

#### New approach to DM searches with mono-photon signature

J. Kalinowski<sup>\*</sup>, W. Kotlarski<sup>+</sup>, <u>P. Sopicki</u><sup>\*</sup>, K.Mękała, A. F. Żarnecki<sup>\*</sup> <sup>\*</sup>University of Warsaw <sup>+</sup> Technische Universität Dresden

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Previously on Snowmass...

#### https://indico.fnal.gov/event/43959/contributions/190182/

**Details: Eur. Phys. J. C (2020) 80: 634** arXiv:2004.14486

**Common approach:** 

- setting limits on the mediator mass for a given coupling structure/values

- This approach is suitable, if one can assume  $\mathbf{M}_{ ext{med}} \gg \sqrt{\mathbf{s}}$ 

Different approach needed for mediator masses of or below  $\sim \sqrt{s}$ 

#### Experimental-like approach:

- set limits on the DM production cross section as a function of the assumed **mediator mass and mediator width** 

- depends on the assumed **coupling structure**, but exact coupling values not relevant

- model independent upper limits on  $\sigma$  in this method can be then interpreted in different models

#### arXiv:1910.11775









• Experimental approach: consider 2-D distribution of mono-photon event

Event selection: single  $\gamma$  with 7°<  $\theta$  <173° and pT > 5 GeV, no other detector activity

• Main backgrounds @ 3TeV CLIC:





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• Main backgrounds  $e^+e^- \rightarrow vv + e^+e^- \rightarrow e^+e^-$  vs signal @ 3TeV CLICM<sub>DM</sub>=50GeV M<sub>med</sub>=1TeV  $\Gamma_{med}$ =50GeV







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Upper Limit on  $\sigma(e^+e^- \rightarrow \chi \chi + \gamma)$ as a function of a vector mediator mass for 3 TeV collisions with +80% polarisaton for electrons

- calculations with CLs method

- with smaller width of the mediator the limit gets stronger at the threshold



#### Conclusions



Applied to Dark Matter analysis one can get model independent limits on cross sections of various scenarios

- Wider range of mediator masses available
- Exact coupling values not needed
- Nice signal-background separation (i.e. RooStat test with hipothesis signal+bckg) and impact of the mediator width

Upper limits on cross section  $\sigma(e^+e^- \rightarrow \chi \chi + \gamma)$  are quite stable and flat even for lighter mediators

More to come:

- different coupling structures
- different collider energies and polarisation
- include detector effects using Delphes



# Backup





#### WIHZARD mono-photons: Bhabha

$\sqrt{s}[GeV]$	Whizard-2.8 $\sigma(e^+e^- \rightarrow e^+e^- + N\gamma)$ [fb]			$\sigma(e^+e^- \rightarrow e^+e^- + N\gamma)$ [fb]
	$e^+e^- + \gamma_{ME}$	$e^+e^- + 2\gamma_{ME}$	$e^+e^- + 3\gamma_{ME}$	after ISR rejection
240	236000	26100	1500	220000
250	224000	24900	1400	209000
380	140000	18400	1200	128000
500	100000	14600	1100	89800
1000	39800	7700	700	34400
1500	23000	5300	500	19400
3000	8800	2700	300	7200

$\sqrt{s}[GeV]$	$q_{min}[GeV]$	Whizard-2.8 $\sigma(e^+e^- \rightarrow e^+e^- + N\gamma)$ [fb]			$\sigma(e^+e^- \rightarrow e^+e^- + N\gamma))$ [fb]
		$e^+e^- + \gamma_{ME}$	$e^+e^-+2\gamma_{ME}$	$e^+e^-+3\gamma_{ME}$	after ISR rejection
	$q_{min}=0.1$	141000	29500	3200	123000
380	$q_{min}=0.5$	140000	21700	1500	126000
	$q_{min}=1$	140000	18400	1200	128000
	$q_{min}=5$	141000	10000	400	131000
	$q_{min}=10$	140000	5100	170	145000
	$q_{min}=50$	11200	160	2	11000
	$q_{min}=0.1$	8900	3700	420	6400
3000	$q_{min}=0.5$	9000	3000	230	7000
	$q_{min}=1$	8800	2700	300	7200
	$q_{min}=5$	8900	2000	140	7600
	$q_{min}=10$	8900	1300	90	7500
	$q_{min}=50$	7000	300	10	6100



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Which was a follow up: Eur. Phys. J. C (2020) 80: 634