

Xenon Doping

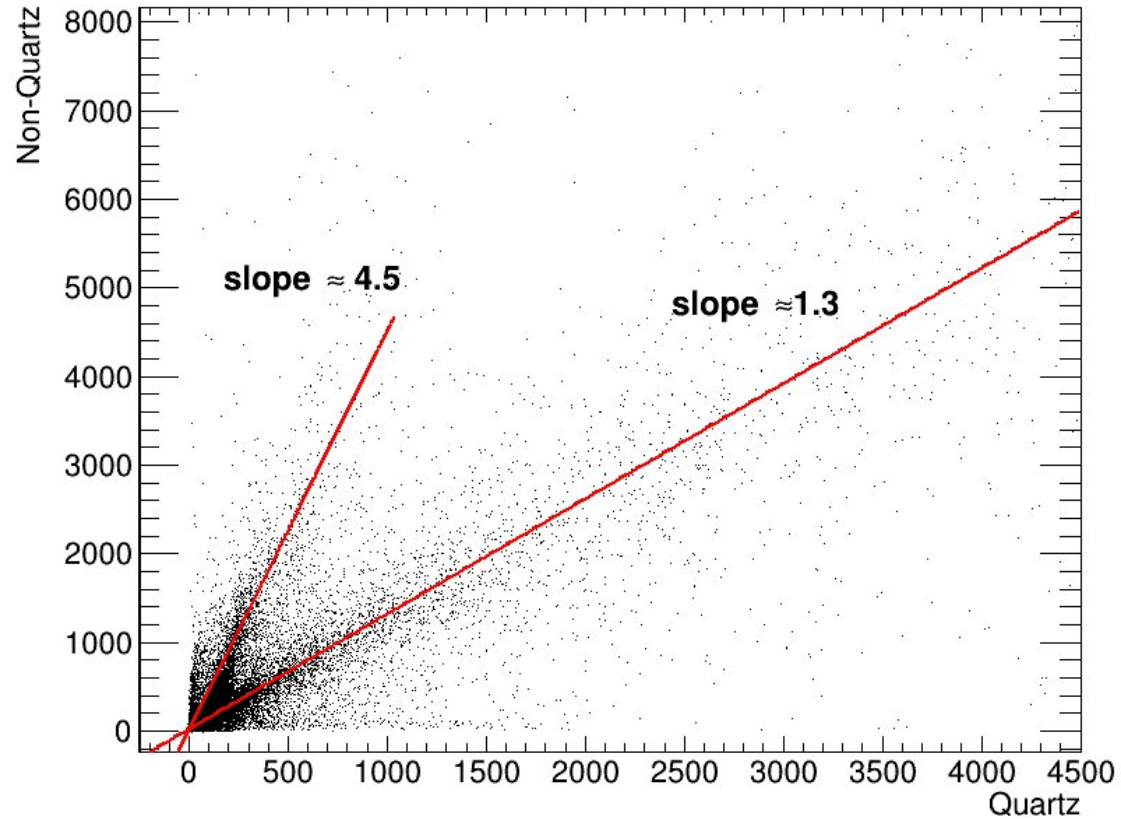
Niccolo' Gallice, Henrique Souza
07/08/2020



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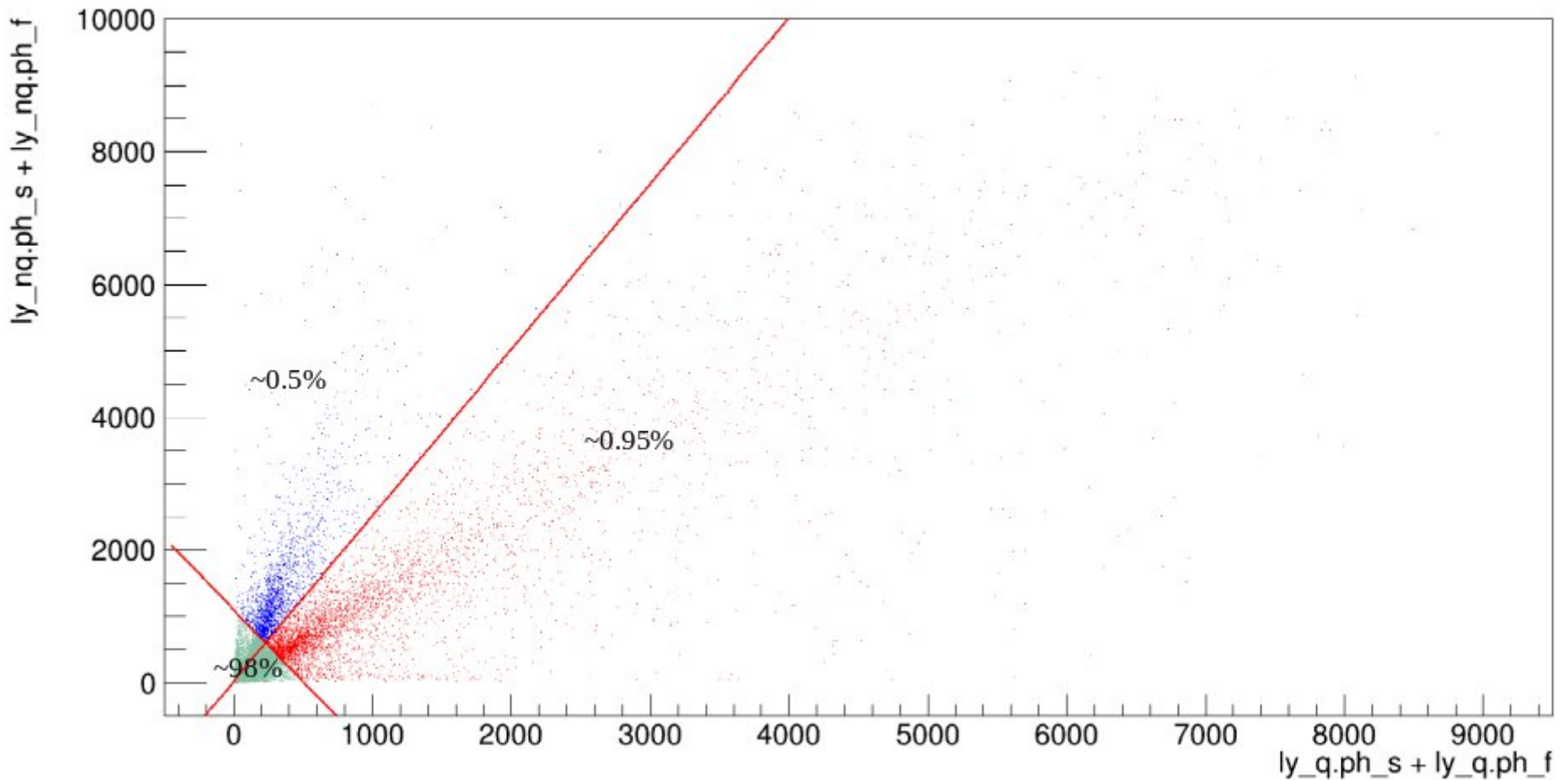


Scatter-plot



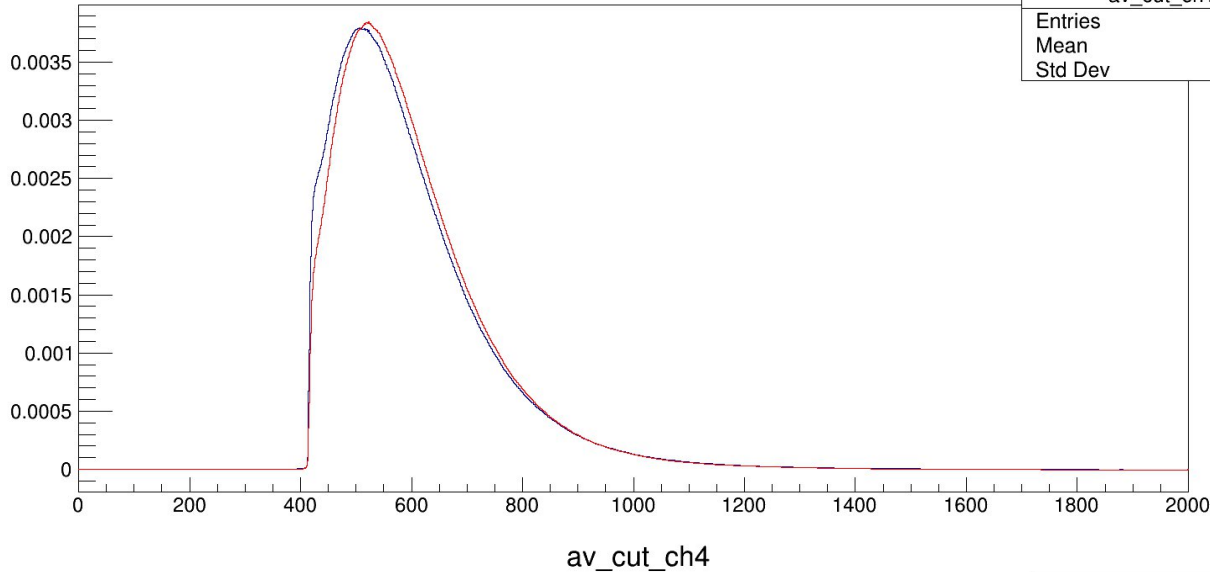
- Scatter-plot of NQ-Arapuca vs Q-Arapuca
- 2 different contributions (lobes)

Scatter-plot



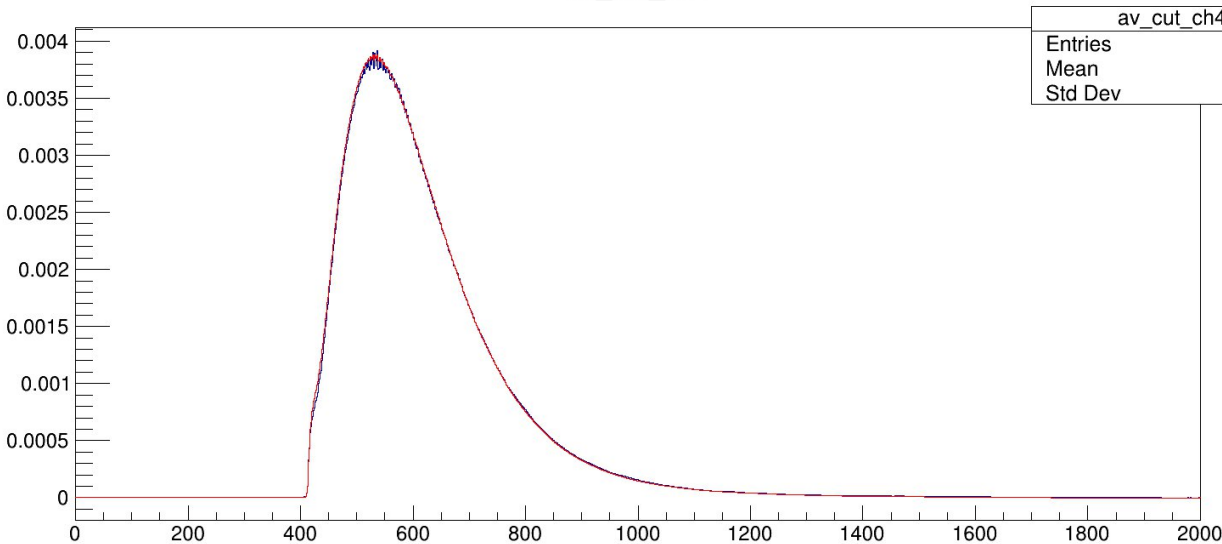
Scatter-plot

av_cut_ch1



Shape seems very similar despite the a subtle difference in the fast component for NQ

av_cut_ch4



Average light detected (Light yield)

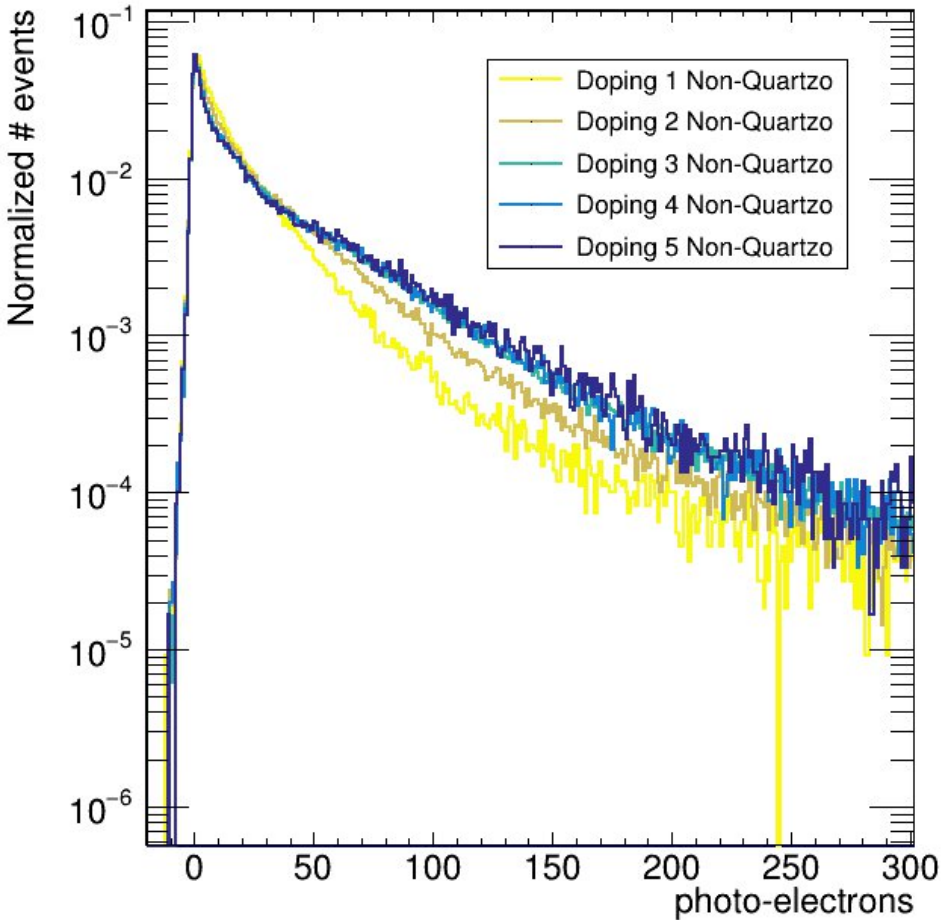
- Average number of photons per trigger was retrieved for each doping.
- A growth in the average is related to the overall efficiency (emission and/or collection)
- Is the ratio affected by geometry ?

