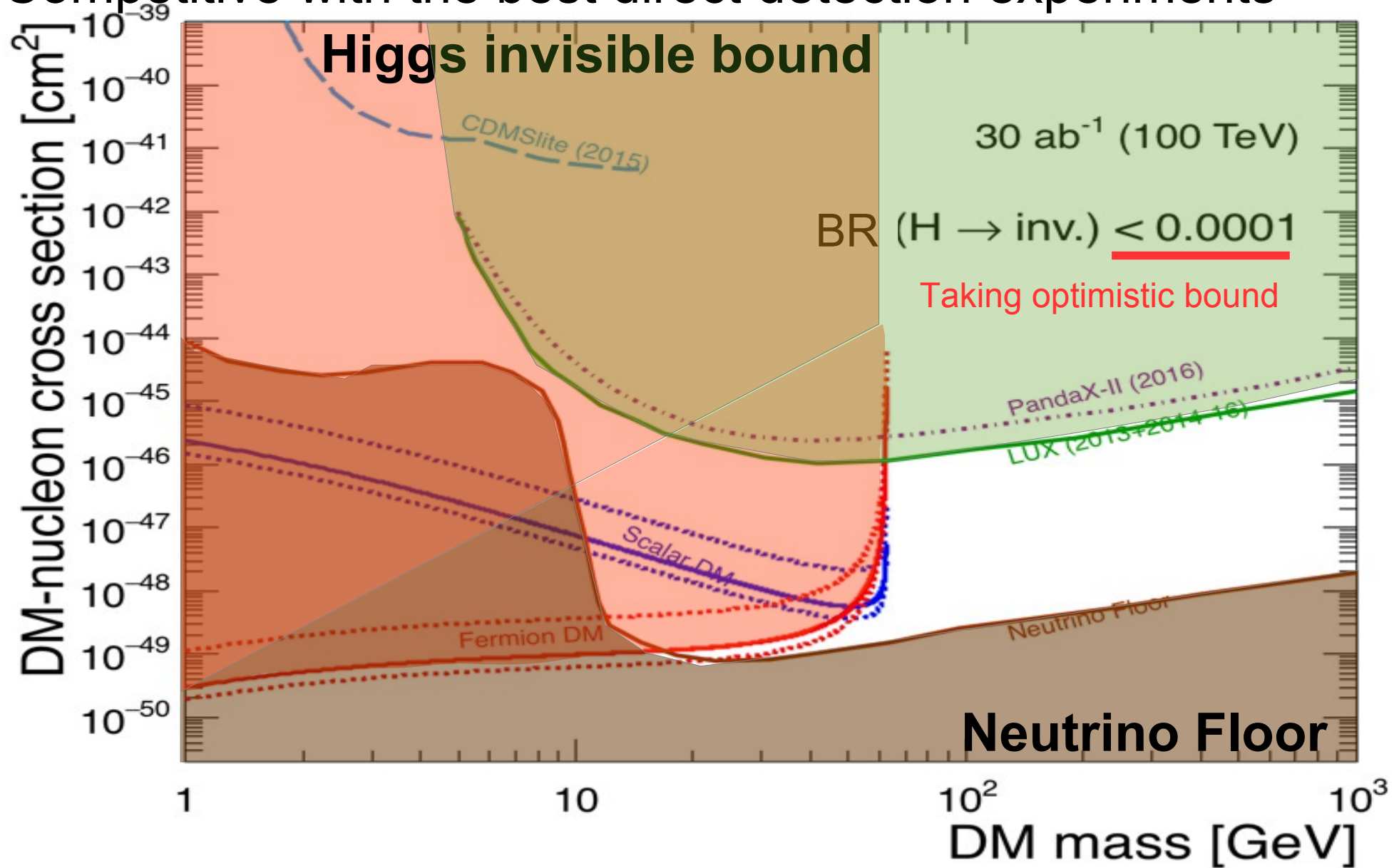
How do things scale?



