

ILC perspective on SUSY searches

Mikael Berggren¹

¹DESY, Hamburg

Snowmass EF08 , June 25, 2020



Outline

- 1 SUSY: What *do* we know ?
- 2 SUSY In The Briefing-book
 - Bino LSP
 - Wino/Higgsino LSP
- 3 SUSY at ILC
 - Picture gallery
- 4 Conclusions

SUSY: What *do* we know ?

Naturalness, hierarchy, DM, $g-2$ all prefers **light electro-weak** sector.

- Except for 3d gen. squarks, **the coloured sector** - where pp machines excel - **doesn't enter the game**.
- Both if the LSP is mainly higgsino or mainly wino, electro-weak sector is "compressed".
- Then, most sparticle-decays are **via cascades**. At the end of these cascades, the mass difference is small.
- For this, current limits from LHC are only for specific models, and LEP2 sets the scene.
- Same goes for sleptons in general, and the $\tilde{\tau}$ in particular.

SUSY: What *do* we know ?

Naturalness, hierarchy, DM, $g-2$ all prefers **light electro-weak** sector.

- Except for 3d gen. squarks, **the coloured sector** - where pp machines excel - **doesn't enter the game**.
- Both if the LSP is mainly higgsino or mainly wino, electro-weak sector is “compressed”.
- Then, most sparticle-decays are **via cascades**. At the end of these cascades, the mass difference is small.
- For this, current limits from LHC are only for specific models, and **LEP2** sets the scene.
- Same goes for sleptons in general, and the $\tilde{\tau}$ in particular.

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
 - Caveat: also CP-conservation. The experimental implication of CP violation needs study
- sfermions not NLSP (**idem**, except $\tilde{\tau}$ but even worse for FCChh...)
- Then: LSP is Bino, Wino, or Higgsino (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, up to values that makes the bosinos out-of-reach for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
 - Caveat: also CP-conservation. The experimental implication of CP violation needs study
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for FCChh...)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
 - Consider **any values**, and combinations of signs, up to values that makes the bosinos out-of-reach for any new facility \sim a few TeV.
 - Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
 - **No other prejudice.**

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
 - Caveat: also CP-conservation. The experimental implication of CP violation needs study
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for FCChh...)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, **up to values that makes the bosinos out-of-reach** for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**

What *would* be seen at colliders in the worst case?

- MSSM, R-parity conservation (R-parity violation **always easier** at e^+e^-)
 - Caveat: also CP-conservation. The experimental implication of CP violation needs study
- sfermions not NLSP (*idem*, except $\tilde{\tau}$ but even worse for FCChh...)
- Then: LSP is **Bino, Wino, or Higgsino** (more or less pure), same for the NLSP
- M_1, M_2 and μ are the main-players.
- Consider **any values**, and combinations of signs, **up to values that makes the bosinos out-of-reach** for any new facility \sim a few TeV.
- Also vary other parameters ($\beta, M_A, M_{sfermion}$) with less impact.
- **No other prejudice.**

Aspects of the spectrum

With a broad brush:

- For Higgsino-like LSP,
- for Wino-like LSP,

and

- for Bino-like LSP

ie. Except for Bino-LSP, the
LSP-NLSP splitting is small !

Aspects of the spectrum

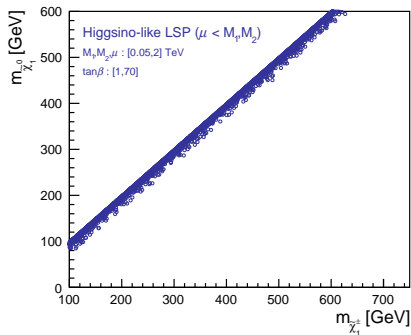
With a broad brush:

- For Higgsino-like LSP,
- for Wino-like LSP,

and

- for Bino-like LSP

ie. Except for Bino-LSP, the
LSP-NLSP splitting is small !



Aspects of the spectrum

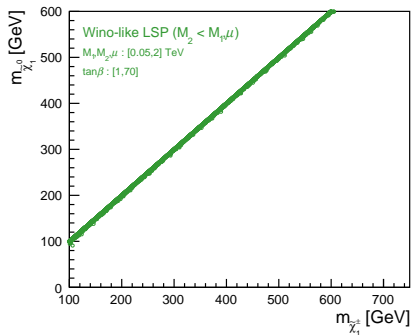
With a broad brush:

- For Higgsino-like LSP,
- for Wino-like LSP,

and

- for Bino-like LSP

ie. Except for Bino-LSP, the
LSP-NLSP splitting is small !

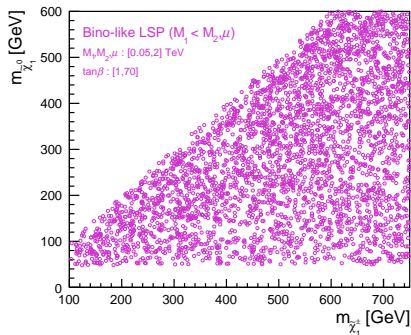


Aspects of the spectrum

With a broad brush:

- For Higgsino-like LSP,
 - for Wino-like LSP,
- and
- for Bino-like LSP

ie. Except for Bino-LSP, the
LSP-NLSP splitting is small !

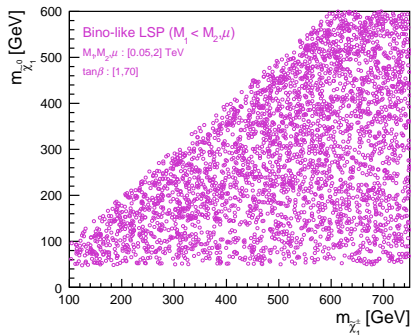


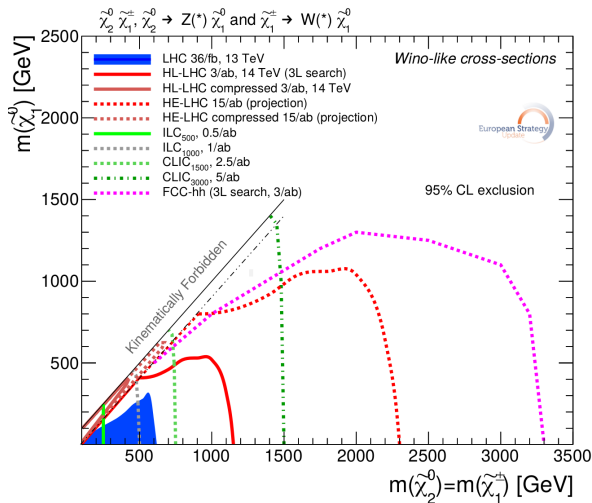
Aspects of the spectrum

With a broad brush:

- For Higgsino-like LSP,
 - for Wino-like LSP,
- and
- for Bino-like LSP

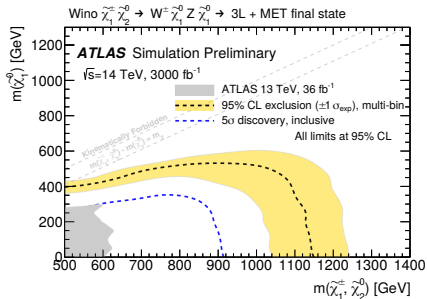
ie. Except for Bino-LSP, the
LSP-NLSP splitting is small !



