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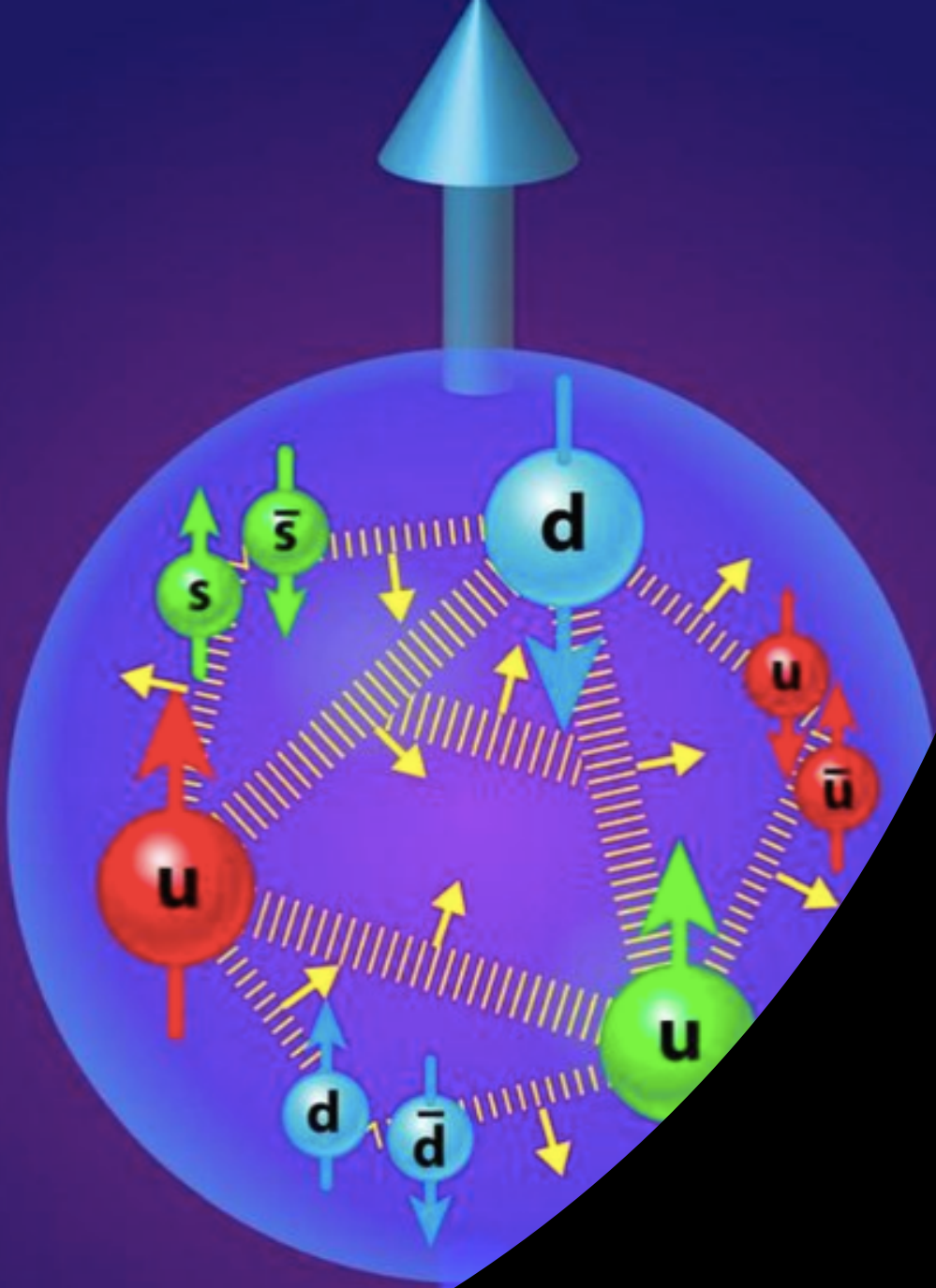
**Capturing the Image of Quarks and Gluons in the Nucleon  
with the JLab to LHC Large Data Array**

Snowmass EF06  
August 19, 2020



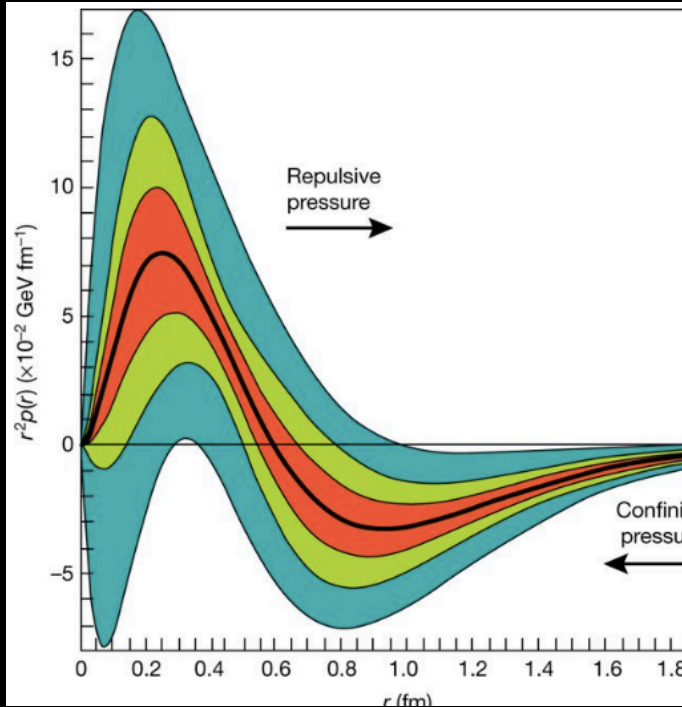
# GPDs and Deeply Virtual Exclusive Experiments

A new paradigm that will allow us to both penetrate and visualize the deep structure of visible matter, answering questions that we couldn't even afford asking before



How does the proton/neutron get its mass and spin and how do we test this dynamics?





“The average peak pressure near the center is about  $10^{35}$  pascals, which exceeds the pressure estimated for the most densely packed known objects in the Universe, neutron stars”

Burkert, Elouadrhiri, Girod, *Nature* 557, 396 (2018)

How do we measure pressure and forces inside the proton?



# The QCD Energy Momentum Tensor

Mass

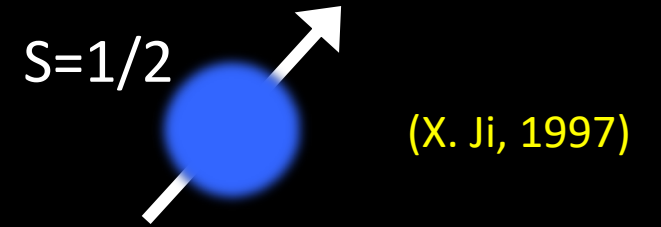
Momentum density

	$\frac{E^2 + B^2}{2}$	$S_x$	$S_y$	$S_z$
$S_x$	$\sigma_{xx}$	$\sigma_{xy}$	$\sigma_{xz}$	
$S_y$	$\sigma_{yx}$	$\sigma_{yy}$	$\sigma_{yz}$	
$S_z$	$\sigma_{zx}$	$\sigma_{zy}$	$\sigma_{zz}$	

Shear stress

Pressure

$\vec{S} = \vec{E} \times \vec{B}$



$$\begin{aligned}
 \langle p', \Lambda | T^{\mu\nu} | p, \Lambda \rangle = & A(t) \bar{U}(p', \Lambda') [\gamma^\mu P^\nu + \gamma^\nu P^\mu] U(p, \Lambda) + B(t) \bar{U}(p', \Lambda') i \frac{\sigma^{\mu(\nu} \Delta^{\nu)} }{2M} U(p, \Lambda) \\
 & + C(t) [\Delta^2 g^{\mu\nu} - \Delta^{\mu\nu}] \bar{U}(p', \Lambda') U(p, \Lambda) + \tilde{C}(t) g^{\mu\nu} \bar{U}(p', \Lambda') U(p, \Lambda)
 \end{aligned}$$

off-forward

q and g not separately conserved

$$\left\{ \begin{aligned}
 P &= \frac{p+p'}{2} \\
 \Delta &= p' - p = q - q' \\
 t &= (p-p')^2 = \Delta^2
 \end{aligned} \right.$$



Direct calculation of EMT form factors

Donoghue et al. PLB529 (2002),

C. Corianò et al. PRD(2018),

A. Freese , QCD Evolution 2019

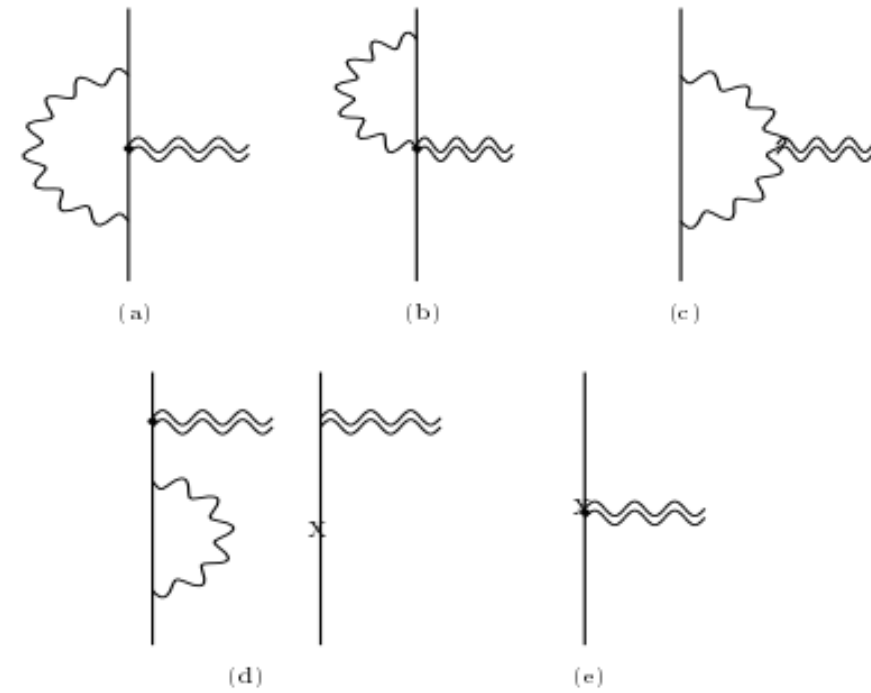
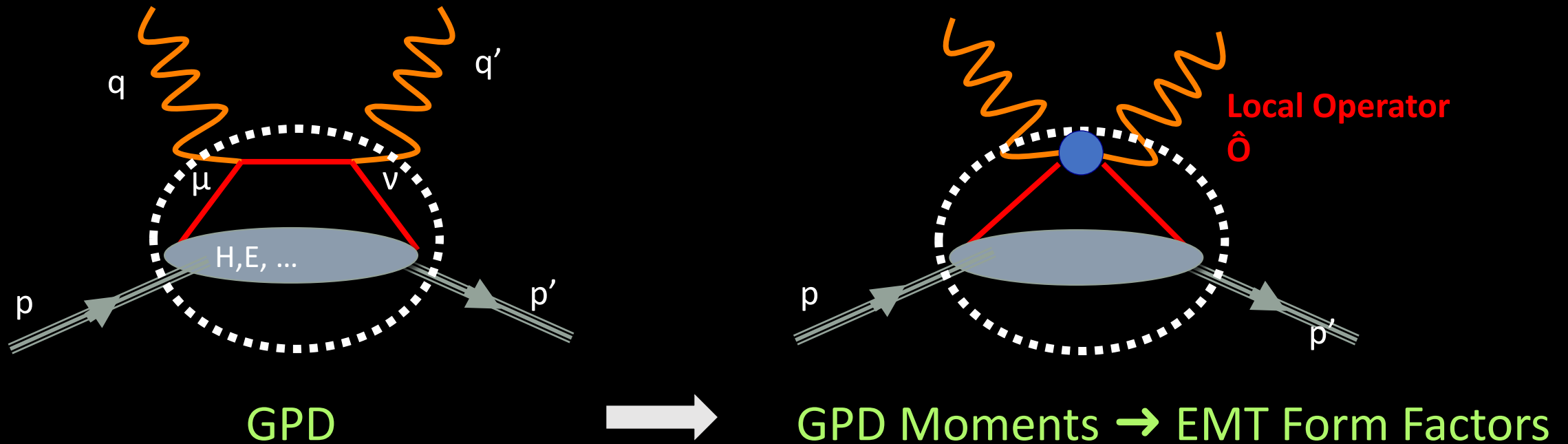


Figure 2: Feynman diagrams for spin 1/2 radiative corrections to  $T_{\mu\nu}$ .

