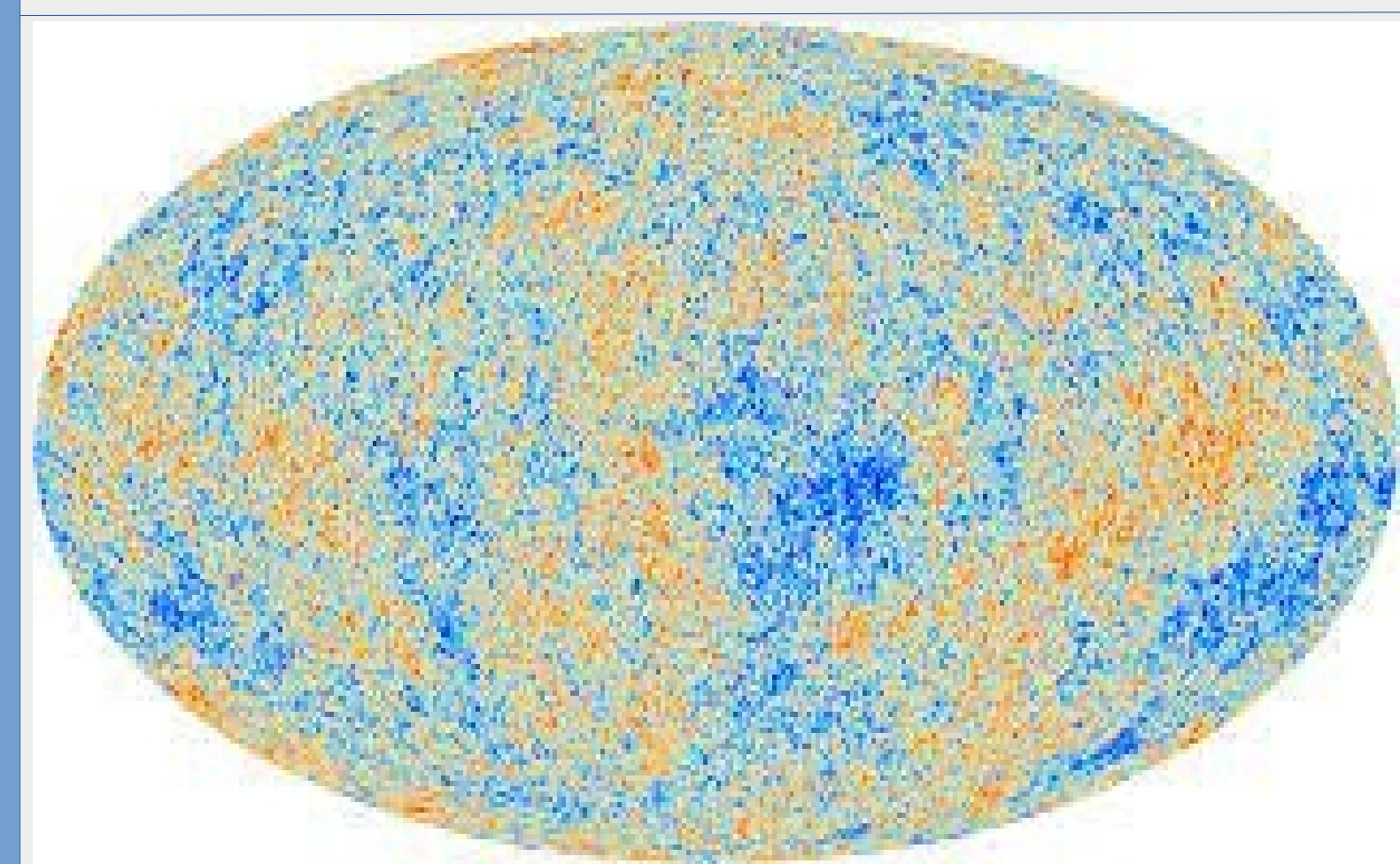


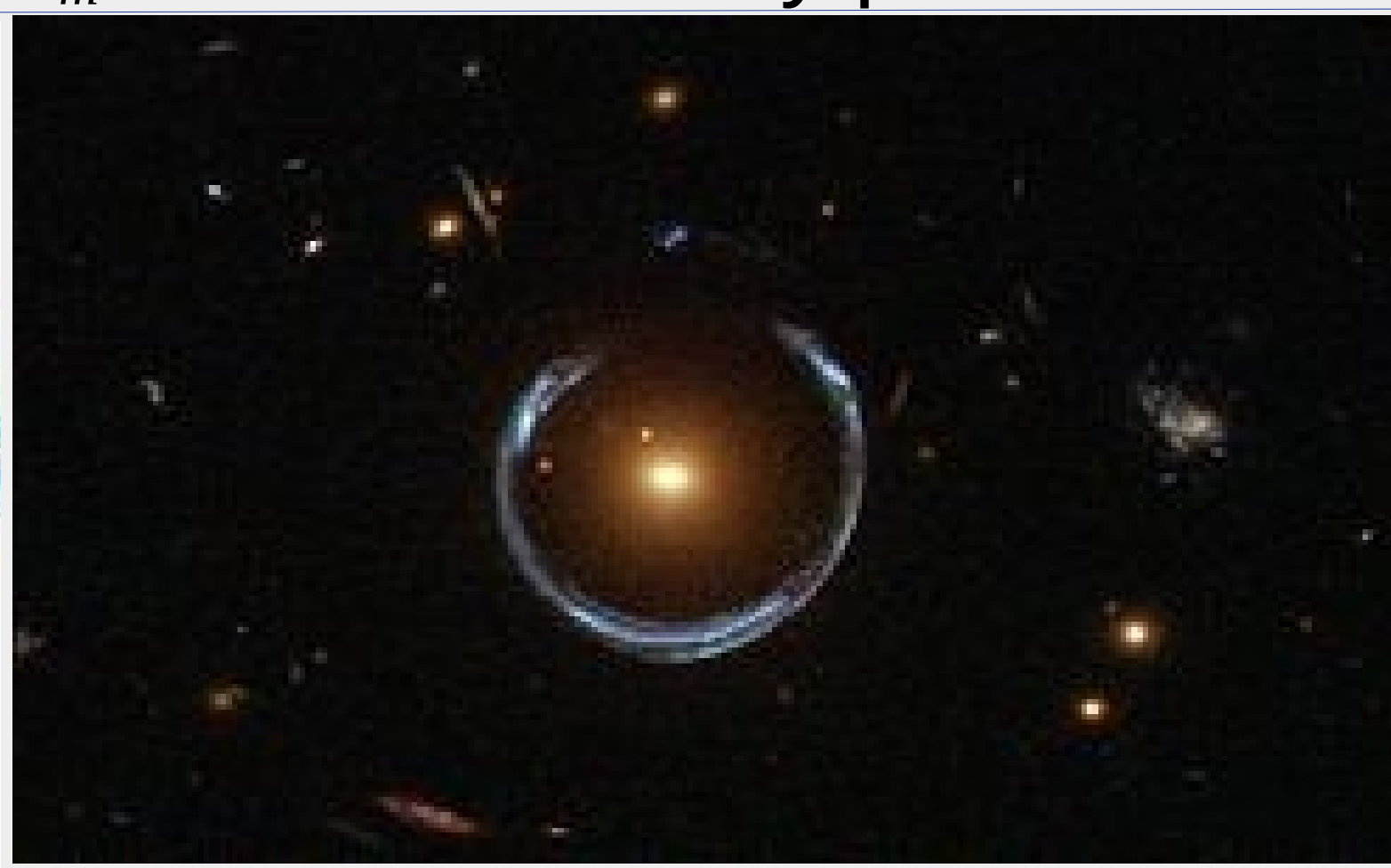
What is S_8 Tension

Defined as $S_8 = \sigma_8 \sqrt{\frac{\Omega_m}{3}}$

Where σ_8 is the root mean square of matter fluctuations on a 8Mpc/h scale
 Ω_m is matter density parameter