There is no systematics wall

Future Bounds

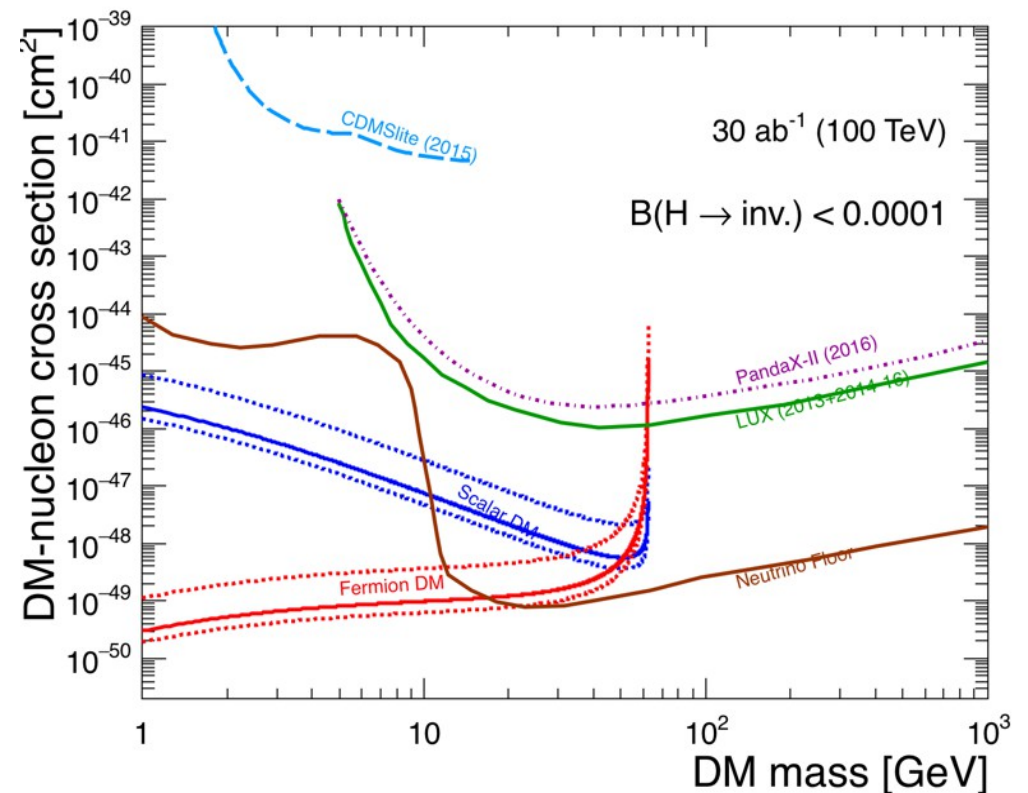
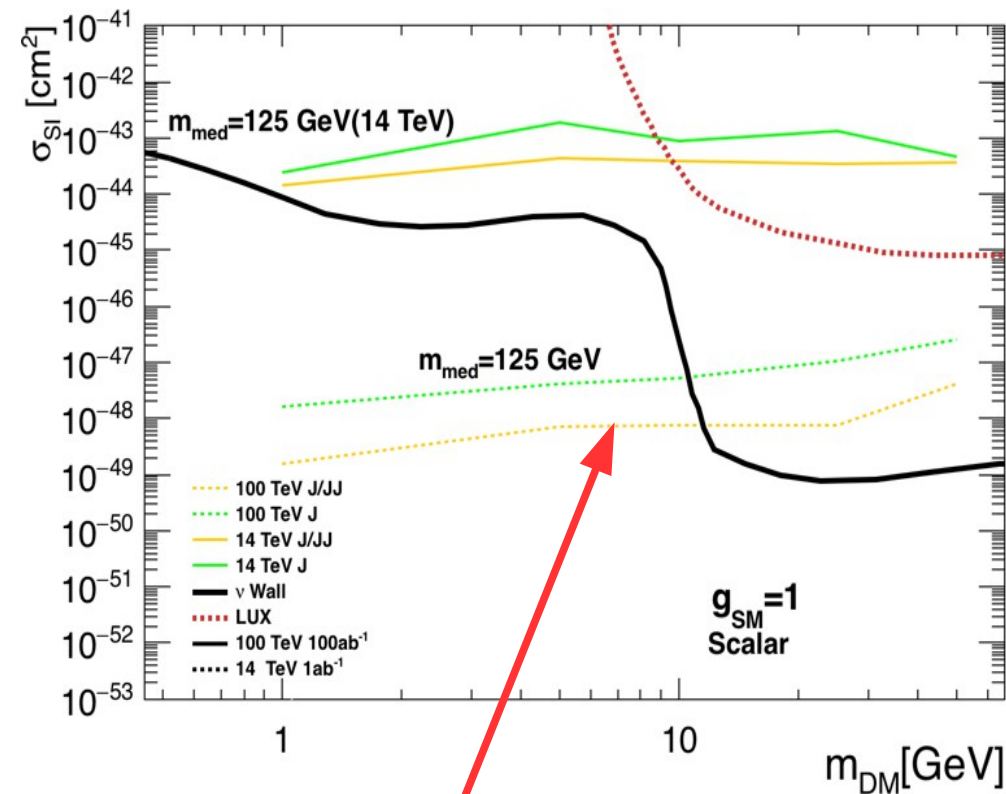
Competitive with the best direct detection experiments



