

A final word on Minimal Dark Matter



We consider $SU(2)$ n -plets

We require STABILITY + CALCULABILITY

We are left with a very short list of cases:

n = 3 stable because of matter-parity

n = 5 stable accidentally

n = 7 stable if a millicharge is assigned

n = 9 already not calculable...

Key Question:

Can we probe all the Minimal Dark Matter WIMPS @ future colliders?

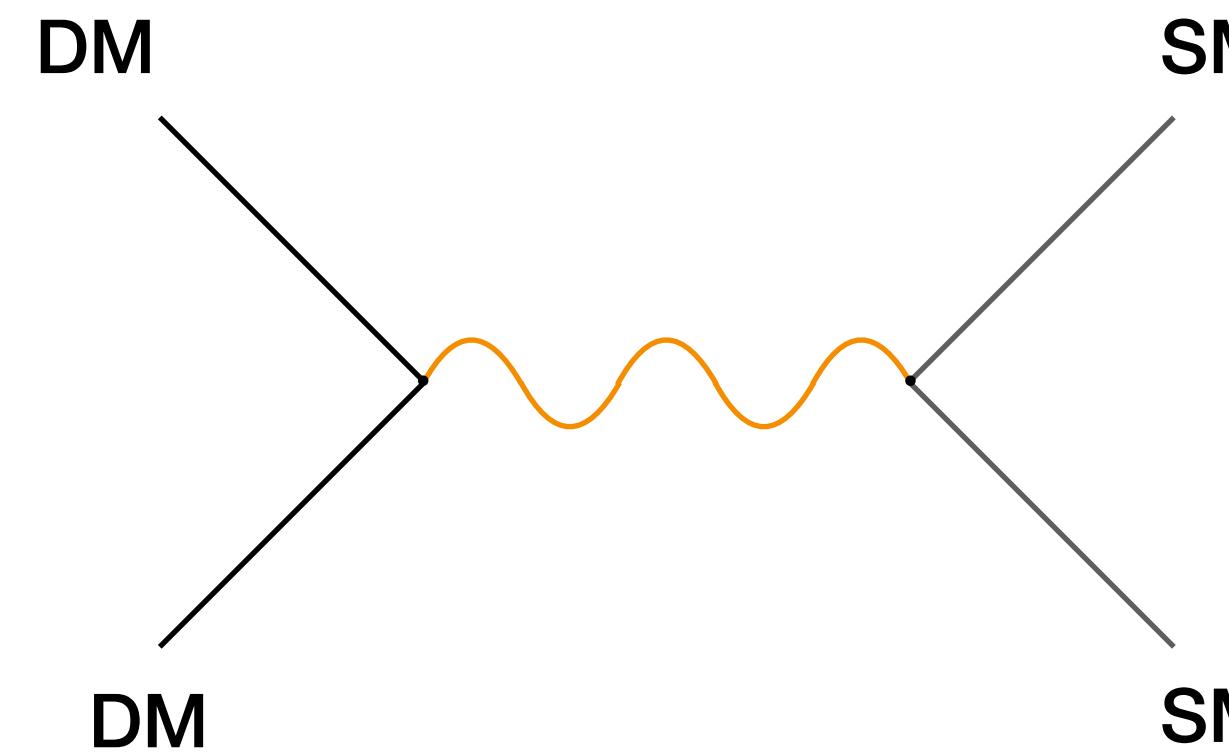


To do List:

- 1. Mass predictions from Freeze-out**
- 2. Reach at future colliders for different C.O.M energies**

Extra: comparison with future indirect detection prospects...

