

# Snowmass 2021 - EF9/10+



THE OHIO STATE  
UNIVERSITY



## **Scintillator Array Probes of the Dark Sector**

Christopher S. Hill  
*The Ohio State University*

# Who? Why?

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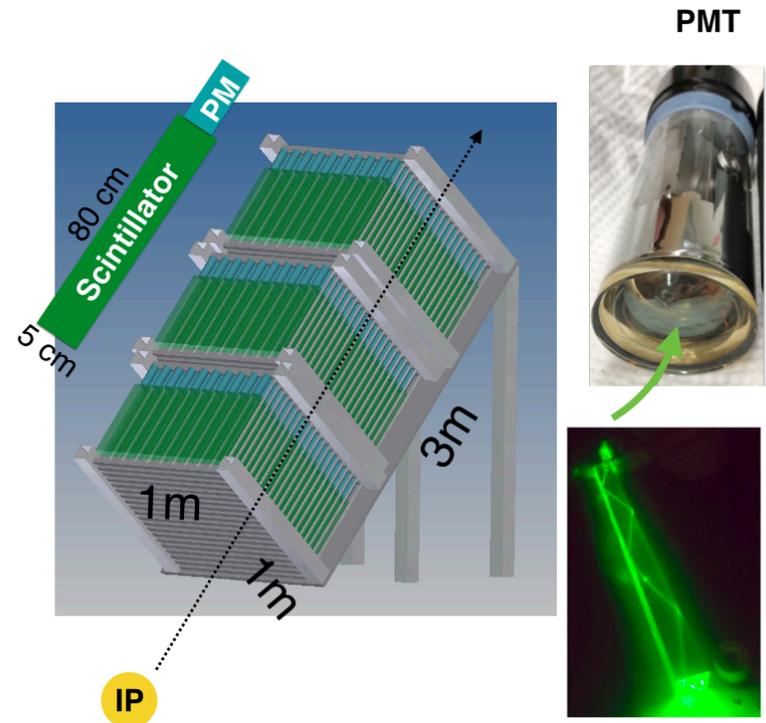
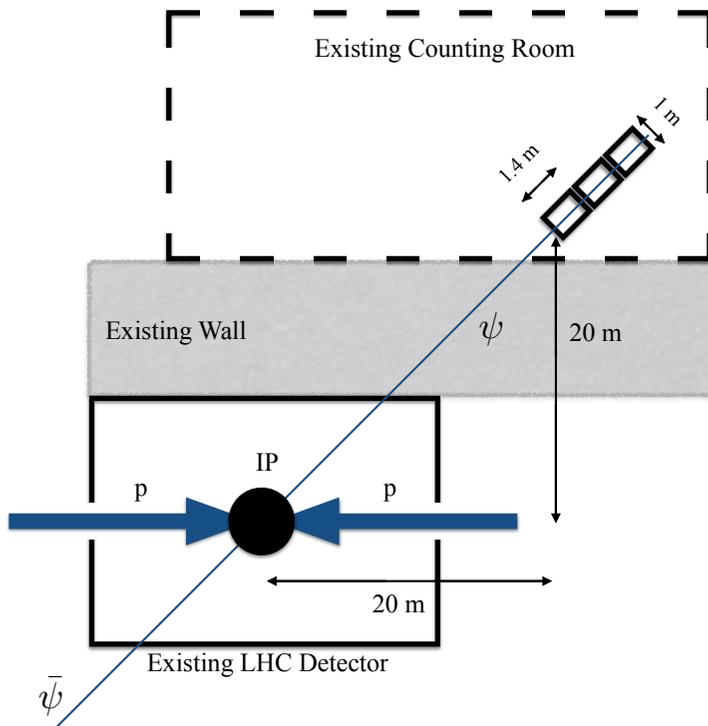
- I am giving this talk on behalf of several people who have submitted similar EOIs and who have self-organized into a group that is planning to submit an LOI to write a whitepaper. Currently we are:
  - *Chris Hill, Brian Francis (OSU)*
  - *Albert De Roeck (CERN)*
  - *Andy Haas (NYU)*
  - *David Stuart, Matthew Citron (UCSB)*
  - *Yu-Dai Tsai (FNAL)*
  - *Jae Hyeok Yoo (Korea U.)*
- **We are all involved in one or more “scintillator array” experiments**
- **In short, for Snowmass we intend to produce an updated, consistent (apples-to-apples) comparison of these experiments to each other (and other exp. where relevant/possible)**
  - *This talk similar to one Matthew gave in previous EF session,*

Others are welcome to join!