Xe (ppm)	Average photons per trigger	
	No Quartz	Quartz
0.8	33.5	7.8
3.3	45.5	19.5
11	56.6	34.5
15	56.9	37.4
20	58.3	38.1

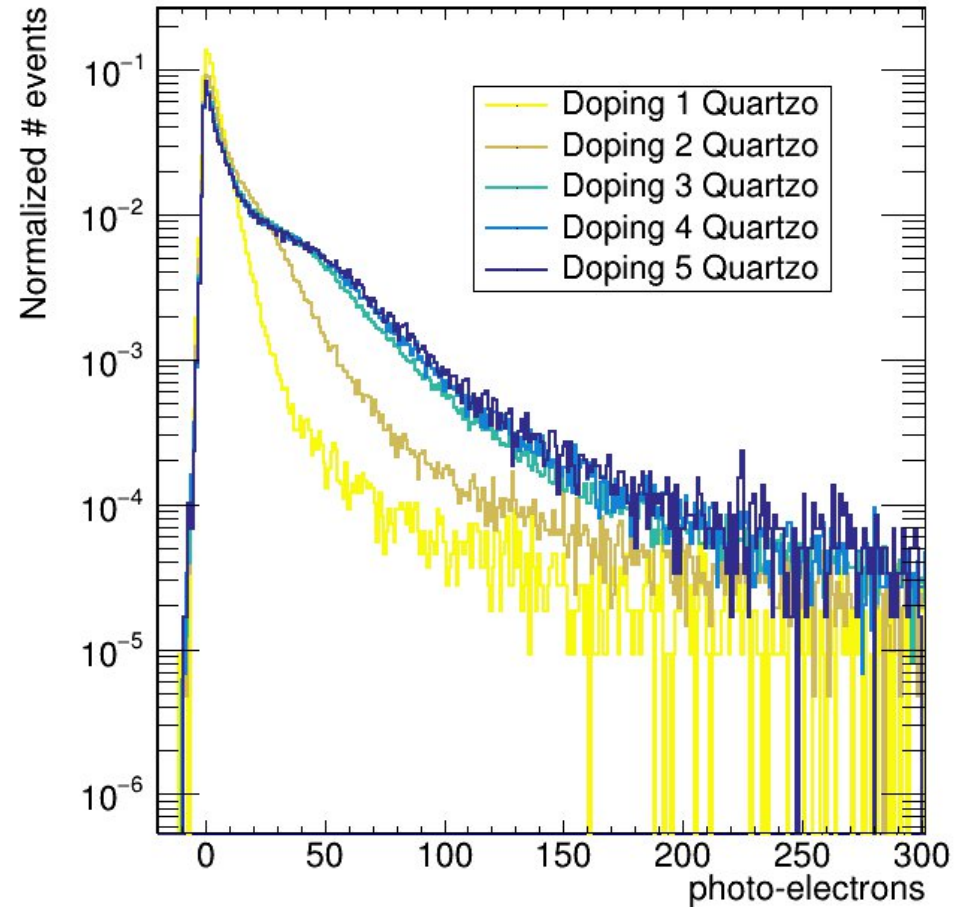
Xe (ppm)	Relative NQ	Q / NQ
0.8	1.000	0.234
3.3	1.360	0.428
11	1.691	0.610
15	1.702	0.657
20	1.741	0.655

Average light detected (Light yield)

NQ - Light distribution



Q - Light distribution



Simulation motivation

A simulation could be helpful to:

- Explain the two lobes observed in the scatter-plot
- Check if the relation between the average number of photons and the light yield is linear
- Verify the muon spectrum response obtained by the two x-arapucas.
- Check the ratio between $Xe/(Ar + Xe)$

Monte Carlo - Toy model

Why a toy model Monte-Carlo?

- Is not a complete **Geant4** simulation. It does not take into account many physical properties, trying only to bring relevant processes and appropriate approximations
- Everything was implemented from scratch, mistakes can be made :)
- Basically only geometrical effects will have an impact on the results

MC - implementation

Muon generation:

Phys. Rev. D58, 05401 (1998)

$$\mathcal{D}_\mu \left(p, h = 1030 \text{ g/cm}^2, \vartheta = 0^\circ \right) = C p^{-(\gamma_0 + \gamma_1 \log p + \gamma_2 \log^2 p + \gamma_3 \log^3 p)}$$

TABLE II. Parameters of the fitting formula (3.4) for the vertical energy spectrum of conventional muons at sea level.

Momentum range (GeV/c)	C ($\text{cm}^{-2}\text{s}^{-1}\text{sr}^{-1}\text{GeV}^{-1}$)	γ_0	γ_1	γ_2	γ_3
$1 \div 9.2765 \times 10^2$	2.950×10^{-3}	0.3061	1.2743	-0.2630	0.0252

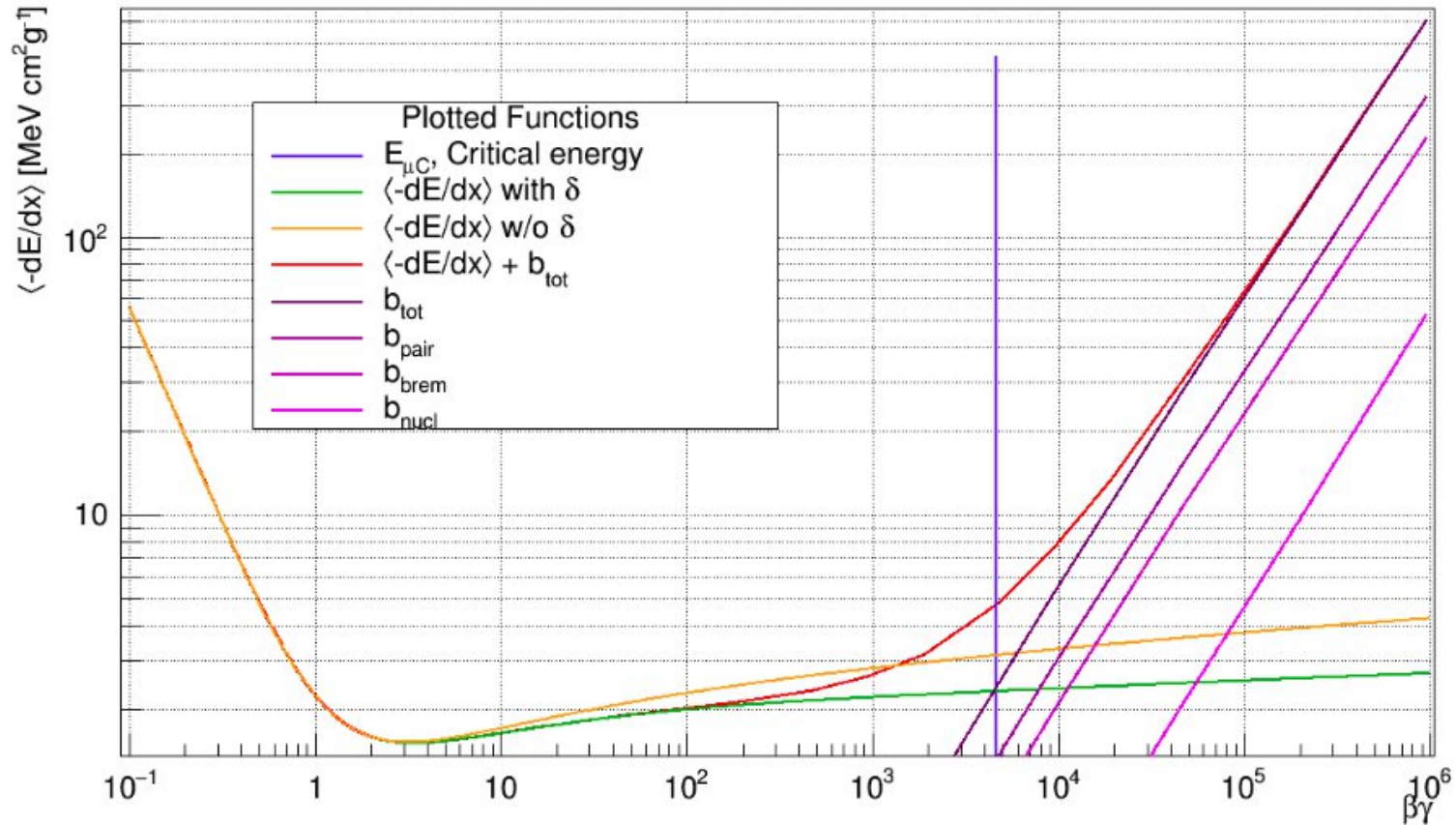
From $p = 1$ to 10 GeV/c

Angular distribution:

- Zenith: $\cos^2\theta$
- Azimuth: uniform

MC - implementation

Muon energy loss:



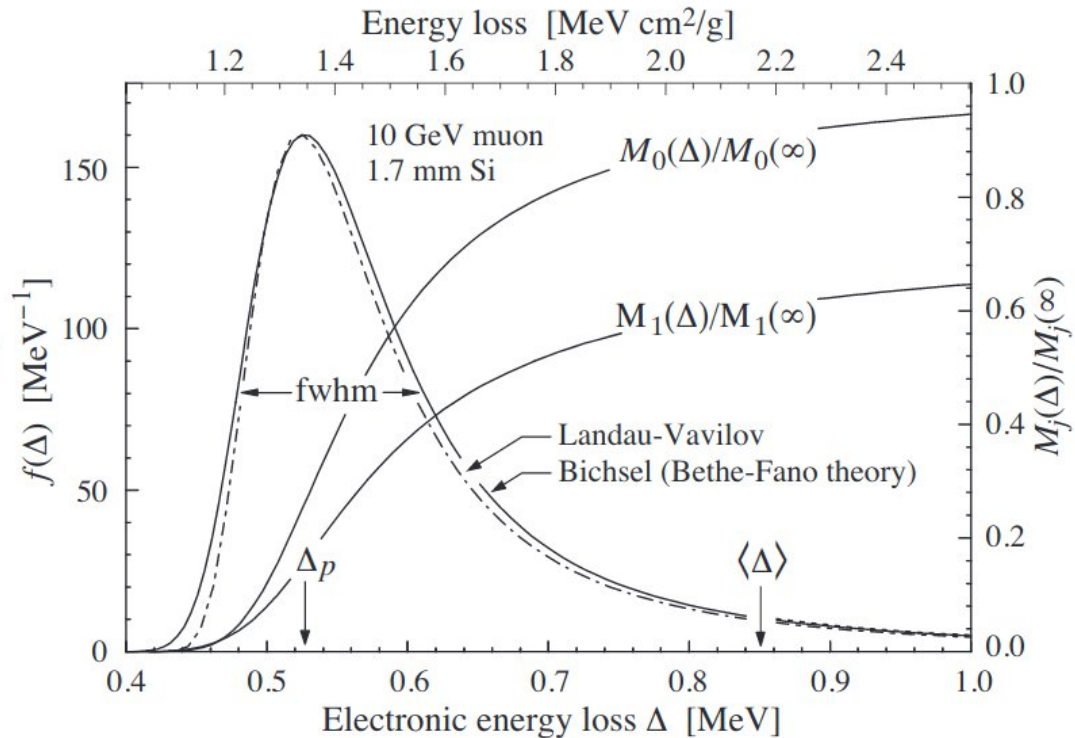
https://indico.fnal.gov/event/14933/contributions/28526/attachments/17961/22583/Final_SIST_Paper.pdf

MC - implementation

Muon energy loss:

fwhm set as 4ξ , where

$$\xi = (K/2) \langle Z/A \rangle z^2 (x/\beta^2) \text{ MeV}$$

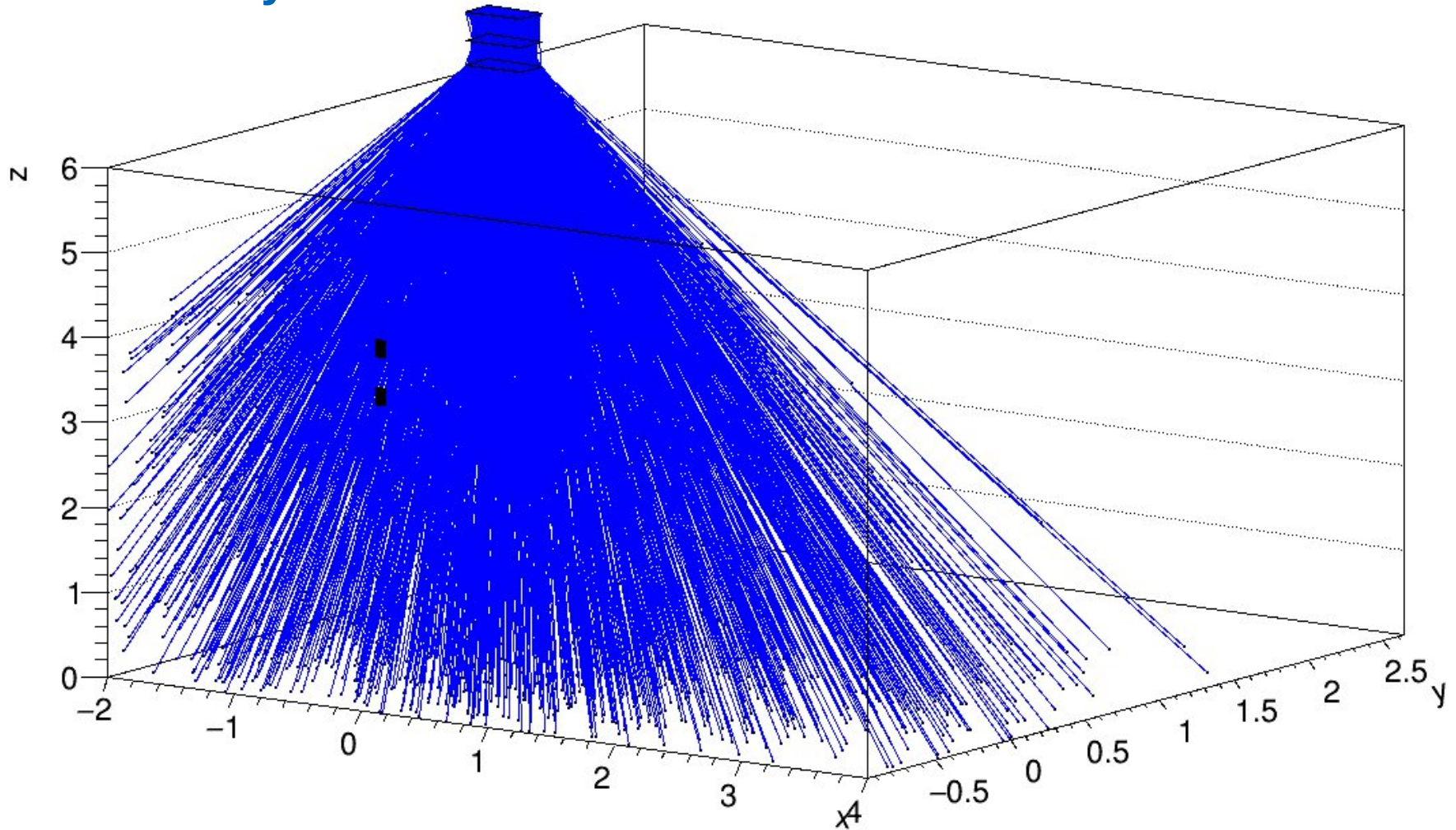


Just an approximation: we still have many muons stopping inside the detector (but this did not affect sensibly the results).