Freeze-out

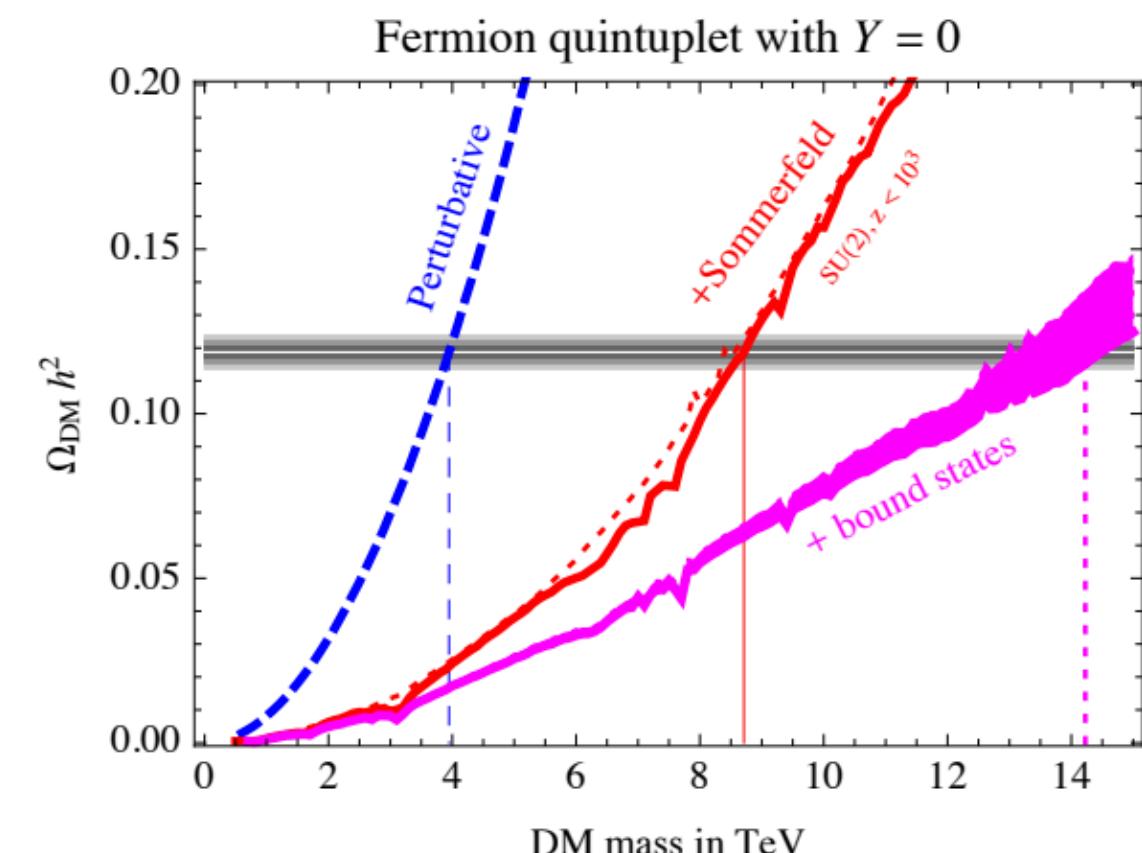
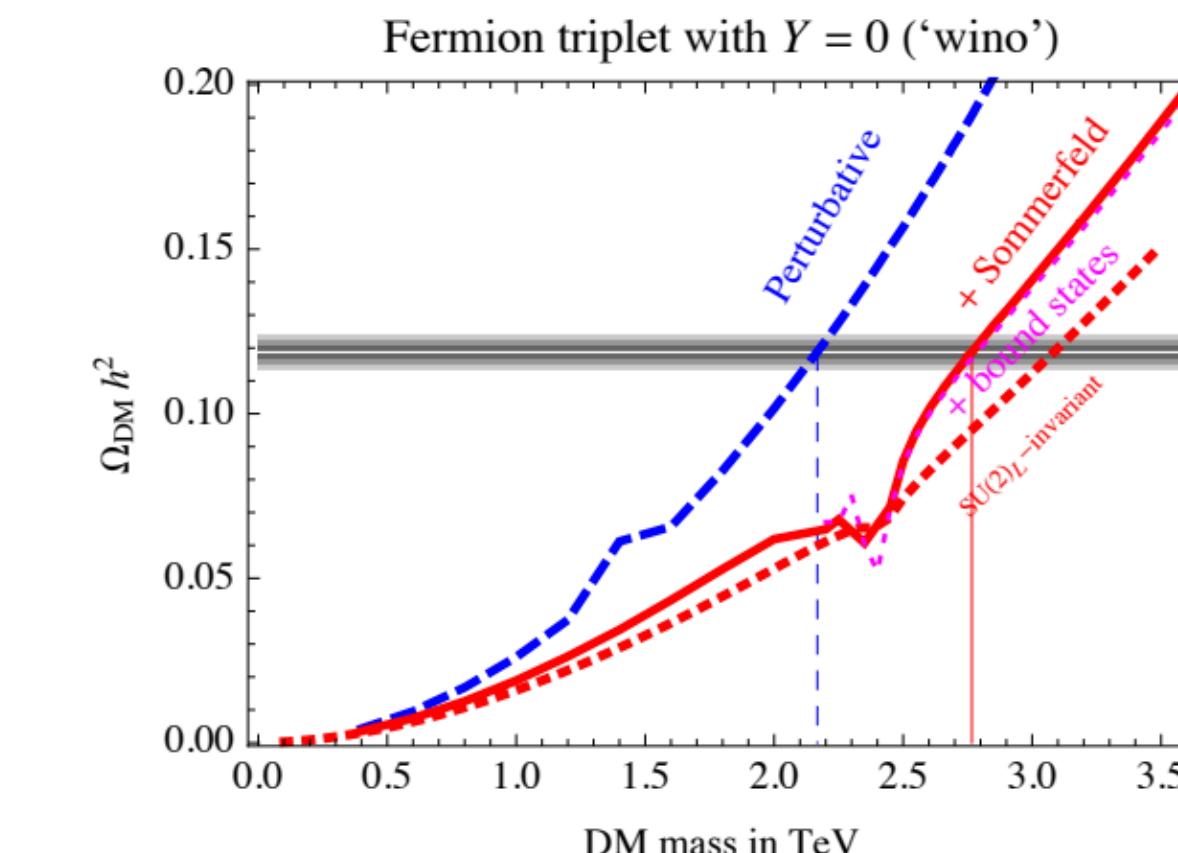


