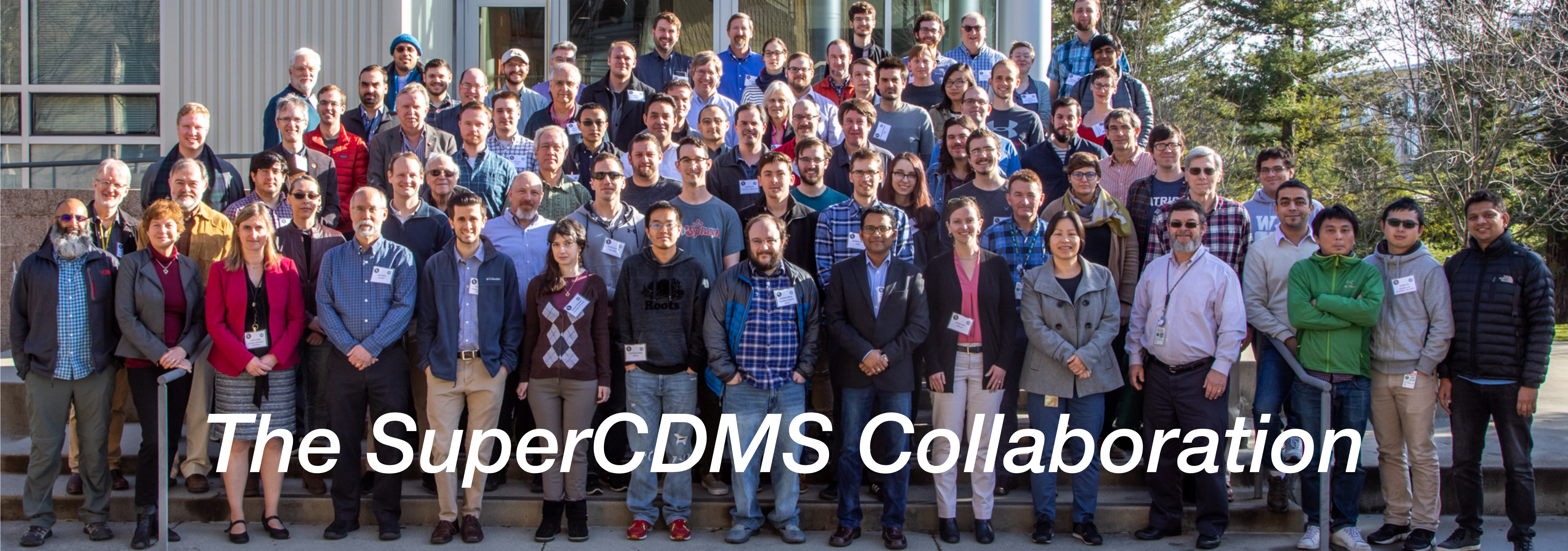


Future Electron Recoil Dark Matter Searches with SuperCDMS Cryogenic Detectors



T. Saab for
SuperCDMS Collaboration

Snowmass 2021 CF1 Meeting
Sept 25 2020



The SuperCDMS Collaboration



 @SuperCDMS

T. Saab \ SNOWMASS 2021 CF1 \ Sep 25, 2020

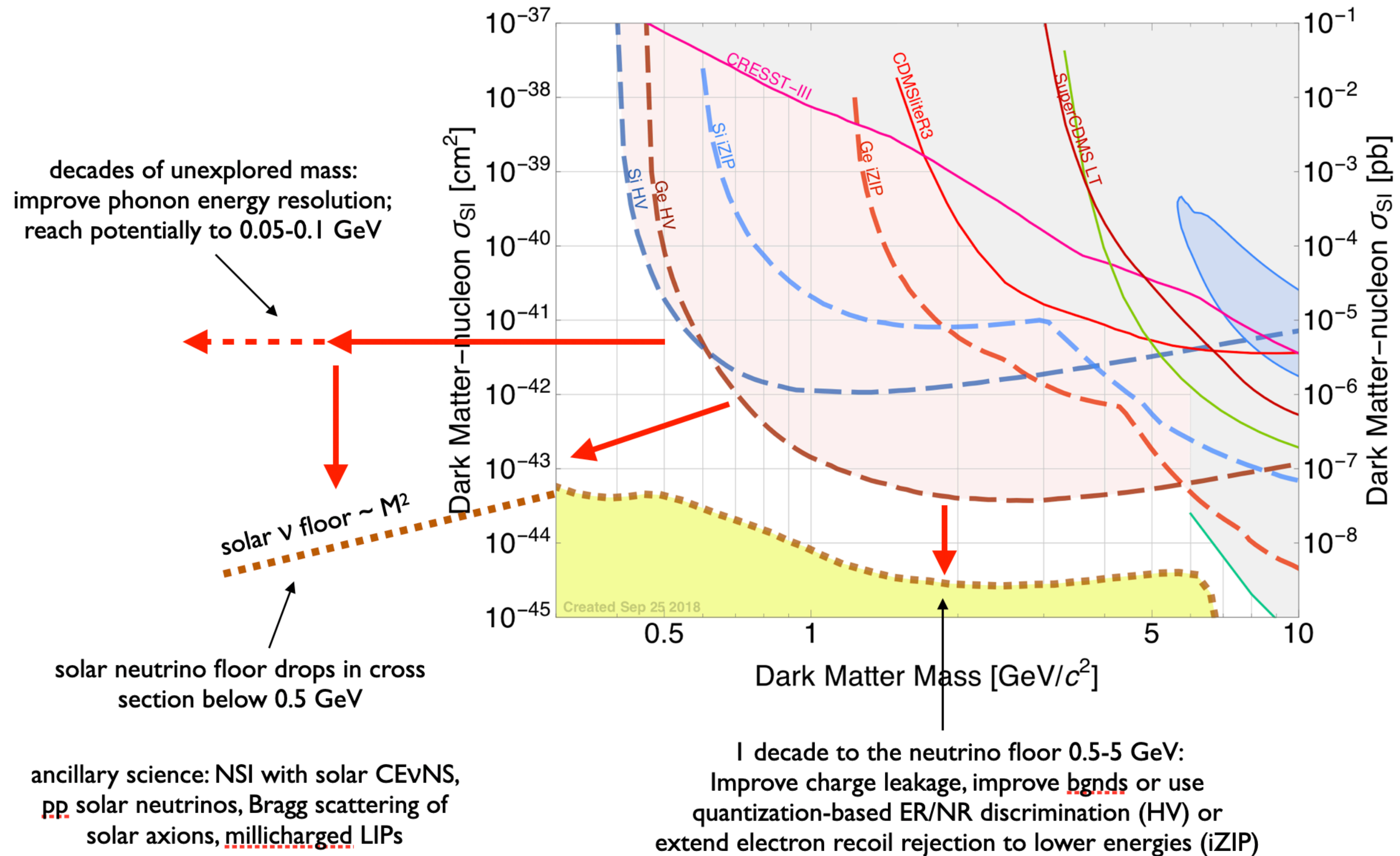
supercdms.slac.stanford.edu

- Although the SuperCDMS SNOLAB G2 Experiment was originally conceived to search for GeV scale dark matter:
 - It has the ability to probe significant regions of sub-GeV dark matter parameter space.
 - Fits in well with the current dark matter landscape which has evolved substantially since the G2 projects began.
- See S. Golwala's [Aug. 21 CF1 presentation](#) for detail re. the SuperCDMS SNOLAB program and facility.