SUSY In The Briefing-book: Bino LSP (ie. large Δ_M)

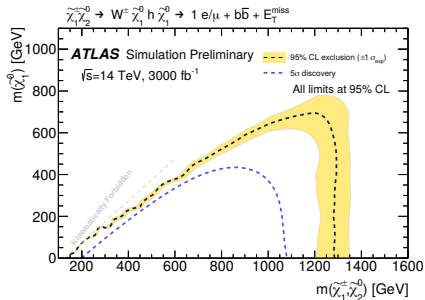
SUSY In The Briefing-book: Bino LSP - Sources

- From PHYS-PUB-2018-04 (ATLAS HL-LHC projection). Then extrapolated (up *and* down)
- Note that the BB curve is exclusion, not discovery!
- This is for the best decay mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .



SUSY In The Briefing-book: Bino LSP - Sources

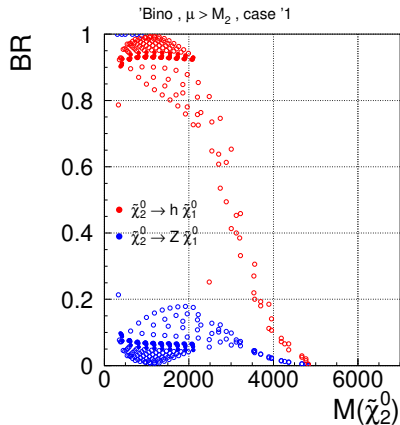
- From PHYS-PUB-2018-04 (ATLAS HL-LHC projection). Then extrapolated (up *and* down)
- Note that the BB curve is exclusion, not discovery!
- This is for the best decay mode!
- The other decay mode
- Better at $M_{LSP}=0$, weaker at lower Δ_M .



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

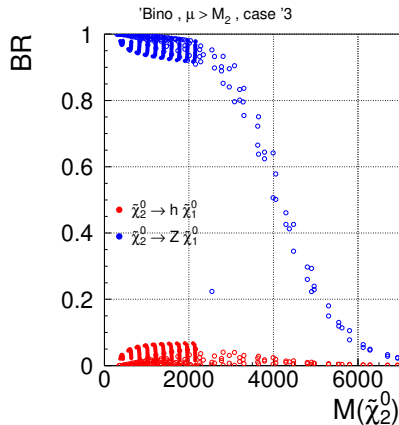
- Vary relative signs of μ , M_1 , and M_2 , for $\mu > M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

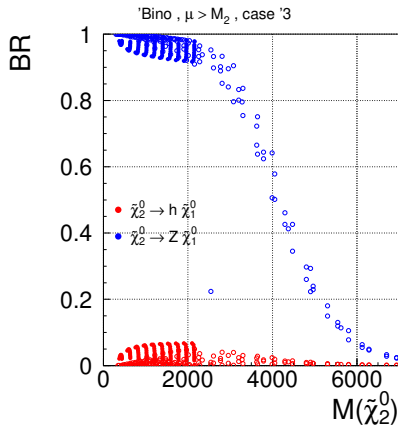
- Vary relative signs of μ , M_1 , and M_2 , for $\mu > M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



Bino LSP: BRs

Why is the decay-mode an issue? Here's why :

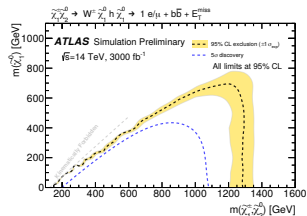
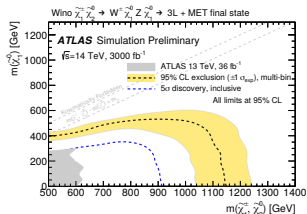
- Vary relative signs of μ , M_1 , and M_2 , for $\mu > M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the *intersection* of the two plots, not the *union*!

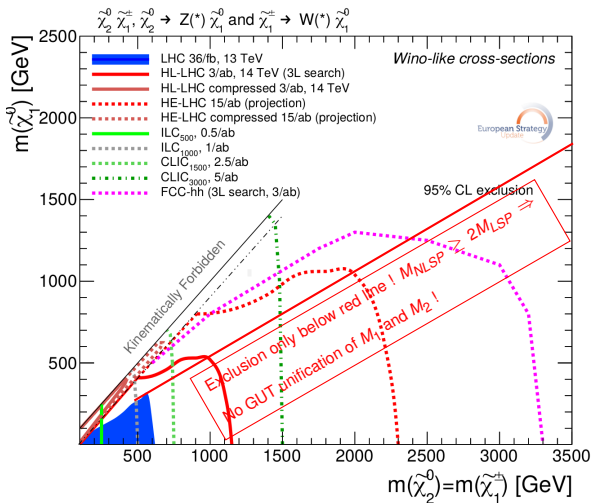


Bino LSP: BRs

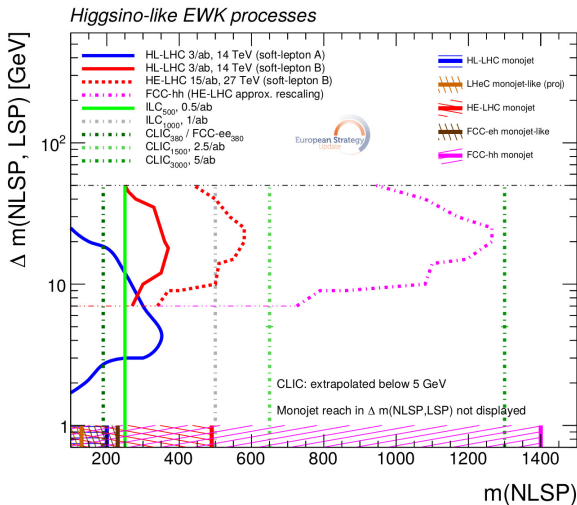
Why is the decay-mode an issue? Here's why :

- Vary relative signs of μ , M_1 , and M_2 , for $\mu > M_2$
- Conclusion: Whether the Z or the H decay-mode of $\tilde{\chi}_2^0$ dominates is **pure speculation** and
- The exclusion-region is the **intersection** of the two plots, not the **union**!



SUSY In The Briefing-book: Bino LSP (ie. large Δ_M)

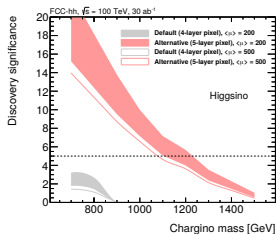
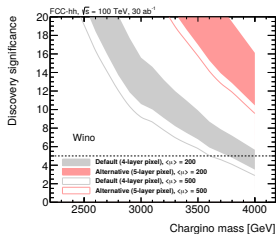
SUSY In The Briefing-book: Wino/Higgsino LSP



SUSY In The Briefing book: Wino/Higgsino LSP - Sources

(Don't look at the pink curves - they correspond to a detector that is never considered anywhere else in the CDR)

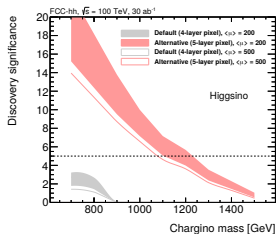
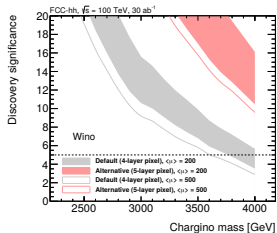
- The “Disappearing tracks” was done by FCChh (in the CDR)
 - FCChh-detector (better than ATLAS in this case: first layer of VD closer.)
 - FCChh-ish PU (but still too small: 500 vs. CDR number 955)
 - For higgsinos: Only *just* reaches 2σ