<https://pdg.lbl.gov/2012/reviews/rpp2012-rev-passage-particles-matter.pdf>

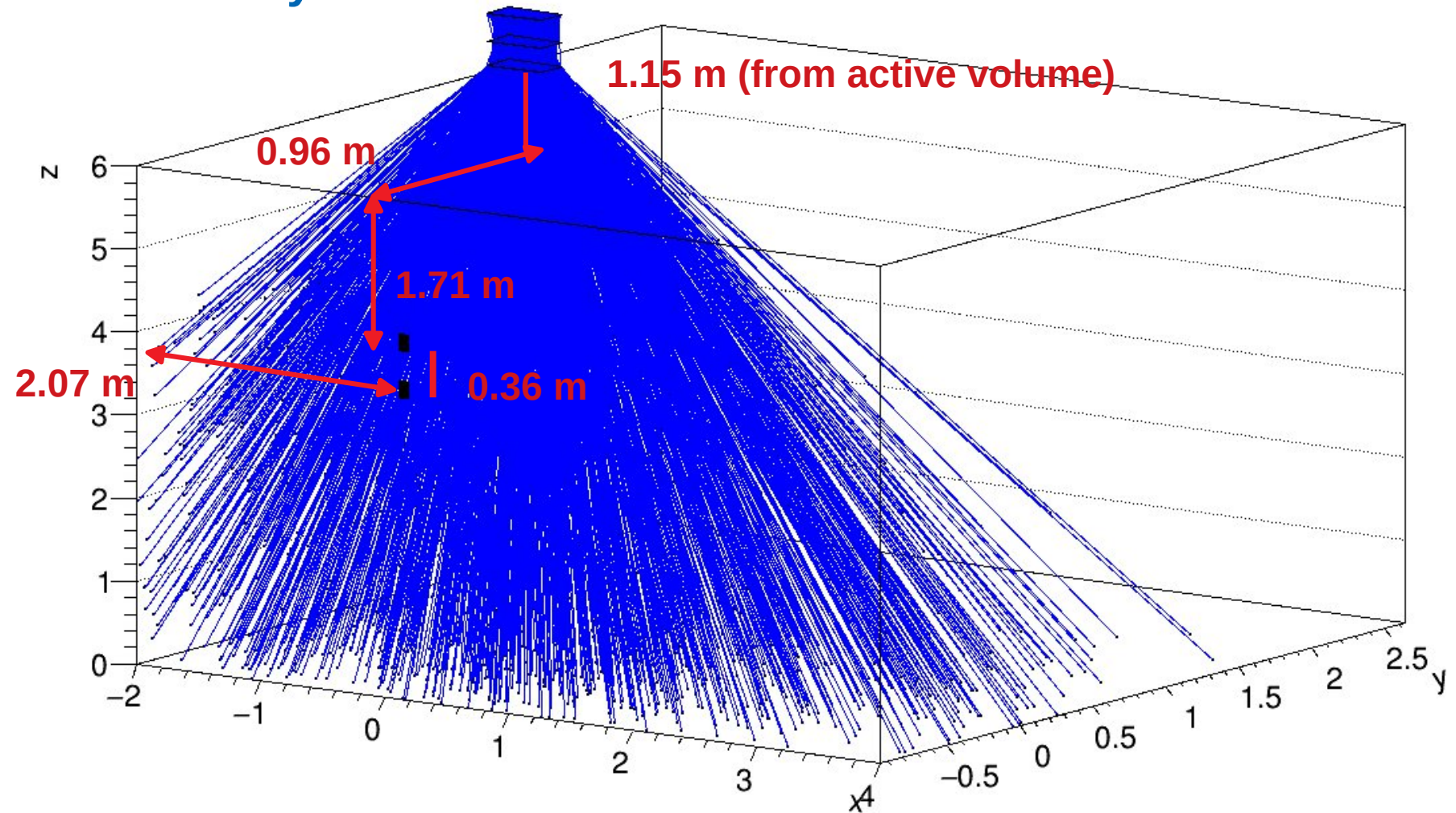
MC - implementation

Geometry:

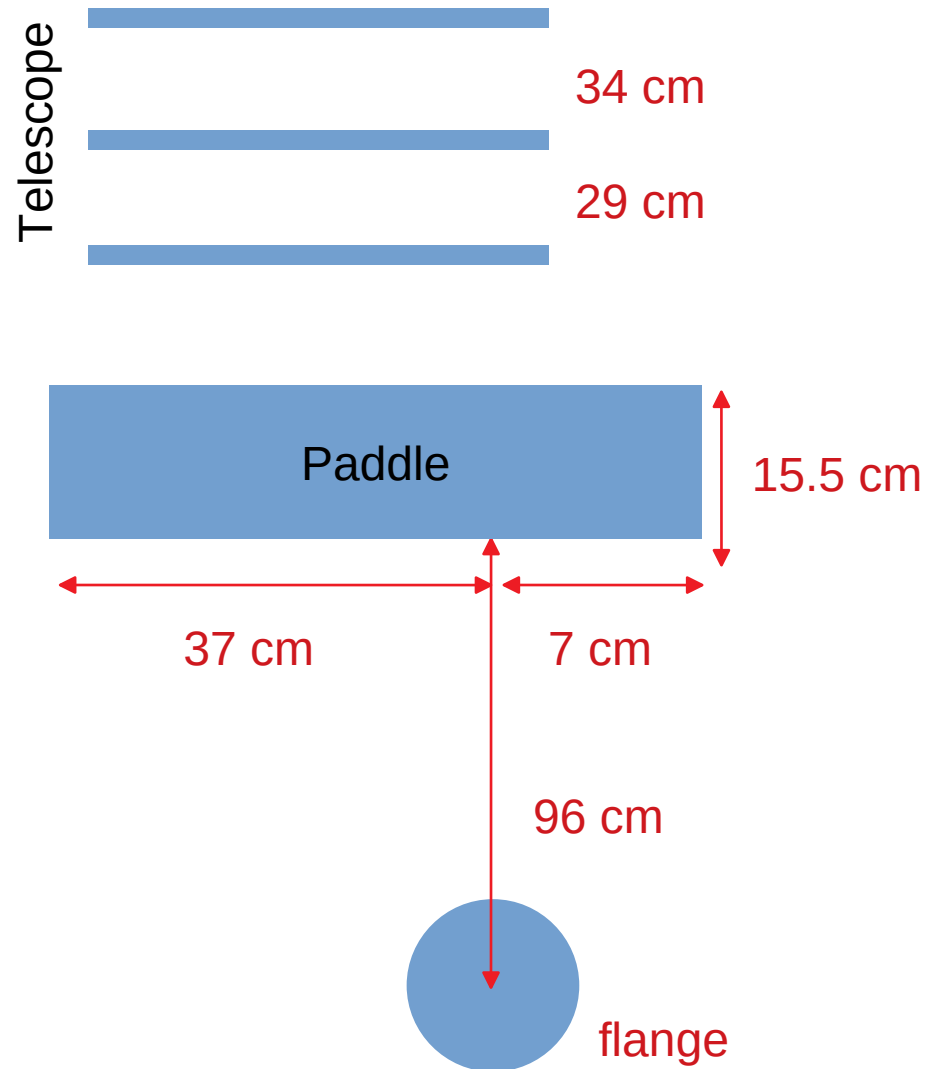
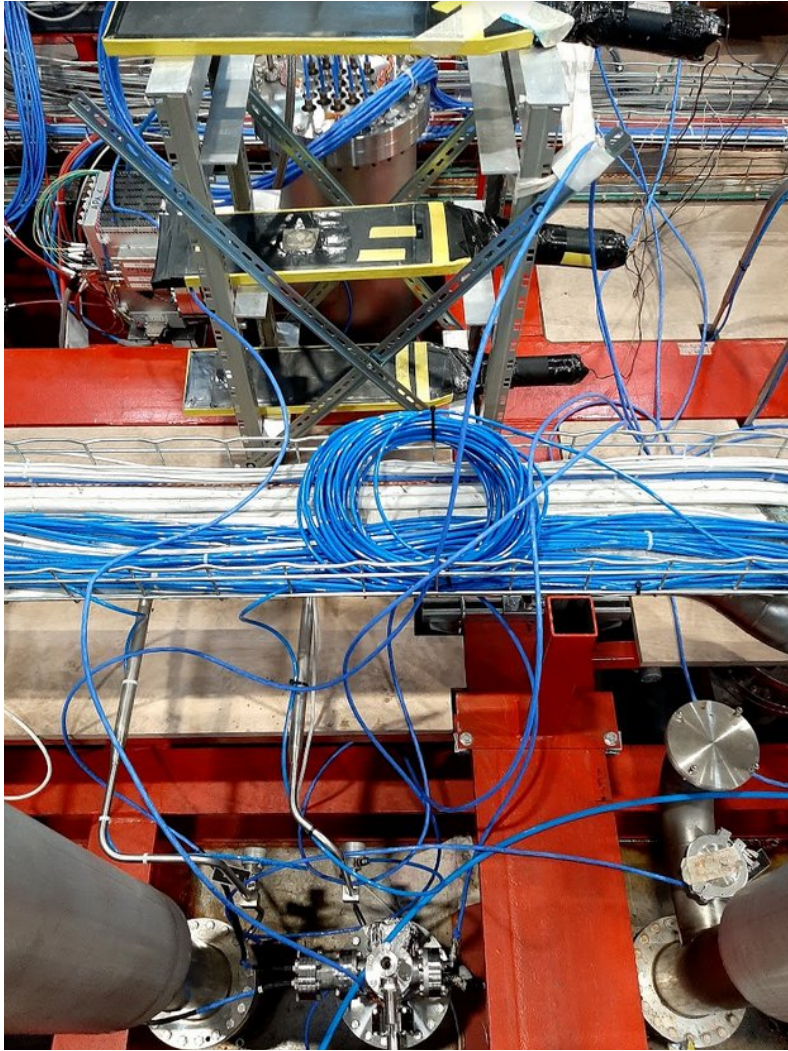


MC - implementation

Geometry:



MC - implementation



MC - implementation

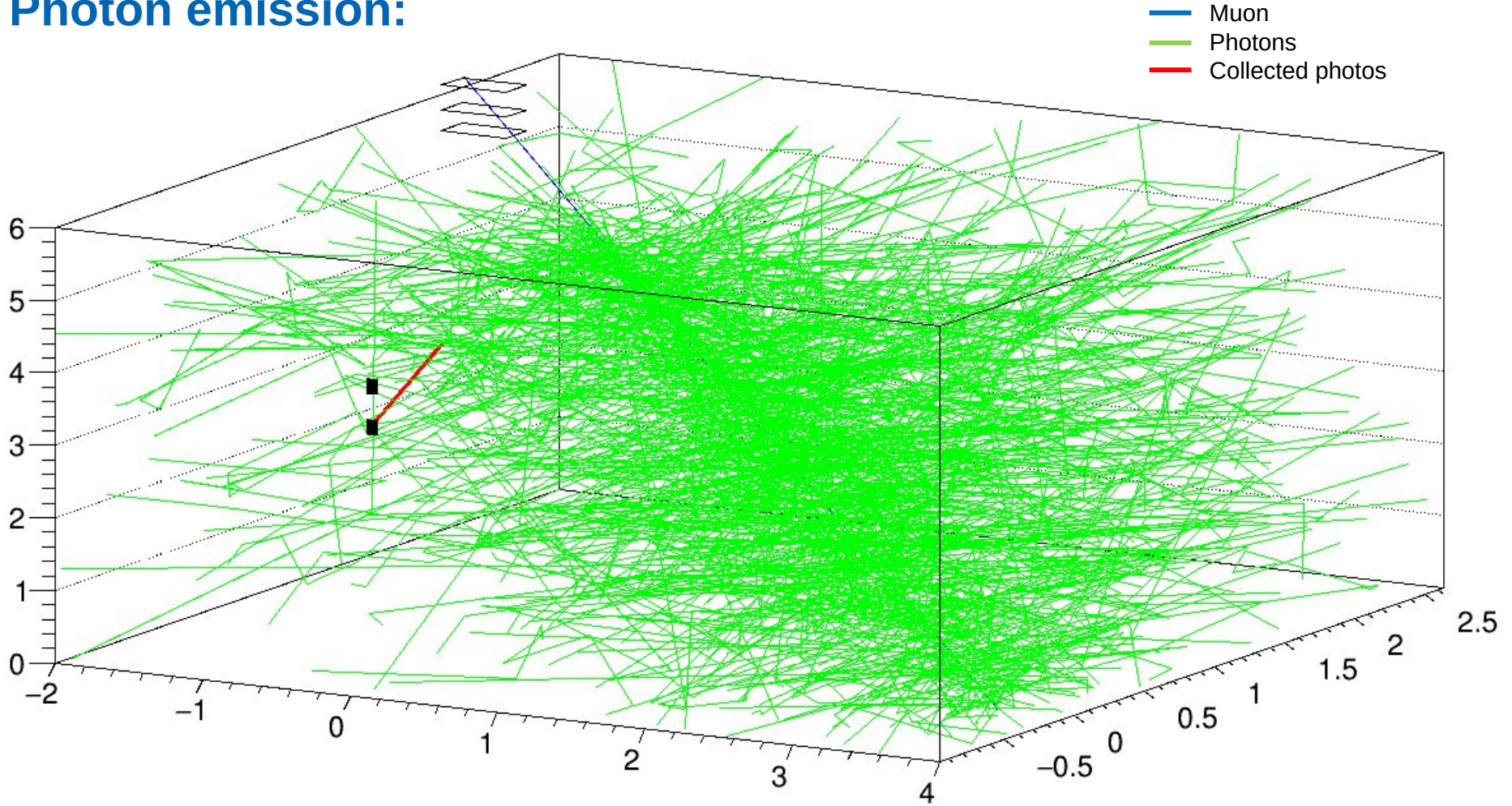
Photon emission:

- Isotropic emission
- Light yield was set as 400 photons/MeV. This (roughly) correspond to a 1% efficiency of the X-ARAPUCA.
- Rayleigh Scattering with $\lambda = 1$ m
- Absorption [1] with $\lambda = 20$ m
- The “type” (128 nm or 175 nm) of photon is decided in the emission, 70% for Xe and 30 % Argon.
- To speed up simulation, we tried to set the photon "type" only if it hit the X-ARAPUCA window. No change in the output was noticed.

