

## Coherent Elastic Neutrino-Nucleus Scattering: Theoretical and experimental impact

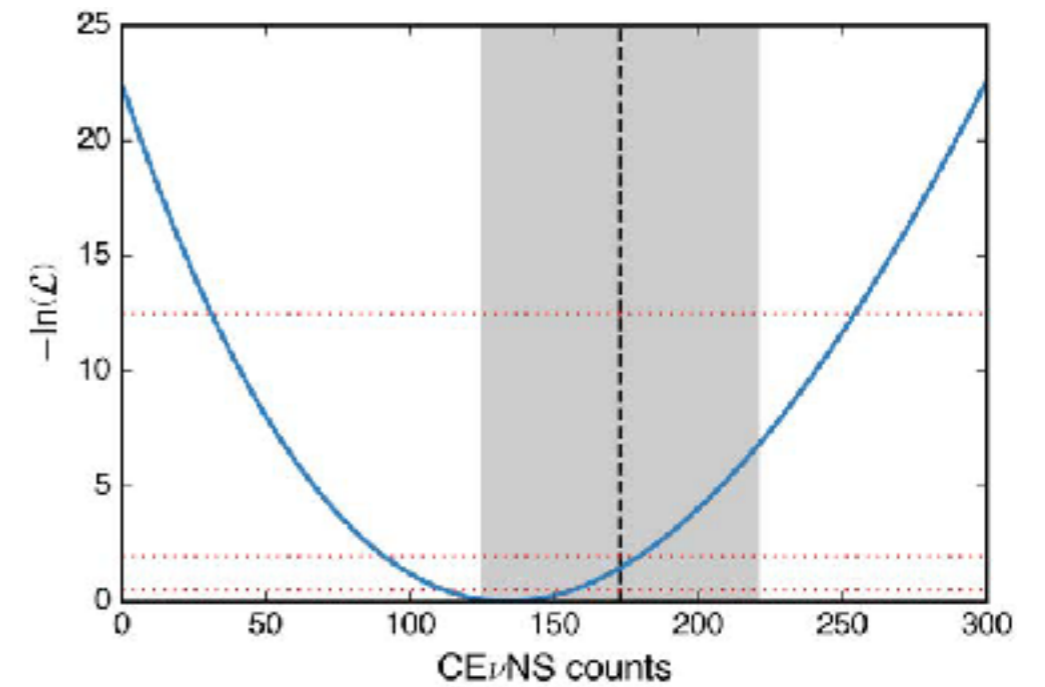
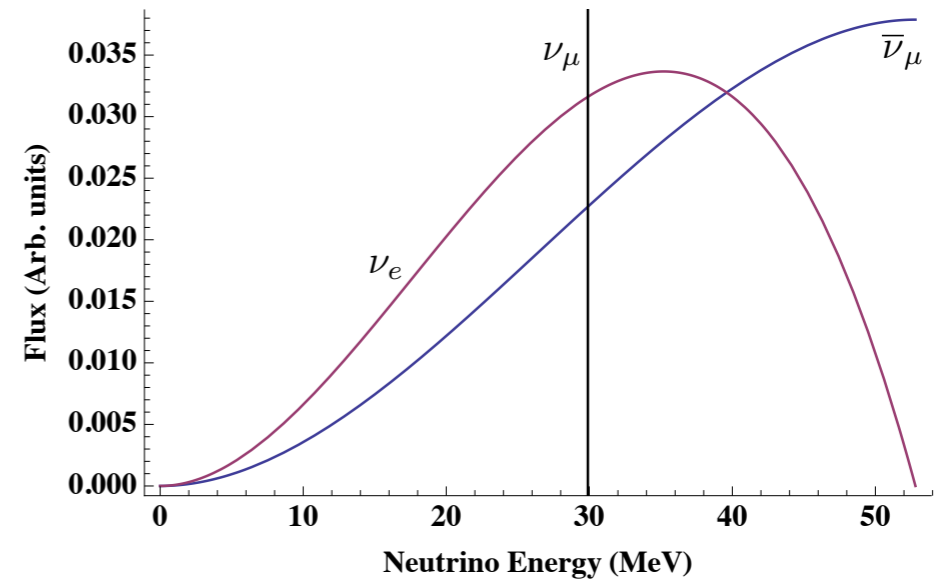
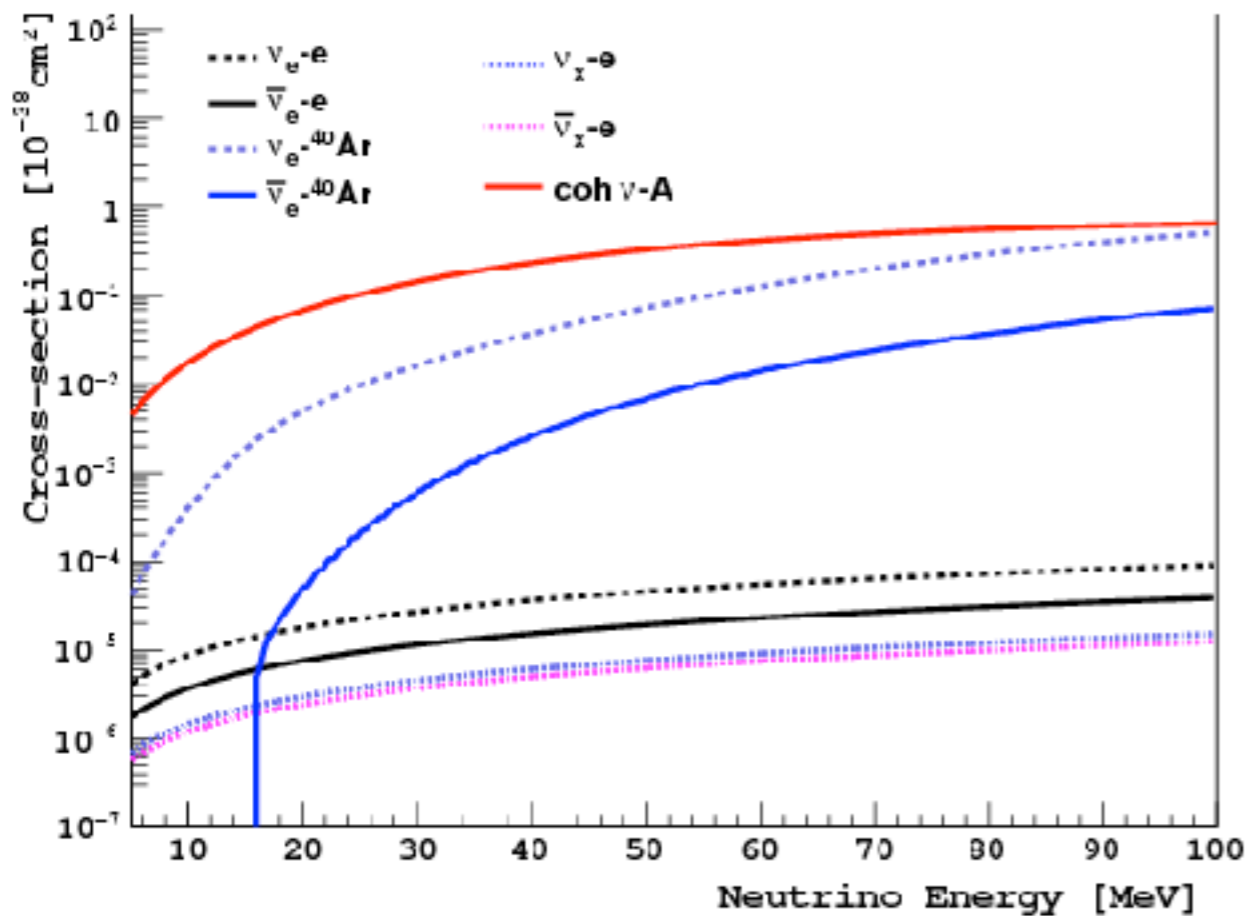
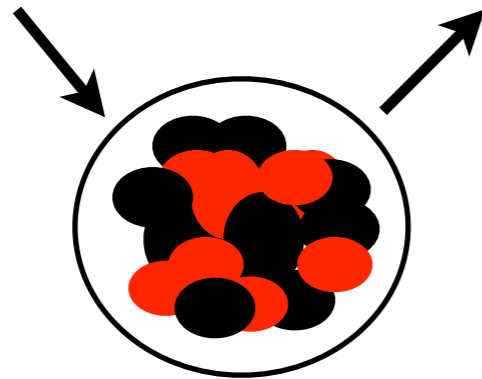
D. Aristizabal Sierra,<sup>1,2</sup> C. Augier,<sup>3</sup> A.B. Balantekin,<sup>4</sup> P. S. Barbeau,<sup>5,6</sup> V. A. Bednyakov,<sup>7</sup> I. A. Bernardi,<sup>8</sup> J. Billard,<sup>3</sup> C. Bonifazi,<sup>9</sup> N. S. Bowden,<sup>10</sup> M. Cadeddu,<sup>11</sup> D. Chernyak,<sup>12</sup> P. Coloma,<sup>13</sup> J. Daughhetee,<sup>8</sup> André de Gouvêa,<sup>14</sup> M. De Jésus,<sup>3</sup> M. Demarteau,<sup>15</sup> J. B. Dent,<sup>16</sup> P.B. Denton,<sup>17</sup> K. Ding,<sup>12</sup> V. De Romeri,<sup>18</sup> F. Dordei,<sup>11</sup> B. Dutta,<sup>19</sup> Yu. Efremenko,<sup>8</sup> J. Estrada,<sup>20</sup> Y. Farzan,<sup>21</sup> A. Fava,<sup>20</sup> M. Febbraro,<sup>15</sup> G. Fernandez Moroni,<sup>20</sup> E. Figueroa-Feliciano,<sup>22</sup> J. A. Formaggio,<sup>23</sup> A. Galindo-Uribarri,<sup>24</sup> F. Gao,<sup>25</sup> E. A. Garcés,<sup>26</sup> J. Gascon,<sup>3</sup> G. Gerbier,<sup>27</sup> J. Gehrlein,<sup>17</sup> G. Giroux,<sup>27</sup> C. Giunti,<sup>28</sup> M.P. Green,<sup>29,24,6</sup> W. Haxton,<sup>30,31</sup> M.R. Heath,<sup>24</sup> S. Hedges,<sup>5,6</sup> M. Hoferichter,<sup>32</sup> N. Jachowicz,<sup>33</sup> I. Jovanovic,<sup>34</sup> T. Katori,<sup>35</sup> I. Katsioulas,<sup>36</sup> Amir N. Khan,<sup>37</sup> D. Kim,<sup>38</sup> H. Kluck,<sup>39</sup> T.S. Kosmas,<sup>40</sup> N.A