

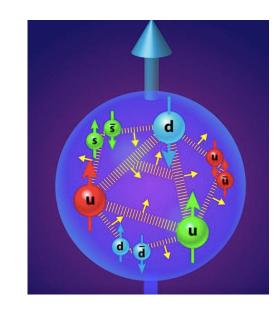
GPDs and Deeply Virtual Exclusive Experiments

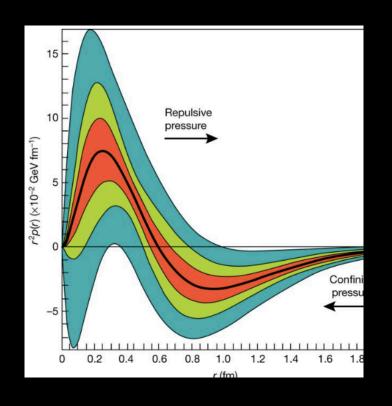
A new paradigm that will allow us to both penetrate and visualize the deep structure of visible matter, answering questions that we couldn't even afford asking before

$$\frac{1}{2} \int_{-1}^{1} dx x \left[H_q(x, 0, 0) + E_q(x, 0, 0) \right] = J_q$$

X. Ji, 1997

How does the proton/neutron get its mass and spin and how do we test this dynamics?







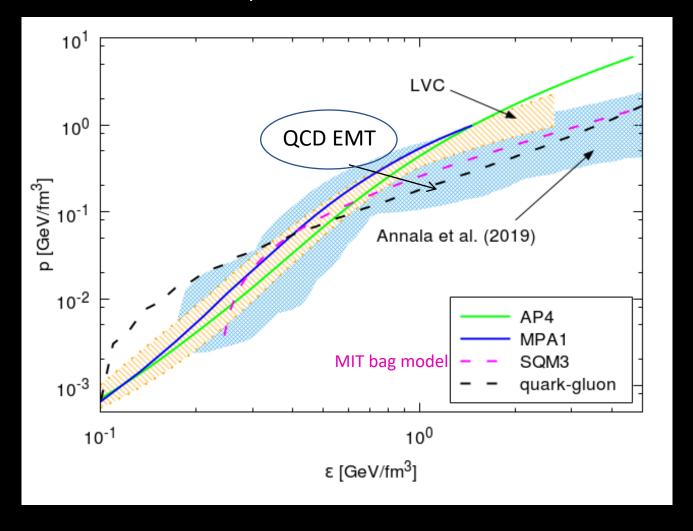
"The average peak pressure near the center is about 10³⁵ pascals, which exceeds the pressure estimated for the most densely packed known objects in the Universe, neutron stars"

Burkert, Elouadrhiri, Girod, Nature 557, 396 (2018)

How do we measure pressure and forces inside the proton?

M. Polyakov, 2003

The Jlab extraction of the pressure measurement helps us understand the EoS of neutron stars



A. Rajan, T. Gorda, SL, K. Yagi, submitted for publication

$$\langle p', \Lambda \mid T^{\mu\nu} \mid p, \Lambda \rangle = \underline{A(t)} \overline{U}(p', \Lambda') [\gamma^{\mu} P^{\nu} + \gamma^{\nu} P^{\mu}] U(p, \Lambda) + \underline{B(t)} \overline{U}(p', \Lambda') i \frac{\sigma^{\mu(\nu} \Delta^{\nu})}{2M} U(p, \Lambda)$$

$$+ \underline{C(t)} [\Delta^{2} g^{\mu\nu} - \Delta^{\mu\nu}] \overline{U}(p', \Lambda') U(p, \Lambda) + \underline{\widetilde{C}(t)} g^{\mu\nu} \overline{U}(p', \Lambda') U(p, \Lambda)$$