Potential Science Goals of Upgrade Program: Nuclear-Scattering DM



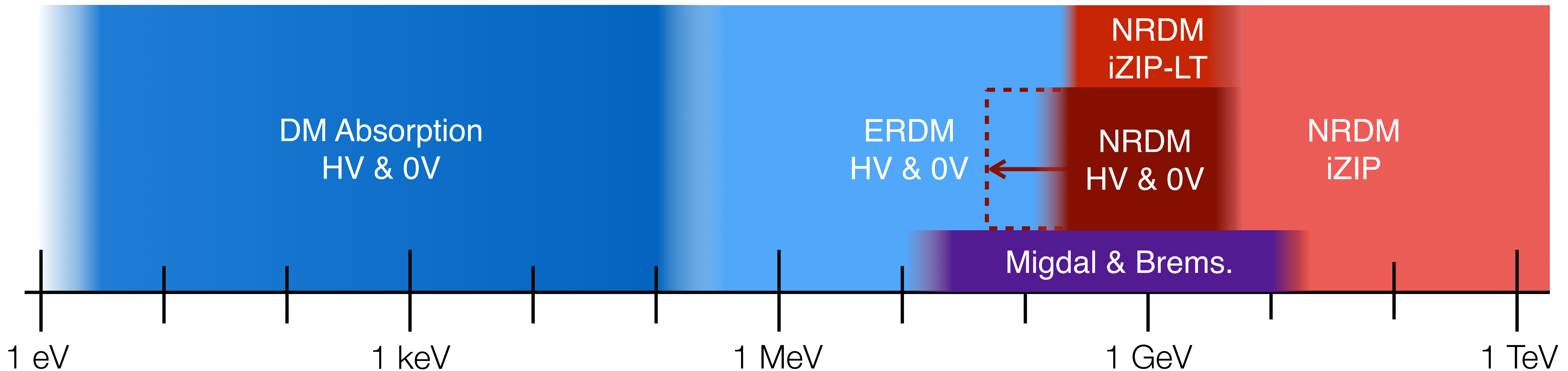
Will discuss electron-scattering DM, dark photon, and ALP constraints in future CFI sessions;
requires modified combination of detector and background improvements



Detectors & DM Mass Scales



- SuperCDMS Detectors able to conduct DM searched over a mass range spanning 1 eV to 1 TeV
 - Nuclear recoil DM sensitivity: $\mathcal{O}(10 \text{ MeV}) - \mathcal{O}(10 \text{ GeV})$
 - Electron recoil DM: $\mathcal{O}(0.1 \text{ MeV}) - \mathcal{O}(10 \text{ GeV})$
 - DM Absorption: $\mathcal{O}(1 \text{ eV}) - \mathcal{O}(0.5 \text{ MeV})$



HV Detector Technology

Low Threshold → Low Mass DM

- Drifting electrons/holes across a potential (V_b) generates a large number of phonons (Luke phonons).
- Enables very low thresholds!

