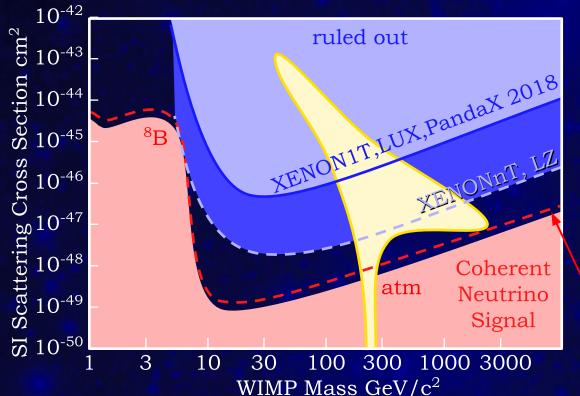
SNOWMASS 2021 Topical Group CF1 Liquid Xenon for a Generation-3 Detector

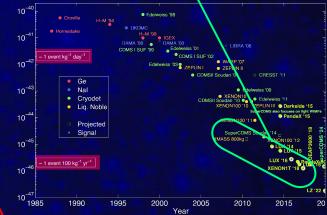
August 7, 2020

Kevin T. Lesko, LZ Berkeley Lab ktlesko@lbl.gov Rafael F. Lang, XENON Purdue University rafael@purdue.edu

