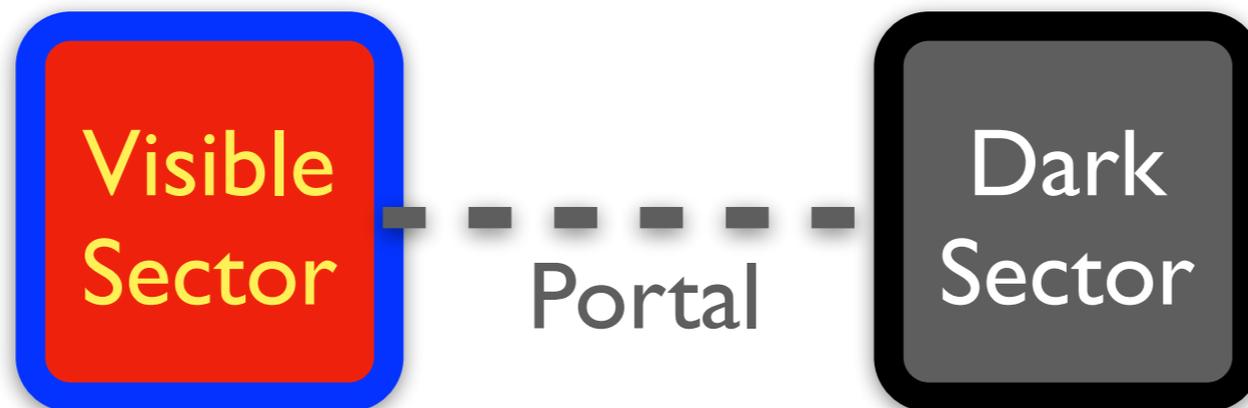


RF6 Dark Sectors at High Intensities

Theory Introduction



Brian Batell (Pitt) and Philip Schuster (SLAC)

Snowmass RF6 Kickoff Meeting
August 12-13, 2020

What is a Dark Sector?

- Set of new particles which do not experience the known forces.
- Weakly coupled to visible sector through a mediator or “portal”.

Why Dark Sectors?

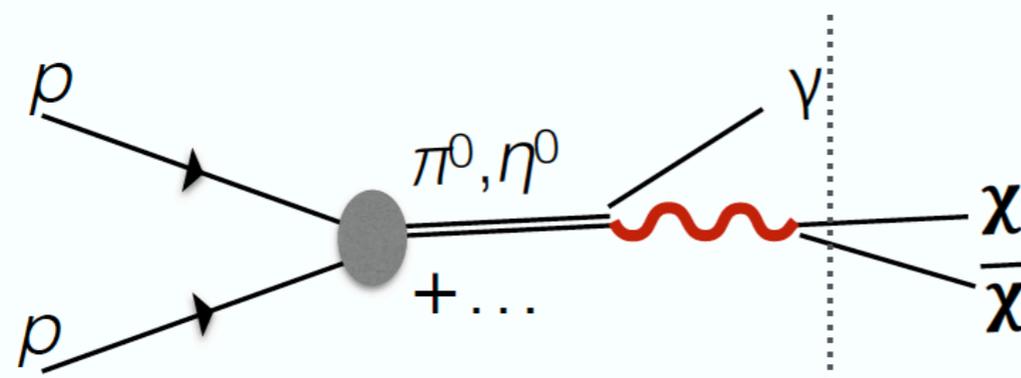
- Dark matter may reside in the dark sector. Simple (thermal) and novel cosmologies are possible.
- Dark sectors may play a role in addressing other puzzles, e.g., neutrino masses, matter-antimatter asymmetry, naturalness, ...
- Experimental anomalies often interpreted in context of dark sectors.

Why high intensities?

- High intensities allow probes of weak portal couplings.
- Past/existing high intensity experiments already provide leading constraints.
- Great potential to discover dark sectors and discern their structure with new searches and experiments in the coming years.

Benchmark studies

- One natural way to organize benchmarks is according to final state signatures
- This scheme aligns well with the science case



“Once you get in to the dark sector, what do you look for?”

