# Nucleon Decay EXPERIMENTS

## **BLV circa 2020**



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think beyond the possible"

## Scientific Perspective of Nucleon Decay

- $\cdot$  Tests a fundamental, but unexplained conservation law: baryon number.
- ❖ Grand Unified Theories make specific predictions: decay modes, lifetimes, branching ratios.
- ❖ Probes scales forever inaccessible to accelerators.
- $\div$  New force carrying particles.
- ❖ Deep connections with other fields: cosmology, inflation, BAU, neutrino mass.
- $\cdot$  Even if no signal, limits are very constraining on theory.



### Theoretical Outlook from Experimental Perspective

- $\cdot$  Numerous and various models exist.
- $\cdot \cdot$  Lifetime predictions are not precise –<br>typically uncertain by 2-3 orders of magnitude.
- $\cdot$  There are two favored and benchmark decay modes:  $e^+\pi^0$  (gauge mediated) and  $\nu$ K<sup>+</sup> (SUSY D=5) good for water good for LAr and Liq. Scint.
- $\cdot$  There are other modes and processes:  $\mu^+\pi^0$  (flipped),  $\mu^+K^0$  (SUSY), invisible modes, dinucleon decay, three-body modes, leptonic modes, B+L conserving modes ...

Ideally, we wish to cover all possibilities

 $\cdot$  Some theories suppress or exclude nucleon decay.



#### Historical Perspective





#### Before we look ahead, let's take a quick look at recent work from Super-K



6

## Improvements to SK analysis

**Nonstop effort since 2013:**

Neutron capture on hydrogen (25% eff.) Expanded fiducial volume Two box search (free proton) Improved reconstruction software



See A. Takenaka, NuPhys 2019 <https://indico.cern.ch/event/818781/>







background reduction with neutron capture on hydrogen  $\frac{1}{8}$  with decent efficiency

Low background achieved with decent efficiency



Studying suitability for  $p \to \nu K^+$  and others



#### Nuclear Modeling is Important

- Effective mass in <sup>16</sup>O
- Correlation with other nucleons
- Fermi motion by shell
- Initial position (Woods-Saxon)
- Nuclear de-excitation  $\gamma$
- pion-nuclear interactions
	- Elastic Scattering
	- Charge Exchange
	- Absorption
- Similar issues cause uncertainty in atm. Bkg.

Trouble for decay in the nucleus





#### Exotic Nucleon and Dinucleon Decay Modes



#### <https://arxiv.org/abs/1811.12430> SK (0.36 Mton yrs, preliminary)

**Frejus** 

**IMB** 



2015: see also <https://arxiv.org/abs/1504.01041> for dinucleon decay to pions and <https://arxiv.org/abs/1409.1947> and<https://arxiv.org/abs/1508.05530>



## $nn$

New: SK 1 (old) + 2/3/4 (new) Refined analysis (multivariate) Updated intranuclear absorption (more). Affects signal and background



Unfortunately, search is very much background limited.



Linyan Wan, Neutrino 2020, <https://nusoft.fnal.gov/nova/nu2020postersession/pdf/posterPDF-43.pdf>

and upcoming  $\Delta$ B=2 [workshop:](https://www.physics.umass.edu/acfi/seminars-and-workshops/theoretical-innovations-for-future-experiments-regarding-baryon-number)



#### New experiments that can have an impact



most massive – superior for  $e^{\dagger}\pi^0$ broad search capabilities free proton advantage kaons below Cherenkov threshold

clean timing signature specialize in charged kaon (also invisible mode)

 $10<sup>3</sup>$ 

 $2.2 \mu s$ 

 $10<sup>4</sup>$ 

Hit Time [ns]

