

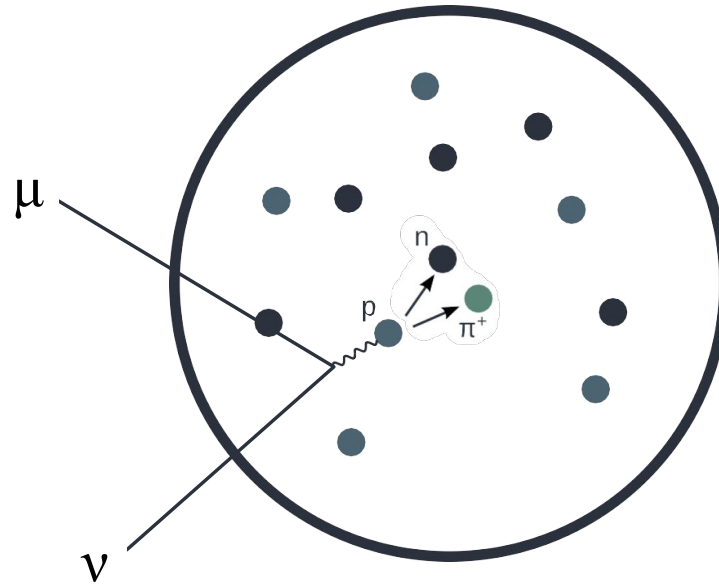
# Neutrino Interaction Physics & the DUNE Near Detector

Mateus F. Carneiro on behalf of the DUNE Collaboration

## Why is it important to measure Cross Sections?

Knowledge of **neutrino-nucleus scattering cross sections** is crucial to the global neutrino physics program

We may understand neutrino-nuclei interactions but we still have a long way to understand the **nuclear effects** that define what we see in our detectors

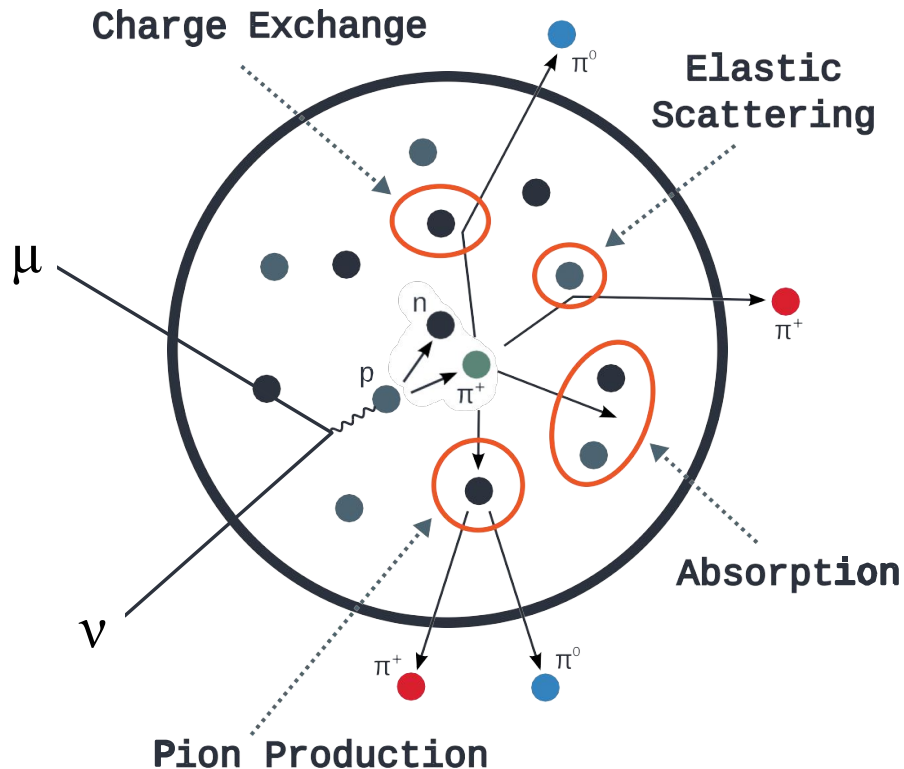


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**Final State Interactions (FSI)** and other **nuclear effects** make different interaction channels have the **same final topology**



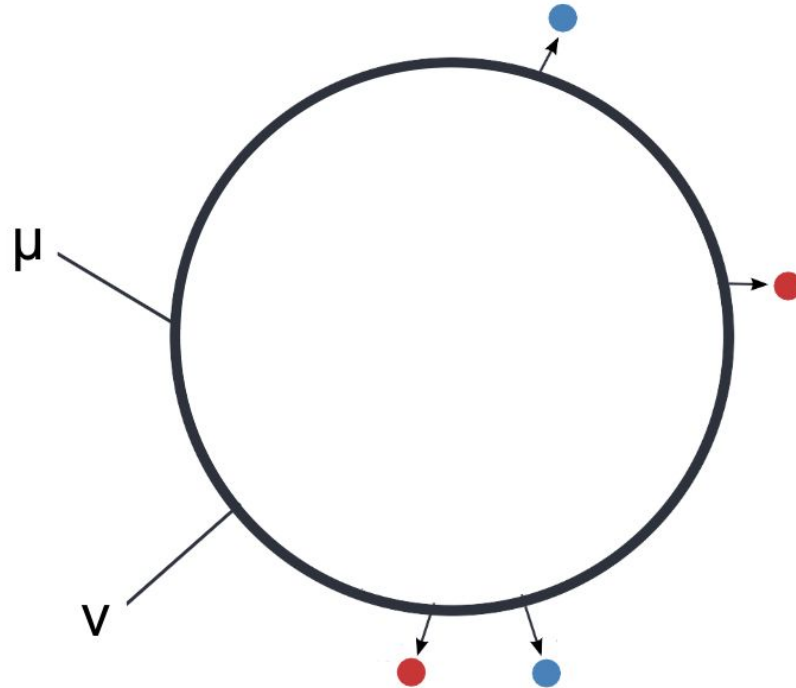
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**Most detectors are made with heavy nuclei**





# Effects in Neutrino Oscillation Systematics

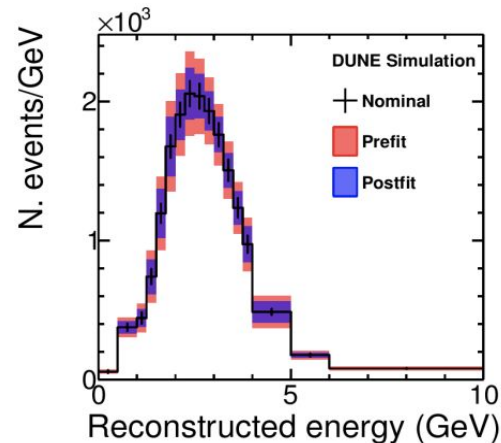
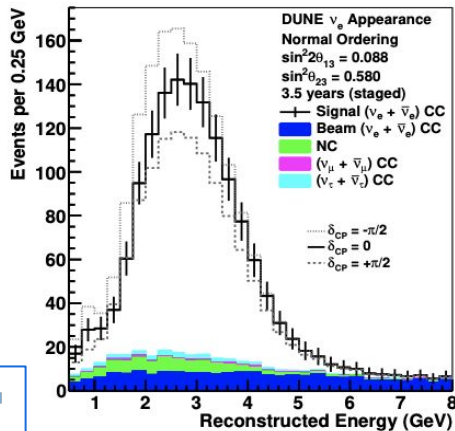
Cross sections and interaction models are **directly used to reconstruct Neutrino energy**, affecting the **event rate** in different stages of the process of measuring oscillation parameters