Neutrino Floor is far, far away

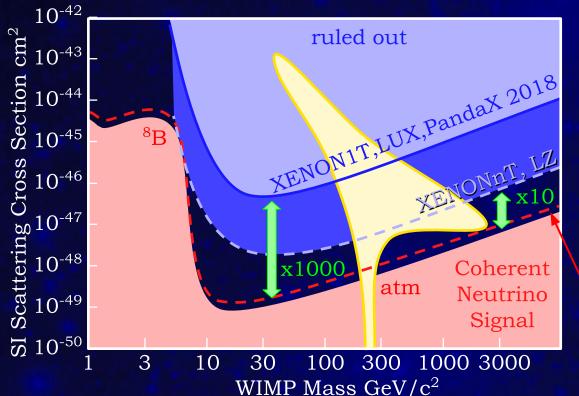


Leading SI limits all from LXe TPCs, for the past decade

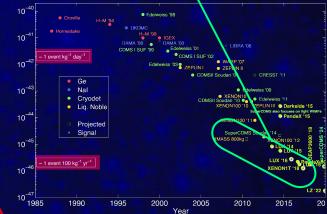


1 Neutrino Event

Neutrino Floor is far, far away

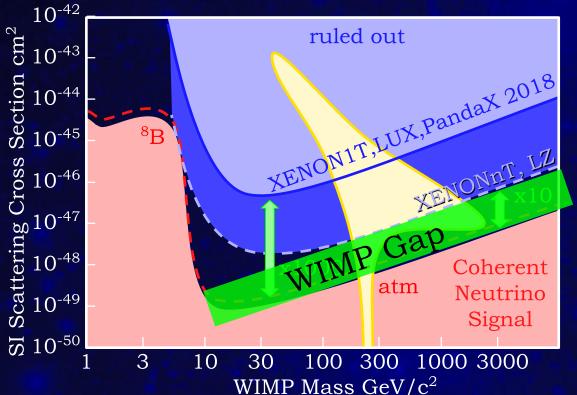


Leading SI limits all from LXe TPCs, for the past decade

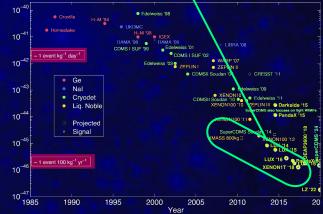


1 Neutrino Event

^{atm}Neutrino Floor is far, far away

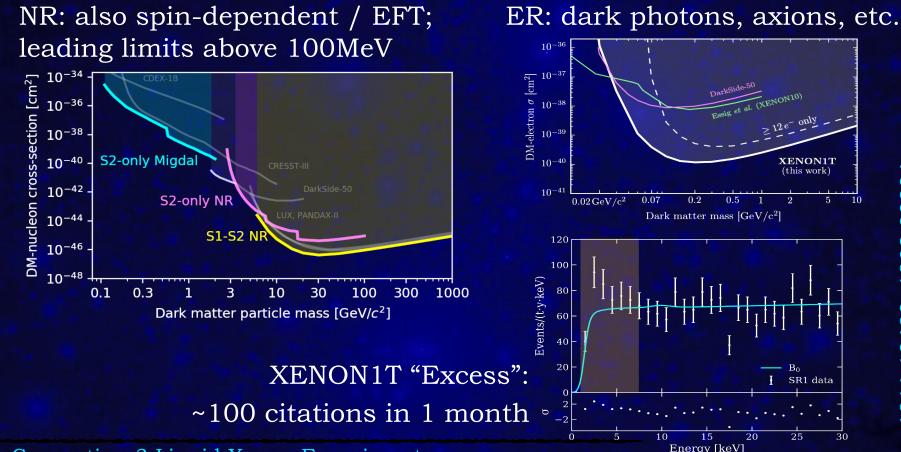


Leading SI limits all from LXe TPCs, for the past decade



Current program leaves well-motivated WIMP Gap

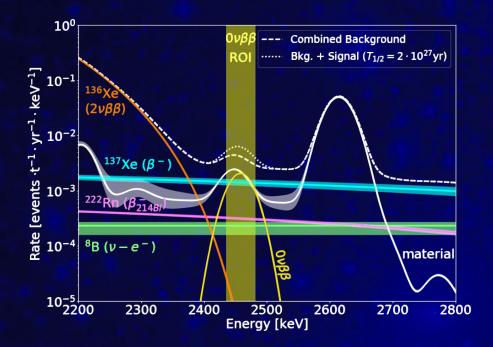
Xenon: Sensitivity to Many Dark Matter Candidates



Generation-3 Liquid Xenon Experiment

006.0972

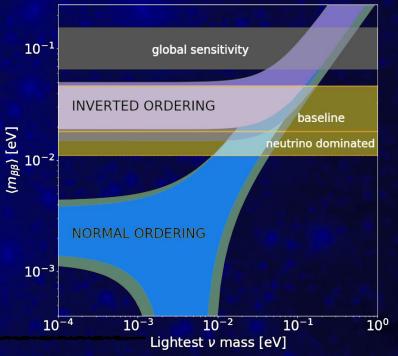
¹³⁶Xe 0v2β with ^{nat}Xe Target Abundance 8.9%: No (expensive) enrichment



plus additional ββ channels Generation-3 Liquid Xenon Experiment

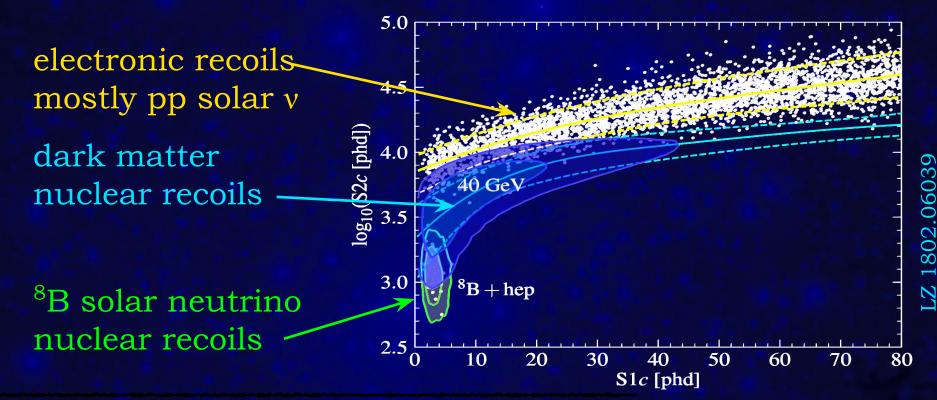
potentially probe entire inverted hierarchy