off-forward

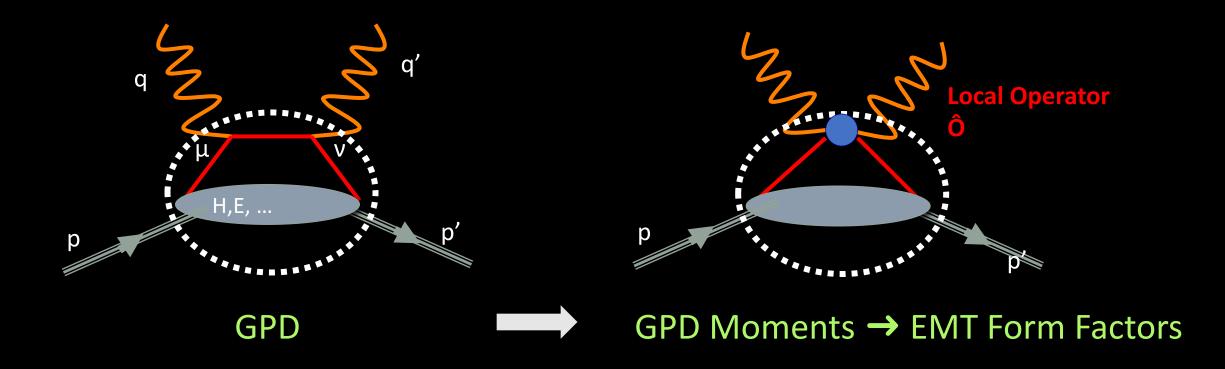
q and g not separately conserved

$$P = \frac{p+p'}{2}$$

$$\Delta = p'-p = q - \underline{q'}$$

$$t = (p-p')^2 = \Delta^2$$

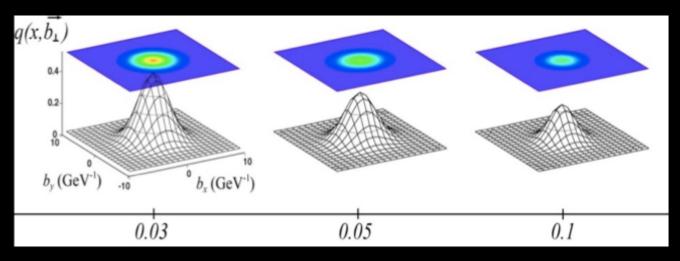
EMT matrix elements from Generalized Parton Distributions Moments (X.Ji,1997)



- Large momentum transfer Q²>>M² → "deep"
- Large Invariant Mass W²>>M²

 equivalent to an "inelastic" process

Measuring the Nucleon Gravitomagnetic Form Factors



graph from M. Defurne

A multi-step, multi-prong process that compares to imaging a black hole

M87*

Event Horizon Telescope (EHT)

- ✓ Main idea: Very Long Baseline Interferometry (VLBI), an array of smaller telescopes synchronized to focus on the same object and act as a giant telescope
- ✓ Precision: large aperture (many telescopes widely spaced) and high frequency radio waves
- ✓ Data Management: 5 petabytes physically transported to a central location. Data from all eight sites were combined to create a composite set of images, revealing for the first time M87*'s event horizon.

It took nearly two decades to achieve!

Jefferson Lab@12 GeV

- Main idea: use DVCS, TCS, DVMP... and related processes as probes
- Precision: high luminosity in a wide kinematic range is key!
- ✓ Data Management: unprecedented amount of data need new AI based techniques to handle the image making

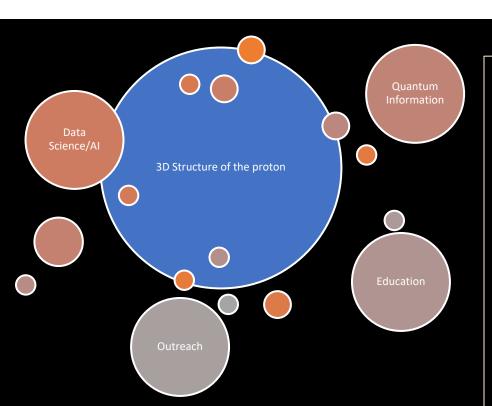
In the course of 10 years, first proton image.

Topic	Hall A	Hall B	Hall C	Hall D	Other	Total
Hadron spectra as probes of QCD	0	2	1	3	0	6
Transverse structure of the hadrons	6	3	3	1	0	13
Longitudinal structure of the hadrons	2	3	7	0	0	12
3D structure of the hadrons	5	9	6	0	0	20
Hadrons and cold nuclear matter	8	5	7	0	1	21
Low-energy tests of the Standard Model and Fundamental Symmetries	3	1	0	1	2	7
Total	24	23	24	5	3	79



Center for Femtography (CNF)