Planck CMB
 $S_8 = 0.832 \pm 0.013$



KIDS/VIKING+DES
 $S_8 = 0.755^{+0.019}_{-0.021}$

Tension of 3σ

Solution

Just a massive thermal neutrino not enough

Nonthermal Dark Matter

Nonthermal sterile particle produced from moduli decay

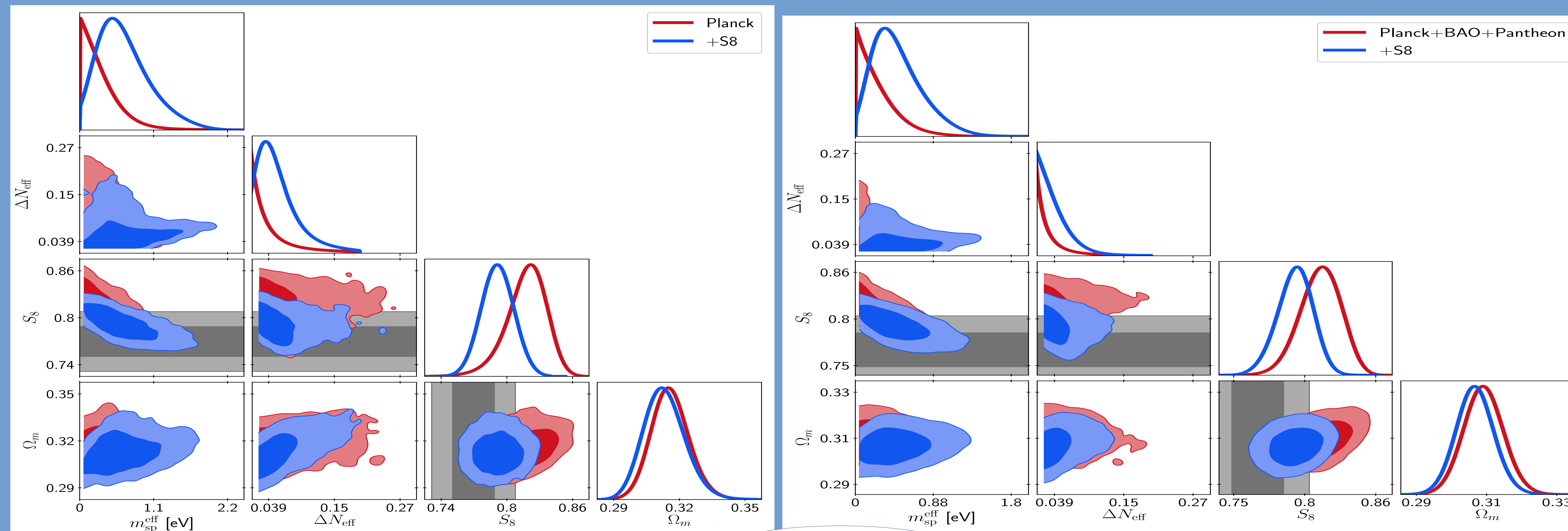
Sterile particle obeying Nonthermal Dodelson widrow distribution

Thermal sterile particle with a different temperature in the hidden section

All three Non-thermal models have same cosmological implications

Model	Non-thermal		Thermal		Dodelson Widrow	
Data set	m_{sp} [eV]	B_{sp}	m_{sp} [eV]	$\frac{T_s}{T_\nu}$	m_{sp} [eV]	χ
Planck	0.05	0.01	0	0.40	0	0.03
Planck+ S_8	38.62	0.012	11.36	0.43	26.43	0.03
Planck+Ext	18.98	0.01	04.59	0.36	12.85	0.02
Planck+Ext+ S_8	39.81	0.01	11.75	0.43	27.49	0.03