Higgs invisible of 10^{-4} corresponds to g_{SM} from 10^{-3} to 10^{-2}

Back to s-channel

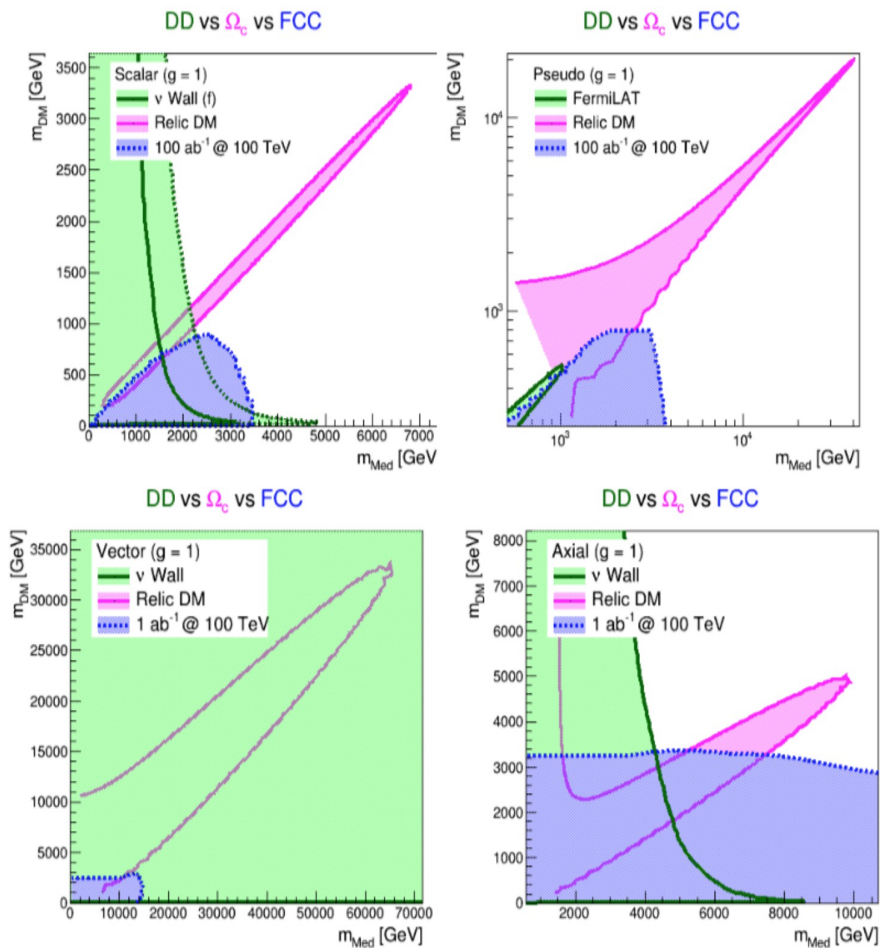
- Original studies used a worse uncertainty scheme
 - Fit strategy and setup was not as sophisticated
 - Updating with the full procedure is a good idea