SURA

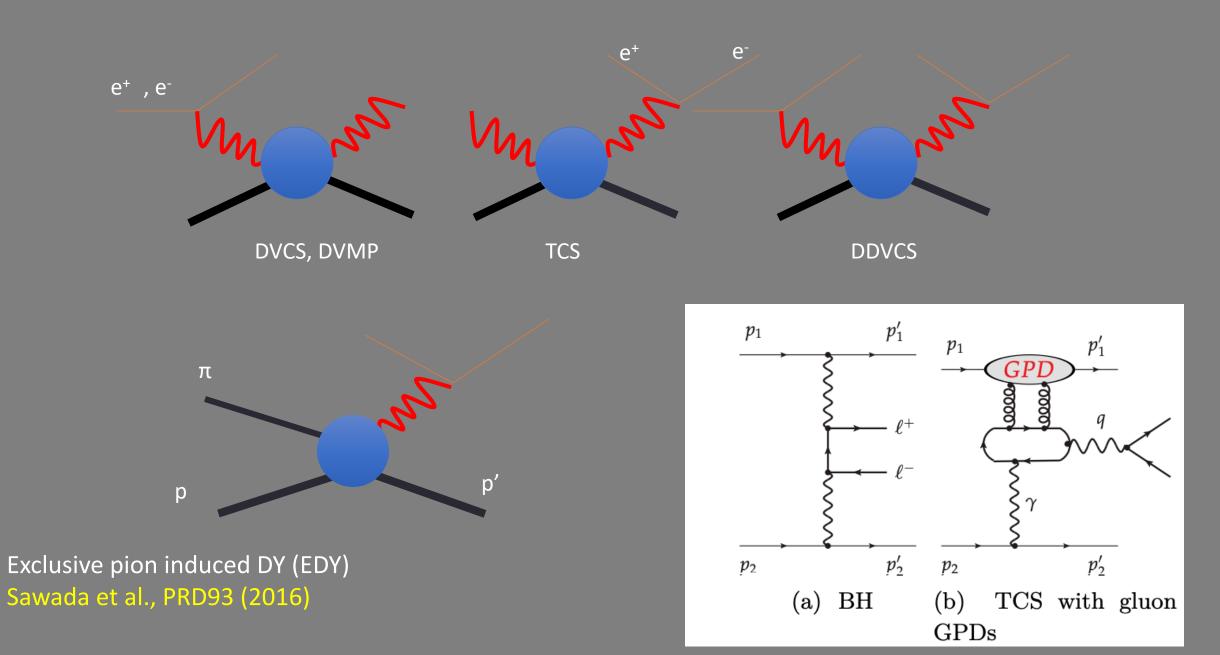


CNF has funded multiple projects on Femtography

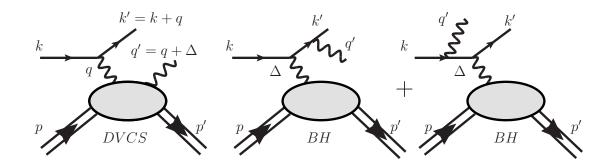
efferson Science Associates, LLC

- Covering different areas
 - Experimental data
 - ML & Al
 - Inverse problems
 - Lattice calculations
 - ...
- Total funding exceeding \$0.5M per year
- Presently moving into a larger collaboration, involving people outside VA.

Harnessing/coordinating information from all channels



We need a robust framework for the cross section, where kinematic limits are under control (beyond "harmonics" model)