# EMT matrix elements from Generalized Parton Distributions Moments (X.Ji,1997)



- Large momentum transfer  $Q^2 \gg M^2 \rightarrow$  “deep”
- Large Invariant Mass  $W^2 \gg M^2 \rightarrow$  equivalent to an “inelastic” process



## Physical interpretation of EMT form factors

$$\frac{1}{2} (A_q + B_q) = J_q = \frac{1}{2} (A_{20} + B_{20})$$

$$J_q^i = \int d^3r \epsilon^{ijk} r_j T_{0k}$$

Angular Momentum

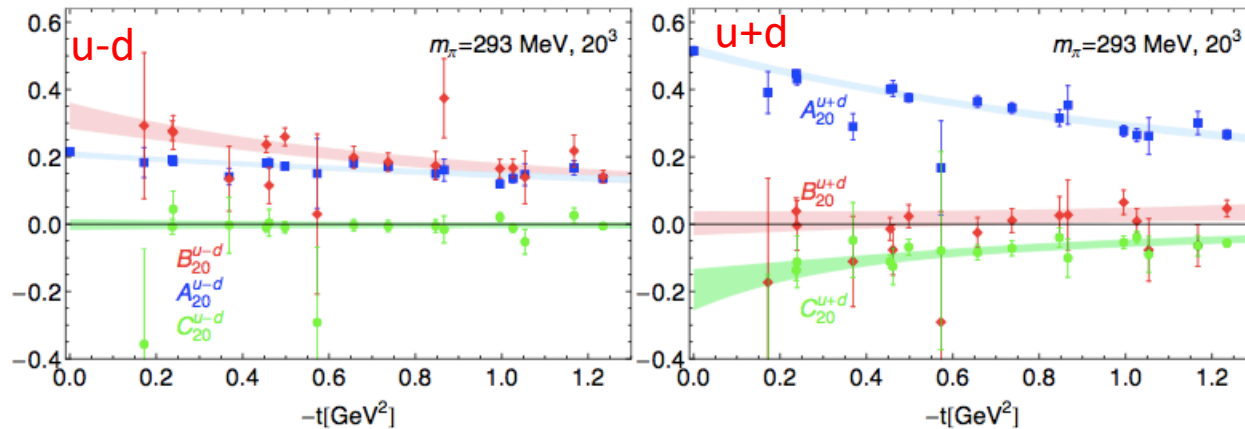
$$A_q = \langle x_q \rangle = A_{20}$$

Momentum

$$C_q = \text{Internal Forces} = C_{20}$$

$$\int d^3r (r^i r^j - \delta^{ij} r^2) T_{ij}$$

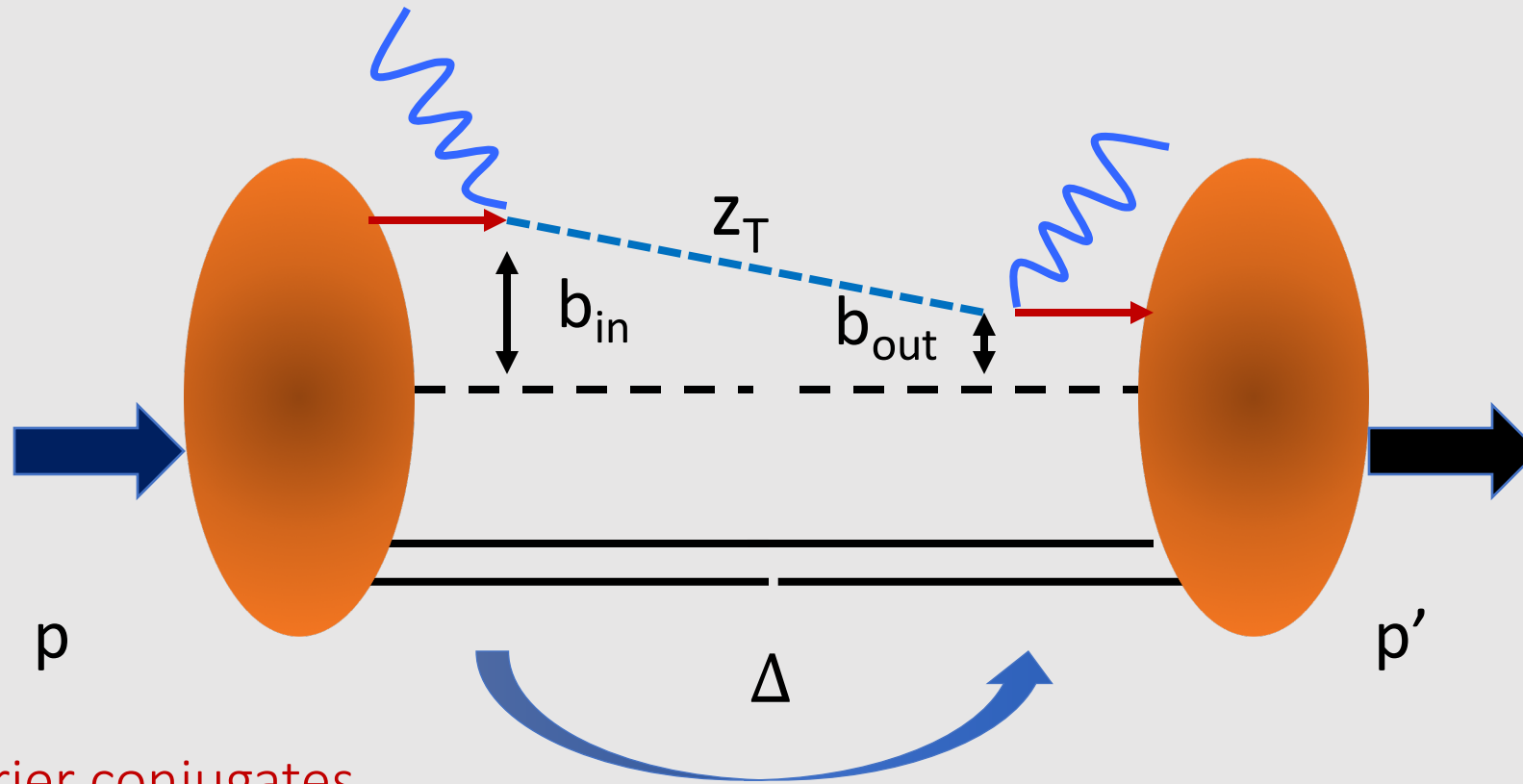
Pressure



Ph. Haegler, JoP: **295** (2011) 012009

More recent results in review paper  
Huey Wen Lin et al., arXiv [2006.08636](https://arxiv.org/abs/2006.08636)

# Accessing transverse distances through Fourier transformation



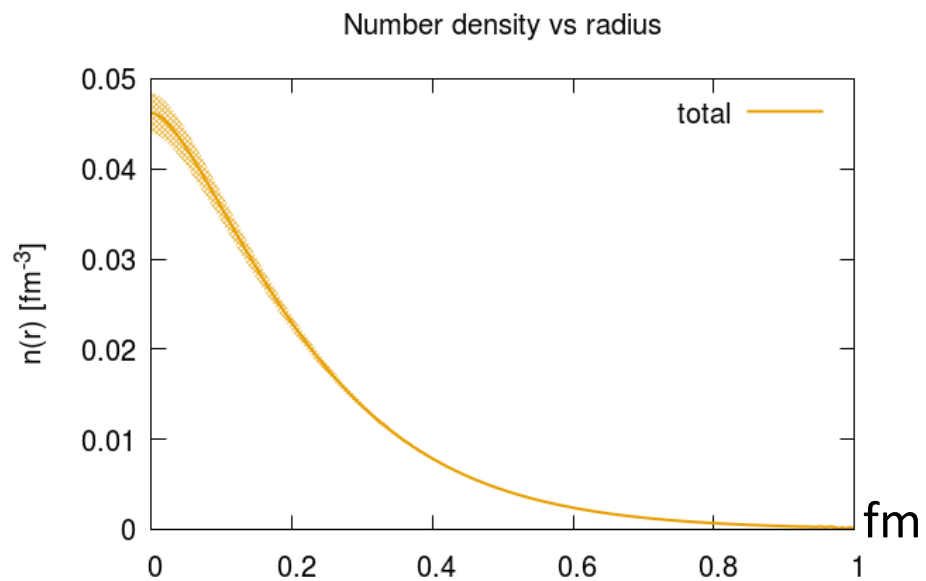
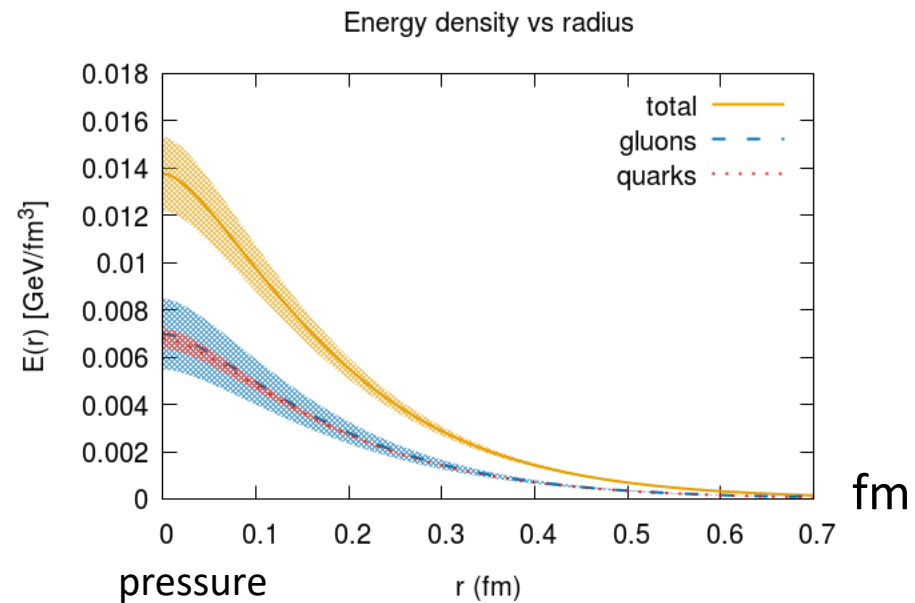
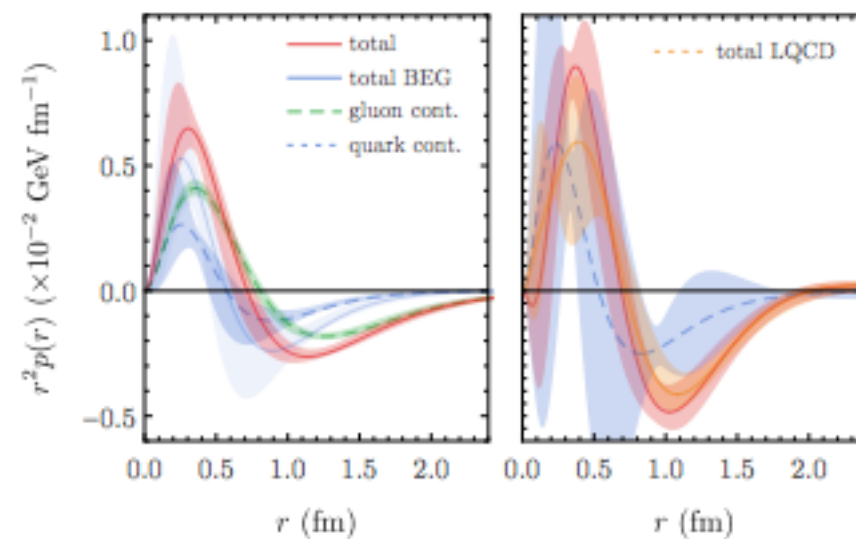
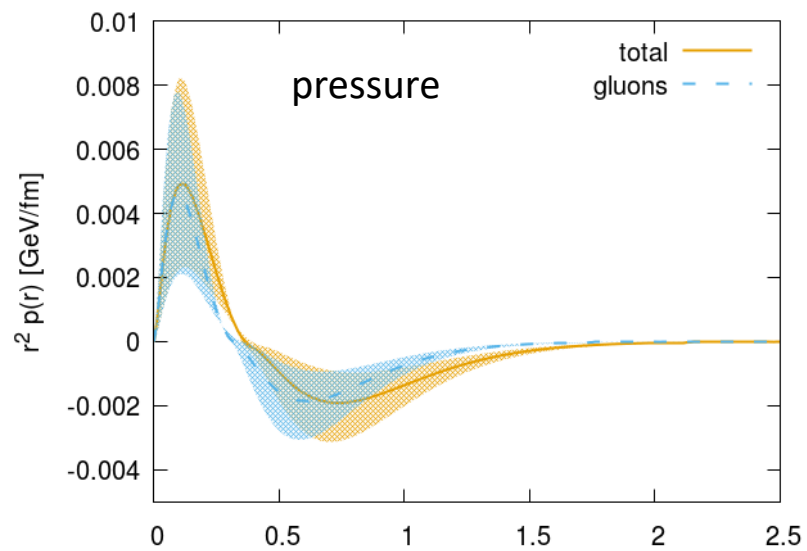
Two types of Fourier conjugates

$$\left\{ \begin{array}{l} b = \frac{b_{in} + b_{out}}{2} \\ \Delta = p - p' \end{array} \right.$$

$$\left\{ \begin{array}{l} z_T = b_{in} - b_{out} \\ k_T \end{array} \right.$$

Related to Ioffe "time"