$$E_t = E_r + N_{eh} e V_b$$

total phonon energy
primary recoil energy
Luke phonon energy

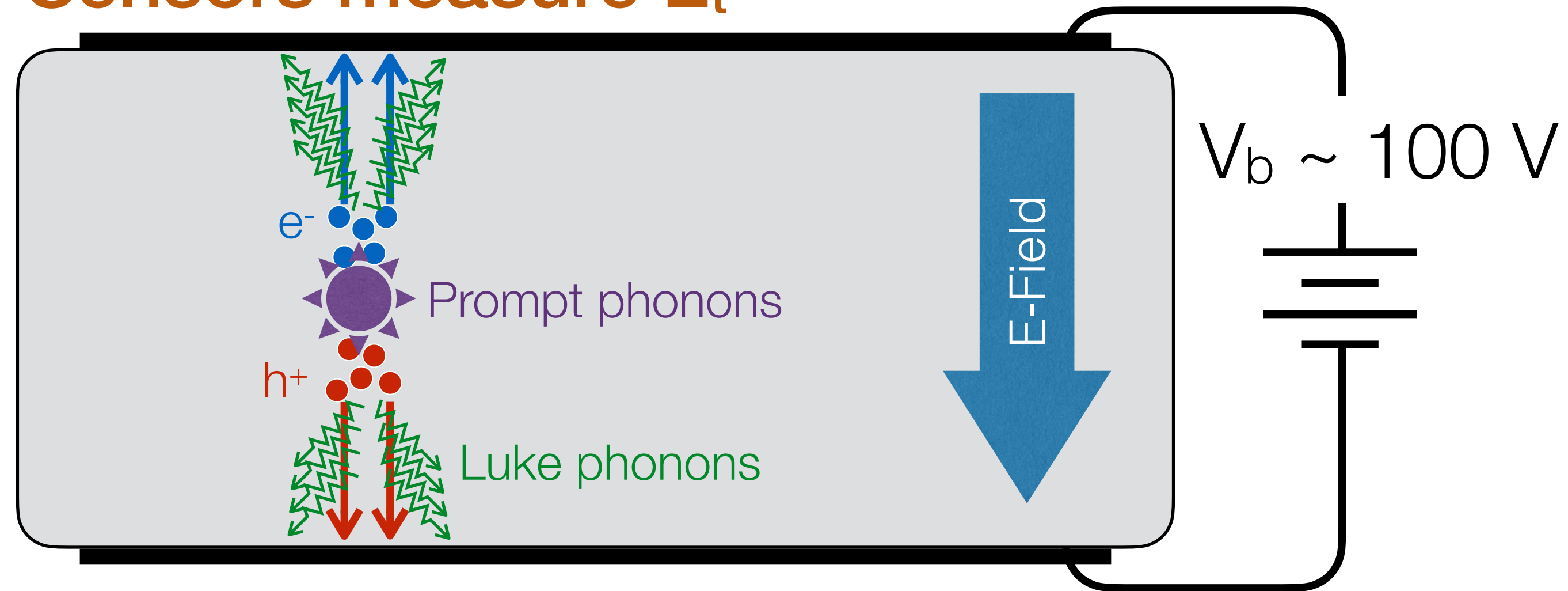
$$\sigma_{ER-eh} = \sigma_{Phonon} / V_b$$

$$\sigma_{ER-effective} = \sigma_{Phonon} / (V_b / \mathcal{E})$$

- Retains ability to discriminate between NR/ER events if:

$$eV_b > \sigma_{Phonon}$$

Sensors measure E_t

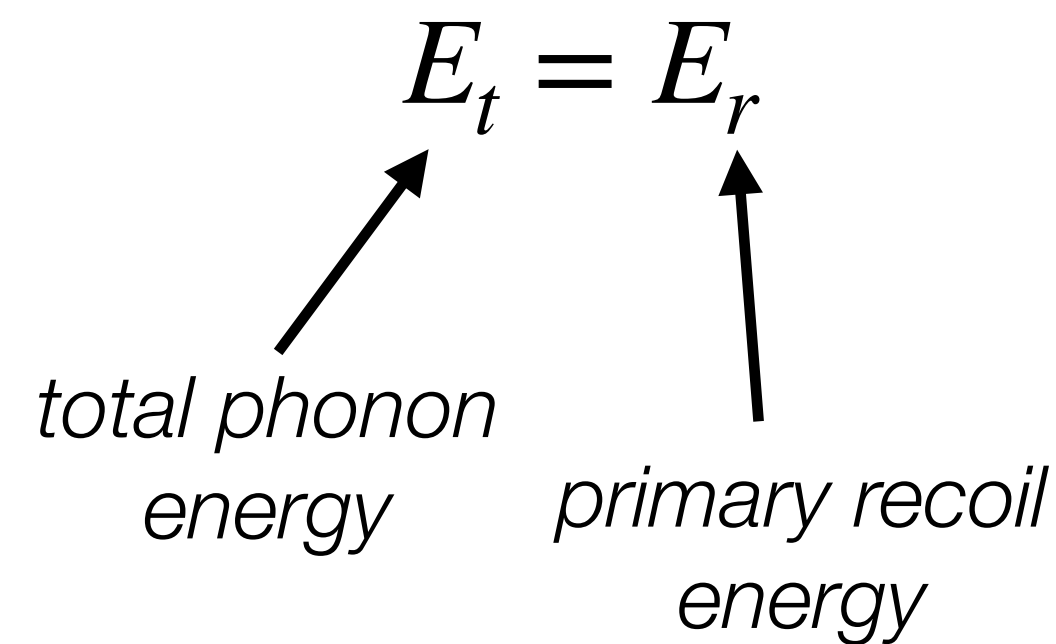


- Signal amplification / lower threshold benefit from high V_b
- V_b dependent charge leakage can be a significant background
- Technology challenge lies in keeping charge leakage under control for needed values of V_b

0V Detector Technology

Low Threshold → Low Mass DM

- Does not drift electrons/holes across the crystal. Measures recoil phonons.