SUSY In The Briefing book: Wino/Higgsino LSP - Sources

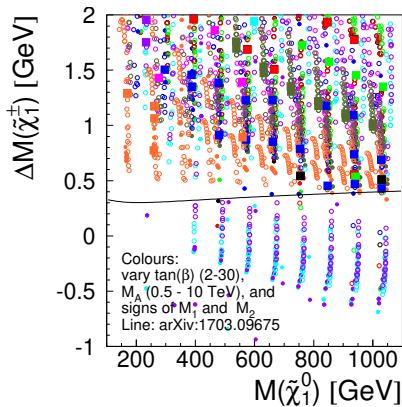
(Don't look at the pink curves - they correspond to a detector that is never considered anywhere else in the CDR)

- The “Disappearing tracks” was done by FCChh (in the CDR)
 - FCChh-detector (better than ATLAS in this case: first layer of VD closer.)
 - FCChh-ish PU (but still too small: 500 vs. CDR number 955)
 - For higgsinos: Only *just* reaches 2σ



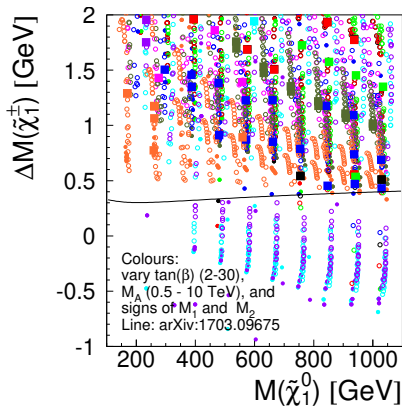
Key element for “Disappearing tracks”: $\Delta(M)$

- Higgsino LSP. The line is the absolute limit mentioned in the BB.
- Let other parameters vary, any signs, M_1 and M_2 close to μ Note that the LSP often would be the $\tilde{\chi}_1^\pm$!
- Reason: 1703.09675 considers *only SM* effects on the mass-splitting, ie. that M_1 and $M_2 \gg \mu$
- Same for Wino LSP.



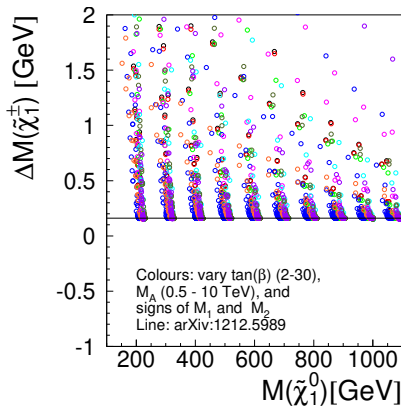
Key element for “Disappearing tracks”: $\Delta(M)$

- Higgsino LSP. The line is the absolute limit mentioned in the BB.
- Let other parameters vary, any signs, M_1 and M_2 close to μ Note that the LSP often would be the $\tilde{\chi}_1^\pm$!
- Reason: 1703.09675 considers *only SM* effects on the mass-splitting, ie. that M_1 and $M_2 \gg \mu$
- Same for Wino LSP.



Key element for “Disappearing tracks”: $\Delta(M)$

- Higgsino LSP. The line is the absolute limit mentioned in the BB.
- Let other parameters vary, any signs, M_1 and M_2 close to μ Note that the LSP often would be the $\tilde{\chi}_1^\pm$!
- Reason: 1703.09675 considers *only SM* effects on the mass-splitting, ie. that M_1 and $M_2 \gg \mu$
- Same for Wino LSP.



Key element for “Disappearing tracks”: $\Delta(M)$

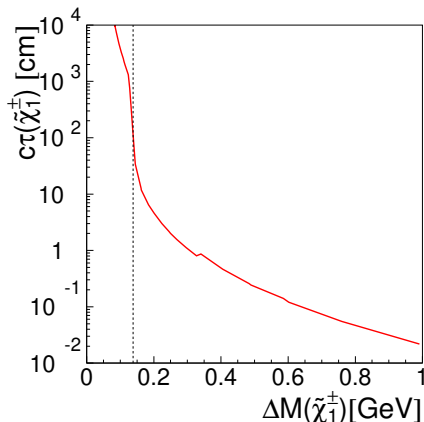
Why is this important?

- Because $c\tau$ depends on $\Delta(M)$, and $c\tau$ needs to be macroscopic to get “Disappearing tracks”
- $c\tau$ for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.

Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

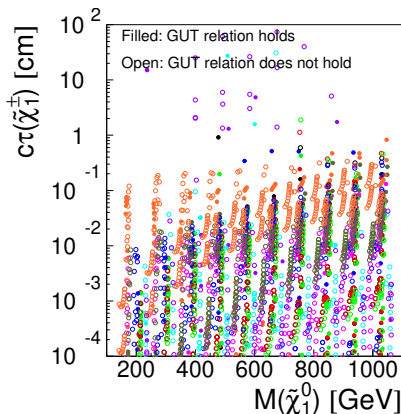
- Because $c\tau$ depends on $\Delta(M)$, and $c\tau$ needs to be macroscopic to get “Disappearing tracks”
- $c\tau$ for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.



Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

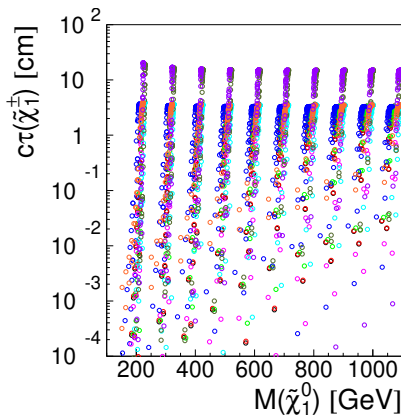
- Because $c_{\mathcal{T}}$ depends on $\Delta(M)$, and $c_{\mathcal{T}}$ needs to be macroscopic to get “Disappearing tracks”
- $c_{\mathcal{T}}$ for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.



Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

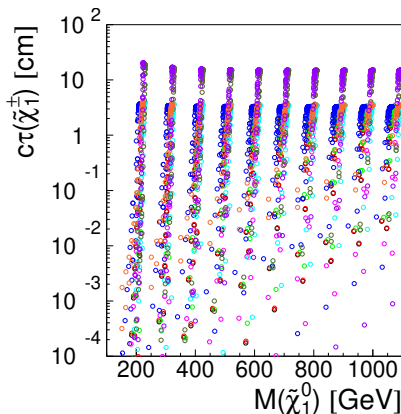
- Because $c\tau$ depends on $\Delta(M)$, and $c\tau$ needs to be macroscopic to get “Disappearing tracks”
- $c\tau$ for Higgsino LSP
- ... and Wino LSP
- Conclusion: Not at all sure that that lifetime will be large. Good chances - no guarantee - for Wino, unlikely for Higgsino.



