

*may include some non-standard interactions with other fields...

Enectalí Figueroa-Feliciano Northwestern

From BRN Report:

Ultra-low temperature test stands

As described in the dark matter and quantum sensor sections, these areas of research will impose exacting requirements on ultra-low temperature (mK) environments and significant testing will be required to realize future experiments. Yet, test stands for operating at these temperatures are costly and limited in number. These high value assets are dispersed throughout the community, primarily at national labs and large research universities. Some were funded by HEP while many have been funded by institutional investments and other government agencies.

The expansion of interest in such facilities is a relatively recent phenomenon, driven largely by the quickly growing HEP-QIS overlap, and thus there is not yet a consensus on the scale of community need and what siting and operating models ensure appropriately broad access and sustainable operation. Hand-in-hand with facilities supply and access is the question of what explicit efforts ought to be made to disseminate expertise in areas relevant to work that use such facilities—in particular device design and fabrication and experimental techniques—and to facilitate access. Further study by the relevant communities would clarify the situation and better define a path forward. Such study would also determine the potential impact of technical or manufacturing advances that reduce the cost of such test stands, though it is not clear that direct HEP funding for such work would substantially augment the much larger commercial and governmental interest driven by quantum computing.







Cryogenics are Everywhere!

Dark Matter

Neutrinos

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QIS



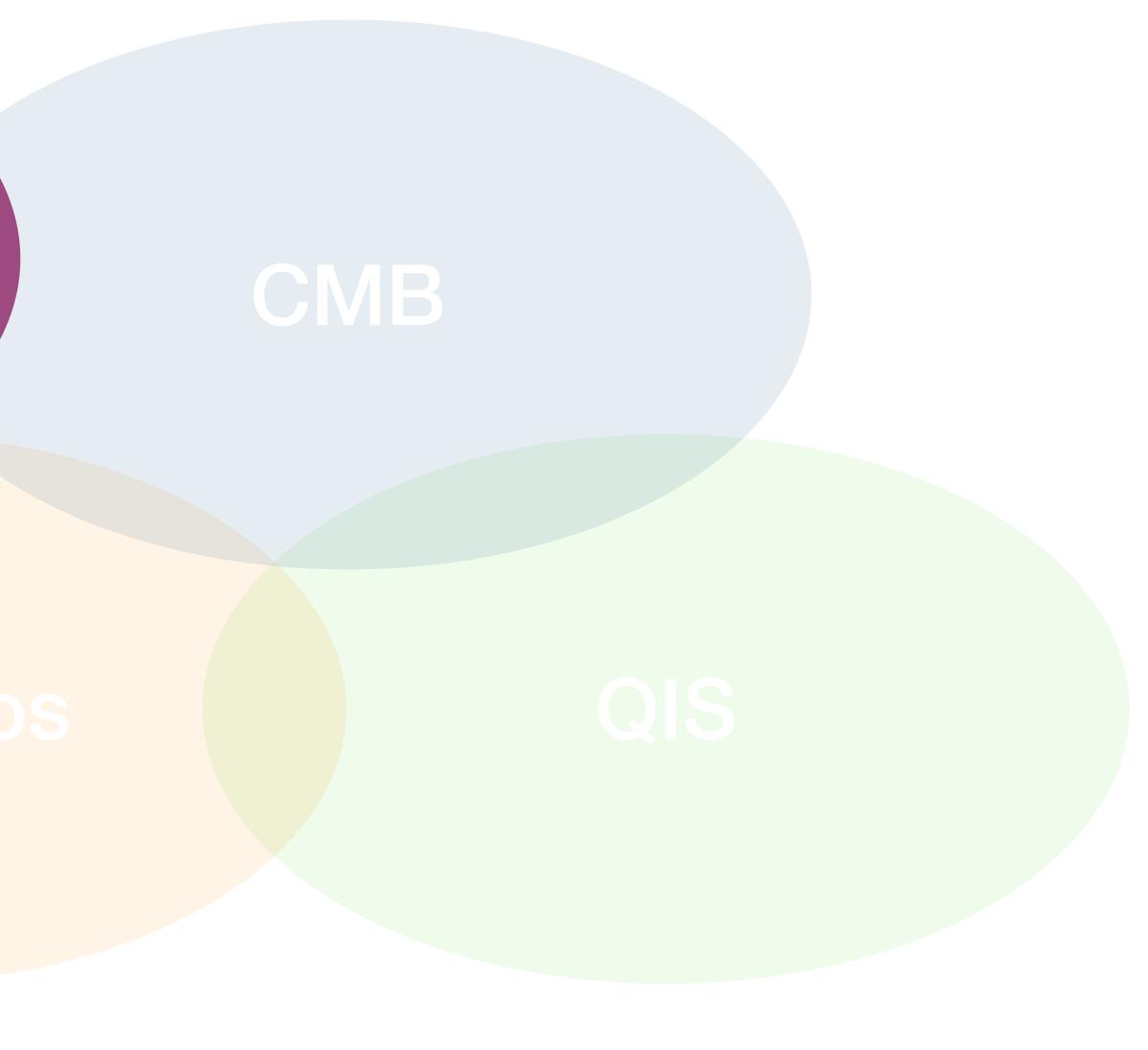


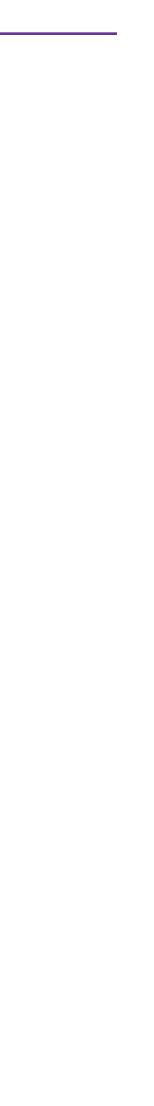
Cryogenics are Everywhere!