Kurinsky,<sup>20</sup> R.F. Lang,<sup>41</sup> S.Y. Lee,<sup>42</sup> B.G. Lenardo,<sup>43</sup> I. Levine,<sup>44</sup> Y. F. Li,<sup>45</sup> J. Liu,<sup>12</sup> L. Li,<sup>5,6</sup> P.A.N. Machado,<sup>20</sup> D.M. Markoff,<sup>46</sup> J. Mattingly,<sup>47</sup> B. Mauri,<sup>48</sup> J. Menéndez,<sup>49</sup> O. G. Miranda,<sup>50</sup> D. V. Naumov,<sup>7</sup> R. Neilson,<sup>51</sup> J. Newby,<sup>15</sup> J.L. Newstead,<sup>52</sup> K. Ni,<sup>53</sup> K. Nikolopoulos,<sup>36</sup> C. Nones,<sup>48</sup> D. Norcini,<sup>54,55</sup> K.J. Palladino,<sup>4</sup> V. Pandey,<sup>56</sup> D.K. Papoulias,<sup>40</sup> A. Parada,<sup>57</sup> J.C. Park,<sup>58</sup> D.S. Parno,<sup>59</sup> L. Pattavina,<sup>60,61</sup> E. Picciau,<sup>62,63</sup> M.-C. Piro,<sup>64</sup> J. Qi,<sup>53</sup> K. Ramanathan,<sup>54,55</sup> R. Rapp,<sup>59</sup> H. Ray,<sup>56</sup> J. Raybern,<sup>5</sup> G.C. Rich,<sup>54</sup> A. Ritz,<sup>65</sup> D. Rodrigues,<sup>66</sup> G. Sanchez Garcia,<sup>50</sup> T. Salagnac,<sup>3</sup> D.J. Salvat,<sup>67</sup> O. Sanders,<sup>50</sup> J. Schieck,<sup>39,68</sup> K. Scholberg,<sup>5</sup> A. Schwenk,<sup>69</sup> S. Shin,<sup>42</sup> I.M. Shoemaker,<sup>70</sup> V. Sibille,<sup>23</sup> N.J.C. Spooner,<sup>71</sup> R. Strauss,<sup>61</sup> L.E. Strigari,<sup>38</sup> B.D. Suh,<sup>67</sup> J. Suhonen,<sup>72</sup> A.M. Suliga,<sup>73</sup> Z. Tabrizi,<sup>70</sup> V. Takhistov,<sup>74</sup> R. Tayloe,<sup>67</sup> M. Toups,<sup>20</sup> M. Tórtola,<sup>18</sup> M. Tripathi,<sup>56</sup> José W. F. Valle,<sup>18</sup> M. Vidal,<sup>27</sup> M. Vignati,<sup>75</sup> M. Vivier,<sup>48</sup> V. Wagner,<sup>61</sup> J. W. Walker,<sup>16</sup> J. Xu,<sup>10</sup> Y. Y. Zhang,<sup>45</sup> J. Zettlemoyer,<sup>67</sup> and I. Savvidis<sup>76</sup>