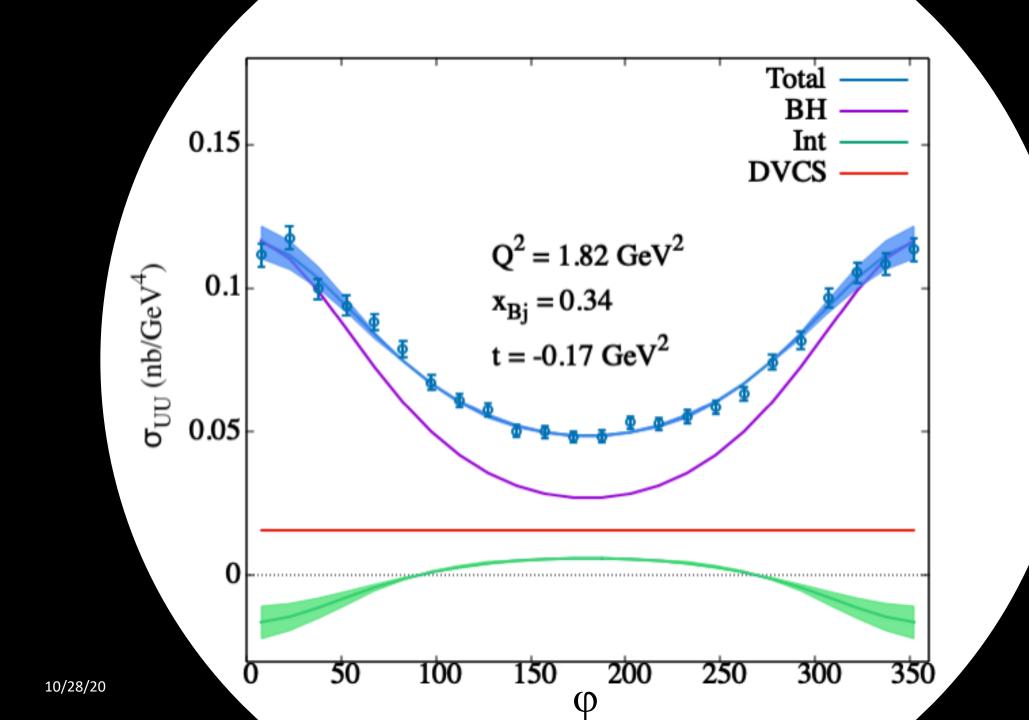


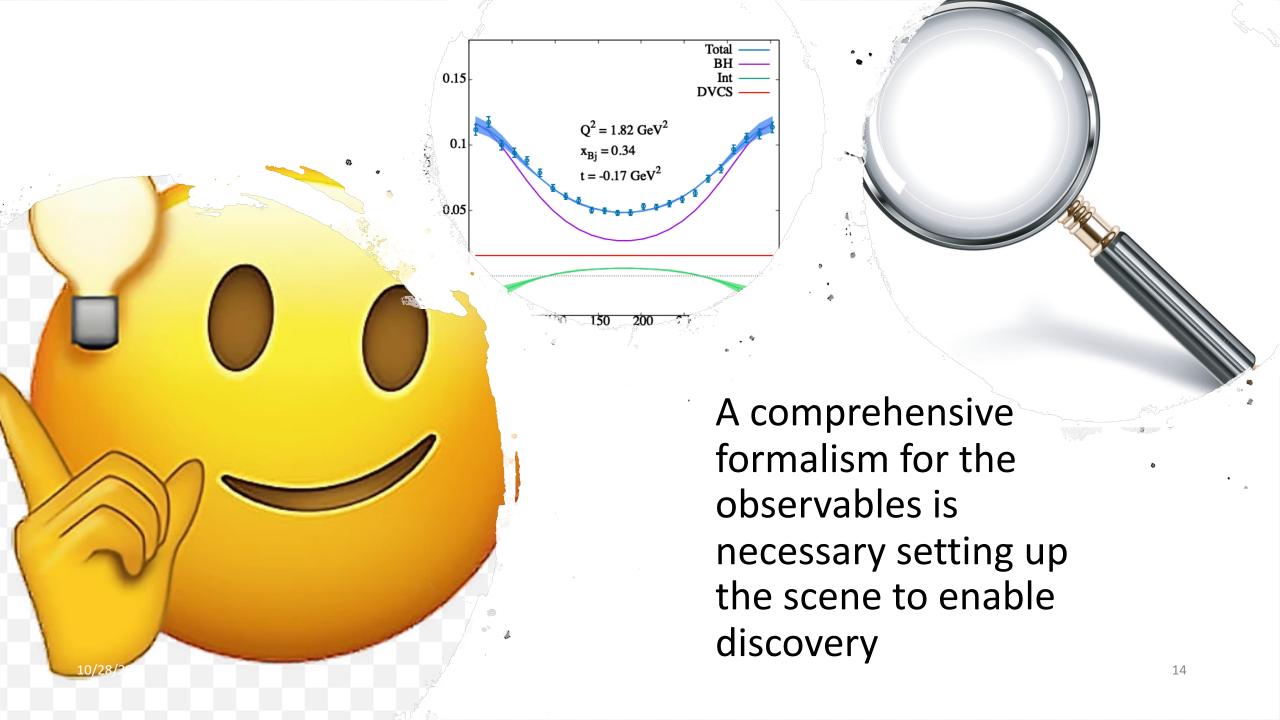
- **B.** Kriesten et al, *Phys.Rev. D* 101 (2020)
- **B. Kriesten and S. Liuti,** arXiv <u>2004.08890</u>