 $^{136}\text{Xe} \rightarrow ^{136}\text{Ba} + 2e^{-1}$



6

Solar ⁸B CEvNS ~2025: Guaranteed Science Here: simulation of 1000 days LZ



Generation-3 Liquid Xenon Experiment

 $\nu_e + N \rightarrow \nu_e + N$

<100MeV Atmospheric Neutrino CEvNS

10

8

6

4

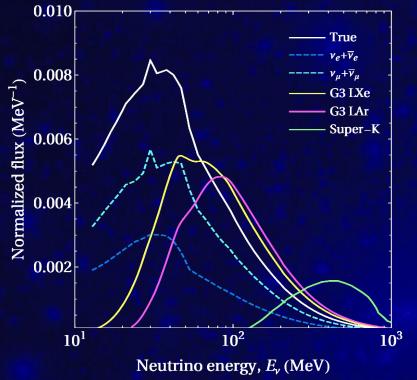
2

0

0

Significance, $Z(\sigma)$

measure low energies:

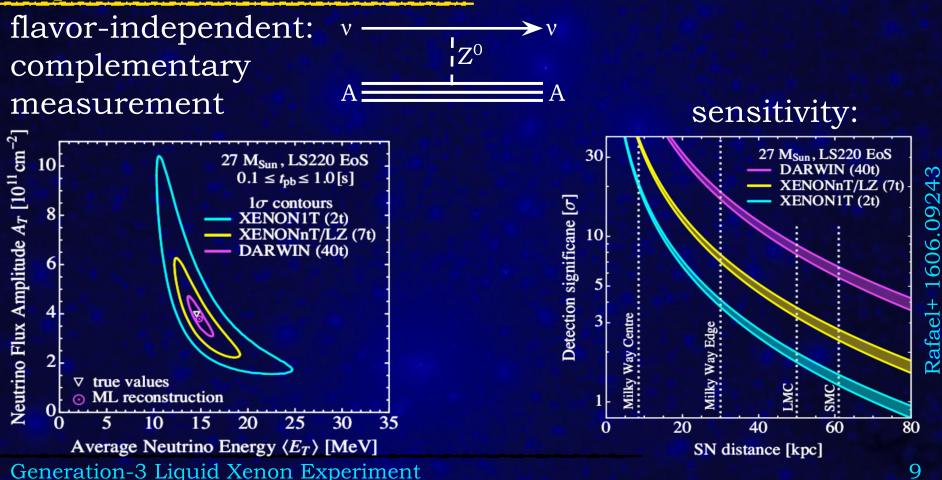




need >500 t years: Atm. ν discovery in Xenon ($g_1=0.3$) 68% CI (statistical) ----- $\mu=1, Med[Z]$ ±20% (systematic) 500 1000 1500 2000

Exposure (tonne.yr)