XSec systematics are often the **largest contribution to the error summary** of neutrino oscillation experiments

The **current level of precision** we have in the ongoing neutrino oscillation program is a **problem for the measurements** we plan to do

"After the constraint from the ND is applied, cross section uncertainties currently comprise 30-50% of the systematic budget on the most important oscillation parameter measurements"

*Jeremy Wolcott, for the NOvA Collaboration (NuFACT 2018) arXiv:1812.05653v1*

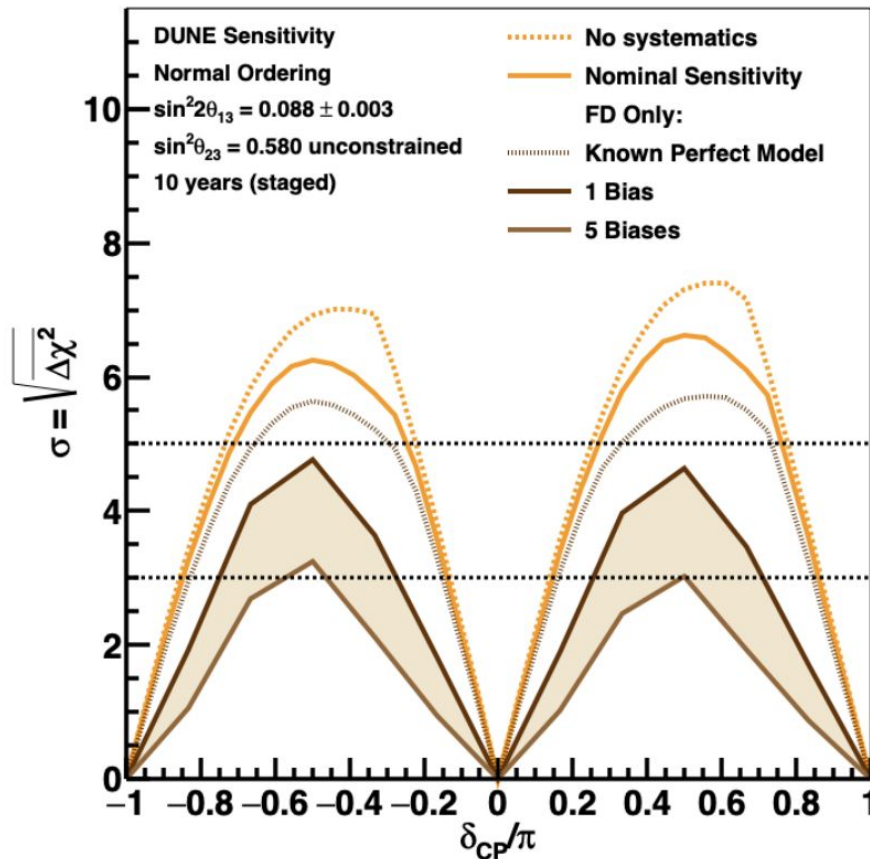


# Effect on CP Violation Sensitivities

Nuclear models also play a big role in CP Violation Sensitivity

For this **sensitivity we are dependent on the event generator assumptions** of the cross section, but what if those assumptions are wrong?

## CP Violation Sensitivity



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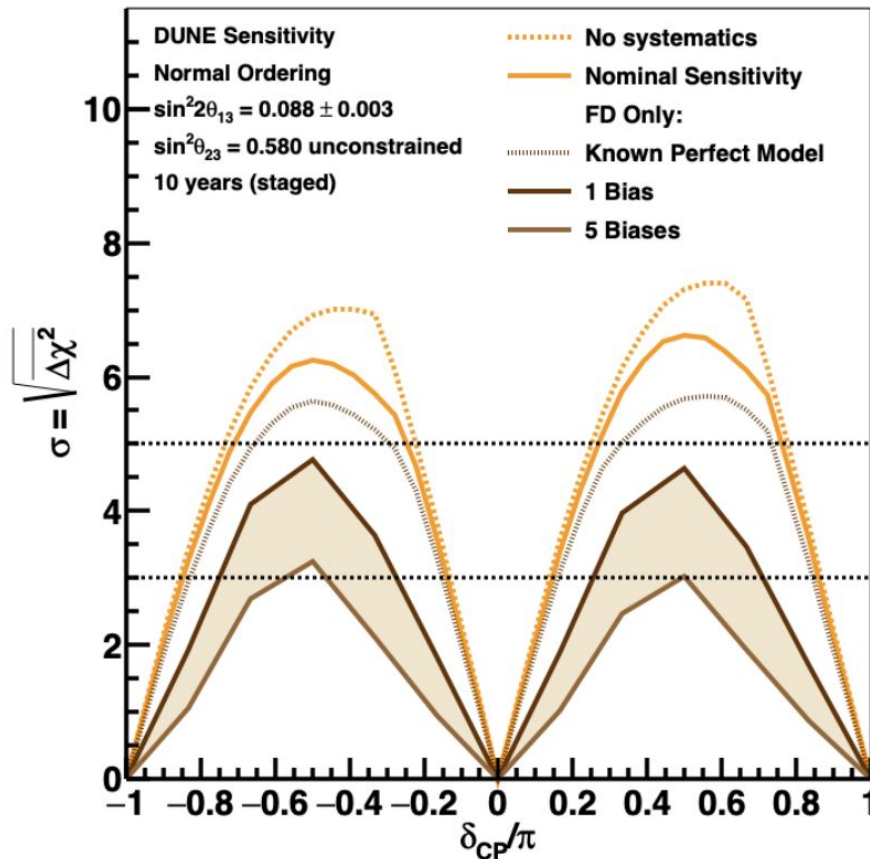
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- **Sensitivity using GENIE**