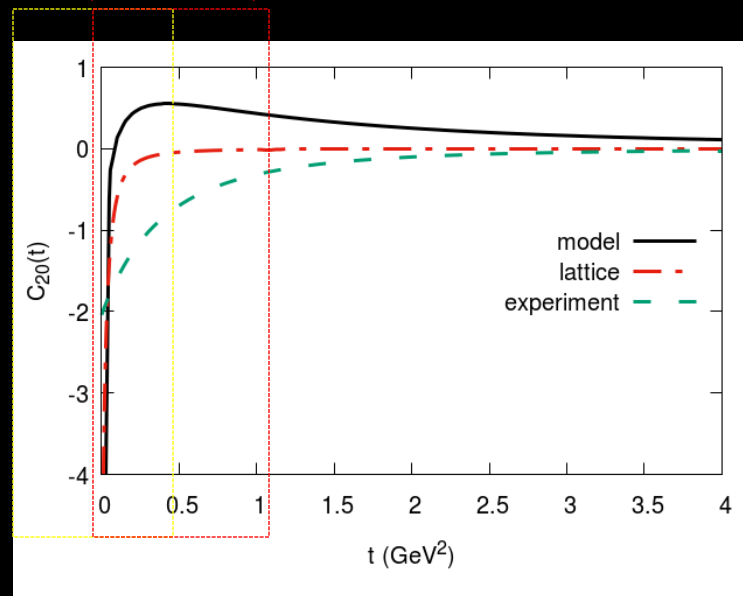


$A_{10}$  $A_{20}$  $C_{20}$ 

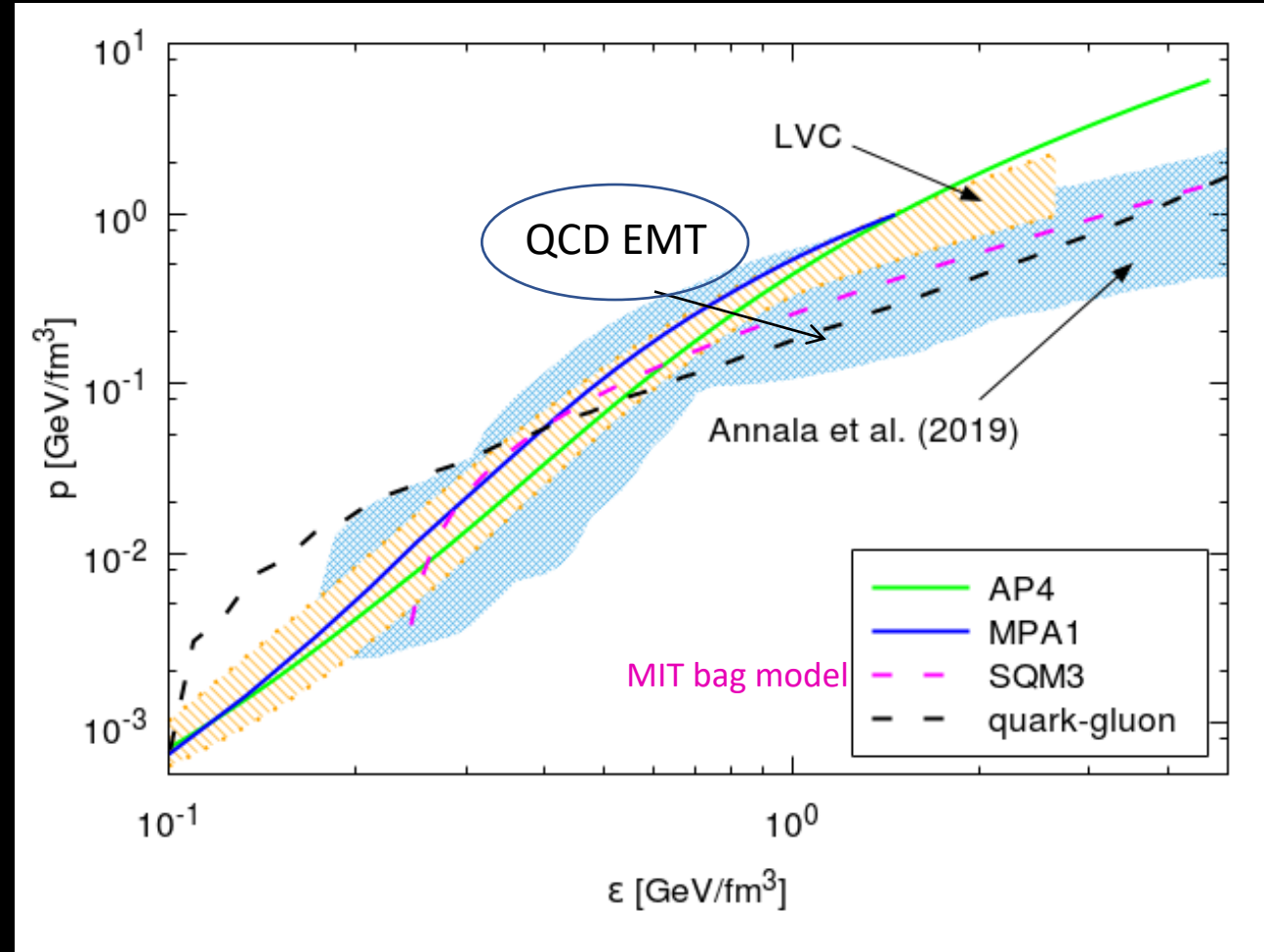
$C_{20}$

experimental range

Lattice



The Jlab extraction of the pressure measurement helps us understand the EoS of neutron stars

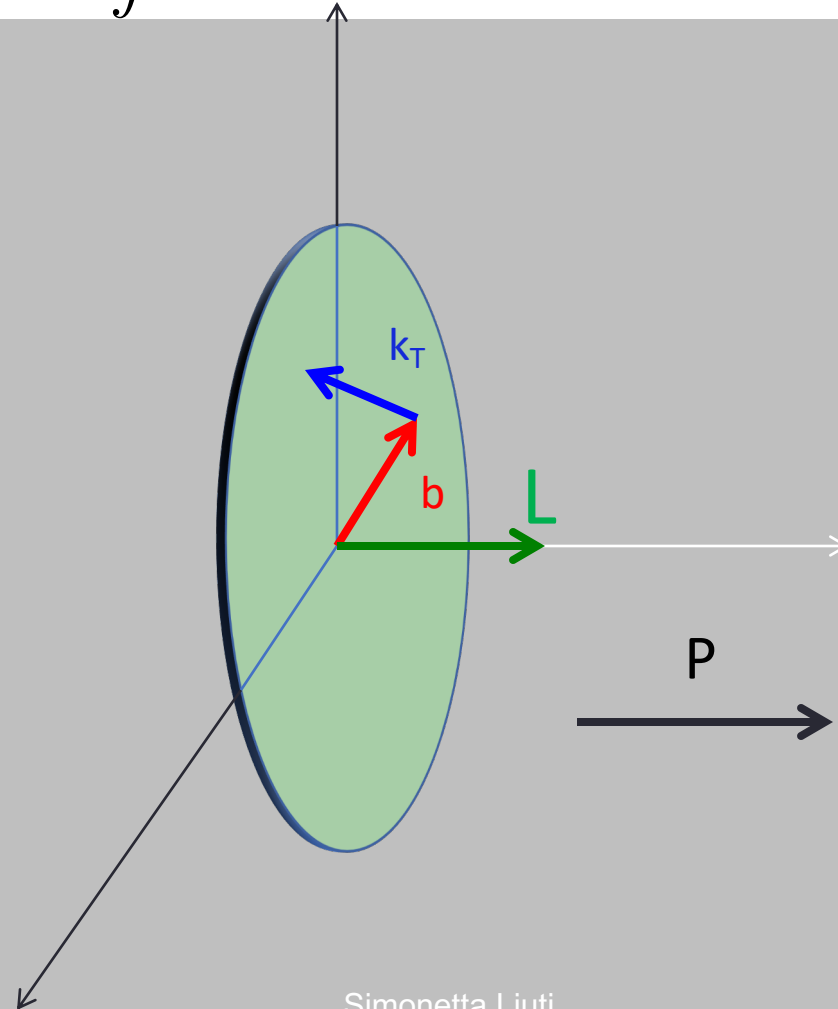


A. Rajan, T. Gorda, SL, K. Yagi, submitted for publication

# Orbital angular momentum definition through Wigner Distributions

$$L_q^{\mathcal{U}} = \int dx \int d^2 \mathbf{k}_T \int d^2 \mathbf{b} (\mathbf{b} \times \mathbf{k}_T)_z \mathcal{W}^{\mathcal{U}}(x, \mathbf{k}_T, \mathbf{b})$$

Hatta  
Burkardt  
Lorce, Pasquini,  
Ji, Xiong, Yuan  
Mukherjee,  
Courtoy, Engelhardt, Rajan, SL



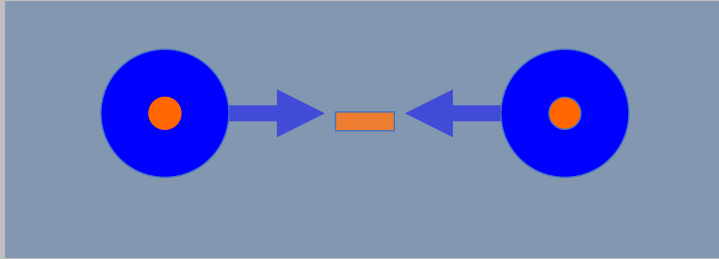


# Angular Momentum Sum Rule

A. Rajan, A. Courtoy, M. Engelhardt, S.L., PRD (2016)

A. Rajan, M. Engelhardt, S.L., PRD (2018)

$$J_q = L_q + \frac{1}{2} \Delta \Sigma_q$$



Beam Target Spin Correlation: unpolarized quark density in a longitudinally polarized proton

Twist – 3 GPD

$$\tilde{E}_{2T} = - \int_x^1 \frac{dy}{y} (H + E) + \left[ \frac{\tilde{H}}{x} - \int_x^1 \frac{dy}{y^2} \tilde{H} \right] + \left[ \frac{1}{x} \mathcal{M}_{F_{14}} - \int_x^1 \frac{dy}{y^2} \mathcal{M}_{F_{14}} \right]$$

New! From DVCS

DVCS (Ji,'97)