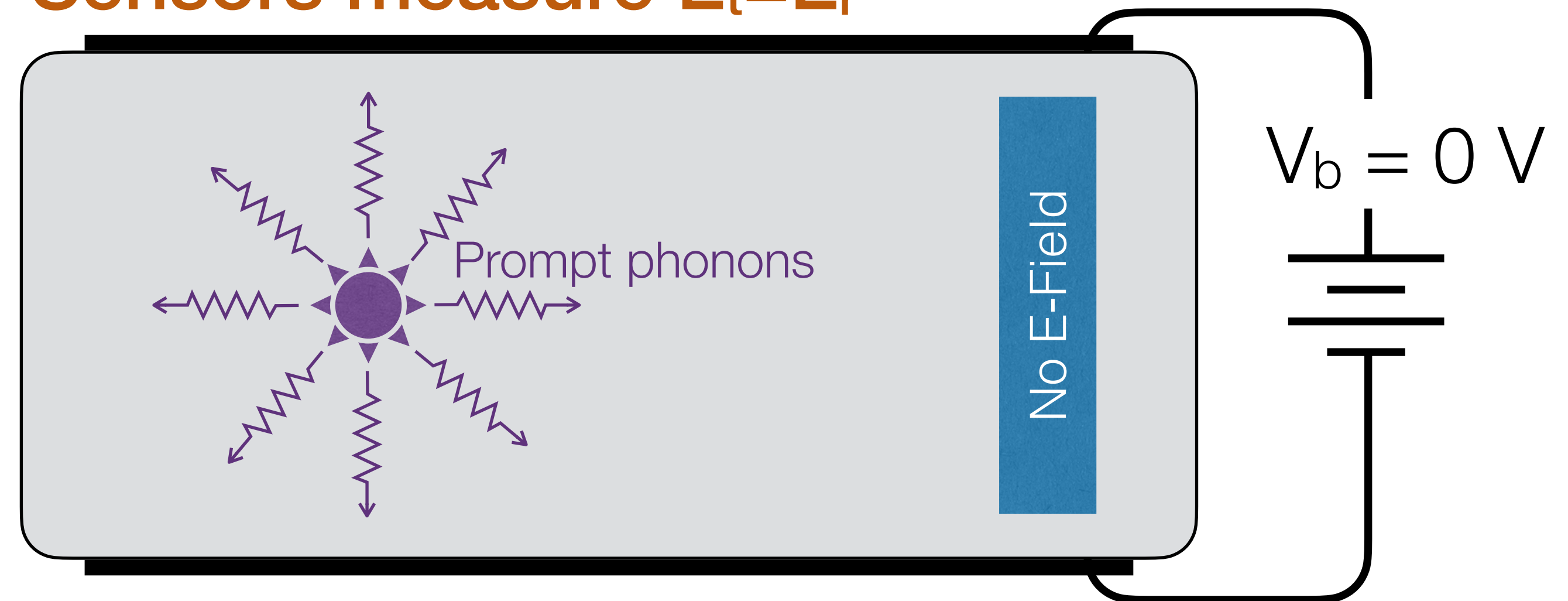


- Thresholds determined by σ_{Ph}

$$\sigma_{ER-eh} = \sigma_{Phonon}$$

$$\sigma_{ER-effective} = \sigma_{Phonon}$$

Sensors measure $E_t = E_r$



- Measurement insensitive to details of NR or ER ionization yield
- Charge leakage background is non-existent
- Technology challenge lies in achieving needed phonon resolution while maintaining large detector masses