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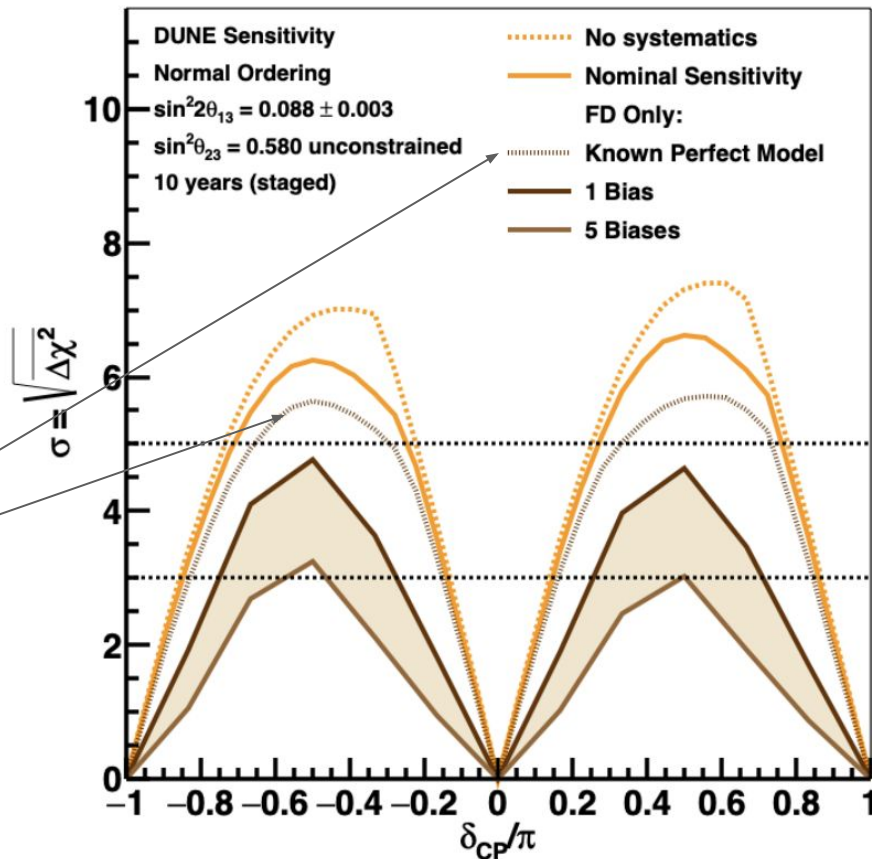
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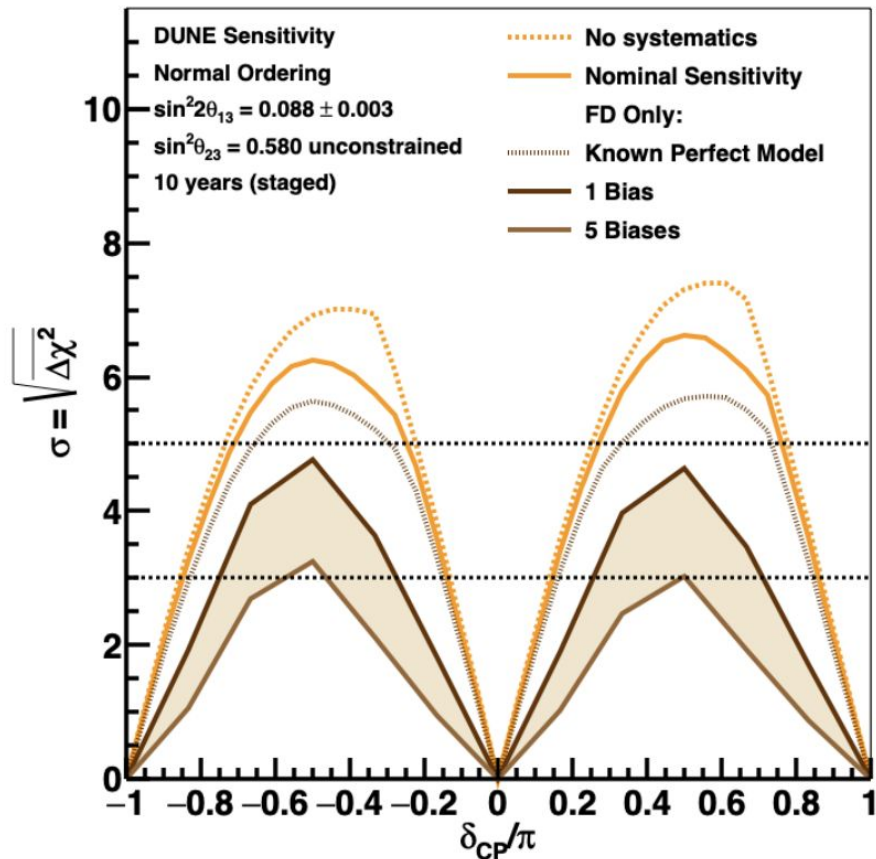
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- **Then we use a different (but plausible) model and input it as fake data**

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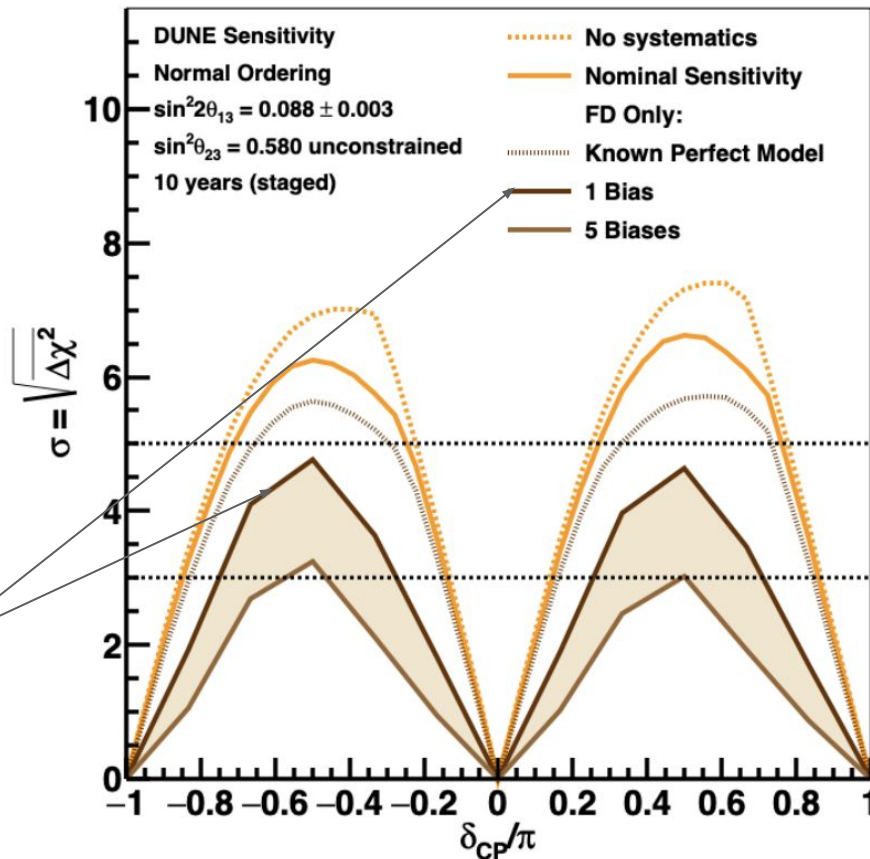
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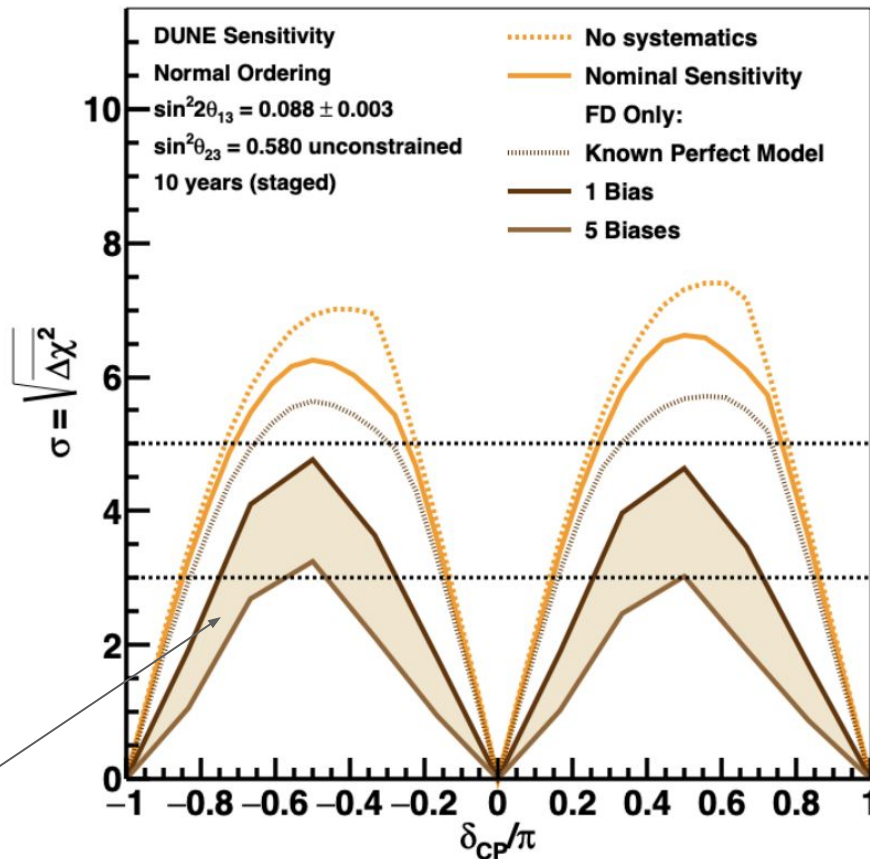
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These are not the only possible viable models, more uncertainties due to **model bias** will need to be included, **reducing sensitivity even further**

