

# Neutrino Self-interaction as a solution(?) to the Hubble Tension

Anirban Das

SLAC National Accelerator Laboratory

Based on 2011.12315  
with Subhajit Ghosh

WIv 2021

# Outline

- **Cosmology of self-interacting neutrino**
- **Flavor-universal  $S_{\nu\nu}$  & the Hubble tension**
- **Laboratory constraints on  $S_{\nu\nu}$**
- **Flavor-specific  $S_{\nu\nu}$**

# A phenomenological model

$$\mathcal{L} \supset \frac{1}{2} g_{ij} \bar{\nu}_i \nu_j \phi, \quad g_{ij} = g \delta_{ij}$$



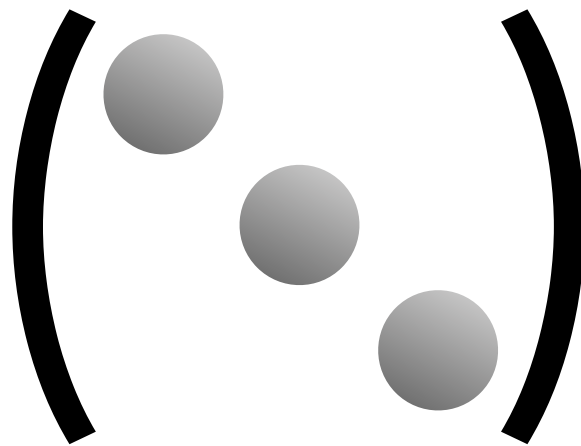
$$\mathcal{L} \supset G_{\text{eff}} \bar{\nu} \nu \bar{\nu} \nu, \quad G_{\text{eff}} = \frac{g^2}{M_\phi^2}$$

1306.1536

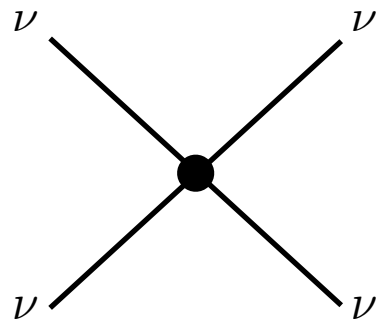
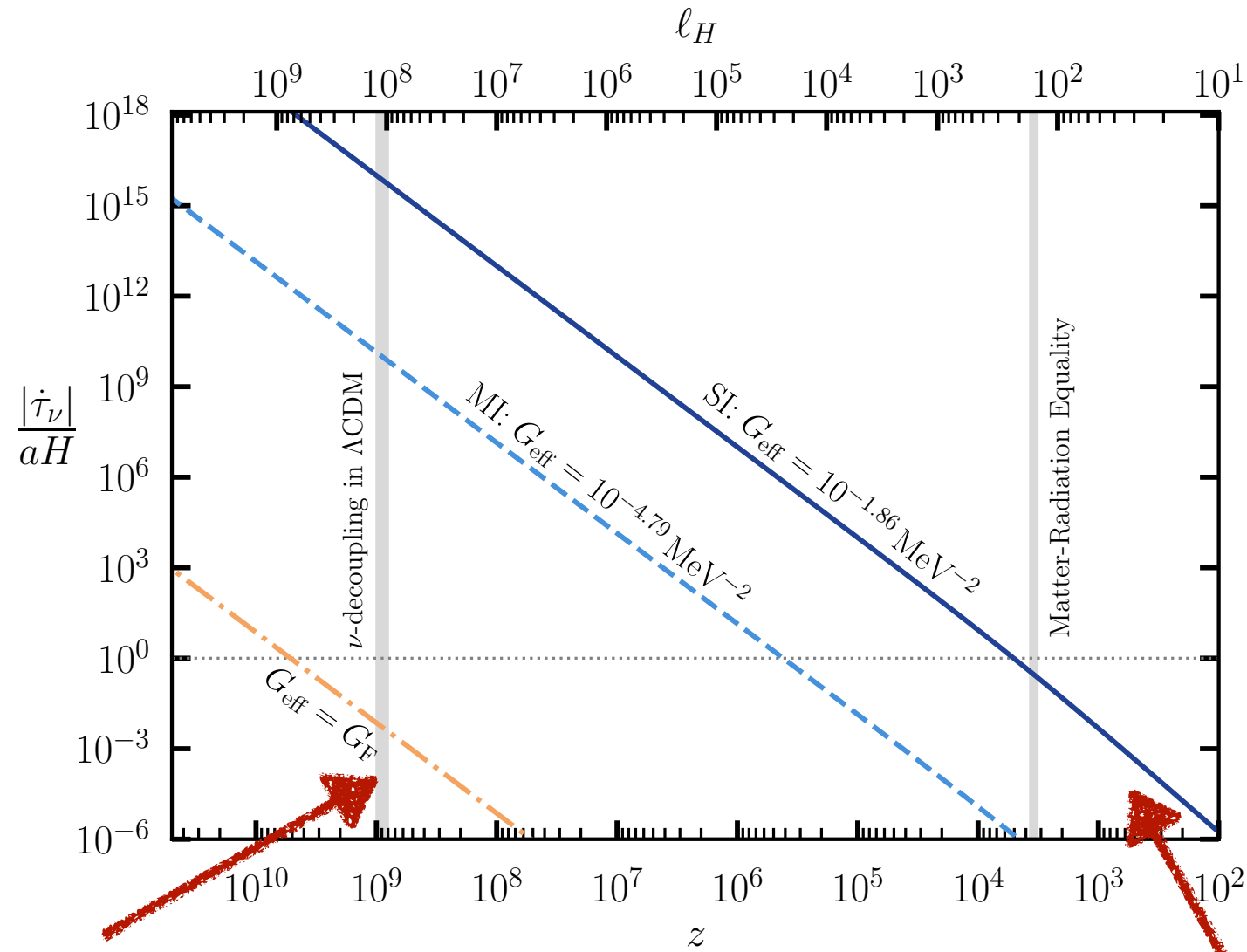
1704.06657