MCMC Analysis Results

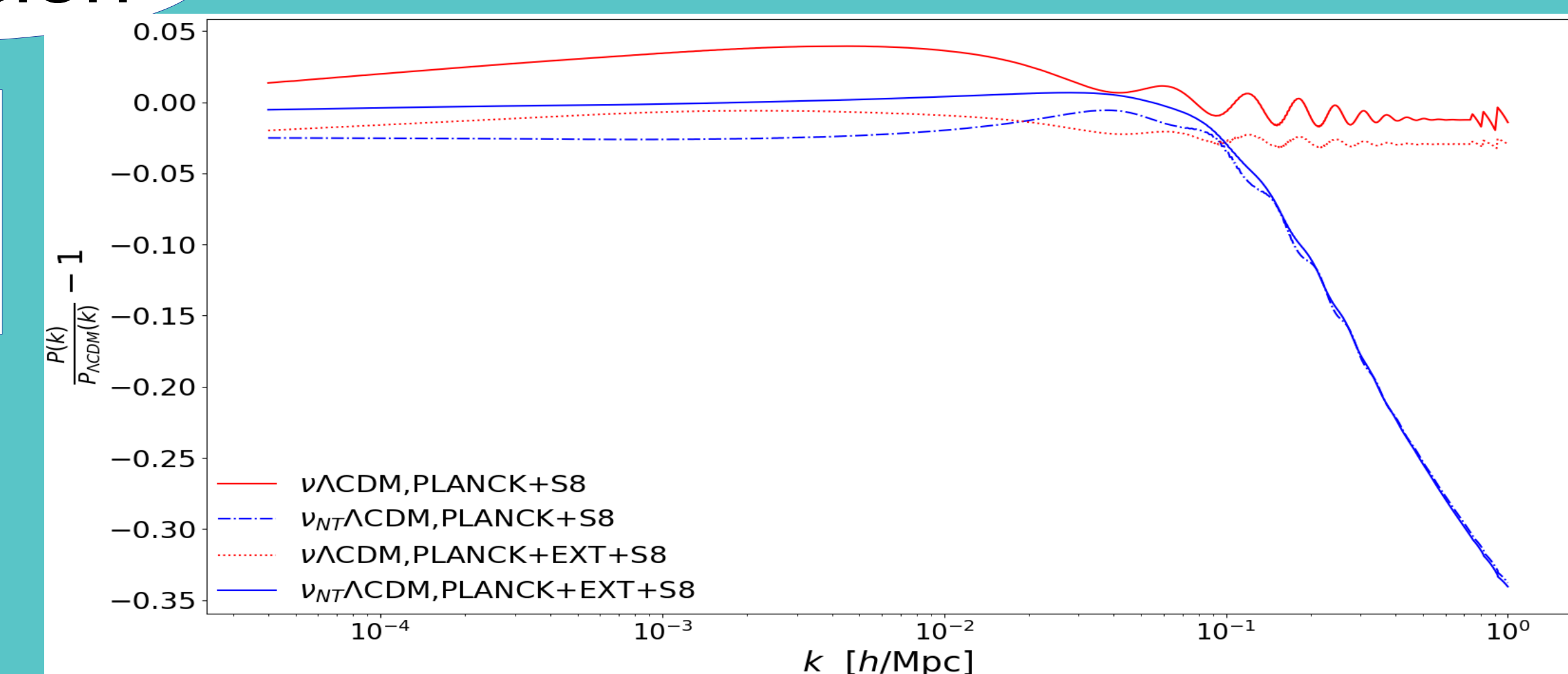


Conclusion

Planck data+S8 Prior we get $S_8(v_{NT} \Lambda CDM) = 0.816^{+0.022}_{-0.016}$ i.e. value get 1 sigma downward shift compared to LCDM. The chi2 is also improved by -4.8.

Inclusion of BOSS+Pantheon with Planck and S8 prior we get. $S_8(v_{NT} \Lambda CDM) = 0.795^{+0.015}_{-0.013}$. The tension get alleviated. However the inclusion of S8 prior in with these additional data sets are not favoured by chi2 numbers

We have got constraints on other nonthermal models mainly Dodelsonwidrow model and thermal hidden sector with a different temperature. Constraint shown in table



Compared to standard thermal neutrinos, $v_{NT} \Lambda CDM$ leads to a much stronger suppression in the matter power spectrum at late-times thus significant decrease in S_8 .

Future Goals

In future ,improving BAO data sets may show better consistency with our model

Analysis including results of short base line neutrino oscillation

In future N-body simulations may help to distinguish among all three Nonthermal models

Acknowledgement

I would like to thank my collaborators Prof subinoy Das , Prof Anshuman Maharana and Prof Vivian Poulin.