Background and technical challenges of HV and 0V detectors are orthogonal

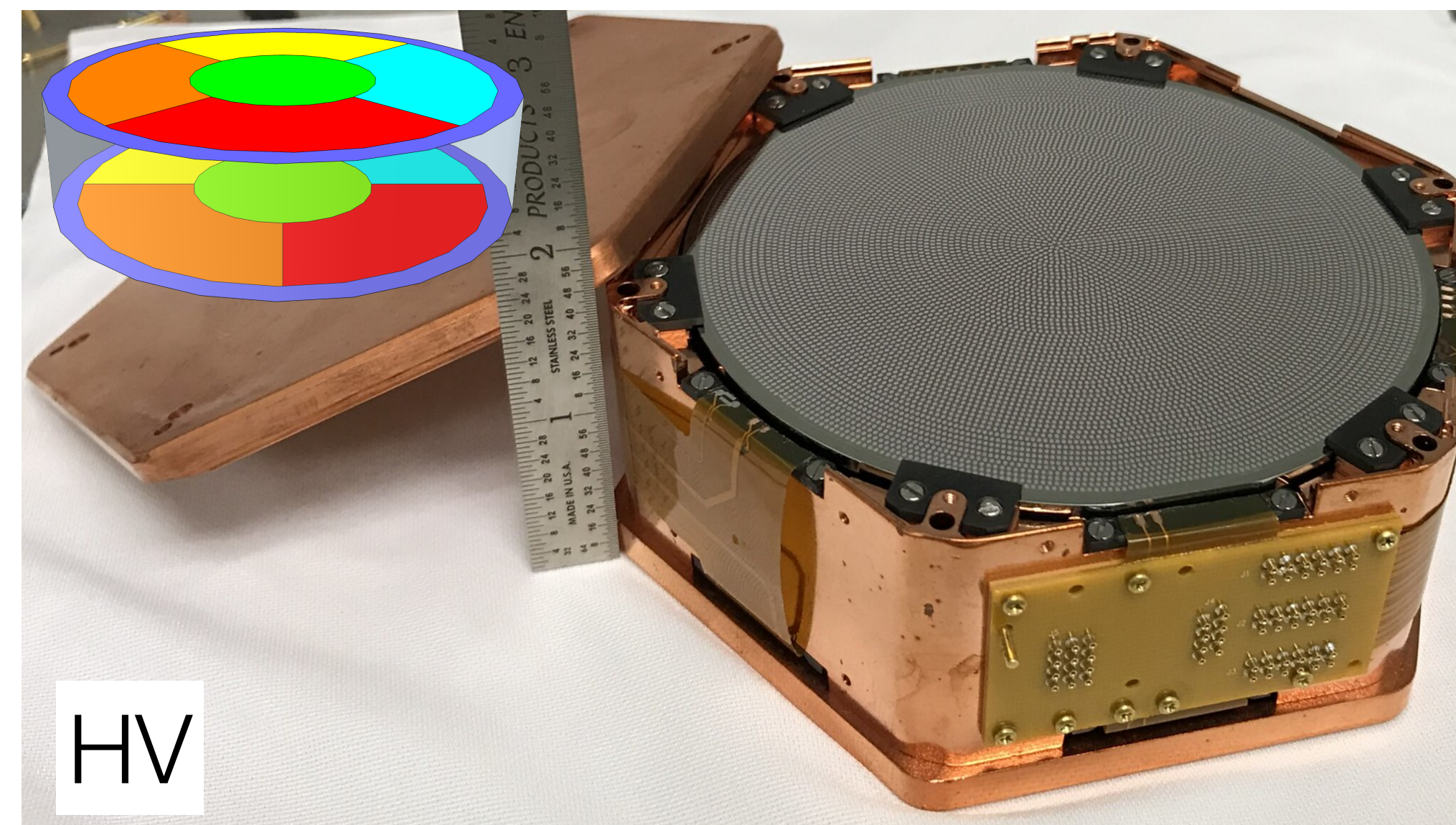
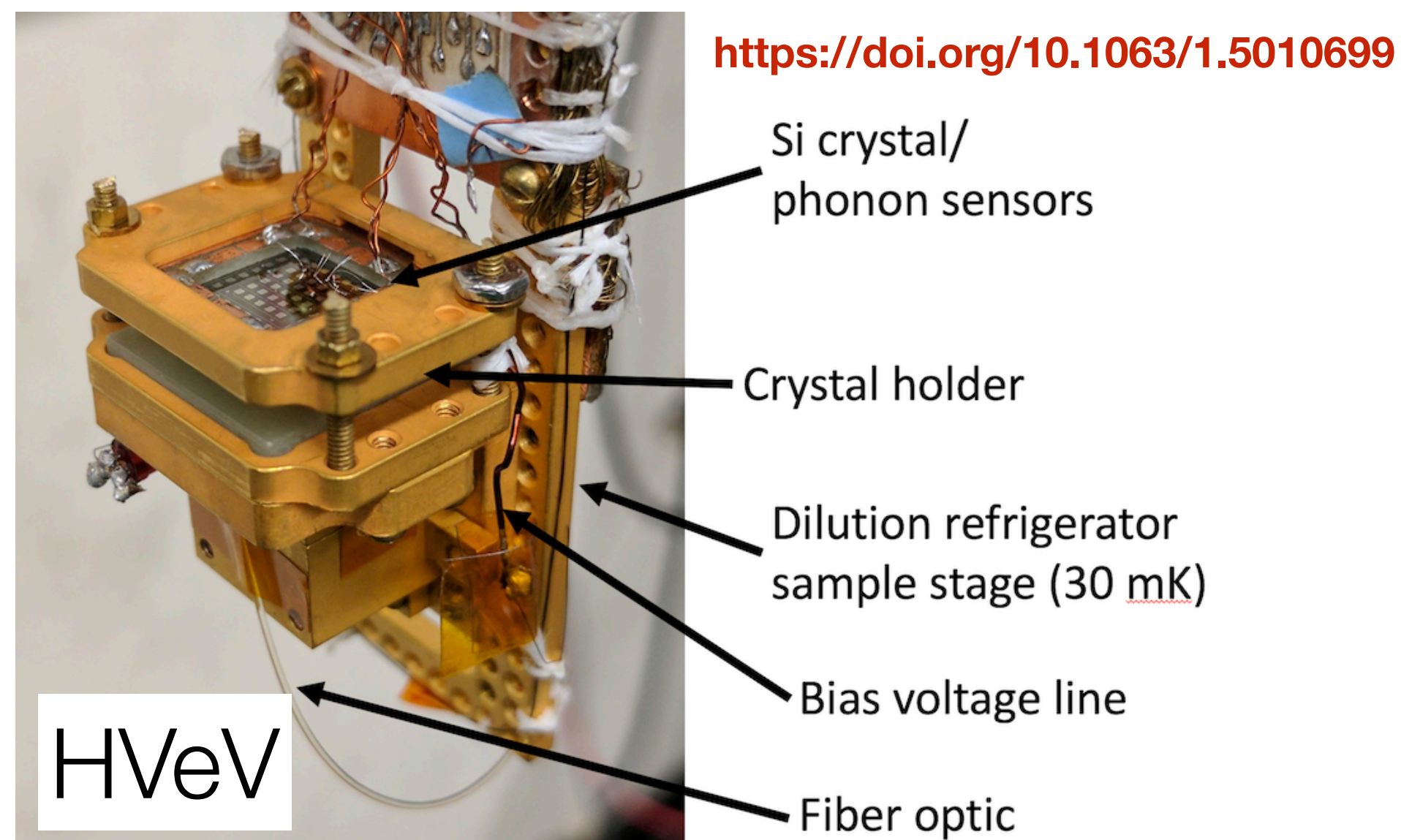
ER Detection Energy Scales