1902.00534

**All neutrino flavors interact with the same strength**

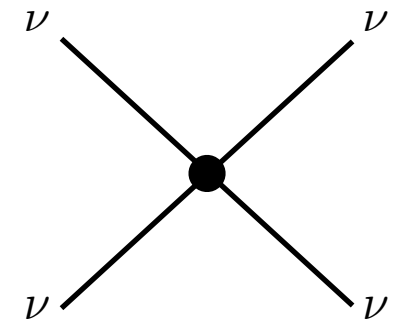


# Cosmology of Slnu



$$\sigma \propto G_F^2$$

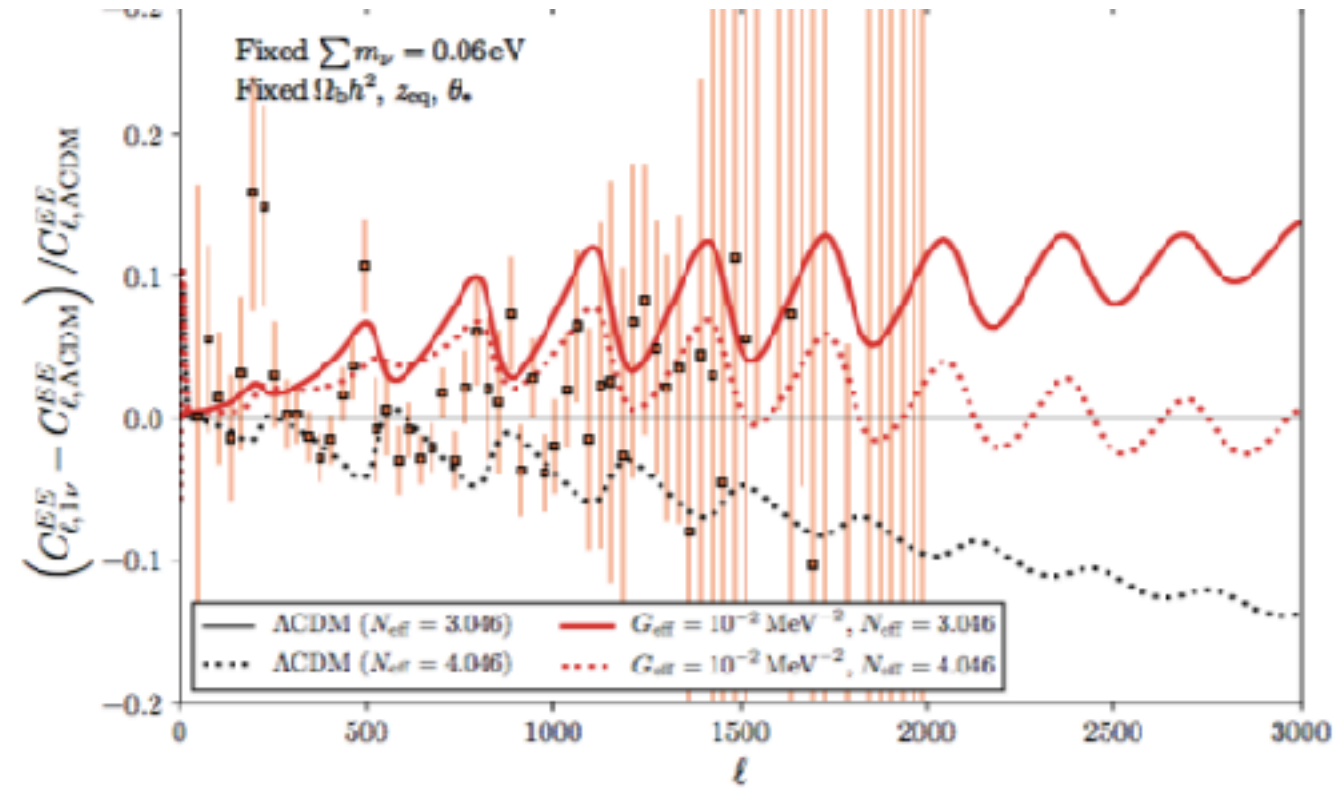
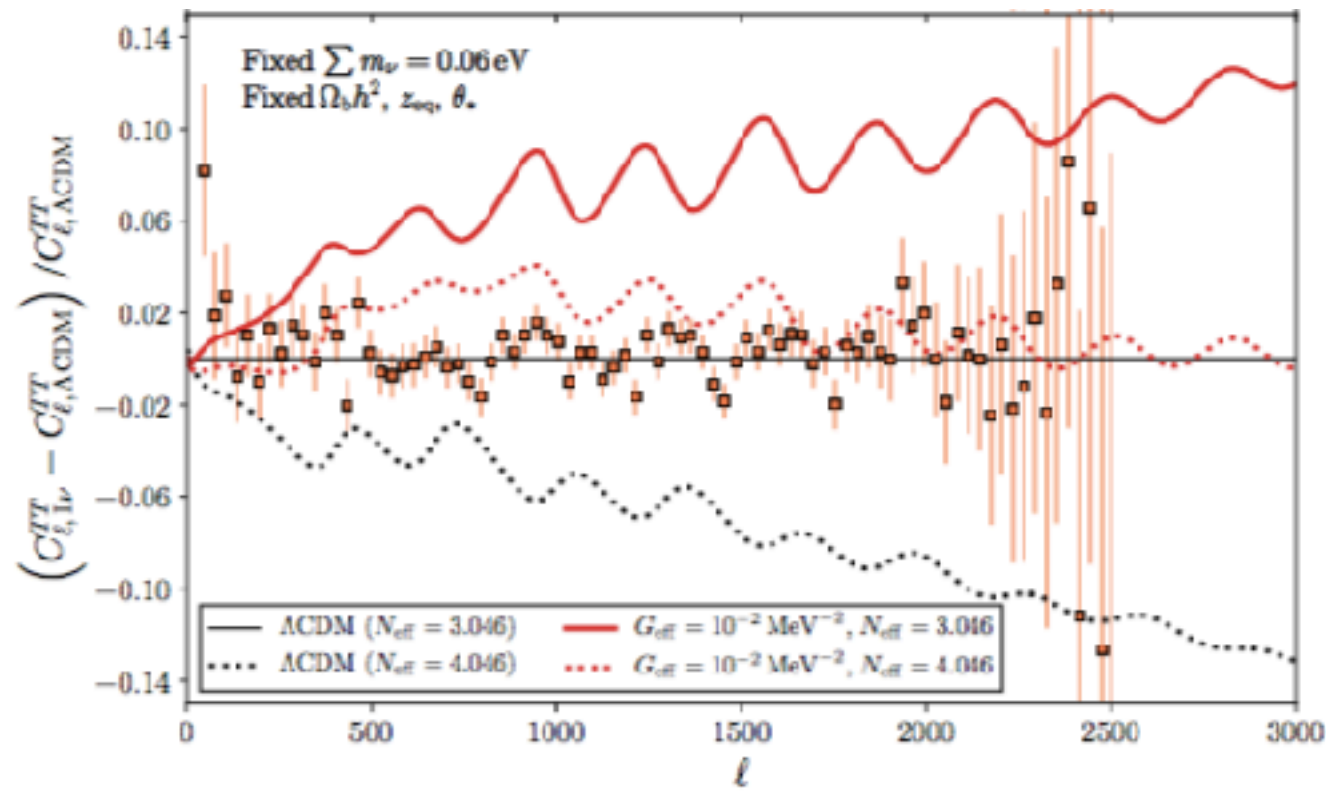
$$G_{\text{eff}} \simeq 10^9 G_F$$