**Liquid Scintillator**

JUNO

LENA

THEIA

 $\mu^+ + \nu_\mu$ 

 $10<sup>2</sup>$ 

 $+V$  $12<sub>ns</sub>$ 

10



fine grained detail visible kaon track heavy nucleus, no free protons





 $\frac{B}{20.20}$ 

### $p \rightarrow \nu K^+$  in Liquid Scintillator



K. Asakura,<sup>1</sup> A. Gando,<sup>1</sup> Y. Gando,<sup>1</sup> T. Hachiya,<sup>1</sup> S. Hayashida,<sup>1</sup> H. Ikeda,<sup>1</sup> K. Inoue,1, 2 K. Ishidoshiro,<sup>1</sup>  $\frac{1}{\sqrt{3}}$  and  $\frac{1}{\sqrt{3}}$  and  $\frac{1}{\sqrt{3}}$  or  $\frac{1}{\sqrt{3}}$  is  $\frac{1}{\sqrt{3}}$  is  $\frac{1}{\sqrt{3}}$  or  $\frac{1}{\sqrt{3}}$  or  $\frac{1}{\sqrt{3}}$  is  $\frac{1}{\sqrt{3}}$  $\frac{1}{3}$   $\frac{1}{2}$   $\frac{1}{1}$   $\frac{1}{1}$   $\frac{1}{2}$   $\frac{1}{1}$   $\frac{1}{3}$   $\frac{1}{3}$   $\frac{1}{2}$   $\frac{1}{3}$   $\frac{1}{2}$   $\frac{1}{3}$   $\frac{1}{2}$   $\frac{1}{3}$   $\frac{1}{2}$   $\frac{1}{3}$  $\begin{array}{ccccccc} \mathsf{K} & 4 & & \downarrow & & \downarrow & & \downarrow & & \downarrow & & \mathsf{K} \end{array}$  $\begin{bmatrix} (a) p \rightarrow v \text{ A} & \text{NIC} \\ \vdots & \vdots & \ddots & \vdots \\ 0 & \text{B} & \text{B} & \text{B} & \text{A} \end{bmatrix}$   $\begin{bmatrix} (b) p \rightarrow v \text{ A} & \text{NIC} \\ \vdots & \vdots & \ddots & \vdots \\ 0 & \text{B} & \text{B} & \text{B} & \text{B} \end{bmatrix}$  $\begin{bmatrix} 1.500 & \cdots & 1.500 \\ \vdots & \cdots & \vdots \\ 0 & \cdots & 0 & 0 \end{bmatrix}$   $\begin{bmatrix} E_1 - 99 \text{ MeV} \\ \vdots & E_n = 160 \text{ MeV} \end{bmatrix}$   $\begin{bmatrix} E_1 - 112 \text{ MeV} \\ \vdots & E_n = 160 \text{ MeV} \end{bmatrix}$   $\begin{bmatrix} 1.00 & \cdots & 1.00 \\ \vdots & \vdots & \vdots \\ 0 & \cdots & 0.00 \end{bmatrix}$ 

Residual (σ)

Residu  $\overline{\phantom{a}}$ 2  $\epsilon$ 

(The KampLand Collaboration) is  $\mathbb{R}^n$  $\mathbf{F} = \mathbf{W} \mathbf{N} = \mathbf{I} \qquad \qquad \mathbf{1} \qquad \mathbf{1} \qquad \mathbf{1} \qquad \mathbf{M} \mathbf{N} = \mathbf{1} \qquad \qquad \mathbf{1} \qquad \$  $2\cos\left[\begin{array}{c|c} 1 & 0 \end{array}\right]$  in the UNIVERSITY of the Unive  $T$ okyo, Kashima, Tokyo, Japan 3Graduate School of Tokyo, Kashima, Osaka 560-0043, Japan 3Graduate School of Science, Osaka 560-0043, Japan 3Graduate School of Science, Osaka 560-0043, Japan 3Graduate Science, Toyonaka, J

 $\frac{2}{9}$ <br> $\frac{6}{1000}$ 

2000 -4

-4 -2 0 2 4

0

 $T(\text{ns})$  of 100 200 200  $T(\text{ns})$  100 200  $T(\text{ns})$ 

 $T (ns)$ 

 $T$ (ns)

0 100 200

0 100 200

 $E_1 = 99$  MeV  $E_2 = 169$  MeV  $\Delta T = 19.3$  ns

 $\mathcal{F} = \frac{1}{2} \mathcal{F} \left( \frac{1}{2} \mathcal{F} \right)$ 

Residual (σ)