$$\frac{d^5\sigma}{dx_{Bj}dQ^2d|t|d\phi d\phi_S} = \frac{\alpha^3}{16\pi^2(s-M^2)^2\sqrt{1+\gamma^2}} |T|^2$$

$$T(k, p, k', q', p') = T_{DVCS}(k, p, k', q', p') + T_{BH}(k, p, k', q', p'),$$

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DVCS

$$\frac{d^5\sigma_{DVCS}}{dx_{Bj}dQ^2d|t|d\phi d\phi_S} = \\ \text{twist two GPDs} \\ = \\ \frac{\Gamma}{Q^2(1-\epsilon)} \left\{ \begin{bmatrix} F_{UU,T} + F_{UU,L} + \epsilon \cos 2\phi F_{UU}^{\cos 2\phi} \\ F_{UU} + F_{UU,L} + \epsilon \cos 2\phi F_{UU}^{\cos 2\phi} \\ F_{UU} + F_{UU,L} + \epsilon \cos 2\phi F_{UU}^{\cos 2\phi} \\ F_{UU} + F_{UU,L} + \epsilon \cos 2\phi F_{UU}^{\cos 2\phi} \\ F_{UU} + F_{UU,L} + F_$$

Twist 3 GPDs Physical Interpretation

GPD	$P_q P_p$	TMD	Ref. 1
H^{\perp}	UU	f^{\perp}	$2\widetilde{H}_{2T} + E_{2T}$
\widetilde{H}_L^\perp	LL	g_L^\perp	$2\widetilde{H}_{2T}' + E_{2T}'$
H_L^\perp	UL	$f_L^{\perp (*)}$	$\widetilde{E}_{2T} - \xi E_{2T}$
\widetilde{H}^{\perp}	LU	$g^{\perp (*)}$	$\widetilde{E}'_{2T} - \xi E'_{2T}$
$H_T^{(3)}$	UT	$f_T^{(*)}$	$H_{2T}+ au\widetilde{H}_{2T}$
$\widetilde{H}_{T}^{(3)}$	LT	g_T'	$H'_{2T} + au \widetilde{H}'_{2T}$



 \longrightarrow 1/Q correction to \widetilde{H}

NEW!! Orbital Angular Momentum L

NEW!! Spin Orbit correlation L · S

1/Q correction to E

 \longrightarrow 1/Q correction to \widetilde{E}

(*) T-odd

[1] Meissner, Metz and Schlegel, JHEP(2009)

BH

$$\frac{d^5 \sigma_{unpol}^{BH}}{dx_{Bj} dQ^2 d|t| d\phi d\phi_S} \equiv \frac{\Gamma}{t} F_{UU}^{BH} = \frac{\Gamma}{t} \left[A(y, x_{Bj}, t, Q^2, \phi) \left(F_1^2 + \tau F_2^2 \right) + B(y, x_{Bj}, t, Q^2, \phi) \tau G_M^2(t) \right]$$

$$A = \frac{16 M^2}{t(k q')(k' q')} \left[4\tau \left((k P)^2 + (k' P)^2 \right) - (\tau + 1) \left((k \Delta)^2 + (k' \Delta)^2 \right) \right]$$

$$B = \frac{32 M^2}{t(k q')(k' q')} \left[(k \Delta)^2 + (k' \Delta)^2 \right],$$

$$\epsilon_{BH} = \left(1 + \frac{B}{A}(1+\tau)\right)^{-1}$$

...compared to ELASTIC SCATTERING

$$\left(\frac{d\sigma}{d\Omega}\right)_0 = \left(\frac{d\sigma}{d\Omega}\right)_{\text{Mott}} \frac{\epsilon (G_E^N)^2 + \tau (G_M^N)^2}{\epsilon (1+\tau)},\,$$

where N = p for a proton and N = n for a neutron, (the recoil-corrected relativistic point-particle (Mott) and τ , ϵ are dimensionless kinematic variables:

$$\tau = \frac{Q^2}{4m_N^2}, \quad \epsilon = \left[1 + 2(1+\tau)\tan^2\frac{\theta}{2}\right]^{-1},$$

J. Arrington, G. Cates, S. Riordan, Z. Ye, B. Wojsetowski, A. Puckett ...

$$F_{UU}^{\mathcal{I}} = F_{UU}^{\mathcal{I},tw2} + \frac{K}{\sqrt{Q^2}} F_{UU}^{\mathcal{I},tw3}$$