Dark Matter

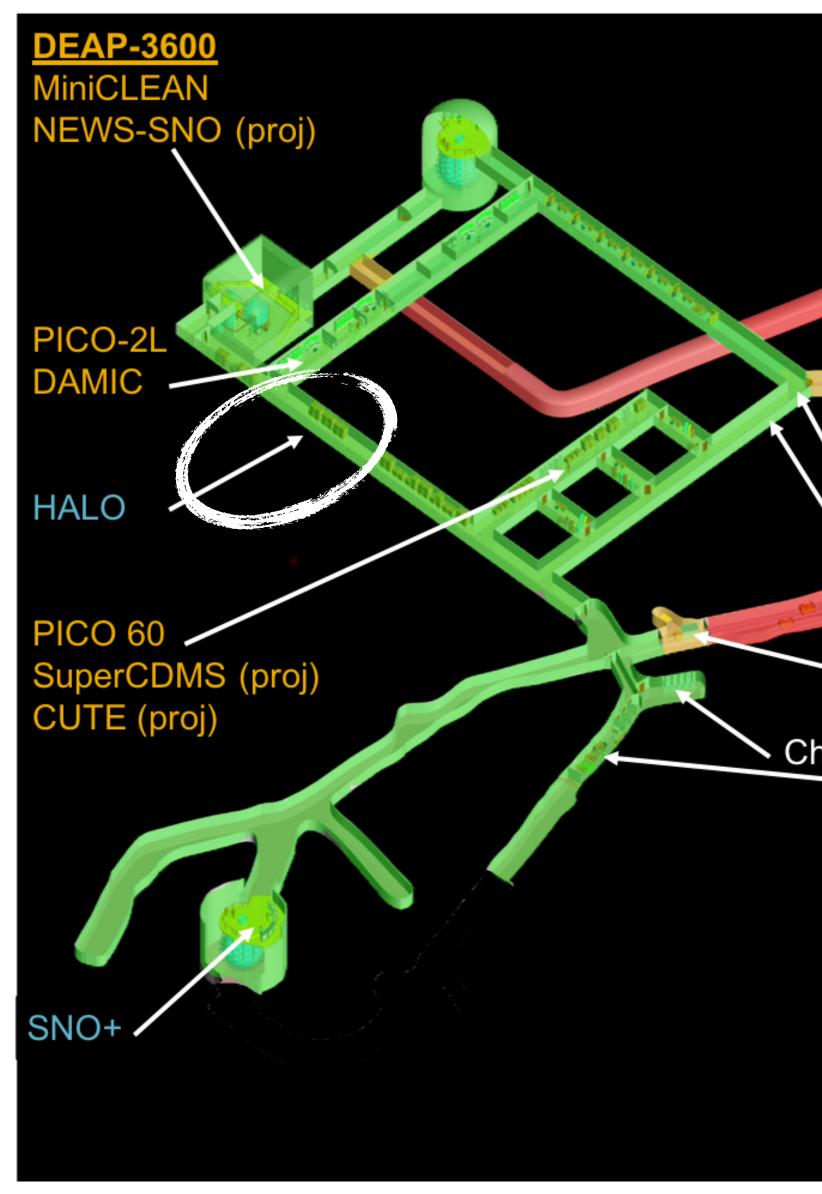
Neutring

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SNOLAB Cryogenic Facilities



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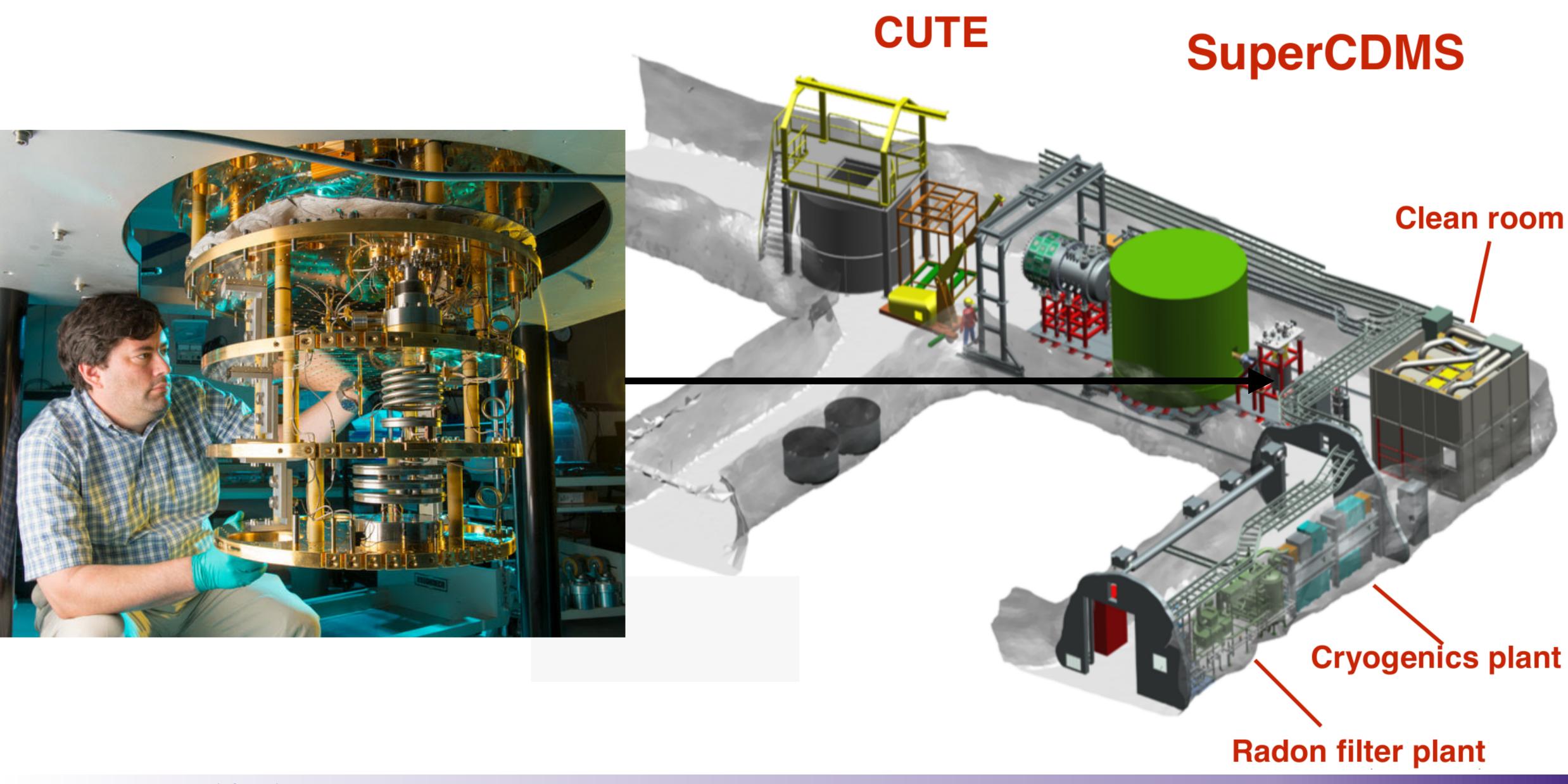


Mine Airlock Clean room (Class 2000)

Machine Shop Chemistry lab Car wash Change rooms/Showers Lunch room Chiller/air filters

5

SNOLAB Cryogenic Facilities



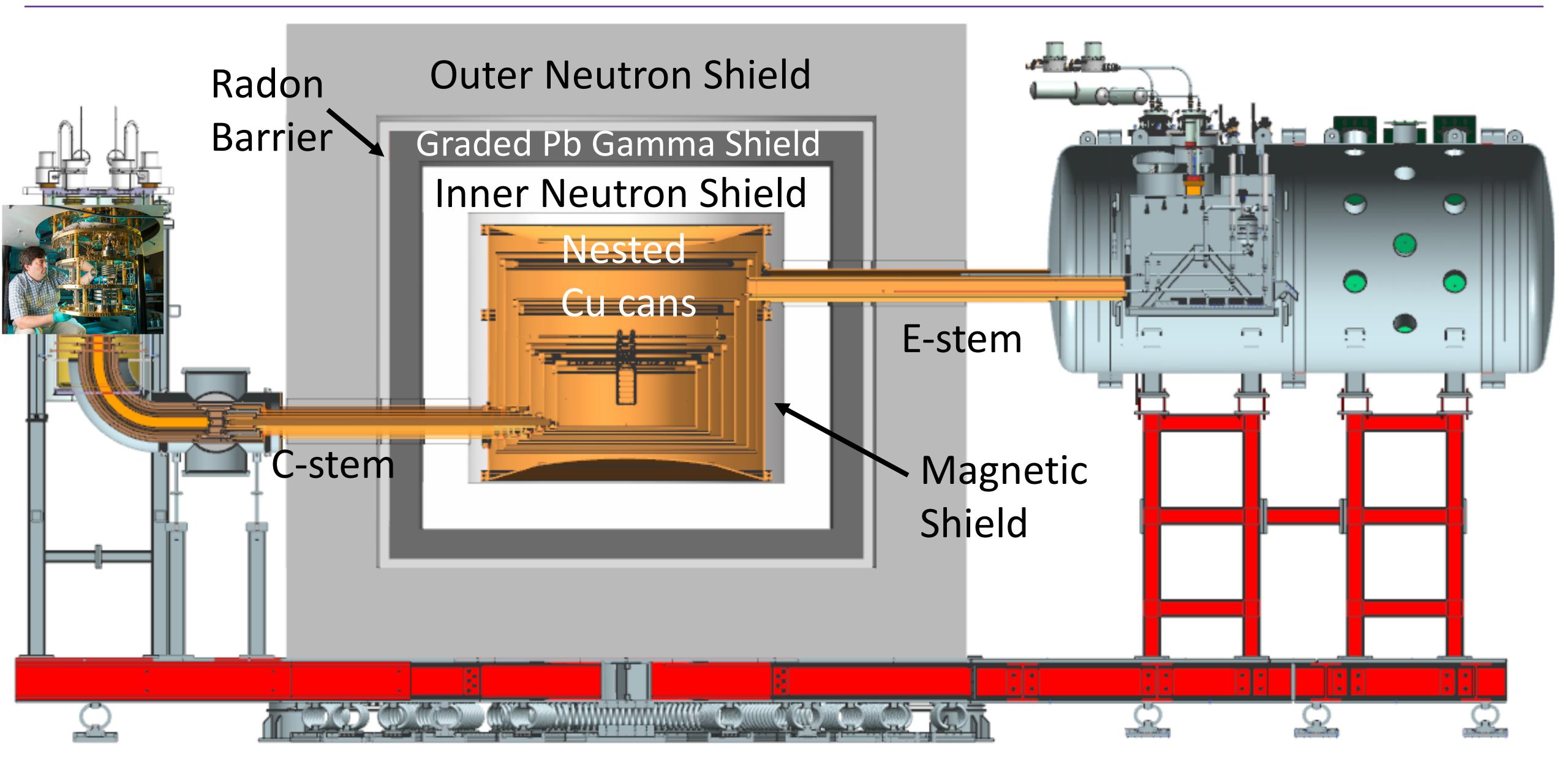
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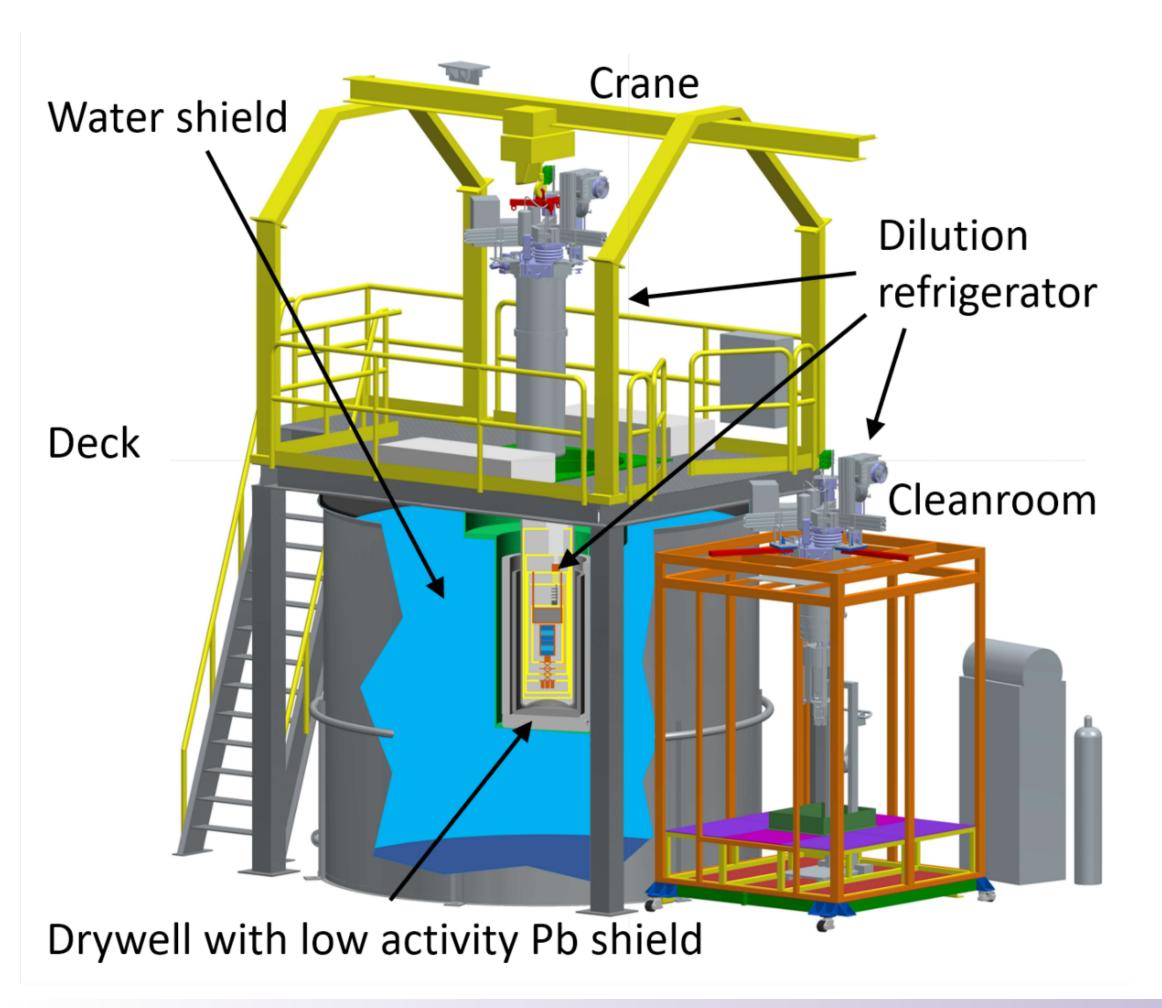
SuperCDMS SNOLAB: Premier Cryo Facility for the Future