# What?

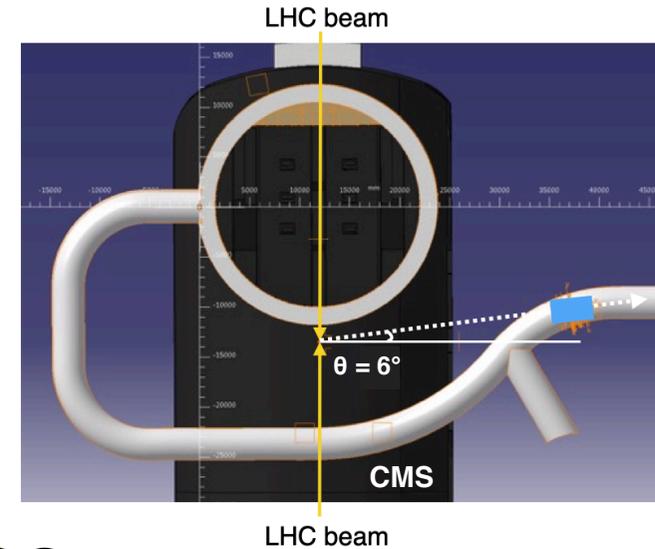
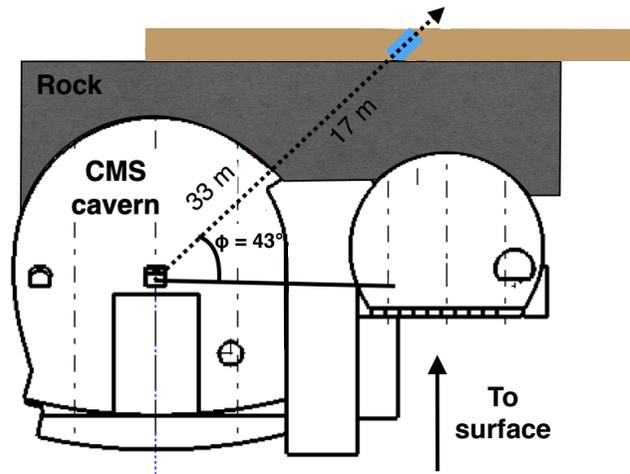
- By scintillator array experiments, we mean something like that which Andy and I advocated in *Phys.Lett.B* 746 (2015) 117-120
  - *Dark sector probe using distinctive signature*
    - **“millicharged” particles**

- Inexpensive dedicated detector
- With  $Q$  down to  $\sim 10^{-3}e$ ,  $dE/dx$  is  $10^{-6}$  MIP  $\rightarrow$  need long, sensitive, active length to see signal,  $\mathcal{O}(1)$  PE.
- $\sim 1\text{ m} \times 1\text{ m} \times 3\text{ m}$  scintillator + PMT array, pointing back to IP, in well shielded area near CMS/ATLAS
- With triple coincidence, thought expected dominant random background would be controlled



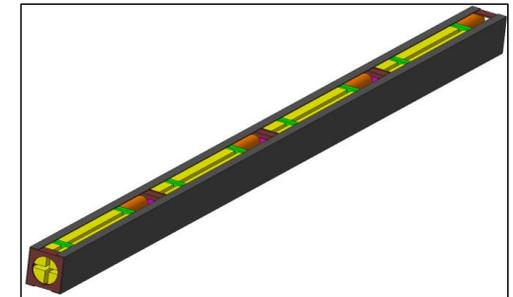
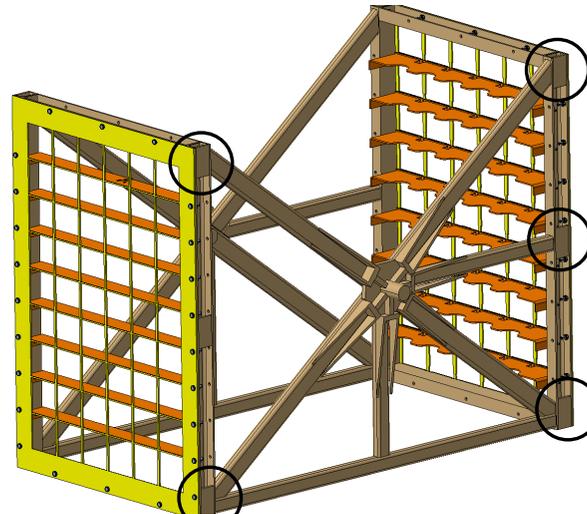
# What? (cont.)