[1] Jones, B J P et al. “A Measurement of the Absorption of Liquid Argon Scintillation Light by Dissolved Nitrogen at the Part-Per-Million Level.” Journal of Instrumentation 8.07 (2013): P07011–P07011. <https://arxiv.org/abs/1306.4605v2>.

MC - implementation

Photon emission:



MC - implementation

Shadow due to grids:

In front of the x-arapucas there are:

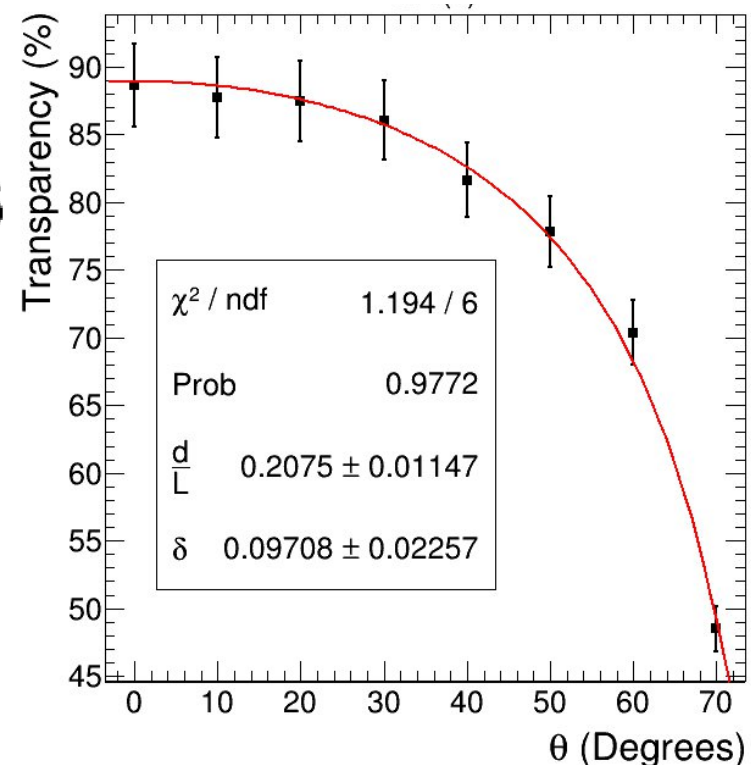
- 6 TPC wires \rightarrow $d = 0.15$ mm (diameter) and $L = 4.75$ mm (pitch)
- Two sets of ground grids (transparency measured at UNICAMP)

Fitting function

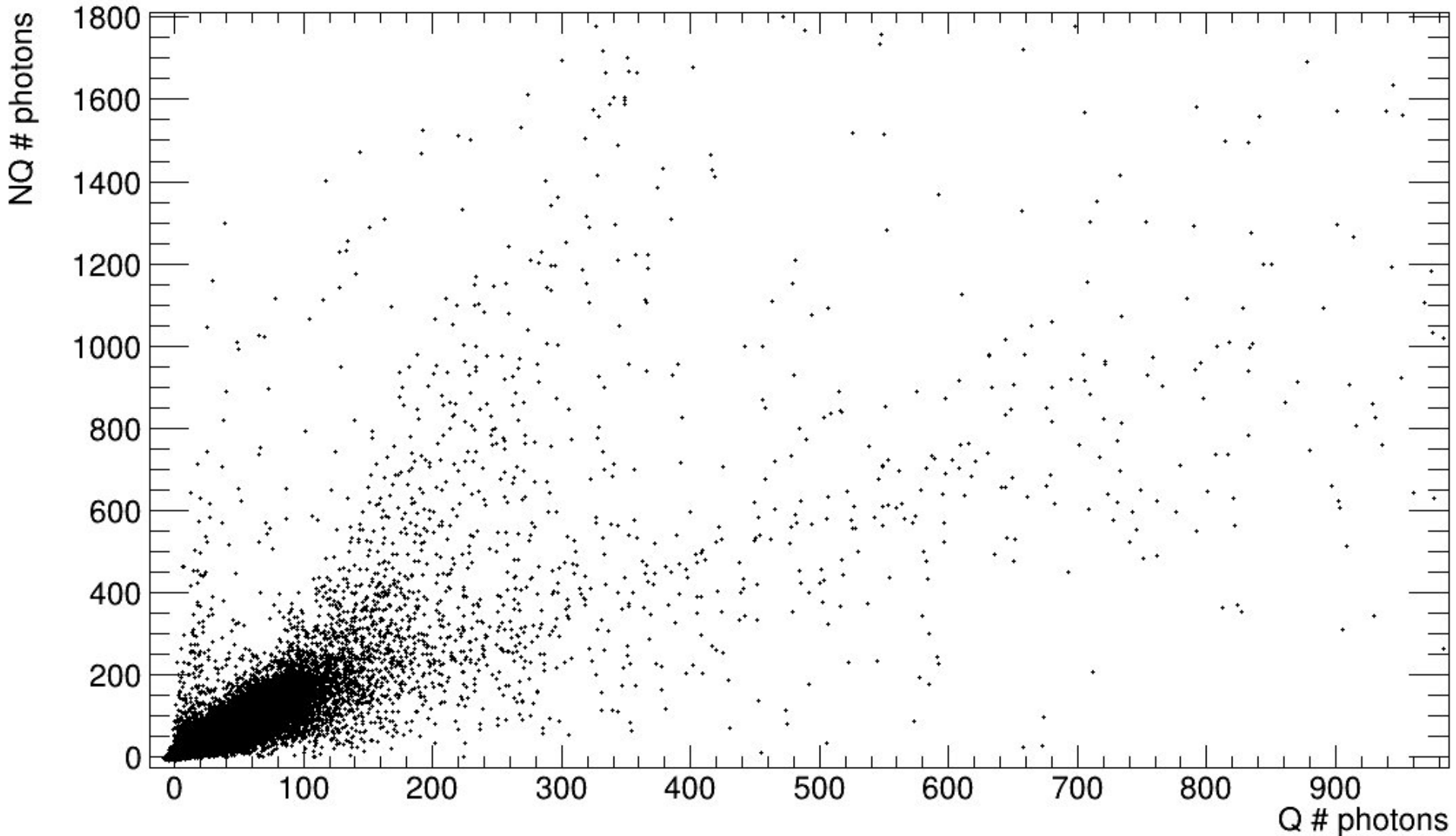
$$1 - \frac{d}{L} \frac{1}{\cos(\theta)} + \delta$$

For each photon, we save the angle relative to the normal of the X-ARAPUCA.

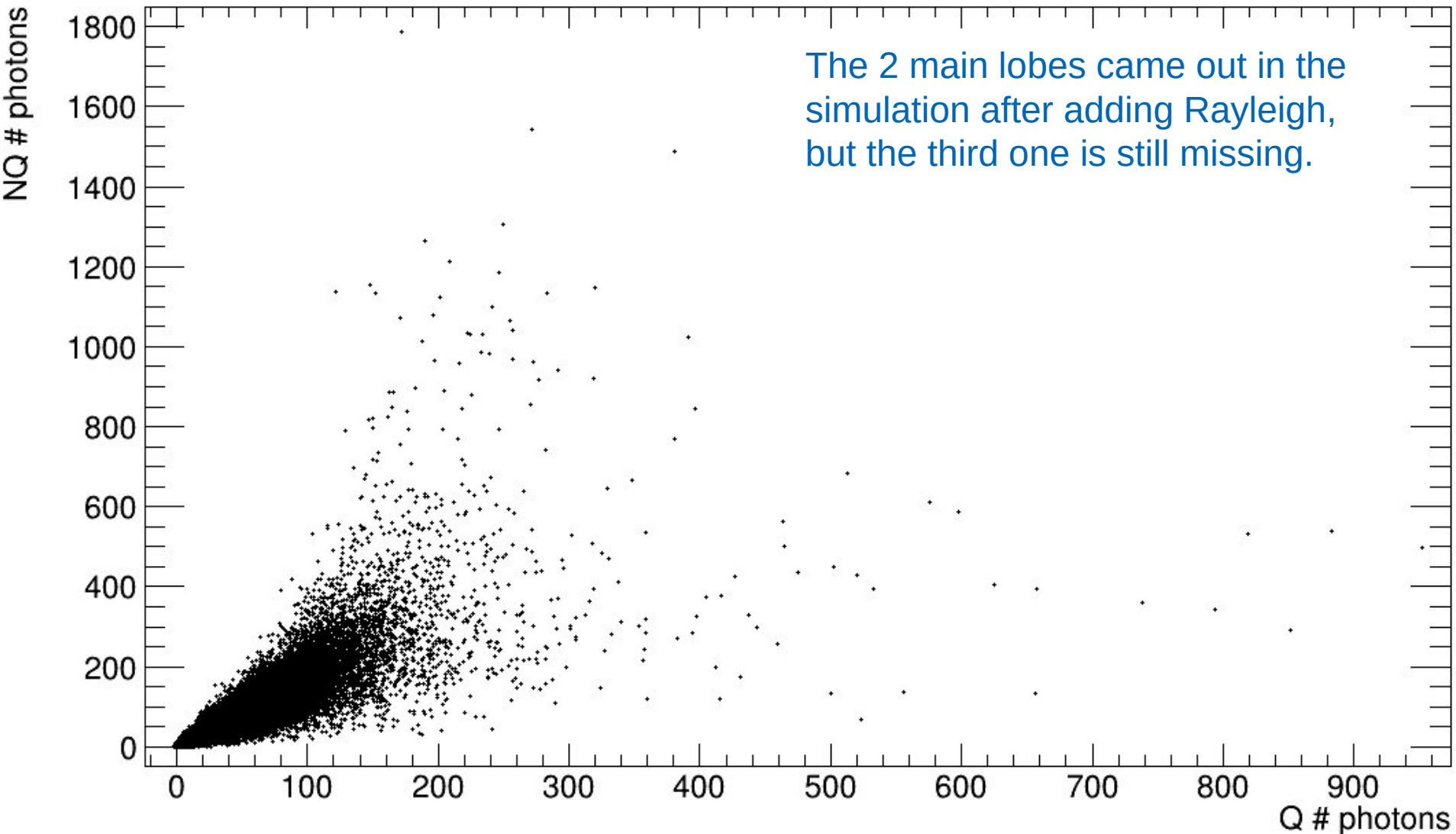
The photon must go through all the 6 + 2 grids with a survival probability given by the formula above.