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NF03 Workshop  
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# Coherent elastic neutrino-nucleus scattering (CEvNS)

neutrino                      neutrino

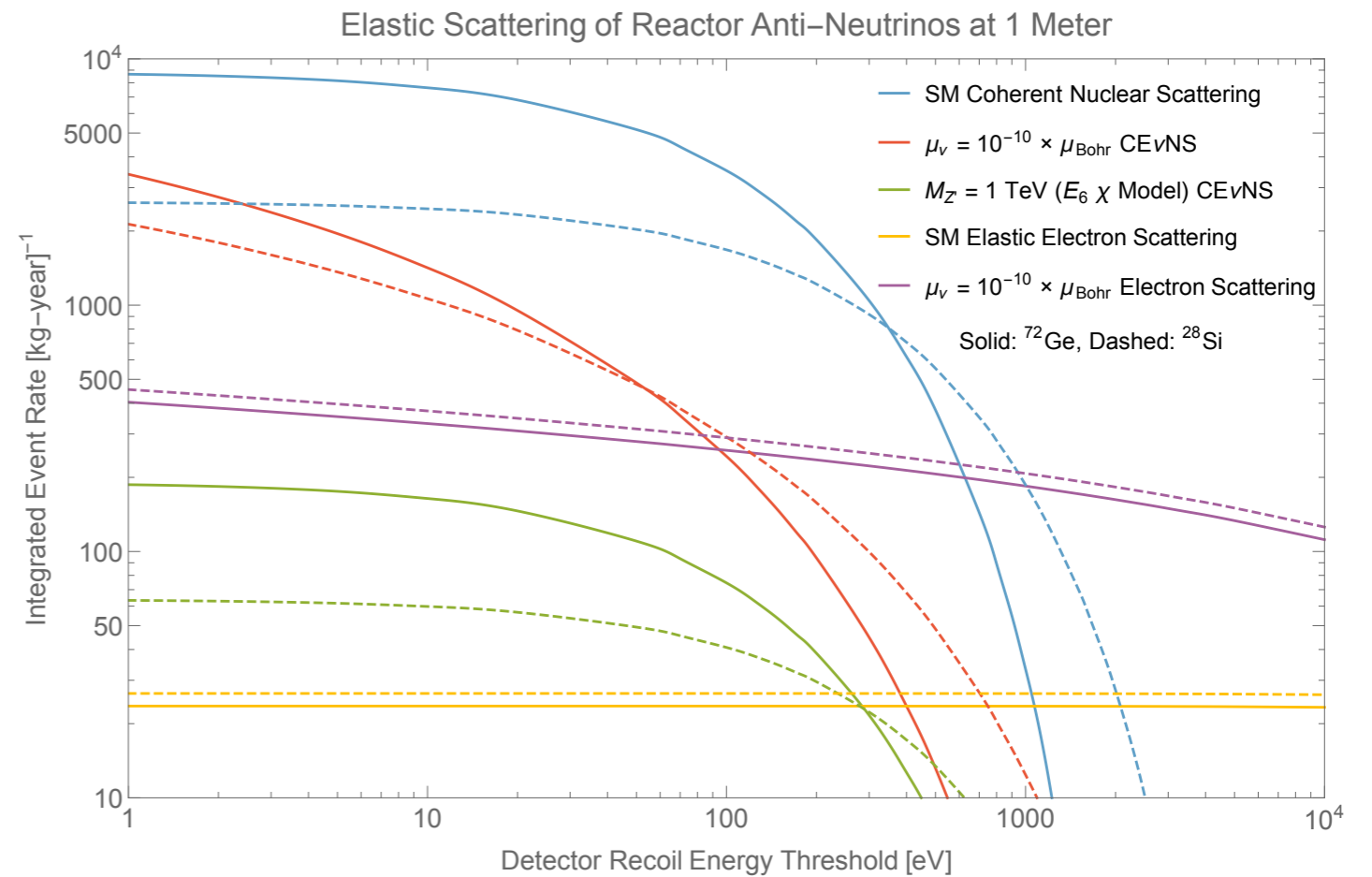
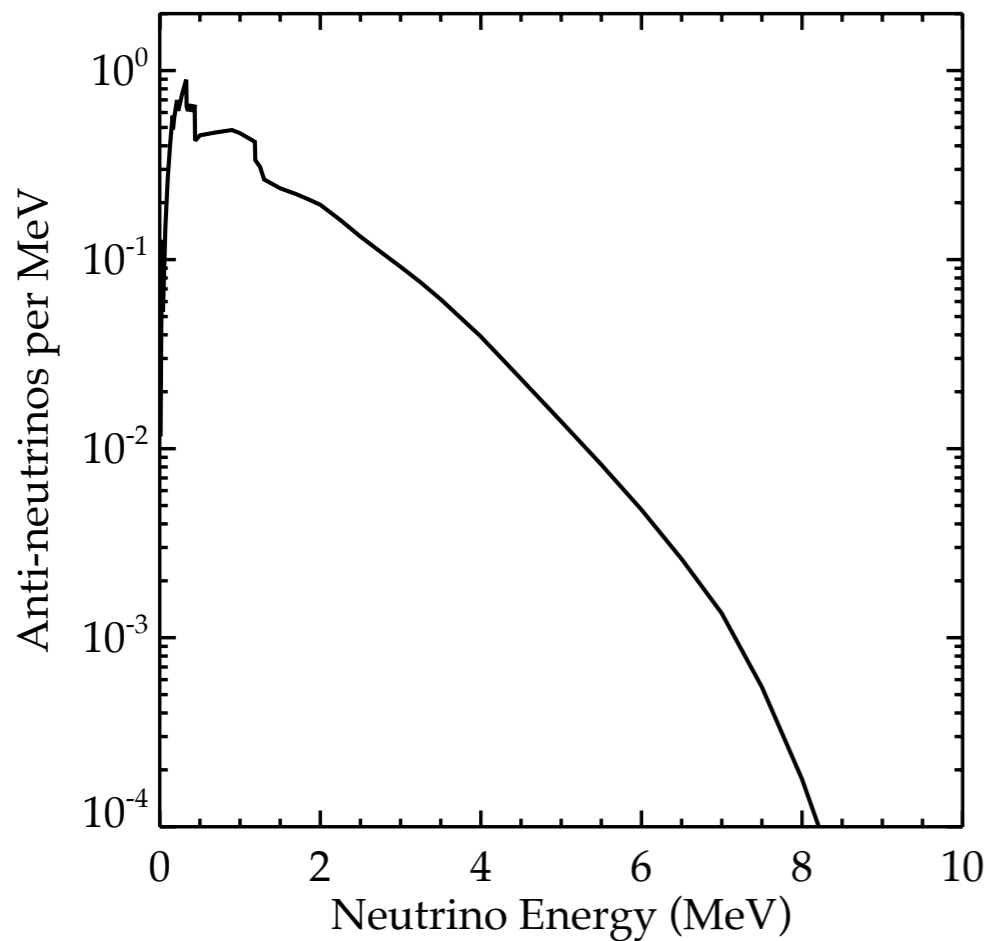