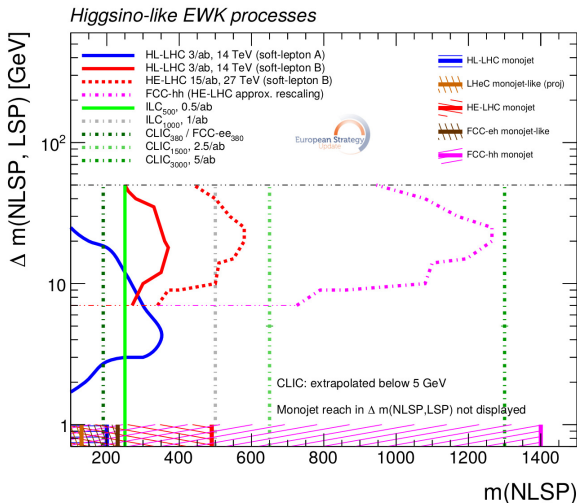
Key element for “Disappearing tracks”: $\Delta(M)$

Why is this important?

- Because $c\tau$ depends on $\Delta(M)$, and $c\tau$ needs to be macroscopic to get “Disappearing tracks”
- $c\tau$ for Higgsino LSP
- ... and Wino LSP
- Conclusion: **Not at all sure that that lifetime will be large.** Good chances - no guarantee - for Wino, unlikely for Higgsino.

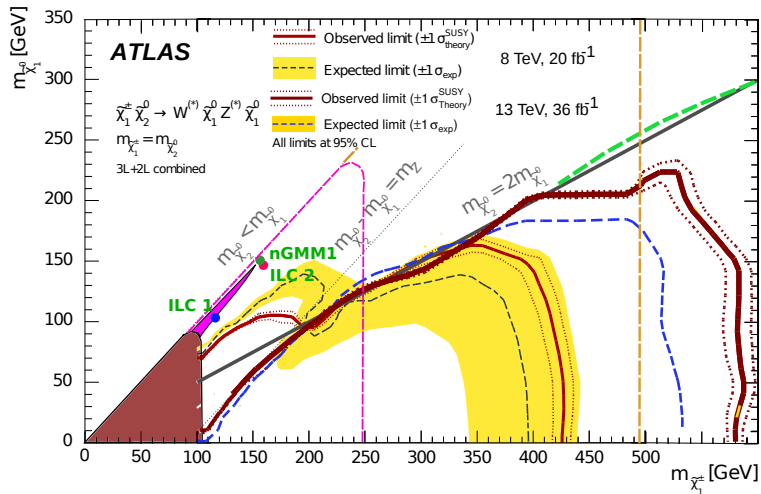


SUSY In The Briefing-book: Wino/Higgsino LSP



So: Disappearing tracks exclusion is actually off the scale !

SUSY: All-in-one

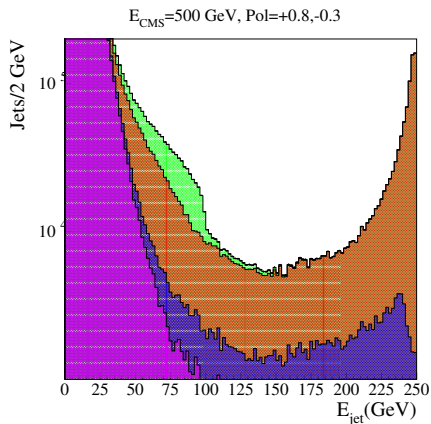


SUSY at ILC

- Exclusion = Discovery !
- Exclusion **is** Exclusion
- SUSY search = SUSY measurement
- Model testing
- Experimentally, SUSY at ILC more like a B-factory than PP

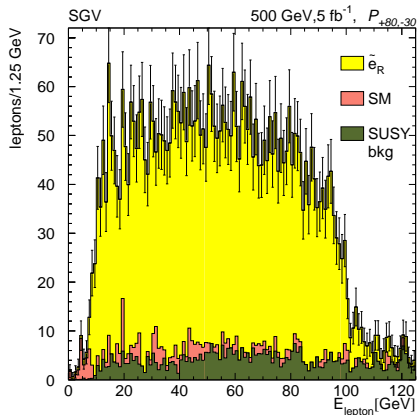
SUSY at ILC: Exclusion = Discovery !

- Just select low multiplicity events...
- Selectrons after one week ...
- Exclusion and discovery
 - Smuon NLSP
 - Stau NLSP
- Chargino NLSP with low $\Delta(M)$, extrapolated from LEP II



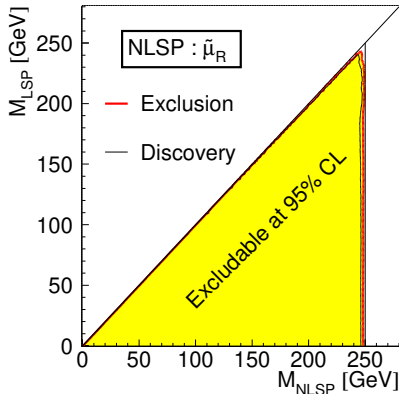
SUSY at ILC: Exclusion = Discovery !

- Just select low multiplicity events...
- Selectrons after one week ...
- Exclusion and discovery
 - Smuon NLSP
 - Stau NLSP
- Chargino NLSP with low $\Delta(M)$, extrapolated from LEP II



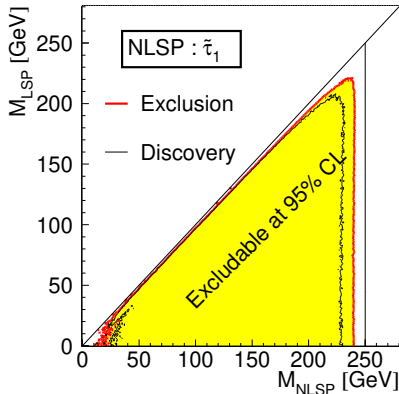
SUSY at ILC: Exclusion = Discovery !

- Just select low multiplicity events...
- Selectrons after one week ...
- Exclusion and discovery
 - Smuon NLSP
 - Stau NLSP
- Chargino NLSP with low $\Delta(M)$, extrapolated from LEP II



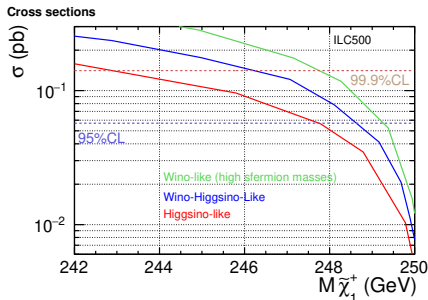
SUSY at ILC: Exclusion = Discovery !

- Just select low multiplicity events...
- Selectrons after one week ...
- Exclusion and discovery
 - Smuon NLSP
 - Stau NLSP
- Chargino NLSP with low $\Delta(M)$, extrapolated from LEP II



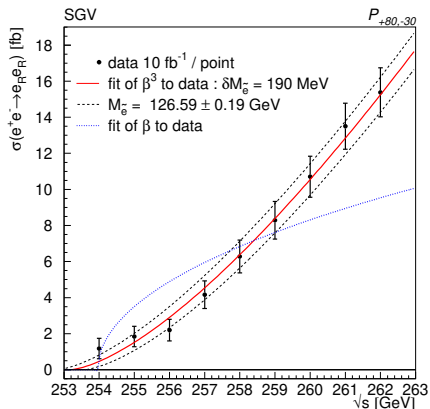
SUSY at ILC: Exclusion = Discovery !