Residu

1500

1<br>1500a

 $\alpha$  $\overline{\phantom{a}}$ 2  $\overline{\phantom{a}}$ 

 $2000 - (a) p \rightarrow \overline{v} K^+ M C$ 

 $\mathcal{L}_{1500}$   $\mathcal{L}$   $\mathcal{N}$ 

 $\sum_{\substack{10 \ 1000}}^{\infty} 1000$ 

 $\overline{a}$ 

500 1000

500

0

0

 $T (ns)$ 

 $T$ (ns) 0 100 200

(c) atmospheric neutrino MC

(c) atmospheric neutrino MC

0 100 200

 $E_1 = 213 \text{ MeV}$ 

5

E1 = 115 MeV

 $E_1 = 115 \text{ MeV}$ 

#### Proof of principle: KamLAND

Efficiency =  $0.44 \pm 0.05$ Background =  $0.9 \pm 0.02$  events Events  $= 0$ 

$$
\frac{\tau}{B} > 5.4 \times 10^{32} \text{ y}
$$

FIG. 4: Typical waveform shape (shaded) expected from: (a) and (b)  $p \to \overline{\nu} K^{+}$  MC example events and (c) an atmospheric rio. 4. Typical waveform shape (shaded) expected from. (a) and (b)  $p \to pK$  with example events and (c) an atmosphere<br>neutrino MC event, together with the best-fit curves from the multi-pulse fit (solid thick line) and fro the  $\mu^+$  decay daughter of the  $K^+$  (blue dashed line) are also shown. The shape parameters (see text) for each fit are (a)  $log(L_{shape})=0.06$  and (b)  $log(L_{shape})=0.02$ . (solid thin line). In the upper panel, the residuals from the best-fit are shown to identify multi-pulse events. The fitted energies and, for the multi-pulse fits, time differences between the first pulse from the  $K^+$  (red dashed line) and the second pulse from  $\log(L_{\text{other}}) = 0.06$  and (b)  $\log(L_{\text{other}}) = 0.02$ .  $\log (L_{shape}) = 0.00$  data (b)  $\log (L_{shape}) = 0.02$ .

 $\frac{1}{\text{HS}}$  (ns) and (b) muon  $\frac{1}{\text{HS}}$  and  $\frac{1}{\text{HS}}$  and  $\frac{1}{\text{HS}}$  and  $\frac{1}{\text{HS}}$ 

T (ns)

 $T$ (ns) 0 100 200

 $\frac{1}{\Gamma(\text{ns})}$ 

0 100 200

<sup>11</sup> W. Tornow,11, 2 J.A. Detwiler,<sup>12</sup> S. Enomoto,12, 2 and M.P. Decowski13, 2

 $E_1 = 112$  MeV  $E_2 = 160$  MeV  $\Delta T = 9.9$  ns

 $\frac{2}{9}$  1000

 $\overline{a}$ 

500

1500

-4 -2 0 2 4

Residual (σ)

Residu  $-4$ 2  $\epsilon$ 

0

0

(b)  $p \rightarrow \overline{v} K^+ M C$ 

MC *<sup>+</sup>* (b) *p* → ν *K*



0

N.B. Efficiency is near required level but background rate must be reduced

## JUNO arXiv:1508.07166v2<br>JUNO Jiangmen Underground Neutrino Observatory





Cavern excavation in progress through 2020 Detector ready in 2022

Yue Meng Neutrino 2020





JUNO

Primary physic goal is determination of neutrino mass ordering using reactor neutrino oscillation. Also geo, supernova, solar, …

#### **Proton decay**

 $\rightarrow$  K<sup>+</sup> + v

 $12<sub>ns</sub>$ 

 $K^+$ 

10

Number of PE

D

Competitive sensitivity to proton decay searches

 $10<sup>3</sup>$ 

Triple coincidence signal  $\bullet$ 

 $\mu^+ + \nu_\mu$ 

 $10<sup>2</sup>$ 



Yue Meng Neutrino 2020

#### Should enter new territory before DUNE/HK turn on

 $6.75 \times 10^{33}$  protons 85% kaon decay modes x 65% signal efficiency Background 0.5 events in 10 years  $\tau$ 

 $\frac{c}{B} > 1.9 \times 10^{34}$ 

$$
\sum_{\text{University}} \text{Ed Kearns Boston}
$$

#### THEIA: an advanced optical neutrino detector

<https://doi.org/10.1140/epjc/s10052-020-7977-8>



<https://www.bnl.gov/dmo2019/>

Ed Kearns Boston

**Broad program:** Beam neutrinos at SURF Solar, reactor, geo, supernova neutrinos Neutrinoless double beta decay

For  $p \to K^+ \nu$  assume same efficiency and backround rate as JUNO.