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Motivation

- Above-ground testing exposes detectors to cosmogenic radiation / activation
- Large rate at surface may lead to constant pile-up
- Discrimination power (ER/NR) measurement requires absence of neutrons
- need for a well shielded underground test facility

Design

- 3.6 m water tank (n and y shielding), w/ drywell for fridge
- ~12 cm of low-activity Pb inside drywell; 15 cm of Pb inside the fridge (between active components and detectors) to shield direct line of sight from top, 20 cm PE above cryostat
- Low-Rn purge gas expels lab air from drywell
- Clean room for payload change (filtered low Rn air)
- Cryostat mounted on soft dampers to minimize vibration transmission.
- Holds up to 6 SuperCDMS SNOLAB detectors













Status

- Completed / commissioned in 2019
- Performance validation measurements:
 - Cryogenics: Base temperature w/ detectors: 12 mK
 - Background: <10 evts/keV/kg/day above 10 keV
 - Vibration mitigation effective
- Detector testing for SuperCDMS:
 - Small devices
 - X-ray / y calibrations (55Fe, 133Ba)
 - SuperCDMS Pathfinder testing upcoming
- SuperCDMS Data Challenge
 - Same DAQ system as SuperCDMS
 - Full test of data transfer/processing

Plans

- 2021/22:
 - SuperCDMS Pathfinder and Tower testing •
 - First DM search with SuperCDMS detectors as time allows
- Later: open for future projects
 - SuperCDMS upgrades
 - Cryogenic detector testing for other projects
 - Small scale rare event searches with cryogenic • detectors
 - Other cryogenic low-background tests (e.g. cryogenic Q-bits)
- Continued facility improvements
 - Facility reliability / ease of operation
 - **Background improvements**
 - Noise reductions
 - ...











Moving fridge between drywell and cleanroom

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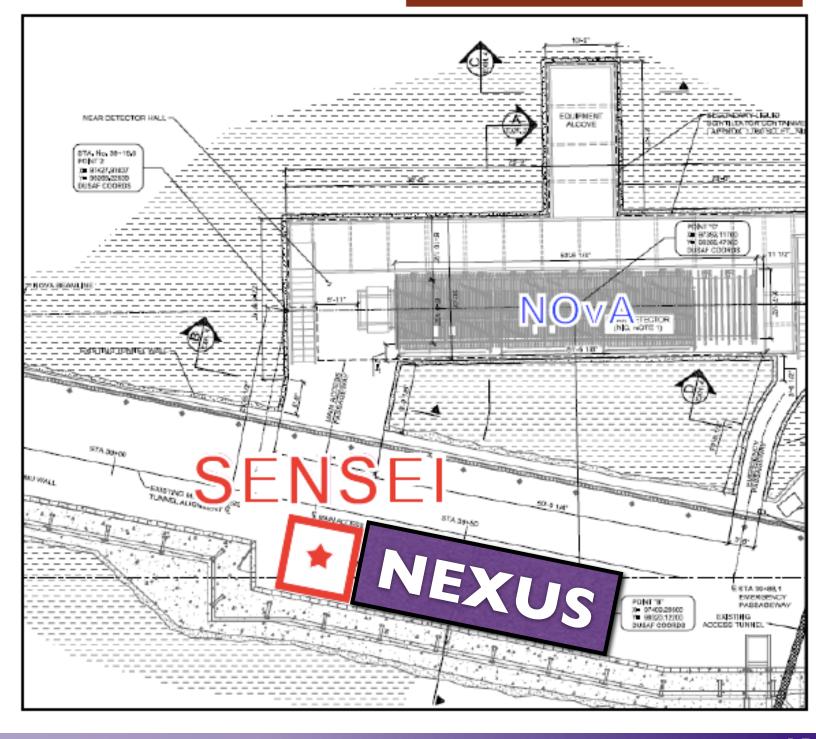


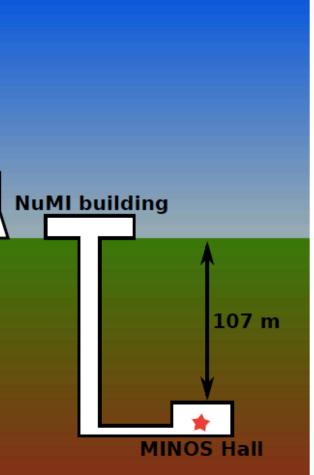


NEXUS@FNAL Facility

- Northwestern Experimental Underground Site at Fermilab
- Underground cryogenic facility in class 10,000 clean room
- Vibration-isolated dry dilution refrigerator with 8mK base temperature
- 107 m depth (300 meters water equivalent) + lead shielding
 - Design background < 100 events/keV/kg/day
- Optical fiber, neutron, gamma calibration sources