- Just select low multiplicity events...
- Selectrons after one week ...
- Exclusion and discovery
 - Smuon NLSP
 - Stau NLSP
- Chargino NLSP with low $\Delta(M)$, extrapolated from LEP II



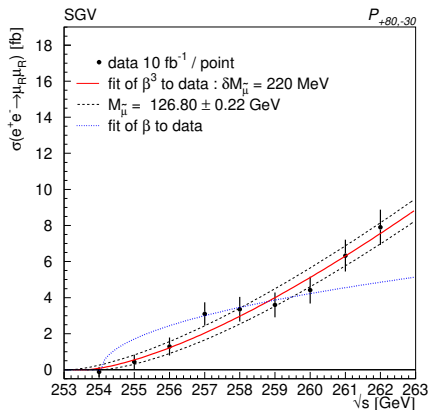
SUSY at ILC: SUSY is SUSY

- Scan of the selectron threshold
- ... and the smuon.
- \Rightarrow P-wave, not S-wave \Rightarrow the new states are scalars!



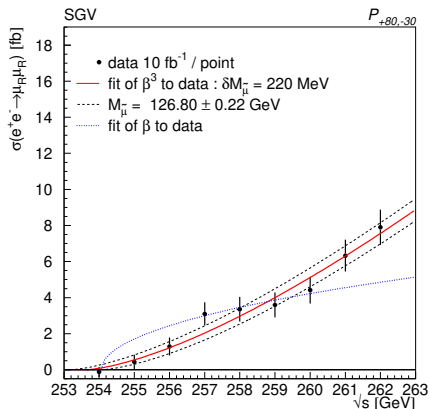
SUSY at ILC: SUSY is SUSY

- Scan of the selectron threshold
- ... and the smuon.
- \Rightarrow P-wave, not S-wave \Rightarrow the new states are scalars!



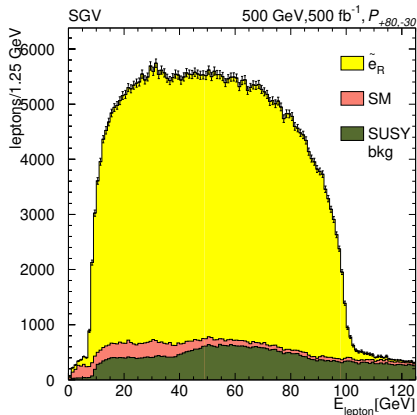
SUSY at ILC: SUSY is SUSY

- Scan of the selectron threshold
- ... and the smuon.
- \Rightarrow P-wave, not S-wave \Rightarrow the new states **are** scalars!



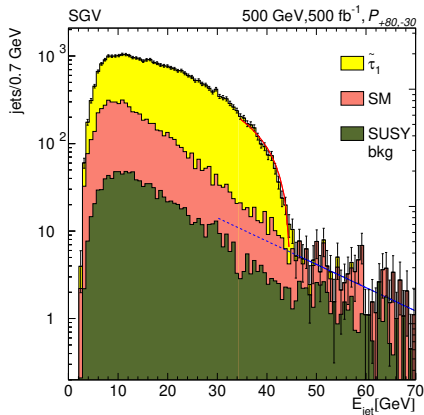
SUSY: Measure properties

- Selectron spectrum after \sim one year....
- Same for staus...
- and charginos.
- Masses determined to ~ 100 MeV = permil-level.



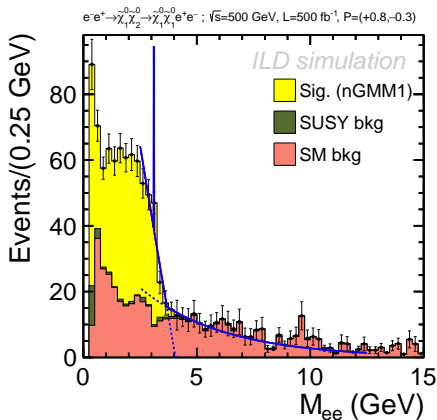
SUSY: Measure properties

- Selectron spectrum after \sim one year....
- Same for staus...
- and charginos.
- Masses determined to ~ 100 MeV = permil-level.



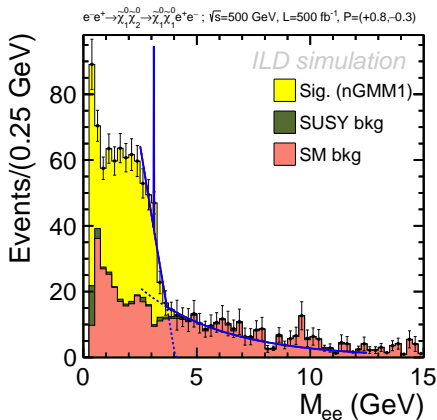
SUSY: Measure properties

- Selectron spectrum after \sim one year....
- Same for staus...
- and charginos.
- Masses determined to ~ 100 MeV = permil-level.



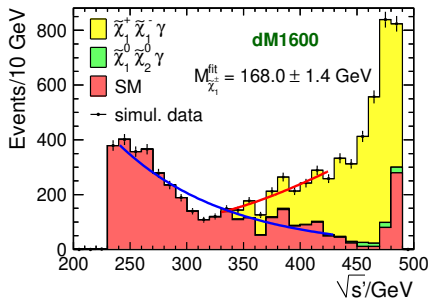
SUSY: Measure properties

- Selectron spectrum after \sim one year....
- Same for staus...
- and charginos.
- Masses determined to ~ 100 MeV = permil-level.



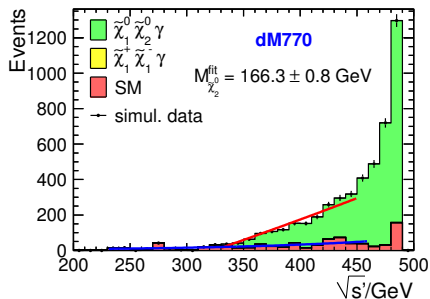
SUSY at ILC: Measure properties, even in tough cases

- Chargino mass determination with $\Delta(M) = 1.6$ GeV ...
- Or neutralino mass with $\Delta(M) = 770$ MeV.



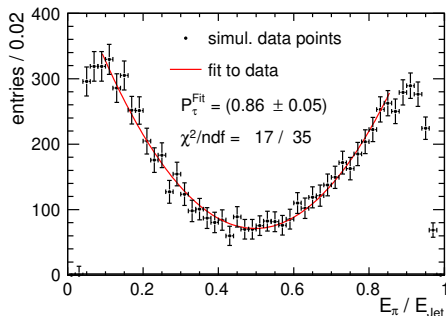
SUSY at ILC: Measure properties, even in tough cases

- Chargino mass determination with $\Delta(M) = 1.6$ GeV ...
- Or neutralino mass with $\Delta(M) = 770$ MeV.



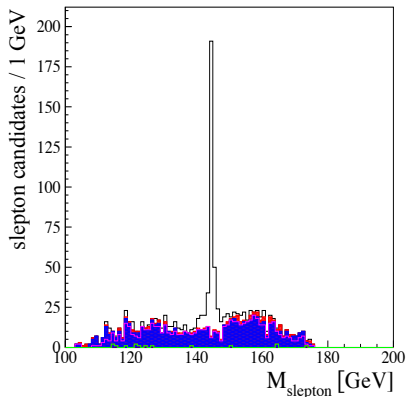
SUSY at ILC: Measure SUSY, not just mass

- Also other aspects. Here
 - Determine the polarisation of taus in stau decays
 - Direct handle on stau mixing.