La Reams Boston **Core Core Community included** the City of the Cit Aspirational 100 kt detector may cover many modes like SK/HK



Enabled due to great depth if at SURF (reduced cosmic ray spallation) Triple coincidence:  ${}^{12}C \rightarrow {}^{11}C \rightarrow {}^{10}C + \gamma (\tau = 19 \text{ s})$ 

#### DUNE (Deep Underground Neutrino Experiment)



1.5 km deep in SURF (South Dakota) modular … up to 40 kt total fiducial mass Single and dual phase modules 4th module under open study





ProtoDUNE surface data

Liquid Argon Time Projection Chamber (LArTPC)





#### LArTPC Shines for Many Modes

- **EXADED Modes with charged kaon** in final state (SUSY)
- $\bullet\bullet\bullet$  Modes with displaced vertices  $\Rightarrow\bullet\bullet\bullet\bullet\bullet$
- $\cdot$  Multi-prong modes with no neutrino
- $\bullet$  nnbar background rejection
	- No recoil proton allowed
	- No CC electron (or muon)
- $\triangle$  Lepton + light meson likely no better than water due to nuclear absorption of the light meson.





#### $p \rightarrow \nu K^+$ in Liquid Argon Time Projection Chamber (LArTPC)



Figure 4.3: Event display for a decaying kaon candidate  $K \to \mu \nu_{\mu} \mu \to e \nu_{e} \nu_{\mu}$  in the ICARUS T600 detector observed in the CNGS data (K: 90 cm, 325 MeV;  $\mu$ : 54 cm, 147 MeV;  $e$ : 13 cm, 27 MeV). The top figure shows the signal on the collection plane, and the bottom figure shows the signal on the second induction plane [102].

3 mm pitch



circa 2013: Hand scanning suggested efficiency of 97% With background rate of 1 evt/Mty was achievable hep-ph/0701101 (Bueno et al.)



#### Progress in Event Reconstruction

At this time, DUNE is taking the efficiency hit of 30%, with a background rate of 1/Mt y (fully automated analysis, 10 y, 40 kt fiducial mass)

 $\tau/B(K^+\nu) > 1.3 \times 10^{34}$  years



#### Proton decay signal (at SK limit)





J. Klein, Neutrino 2020



 $20^{\sqrt{2}}20$ 

## When ?? Everyone wants to know.



2022 **Detector ready for** data taking

#### 2019-2021

- **Electronics production** starts
- **Civil construction and** lab preparation completed
- **Detector construction**

### **DUNE 2026**

#### **DUNE - Timeline**



Start installing first far detector module Start installing second far detector

DUNE physics data talking starts: Atmospheric neutrinos

Beam operational at 1.2MW

DUNE physics data with beam starts!

- Fiducial mass of 20kt

Third far detector module added (30kt fiducial mass)

Add fourth far detector module (40kt fiducial mass)

Upgrade to 2.4MW beam

25 18/12/2019 Nicola McConkey | DUNE

### **Hyper-K 2027**



Ed Kearns Boston University

But these estimated dates are subject to … change, of course!

#### from the published design reports or recent talks …



We probably should capture some neutral, correct, and up-to-date curves





Likewise, this figure or others like it could be updated



## Messages to Snowmass

Derived from 2013

#### $\triangle$  Testing Baryon Number Violation remains an essential and valued objective of particle physics.

- ❖ Proton decay experiments have been negative so far, and severely constraining of theory. But the ongoing searches are in potentially fruitful territory.
- $\cdot$  The next generation nucleon decay experiments are tied into large neutrino detectors and together with neutrino physics establish a broad science program.
- $\cdot$  The next generation experiments are approved and being constructed.









Reasonable reconstruction result.

#### PRD 95, 012004 (2017)

- No new candidates incl. in enlarged region. ٠
- No significant data excess compared to expected BG (0.89 in BOX2).
- Lower lifetime limit @90%C.L.
- Ed Kearns Bo:  $\tau$  /B<sub>p→</sub> $\mu$ + $\pi$ 0 > 1.6\*10<sup>34</sup> years (published: 7.7\*10<sup>33</sup> years)<br>University Most stringent constraint. **~2 times longer than published.** 32
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