## CP Violation Sensitivity



# The DUNE ND Complex

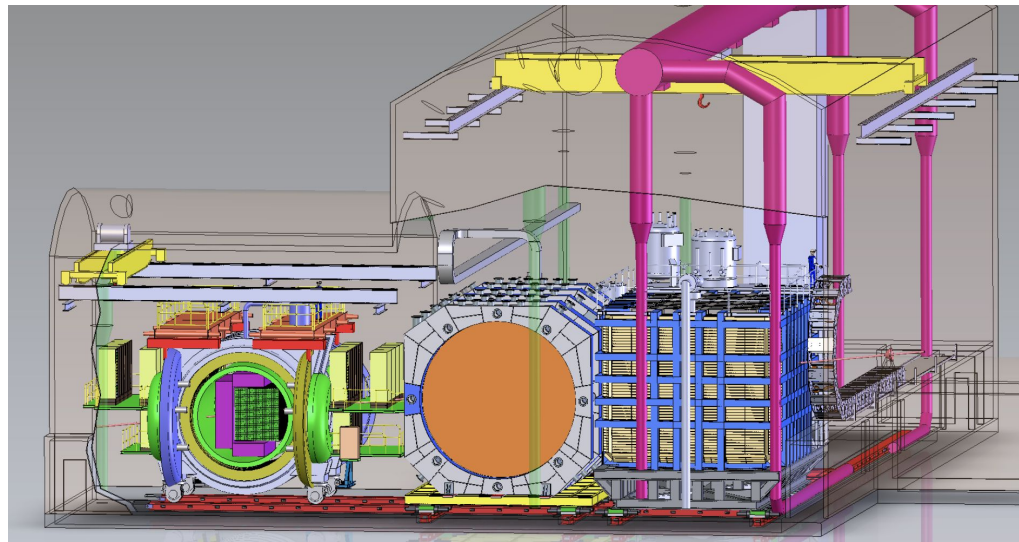
## 3 subsystems:

**ND-LAr:** modular, pixelated LArTPC  
Primary target, similar to FD

**ND-GAr:** high-pressure GARTPC surrounded by  
ECAL and magnet  
muon spectrometer  
Constraint nuclear interaction model,

**SAND:** tracker surrounded by ECAL and magnet  
On-axis monitor of beam spectrum

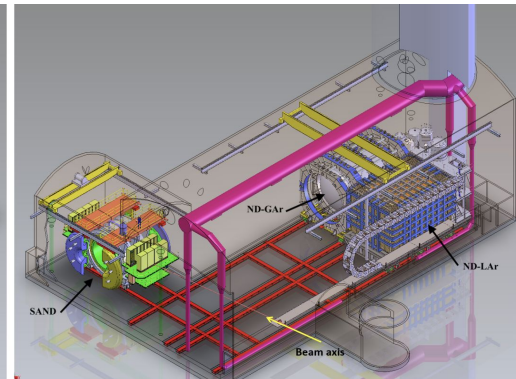
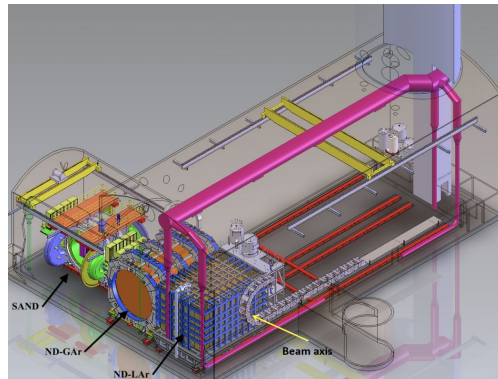
ND-LAr/ND-GAr can move off-axis  
(**DUNE-PRISM**)



SAND

ND-GAr

ND-LAr



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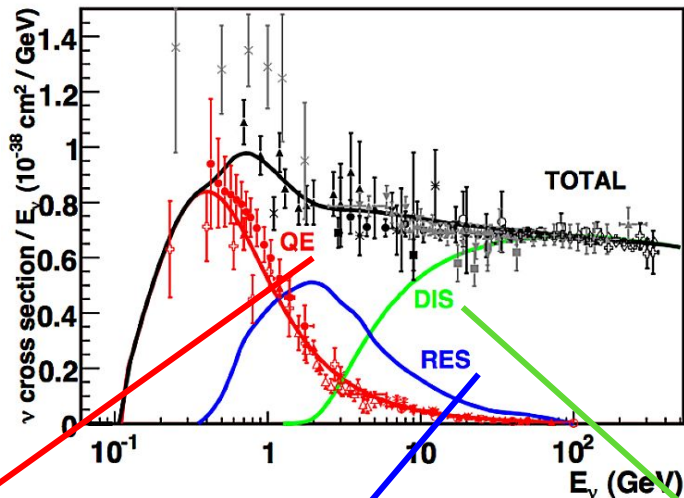


# Separating interaction channels by pion multiplicity in ND-GAr

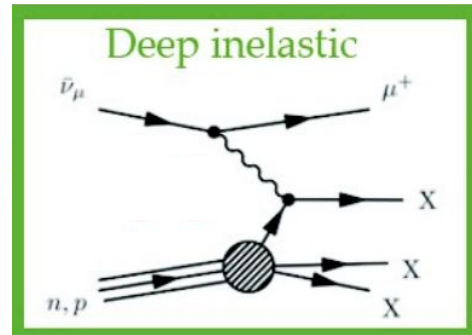
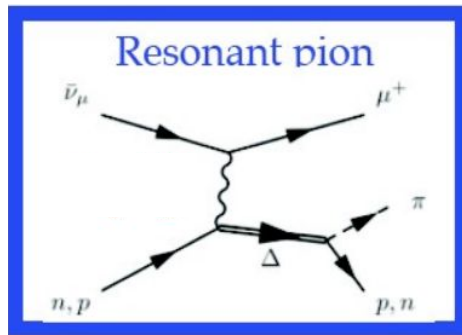
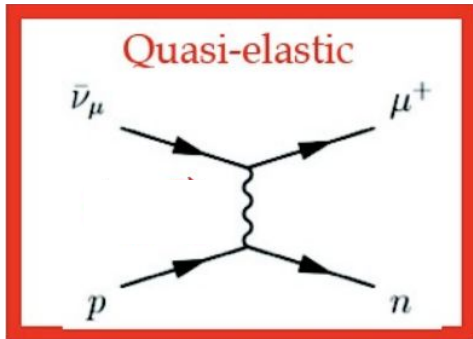
A rich spectrum of interaction types will take place in the near detectors

Simulation show the expected distribution of events in the **ND-GAr by pion multiplicity**

We **can't tell the initial interaction** channels



J.A. Formaggio and G.P. Zeller, Rev. Mod. Phys. 84 (2012)