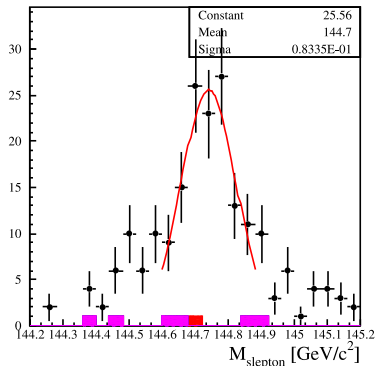
SUSY at ILC: Full reconstruction

- Or, in a case with cascade-decays $\tilde{\chi}_2^0 \rightarrow \tilde{l} \rightarrow l \tilde{\chi}_1^0$: Enough constraints to do full kinematic reconstruction.
- Smuon/Selectron
- Zoom in: Mass to less than 10 MeV
- Even for staus.



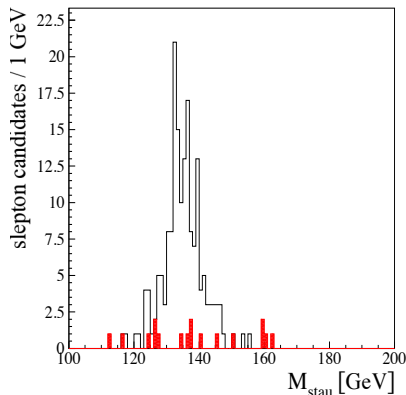
SUSY at ILC: Full reconstruction

- Or, in a case with cascade-decays $\tilde{\chi}_2^0 \rightarrow \tilde{\ell} \ell \rightarrow \ell \tilde{\chi}_1^0$: Enough constraints to do full kinematic reconstruction.
- Smuon/Selectron
- Zoom in: Mass to less than 10 MeV
- Even for staus.



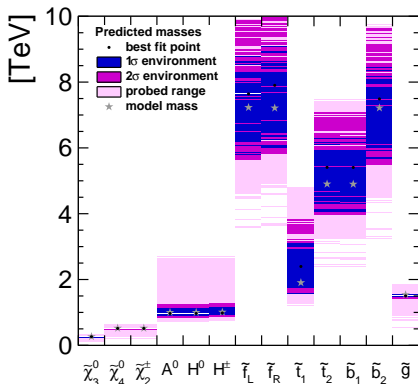
SUSY at ILC: Full reconstruction

- Or, in a case with cascade-decays $\tilde{\chi}_2^0 \rightarrow \tilde{l} \rightarrow l \tilde{\chi}_1^0$: Enough constraints to do full kinematic reconstruction.
- Smuon/Selectron
- Zoom in: Mass to less than 10 MeV
- Even for staus.



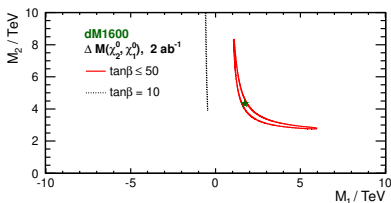
SUSY at ILC: Predict unseen states

- Assume a 10 parameter pMSSM and use constraints from ILC higgs, in addition to measured SUSY properties:
- Can predict the masses of the rest of the spectrum



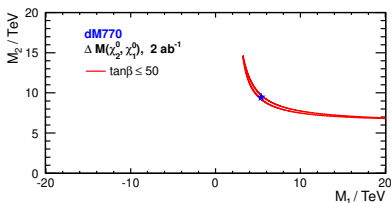
SUSY at ILC: Test models

- Determine constraints on (multi-TeV) M_1 and M_2 from observations of higgsinos only
- Or test M_1 and M_2 unification at the GUT-scale
- ... and from that predict the gluino mass.



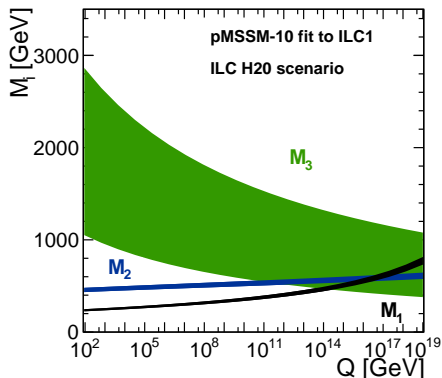
SUSY at ILC: Test models

- Determine constraints on (multi-TeV) M_1 and M_2 from observations of higgsinos only
- Or test M_1 and M_2 unification at the GUT-scale
- ... and from that predict the gluino mass.



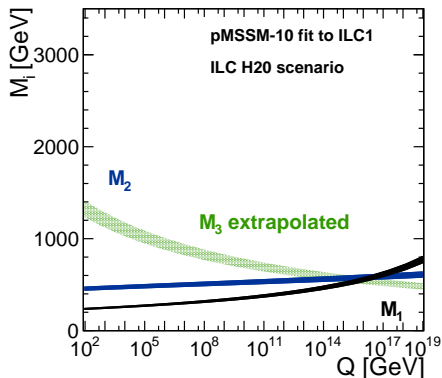
SUSY at ILC: Test models

- Determine constraints on (multi-TeV) M_1 and M_2 from observations of higgsinos only
- Or test M_1 and M_2 unification at the GUT-scale
- ... and from that predict the gluino mass.



SUSY at ILC: Test models

- Determine constraints on (multi-TeV) M_1 and M_2 from observations of higgsinos only
- Or test M_1 and M_2 unification at the GUT-scale
- ... and from that predict the gluino mass.



Conclusions

- **Separate:**
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - discovery potential to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
- Future TeV-scale ee machines have
 - Full discovery **and** exclusion potential up to the kinematic limit
- For the RPC-MSSM, most studies exist, usually with FullSim. Not much to be added for Snowmass. But, little has been done for:
 - **RPV SUSY**.
 - Extensions to MSSM - **nMSSM**, **GMSB**, ...
 - **Squarks**. Might there be holes in squark coverage from HL-LHC?
 - **Long-lived particles**: Much *shorter* lifetimes for long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section (also polarised ones) to the percent level.

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - discovery potential to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
- Future TeV-scale ee machines have
 - Full discovery **and** exclusion potential up to the kinematic limit
- For the RPC-MSSM, most studies exist, usually with FullSim. Not much to be added for Snowmass. But, little has been done for:
 - **RPV SUSY**.
 - Extensions to MSSM - **nMSSM**, **GMSB**, ...
 - **Squarks**. Might there be holes in squark coverage from HL-LHC?
 - **Long-lived particles**: Much *shorter* lifetimes for long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section (also polarised ones) to the percent level.

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - discovery potential to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
- Future TeV-scale ee machines have
 - Full discovery **and** exclusion potential up to the kinematic limit
- For the RPC-MSSM, most studies exist, usually with FullSim. Not much to be added for Snowmass. But, little has been done for:
 - RPV SUSY.
 - Extensions to MSSM - nMSSM, GMSB, ...
 - Squarks. Might there be holes in squark coverage from HL-LHC?
 - Long-lived particles: Much *shorter* lifetimes for long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section (also polarised ones) to the percent level.