- This paper soon led to the formulation of the **milliQan collaboration**, with proposed detector sited at P5 (near CMS)



## milliQan

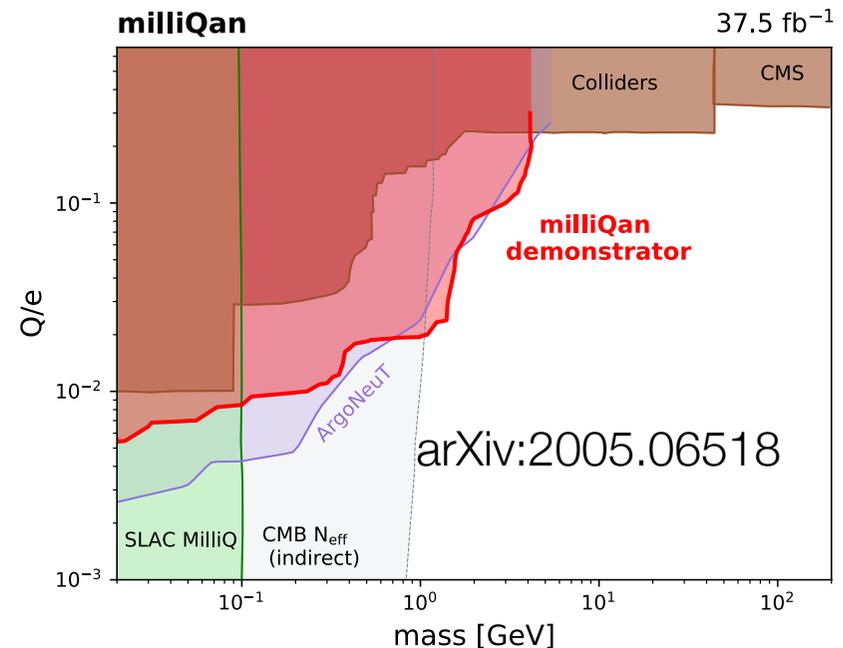
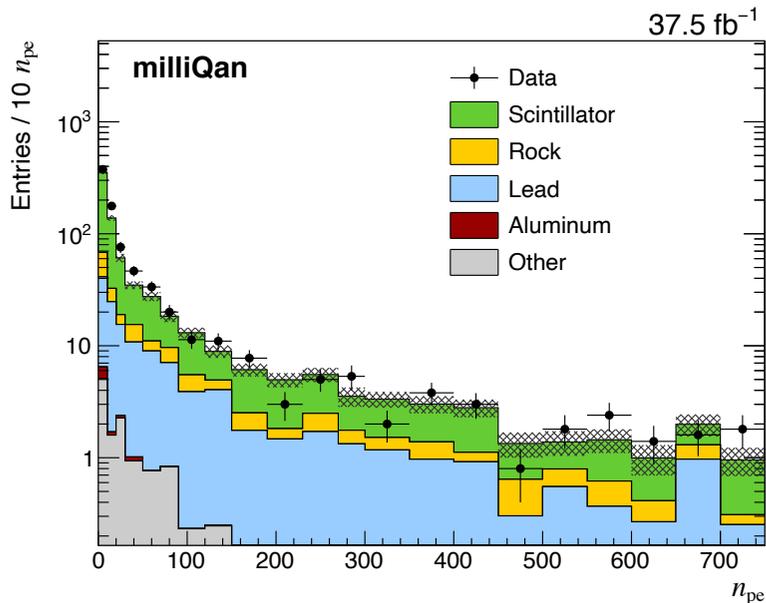
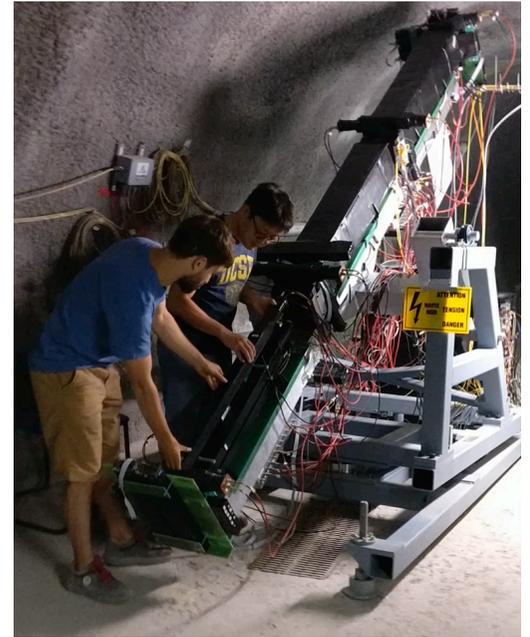
arXiv:1607.04669