Polarized ep  
 $g_1(x)$

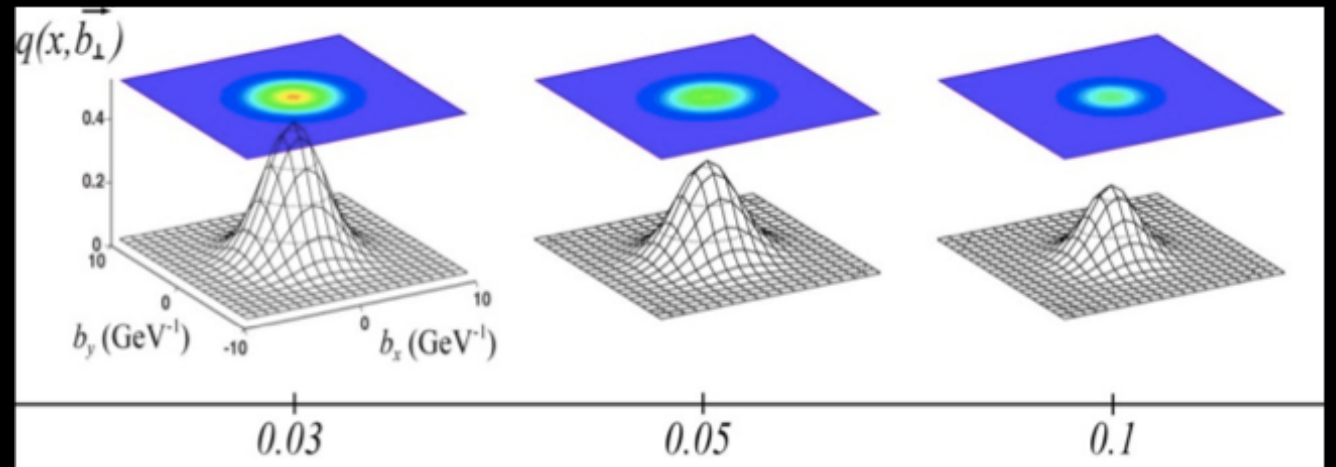
Burkardt  
Hatta

Integrated in x

Color force/gauge link

$$L = J - S + 0$$

# Measuring the Nucleon Gravitomagnetic Form Factors



graph from M. Defurne

# A multi-step, multi-prong process that compares to imaging a black hole

Images courtesy of Kent Yagi, UVA

**Event Horizon Telescope Array**  
To get a good look at the light show coming from our galaxy's black hole, astronomers will combine the data from telescopes the world over. Here's a sample of the dozen telescopes that may one day be part of the Event Horizon Telescope.

**CSO**  
The Caltech Submillimeter Observatory  
10.4-meter telescope on Mauna Kea, Hawaii



**CARMA**  
The Combined Array for Research in Millimeter-Wave Astronomy  
15 antennae near Bishop, Calif.



**ARO/SMT**  
The Arizona Radio Observatory's Submillimeter Telescope  
10-meter telescope near Safford, Ariz.



**IRAM 30M**  
The Instituto for Radio Astronomy in the millimeter range's 30M scope  
30-meter telescope on Pico Veleta, Spain



**JCMT**  
The James Clerk Maxwell Telescope  
15-meter telescope on Mauna Kea, Hawaii



**SMA**  
The Submillimeter Array  
8 antennae on Mauna Kea, Hawaii



**ALMA**  
The Atacama Large Millimeter/sub-millimeter Array  
66 antennae on the Chajnantor plain of Chile



**APEX**  
Atacama Pathfinder Experiment  
12-meter telescope on the Chajnantor plain of Chile

## Event Horizon Telescope (EHT)

## In a nutshell.....

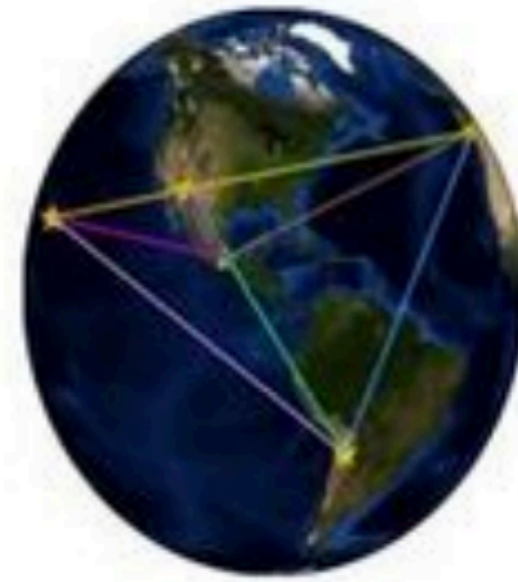
$$\text{Telescope size} = \frac{\text{wavelength}}{\text{dimension}} \quad \longrightarrow \quad \text{Size of Earth} = \frac{\text{wavelength}}{40 \mu \text{ arcsec}} \quad \longrightarrow \quad \text{wavelength} = 1\text{mm}$$

- Detector: a network of telescopes (EHT)
- Imaging technique: Very Long Baseline Interferometry (VLBI)
- Computational processing/simulation/validation of the derived image

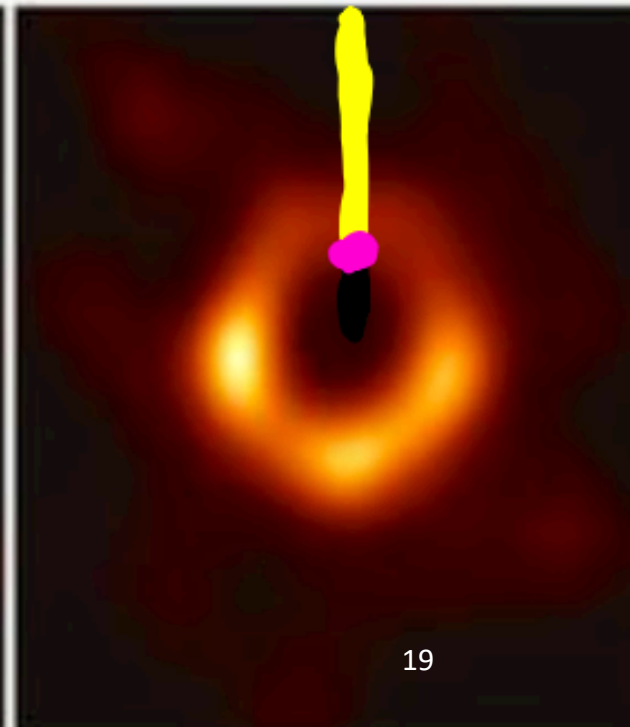
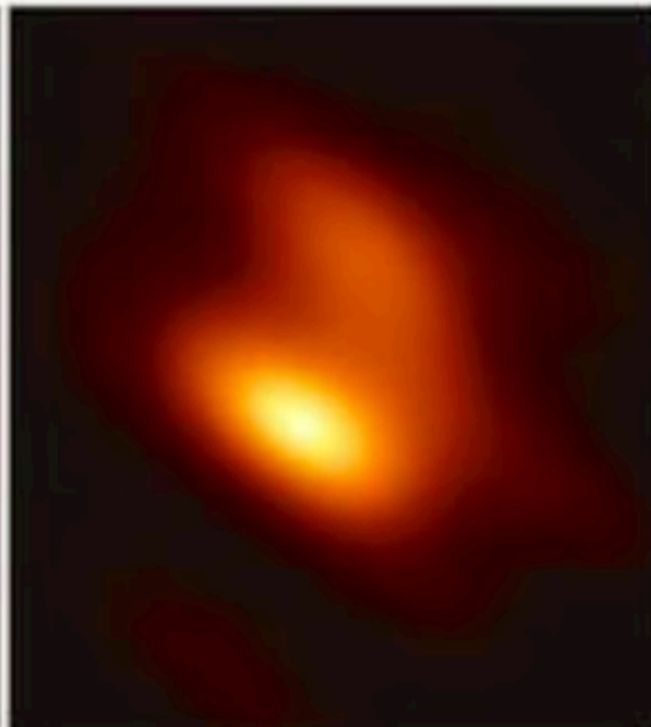
S. Doeleman et al., APJL Focus issue, 2019

[https://iopscience.iop.org/journal/2041-8205/page/Focus\\_on\\_EHT](https://iopscience.iop.org/journal/2041-8205/page/Focus_on_EHT)





8/18/20



19

M87\*



## Jefferson Lab@12 GeV

### Event Horizon Telescope (EHT)

- ✓ **Main idea:** Very Long Baseline Interferometry (VLBI), an array of smaller telescopes synchronized to focus on the same object and act as a giant telescope
- ✓ **Precision:** large aperture (many telescopes widely spaced) and high frequency radio waves
- ✓ **Data Management:** 5 petabytes physically transported to a central location. Data from all eight sites were combined to create a composite set of images, revealing for the first time M87\*'s event horizon.

It took nearly two decades to achieve !

- ✓ **Main idea:** use DVCS, TCS, DVMP... and related processes as probes
- ✓ **Precision:** high luminosity in a wide kinematic range is key!
- ✓ **Data Management:** unprecedented amount of data need new AI based techniques to handle the image making

In the course of 10 years, first proton image!

Topic	Hall A	Hall B	Hall C	Hall D	Other	Total
Hadron spectra as probes of QCD	0	2	1	3	0	6
Transverse structure of the hadrons	6	3	3	1	0	13
Longitudinal structure of the hadrons	2	3	7	0	0	12
3D structure of the hadrons	5	9	6	0	0	20
Hadrons and cold nuclear matter	8	5	7	0	1	21
Low-energy tests of the Standard Model and Fundamental Symmetries	3	1	0	1	2	7
<b>Total</b>	<b>24</b>	<b>23</b>	<b>24</b>	<b>5</b>	<b>3</b>	<b>79</b>

# Center for Femtography (CNF)



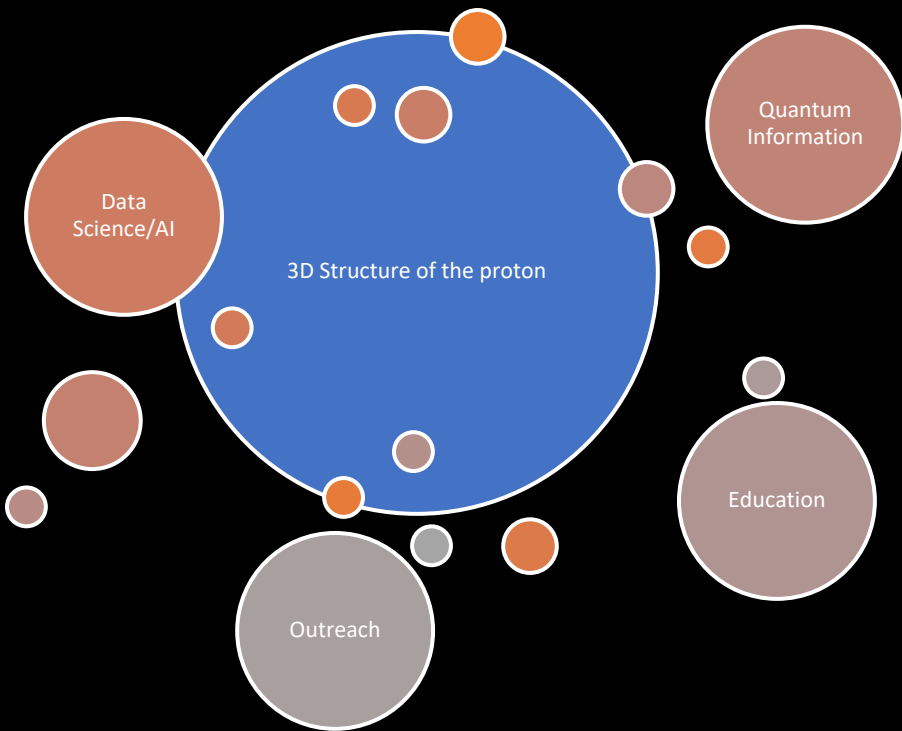
## SURA

- December 2018 -- Founded at UVA meeting
- Summer 2019 -- Pilot projects started funded by SURA
- Summer 2020 -- Xiangdong Ji is appointed director



CNF Talks series starts on September 2<sup>nd</sup> and will continue every Wednesday at 11am, please join our discussions!  
(Newly designed website will have all information)

sl4y@virginia.edu



- CNF has funded multiple projects on Femtography
- Covering different areas
  - Experimental data
  - ML & AI
  - Inverse problems
  - Lattice calculations
  - ...
- Total funding exceeding \$0.5M per year
- Now it is time to organize different efforts into a larger collaboration, involving people outside VA.