ERDM Mass [MeV]	Required Energy Threshold [eV]	HV σ_{Ph} @ 100V	0V σ_{Ph}	ALP/DP Mass [eV]	Required Energy Threshold [eV]	σ_{Ph} @ 100V	σ_{Ph} @ 0V
0.5	1	15 eV	0.15 eV	1	1	15 eV	0.15 eV
5	15	200 eV	2 eV	10	10	150 eV	1.5 eV
10	30	400 eV	4 eV	100	100	1500 eV	15 eV

ERDM values presented for Si target

SuperCDMS HVeV & HV Detectors

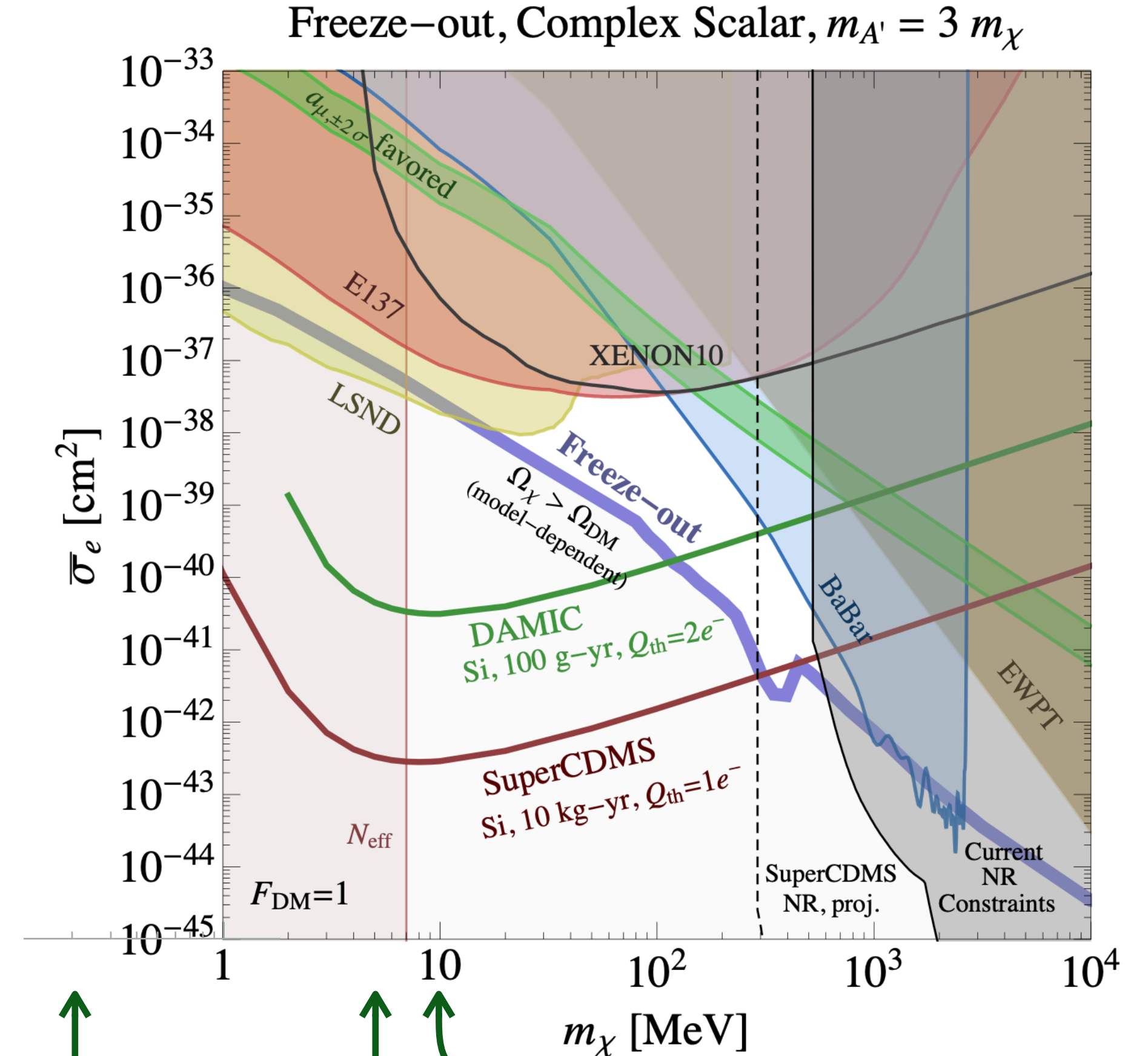
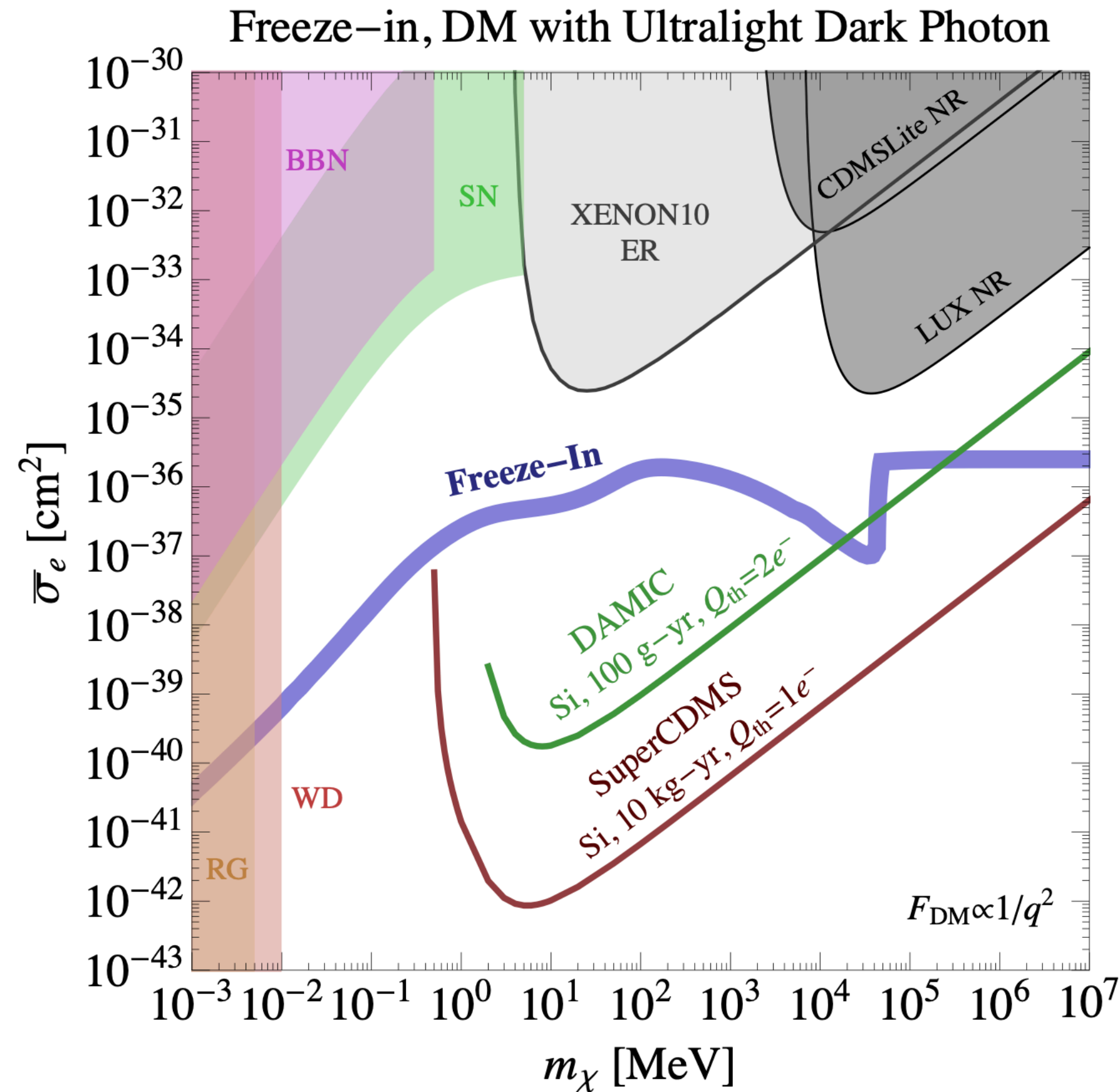