- Experimental design at an **advanced stage** of experimental design (validated. full simulation, in-situ experience with alignment, calibrations, backgrounds, fully engineered mechanics/electronics) ... **only need \$**

# Physics obtained from 1% milliQan demonstrator!

- One reason at advanced stage is we built, installed, and operated a 1% prototype
- “Physics” run in 2018, collecting  $\sim 35/\text{fb}$ , 2000h of data
  - *Operational experience (triggering/DAQ/DQM)*
  - *Used for range of studies to prove feasibility of full detector: alignment, calibrations, background measurements*
    - **e.g. modified full milliQan design to have 4-fold coincidence**
  - *Fully simulated in GEANT4 (validated with data)*
    - **Able to produce competitive limits with only 1%**

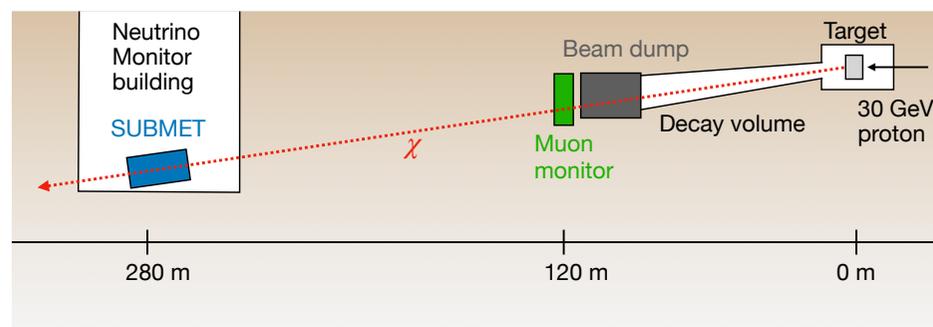
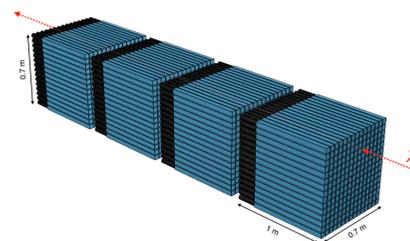


# New milliQan-like proposals

- Success of this simple scintillator approach has (more recently) led to proposals to install milliQan-like detectors at other accelerator complexes
  - *FerMINI at FNAL*
  - *SUBMET at J-PARC*