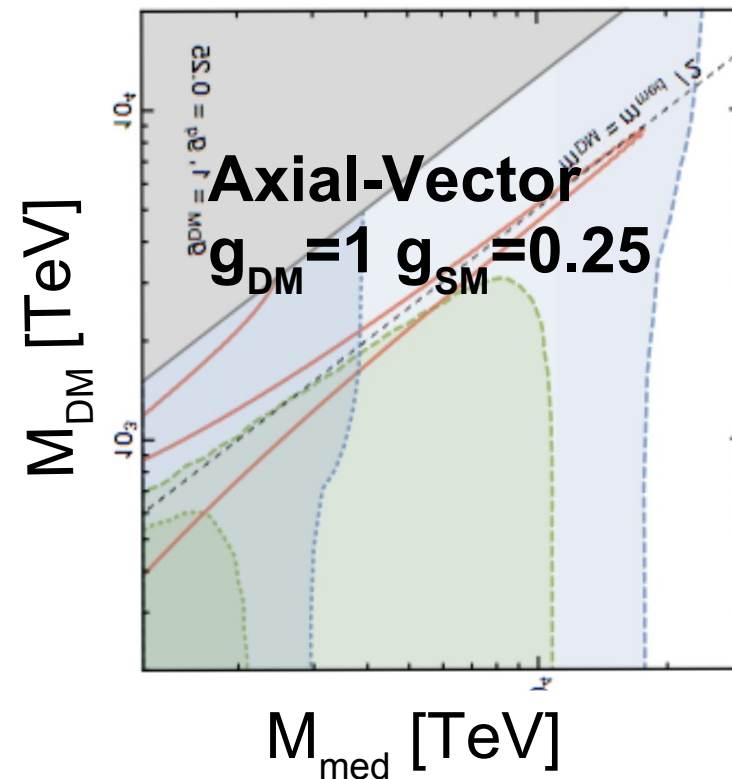
Original studies are worse

Back to s-channel

- Note the original Spin-1 used wrong coupling
 - Other groups have done this with DMWG recommends
- High mass bounds: have a good feel for the range of performance