$$\sigma_{\text{ann}}(\text{DM} + \text{DM} \leftrightarrow \text{SM} + \text{SM}) \propto \frac{g^4}{M^2} n^3 \quad \xrightarrow{\sigma_{\text{ann}} = \sigma_{f.o.}} \quad M \propto n^{3/2}$$

Low-velocity corrections modify substantially σ_{ann}

- Sommerfeld enhancement
- Bound-state formation
(large corrections for $n \geq 5$)

taken from Mitridate, Redi, Smirnov, Strumia 2018



Summary of the WIMPs mass predictions

n = 3 (thermal Wino)

Perturbative: $M_3 = 2.2 \text{ TeV}$

w/ Sommerfeld: $M_3 = 2.8 \text{ TeV}$

No bound state formation!

n = 5

Perturbative: $M_5 = 4 \text{ TeV}$

w/ Sommerfeld: $M_5 = 9.4 \text{ TeV}$

w/ Bound-states: $M_5 = 14.4 \text{ TeV}$



n = 7

Perturbative: $M_7 = 9 \text{ TeV}$

Sommerfeld: $M_7 = 23 \text{ TeV}$

(in SU(2)-invariant limit – 1512.03332)

No calculation of bound-state effects yet!

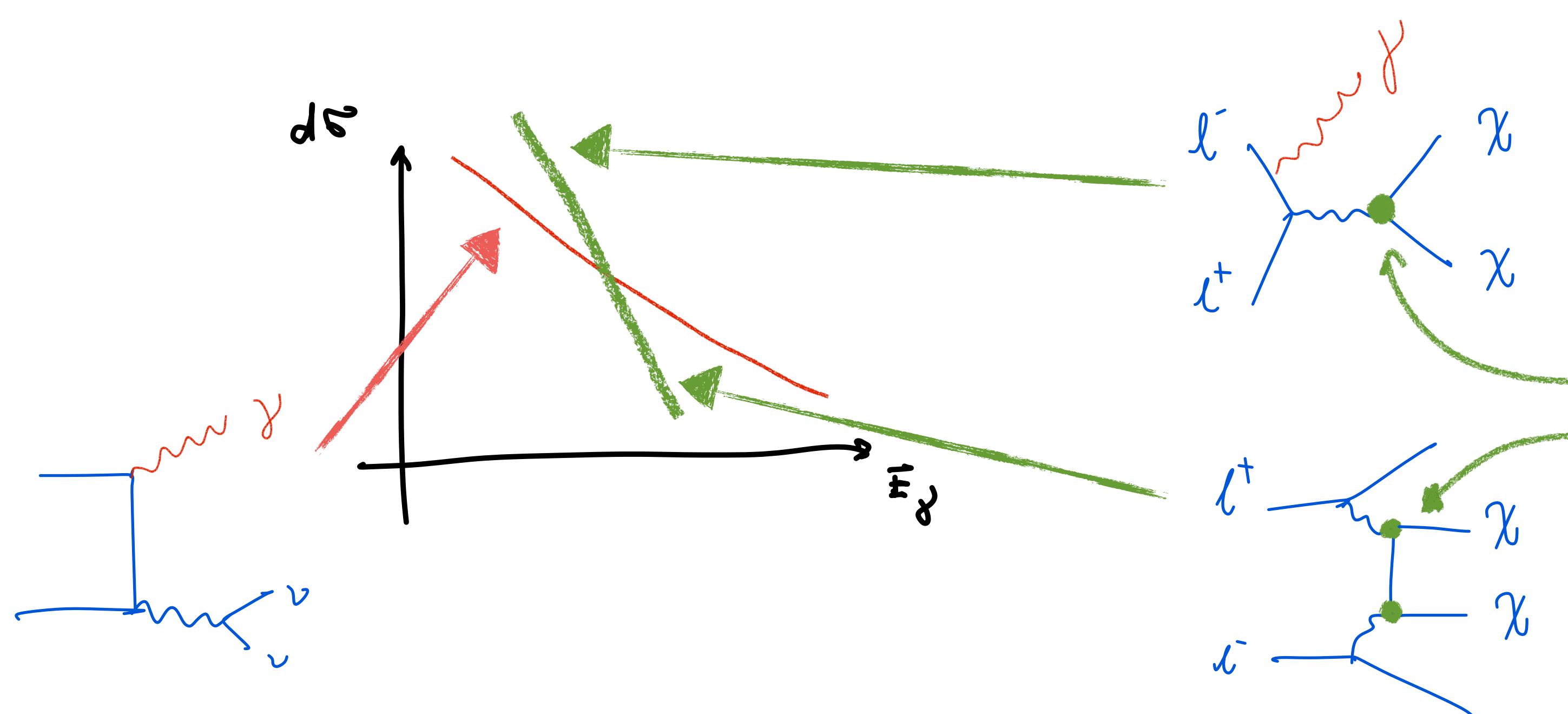


FUTURE COLLIDER REACH

Recoil on “nothing”

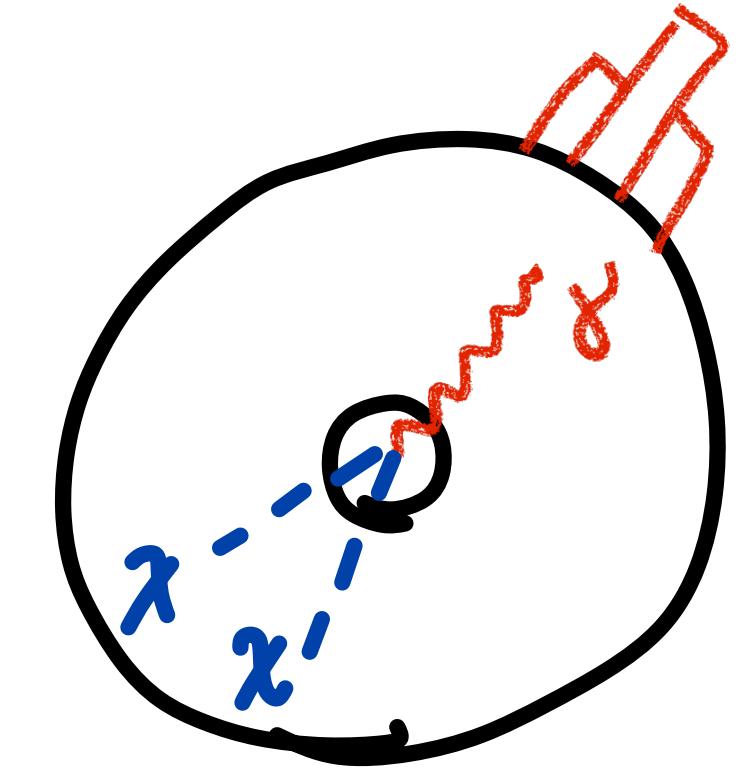
GENERIC

SEARCH INTERPRETED FOR DARK MATTER



+ “nothing”