$$\sigma \propto \frac{g_\phi^4}{M_\phi^4} \equiv G_{\text{eff}}^2$$

# Cosmology of Slnu

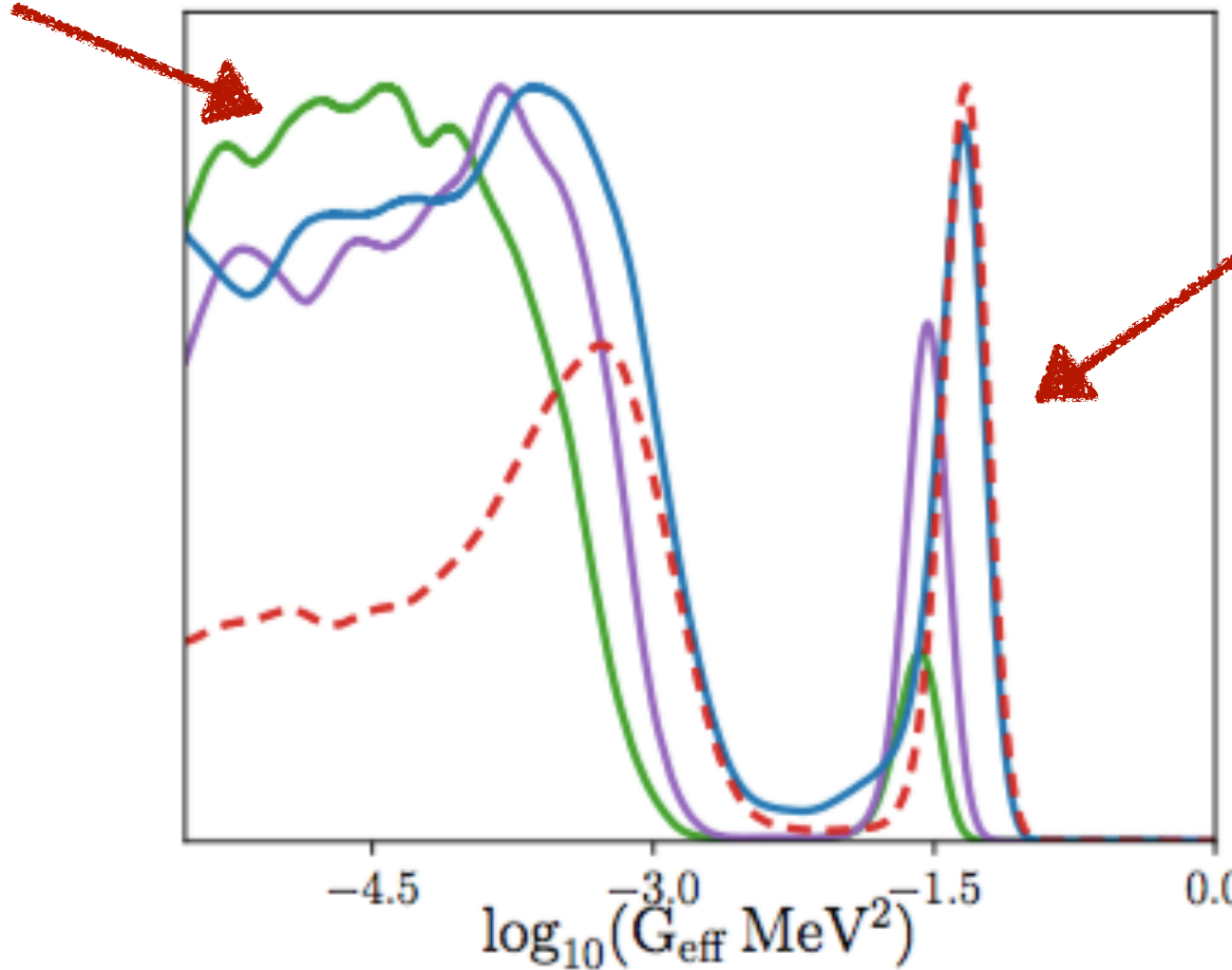
Potentials enhance photon perturbation  $\implies$  larger  $C_l$



Kreisch et al 1902.00534

# Cosmology of Slnu

Moderately Interacting  
(MI)



Strongly Interacting  
(SI)

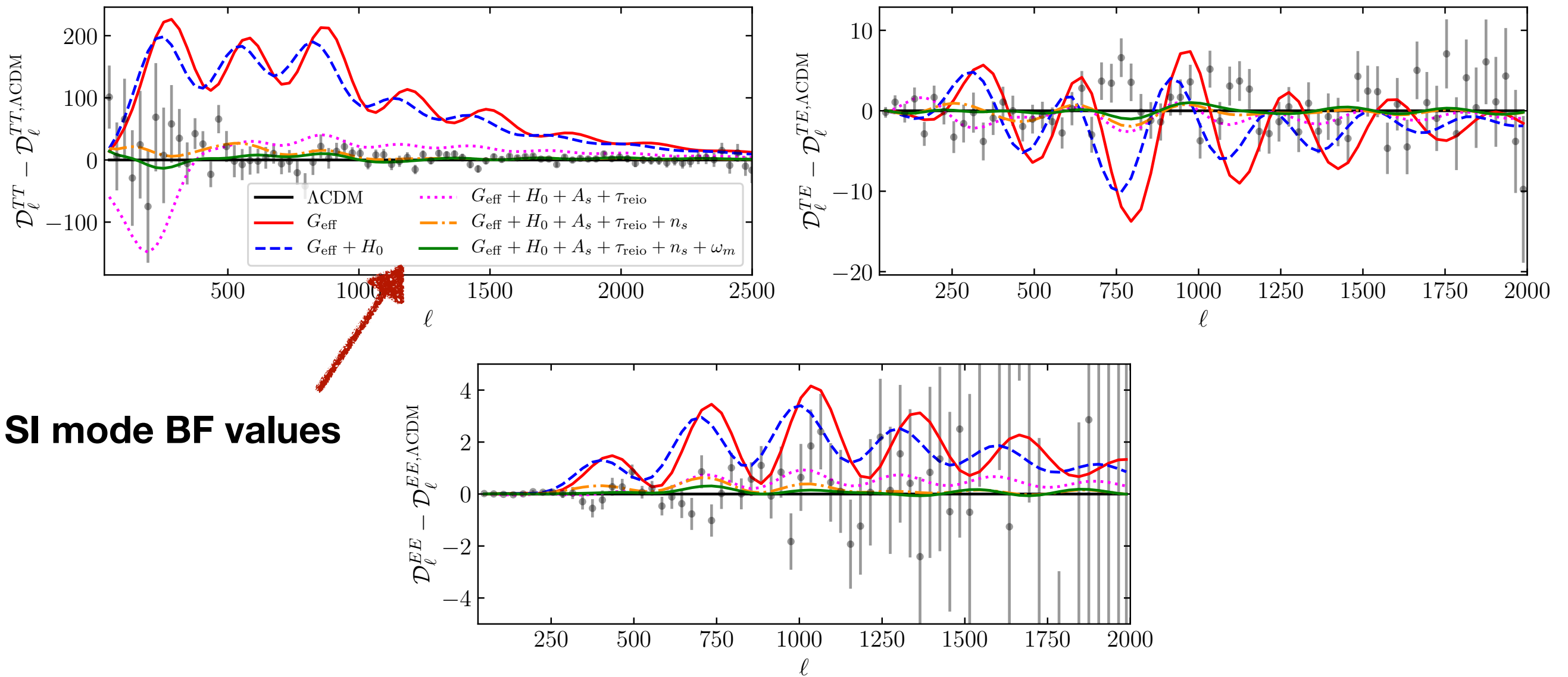
Kreisch et al 1902.00534

Two questions-

- Origin of the SI mode?
- Why is there a valley in between?