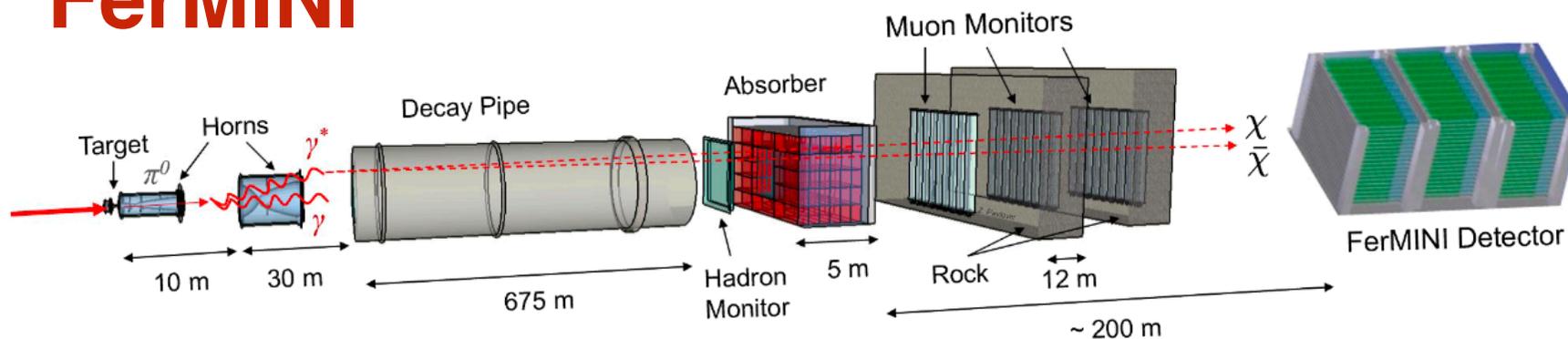
## SUBMET

arXiv:2007.06329



**J-PARC** with sensitivity for  $m < \sim 1.5$  GeV

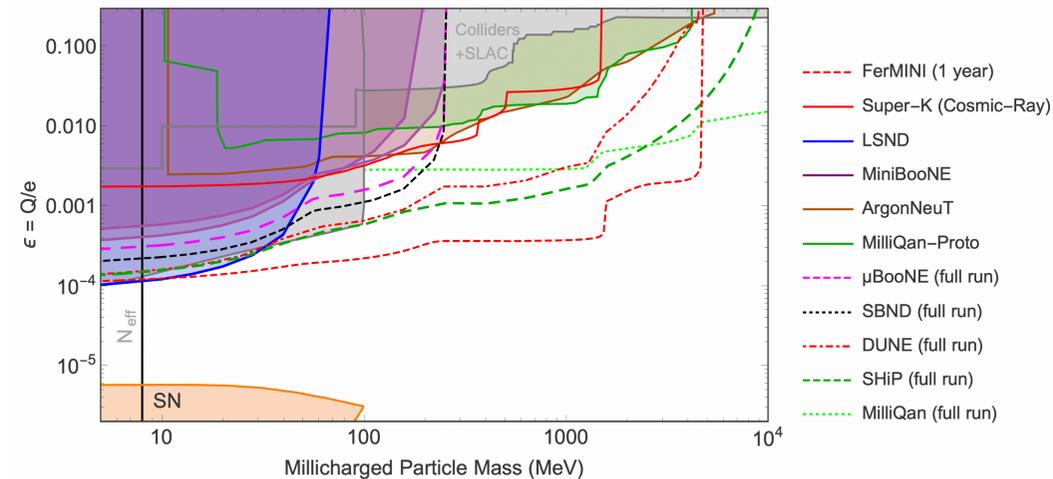
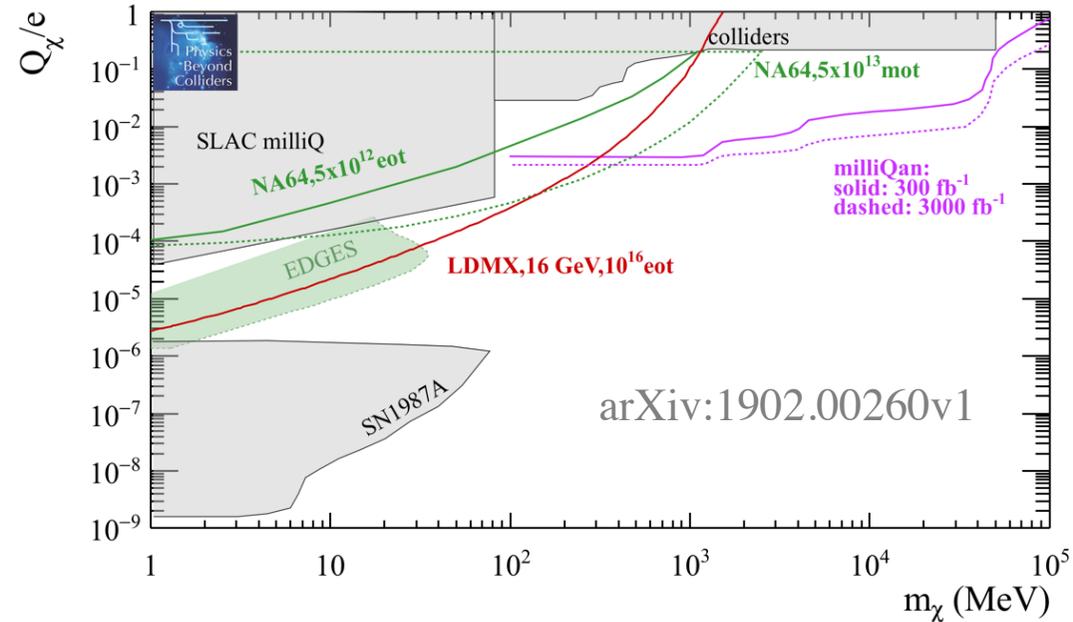
## FerMINI