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
- Future TeV-scale ee machines have
 - Full **discovery and exclusion potential** up to the kinematic limit
- For the RPC-MSSM, most studies exist, usually with FullSim. Not much to be added for Snowmass. But, little has been done for:
 - **RPV SUSY**.
 - Extensions to MSSM - **nMSSM, GMSB, ...**
 - **Squarks**. Might there be holes in squark coverage from HL-LHC?
 - **Long-lived particles**: Much *shorter* lifetimes for long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section (also polarised ones) to the percent level.

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
- Future TeV-scale ee machines have
 - Full **discovery and exclusion potential** up to the kinematic limit
- For the RPC-MSSM, most studies exist, usually with FullSim. Not much to be added for Snowmass. But, little has been done for:
 - **RPV SUSY**.
 - Extensions to MSSM - **nMSSM, GMSB, ...**
 - **Squarks**. Might there be holes in squark coverage from HL-LHC?
 - **Long-lived particles**: Much *shorter* lifetimes for long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section (also polarised ones) to the percent level.

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
- Future TeV-scale ee machines have
 - Full **discovery and exclusion** potential up to the kinematic limit
- For the RPC-MSSM, most studies exist, usually with FullSim. Not much to be added for Snowmass. But, little has been done for:
 - **RPV SUSY**.
 - Extensions to MSSM - **nMSSM, GMSB, ...**
 - **Squarks**. Might there be holes in squark coverage from HL-LHC?
 - **Long-lived particles**: Much *shorter* lifetimes for long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section (also polarised ones) to the percent level.

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
- Future TeV-scale ee machines have
 - Full **discovery and exclusion** potential up to the kinematic limit
- For the **RPC-MSSM**, most studies exist, usually with FullSim. Not much to be added for Snowmass. But, little has been done for:
 - **RPV SUSY**.
 - Extensions to MSSM - **nMSSM, GMSB, ...**
 - **Squarks**. Might there be holes in squark coverage from HL-LHC?
 - **Long-lived particles**: Much *shorter* lifetimes for long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section (also polarised ones) to the percent level.

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
- Future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.

Take-home message

- Future TeV scale lepton-colliders
 - Full coverage of the parameter space
 - **Without a TeV scale lepton-collider**, we would not be able to exclude SUSY further than today at the end of this century. **LEP2++ would be the final word.**
 - Except if a future pp machine discovers SUSY, which is a **problem we'd like to have!**
 - **Long-lived particles** from other machines running in the same states can be searched for than at pp.
- And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section

Conclusions

- Separate:
 - Discovery potential: Could discover **some** model.
 - Exclusion potential: Can exclude **all** models.
 - Future pp machines have
 - **discovery potential** to very high masses
 - but - to put it bluntly - **NO** exclusion potential: always loopholes.
 - Future TeV lepton colliders
 - Full coverage of the parameter space
 - For the RPV SUSY models, much to be learned
 - **RPV** models
 - Extended SUSY
 - **Squark** masses
 - **Long-lived particles** from either direct or long-lived states can be searched for than at pp.
 - And of course, any **new theory** ideas. Remember: From the ILC, masses can be measured to the sub-percent level, cross-section
- Take-home message**

 - **Without a TeV scale lepton-collider**, we would not be able to exclude SUSY further than today at the end of this century. **LEP2++ would be the final word.**
 - **Except** if a future pp machine discovers SUSY, which is a **problem we'd like to have!**

Thank You !

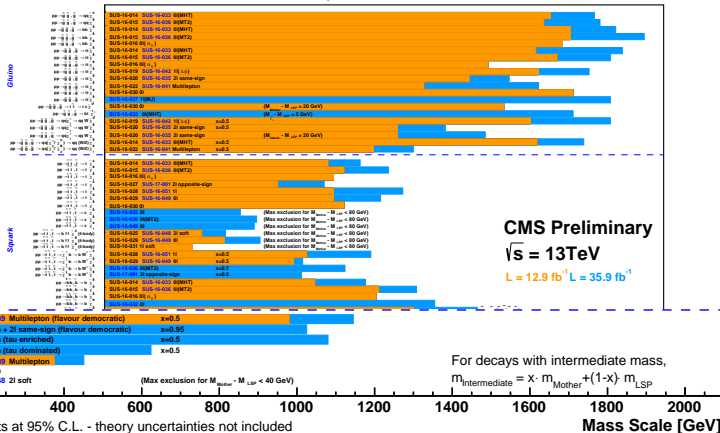
Backup

BACKUP SLIDES

SUSY@LHC: No! Read the fine-print!

Selected CMS SUSY Results* - SMS Interpretation

ICHEP '16 - Moriond '17



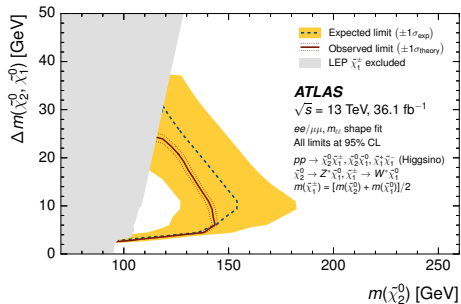
CMS Preliminary
 $\sqrt{s} = 13\text{TeV}$
 $L = 12.9 \text{ fb}^{-1} \quad L = 35.9 \text{ fb}^{-1}$

*Observed limits at 95% C.L. - theory uncertainties not included

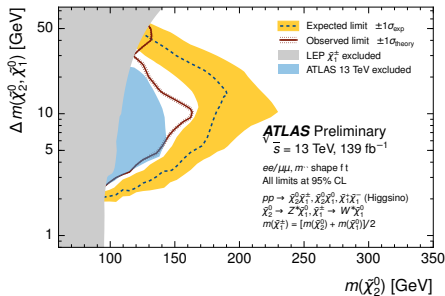
Only a selection of available mass limits. Probe *up* to the quoted mass limit for $m_{\text{LSP}} = 0 \text{ GeV}$ unless stated otherwise

Latest Atlas (13 TeV, 36 and 139 fb⁻¹) on higgsinos

arXiv:1803.02762

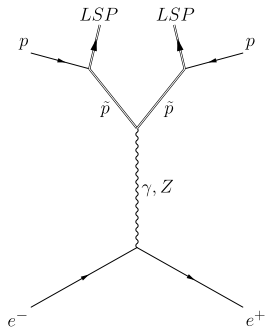


ATLAS-CONF-2019-01



Loop-hole free SUSY searches

- All is **known** for given masses, due to SUSY-principle: “sparticles couples as particles”.
- This doesn't depend on the SUSY breaking mechanism !
- Obviously: There is **one** NLSP.

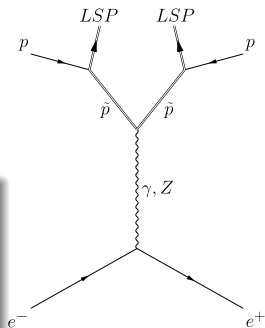


Loop-hole free SUSY searches

- All is **known** for given masses, due to SUSY-principle: “sparticles couples as particles”.
- This doesn't depend on the SUSY breaking mechanism !
- Obviously: There is **one** NLSP.

So, at an LC :

- Model **independent** exclusion/ discovery reach in $M_{NLSP} - M_{LSP}$ plane.
- Repeat for **all** NLSP:s.
- **Cover entire parameter-space in a hand-full of plots**
- NLSP search \leftrightarrow “simplified models” @ LHC!