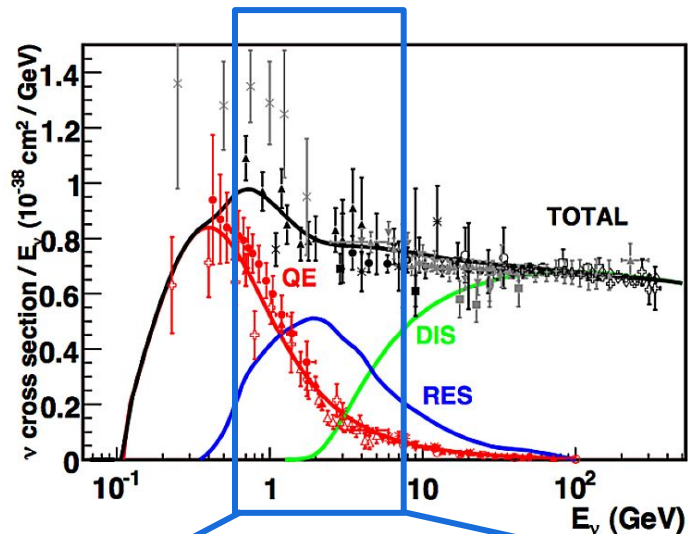


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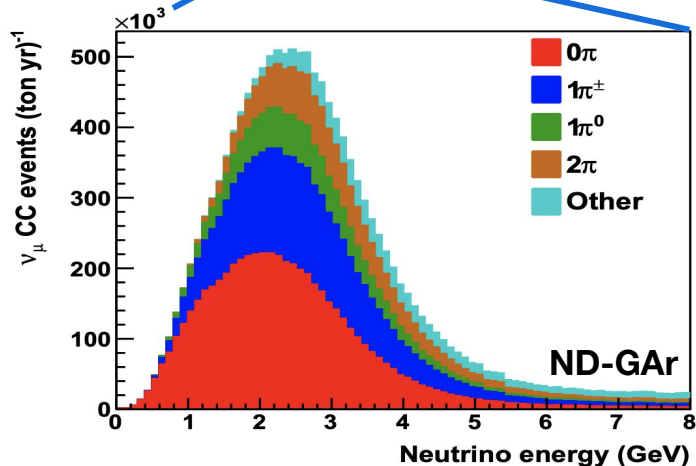
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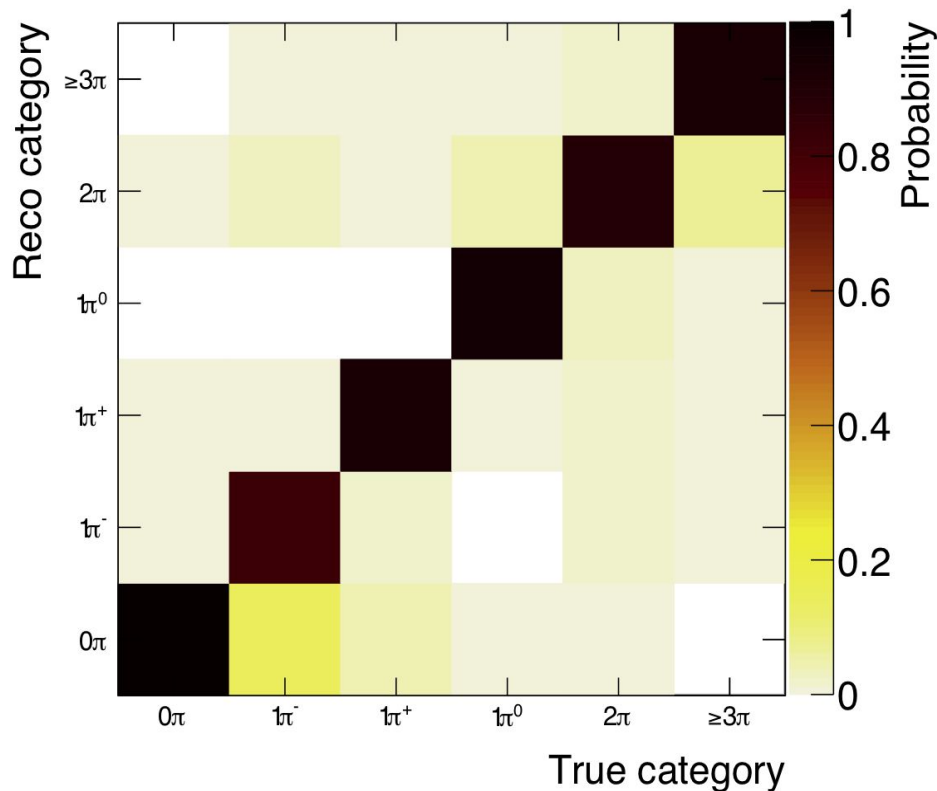
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We **can't tell the initial interaction** channels, but we can **identify the multiplicity** quite well

Final state confusion matrix in **ND-GAr**





## Separating interaction channels by pion multiplicity in ND-GAr

A rich spectrum of interaction types will take place in the near detectors

Simulation show the expected distribution of events in the **ND-GAr by pion multiplicity**

We **can't tell the initial interaction** channels, but we can **identify the multiplicity** quite well and with **great statistics**

More recent MINERvA **high statistics** sample have  $6 \times 10^5$  events

		Interaction Channel		Event Rate	
				ND-LAr	ND-GAr
CC	$\nu_\mu$			$8.2 \times 10^7$	$1.64 \times 10^6$
		$0\pi$		$2.9 \times 10^7$	$5.8 \times 10^5$
		$1\pi^\pm$		$2.0 \times 10^7$	$4.1 \times 10^5$
		$1\pi^0$		$8.1 \times 10^6$	$1.6 \times 10^5$
		$2\pi$		$1.1 \times 10^7$	$2.1 \times 10^5$
		$3\pi$		$4.6 \times 10^6$	$9.3 \times 10^4$
		other		$9.2 \times 10^6$	$1.8 \times 10^5$
	$\bar{\nu}_\mu$			$3.6 \times 10^6$	$7.1 \times 10^4$
	$\nu_e$			$1.45 \times 10^6$	$2.8 \times 10^4$
NC				$5.3 \times 10^5$	$5.5 \times 10^5$
$\nu + e$				$8.3 \times 10^3$	$1.7 \times 10^2$

Events per year ( $1.1 \times 10^{21}$  POT)

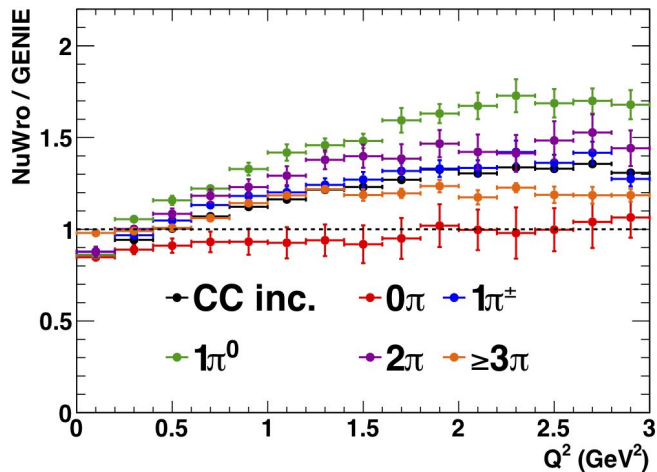
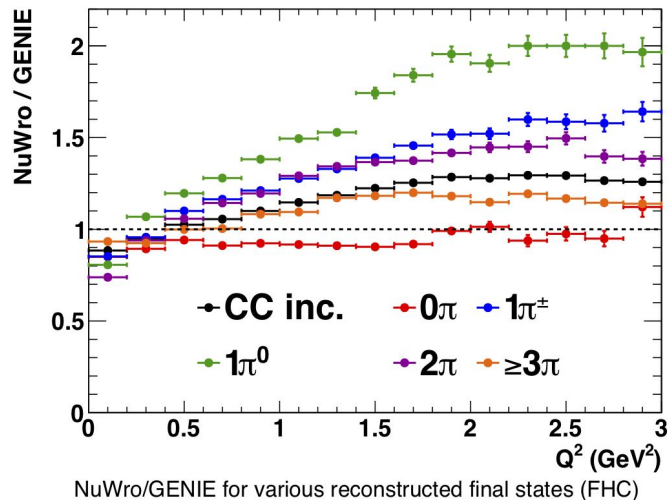
# Separating interaction channels by pion multiplicity in ND-GAr