Generator

DD

. T



Location for

Backing Array

Dilution Refrigerator



Detector Calibration via Scattering Angle Measurement

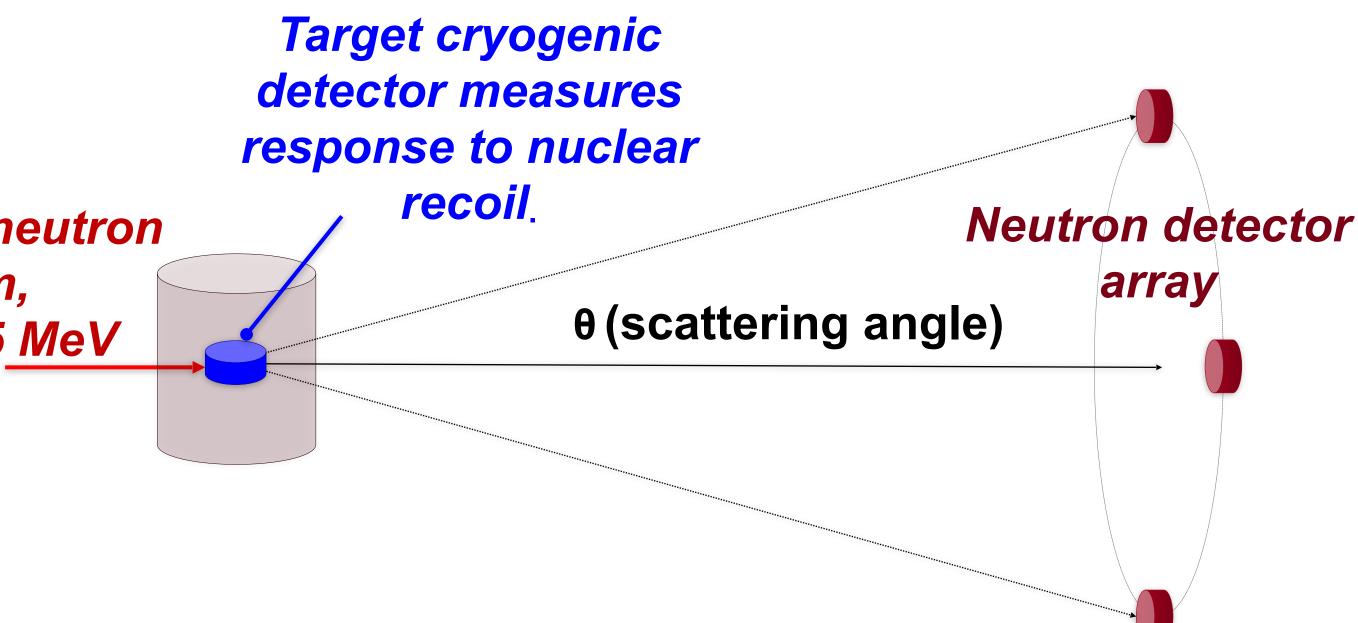
- A standard way characterize detector response to nuclear recoils is with a neutron scattering setup
- Precise knowledge of the scattering angle provides the recoil energy in the detector
- The detector signal can then be calibrated as a function of neutron recoil energy



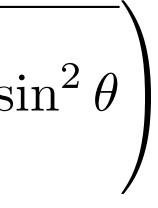
Incoming neutron beam, $E_n = 2.45 \, MeV$



$$E_{\text{recoil}} = 2E_n \frac{M_n^2}{(M_n + M_T)^2} \left(\frac{M_T}{M_n} + \sin^2\theta - \cos\theta\sqrt{\left(\frac{M_T}{M_n}\right)^2 - s}\right)$$



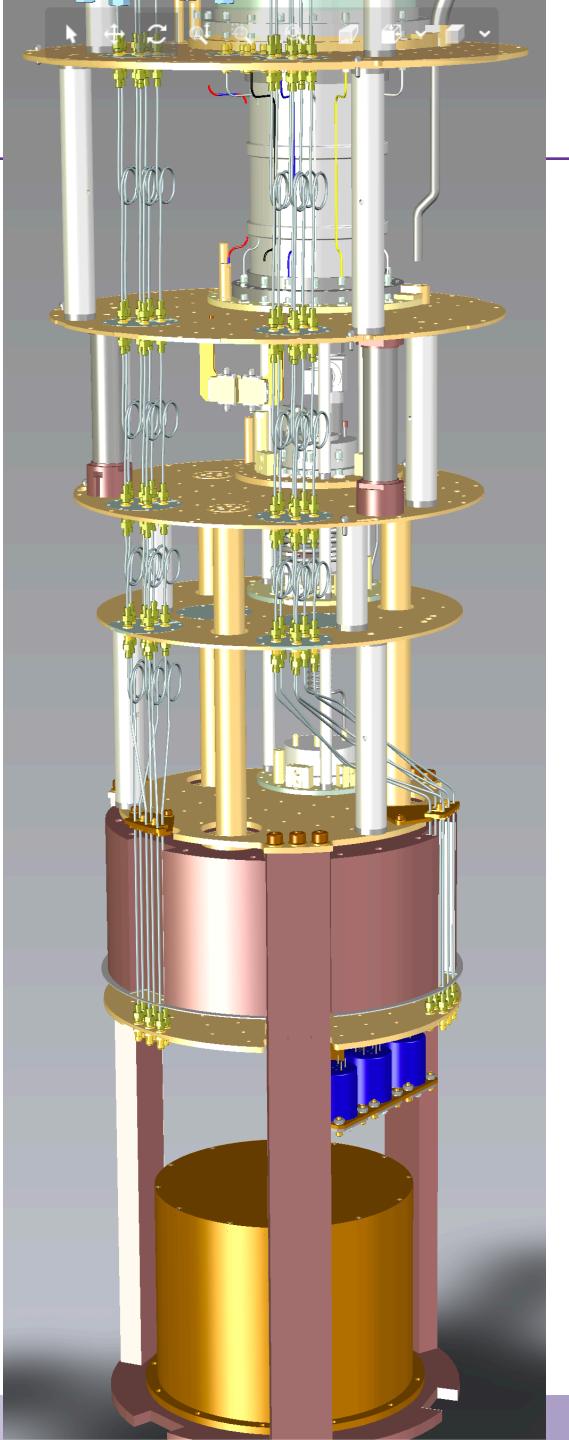






Experimental Space

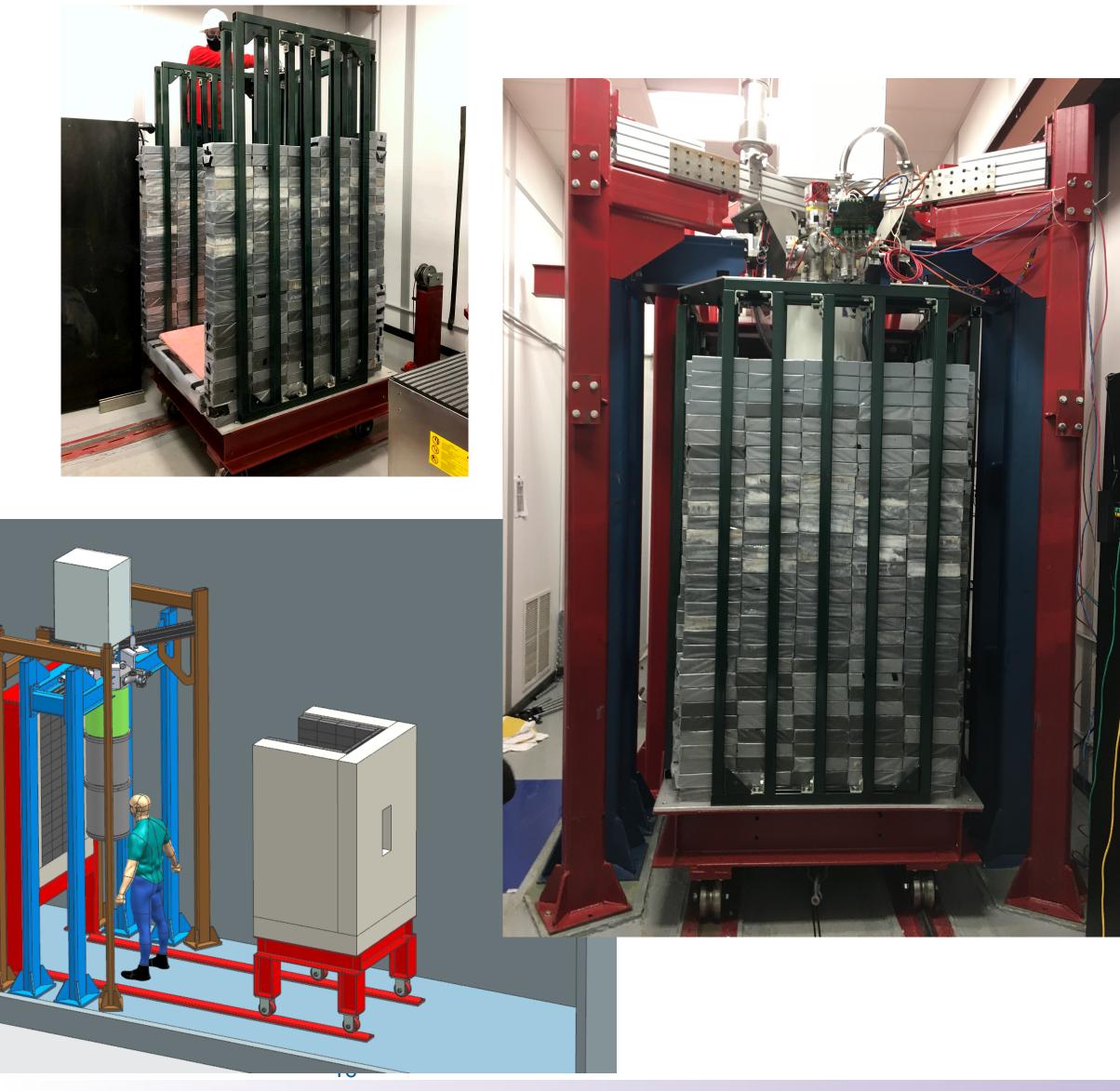
- Experimental space: 500 x 300 mm
- Readout: 8 DC SQUID channels (expanding to 12)
- 6 4-wire measurement channels
- 10 RF channels being installed right now!