- R&D Detectors:
 - $\varnothing 10\text{mm} \times 4\text{mm}$;
 - Si: 0.93 g
 - σ_{Ph} : 3 eV <https://doi.org/10.1016/j.nima.2020.163757>
[arXiv:2005.14067](https://arxiv.org/abs/2005.14067)
 - $E_{\text{th}} \sim 21\text{ eV}$
 - Single e^-/h^+ pair resolution: $\sigma_{\text{eh}} \sim 0.03\text{ e}^-/h^+$.
 - Has sensitivity to a variety of sub-GeV DM models with gram*day exposures

- Payload: 4 Si and 8 Ge detectors
 - $\varnothing 100\text{mm} \times 33\text{mm}$;
 - Si: 0.6 kg, Ge: 1.4 kg
 - SNOLAB Goal σ_{Ph} : 7 eV (Si), 10 eV (Ge)
 - $E_{\text{th}} \sim 50\text{ eV}$ (Si), 70 (Ge)
 - Single e^-/h^+ pair resolution σ_{eh} @100V bias: Si: $\sim 0.07\text{ e}^-/h^+$, Ge: $\sim 0.1\text{ e}^-/h^+$
 - Has sensitivity to various relic abundance DM models of sub-GeV DM with kg*year exposures

← Better resolution Tradeoff → Greater mass

Large Exposure Sensitivity



- Phonon resolution determines minimum accessible DM mass.
- SNOLAB HV detectors at SNOLAB in principle have sensitivity and exposure to reach ERDM relic models.
- Current great uncertainty is magnitude of leakage currents in the lowest e⁻/h⁺ bins
- R&D is needed to achieve
 - necessary thresholds shapes for 0V detectors and
 - control charge leakage for 100 V detectors

$Q_{th} = 1e^-$
 $\sigma_{Ph} @ 0V = 0.15 eV$
 $\sigma_{Ph} @ 100V = 15 eV$

$Q_{th} = 5e^-$
 $\sigma_{Ph} @ 0V = 2 eV$
 $\sigma_{Ph} @ 100V = 200 eV$

$Q_{th} = 10e^-$
 $\sigma_{Ph} @ 0V = 4 eV$
 $\sigma_{Ph} @ 100V = 400 eV$

Detector R&D to Realize the Science Program



Science Target		Potential Detector R&D Topics – to be narrowed							
		Improved Phonon Resolution	Different Substrate Size/ Geometry	Improved Charge Leakage	Improved Phonon Position Information	Improved Charge Noise	Different Targets	Active Veto of Photons	Active Veto of Betas
<0.5 GeV nuclear-scattering DM Dark γ via Optical Phonon		✓	✓		desirable		✓	✓	desirable
Electron-scattering DM Dark γ ALPs	with NTL gain (HV)	for large substrates		✓					
	0V	✓	✓		desirable		desirable	✓	desirable
>0.5 GeV nuclear-scattering DM	Quantization via NTL gain (HV)	for large substrates	potentially	✓					
	ER rejection to lower energies (iZIP)	for large substrates	potentially			✓			desirable
	Phonon Position Information	✓	potentially	✓	✓				
	Photon/Phonon Discrimination	✓	potentially				✓		✓