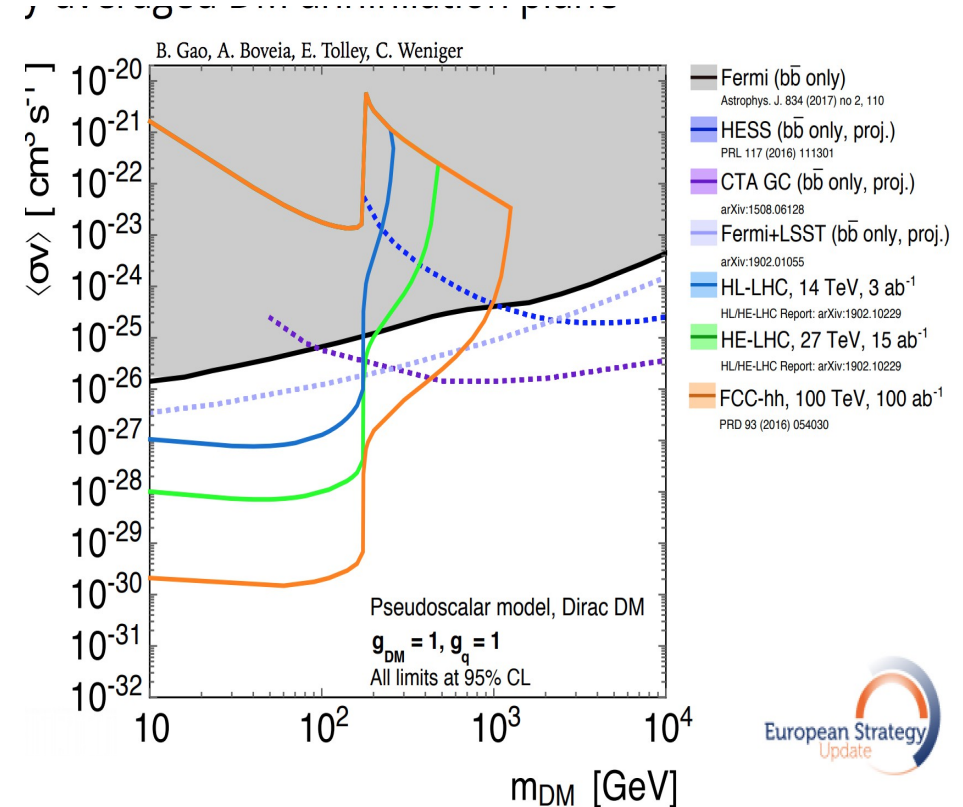
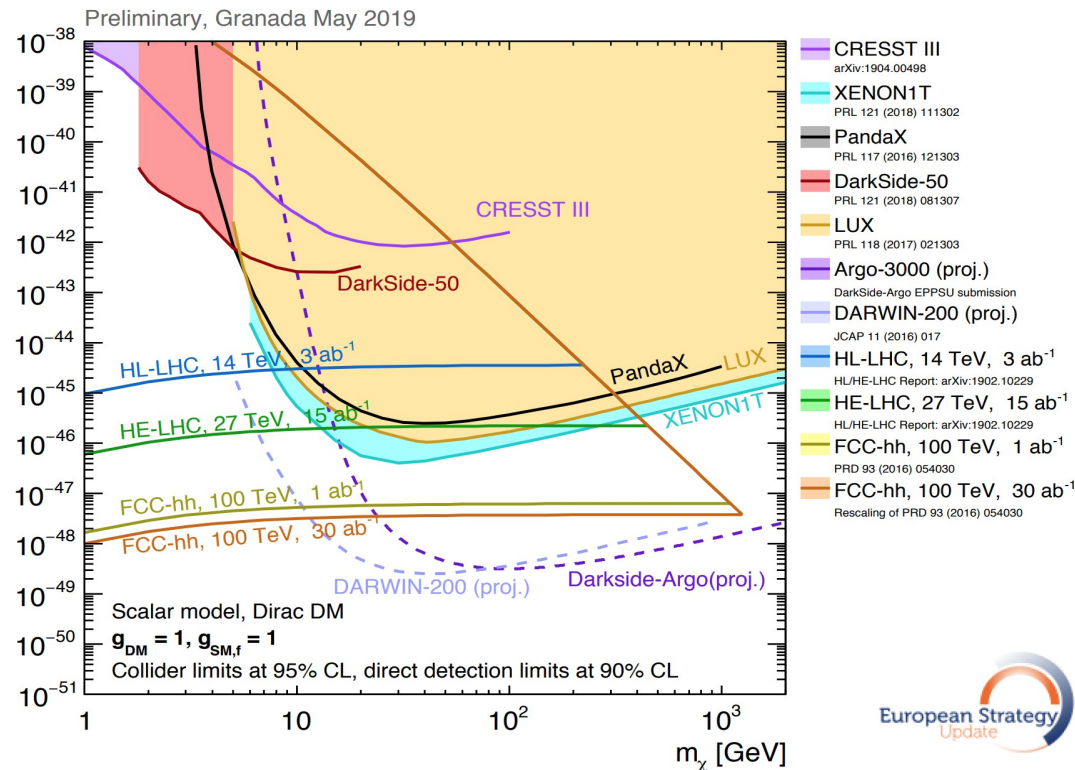