# Cosmology of Slnu: $G_{\text{eff}}$ degeneracy w/ other params

P18 : TTTEEE + lowE | 3c + 0f



SI mode BF values

SI mode changes  $l > 200$

# SInu & Hubble

Phase shift  $\rightarrow \phi_\nu \simeq 0.19\pi R_\nu \leftarrow$  Free-streaming neutrino

$$\ell \approx (m\pi - \phi_\nu) \frac{D_A^*}{r_s^*}$$

$$D_A^* = \int_0^{z^*} \frac{1}{H(z)} dz, \quad r_s^* = \int_{z^*}^{\infty} \frac{c_s(z)}{H(z)} dz$$

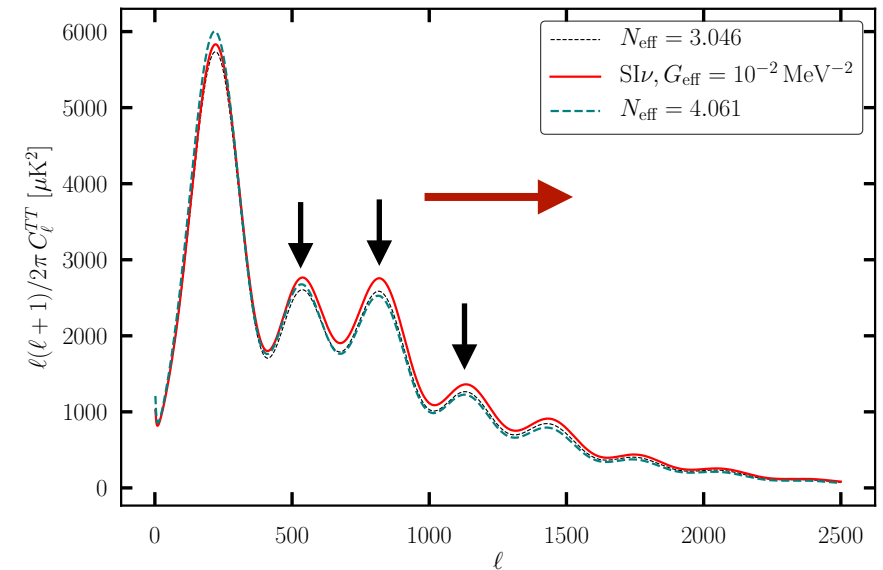
SInu phase shift increases  $\ell$



To keep the spectrum fixed,  $\theta_* = \frac{r_s^*}{D_A^*}$  needs to increase