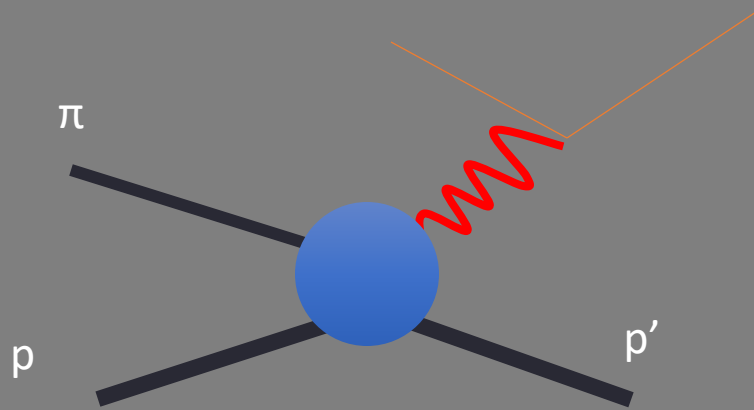
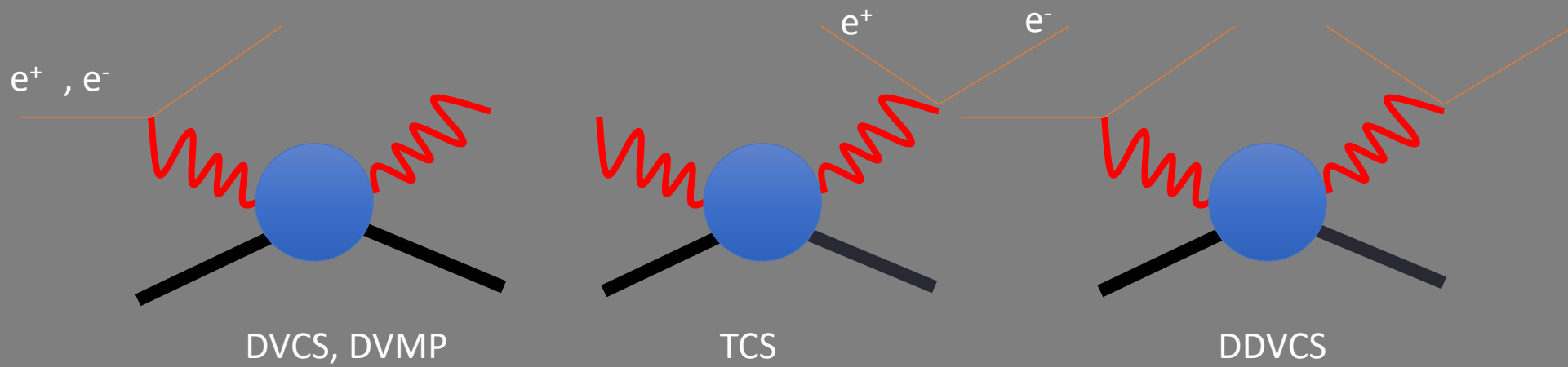
Courtesy of X. Ji



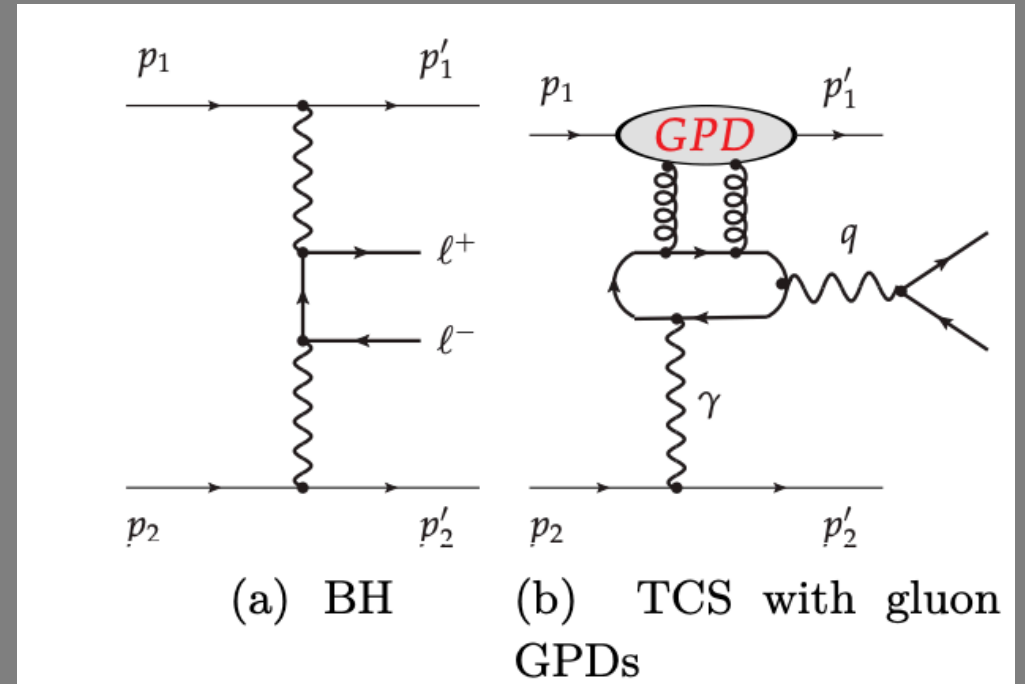
# Coordinated directions

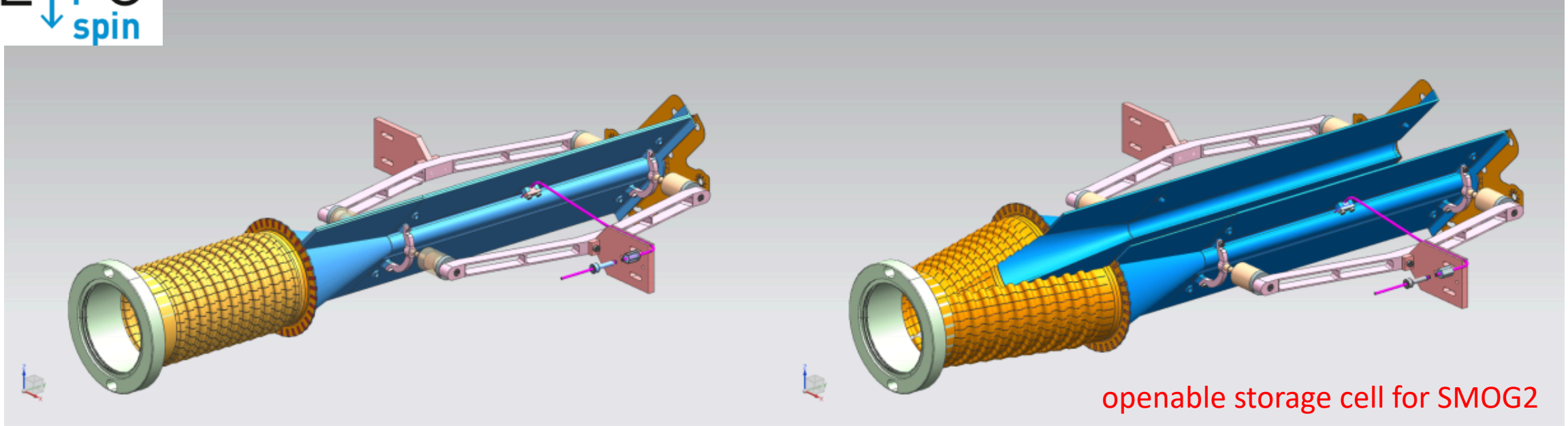
1. GPD related experimental data base (+form factors)
2. GPD models and parametrizations
3. QCD scale evolution to NLO or NNLO
4. Lattice QCD calculations on GPDs
5. GPD fitting (inverse problem), ML and AI technique.
6. GPD data base, 3D Femtography, outreach.

# Harnessing/coordinating information from all channels



Exclusive pion induced DY (EDY)  
Sawada et al., PRD93 (2016)





## Unpolarized

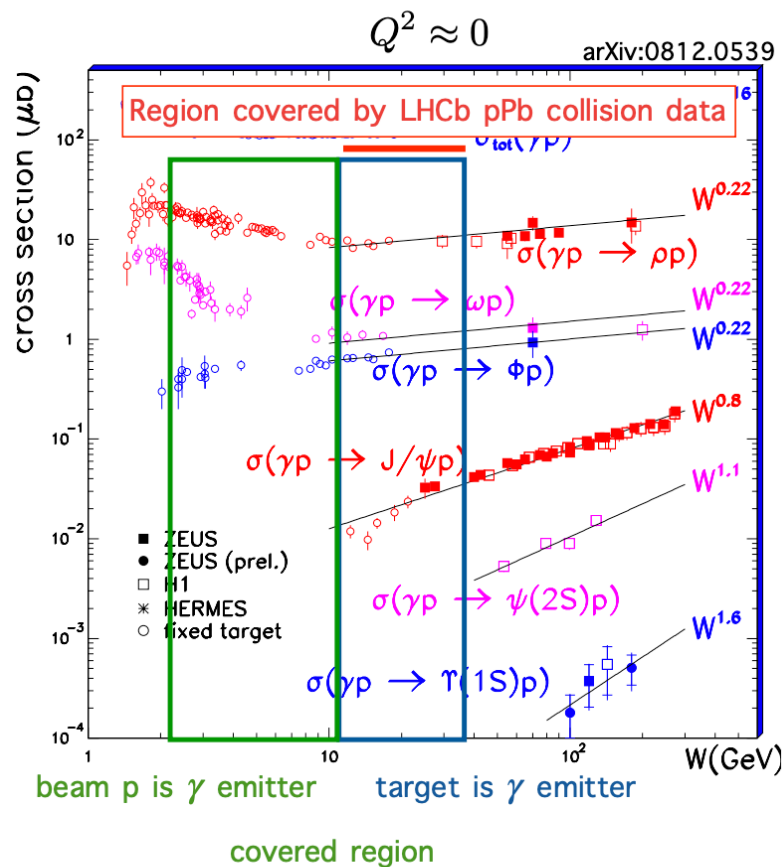
LHCb experiment pioneered fixed-target physics with LHC beams, thanks to the SMOG internal gas target

## Polarized

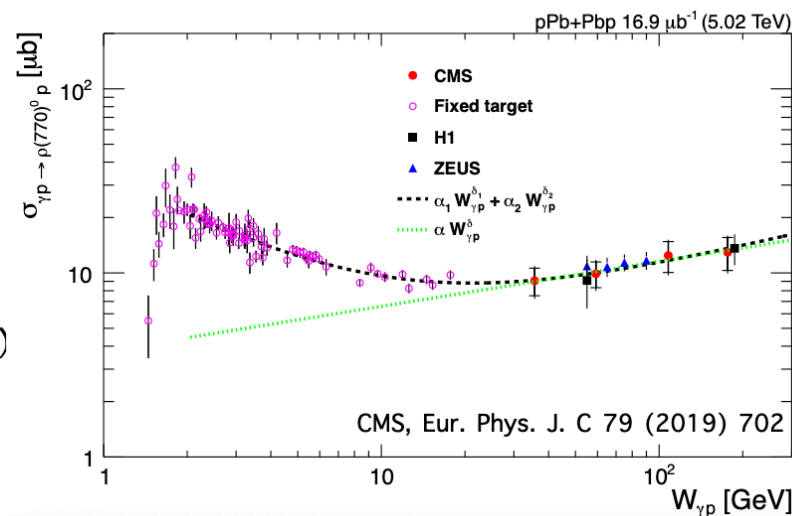
## The LHCSpin Project

C.A. Aidala, A. Bacchetta, M. Boglione, G. Bozzi, V. Carassiti, M. Chiosso, R. Cimino, G. Ciullo, M. Contalbrigo, U. D'Alesio, P. Di Nezza, R. Engels, K. Grigoryev, D. Keller, P. Lenisa, S. Liuti, A. Metz, P.J. Mulders, F. Murgia, A. Nass, D. Panzieri, L.L. Pappalardo, B. Pasquini, C. Pisano, M. Radici, F. Rathmann, D. Reggiani, M. Schlegel, S. Scopetta, E. Steffens, A. Vasilyev