<https://arxiv.org/pdf/1503.05916.pdf>



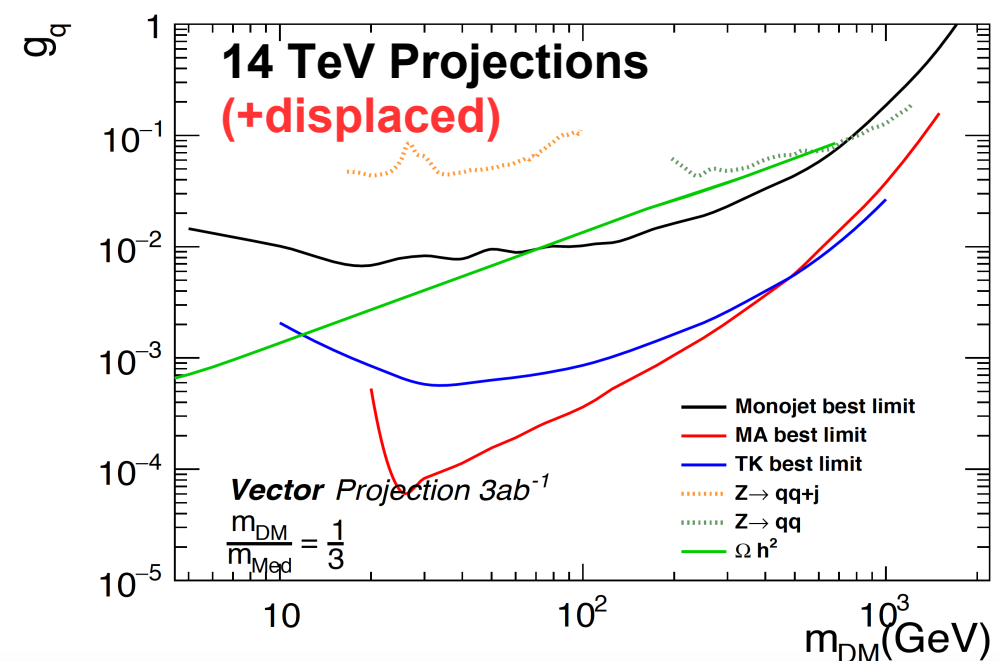
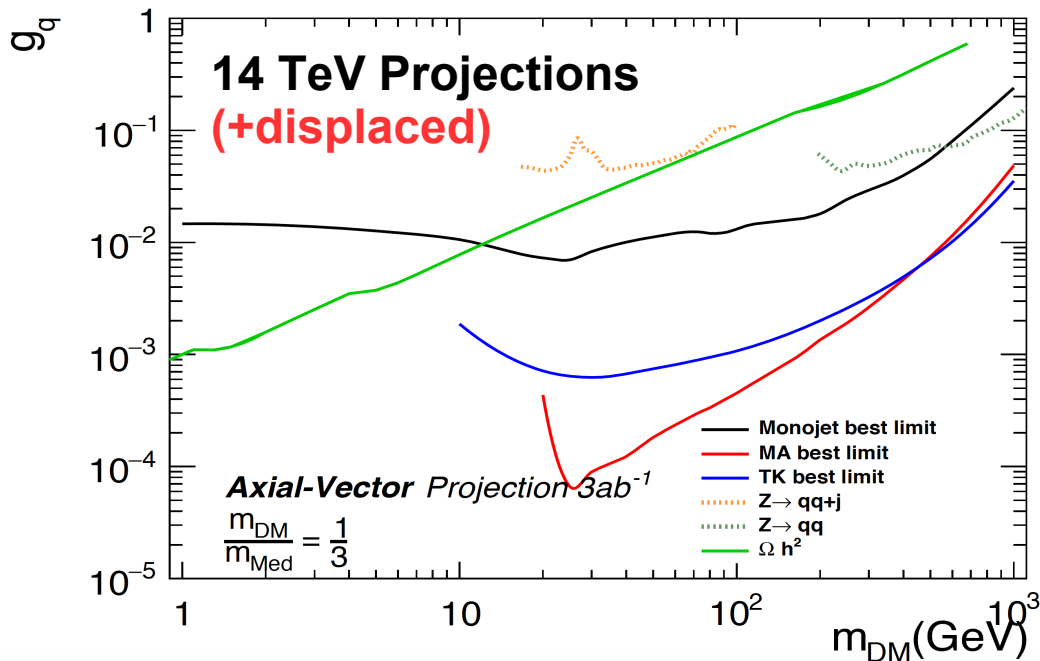
Updated Now

- Thanks to Caterina and Co for updated plots
 - Original ones used non-standard scheme
 - More on these plots in Antonio's talk