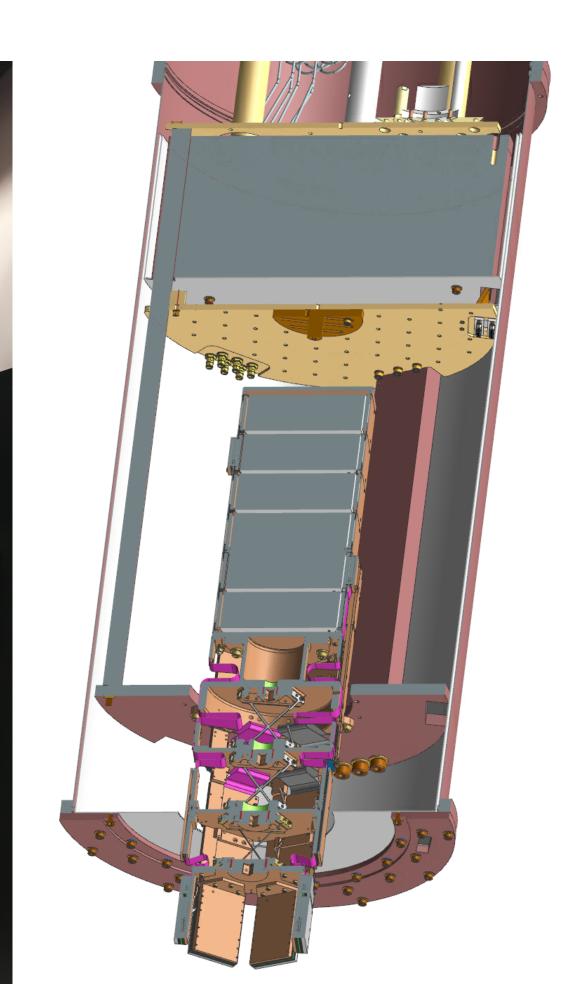


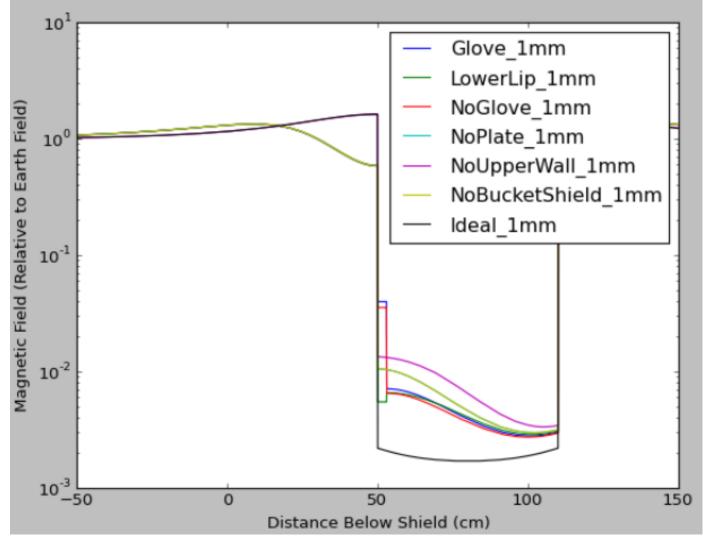


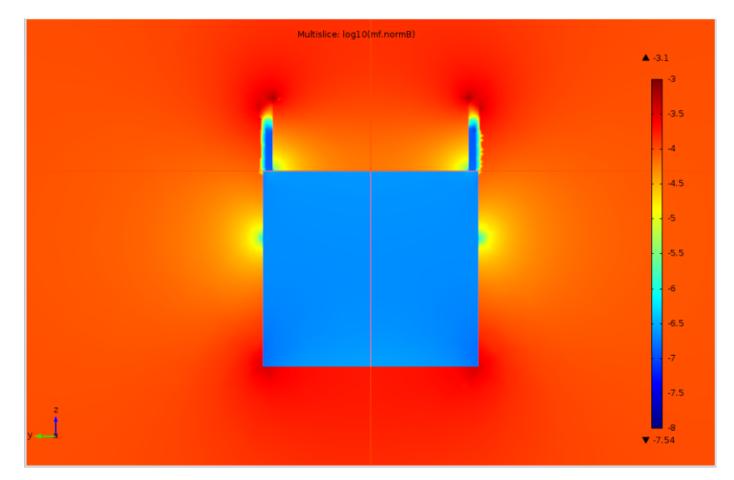
NEXUS Gamma and Magnetic Shielding



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Northwestern



Other Particle Dark Matter Experiments Using Cryogenics

- EDELWEISS (Ge) / CRESST (CaWO+)
- SPICE (Polar Crystals) / HeRALD (Superfluid He)
- Cryogenic needs:

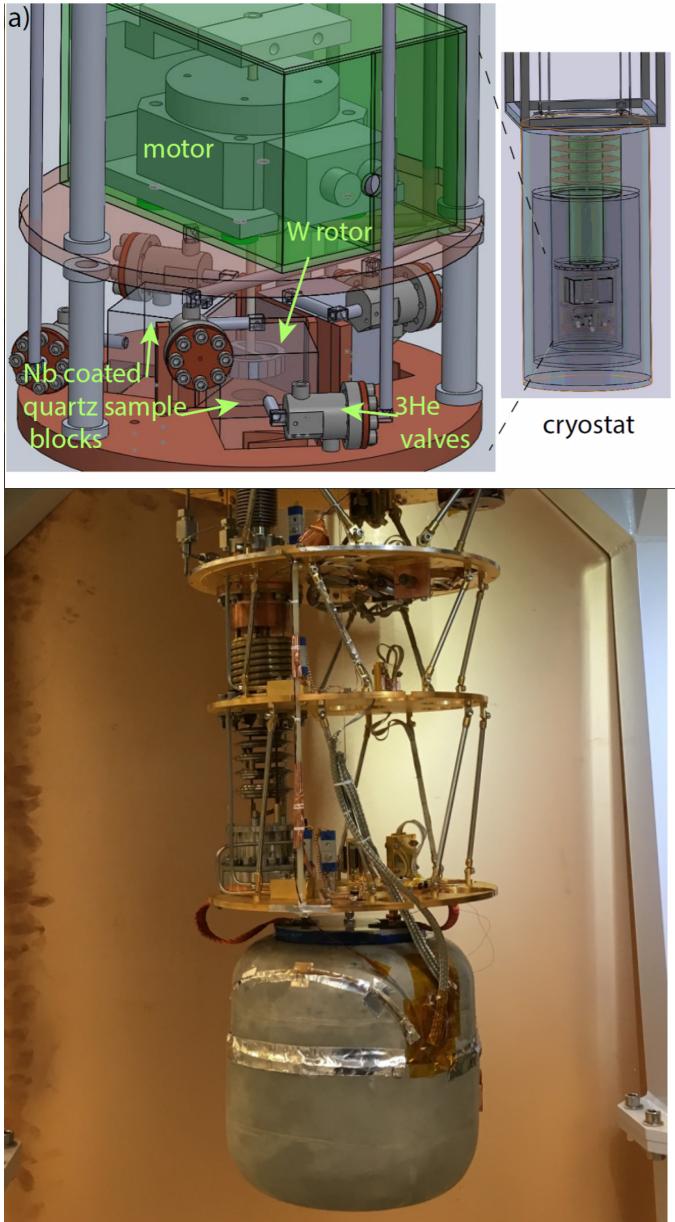
 - Vibration Isolation from Dry Dilution Refrigerators is key!
 - RF shielding important as well

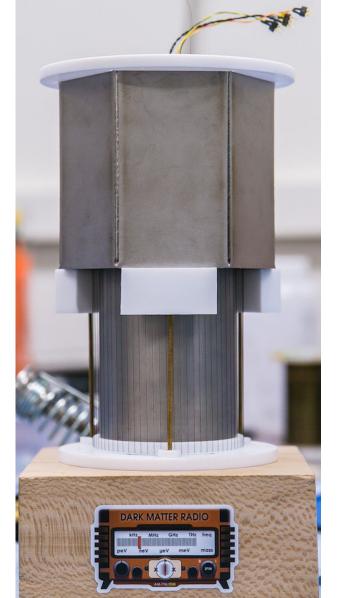
Volume and cooling power are adequate with current existing designs.