non-SM

Dark Matter Production

Producing stable particles that could be (all or part of) the Dark Matter

SM

Portal-Mediator Decays to SM

Systematically explore the minimal couplings of SM to dark sectors

mixed

Structure of the Dark Sector

There could be a rich sector of new physics under our noses

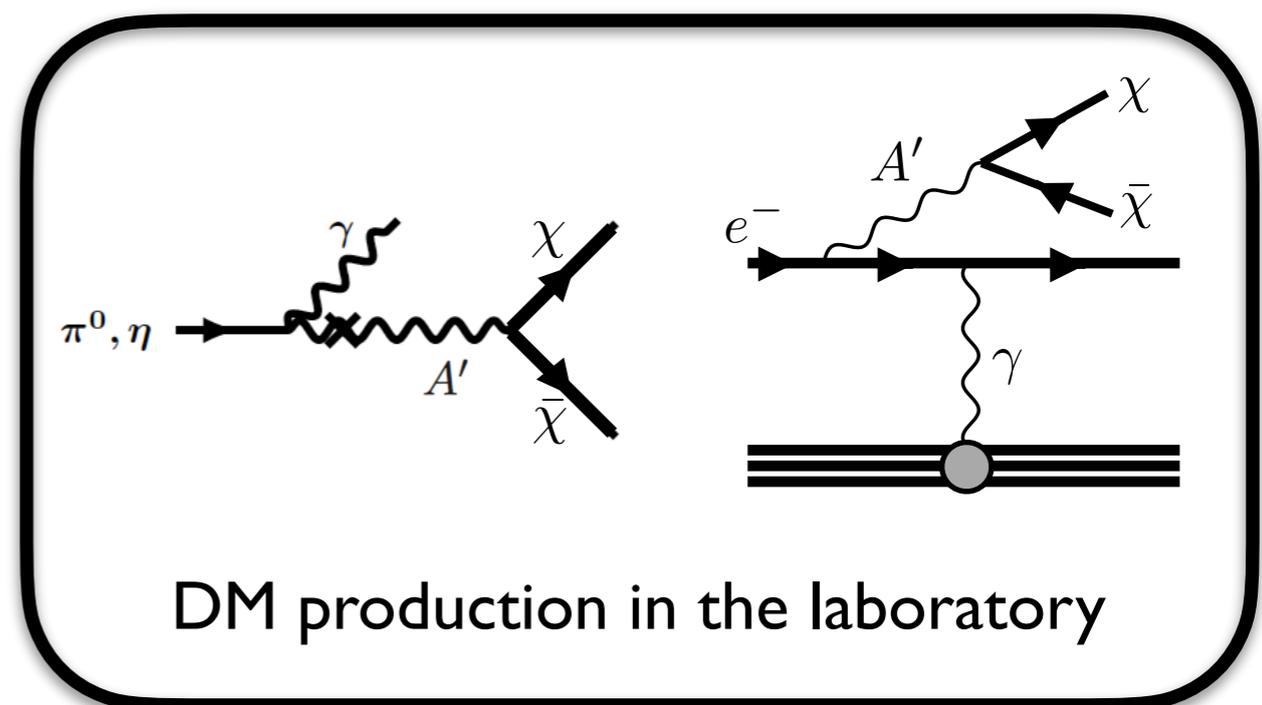
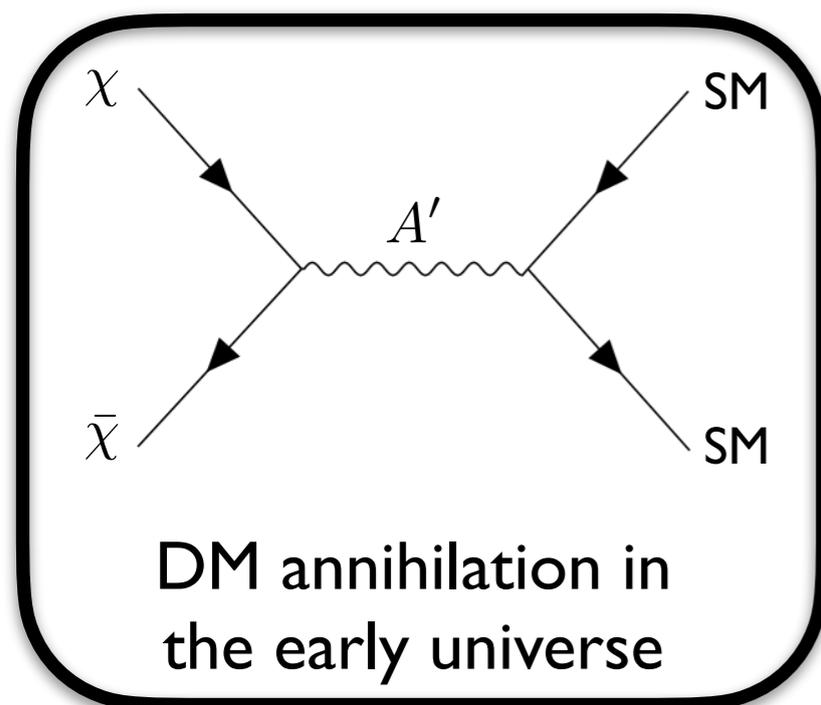
Dark matter production at high intensities

see talk by G. Krnjaic

- Cosmological DM abundance generated via thermal freeze-out
- Extension of the WIMP below Lee-Weinberg bound



- Direct DM annihilation to SM leads to predictive targets for experiment
- Same interaction governs DM annihilation and laboratory DM production



Portal/mediator production/decay at high intensities

see talk by A. Ritz

- The mediator (or “portal”) plays critical role in dark sector physics

• Renormalizable portals:	$\frac{\epsilon}{2 \cos \theta_W} F'_{\mu\nu} B^{\mu\nu}$	Vector Portal
	$(A S + \lambda S^2) H^\dagger H$	Higgs Portal
	$y N L H$	Neutrino portal