Larger  $H_0$  decreases  $D_A^*$  & increases  $\theta_*$



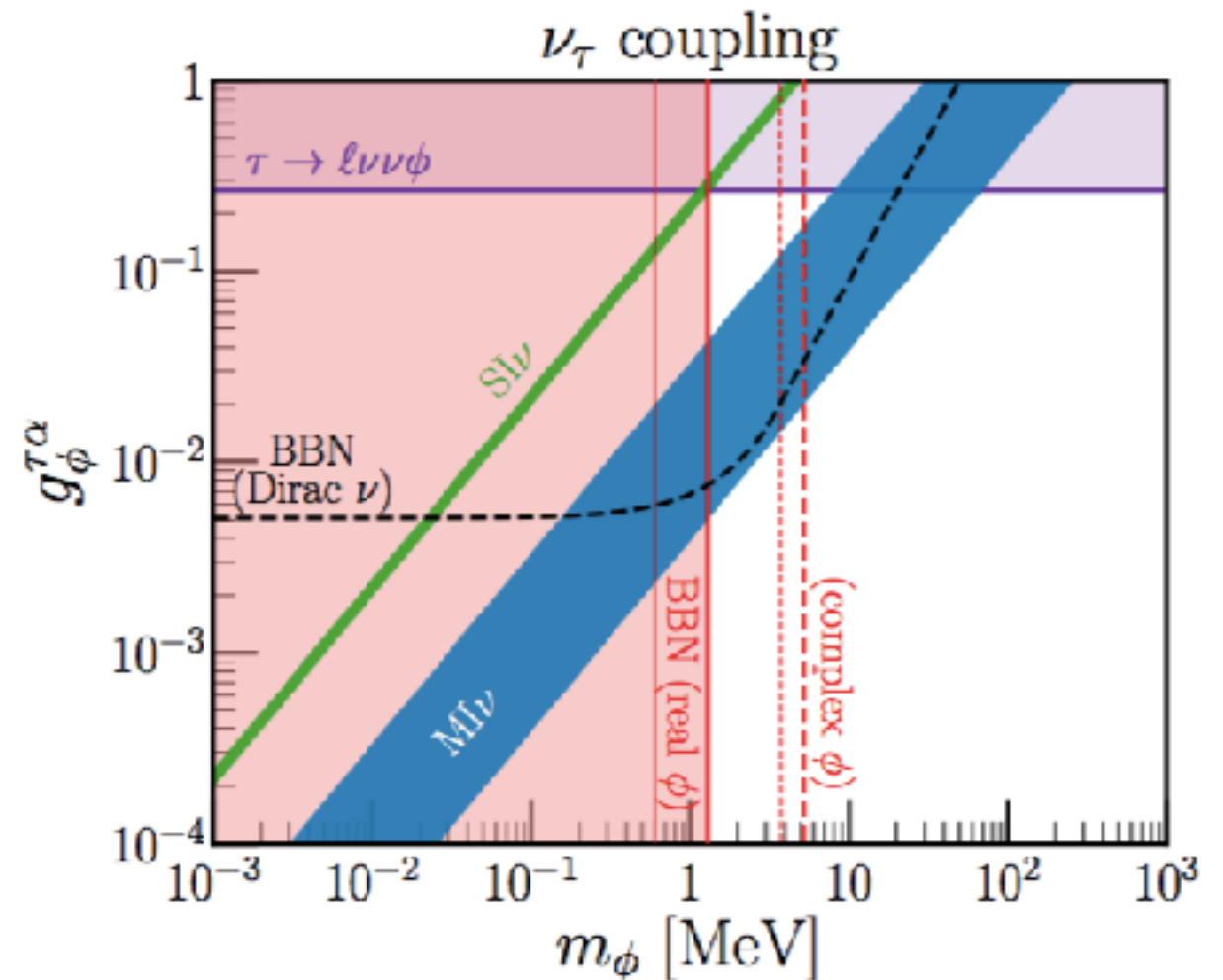
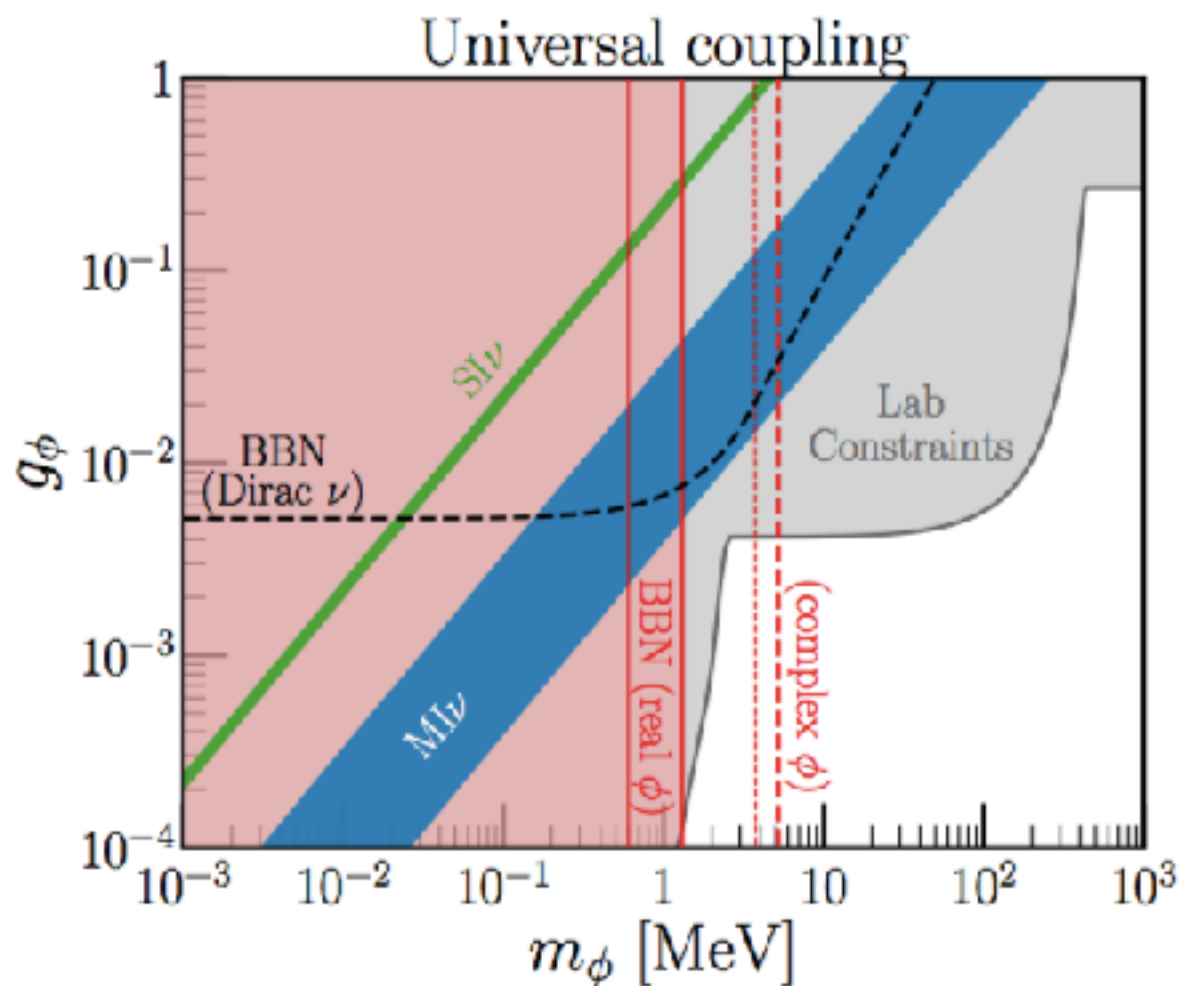


# Laboratory Constraints

Example:  $\nu$  self-interaction could come from a scalar (1905.02727)

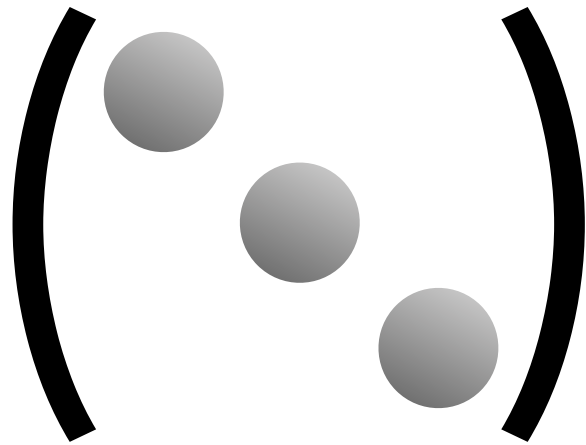
$$\mathcal{L} \supset -\frac{1}{2}m_\phi^2\phi^2 + \frac{1}{2}(g_\phi^{\alpha\beta}\nu_\alpha\nu_\beta\phi + \text{h.c.})$$

$$G_{\text{eff}} \equiv \frac{g_\phi^2}{m_\phi^2} = (10 \text{ MeV})^{-2} \left(\frac{g_\phi}{10^{-1}}\right)^2 \left(\frac{\text{MeV}}{m_\phi}\right)^2$$