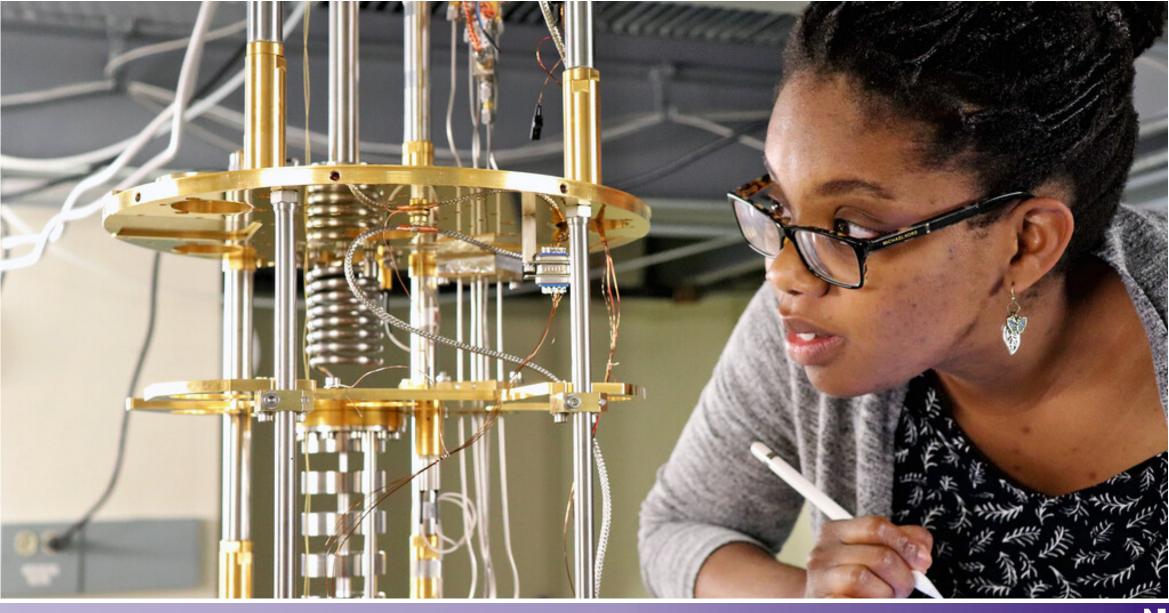
Wave-like Dark Matter Searches using Cryogenics

- ADMX / HAYSTAC
- DM Radio
- ABRACADABRA
- CASPEr
- ARIADNE
- Qubit searches
- Others...













Wave-like Dark Matter Searches using Cryogenics

- ADMX / HAYSTAC
- DM Radio
- ABRACADABRA
- CASPEr
- ARIADNE
- Qubit searches
- Others...

- Need to build expertise and best practices in integrating pulse tubes, fast cooling, and dilution units with custom thermal shields and vibration isolation.
- Need to build expertise in custom high field large volume superconducting magnets.
- Facility contribution could be millikelvin cryogenic engineering expertise at the labs.
- Perhaps a very large volume magnet facility that multiple axion experiments could utilize?





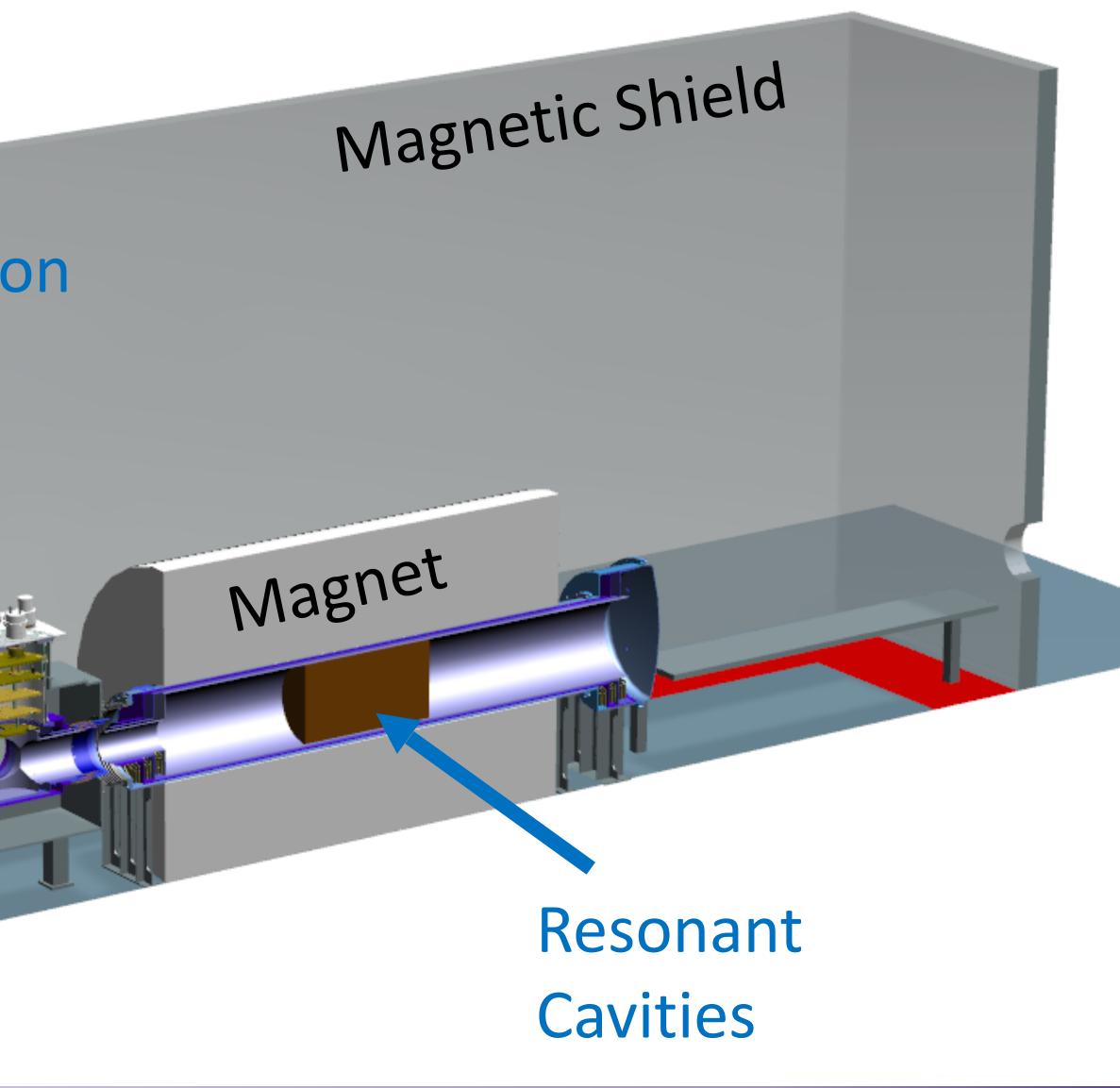




Electronics dilution fridge

Cavity dilution fridge

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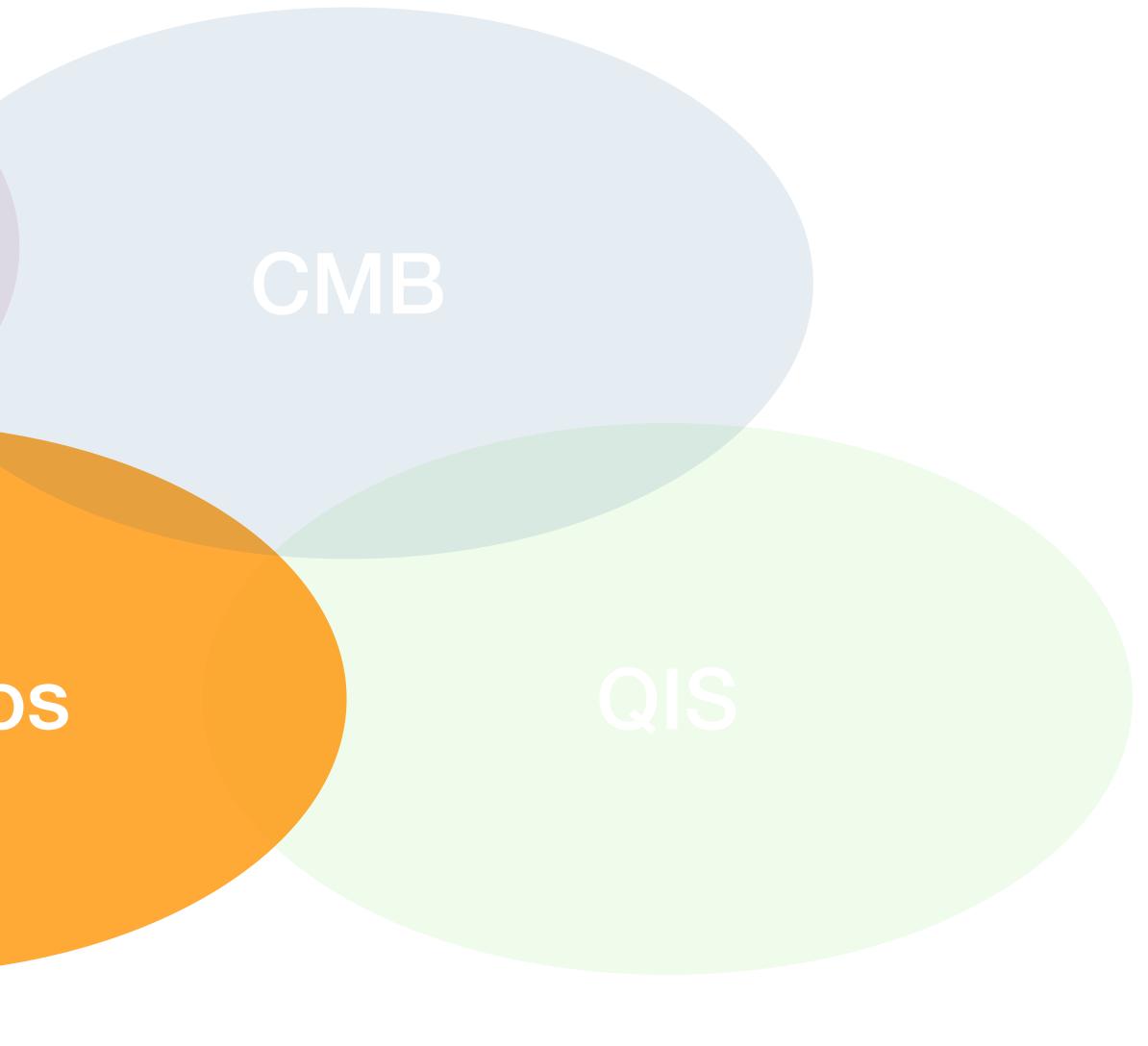


