Dark Matter, Baryogenesis, and Gravitational Waves from Axion Rotations

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Collaborators:

arXiv: 1910,02080 Keisuke Harigaya Phys. Rev. Lett. 124, 111602 (2020) Phys. Rev. Lett. 124, 251802 (2020) arXiv: 1910.14152 Lawrence Hall, Keisuke Harigaya JHEP 01 (2021) 172 arXiv: 2006,04809 Lawrence Hall, Keisuke Harigaya arXiv: 2006.05687 Nicolas Fernandez, Akshay Ghalsasi, Lawrence Hall, Keisuke Harigaya arXiv: 2104.02077 Keisuke Harigaya and Aaron Pierce



storyline

Axion

(0) Misalignment mechanism Preskill, Wise, Wilczek 1983, Abbott, Sikivie 1983, Dine, Fischler 1983
(1) Parametric resonance RC, L. Hall, K. Harigaya 2017 K. Harigaya, J. Leedom 2019
(2) - Kinetic misalignment mechanism RC, L. Hall, K. Harigaya 2019
- Axiogenesis RC, K. Harigaya 2019
- ALP cogenesis RC, L. Hall, K. Harigaya 2020
- Lepto-Axiogenesis RC, N. Fernandez, A. Ghalsasi, L. Hall, K. Harigaya 2020 Possible initial conditions

(0)

Peccei-Quinn symmetry

Dark Matter

Baryon Asymmetry

storyline

Axion

Misalignment mechanism (0)Preskill, Wise, Wilczek 1983, Abbott, Sikivie 1983, Dine, Fischler 1983 (1)Parametric resonance RC, L. Hall, K. Harigaya 2017 K. Harigaya, J. Leedom 2019 (2)- Kinetic misalignment mechanism RC, L. Hall, K. Harigaya 2019 - Axiogenesis RC, K. Harigaya 2019 - ALP cogenesis RC, L. Hall, K. Harigaya 2020 - Lepto-Axiogenesis RC, N. Fernandez, A. Ghalsasi, L. Hall, K. Harigaya 2020



lisalignment

see later slides for axion-electron and axion-nucleon couplings

Axions



Rotation



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Why Robation?

Large field value : Flat potential

For example, as an initial condition or set dynamically by the Hubble-induced mass

 $V(|P|) \sim -H_I^2 |P|^2 + \frac{|P|^{2d}}{M^{2d-4}}$

Initial condition

 $P = \frac{S + f_a}{\sqrt{2}} e^{i\frac{a}{f_a}}$

Why Robation?

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 $V(|P|) \sim -H_I^2 |P|^2 + \frac{|P|^{2d}}{M^{2d-4}}$

Angular motion : Explicit PQ breaking

$$V(P) \sim \frac{P^n}{M^{n-4}} + \text{h.c.}$$

expected from quantum gravity or PQ as an accidental symmetry

S. Giddings et al. 1988, S. Coleman 1988, G. Gilbert 1988, D. Harlow et al. 2019 R. Holman 1992, S. Barr 1992, M. Kamiokowski 1992, D. Dine 1992

Dynamics analogous to that in Affleck-Dine baryogenesis

I. Affleck and M. Dine 1991

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PRL 124, 111602 (2020) RC and K. Harigaya

Initial condition $P = \frac{S + f_a}{\sqrt{2}} e^{i \frac{a}{f_a}}$

Asymmetry of PQ Charge

Noether charge associated with the shift symmetry

 $n_{\rm PQ} = iP\dot{P^*} - iP^*\dot{P}$ $n_{\rm PQ} = S^2\dot{\theta}$

PQ asymmetry PQ charge density = Rotation of PQ field

PQ charge is conserved soon after the onset.

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 $P = \frac{S + f_a}{\sqrt{2}} e^{i\frac{a}{f_a}}$

Why a large angular speed?

Reason:

 $n_{\rm PQ} = S^2 \dot{\theta}$ $n_{\rm PQ} R^3 = \text{conserved charge}$

Conventional:

$$S^2 = f_a^2 \qquad \dot{\theta} \propto R^{-3}$$

Our scenario (S >> fa):

 $\begin{cases} \text{quartic} & S^2 \propto R^{-2} & \dot{\theta} \propto R^{-1} \\ \text{quadratic} & S^2 \propto R^{-3} & \dot{\theta} = \text{constant} \end{cases} \\ \end{cases} \\ \end{cases} \\ \end{cases} \\ \text{Slower redshift!}$

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(Misalignment + non-zero kinetic energy)

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Kinetic Misalignment Mechanism

Consequence: delaying usual $T_{
m osc}$ until KE = PE, enhancing the dark matter abundance: Abundance: $\Omega_a h^2 \propto \dot{ heta}$

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Kinetic Misalignment Mechanism



(QCD axion + baryogenesis)

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 $n_B = c_B \dot{\theta} T^2$ $c_B \simeq 0.1 - 0.15 \, c_W$

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 $\begin{vmatrix} n_B = c_B \dot{\theta} T^2 \\ c_B \simeq 0.1 - 0.15 c_W \end{vmatrix} Y_B \equiv \frac{n_B}{s} = \frac{c_B \dot{\theta} T^2}{s} \bigg|_{T = T_{\rm EW}} = c_B Y_{\rm PQ} \left(\frac{T_{\rm EW}}{f_a}\right)^2$

Baryon asymmetry fixes rotational speed, equivalently Y_{PQ} . R. Co Minnesota PRL 124, 111602 (2020) RC and K. Harigaya



$$Y_B \simeq 10^{-10} \left(\frac{c_B}{0.1}\right) \left(\frac{T_{\rm EW}}{130 \text{ GeV}}\right)^2 \left(\frac{10^8 \text{ GeV}}{f_a}\right)^2 \left(\frac{Y_{\rm PQ}}{500}\right)^2$$