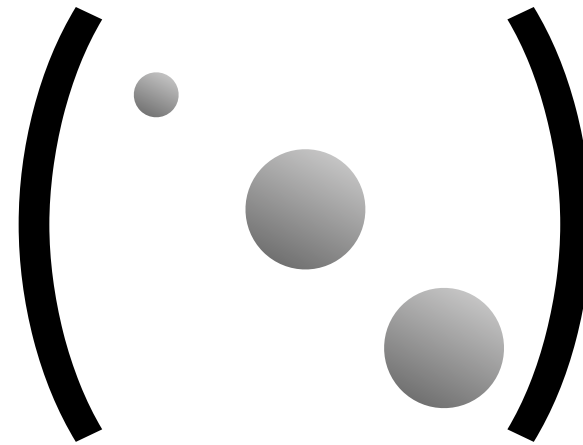


# Flavor-specific $S\nu\nu$

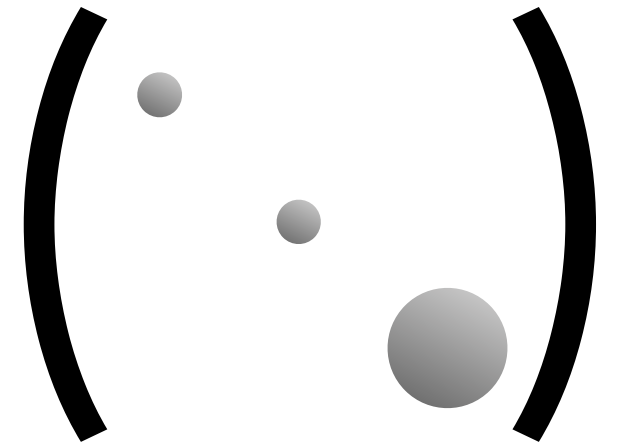
Flavor  
Space



$3c+0f$



$2c+1f$

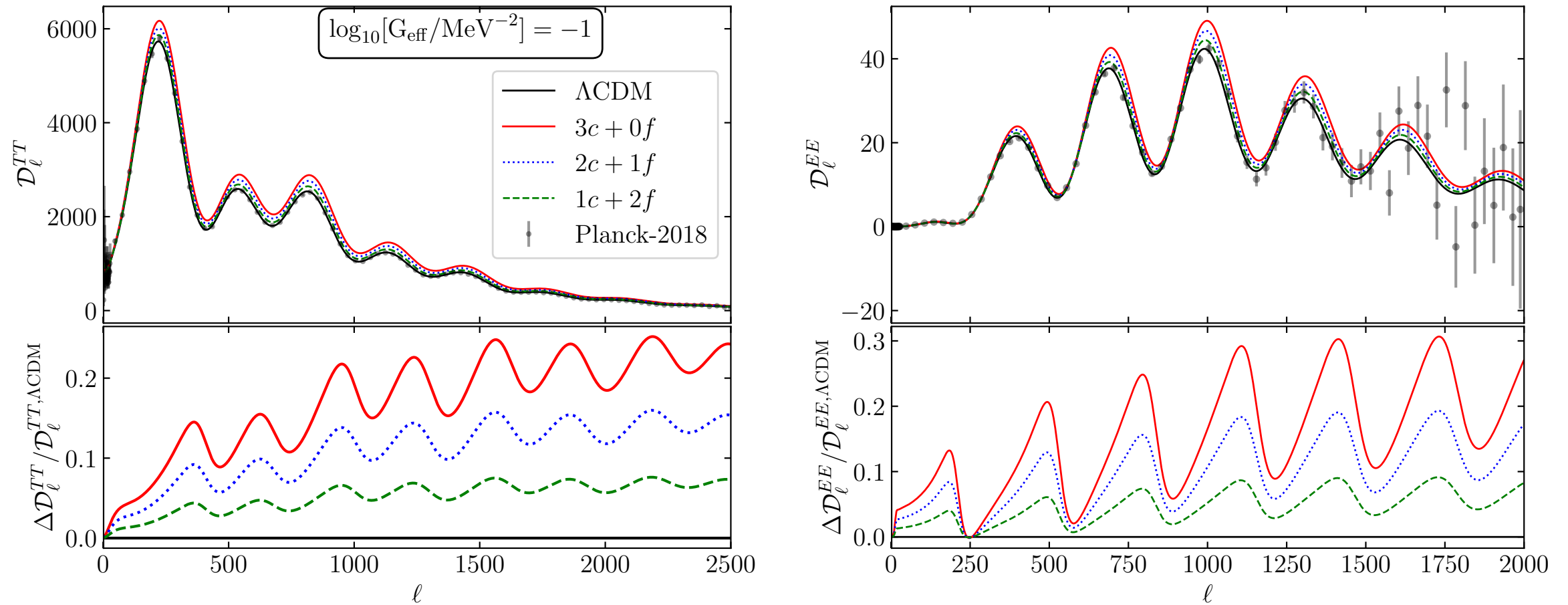


$1c+2f$

$$\mathcal{L} \supset G_{\text{eff}}^{(ijkl)} \bar{\nu}_i \nu_j \bar{\nu}_k \nu_l,$$

