Neutrino mass models at colliders in the post-ESU 2020 era

Energy Frontier Biweekly Meeting

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Acknowledgments, Apologies, and Disclaimers

finite time constraints \implies many omissions

- Purpose: a snapshot of activities since European Strategy Update
- Main focus is on Type I (N) and Type II ($\Delta^{\pm\pm}$) Seesaws
- See references below for details + other Seesaws

source material:

- Review on ν mass models at colliders w/ Y. Cai, T. Han, T. Li [1711.02180]
- 2 European Strategy Update Chapter on ν mass models

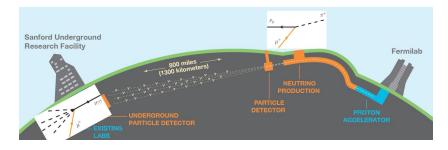
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w/ T. Han, T. Li, X. Marcano, S. Pascoli, C. Weiland [1812.07831]
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Other community documents and some newer results

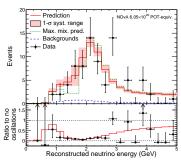
humble reminder: RH neutrinos (ν_R) are **not** the only explanation for tiny m_{ν} nor are they necessary (e.g., Type II Seesaw)

- ullet Lack of guidance from data \Longrightarrow broad approach needed
- E.g., models without ν_R , UV completions of NSI,

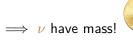
the big physics picture



In neutrino fixed-target expts, ν_{μ} beams from collimated π^{\pm} , then studied at near and far detectors



Deficit/disappearance of expected ν_{μ} (+apperance of $\nu_e/\nu_{\tau})$ interpreted as $\nu_{\ell_1} \rightarrow \nu_{\rm mass} \rightarrow \nu_{\ell_2}$ transitions/oscillations [E.g. NO ν A ν_{μ} disapp., 1701.05891]





So, neutrinos have masses $\lesssim \mathcal{O}(0.1)$ eV

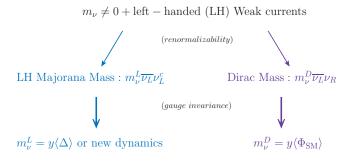
Is this a problem?

Yes.

Neutrinos Masses and New Particles?

Nonzero neutrino masses implies new degrees of freedom exist

[Ma'98]



$m_{\nu} \neq 0$ + renormalizability + gauge inv. \implies new particles!

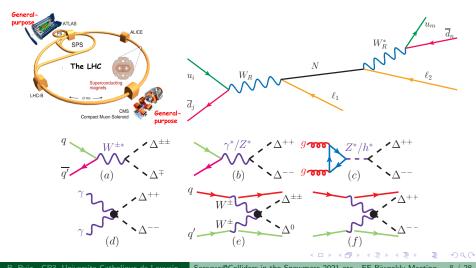
- New particles might be charged under new or old gauge symm., E.g., ν_R may have $U(1)_{B-L}$ charge and Δ_L is an $SU(2)_L$ triplet
- Particles must couple to *h* or *L*, often inducing LNV/cLFV!

the slightly-less-big picture

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models that explain tiny neutrino masses (Seesaw models) are testable, especially at colliders

for a review, see w/ Y. Cai, T. Li, and T. Han [1711.02180] as well as w/ Pascoli, et al [1812.08750]



the little picture

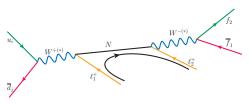
(our part!)

Snowmass 2013 inspired an effort to systematically modernize the collider phenomenology for Seesaw models

for example

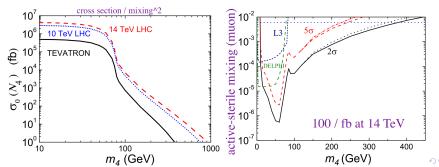
Historically, searches for N with $m_N > M_W$ relied on $(q\bar{q})$ annihilation

Keung & Senjanovic (PRL'83)



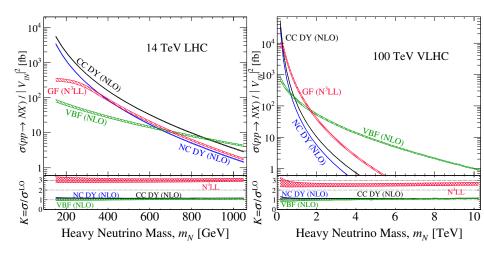
At the LHC, a canonical signature for N: $pp \rightarrow \ell_i^{\pm} \ell_i^{\pm} + nj + \text{ no MET}$

based on seminal works by K&S, del Aguila & Aguilar-Saavedra [0808.2468], and Atre, et al [0901.3589]