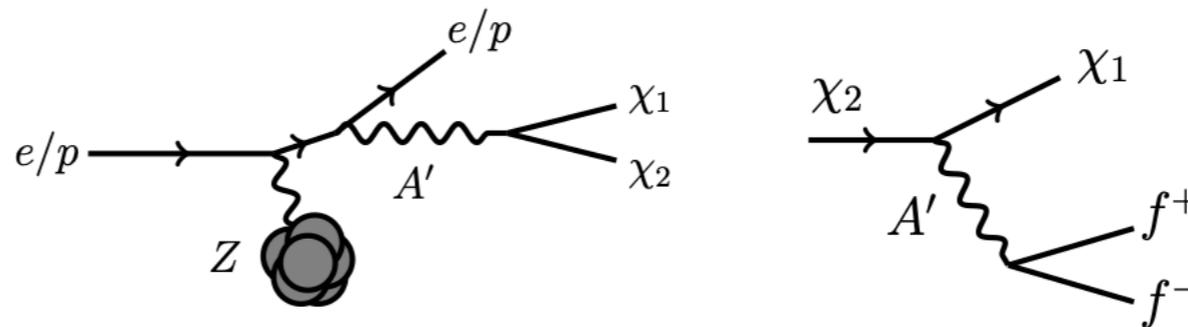
- Other possibilities include higher dimension operators (e.g., axion portal), gauging conserved currents ($B - L, L_\mu - L_\tau, \dots$), etc.
- The mediator can decay back to visible SM final states through the same portal coupling (can be prompt or displaced/long-lived)
- Dark matter can still be thermally produced through “secluded” annihilation to mediators
- Typically very simple parameter space (mass — coupling plane)

Rich dark sectors at high intensities

see talk by A. Berlin

- The dark sector may be as rich as the Standard Model. It could feature many feature myriad novel phenomena such as spontaneous symmetry breaking, confinement, supersymmetry, ...
- UV considerations may call for additional structure (e.g., a dark Higgs generates dark photon mass, which may further cause DM mass splittings)
- Novel DM cosmology may suggest rich structure as in e.g., SIMP
- Rich dark sector structure often leads to novel experimental signatures

e.g., inelastic DM
from 1703.06881



- Much work remains to characterize and fully explore rich dark sectors

Benchmarks in Final State x Portal Organization

	DM Production	Mediator Decay Via Portal	Structure of Dark Sector
Vector	m_{χ} vs. y [$m_A/m_{\chi}=3, \alpha_D=.5$] $m_{A'}$ vs. y [$\alpha_D=0.5, 3 m_{\chi}$ values] <u>m_{χ} vs. α_D</u> [$m_A/m_{\chi}=3, y=y_{fo}$] m_{χ} vs. m_A [$\alpha_D=0.5, y=y_{fo}$] <i>Millicharge m vs. q</i>	<u>$m_{A'}$ vs. ϵ</u> [decay-mode agnostic] $m_{A'}$ vs. ϵ [decays]	iDM m_{χ} vs. y [$m_{A'}/m_{\chi}=3, \alpha_D=.5$] (anom connection) SIMP-motivated cascades [slices TBD] $U(1)_{B-L} / \mu-\tau / B-3\tau$ (DM or SM decays)
Scalar	m_{χ} vs. $\sin\theta$ [$\lambda=0, \text{fix } m_S/m_{\chi}, g_D$] (thermal target excluded 1512.04119, should still include) Note secluded DM relevance of $S \rightarrow \text{SM}$ of mediator searches	m_S vs. $\sin\theta$ [$\lambda=0$] m_S vs. $\sin\theta$ [$\lambda=\text{s.t. Br}(H \rightarrow \phi\phi \sim 10^{-2})$]?	Dark Higgsstrahlung (w/vector) scalar SIMP models? Leptophilic/leptophobic dark Higgs?
Neutrino	$e/\mu/\tau$ a la 1709.07001?	m_N vs. U_e m_N vs. U_{μ} m_N vs. U_{τ} Think more about reasonable flavor structures	Sterile neutrinos with new forces?
ALP	m_{χ} vs. f_q/l [$\lambda=0, \text{fix } m_a/m_{\chi}, g_D$] (thermal target excluded) What about f_{ν}, f_G ?	m_a vs. f_{ν} m_a vs. f_G m_a vs. $f_q=f_l$ (separate?) Think more about reasonable coupling relations including f_{WZ}	FV axion couplings

+ Neutron portal? Hidden valleys (or are these out-of-scope?)? See e.g. 2003.02270

Bold = BRN benchmark, italic=PBC benchmark. others are new suggestions. Underline=CV benchmarks that were not used in BRN

Resources

- Dark Sectors 2016 Workshop: Community Report

<https://arxiv.org/abs/1608.08632>

- U.S. Cosmic Visions: New Ideas in Dark Matter 2017: Community Report

<https://arxiv.org/abs/1707.04591>

- Physics Beyond Colliders at CERN: Beyond the Standard Model Working Group Report

<https://arxiv.org/abs/1901.09966>

- Basic Research Needs for Dark Matter Small Projects New Initiatives

https://science.osti.gov/-/media/hep/pdf/Reports/Dark_Matter_New_Initiatives_rpt.pdf