Possible large enhancement
due to multiplicity of states
and size of couplings

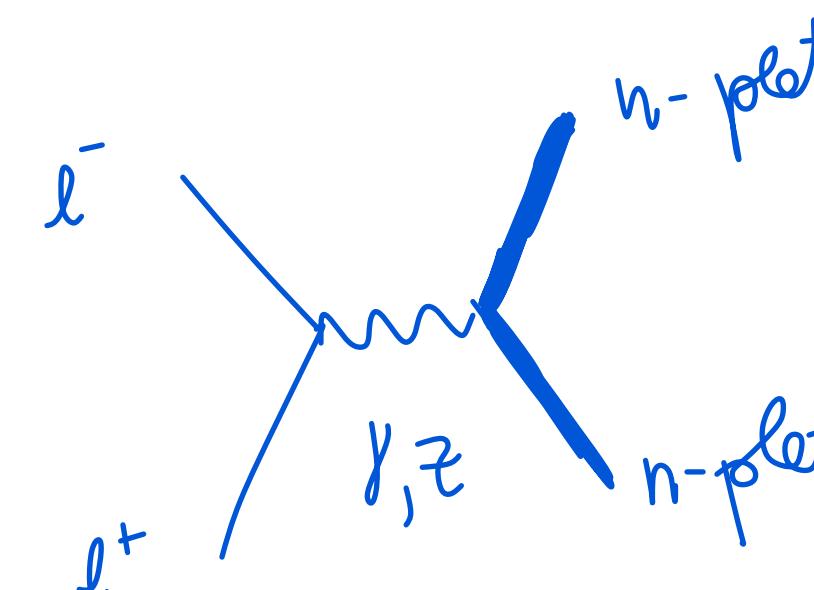


Degenerate EW multiplets

STUB-TRACKS

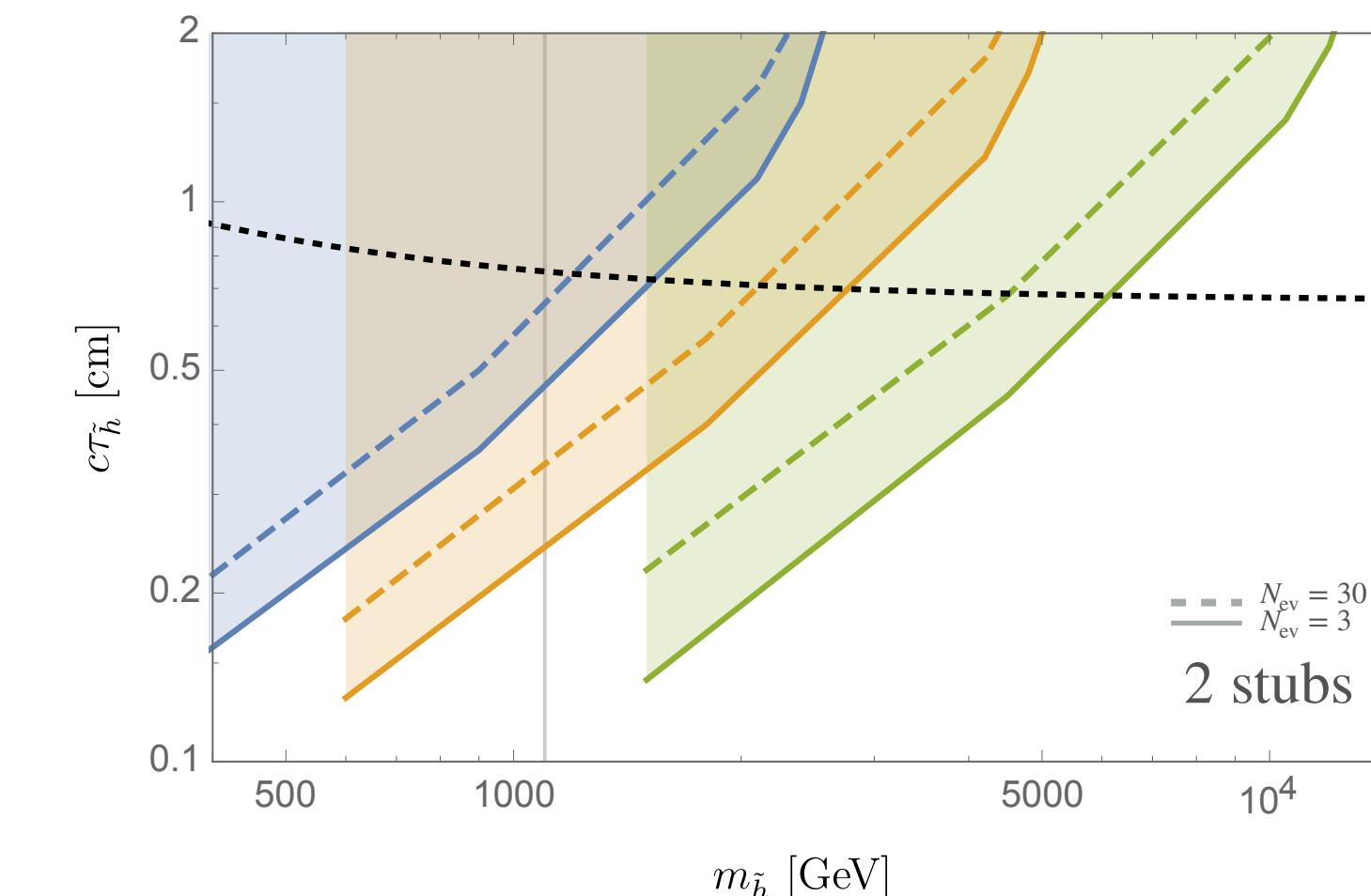
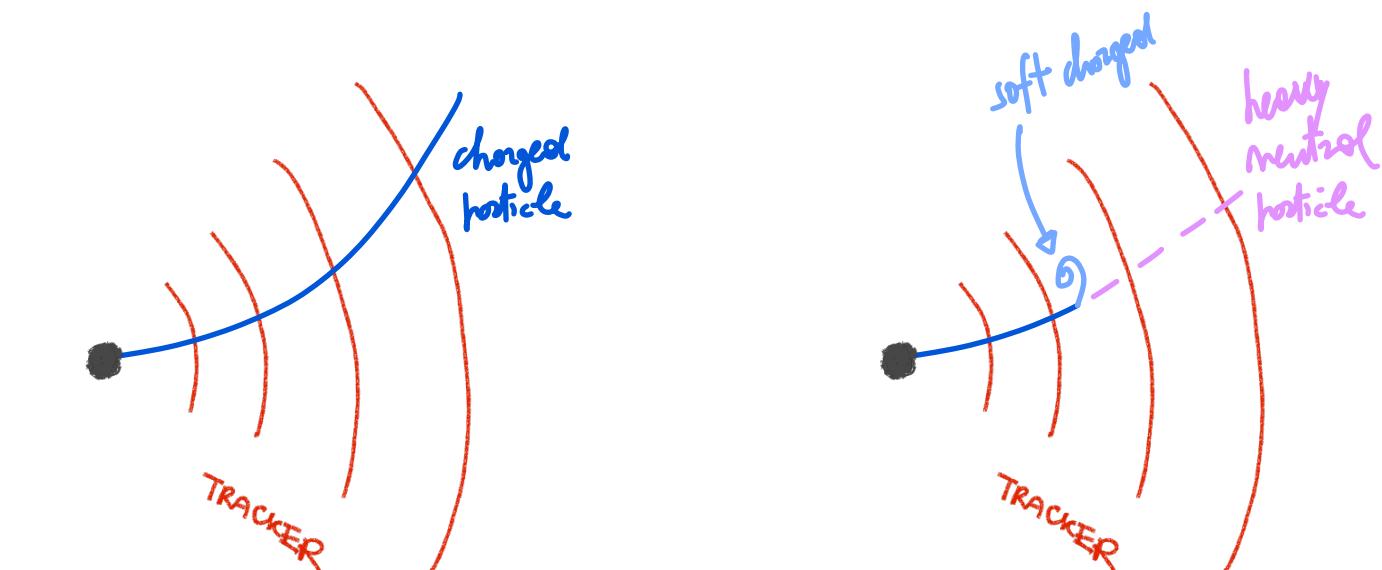
EXTRAPOLATION FROM CLIC

- Heavy n-plet of SU(2)
- Mass splitting $\sim a_w m_W \sim 0.1 \text{ GeV} - \text{GeV}$



LARGE RATES, BUT NEEDS TO LIGHT UP THE
DETECTOR IN A DISCERNIBLE WAY

- Heavily subject to detector design issues
- Even in CLIC needs full detector simulation



Assume track is seen when
 $c\tau > \frac{4.4\text{cm}}{\sin\theta}$ for $\theta \in [19,90]\text{deg}$