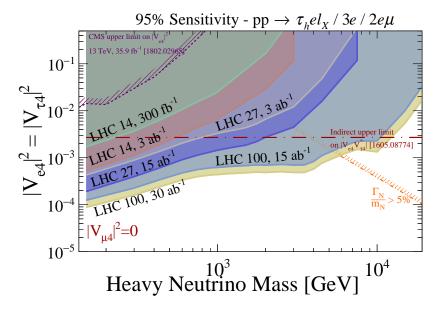


a lot has happened since 2013

Plotted: Normalized production rate $(\sigma/|V|^2)$ vs heavy N mass (m_N)



For $m_N = 10$ TeV and $|V_{\ell N}|^2 \sim 10^{-3}$, then at 100 TeV, one has $\mathcal{O}(30)$ VBF events after 30 ab⁻¹! If BR× ε × \mathcal{A} ~ $\frac{1}{3}$, then $\sqrt{N_{Obs.}}$ > 3σ

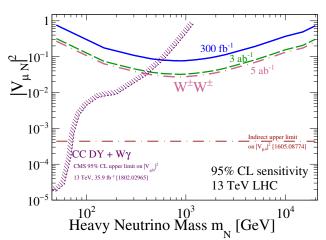


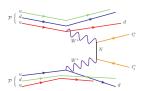
Major improvements $\implies > 10 \times$ better sensitivity to LNV + cLFV

Only one example. See the big paper [1812.08750] for various flavor, Dirac vs Majorana, and sys permutations ?

How heavy is too heavy for the LHC?

Question: is a multi-TeV N too heavy for the LHC?





w/ Fuks, Neundorf, Peters, Saimpert [In Prep.]

what if there are new forces?¹ how heavy can we go?

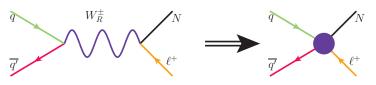


¹See also talk by N. Okada!

Interesting observation: vast literature on collider searches for N coupling to new gauge bosons, e.g., W_R in Left-Right Symmetric Model, nearly everyone assumes that both N and W_R are resonantly produed

If new gauge mediators are too heavy, light N are still accessible

(this is a UV realization of ν_R EFT!)

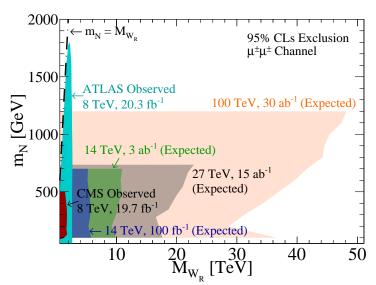


Exmaple: When $M_{W_R} \gg \sqrt{\hat{s}}$ but $m_N \lesssim \mathcal{O}(1)$ TeV, $pp \to N\ell + X$ in the LRSM and phenomenological Type I Seesaw are not discernible

w/ Han, Lewis, Si, [1211.6447]; RR, [1703.04669]

• Same signature: $pp \to \ell^{\pm}\ell^{\pm} + nj + X + p_T^{\ell} \gtrsim \mathcal{O}(m_N) + \text{no MET}$

How about reinterpreting search for phenomenological Type J Seesaw?

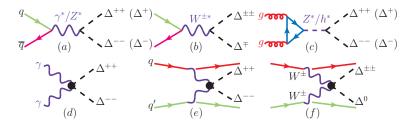


RR [1703.04669]

At 14 (100) TeV with $\mathcal{L}=1$ (10) ab^{-1} , $M_{W_R}\lesssim 9$ (40) TeV probed DO NOT STOP SEARCHING FOR TYPE I LNV

what if ν_R do not exist?

Type II Seesaw²



²Konetschny and Kummer ('77); Schechter and Valle ('80); Cheng and Li ('80); Lazarides, et al ('81); Mohapatra and Senjanovic ('81)

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• Important example that $m_{\nu} \neq 0 \not\Rightarrow$ that ν_R exist

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Hypothesize a scalar $SU(2)_L$ triplet with lepton number L=-2

$$\hat{\Delta} = \frac{1}{\sqrt{2}} \begin{pmatrix} \Delta^+ & \sqrt{2} \Delta^{++} \\ \sqrt{2} \Delta^0 & -\Delta^+ \end{pmatrix}, \quad \text{with} \quad \mathcal{L}_{\Delta\Phi} \ni \mu_{h\Delta} \Big(\Phi^\dagger \hat{\Delta} \cdot \Phi^\dagger + \text{H.c.} \Big)$$