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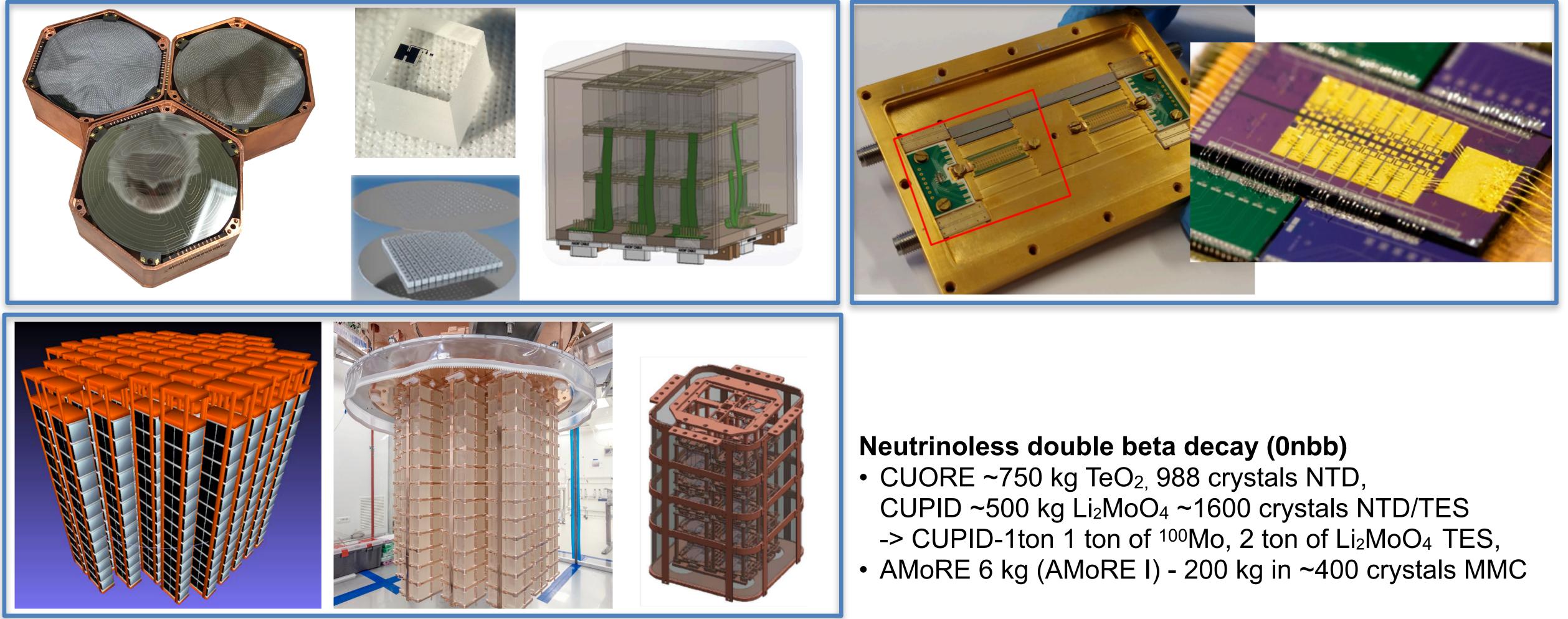




Neutrino Experiments using Cryogenics

CEvNS

- MINER 100-g scale Ge detectors
- NUCLEUS 1-g scale absorbers $(CaWO_4, Al_2O_3 - TES)$ 10 g total,
- Ricochet 40-g scale absorbers (Ge - NTD/TES, Zn with TES readout chip) ~1 kg at ILL reactor 2023



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Neutrino Mass

- HOLMES 300 Bq Ho/Au absorber on array of microMUX TESs
- ECHo-1k 1 kBq Ho implanted in a microMUX array of Au/Er Magnetic microCalorimeters (MMCs)

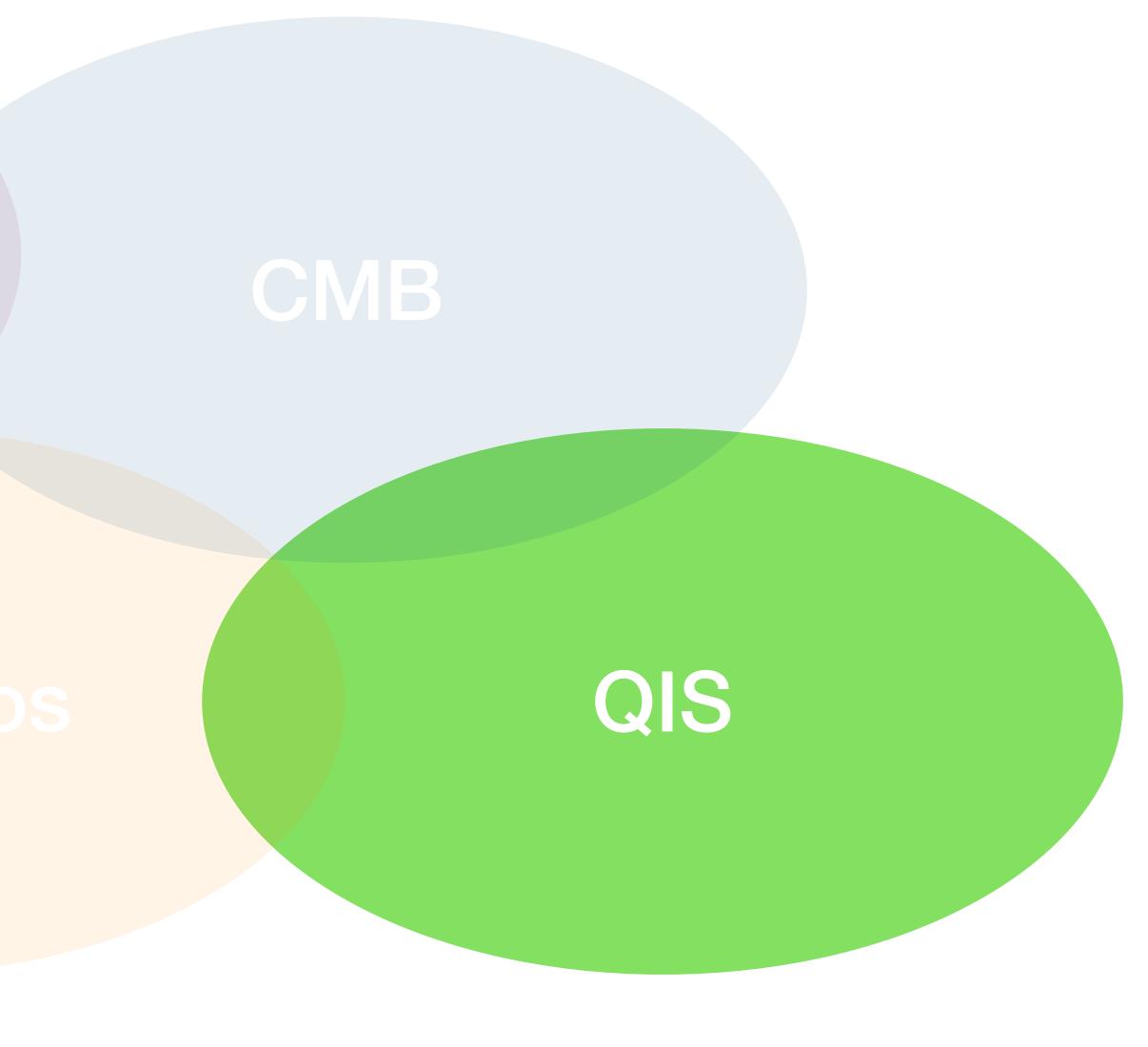


Cryogenics are Everywhere!