.08566 2002. Strigari 8 Rafael Newstead,

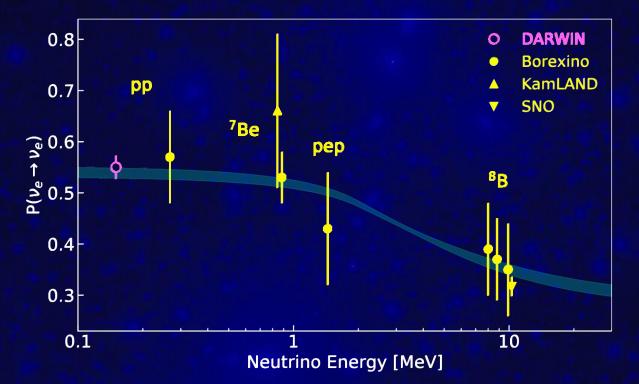


Supernova Neutrino CEvNS

9

 $\overline{\nu_x + N} \rightarrow \overline{\nu_x} + N$

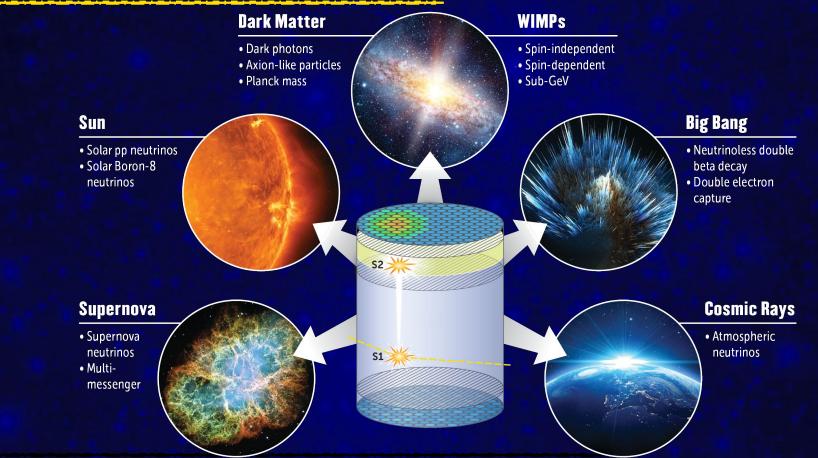
Elastic pp Solar Neutrino Scattering



Also measure $\sin^2 \theta_W$ to a few percent Generation-3 Liquid Xenon Experiment

 $\nu_e + e^- \rightarrow \nu_e + e^-$

Diverse Program with Xenon



The Next Steps for the US Dark Matter Program

Exploit demonstrated & scalable Xe detector designs

- Scale-up existing TPCs by ~ ×10
- Manageable run times: 5 -10 years
- Controlled and mitigated backgrounds
- Capitalize on investments and leadership

 Exploit existing experiments & upgrades
 Expand R&D efforts

 $\stackrel{\text{Particle}}{\longrightarrow}$

SUSTREE F

Engage the Strong International Community

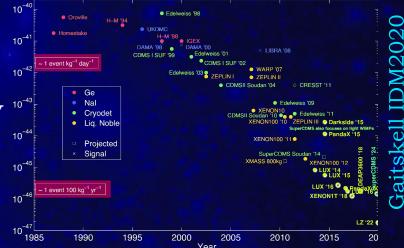
 Large community of (> 500) interested scientists with LXe expertise: LZ, XENON, XMASS, PandaX, ...

• Initial R&D funding in place in Europe

 Excellent opportunities and platforms in the US for essential G3 R&D

Low Risk & Economic Path to Dark Matter Discovery

 ZEPLIN, XENON10, LUX, XENON100, XENON1T/nT and LZ have delivered promised sensitivity within cost envelopes for past decade



- Multiple, <u>demonstrated</u> purification technologies no isotopic enrichment required to reduce radioactive backgrounds
- Backgrounds understood, measured & controlled
- Xenon is an investment, not an expense Generation-3 Liquid Xenon Experiment

Xe: Essential Part of the Global Program in Dark Matter

We recognize the value of pursuing multiple technologies and targets with physics topics of this importance: Dark Matter, $0\nu\beta\beta$, & Neutrinos

Look forward to quantifying the complementarity of different technologies at Snowmass as detector mass, backgrounds, and operational parameters are better understood.

The Xe G3 program builds directly on the Snowmass 2014 Report & P5 Recommendations



Snowmass and 2014 P5 Report

"The results of G2 direct detection experiments and other contemporaneous dark matter searches will guide the technology and design of third-generation experiments. As the scale of these experiments grows to increase sensitivity, the experimental challenge of direct detection will still require complementary experimental techniques, and international cooperation will be warranted. The U.S. should host at least one of the third-generation experiments in this complementary global suite."

Recommendation 20: Support one <u>or more</u> third-generation (G3) direct detection experiments, guided by the results of the preceding searches. Seek a globally complementary program and increased international partnership in G3 experiments."