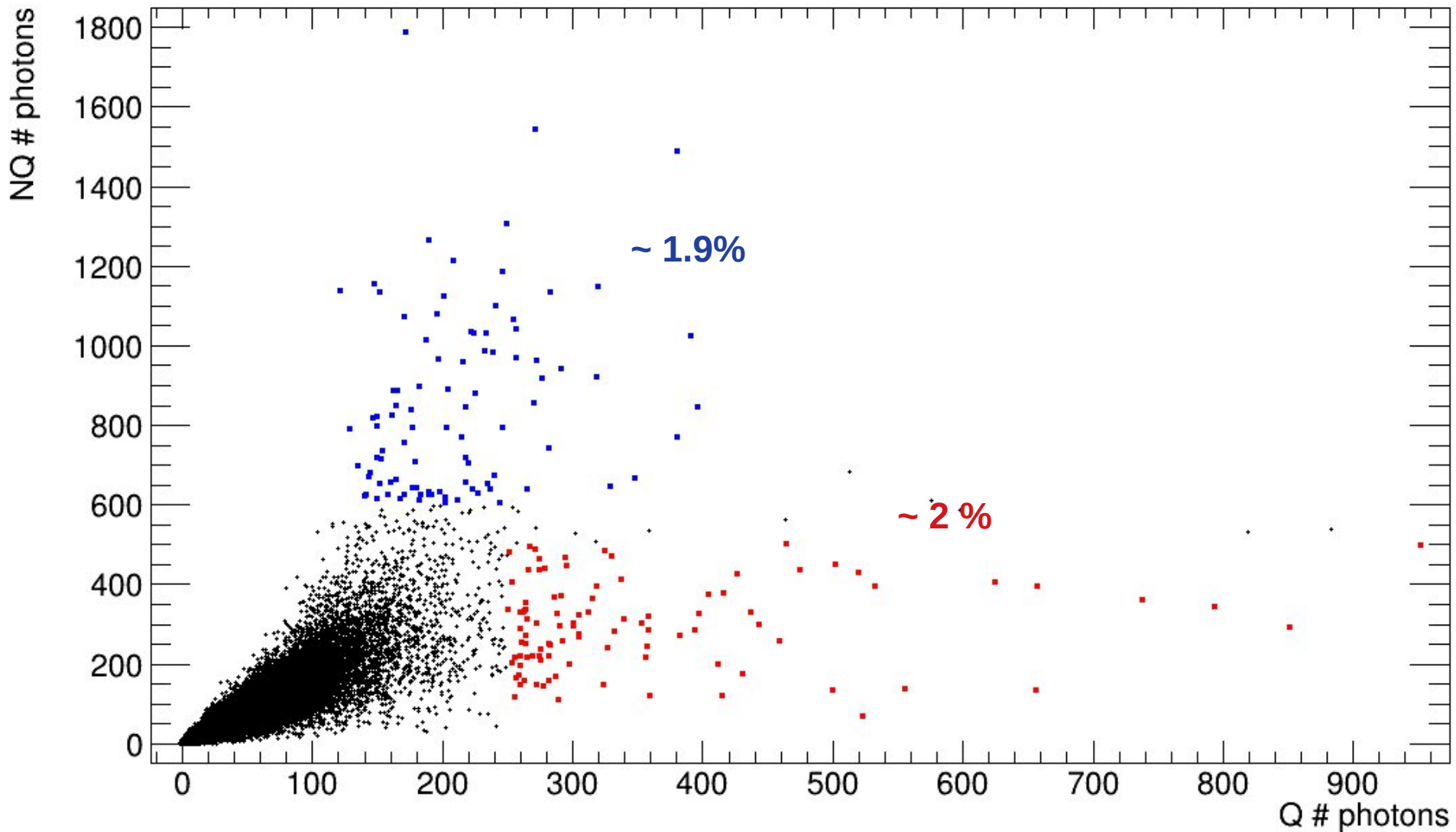
Doping 5 Scatter plot



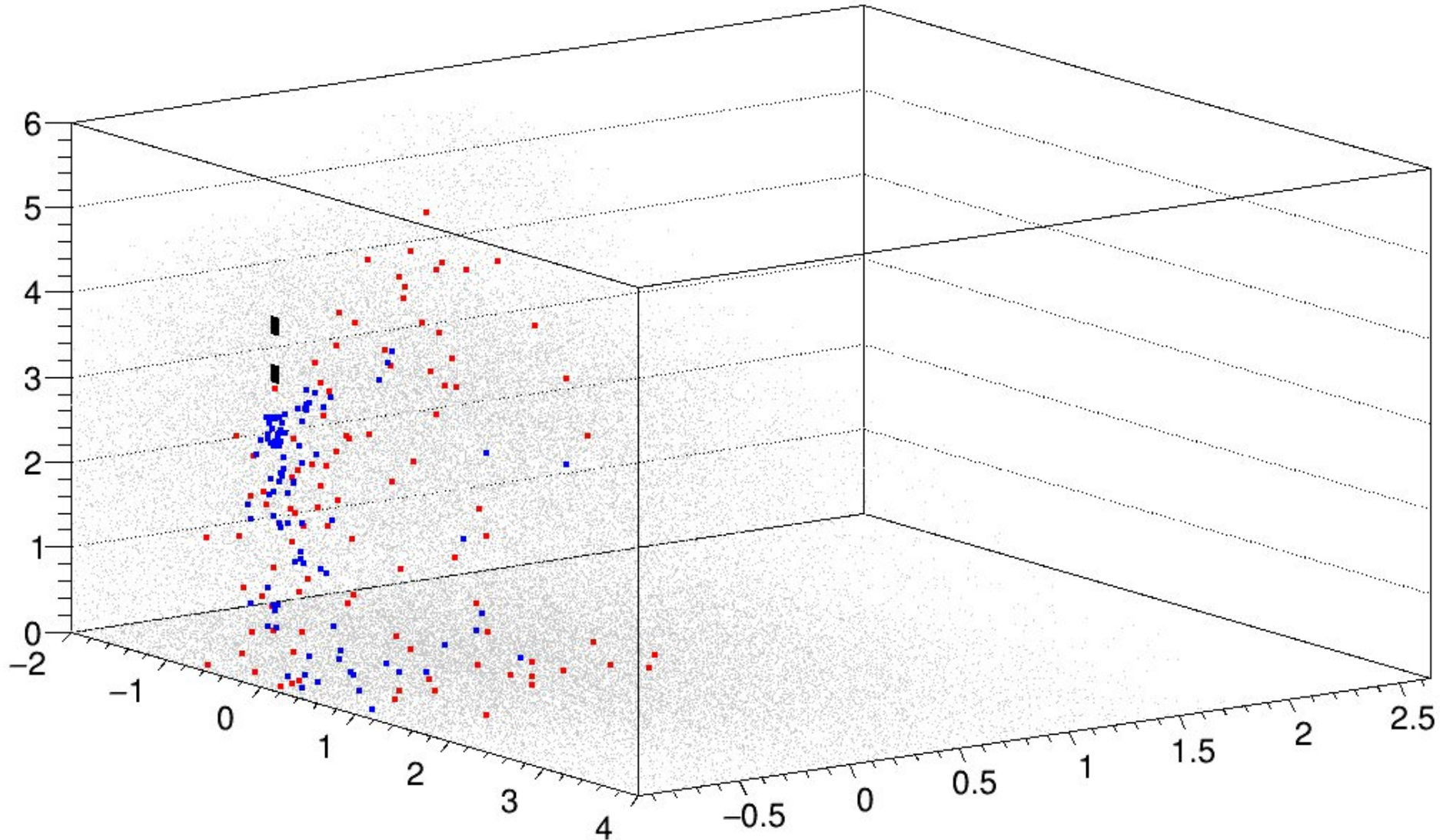
MC - Results



MC - Results



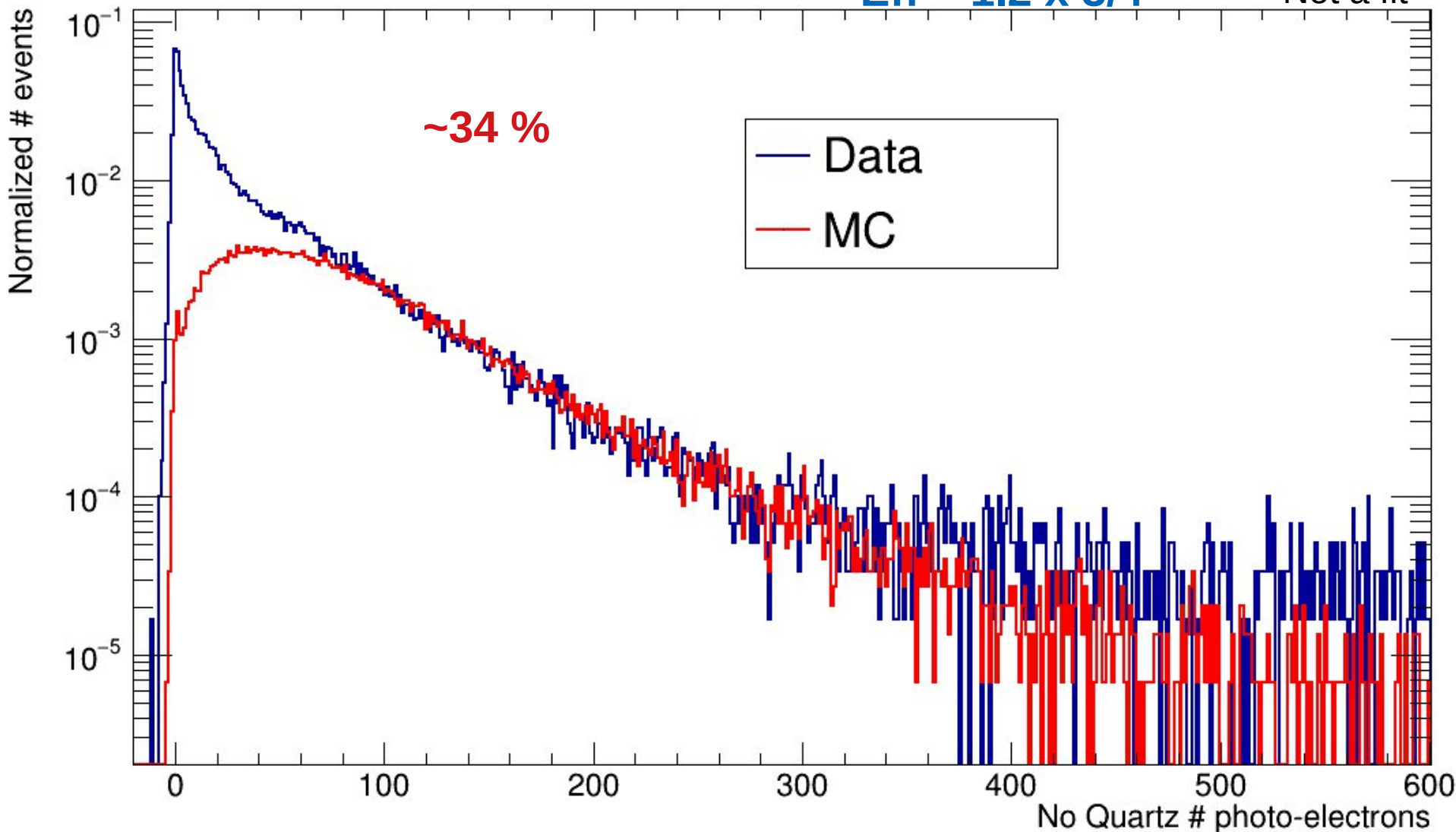
MC - Results



MC - Results

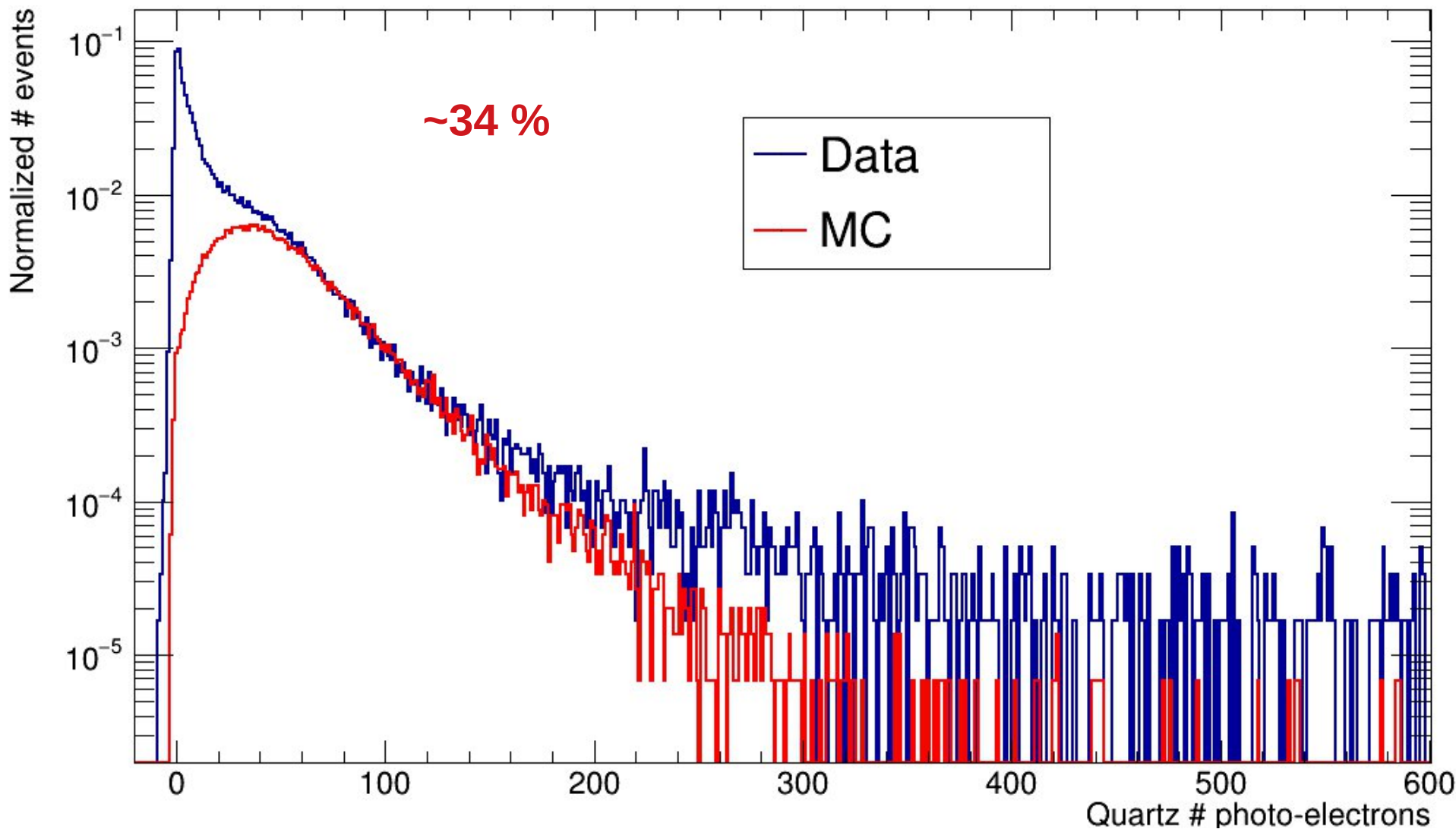
Eff = $1.2 \times 3/4$

Not a fit

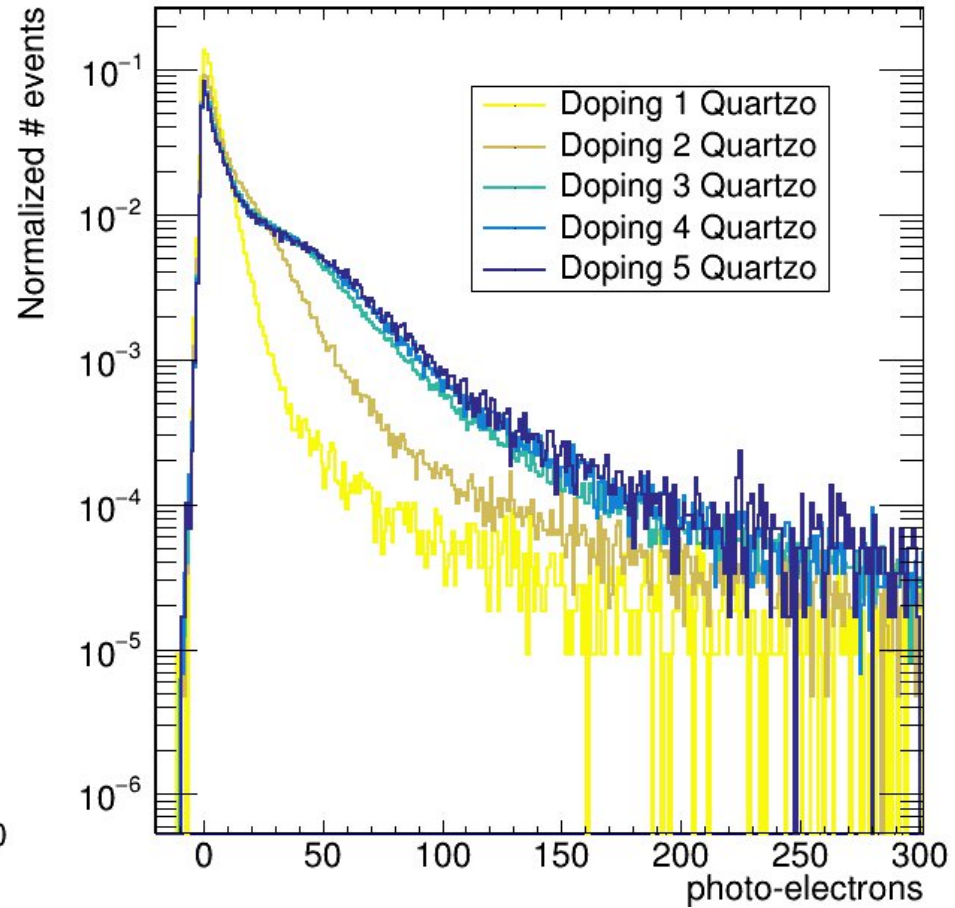
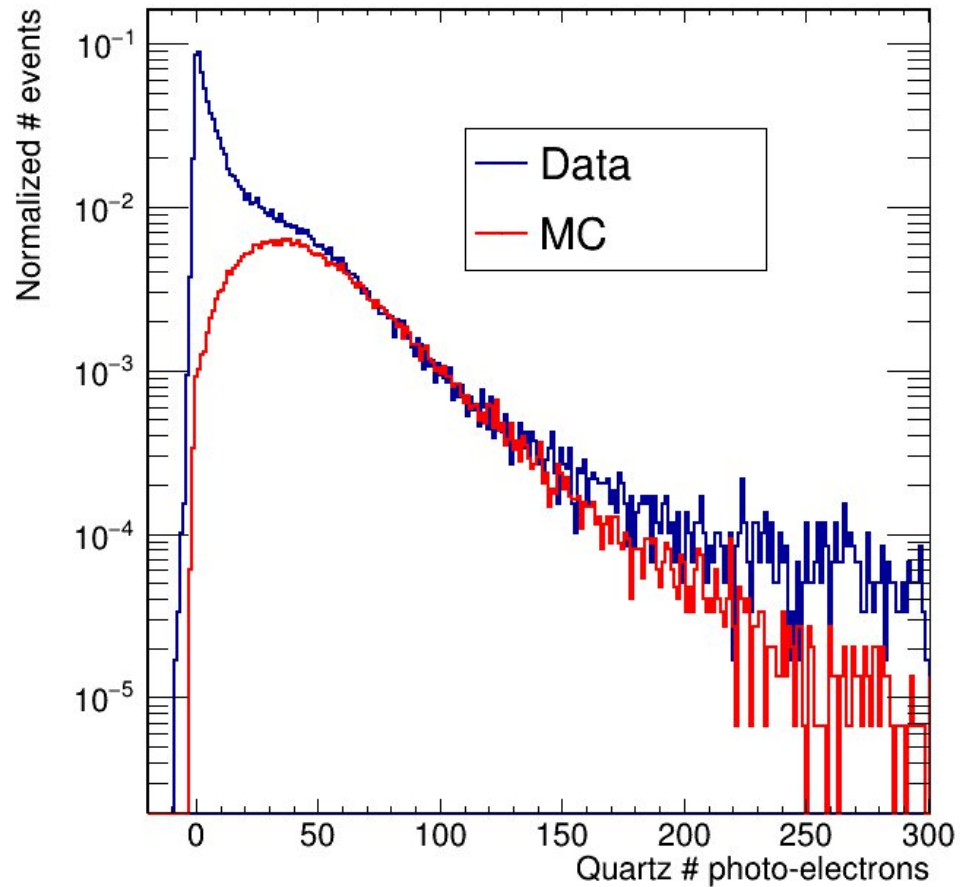


MC - Results

Eff = $1.2 \times 3/4 \times 0.8$ Not a fit



MC vs Data

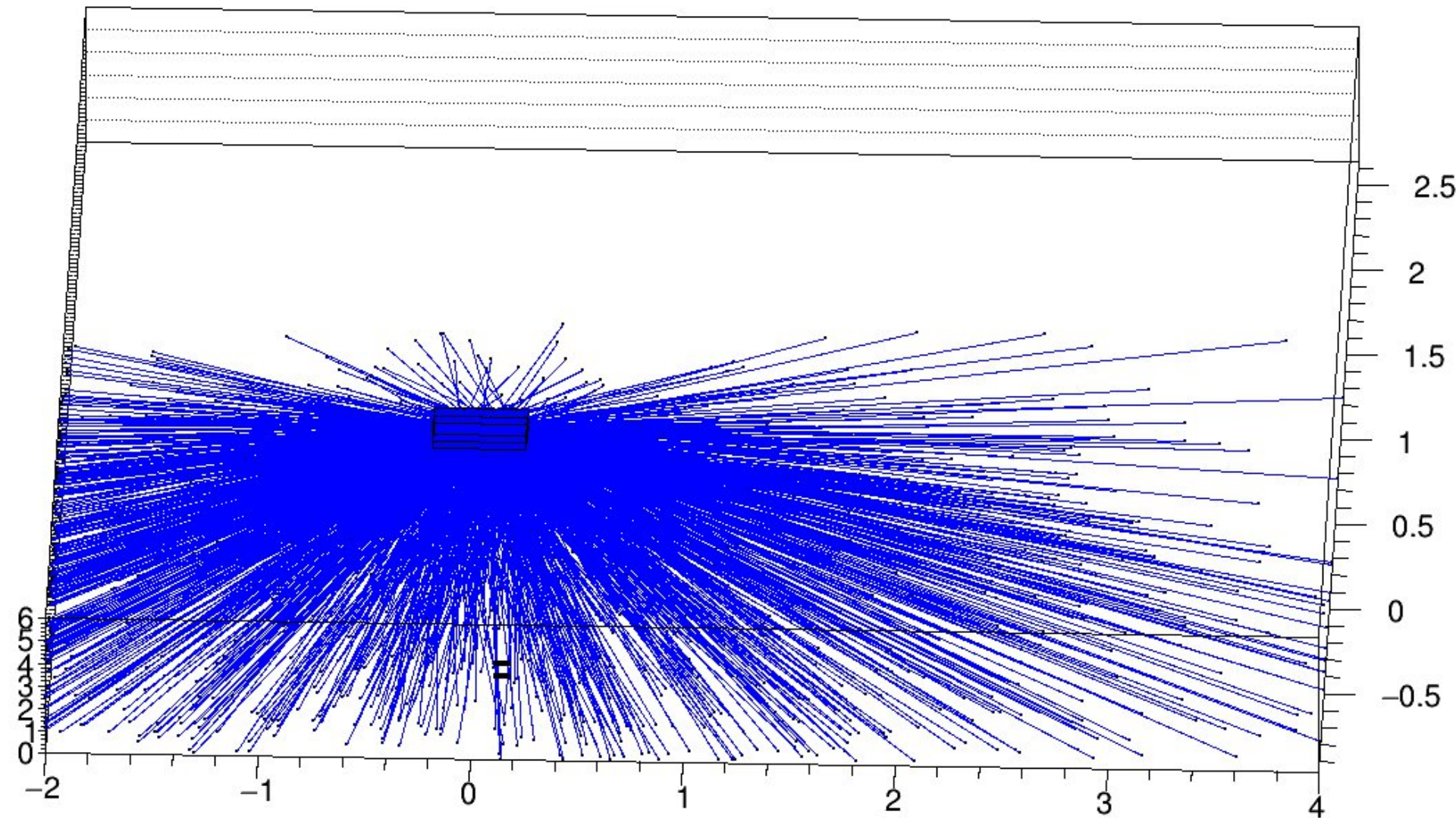


Summary

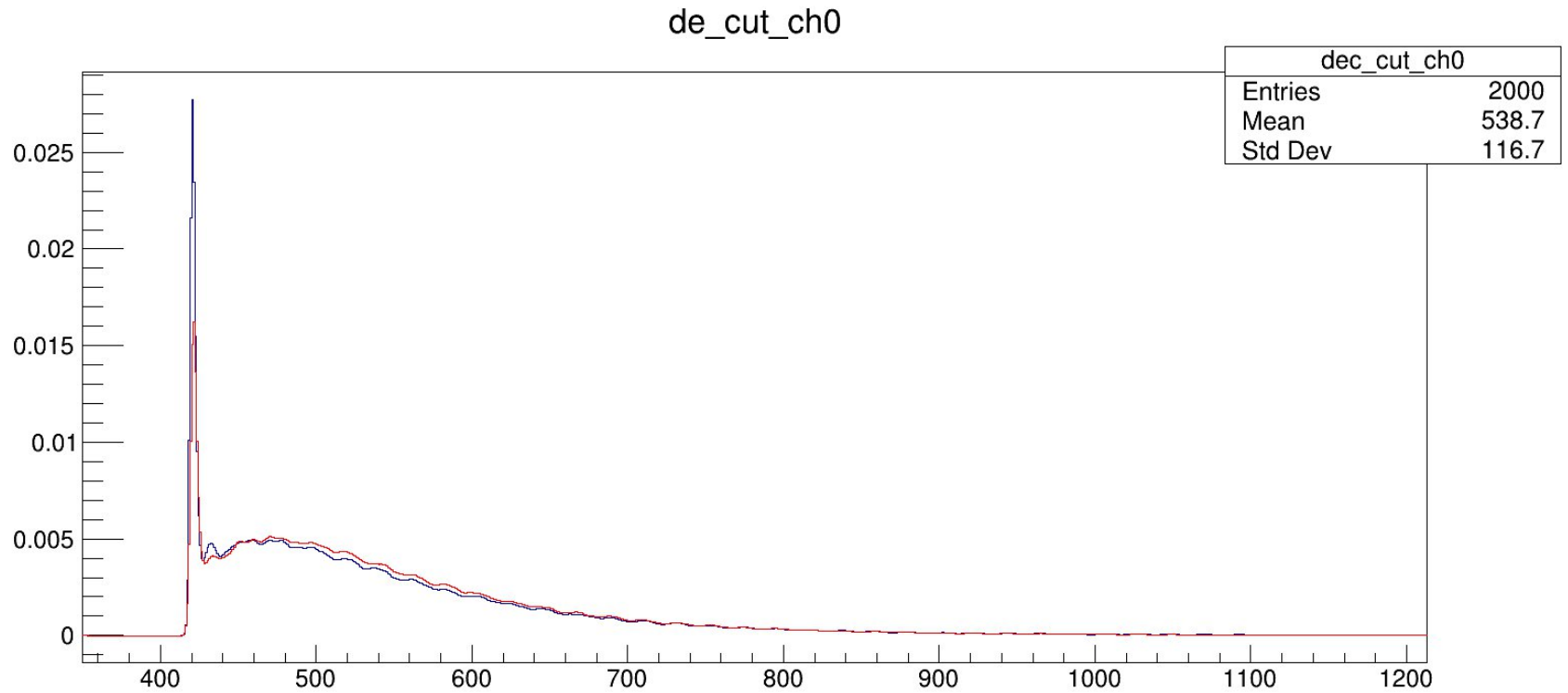
- Start from the problem with the two lobes we decided to investigate with a Toy Monte-Carlo simulation
- Toy model Monte-Carlo indicates that the two lobes are mainly geometrical effects with Rayleigh scattering.
- 3rd (and 4th) lobes did not show up in the simulation. Possible causes are still to be investigated. (Cherenkov, muons passing the APA, APA modules shadowing ?)
- The Biggest difference between data and simulation is the amount of events with zero or few photons (atmospheric showers giving “false triggers” ?)
- For MC simulation, average number of photons per trigger has a linear relation with Light Yield
- If we don't consider the Quartz window blocking light, MC simulation gives a ratio $Xe/(Xe+Ar) \sim 7\%$ higher than the one set. (we set 0.70 and the output was 0.75. With Quartz window the output is $0.6 \rightarrow 0.75 \cdot 0.8$)