## Exclusive $\rho^0$ production

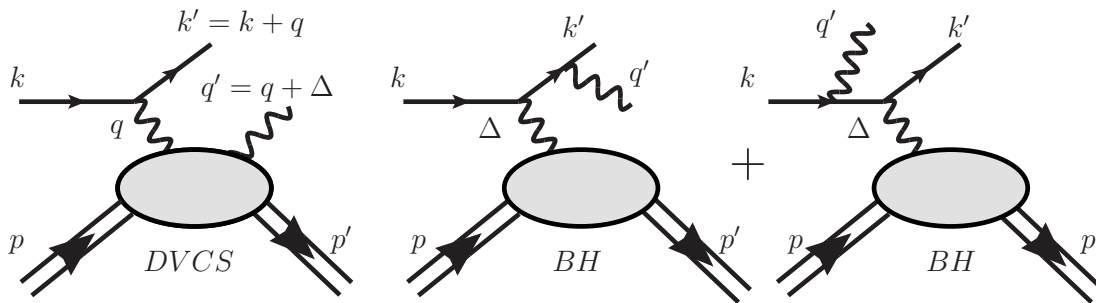


- Measure nuclear ratios
- Compare to pPb collision data
- Theorists contacted for solid theory interpretation
- Measured recently by CMS





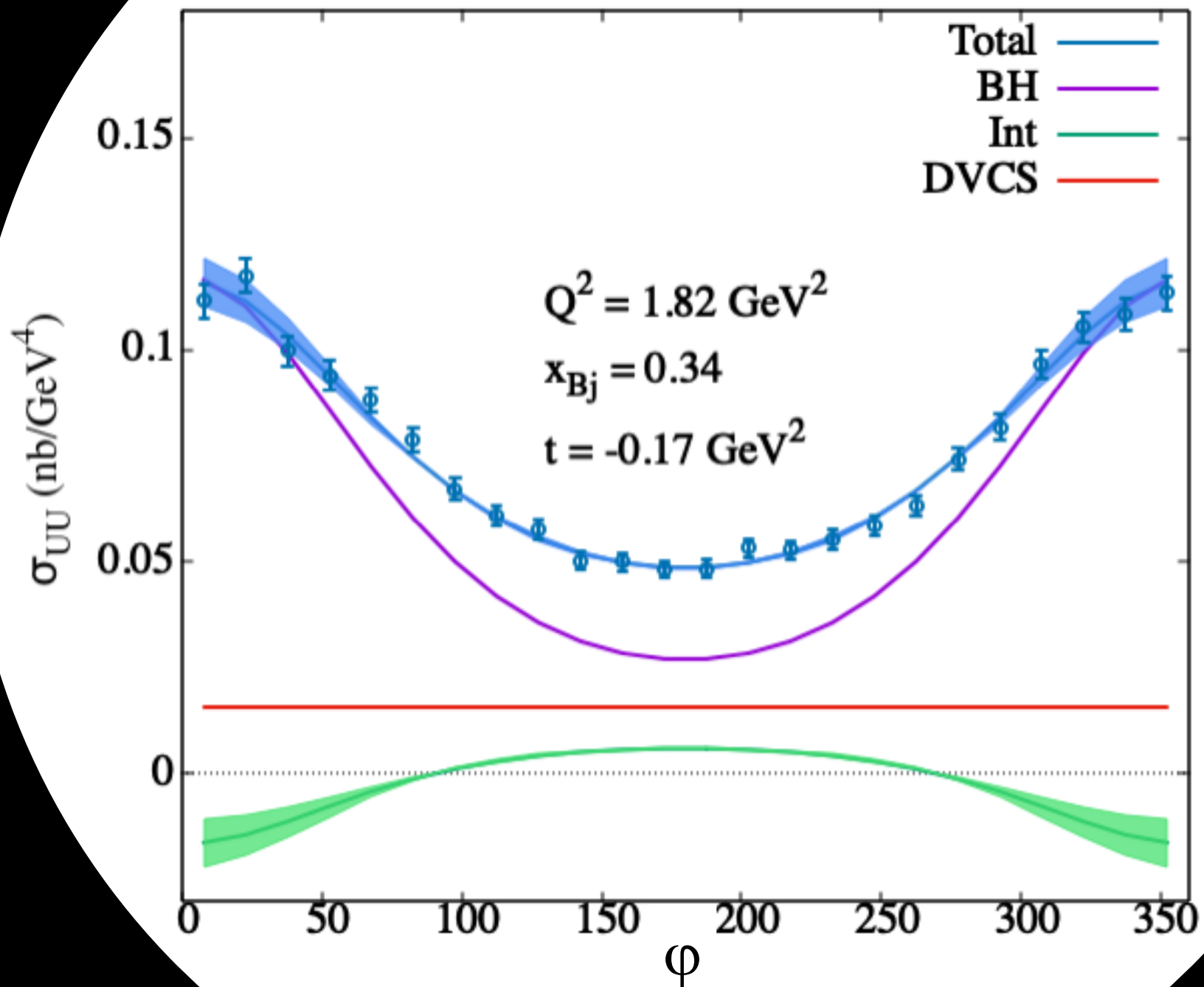
We need a robust framework for the cross section, where kinematic limits are under control (beyond “harmonics” model)



**B. Kriesten et al, *Phys.Rev. D* 101 (2020)**  
**B. Kriesten and S. Liuti, arXiv [2004.08890](https://arxiv.org/abs/2004.08890)**

$$\frac{d^5\sigma}{dx_{Bj}dQ^2d|t|d\phi d\phi_S} = \frac{\alpha^3}{16\pi^2(s - M^2)^2\sqrt{1 + \gamma^2}} |T|^2,$$

$$T(k, p, k', q', p') = T_{DVCS}(k, p, k', q', p') + T_{BH}(k, p, k', q', p'),$$



# DVCS

$$\begin{aligned}
 \frac{d^5\sigma_{DVCS}}{dx_{Bj}dQ^2d|t|d\phi d\phi_S} &= \boxed{\text{twist two GPDs}} \frac{\alpha^3}{16\pi^2(s-M^2)^2\sqrt{1+\gamma^2}} |T_{DVCS}|^2 \\
 &= \boxed{\text{twist three GPDs}} \frac{\Gamma}{Q^2(1-\epsilon)} \left\{ \boxed{F_{UU,T}} - \epsilon \boxed{F_{UU,L}} + \epsilon \cos 2\phi \boxed{F_{UU}^{\cos 2\phi}} \right. \\
 &+ \sqrt{\epsilon(\epsilon+1)} \left[ \cos \phi \boxed{F_{UU}^{\cos \phi}} + \sin \phi \cancel{F_{UU}^{\sin \phi}} \right] \\
 &+ \lambda_e \sqrt{2\epsilon(1-\epsilon)} \sin \phi \boxed{F_{LU}^{\sin \phi}} \\
 &+ \boxed{S_L} \left[ \cancel{F_{UL}^{\sin \phi}} + \sqrt{\epsilon(\epsilon+1)} \sin \phi \boxed{F_{UL}^{\sin \phi}} + \epsilon \sin 2\phi \boxed{F_{UL}^{\sin 2\phi}} \right] \\
 &+ \lambda_e \left[ \sqrt{1-\epsilon^2} \boxed{F_{LL}} + 2 \lambda_e \sqrt{\epsilon(1-\epsilon)} \cos \phi \boxed{F_{LL}^{\cos \phi}} \right] \\
 &+ \boxed{S_T} \left[ \sin(\phi - \phi_S) \left( \boxed{F_{UT,T}^{\sin(\phi - \phi_S)}} + \epsilon \boxed{F_{UT,L}^{\sin(\phi - \phi_S)}} \right) \right. \\
 &\quad \epsilon \sin(\phi + \phi_S) \boxed{F_{UT}^{\sin(\phi + \phi_S)}} + \epsilon \sin(3\phi - \phi_S) \boxed{F_{UT}^{\sin(3\phi - \phi_S)}} \\
 &\quad \left. + \sqrt{2\epsilon(1+\epsilon)} \left( \sin \phi_S \boxed{F_{UT}^{\sin \phi_S}} + \sin(2\phi - \phi_S) \boxed{F_{UT}^{\sin(2\phi - \phi_S)}} \right) \right] \\
 &+ \left( \lambda_e \boxed{S_L} \left[ \sqrt{1-\epsilon^2} \cos(\phi - \phi_S) \boxed{F_{LT}^{\cos(\phi - \phi_S)}} + \sqrt{2\epsilon(1-\epsilon)} \cos \phi_S \boxed{F_{LT}^{\cos \phi_S}} \right. \right. \\
 &\quad \left. \left. + \sqrt{2\epsilon(1-\epsilon)} \cos(2\phi - \phi_S) \boxed{F_{LT}^{\cos(2\phi - \phi_S)}} \right] \right)
 \end{aligned}$$

U

L

T

# BH

$$\frac{d^5 \sigma_{unpol}^{BH}}{dx_{Bj} dQ^2 d|t| d\phi d\phi_S} \equiv \frac{\Gamma}{t} F_{UU}^{BH} = \frac{\Gamma}{t} \left[ A(y, x_{Bj}, t, Q^2, \phi) (F_1^2 + \tau F_2^2) + B(y, x_{Bj}, t, Q^2, \phi) \tau G_M^2(t) \right]$$

$$A = \frac{16 M^2}{t(k' q')(k' q')} \left[ 4\tau \left( (k P)^2 + (k' P)^2 \right) - (\tau + 1) \left( (k \Delta)^2 + (k' \Delta)^2 \right) \right]$$
$$B = \frac{32 M^2}{t(k' q')(k' q')} \left[ (k \Delta)^2 + (k' \Delta)^2 \right],$$

...compared to BKM, NPB (2001)

$$|\mathcal{T}_{\text{BH}}|^2 = \frac{e^6}{x_{\text{B}}^2 y^2 (1 + \epsilon^2)^2 \Delta^2 \mathcal{P}_1(\phi) \mathcal{P}_2(\phi)} \times \left\{ c_0^{\text{BH}} + \sum_{n=1}^2 c_n^{\text{BH}} \cos(n\phi) + s_1^{\text{BH}} \sin(\phi) \right\},$$

$$c_{0,\text{unp}}^{\text{BH}} = 8K^2 \left\{ (2 + 3\epsilon^2) \frac{Q^2}{\Delta^2} \left( F_1^2 - \frac{\Delta^2}{4M^2} F_2^2 \right) + 2x_{\text{B}}^2 (F_1 + F_2)^2 \right\} \\ + (2 - y)^2 \left\{ (2 + \epsilon^2) \left[ \frac{4x_{\text{B}}^2 M^2}{\Delta^2} \left( 1 + \frac{\Delta^2}{Q^2} \right)^2 \right. \right. \\ \left. \left. + 4(1 - x_{\text{B}}) \left( 1 + x_{\text{B}} \frac{\Delta^2}{Q^2} \right) \right] \left( F_1^2 - \frac{\Delta^2}{4M^2} F_2^2 \right) \right. \\ \left. + 4x_{\text{B}}^2 \left[ x_{\text{B}} + \left( 1 - x_{\text{B}} + \frac{\epsilon^2}{2} \right) \left( 1 - \frac{\Delta^2}{Q^2} \right)^2 \right. \right. \\ \left. \left. - x_{\text{B}}(1 - 2x_{\text{B}}) \frac{\Delta^4}{Q^4} \right] (F_1 + F_2)^2 \right\} \\ + 8(1 + \epsilon^2) \left( 1 - y - \frac{\epsilon^2 y^2}{4} \right) \\ \times \left\{ 2\epsilon^2 \left( 1 - \frac{\Delta^2}{4M^2} \right) \left( F_1^2 - \frac{\Delta^2}{4M^2} F_2^2 \right) - x_{\text{B}}^2 \left( 1 - \frac{\Delta^2}{Q^2} \right)^2 (F_1 + F_2)^2 \right\},$$