Brice et al, 1311.5958

COHERENT detection papers:  
 Csl: 1801.05546  
 Argon: 2003.10630

# CEvNS at nuclear reactors

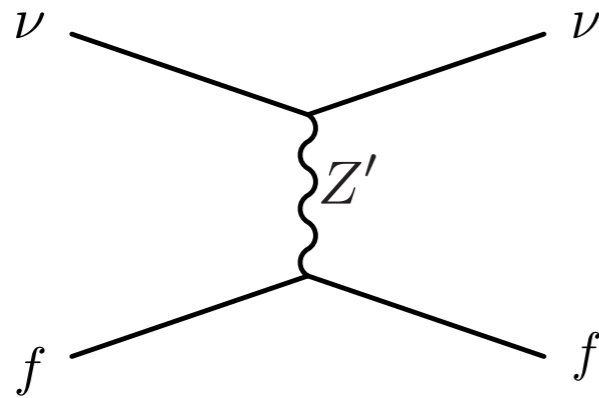
CEvNS at reactors requires low backgrounds, and low threshold detectors



# Searches for new physics with CEvNS

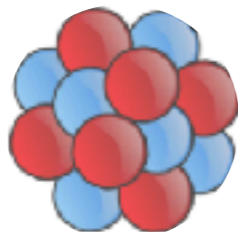
## Non-standard/generalized interactions

Scholberg 2005; Barranco 2005; Coloma et al. 2018;  
Liao & Marfatia 2017; Aristizabal-Sierra et al. 2018



## Nuclear form factors/charge radius

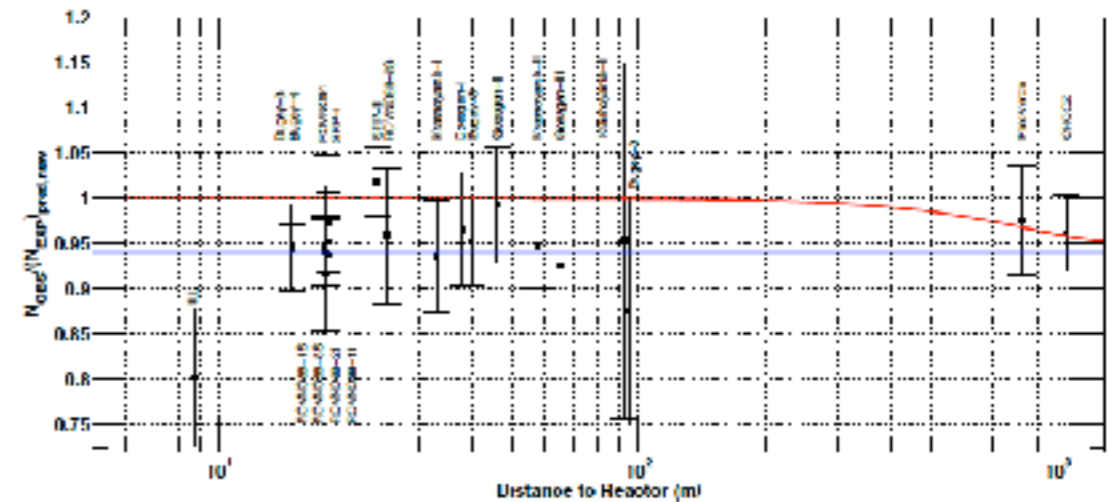
Patton et al. 2013; Cadeddu et al. 2018;  
Ciuffoli et al. 2018