$$G_{\text{eff}}^{(ijkl)} \equiv \frac{g_{ij} g_{kl}}{M_\phi^2}$$

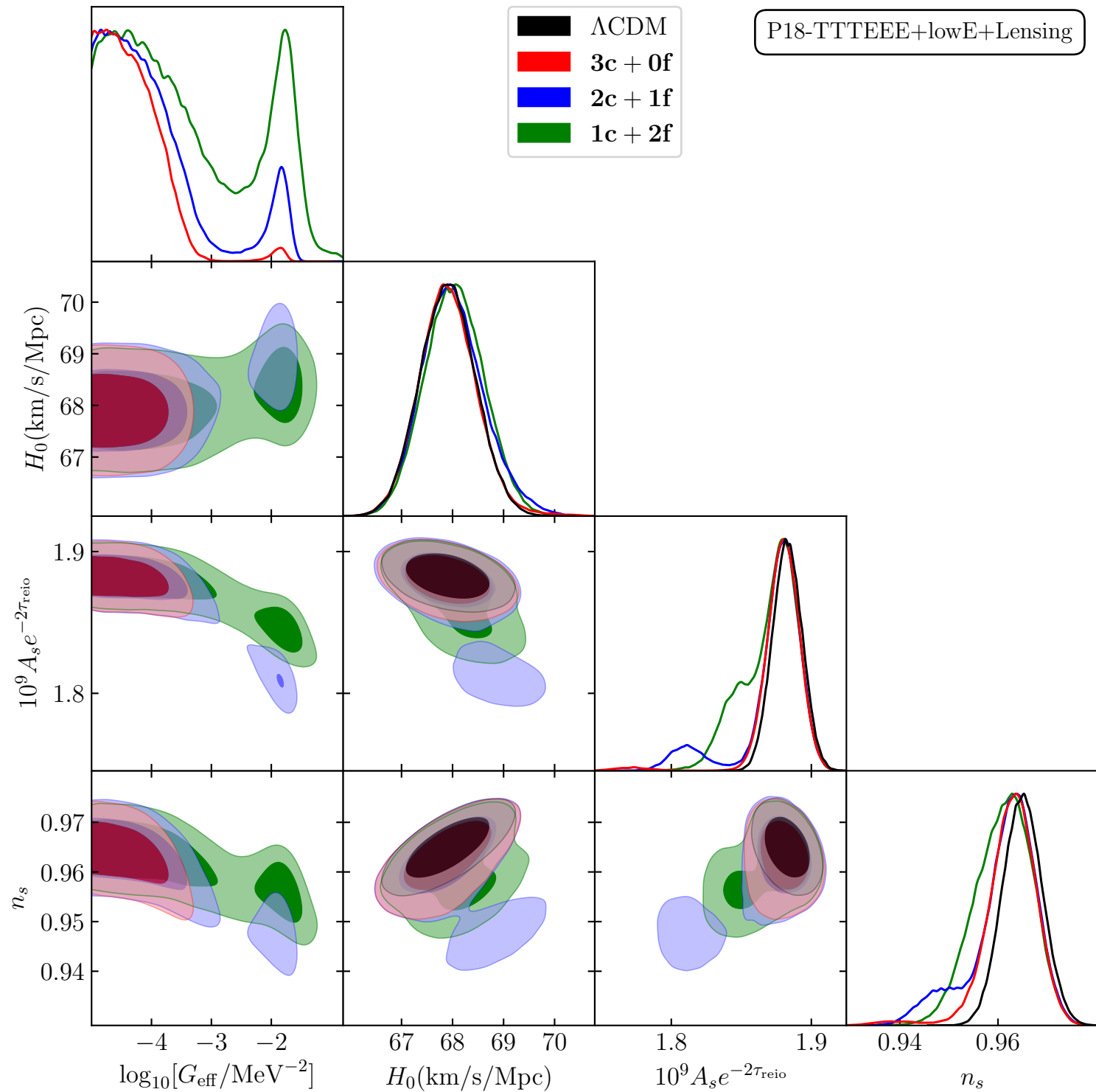
# Flavor-specific $S_{\text{Inu}}$



2011.12315

**Less number of interacting neutrinos reduces the change  
in the CMB power spectra**

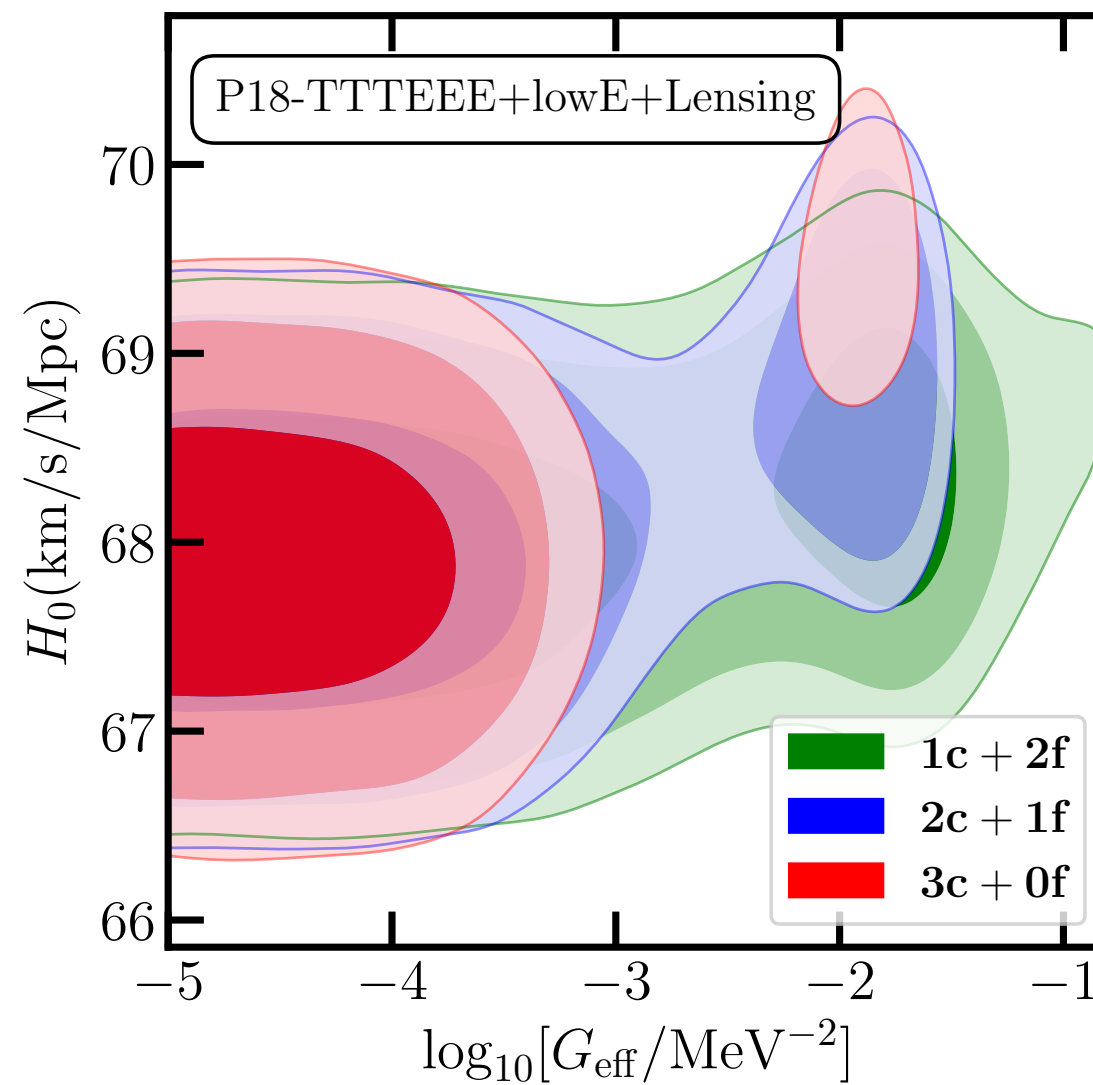
# Flavor-specific Slnu



2011.12315

# Flavor-specific $S_{\text{nu}}$ : effect on $H_0$

Flavor-specific  $S_{\text{nu}}$  yields smaller  $H_0$



$H_0 = 69.5 \pm 0.6$  km/s/Mpc for TTTEEE+lowE+lens (2011.12315)

$H_0 = 69.6 \pm 2.5$  (TRGB, 2002.01550)

$H_0 = 73.2 \pm 1.3$  (Riess et al, 2012.08534)

# Conclusions & Future Outlook

- Flavor-universal  $G_{\text{eff}}$  ( $\sim 10^{-2} \text{ MeV}^{-2}$ ) is incompatible with BBN & laboratory data
- Flavor-specific interaction may still be allowed
- Flavor-specific SI mode coupling does not change
- Changes in CMB spectrum is less, hence, more room to exploit the deg. with  $A_s, n_s, \tau_{\text{reio}}, H_0 \implies$  **enhances the SI mode**
- One could look at massive neutrinos with off-diagonal couplings

Thank you!