A.V. Belitsky et al. / Nuclear Physics B 629 (2002) 323–392

$$c_{1,\text{unp}}^{\text{BH}} = 8K(2 - y) \left\{ \left( \frac{4x_{\text{B}}^2 M^2}{\Delta^2} - 2x_{\text{B}} - \epsilon^2 \right) \left( F_1^2 - \frac{\Delta^2}{4M^2} F_2^2 \right) \right. \\ \left. + 2x_{\text{B}}^2 \left( 1 - (1 - 2x_{\text{B}}) \frac{\Delta^2}{Q^2} \right) (F_1 + F_2)^2 \right\},$$

$$c_{2,\text{unp}}^{\text{BH}} = 8x_{\text{B}}^2 K^2 \left\{ \frac{4M^2}{\Delta^2} \left( F_1^2 - \frac{\Delta^2}{4M^2} F_2^2 \right) + 2(F_1 + F_2)^2 \right\}.$$

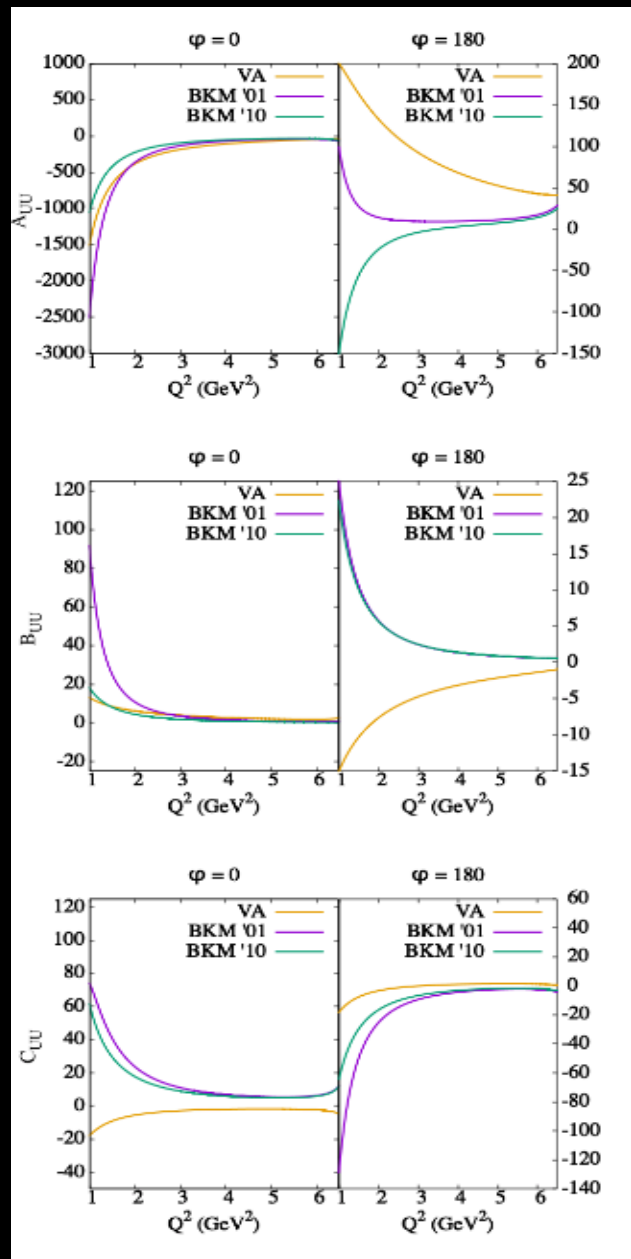


# BH-DVCS interference

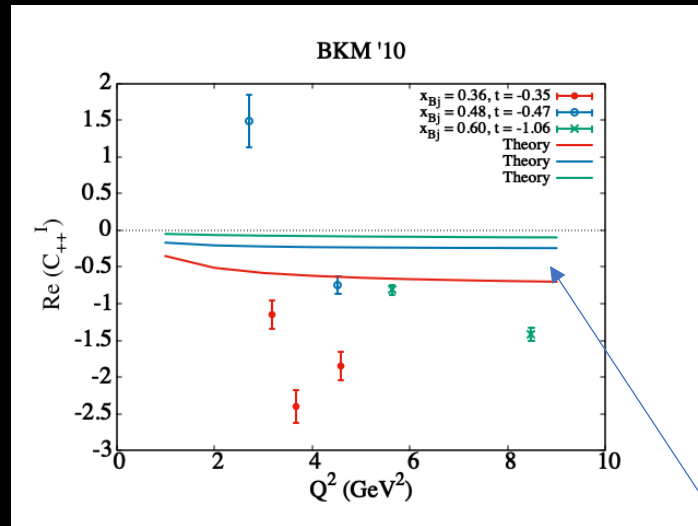
$$F_{UU}^{\mathcal{I}} = F_{UU}^{\mathcal{I},tw2} + \frac{K}{\sqrt{Q^2}} F_{UU}^{\mathcal{I},tw3}$$

$$F_{UU}^{\mathcal{I},tw2} = A_{UU}^{\mathcal{I}} \Re (F_1 \mathcal{H} + \tau F_2 \mathcal{E}) + B_{UU}^{\mathcal{I}} G_M \Re (\mathcal{H} + \mathcal{E}) + C_{UU}^{\mathcal{I}} G_M \Re \tilde{\mathcal{H}}$$

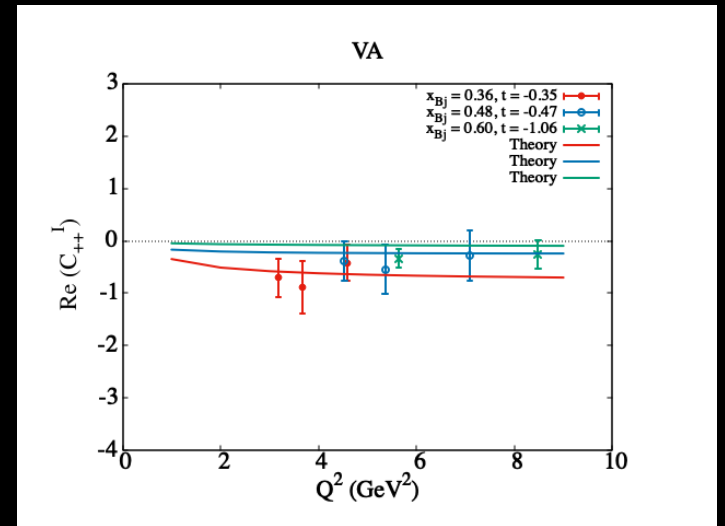
# Impact on $Q^2$ dependence



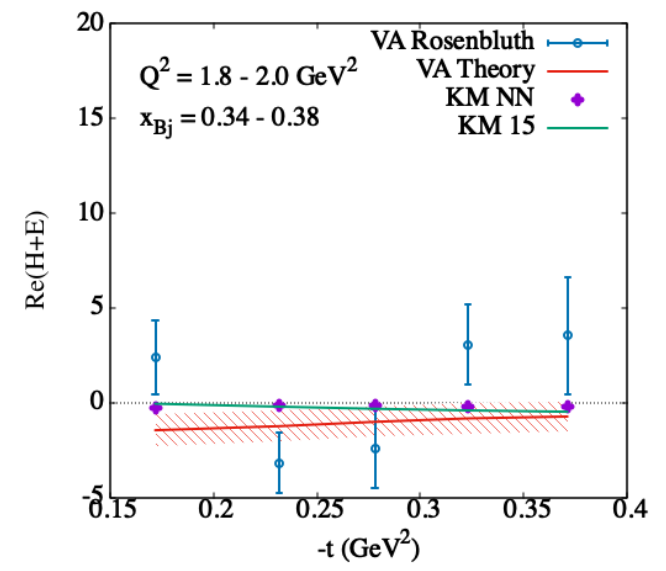
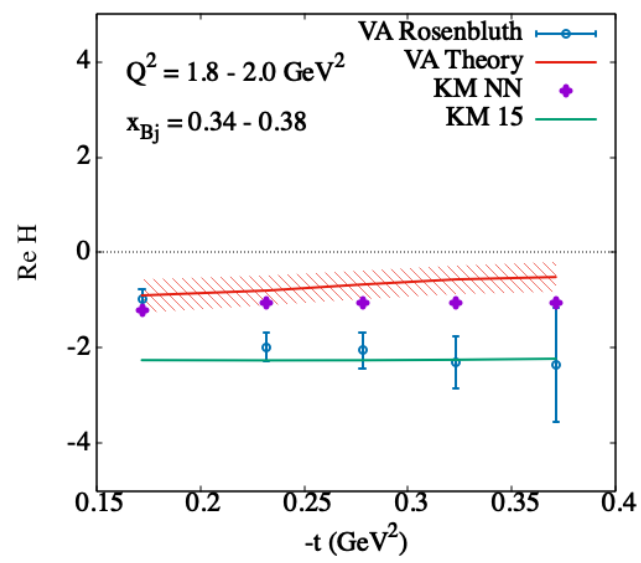
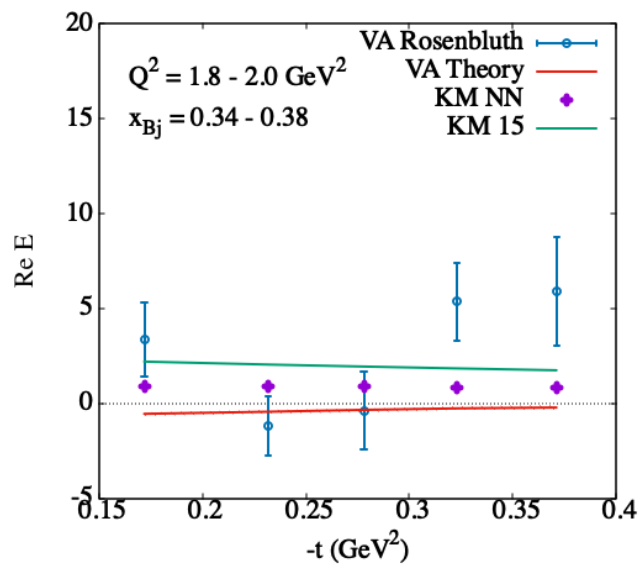
Re H