BH-DVCS interference

$$F_{UU}^{\mathcal{I},tw2} = A_{UU}^{\mathcal{I}} \Re e \left(F_1 \mathcal{H} + \tau F_2 \mathcal{E} \right) + B_{UU}^{\mathcal{I}} G_M \Re e \left(\mathcal{H} + \mathcal{E} \right) + C_{UU}^{\mathcal{I}} G_M \Re e \widetilde{\mathcal{H}}$$

 A_{UU}^{I} B_{UU}^{I} C_{UU}^{I}

are φ dependent coefficients

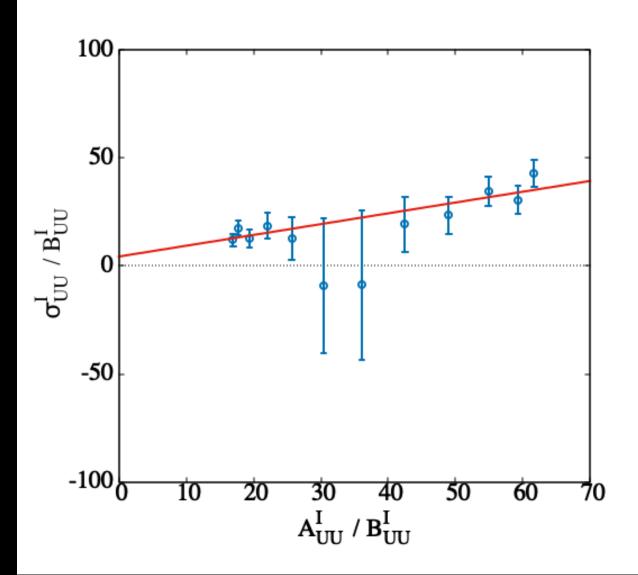
Twist 3 BH-DVCS interference

$$F_{UU}^{\mathcal{I}} = F_{UU}^{\mathcal{I},tw2} + \frac{K}{\sqrt{Q^2}} F_{UU}^{\mathcal{I},tw3}$$

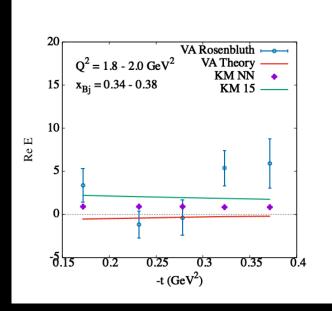
$$\begin{split} F_{UU}^{\mathcal{I},tw3} &= A_{UU}^{(3)\mathcal{I}} \Big[F_1 \Big(\Re e(2\widetilde{\mathcal{H}}_{2T} + \mathcal{E}_{2T}) - \Re e(2\widetilde{\mathcal{H}}_{2T}' + \mathcal{E}_{2T}') \Big) + F_2 \Big(\Re e(\mathcal{H}_{2T} + \tau \widetilde{\mathcal{H}}_{2T}) - \Re e(\mathcal{H}_{2T}' + \tau \widetilde{\mathcal{H}}_{2T}') \Big) \Big] \\ &+ \frac{B_{UU}^{(3)\mathcal{I}} G_M \left(\Re e \widetilde{\mathcal{E}}_{2T} - \Re e \widetilde{\mathcal{E}}_{2T}' \right)}{C} & \text{Orbital Angular Momentum} \\ &+ C_{UU}^{(3)\mathcal{I}} G_M \left[2\xi (\Re e \mathcal{H}_{2T} - \Re e \mathcal{H}_{2T}') - \tau \Big(\Re e(\widetilde{\mathcal{E}}_{2T} - \xi \mathcal{E}_{2T}) - \Re e\left(\widetilde{\mathcal{E}}_{2T}' - \xi \mathcal{E}_{2T}'\right) \Big) \Big] \end{split}$$

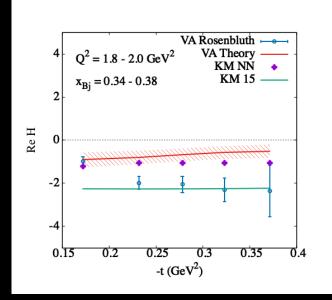
Need different dedicated study of observables!!