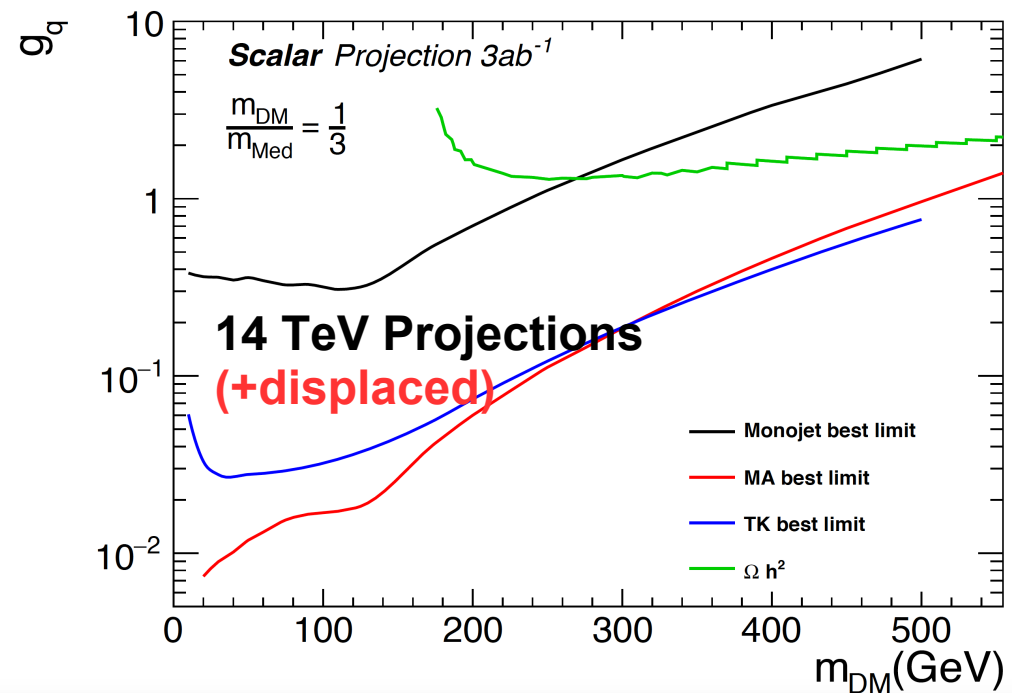
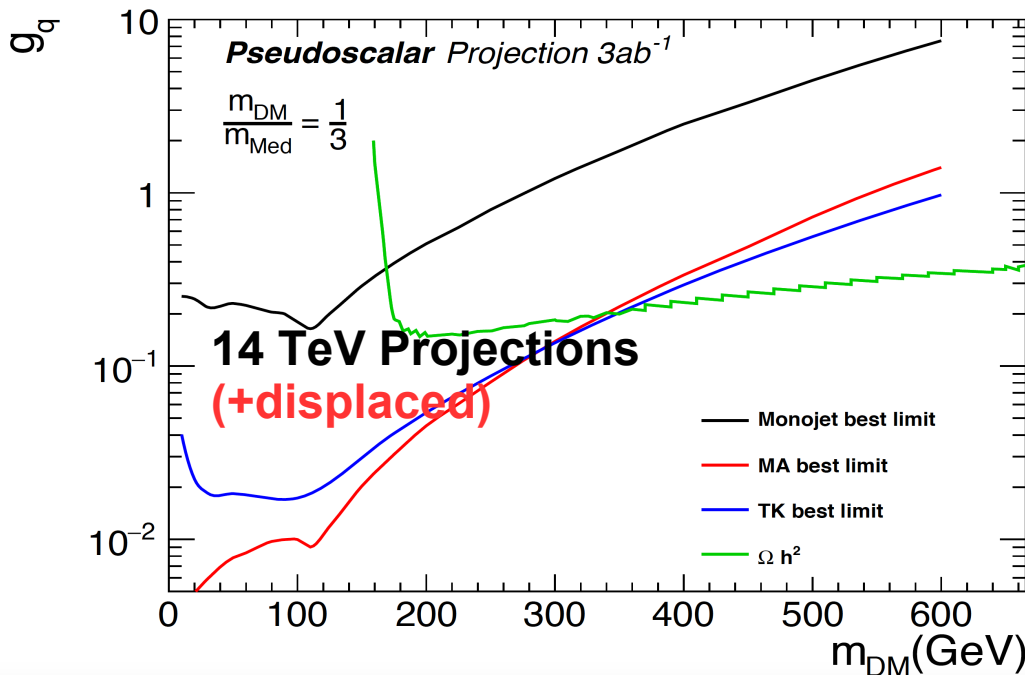
Going Forward

- Original study focused on mass reach
 - Coupling reach (like $H \rightarrow \text{Invisible}$) more interesting now
- Still have Delphes samples and fit framework
 - Can extend this to broad range of models
 - Would be happy to involve/pass on to others



Going Forward

- Original study focused on mass reach
 - Coupling reach (like $H \rightarrow \text{Invisible}$) more interesting now
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Conclusion

- Some old studies on 100 TeV
 - There were some limitations
 - There has been active work since then
 - Has not been extended to s-channel
 - But it could be good
- We should think about how we want to present?
 - Can we focus more on coupling and low masses?
 - **Are there other regions critical to the cosmic frontier**