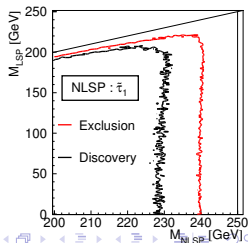
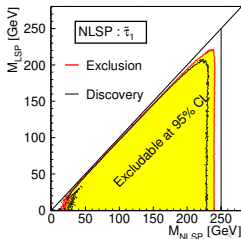
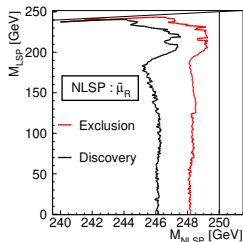
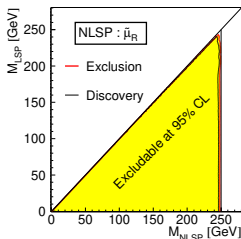


Simplified models

- Simplified methods at hadron and lepton machines are **different beasts**.
- At lepton machines they are quite **model independent**, at LHC **model dependent**.
- A few examples (M.B. arXiv:1308.1461)
 - $\tilde{\mu}_R$ NLSP
 - $\tilde{\tau}_1$ NLSP (minimal σ).

Simplified models

- Simplified methods at hadron and lepton machines are **different beasts**.
- At lepton machines they are quite **model independent**, at LHC **model dependent**.
- A few examples (M.B. arXiv:1308.1461)
 - $\tilde{\mu}_R$ NLSP
 - $\tilde{\tau}_1$ NLSP (minimal σ).



Simplified models

- Simplified methods at hadron and lepton machines are **different beasts**.

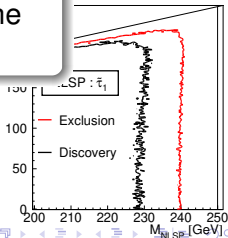
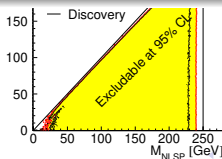
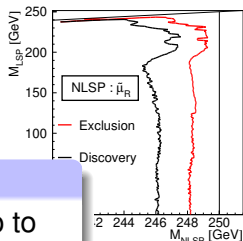
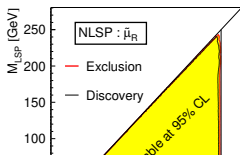
- At lepton machines they are called **At ILC**

independent Both discover and exclude NLSPs up to model dependent **some GeV**:s from the kinematic limit,

- A few examples whatever the NLSP is, and whatever the rest of the spectrum is!

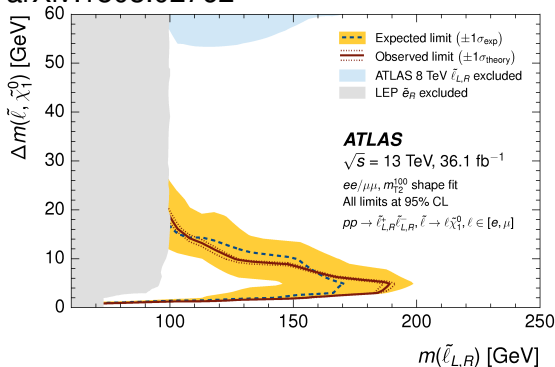
arXiv:1308.1461)

- $\tilde{\mu}_R$ NLSP.
- $\tilde{\tau}_1$ NLSP (minimal σ).

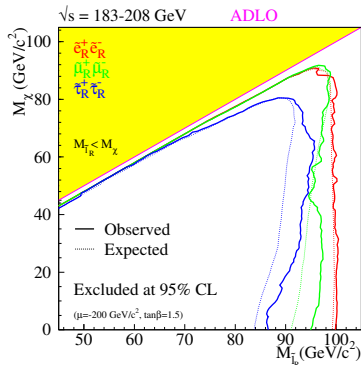


Latest Atlas (13 TeV, 36 fb⁻¹) and LEP on sleptons

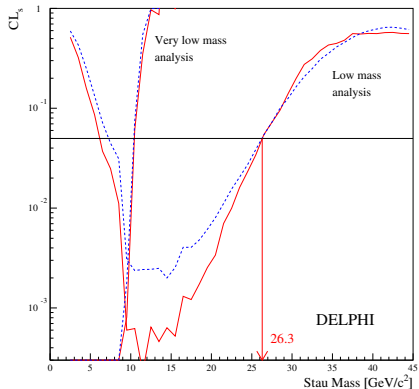
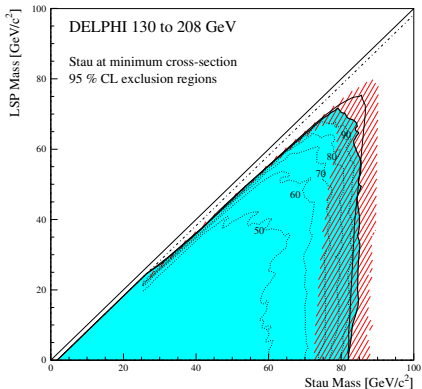
arXiv:1803.02762



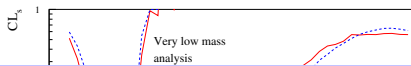
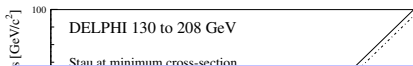
This is a *combined* limit, assuming $\tilde{\mu}_L, \tilde{\mu}_R, \tilde{e}_L$ and \tilde{e}_L all have the **same mass** !!!



This is $\tilde{e}_R, \tilde{\mu}_R$ and $\tilde{\tau}_R$ *only*, separately!

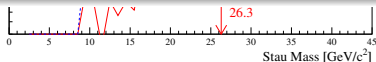
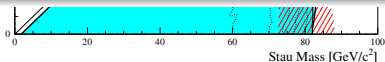
In real life: LEP $\tilde{\tau}$ limits

NB: a $\tilde{\tau}$ as light as 26.3 GeV is *not* excluded!

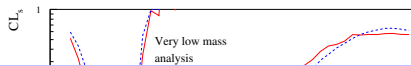
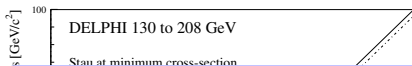
In real life: LEP $\tilde{\tau}$ limits

With 1000 times the luminosity and no trigger, the ILC at 250 will push the limits for all possible NLSPs to close to 125 GeV, and $\Delta(M) \approx 0$. The area covered will \sim double the LEP ones. They are in the most compelling region of parameter-space.

- These will be rock-solid limits.
- Or discoveries!

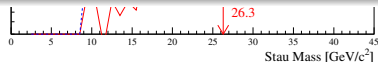
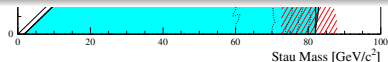


NB: a $\tilde{\tau}$ as light as 26.3 GeV is *not* excluded!

In real life: LEP $\tilde{\tau}$ limits

With 1000 times the luminosity and no trigger, the ILC at 250 will push the limits for all possible NLSPs to close to 125 GeV, and $\Delta(M) \approx 0$. The area covered will \sim double the LEP ones. They are in the most compelling region of parameter-space.

- These will be rock-solid limits.
- Or discoveries!



NB: a $\tilde{\tau}$ as light as 26.3 GeV is *not* excluded!