## Sterile neutrinos

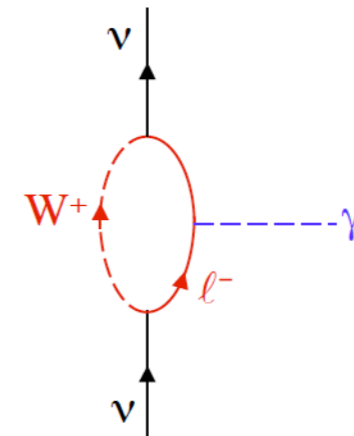
Anderson et al. 2010; Dutta et al. 2015; Kosmas et al. 2017,  
Blanco et al. 2019

Reactor, Gallium anomalies



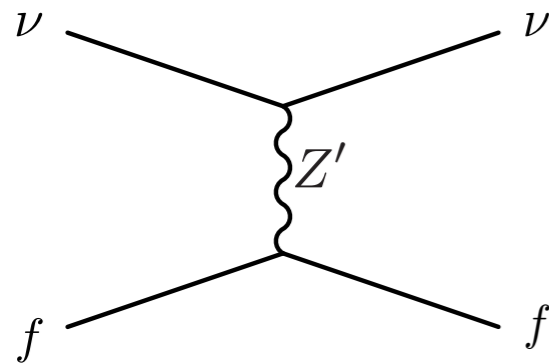
## Magnetic moment

Vogel & Engel 1989



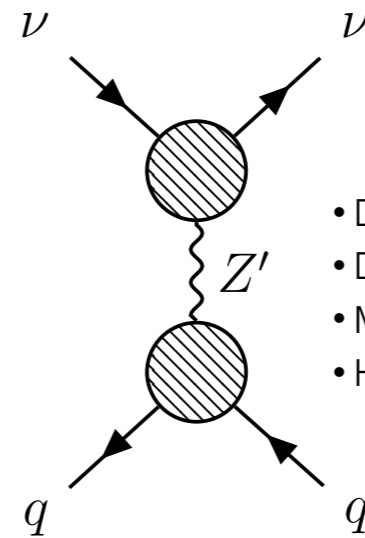
# Non-standard neutrino interactions (NSI)

New physics searches facilities by both energy and timing distributions in COHERENT

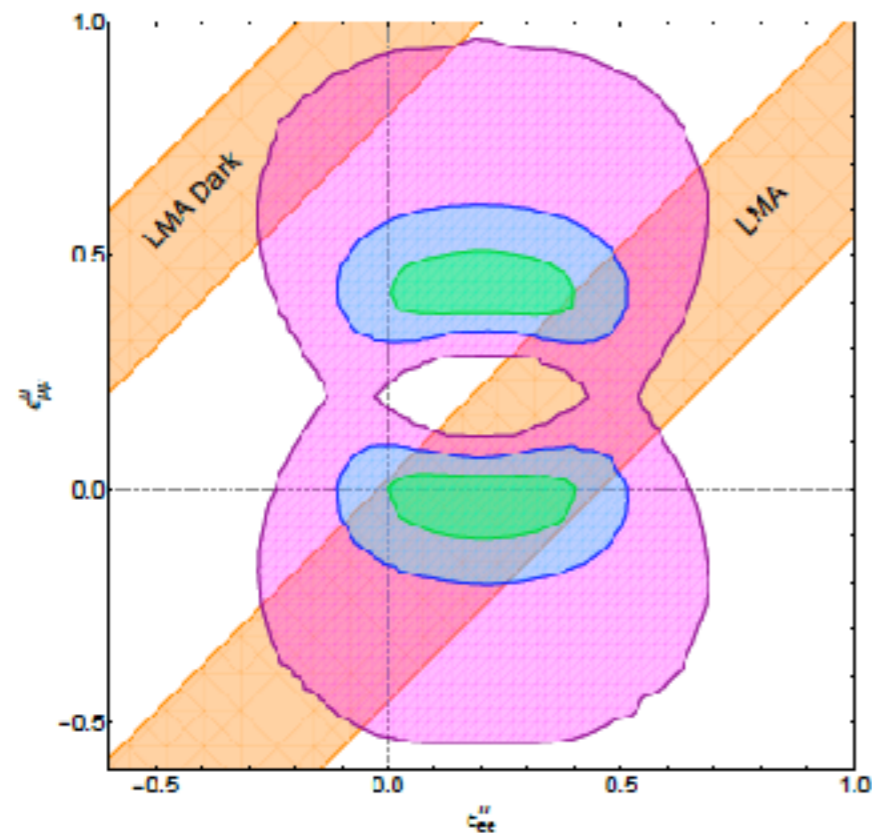


- Heavy mediators parameterized with EFT
- UV complete models can give large NSI
- Independent constraints from oscillation experiments