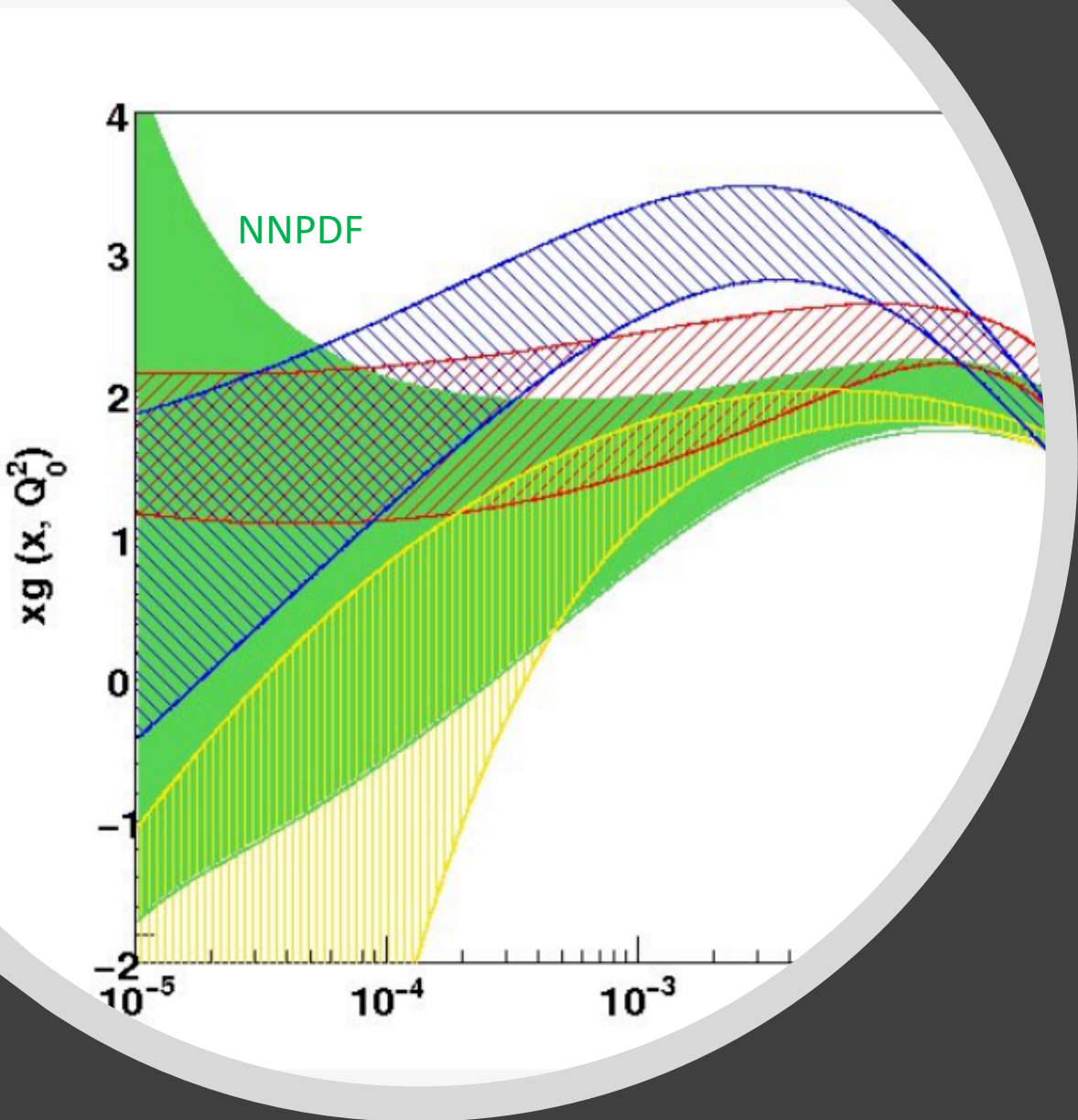


Re H



pQCD Evolution





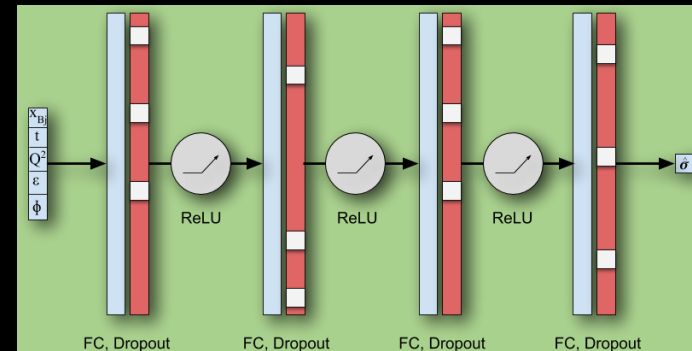
Extracting information from data with standard methods is painstakingly slow: can we use ML/AI?

Brandon Kriesten (UVA), Jake Grigsby (UVA), Pete Alonzi (DS, UVA), Matthias Burkardt (NMSU), Joshua Hoskins (Mary Washington U.), SL

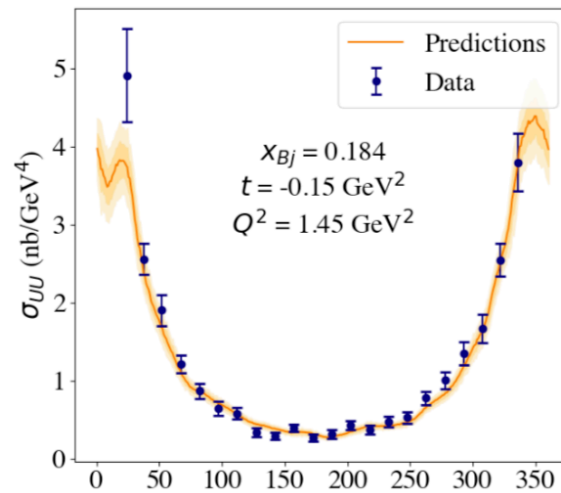
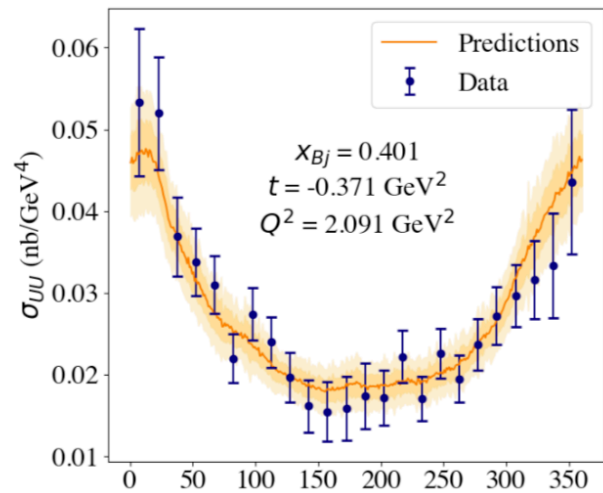
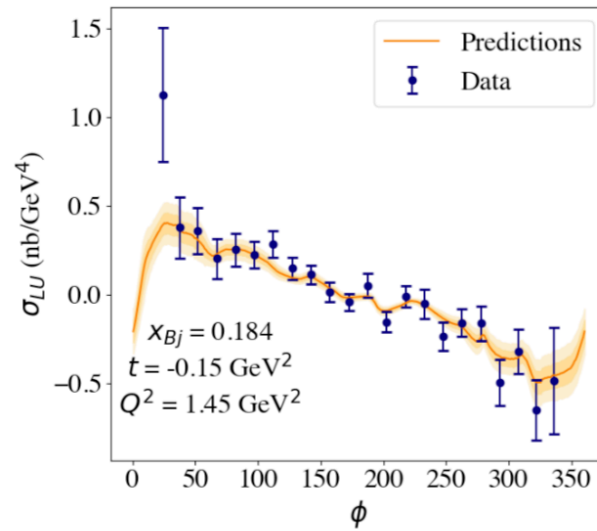
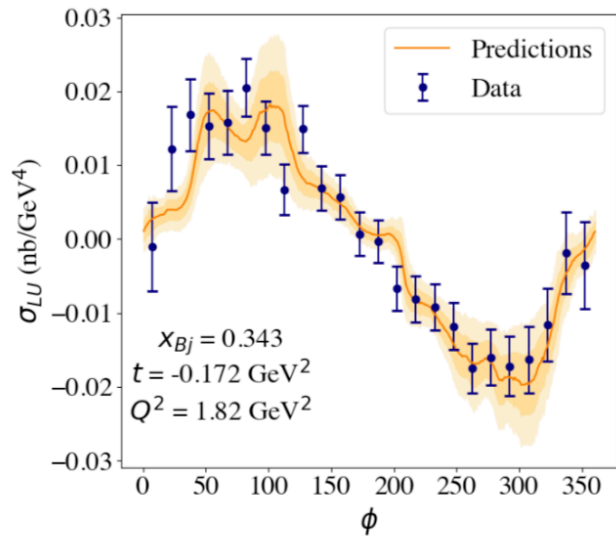
# UVA Strategy: supervised learning algorithm augmented by unsupervised (UMAP) method

## ➤ Using Deep Neural Networks for regression

- The model is a multilayered perceptron (MLP) with 4 hidden layers, each followed by a rectified linear unit (ReLU) activation function
- Once the model is trained, we can use it to make predictions about the value of the cross section in a given kinematic region. The cross section at a point  $(x_{Bj}, t, Q^2, E, \phi)$  is computed by a forward pass of the network. This operation is parallelizable and can run efficiently on a GPU
- We estimate the model's uncertainty by performing many forward passes per prediction point, with Dropout turned on.







# Learning the $\phi$ dependence

# Estimating Uncertainty

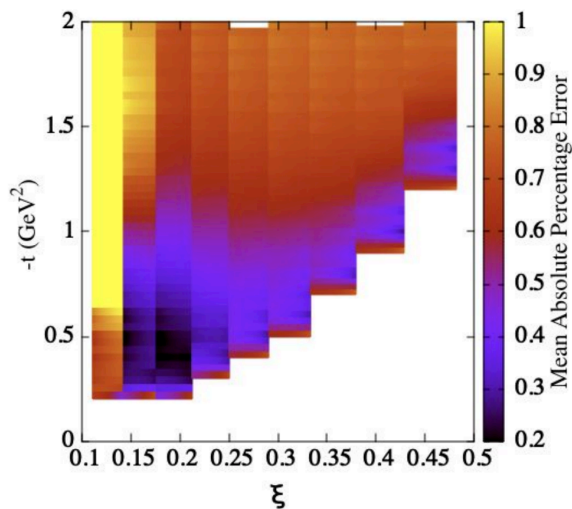
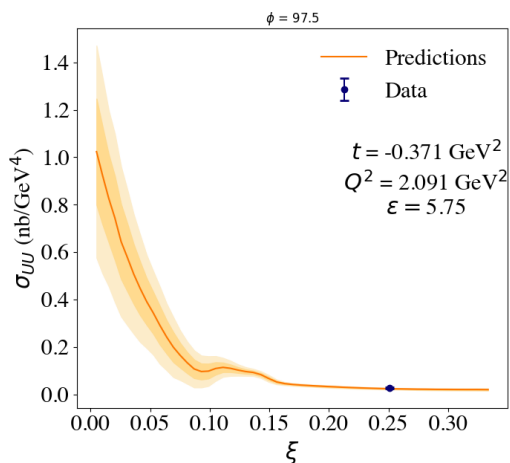
(complementary to PARTON, next talk!)

- We are not restricted to kinematic bins that have already been measured - the model learns to extrapolate between existing bins.

- We can sweep through regions where data do not exist, and use the model's uncertainty estimate to advise our use of its predictions

- This creates a distribution of predictions, and we use the mean and standard deviation of that distribution to estimate uncertainty

- (Gal & This model will be open-sourced

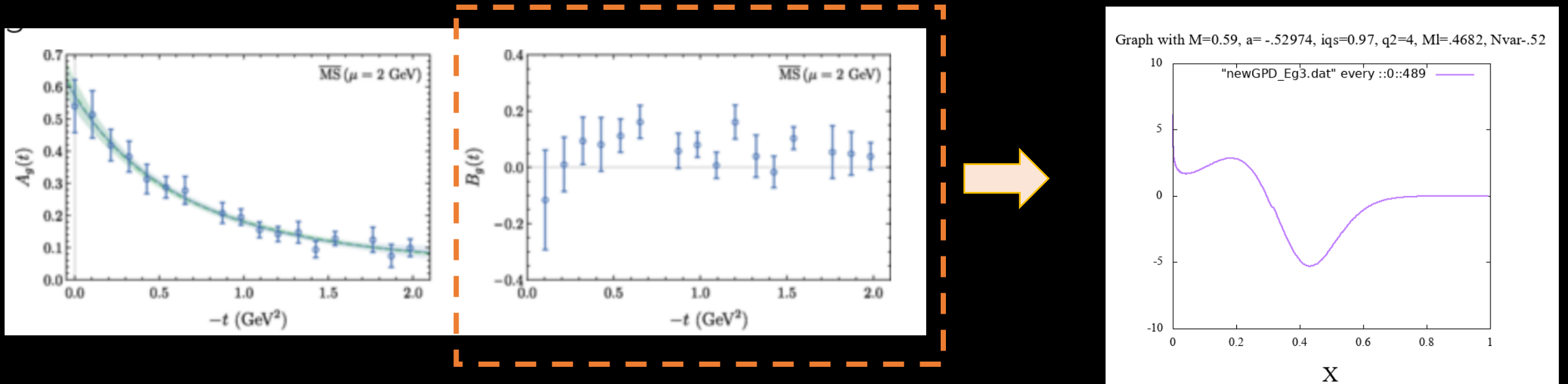


# Moving Towards EIC, LHCSpin

Gluon GPDs  $J_g + \sum_q J_q = \frac{1}{2}$   $H_g, E_g$

Accessed through PQCD evolution and  $J/\psi$  production

Parametrizations/models merging information from lattice QCD



UVA “state of the art results” (B. Kriesten, P. Velie, E. Yeats, F. Yopez)

## What I left out

- **Nuclei:**  $^4\text{He}$  and deuteron GPDs
- $\pi^0$  electroproduction as a means to access the **tensor charge and transversity GPDs**

## CONCLUSIONS

Immense discovery potential as we uncover the mechanical properties the of the proton and observe its spatial images through deeply virtual **exclusive** experiments

To observe, evaluate and interpret GPDs and Wigner distributions at the subatomic level requires stepping up data analyses from the standard methods



developing new numerical/analytic/quantum computing methods



**Center for Nuclear Femtography is a natural entry in the Snowmass Agenda!**