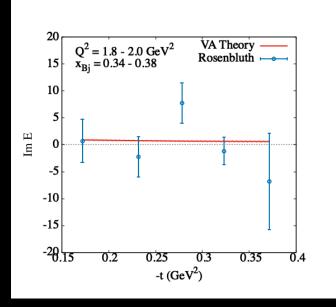
Rosenbluth Separated BH-DVCS interference data

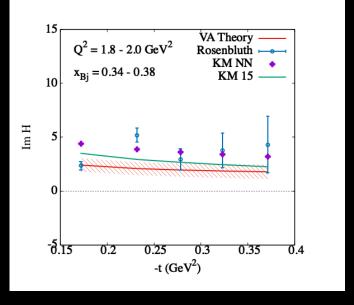


Compton Form Factors



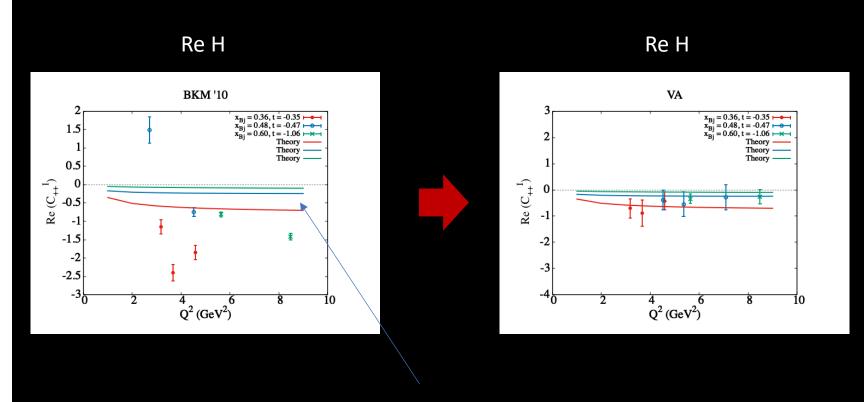






$\phi = 180$ $\varphi = 0$ 1000 VA VΑ BKM '01 BKM '01 500 150 BKM '10 BKM '10 100 -500 ₿-1000 -1500 -50 -2000 -100 -2500 Q^2 (GeV²) Q2 (GeV2) $\phi = 0$ $\phi = 180$ 120 VΑ BKM '01 BKM '01 100 BKM '10 BKM '10 10 \mathbf{B}_{UU} 20 Q^2 (GeV²) Q2 (GeV2) $\phi = 180$ $\varphi = 0$ 120 VA VΑ 40 BKM '01 BKM '01 100 BKM '10 BKM '10 -20 CE 40 20 -80 -100 -20 Q2 (GeV2) Q^2 (GeV²)

Impact on Q² dependence: Brandon Kriesten's talk



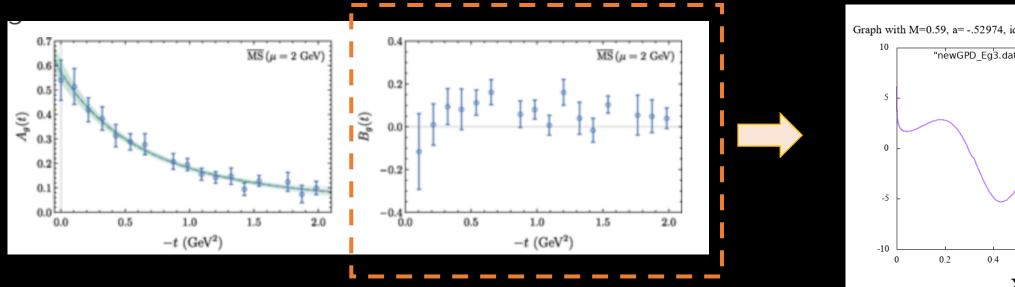
pQCD Evolution

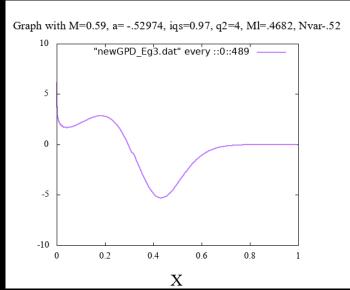
Moving Towards EIC, LHCSpin

Gluon GPDs
$$\int_{g}^{\infty} J_{q} + \sum_{q} J_{q} = \frac{1}{2}$$
 H_{g}, E_{g}