The mass scale $\mu_{h\Delta}$ breaks lepton number, and induces $\langle \Delta \rangle \neq 0$:

$$\sqrt{2}\langle\hat{\Delta}
angle = extstyle extstyle extstyle ag{\mu_{h\Delta} v_{
m EW}^2}{\sqrt{2}m_{\Delta}^2}$$

which leads to left-handed Majorana masses for neutrinos

$$\Delta \mathcal{L} = -\frac{y_{\Delta}^{ij}}{\sqrt{2}} \overline{L^{c}} \hat{\Delta} L = -\frac{y_{\Delta}^{ij}}{\sqrt{2}} \left(\overline{\nu^{jc}} \quad \overline{\ell^{jc}} \right) \begin{pmatrix} 0 & 0 \\ v_{\Delta} & 0 \end{pmatrix} \begin{pmatrix} v^{i} \\ \ell^{i} \end{pmatrix}$$

$$\ni -\frac{1}{2} \left(\sqrt{2} y_{\Delta}^{ij} v_{\Delta} \right) \overline{\nu^{jc}} v^{i}$$

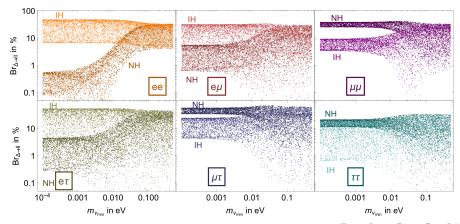
$$= m^{ij}$$

Fewer free parameters \implies richer experimental predictions

Fileviez Perez, Han, Li, et al, [0805.3536], Crivellin, et al [1807.10224], Fuks, Nemevšek, RR [1912.08975] + others

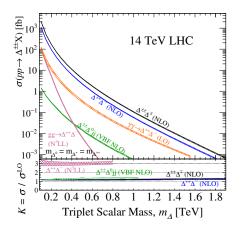
• E.g., Δ branching rates encode inverse (IH) vs normal (NH) ordering of light neutrino masses

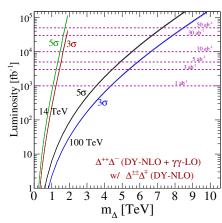
$$\mathsf{BR}(\Delta^{\pm\pm} \to \ell_i^{\pm} \ell_i^{\pm}) \sim y_{\Delta}^{ij} \sim (U_{\mathrm{PMNS}}^* \tilde{m}_{\nu}^{\mathrm{diag}} U_{\mathrm{PMNS}}^{\dagger})_{ij}$$



NEW: a revised outlook for both $\sqrt{s} = 14$ TeV and 100 TeV!

w / Fuks and Nemevšek [1912.08975]





- At LHC with $\mathcal{L}=5$ ab⁻¹, 3σ sensitivity up to $m_{\Lambda}\sim 1.5$ TeV
- At $\sqrt{s} = 100$ TeV with $\mathcal{L} = 30 50$ ab⁻¹ $\implies m_{\Lambda} \approx 8 9$ TeV
- Warning: can be improve for specialized final state / parameter space

Lots of improvement since last Snowmass. What has changed?

Improved outlook for collider tests of LNV and cLFV stems from:

- New channels, e.g., VBF, GF, $W/Z/h/\gamma$ associated production
- New kinematic limits, e.g., off-shell portals, boosted topologies
- Predictions for both Dirac and Majorana particles w/ LNV and cLFV
- Quantitatively reliable descriptions of jets, kinematics, and rates

³UFOs encode Feynman rules for mainstream event generators, e.g. MadGraph, to simulate BSM (not just colliders)

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Part of this stems from improved MC support!

- Ongoing efforts within FeynRules and MadGraph MC collaborations
- Mainstream tools with widespread use and technical support

Available UFOs3

- Type I Seesaw feynrules.irmp.ucl.ac.be/wiki/HeavyN (Requested/used by ATLAS+CMS)
- Type II Seesaw feynrules.irmp.ucl.ac.be/wiki/TypellSeesaw (Requested/used ATLAS)
- Left-Right Symmetry feynrules.irmp.ucl.ac.be/wiki/EffLRSM (Requested/used ATLAS)
- Generic W'/Z' feynrules.irmp.ucl.ac.be/wiki/WZPrimeAtNLO
- ... with more in development (collaborators and friends are welcome!)

Summary

Lack of clear guidance from data and theory means we must take a broad, open approach to uncovering the origin of tiny ν masses.

- Colliders are incredibly complementary to oscillation and $0\nu\beta\beta$ expts
 - Direct production of Seesaw particles
 - ► Test UV realizations of low-scale neutrino EFTs / NSIs
- The European Strategy Update has officially concluded
 - Lots of encouraging projections on collider sensitivity to LNV and cLFV
 - ▶ New analysis techniques ⇒ new territory for cLFV and LNV
 - ightharpoonup N, $H^{\pm\pm}$, W_R , Z_{B-L} , $T^{0,\pm}$ masses up to 10-50 TeV at $\sqrt{s}=100$ TeV
 - Studies aided by publication of user friendly simulation tools
- The Snowmass Process is underway!
 - Community studies are iterative and we plan to keep up the work!
 - ▶ Lots not covered todays, so go check out the review! [1711.02180]



Thank you.