# Parameter values

**Table 4:** Parameter values and 68% confidence limits in 2c + 1f.

Parameters	TT+lowE		TTTEEE+lowE	
	SI	MI	SI	MI
$\Omega_b h^2$	$0.022 \pm 0.00027$	$0.022 \pm 0.00021$	$0.022 \pm 0.00016$	$0.022 \pm 0.00015$
$\Omega_c h^2$	$0.1211 \pm 0.0023$	$0.1203 \pm 0.002$	$0.1205 \pm 0.0014$	$0.1201 \pm 0.0013$
$100\theta_s$	$1.0452 \pm 0.00059$	$1.0419 \pm 0.0005$	$1.045 \pm 0.00076$	$1.0419 \pm 0.00031$
$\ln(10^{10} A_s)$	$2.99 \pm 0.0179$	$3.036 \pm 0.01714$	$3 \pm 0.0167$	$3.042 \pm 0.0161$
$n_s$	$0.9407 \pm 0.0079$	$0.9596 \pm 0.0068$	$0.9473 \pm 0.0046$	$0.9628 \pm 0.005$
$\tau_{\text{reio}}$	$0.0501 \pm 0.008$	$0.0516 \pm 0.0079$	$0.0538 \pm 0.0077$	$0.0538 \pm 0.0077$
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	$-1.69 \pm 0.2$	$-4.03 \pm 0.6$	$-1.93 \pm 0.24$	$-4.24 \pm 0.5$
$H_0(\text{ km s}^{-1}\text{Mpc}^{-1})$	$68.34 \pm 1.00$	$67.57 \pm 0.92$	$68.81 \pm 0.63$	$67.83 \pm 0.6$
$\sigma_8$	$0.823 \pm 0.01$	$0.824 \pm 0.009$	$0.829 \pm 0.0079$	$0.824 \pm 0.0075$

**Table 5:** Parameter values and 68% confidence limits in 1c + 2f.

Parameters	TT+lowE		TTTEEE+lowE	
	SI	MI	SI	MI
$\Omega_b h^2$	$0.022 \pm 0.00023$	$0.022 \pm 0.00021$	$0.022 \pm 0.00015$	$0.022 \pm 0.00015$
$\Omega_c h^2$	$0.1207 \pm 0.0021$	$0.1203 \pm 0.002$	$0.1203 \pm 0.0014$	$0.1201 \pm 0.0013$
$100\theta_s$	$1.0434 \pm 0.00062$	$1.0419 \pm 0.0004$	$1.043 \pm 0.00058$	$1.0419 \pm 0.0003$
$\ln(10^{10} A_s)$	$3.01 \pm 0.0179$	$3.037 \pm 0.01664$	$3.024 \pm 0.0166$	$3.042 \pm 0.016$
$n_s$	$0.9513 \pm 0.0069$	$0.9609 \pm 0.0059$	$0.9553 \pm 0.0049$	$0.963 \pm 0.005$
$\tau_{\text{reio}}$	$0.051 \pm 0.008$	$0.0519 \pm 0.008$	$0.0539 \pm 0.0076$	$0.0539 \pm 0.0077$
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	$-1.75 \pm 0.4$	$-3.94 \pm 0.6$	$-1.9 \pm 0.37$	$-4.06 \pm 0.6$
$H_0(\text{ km s}^{-1}\text{Mpc}^{-1})$	$67.9 \pm 1.00$	$67.56 \pm 0.93$	$68.3 \pm 0.62$	$67.83 \pm 0.61$
$\sigma_8$	$0.821 \pm 0.01$	$0.823 \pm 0.009$	$0.825 \pm 0.0083$	$0.824 \pm 0.0075$



# Parameter values

**Table 11:** Parameter values and 68% confidence limits in **2c + 1f**.

Parameters	TTTEEE+lowE+lens		TTTEEE+lowE+lens+BAO+ $H_0$	
	SI	MI	SI	MI
$\Omega_b h^2$	$0.022 \pm 0.0001$	$0.022 \pm 0.00014$	$0.022 \pm 0.0001$	$0.022 \pm 0.00013$
$\Omega_c h^2$	$0.1202 \pm 0.0013$	$0.1199 \pm 0.0012$	$0.12 \pm 0.001$	$0.1188 \pm 0.0009$
$100\theta_s$	$1.045 \pm 0.0008$	$1.0419 \pm 0.0003$	$1.045 \pm 0.00068$	$1.042 \pm 0.00029$
$\ln(10^{10} A_s)$	$3 \pm 0.0158$	$3.041 \pm 0.014$	$3 \pm 0.0151$	$3.044 \pm 0.0145$
$n_s$	$0.9476 \pm 0.0043$	$0.9629 \pm 0.0048$	$0.9483 \pm 0.004$	$0.966 \pm 0.0046$
$\tau_{\text{reio}}$	$0.0541 \pm 0.0074$	$0.0536 \pm 0.0072$	$0.0544 \pm 0.007$	$0.0565 \pm 0.0071$
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	$-1.96 \pm 0.26$	$-4.22 \pm 0.51$	$-1.91 \pm 0.22$	$-4.22 \pm 0.52$
$H_0(\text{km s}^{-1}\text{Mpc}^{-1})$	$68.87 \pm 0.58$	$67.9 \pm 0.53$	$69.08 \pm 0.42$	$68.47 \pm 0.4$
$\sigma_8$	$0.829 \pm 0.007$	$0.823 \pm 0.006$	$0.827 \pm 0.0065$	$0.821 \pm 0.0059$

**Table 12:** Parameter values and 68% confidence limits in **1c + 2f**.

Parameters	TTTEEE+lowE+lens		TTTEEE+lowE+lens+BAO+ $H_0$	
	SI	MI	SI	MI
$\Omega_b h^2$	$0.022 \pm 0.0001$	$0.022 \pm 0.00014$	$0.022 \pm 0.0001$	$0.022 \pm 0.00013$
$\Omega_c h^2$	$0.1201 \pm 0.0012$	$0.1199 \pm 0.0012$	$0.12 \pm 0.0009$	$0.1188 \pm 0.0009$
$100\theta_s$	$1.043 \pm 0.0006$	$1.0419 \pm 0.0003$	$1.043 \pm 0.00056$	$1.042 \pm 0.00029$
$\ln(10^{10} A_s)$	$3.023 \pm 0.0153$	$3.041 \pm 0.015$	$3 \pm 0.0151$	$3.045 \pm 0.0142$
$n_s$	$0.9555 \pm 0.0046$	$0.9633 \pm 0.0045$	$0.9572 \pm 0.004$	$0.966 \pm 0.0042$
$\tau_{\text{reio}}$	$0.0536 \pm 0.0073$	$0.0536 \pm 0.0074$	$0.0554 \pm 0.007$	$0.0566 \pm 0.0071$
$\log_{10}(G_{\text{eff}}/\text{MeV}^{-2})$	$-1.91 \pm 0.37$	$-4.04 \pm 0.61$	$-1.86 \pm 0.36$	$-4.03 \pm 0.61$
$H_0(\text{km s}^{-1}\text{Mpc}^{-1})$	$68.35 \pm 0.56$	$67.9 \pm 0.54$	$68.75 \pm 0.41$	$68.48 \pm 0.41$
$\sigma_8$	$0.824 \pm 0.007$	$0.823 \pm 0.006$	$0.822 \pm 0.0071$	$0.821 \pm 0.006$