Backup slides

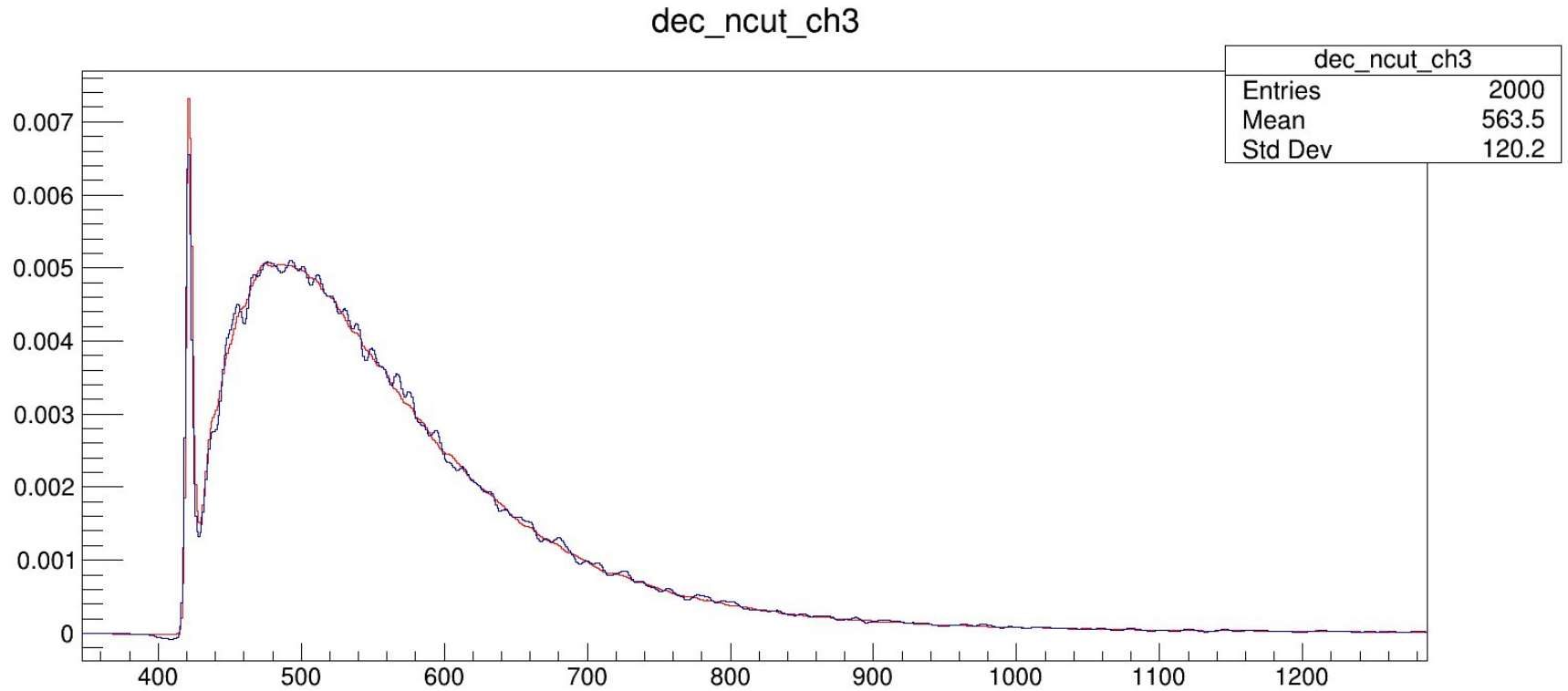
Backup slides



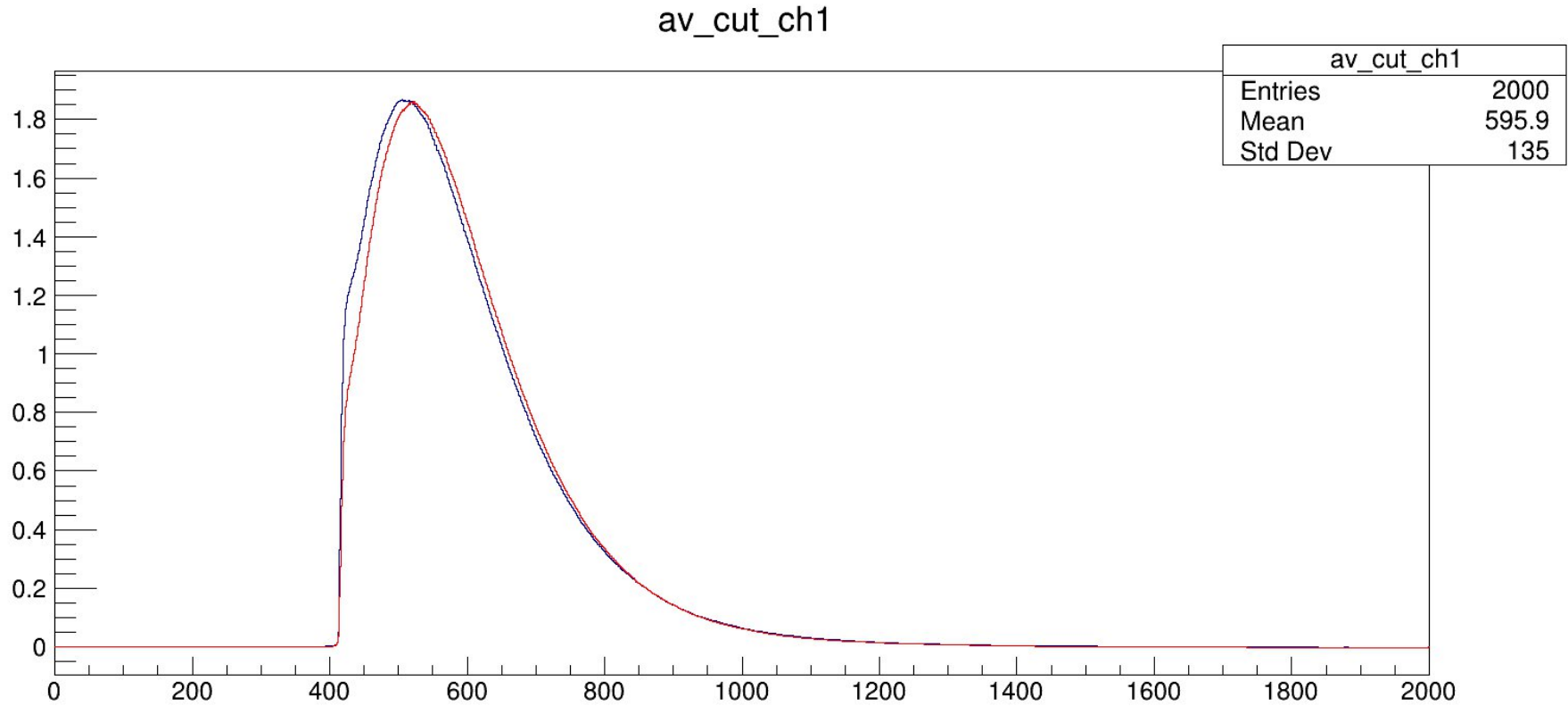
Backup slides



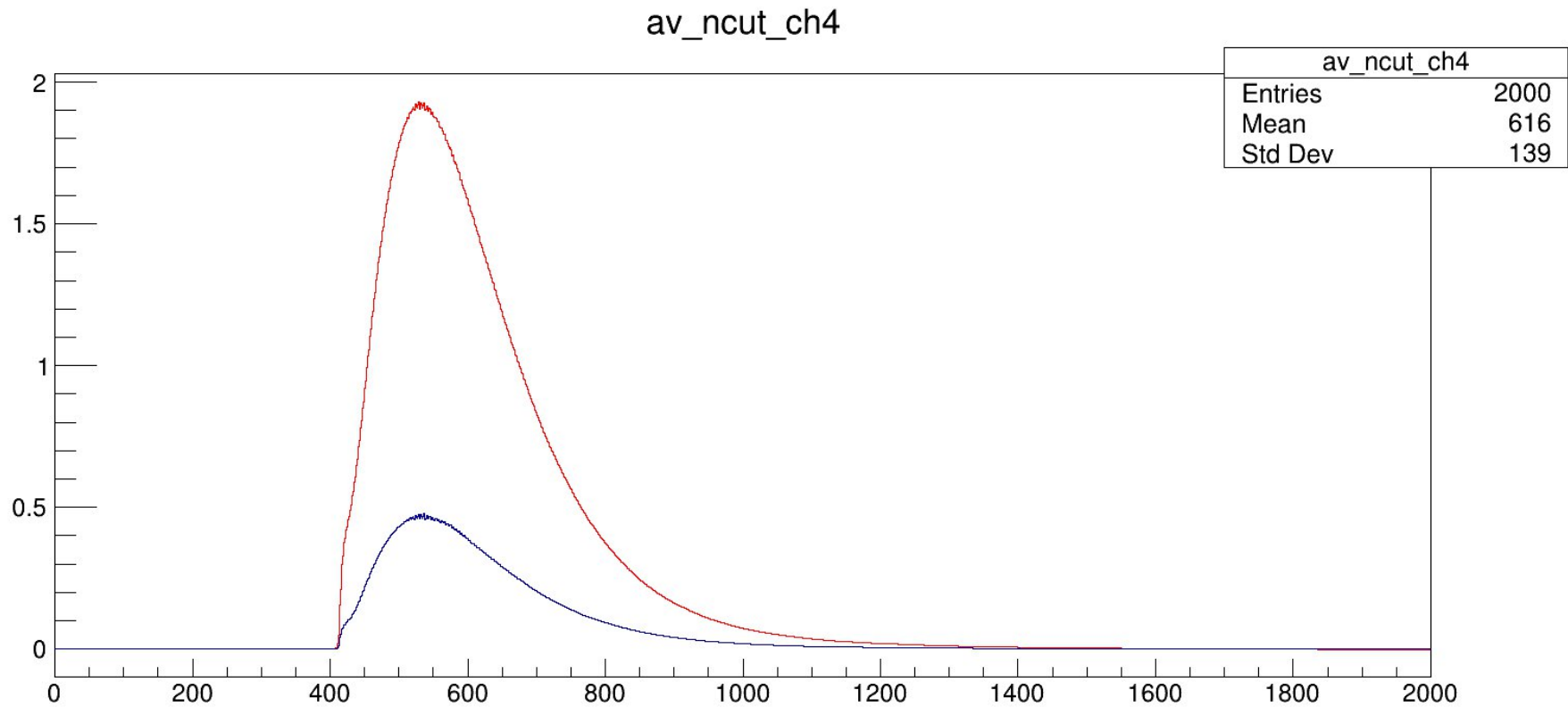
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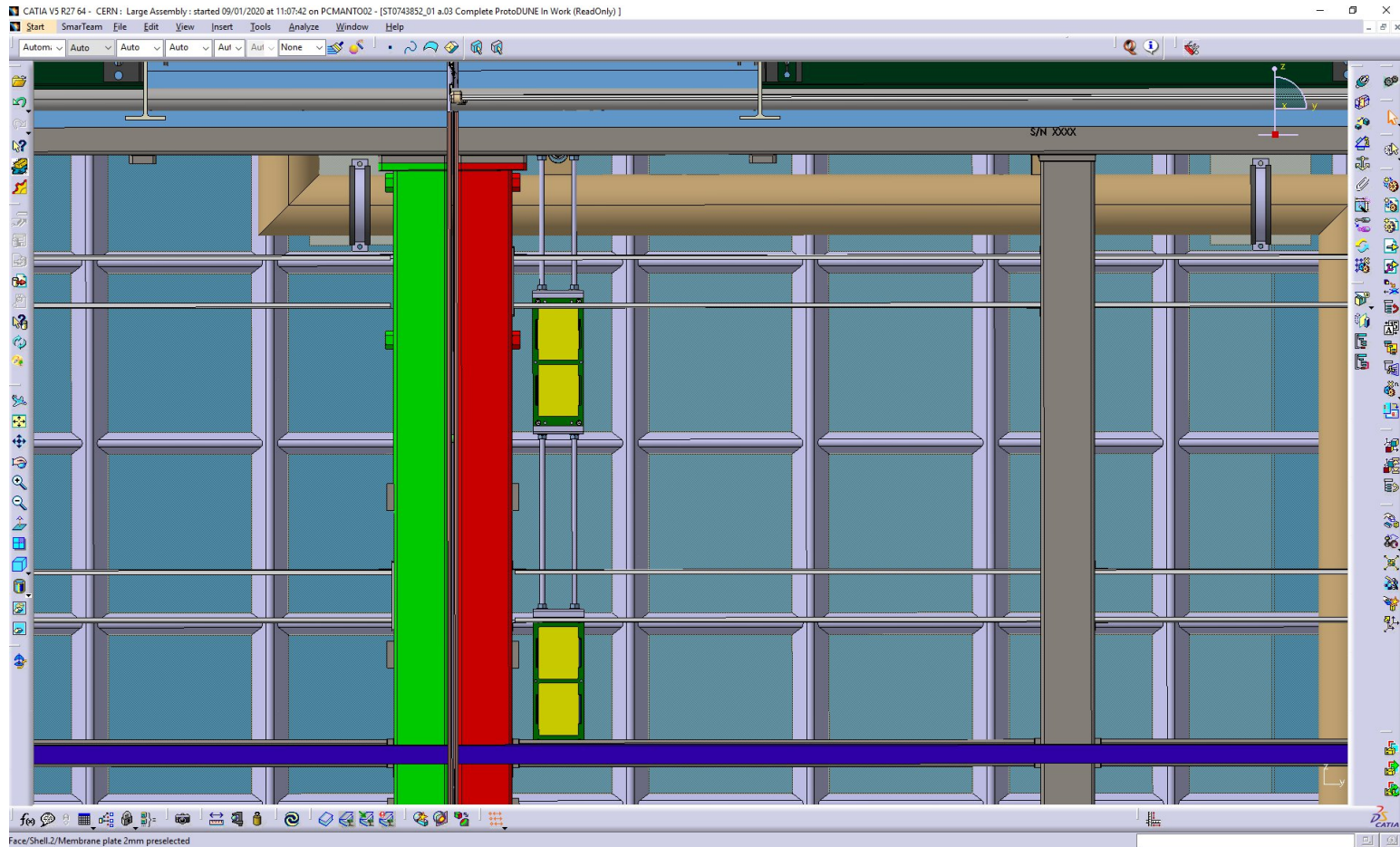
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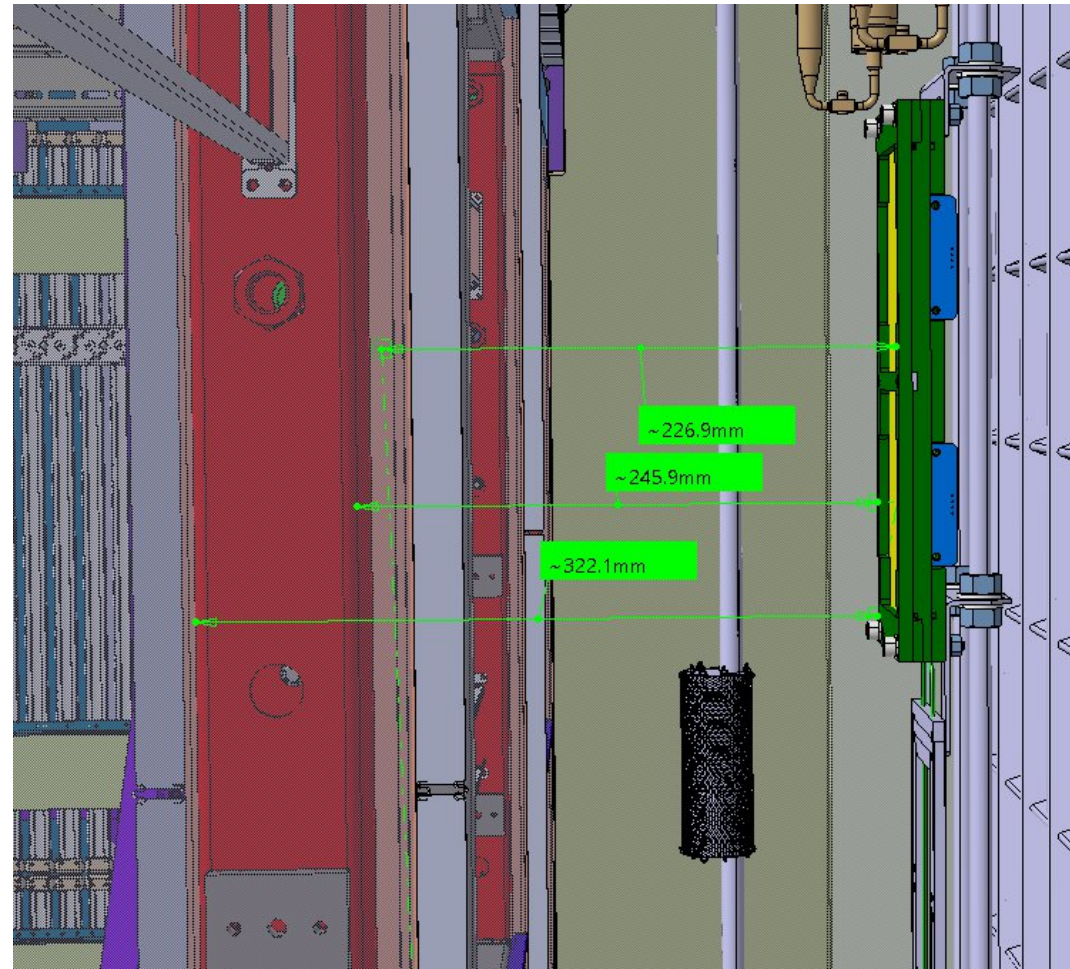
Backup slides