Each **exclusive samples depends** on various parameters that are heavily **dependent on the interaction model** used

Comparing **two different event generators** we can see the spread on predictions resulting from the different viable models

This gives insight to **select, optimize and tune models** using data to **address the bias** and achieve greater **precision**

True NuWro/GENIE (FHC)

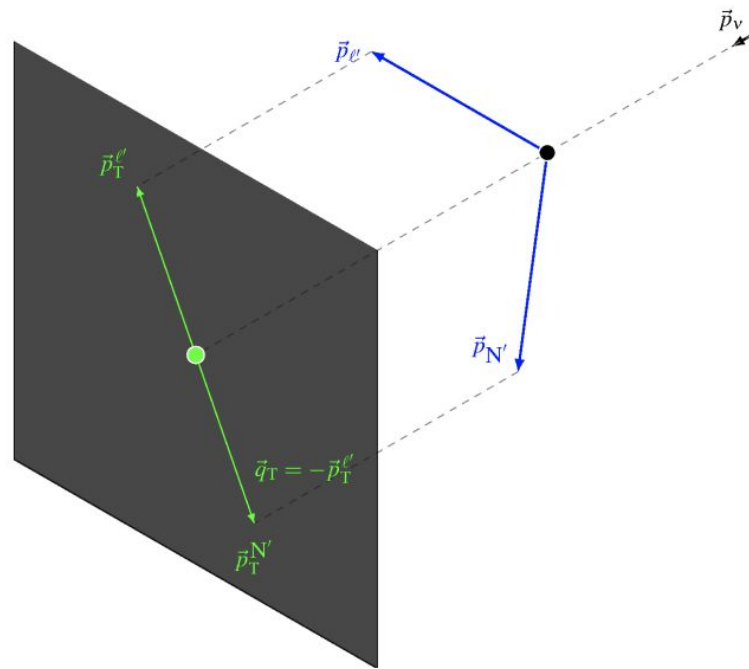


# Investigating Nuclear Effects with Transverse Kinematic Imbalance

TKI is a technique that can use the kinematics of the detected particles to study **the motion of initial-state nucleons** in the nucleus and **FSI**

Considering a final state with only a muon and a proton:

- In a neutrino scattering with a **stationary particle** the **TKI would be zero**



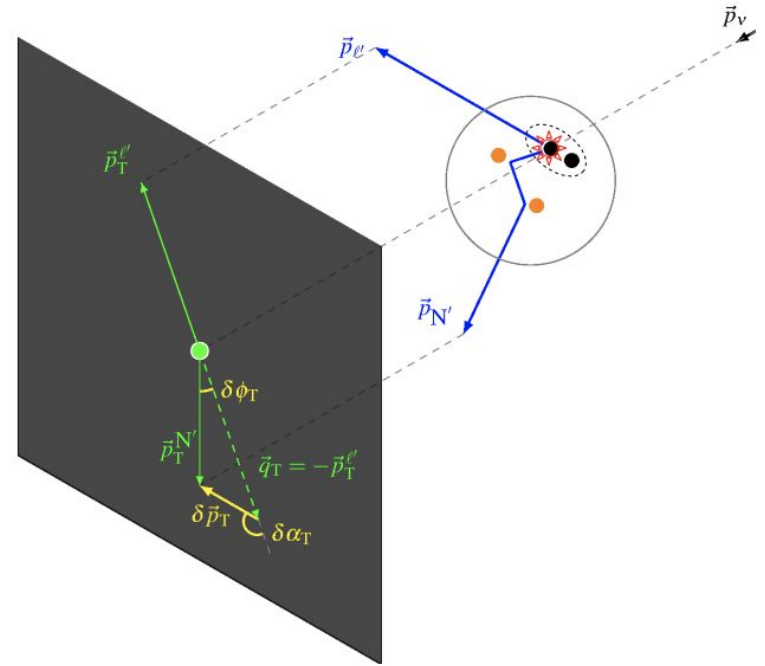
Stationary nucleon target

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Considering a final state with only a muon and a proton:

- In a neutrino scattering with a **stationary particle** the **TKI would be zero**
- thus the **magnitude of the TKI** indicates the **initial momentum of the neutron**
- and the **direction** of the imbalance let us **measure FSI effects**



Nuclear target  
( $A > 1$ )

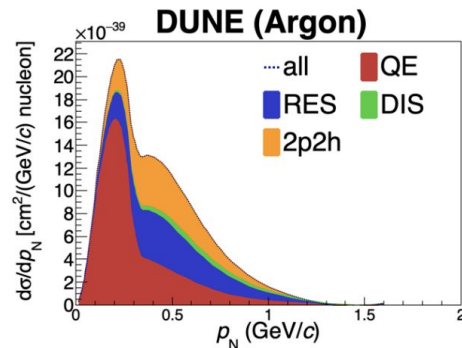
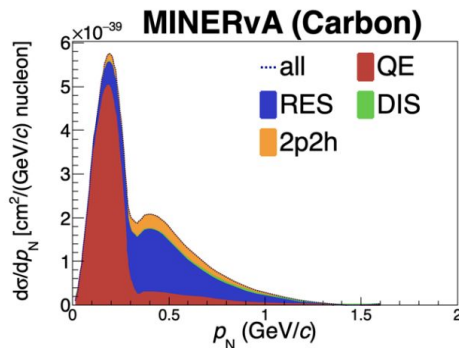
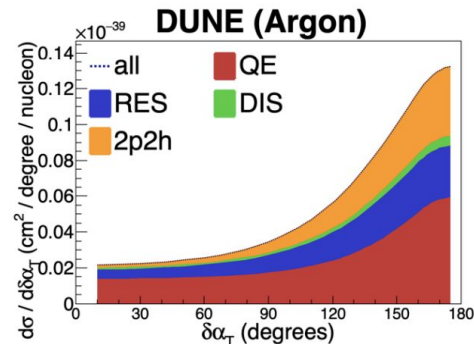
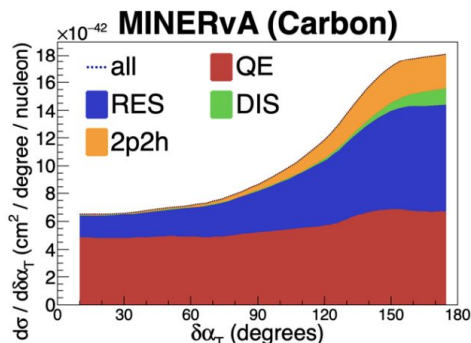
Fermi motion  
Final-state interactions  
Pion absorption  
2p2h  
...