Conclusions for the Xe G3 Program

Excellent Scientific Reach & Discovery Potential

- Variety of Well-motivated Dark Matter Candidates
- Neutrinoless Double-beta Decay
- Astrophysical Neutrinos

Proven Record Delivering Sensitivity Goals & Performance

• 3 multi-tonne detectors being commissioned

Strong Scientific Expertise and Leadership in the US

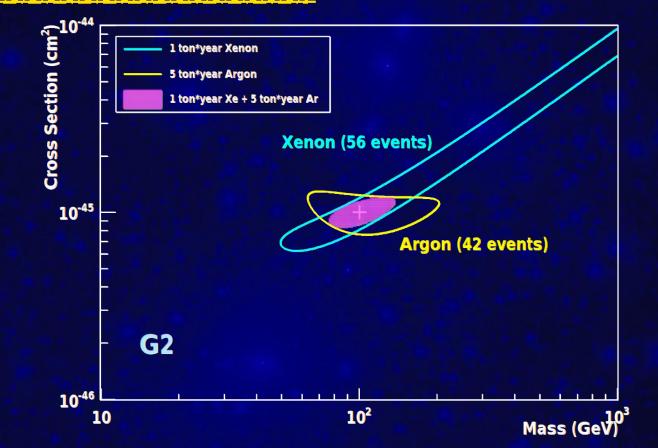
- Exploit Operating Expts and Upgrades to Inform G3 Well-aligned with Existing US and International Plans
 - Complementary Technologies and Targets

Extras

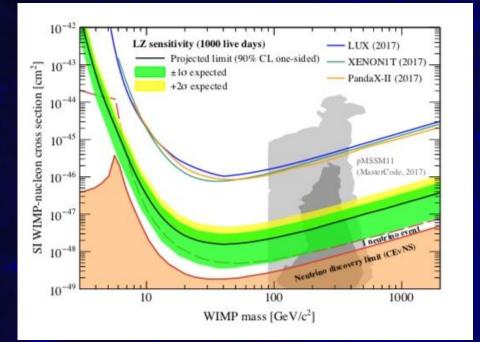
Challenge for Snowmass 2021

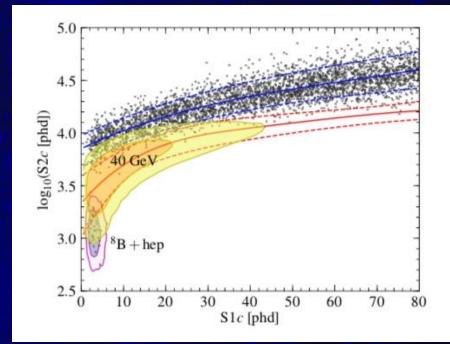


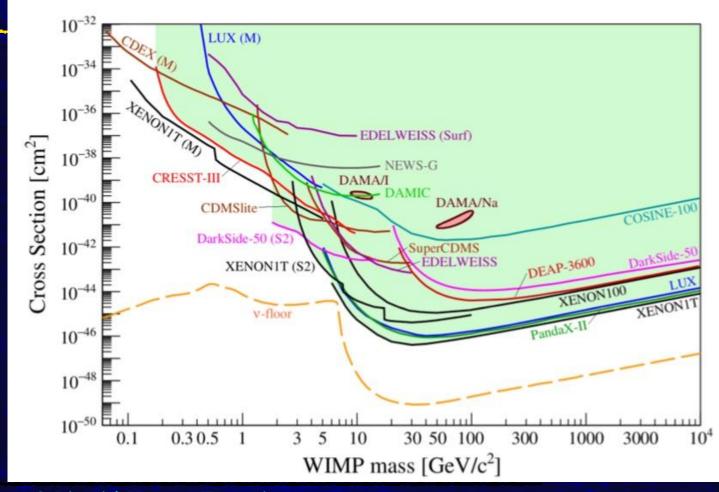
Multi-Target Approach to Identify Signals



1802.06039, Phys. Rev D.







Generation-3 Liquid Xenon Experiment