**Fermilab** with sensitivity for  $m < \sim 5$  GeV

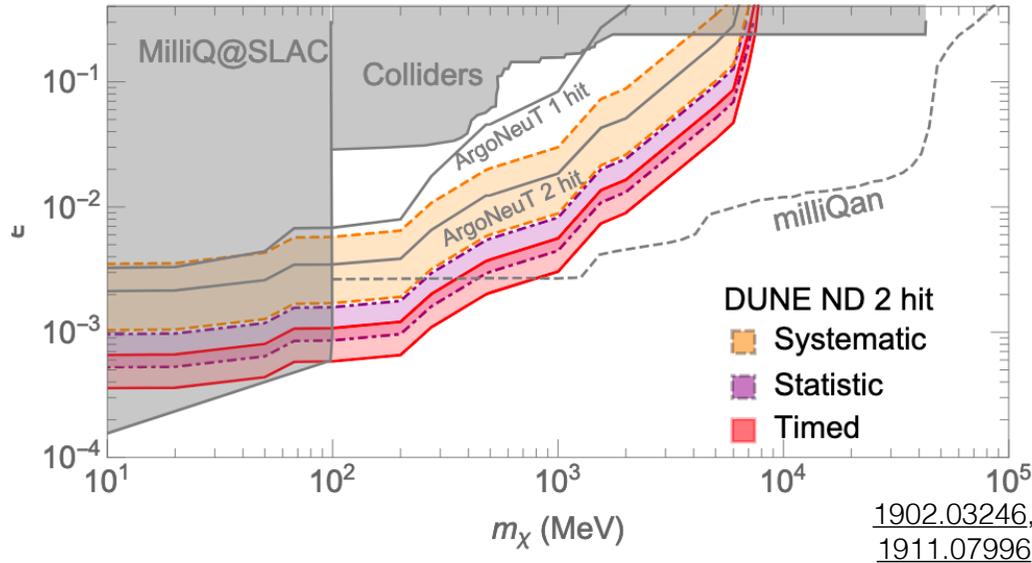
# Snowmass Plans

- Pick up where CERN PBC report (and others) have left off
- Use data and/or lessons learned from demonstrator to update sensitivity projections with “realism” for scintillation based detectors at range of facilities: LHC, J-PARC, FERMILAB
- *Characterization/ understanding of backgrounds (and how to reject them)*
- *Guide design of future detectors based on experience*
- *Add other experiments where relevant? If more expert collaborators join us ...*