# Investigating Nuclear Effects with Transverse Kinematic Imbalance

Applications of this technique in **MINERvA** and **T2K** showed good agreement with current implemented models but **not in the transition region** between Fermi motion and other nuclear effects

DUNE has a significantly **larger phase space**, with a **full angular acceptance**, and a great momentum acceptance

DUNE ND excellent particle identification will allow the use of TKI to investigate these effects in **argon**, shedding light on **A-dependence**.



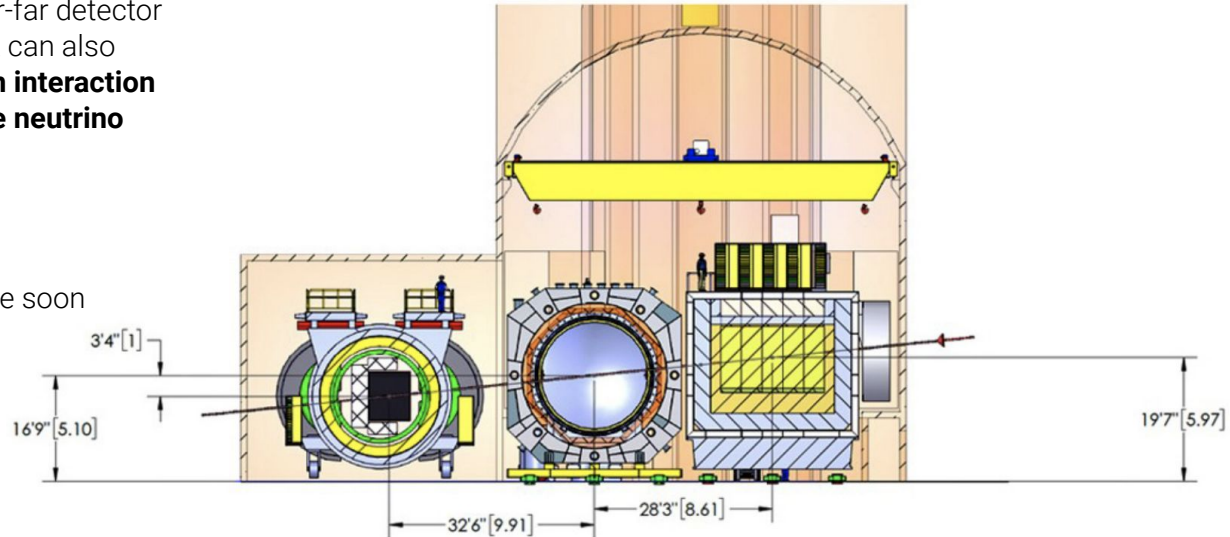
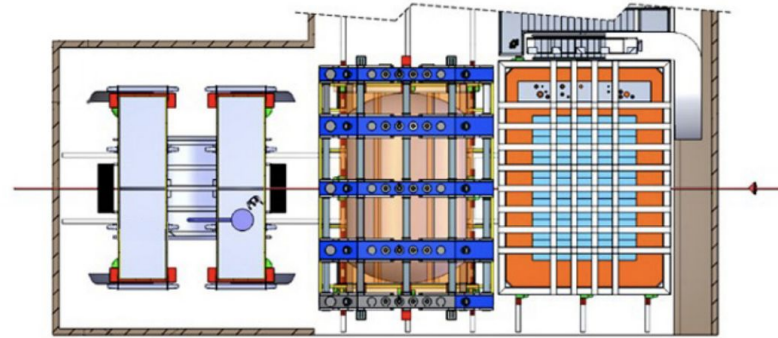
# Summary

Knowing neutrino nucleus cross section is necessary for oscillation experiments.

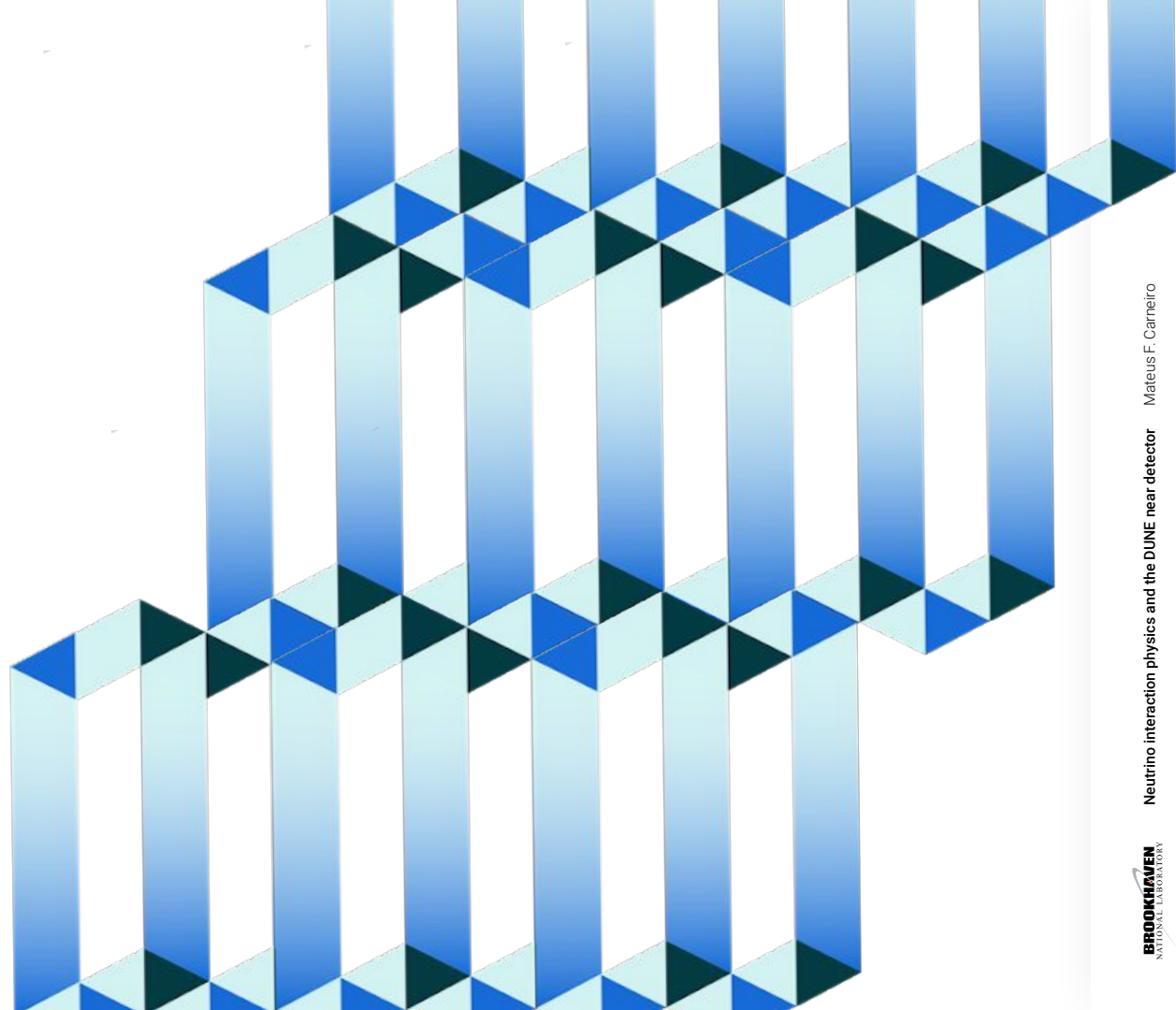
The **DUNE ND** not only provides near-far detector ratios to constrain flux uncertainty, it can also measure cross sections to **constrain interaction models** using **argon** and in the **same neutrino beam** as the FD