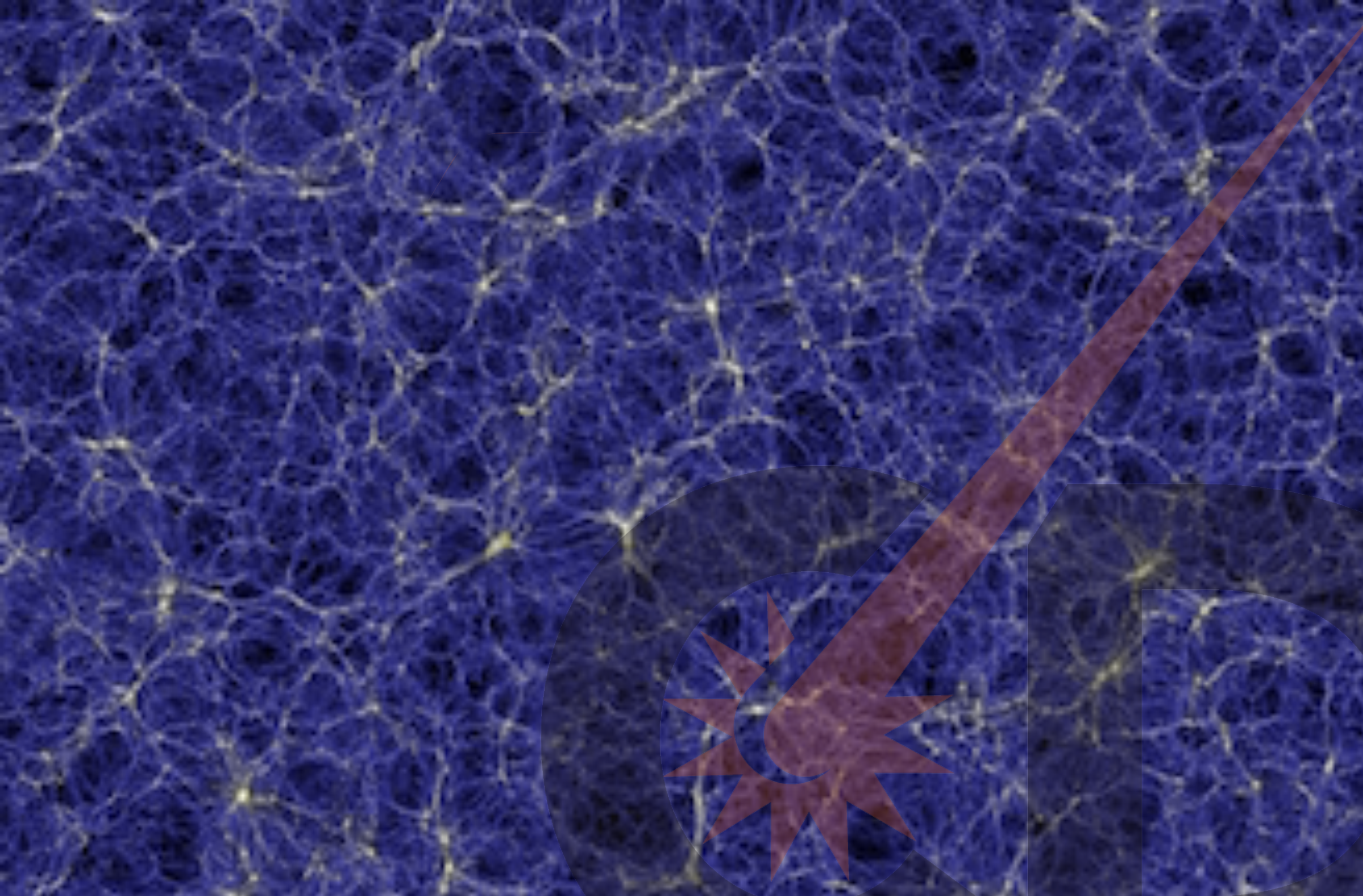
Goals of R&D topics:

- Phonon position information: Allows event fiducialization to reject sidewall leakage and surface backgrounds
- Different targets: Possibility of lower band-gap, sensitivity to direct phonon creation if reach $\mathcal{O}(10 \text{ meV})$ resolution
- Active vetos: Reduction of backgrounds originating at the surfaces nearest the detectors

Conclusions



- SuperCDMS detector technology being developed to search for electron recoil light dark matter, down to
 - $\mathcal{O}(1)$ eV via Dark Photon Absorption and Axion dark matter channels
 - $\mathcal{O}(1)$ MeV via Electron recoil channels and
- This complements nuclear recoil dark matter searches in
 - $\mathcal{O}(10)$ MeV via inelastic Nuclear recoil channels (Migdal, Bremsstrahlung)
 - $\mathcal{O}(1-100)$ GeV via elastic Nuclear recoil channels
 - Cryogenic detectors aiming to reach “neutrino floor” in the 1-10 GeV NR mass range



down

... the end