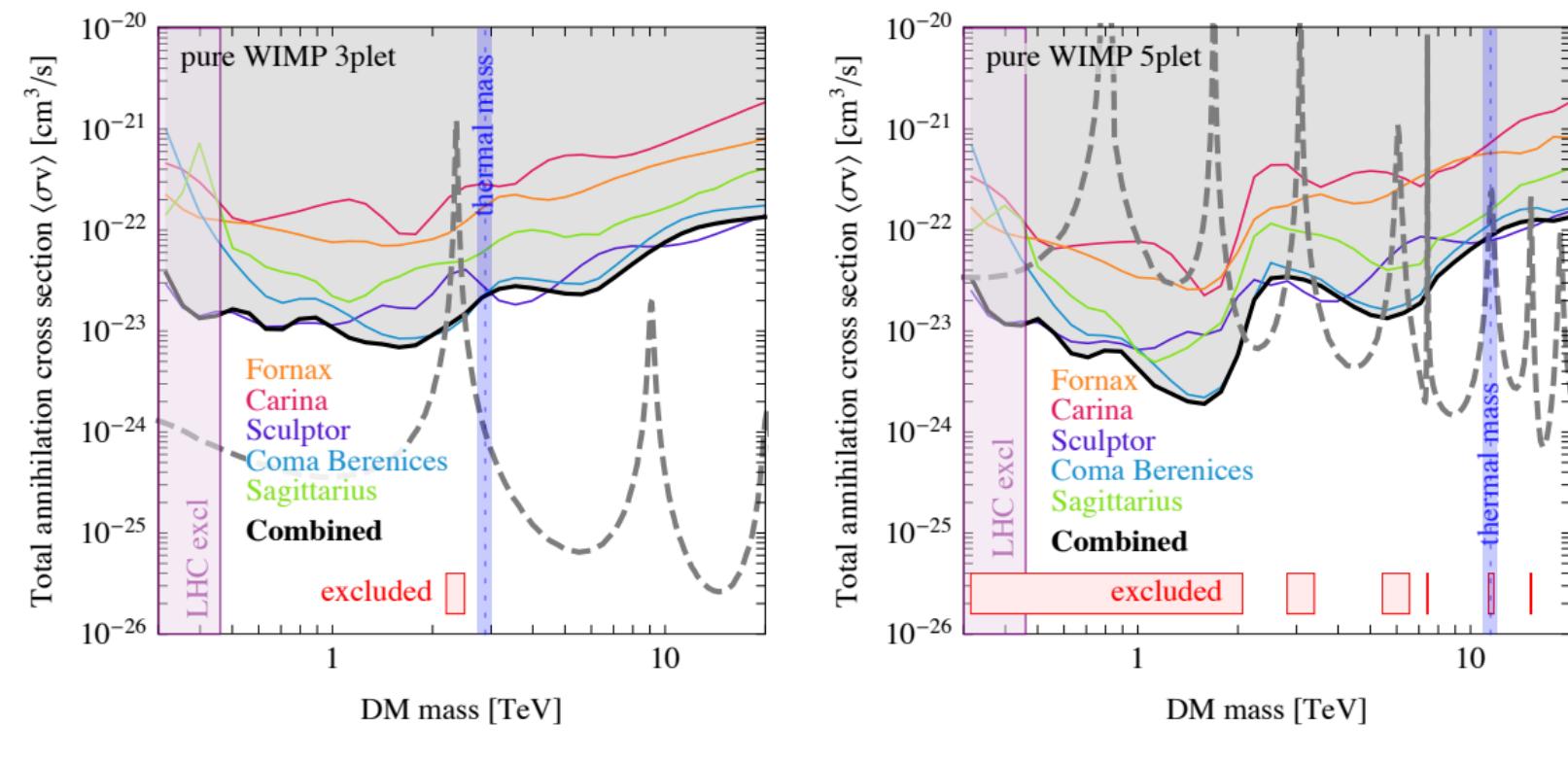
$\sqrt{s} = 6 \text{ TeV } 6/\text{ab}$

$\sqrt{s} = 12 \text{ TeV } 12/\text{ab}$

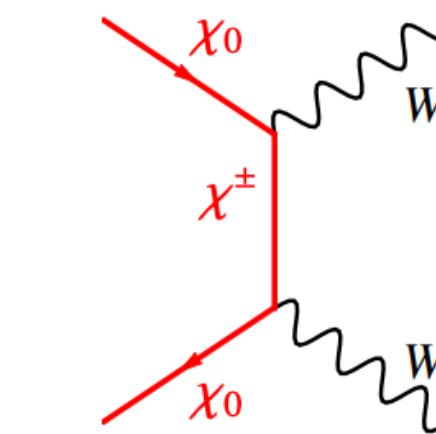
$\sqrt{s} = 30 \text{ TeV } 30/\text{ab}$