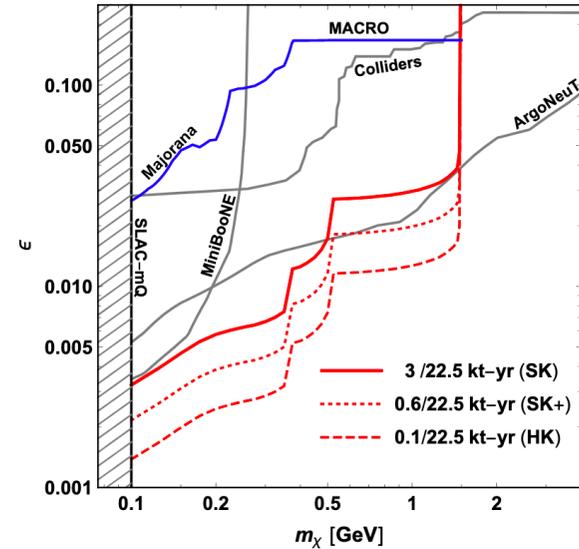


# Additional Material

# Searching for millicharged particles with other experiments

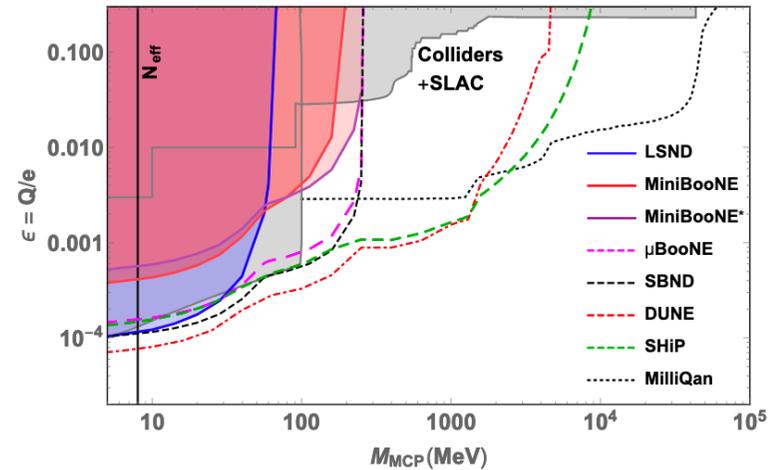


Detection with LarTPC



Production in cosmic ray showers

- Neutrino detectors provide sensitivity to millicharged particles through hard ( $> \text{MeV}$  recoil) electron scattering
- Many recent results!
- Would be interested in collaborating to expand scope of project: **full consideration of all complementary methods to search for mCPs!**



Fixed source neutrino experiments

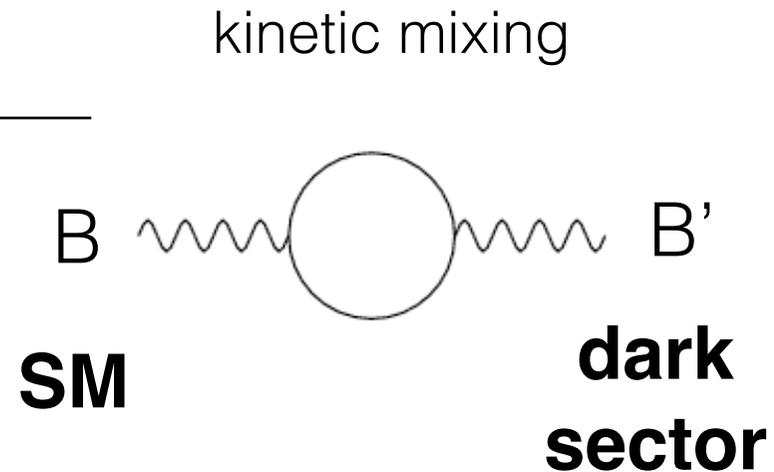
# Why milli-charged?

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{4} B'_{\mu\nu} B'^{\mu\nu} - \frac{\kappa}{2} B'_{\mu\nu} B^{\mu\nu}$$

↙ ↘ ←

massless U'(1) boson in the dark sector

'dark EM'



$$\kappa \sim 10^{-3} - 10^{-2}$$

(naturally  $\sim \alpha/\pi$ )

Kinetic mixing with a new massless 'dark' boson **can provide link between SM and a hidden/dark sector**

# Why milli-charged?

Now add fermion charged under new  $U'(1)$ :

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{4} B'_{\mu\nu} B'^{\mu\nu} - \frac{\kappa}{2} B'_{\mu\nu} B^{\mu\nu} + i\bar{\psi}(\not{\partial} + ie'\not{B}' + iM_{mCP})\psi$$

Standard trick - redefine gauge field  $B'$ :  $B' \rightarrow B' + \kappa B$

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{4} B'_{\mu\nu} B'^{\mu\nu} + i\bar{\psi}(\not{\partial} + \underbrace{ike'\not{B}} + ie'\not{B}' + iM_{mCP})\psi$$

new fermion has small EM charge: **milli-charged particle**