**DUNE ND CDR** will be public available soon with many more details

Thank you



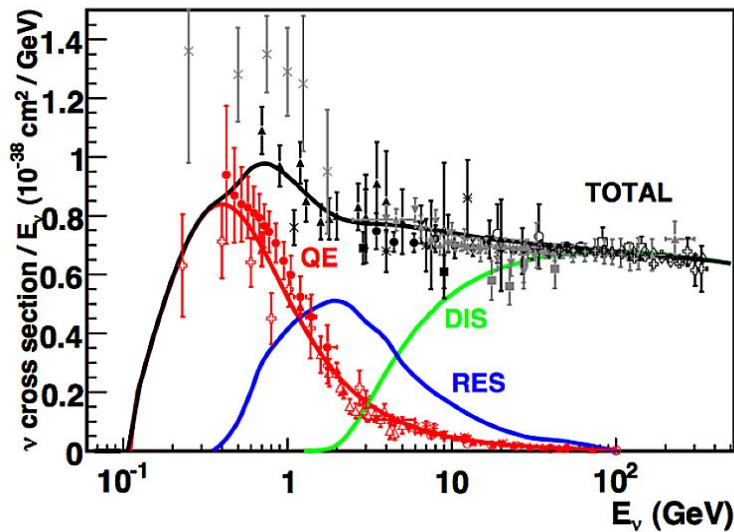
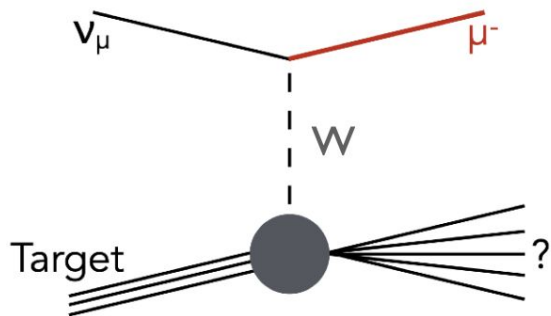
# Backup



# Model validation

Each **generator has a different approach** to almost every interaction channel

Even when considering **exclusive final topologies**, nuclear effects force us to take **every channel into account**



J.A. Formaggio and G.P. Zeller, Rev. Mod. Phys. 84 (2012)

Generator	SBN Experiments	Resonance model	Coherent model	FSI model
NUANCE	MiniBooNE	Rein-Sehgal	Rein-Sehgal KNL-BRS	Cascade
GENIE	MicroBooNE T2K SBN (ND and FD)	Rein-Sehgal KNL-BRS	Rein-Sehgal Bergel-Sehgal	INTRANUKE/hA
NEUT	T2K	Graczyk-Sobczyk	Rein-Sehgal Bergel-Sehgal	Hybrid Oset et al. + exp. Based tune
NuWro	T2K MicroBooNE	Home-grown	Rein-Sehgal Bergel-Sehgal	Cascade



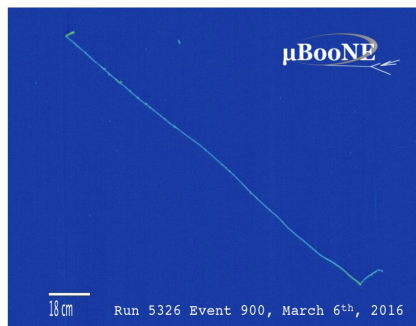
# Particle Identification

Liquid argon TPCs like the **LAr-ND** are sensitive to different final-state hadron topologies and are excellent calorimeters.

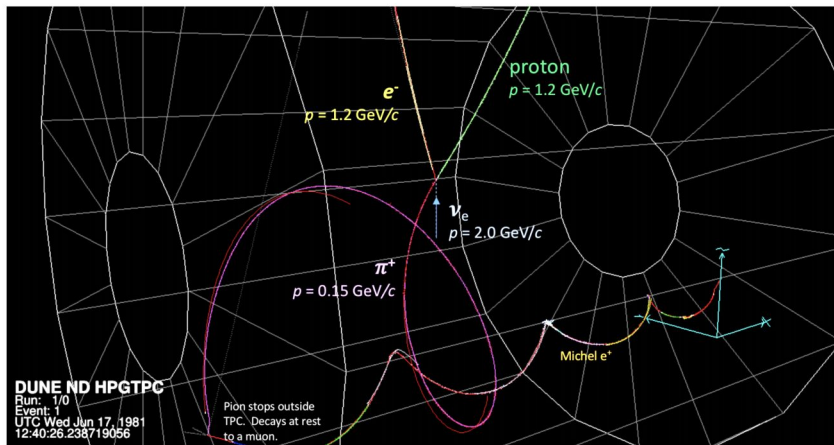
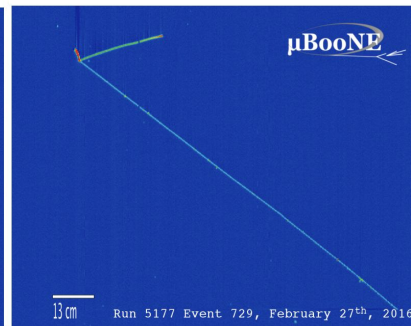
While the magnetized **GAr-ND** can identify the muons that reach the detector and also use its lower density for an even lower energy threshold

**SAND** have a great capacity to identify neutrons which can be very useful for the program

$CC0\pi$



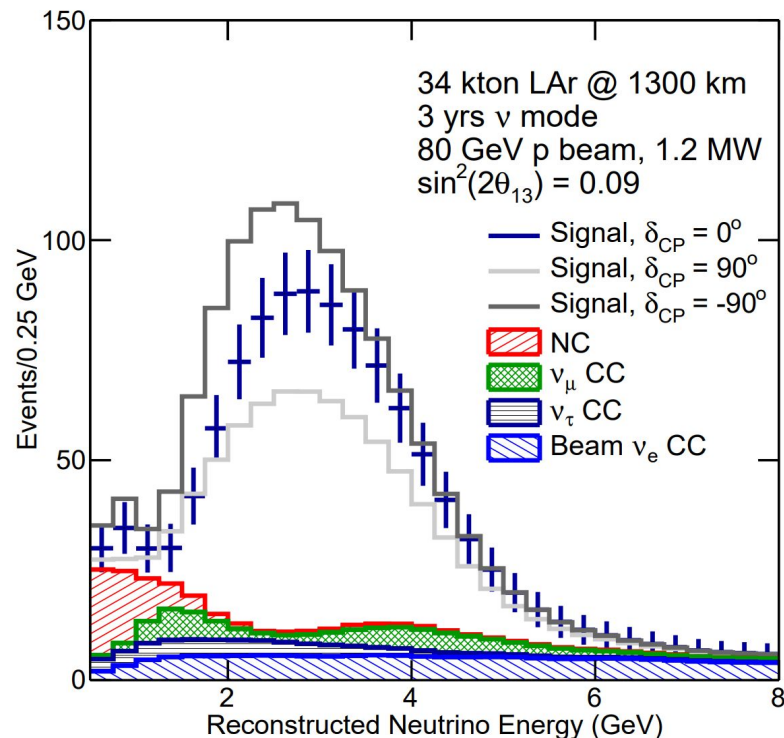
$CC1\pi^\pm$



## How do we find osc parameters?

1. reconstruct topology and energy in the Near Detector
2. use a **nuclear model** to infer the neutrino interaction energy
3. use geometry differences (and oscillation hypothesis) to predict FD flux
4. use the **nuclear model** and the estimated flux to reconstruct topology and energy in the Far Detector
5. compare mc and data and test your hypothesis

LBNE Collaboration  
arXiv:1307.7335



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