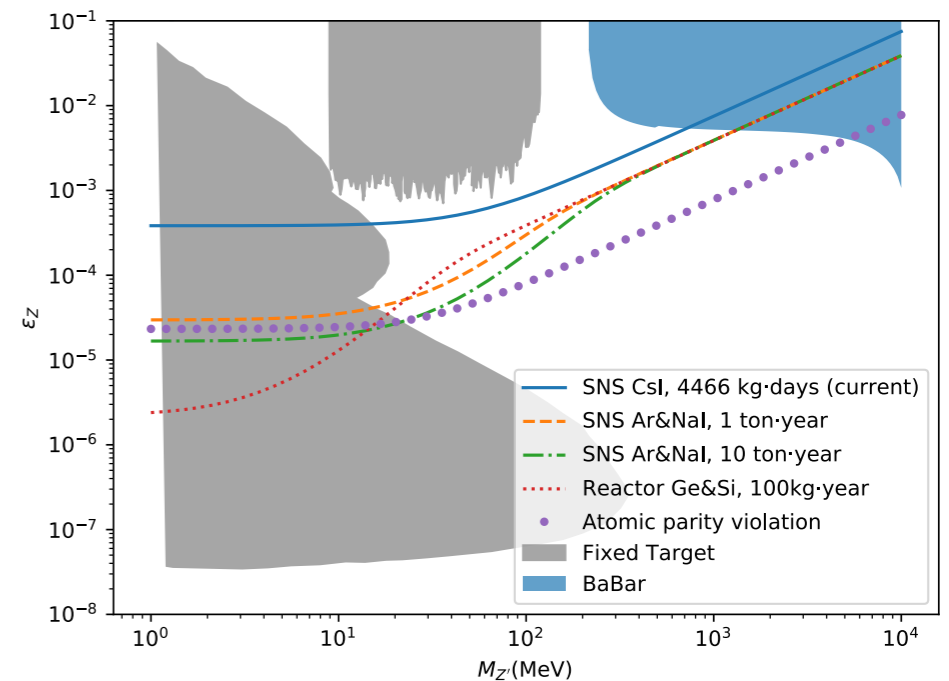
Dutta, Liao, Sinha, Strigari PRL 2019; Giunti PRD 2020



- Dark hypercharge gauge boson
- Dark Z boson
- Model with mu-tau flavor symmetry
- Hidden Sector Fermions

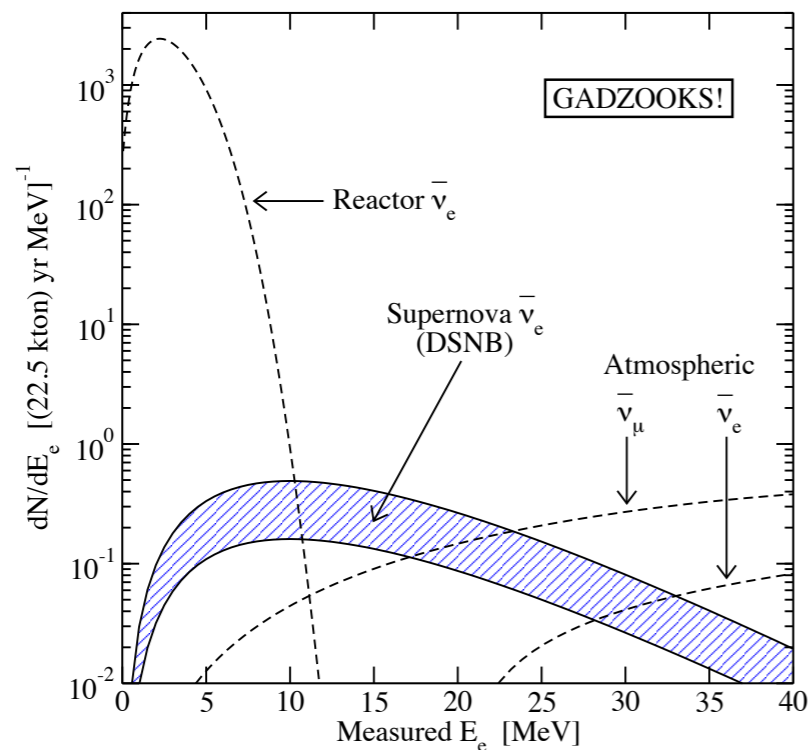
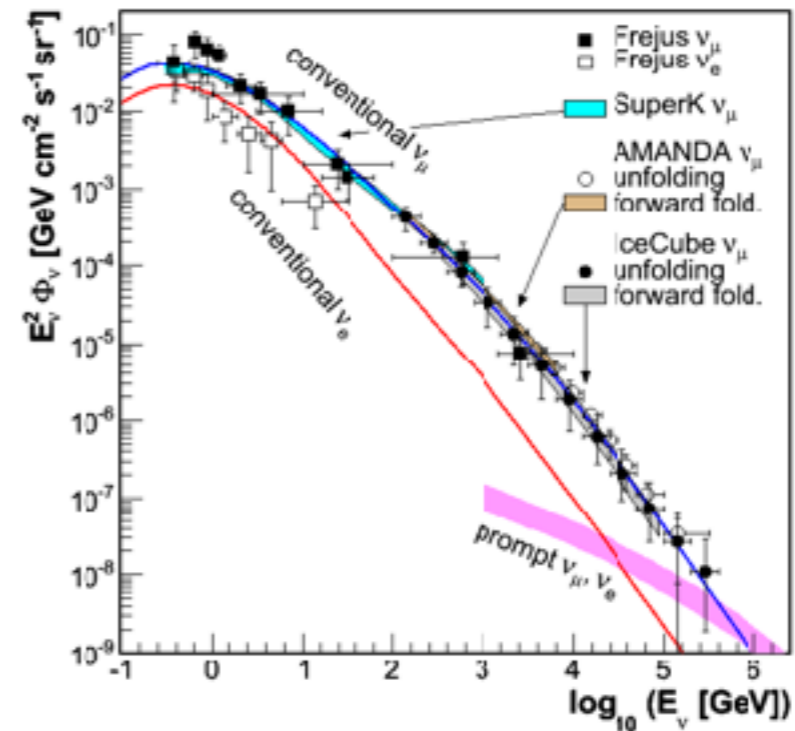
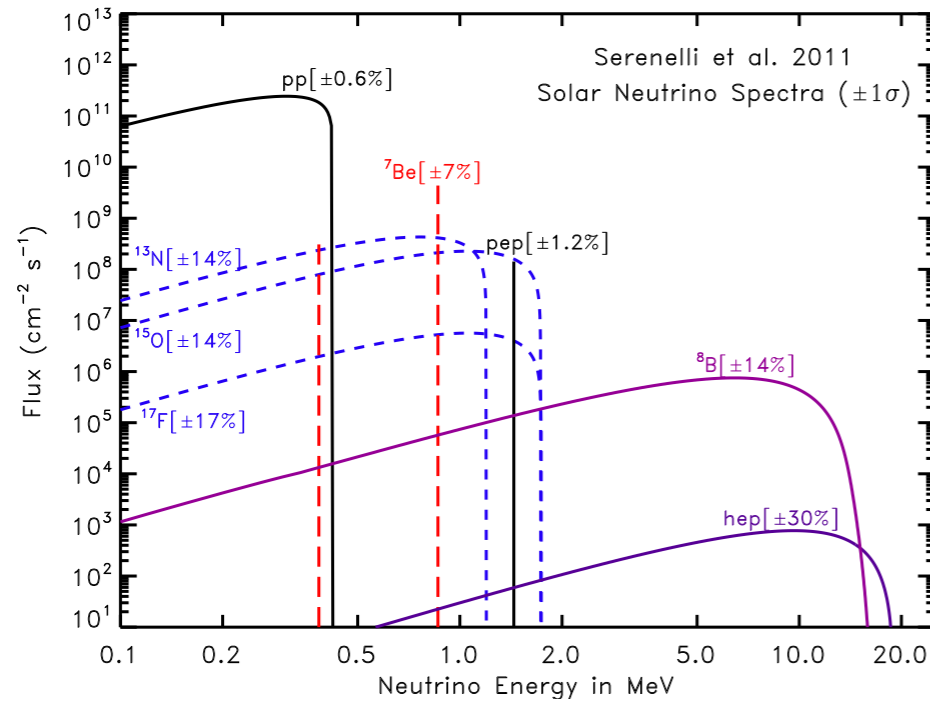