Higgsino Lifetime

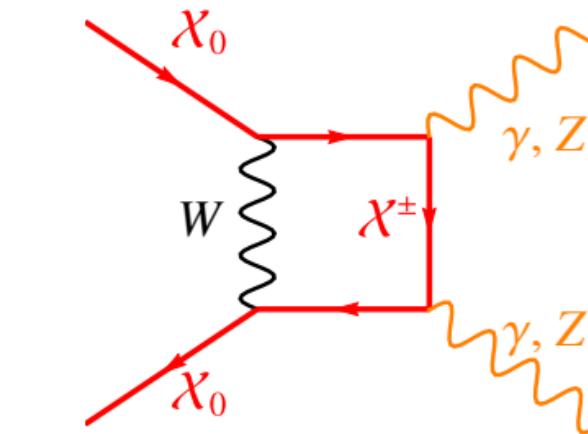
Indirect Detection



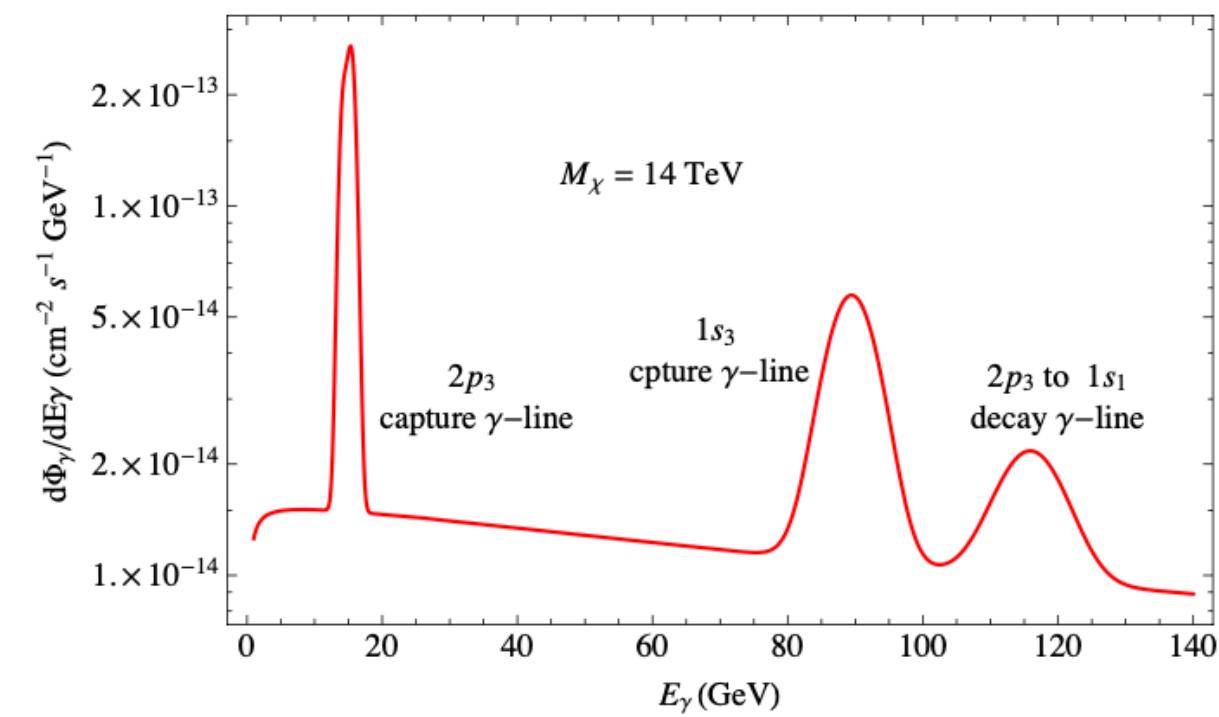
High energy gamma rays in dwarf Spheroidal galaxies



tree-level continuum



loop-induced lines



Mitridate, Redi, Smirnov, Strumia 2018

$$\chi_0 \chi_0 \rightarrow B\gamma$$

extra gamma ray lines from bound state formation!