Baryon asymmetry fixes rotational speed, equivalently Y_{PQ} . R. Co Minnesota PRL 124, 111602 (2020) RC and K. Harigaya

Axiogenesis + Kinetic Misalignment

$$\frac{\Omega_{a}h^{2}}{\Omega_{DM}h^{2}} \simeq 70 \left(\frac{f_{a}}{10^{8} \text{ GeV}}\right) \left(\frac{130 \text{ GeV}}{T_{EW}}\right)^{2} \left(\frac{0.1}{c_{B}}\right)$$

$$T_{EW} = 1 \text{ TeV} \left(\frac{f_{a}}{10^{8} \text{ GeV}}\right)^{\frac{1}{2}}$$

$$\frac{Axion}{physics}$$

$$\frac{Axion}{physics}$$

$$T_{th} = thermatization temperature of saxions$$

$$\xi = 10^{-2} \times \left(\frac{T_{EW}}{130GeV}\right)^{2} \left(\frac{c_{B}}{0.1}\right)$$

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Axion Photon Coupling



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storyline

QCD axion

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(0)

(2)

Peccei-Quinn symmetry

Dark Matter

Baryon Asymmetry

ALP-genesis



$$Y_B \simeq 10^{-10} \left(\frac{c_B}{0.1}\right) \left(\frac{T_{\rm EW}}{130 \text{ GeV}}\right)^2 \left(\frac{10^8 \text{ GeV}}{f_a}\right)^2 \left(\frac{Y_{\rm PQ}}{500}\right)$$

Baryon asymmetry fixes rotational speed, equivalently Ypg. PRL 124, 111602 (2020) RC and K. Harigaya

Axion Photon Coupling



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Axion Nucleon Coupling



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Axion Electron Coupling



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Axion with Majorana neutrinos

 $\dot{\theta} \neq 0$



 $n_B \propto \dot{\theta} T^2$

 $\mathcal{L} = \frac{m_{\nu}}{2v_{\rm EW}^2} \ell \ell H^{\dagger} H^{\dagger}$

2006.03148 V. Domcke, Y. Ema, K. Mukaida, M. Yamada 2006.05687 RC, N. Fernandez, A. Ghalsasi, L. Hall, K. Harigaya

Lepto-Axiogenesis



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Lepto-Axiogenesis

Producing L at high temperatures

Converting to B at TEW



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Axion Photon Coupling



Kinetic Misalignment Mechanism

StoryLine

Axion and dark photon coupling

> Gravitational waves

> Dark photon dark matter

Tachyonic Instability

 ψ m A' ϕ

 $\mathcal{L}_{dark} = \frac{\alpha_D}{8\pi} \frac{\phi}{f_{\phi}} F'_{\mu\nu} \tilde{F'}^{\mu\nu}$

 $F'_{\mu\nu} = \partial_{\mu}A'_{\nu} - \partial_{\nu}A'_{\mu}$ $\widetilde{F'}^{\mu
u} = \overline{\epsilon^{lphaeta\mu
u}F'_{lphaeta}}/2$

Tachyonic Instability

Carroll and Field 1991 + Garretson 1992, Ratra 1992 Felder, Garc´ıa-Bellido, Greene, Kofman, Linde, Tkachev 2001

$$\mathcal{L}_{\text{dark}} = \frac{\alpha_D}{8\pi} \frac{\phi}{f_\phi} F'_{\mu\nu} \tilde{F'}^{\mu\nu}$$

$$\frac{\partial^2 A'_{\pm}}{\partial t^2} + H \frac{\partial A'_{\pm}}{\partial t} + \left(m_{A'}^2 + \frac{k^2}{R^2} \pm \frac{e_D^2}{8\pi^2} \frac{k}{R} \dot{\theta} \right) A'_{\pm} = 0$$

storyline

Axion and dark photon coupling

> Gravitational waves

> Dark photon dark matter

Possible initial conditions

(0

Peccei-Quinn symmetry

(3)

(1)

Gravitational Waves

 $ds^{2} = R(\tau)^{2} \left(d\tau^{2} - \left(\delta_{ij} + h_{ij} \right) dx^{i} dx^{j} \right)$ $\ddot{h}_{ij} + 3H\dot{h}_{ij} - \frac{1}{R^2}\nabla^2 h_{ij} = \frac{2}{M_{\rm Pl}^2}\Pi_{ij}^{TT}$

 $\frac{\rho_{\rm GW}}{\rho_R} = \frac{r_p^2 \rho_{A'}^2}{3k_{\rm TI}^4 M_{\rm Pl}^4}$

the transverse-traceless anisotropic stress tensor

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C. Machado, W. Ratzinger, P. Schwaller, and B. Stefanek arXiv:1811.01950

Gravitational Waves



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arXiv:2104.02077 RC, K. Harigaya, A. Pierce,

storyline

Axion and dark photon coupling

- > Gravitational waves
- > Dark photon dark matter

Possible initial conditions

(0

Peccei-Quinn symmetry

(3)

(1)

Dark Photon Dark Matter

Correlation between GWs and dark photons







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arXiv:2104.02077 RC, K. Harigaya, A. Pierce,

Conclusions

New axion dynamics allows the QCD axion to simultaneously explain

✓ the Strong CP problem
✓ the dark matter abundance
✓ the baryon asymmetry

Possible signatures:
 (QCD) axion searches
 Saxion-Higgs mixing
 Warm axion dark matter
 Gravitational waves

New model building opportunities