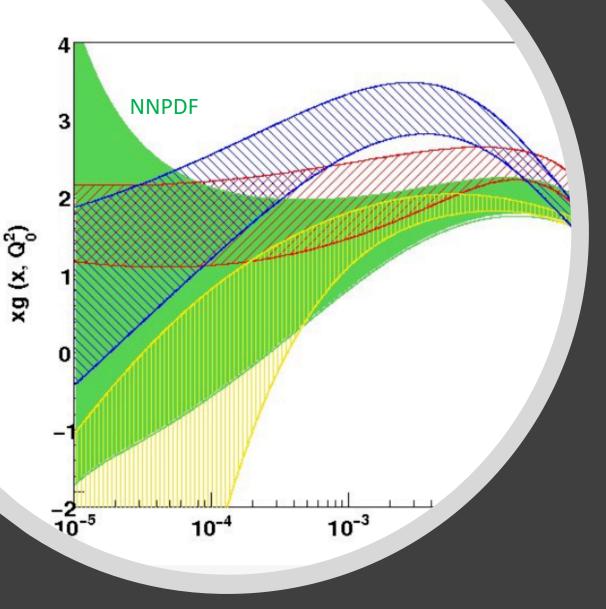
Accessed through PQCD evolution and J/ψ production

Parametrizations/models merging information from lattice QCD





UVA "state of the art results" (B. Kriesten, P. Velie. E. Yeats, F. Yepez)



Extracting information from data with standard methods is painstakingly slow: can we use ML/AI?

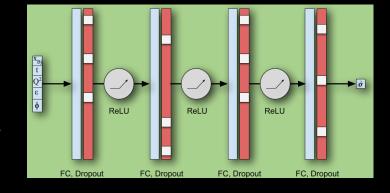
Brandon Kriesten (UVA), Jake Grigsby (UVA), Pete Alonzi (DS, UVA), Matthias Burkardt (NMSU), Joshua Hoskins (Mary Washington U.), SL

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UVA Strategy: supervised learning algorithm augmented by unsupervised (UMAP) method

Using Deep Neural Networks for regression

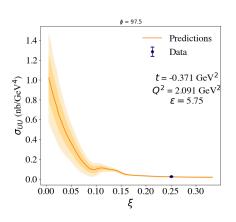
- The model is a multilayered perceptron (MLP) with 4 hidden layers, each followed by a rectified inear unit (ReLU) activation function
- Once the model is trained, we can use it to make predictions about the value of the cross section in a given kinematic region. The cross section at a point (x_s, t, Q₂, E, φ) is computed by a forward pass of the network. This operation is parallelizable and can run efficiently on a GPU

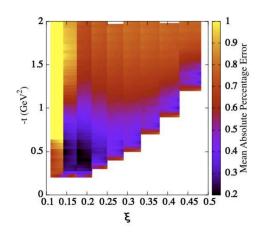


 We estimate the model's uncertainty by performing many forward passes per prediction point, with Dropout turned on.

Role of Noise and Uncertainty

- We are not restricted to kinematic bins that have already been measured the model learns to extrapolate between existing bins.
- We can sweep through regions where data do not exist, and use the model's uncertainty estimate to advise our use of its predictions
- This creates a distribution of predictions, and we use the mean and standard deviation of that distribution to estimate uncertainty
- (Gal & Ghahramani, 2016)





	UU			LU				
	Standa	ırd	Feature Engineering		Standard		Feature Engineering	
Method	Median Error (%)	Accuracy (%)	Median Error (%)	Accuracy (%)	Median Error (%)	Accuracy (%)	Median Error (%)	Accuracy (%)
Linear	238.87	1.61	347.77	0.53	239.37	22.44	289.60	24.02
SVR	57.70	15.86	58.92	13.17	94.92	43.70	68.55	59.06
Random Forest	32.90	24.73	32.27	26.88	62.87	59.45	63.65	59.06
KNN	19.84	37.63	20.93	34.41	62.36	60.24	68.09	56.69
DNN	19.11	61.3	21.56	60.5	145.86	62.0	150.62	65.8

This model will be open-sourced

What I left out

• Nuclei: ⁴He and deuteron GPDs

• π^0 electroproduction as a means to access the tensor charge and transversity GPDs

CONCLUSIONS

Immense discovery potential as we uncover the mechanical properties the of the proton and observe its spatial images through deeply virtual exclusive experiments

To observe, evaluate and interpret GPDs and Wigner distributions at the subatomic level requires stepping up data analyses from the standard methods



developing new numerical/analytic/quantum computing methods



Center for Nuclear Femtography