Coloma, Denton, Gonzalez-Garcia, Maltoni 2017;  
Denton, Farzan, Shoemaker 2018; Denton & Gehrlein 2020



Lindner et al. 2017; Farzan et al. 2018; Abdullah et al. 2018, Brdar et al. 2018

# CEvNS from astrophysical sources



## Sun

Neutral current  $^8\text{B}$  energy spectrum  
 CEvNS + electron scattering events (Recent Xenon nT)

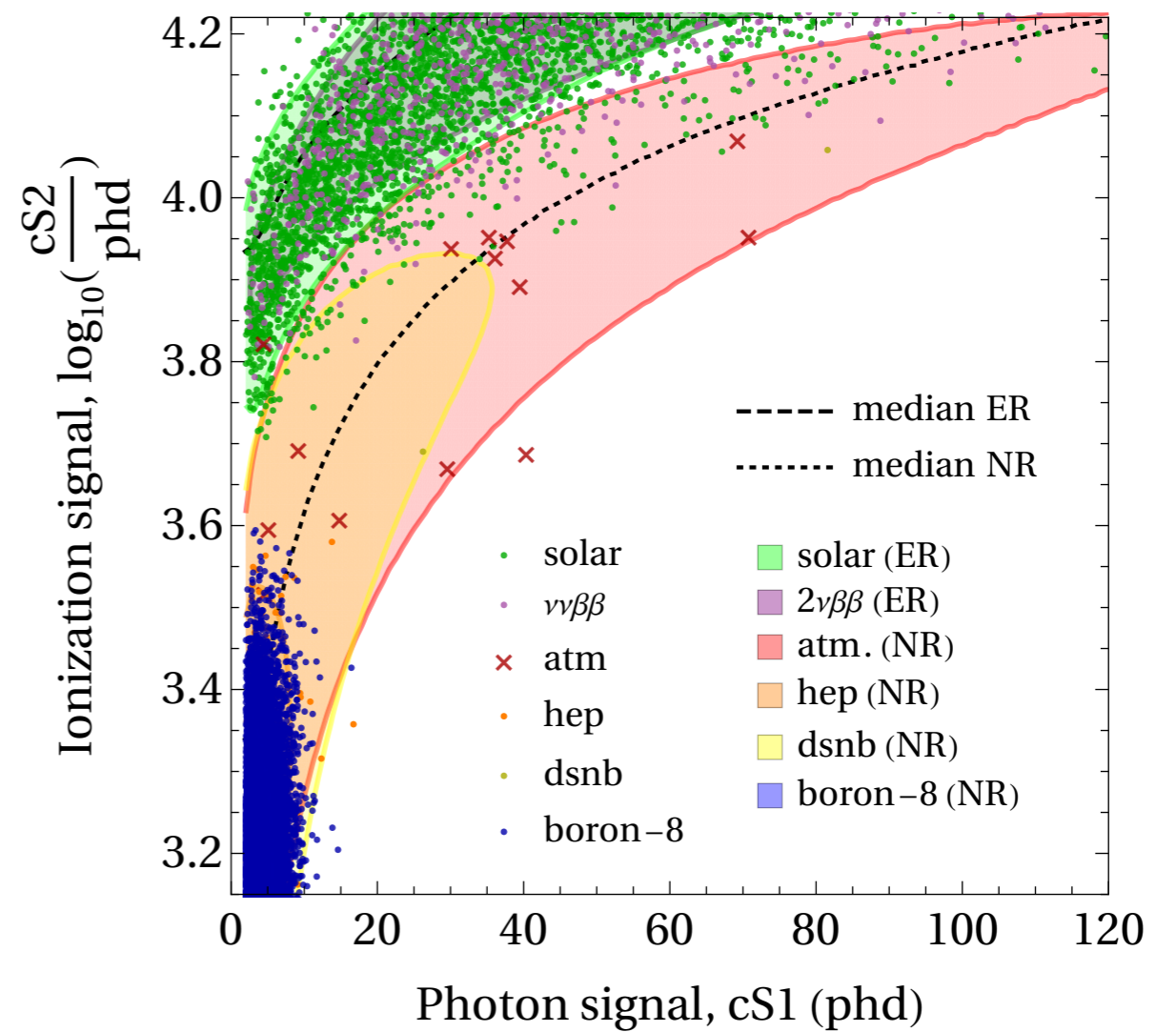
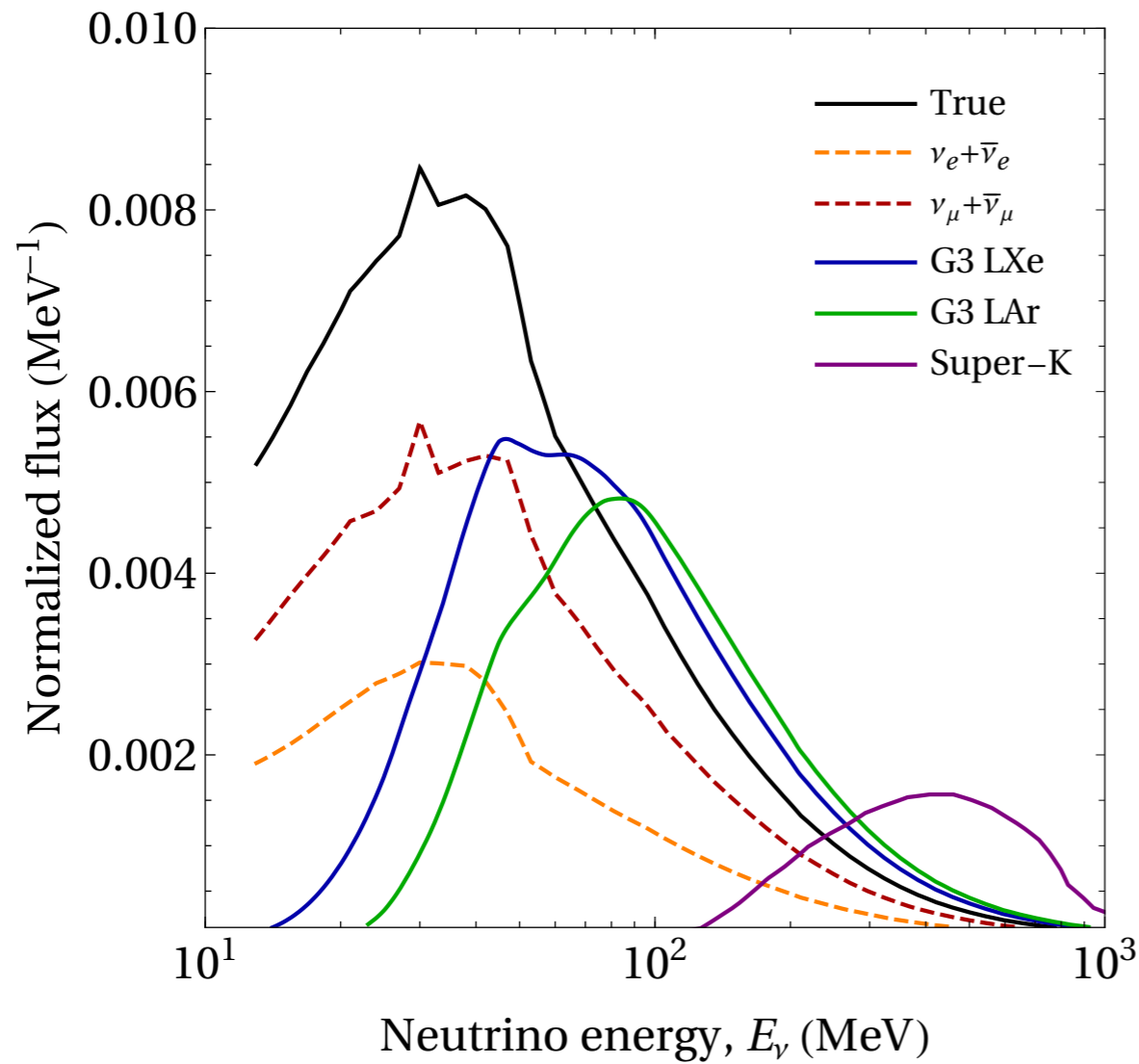
## Atmosphere

Low energy ( $< 50$  MeV) atmospheric neutrinos  
 Flux uncertainties, dependence on detector location  
 Neutral current interaction; nuclear cross section uncertainties

## Supernovae

Neutral current sensitivity to all neutrino flavor components  
 Sensitivity to both Galactic supernova burst (Horiwitz et al. 2003; Lang et al. 2016) and diffuse supernova neutrino background (DSNB)

# CEvNS with atmospheric neutrinos



# CEvNS community white-paper

Contributions/signatures welcome from all members of community

<https://www.overleaf.com/2859619433mgdvwgrxqddt>

## CEvNS whitepaper

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## 7 Experimental efforts

7.1	Stopped-pion beams	.....
7.1.1	SNS	.....
7.1.2	Lujan	.....
7.1.3	ESS	.....
7.1.4	JSNS	.....
7.2	Reactors	.....
7.2.1	CONNIE	.....
7.2.2	CONUS	.....
7.2.3	MINER	.....
7.2.4	NEON	.....
7.2.5	NUCLEUS	.....
7.2.6	RICOCHET	.....
7.2.7	RED-100	.....
7.2.8	NuGen	.....
7.2.9	TEXONO	.....
7.2.10	NEWSG	.....
7.3	Dark matter & CEvNS detectors	.....
7.3.1	XENON nT	.....
7.3.2	LZ	.....
7.3.3	DARWIN	.....
7.4	SuperNova Early Warning System (SNEWS)	.....
7.5	Directional detectors	.....
7.5.1	CYGNUS	.....
7.5.2	DRIFT	.....