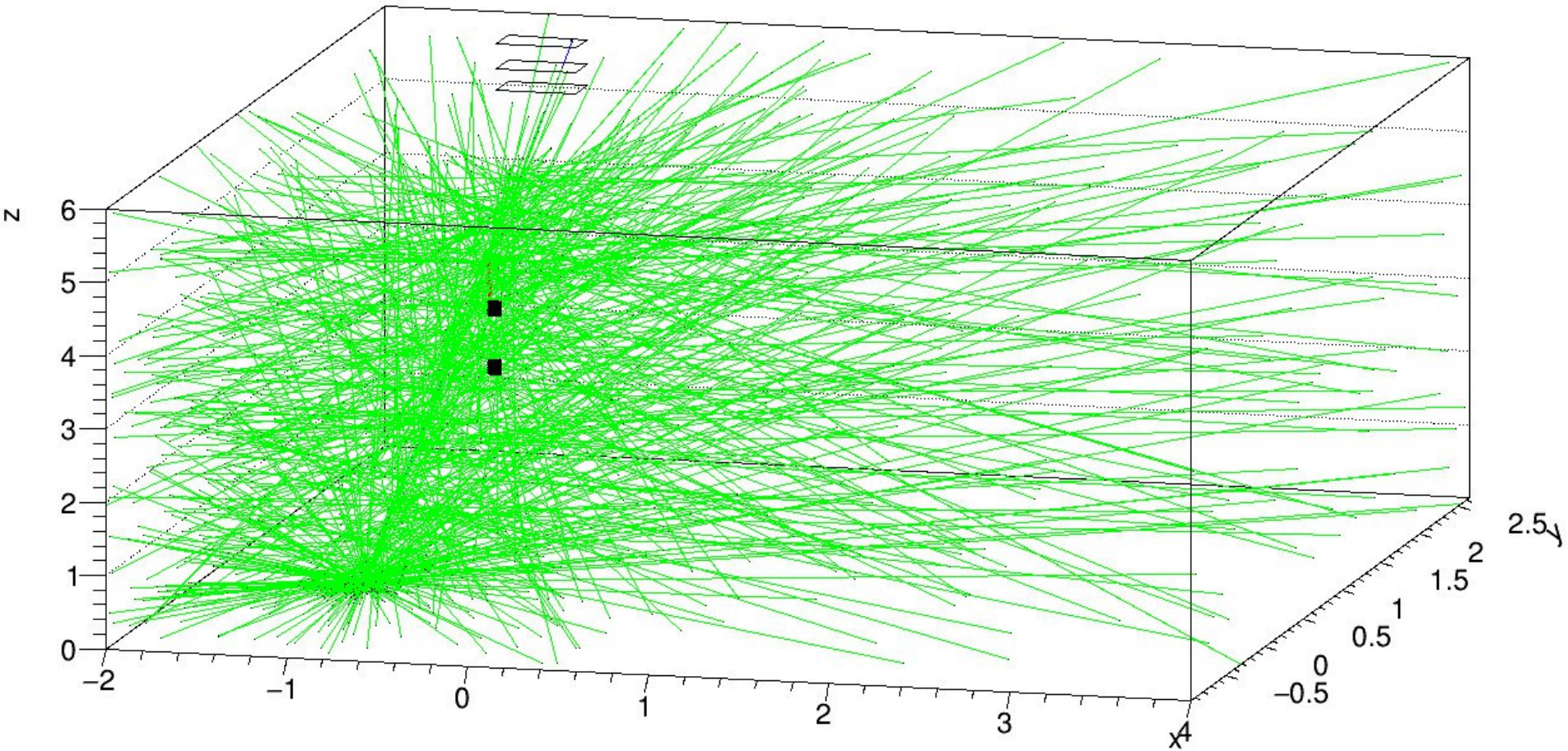
Backup slides



Backup slides

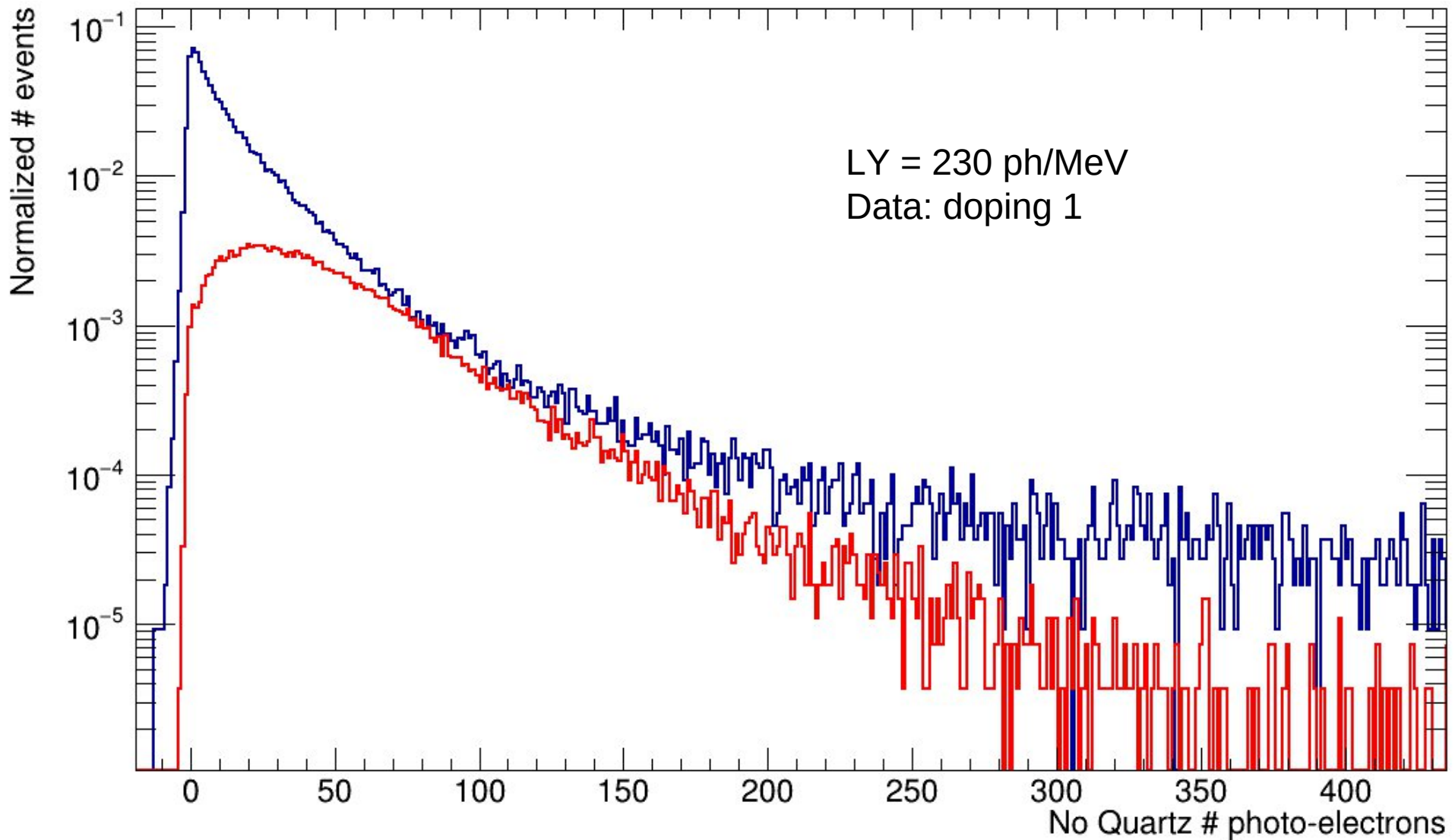


Backup slides

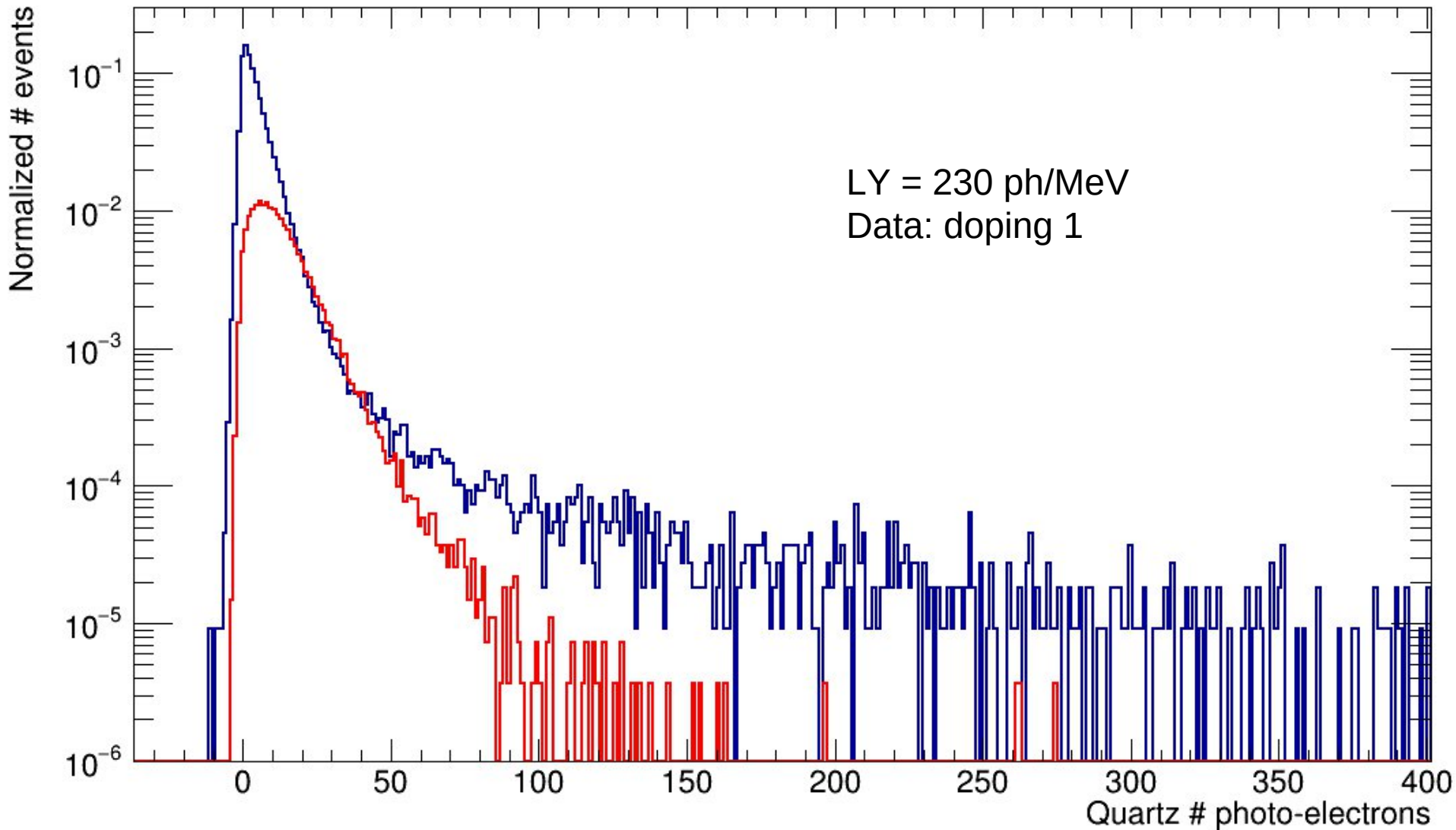


No Rayleigh scattering

Backup slides



Backup slides



Backup slides