Dark Matter

Neutrino

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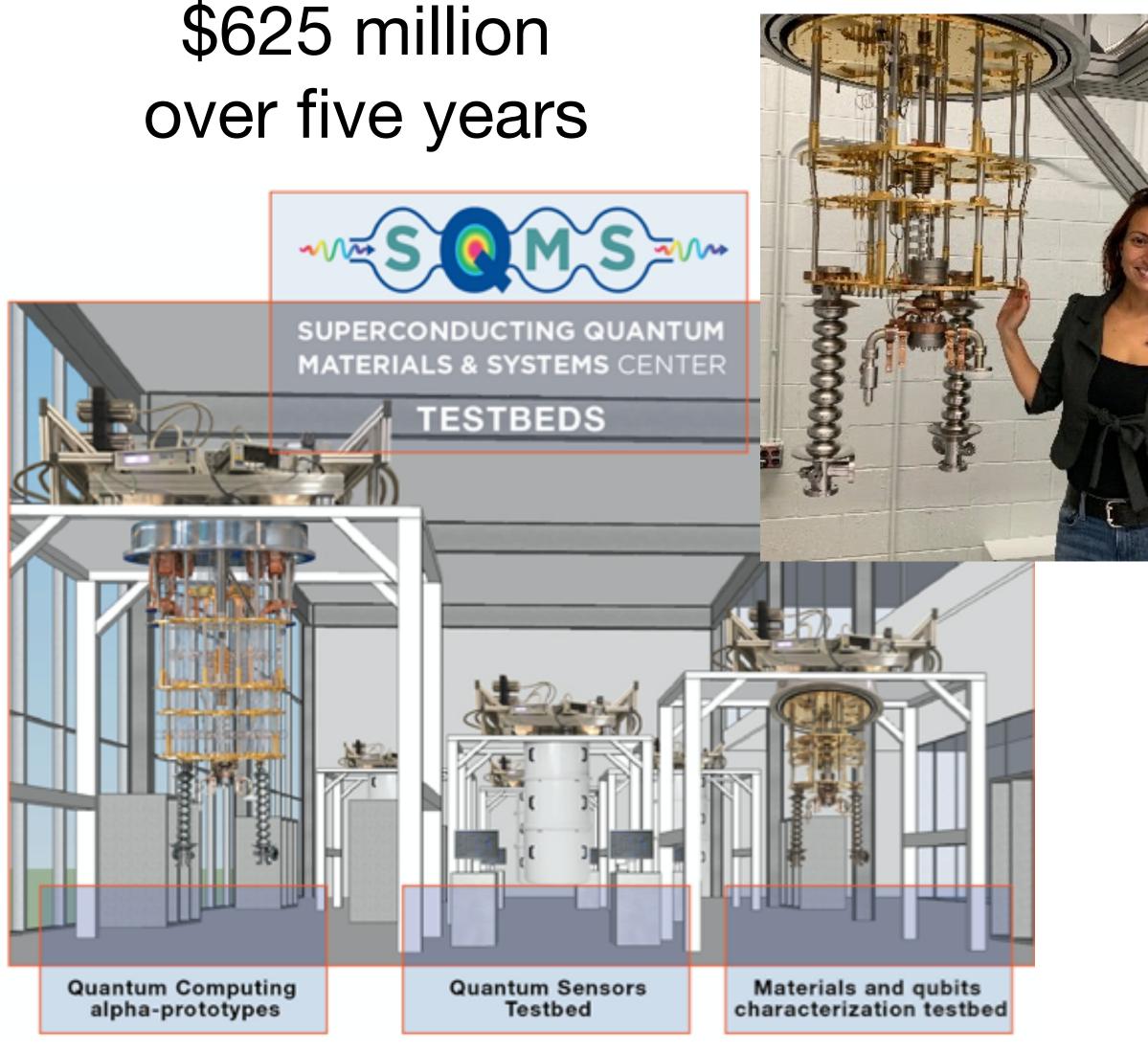




QIS Facilities



Catalyzing the Quantum Ecosystem



SQMS @ FNAL, QSA @ LBNL, C²QA @ Brookhaven, Q-NEXT @ Argonne, QSC @ Oakridge

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QUANTUM SYSTEMS ACCELERATOR















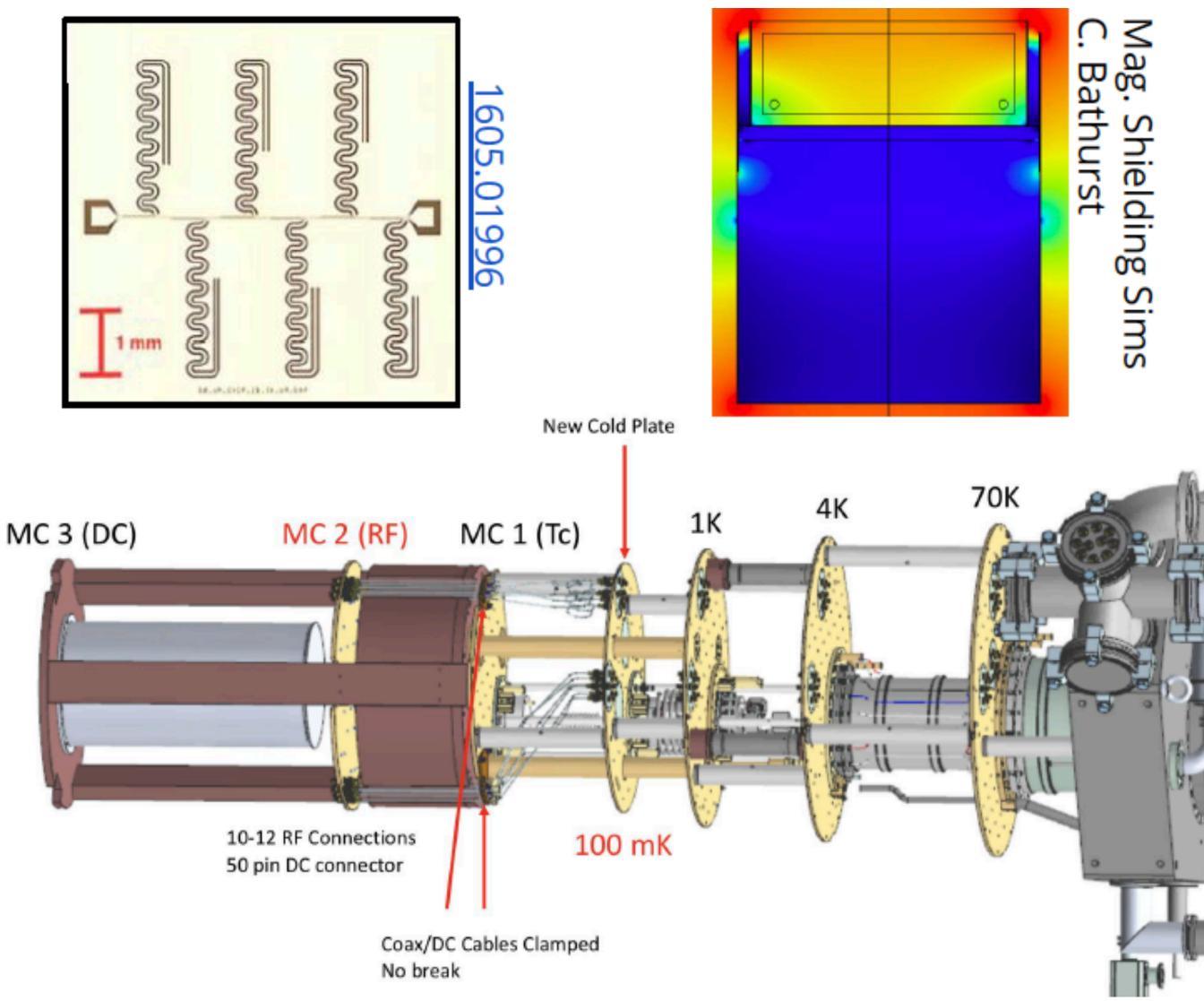


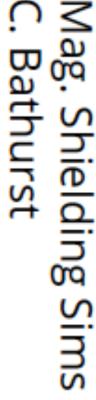
Quantum Information Science

- Collaboration with R. McDermott (UW-Madison), D. Bowring (FNAL), et al.
 - Study Quasiparticle Poisoning in Ο Superconducting Microwave **Resonators** (<u>1610.09351</u>)
- Quantum coherence improved underground (<u>2005.02286</u>)
- Will use NEXUS to study coherence time in low background environments
- Fridge upgrades (2020 FNAL LDRD, Bowring DOE ECA)
 - Superconducting coax wiring for RF Ο signals
 - Additional MC plate Ο
 - Magnetic shielding (@ 1K stage) Ο
 - Improved light and EMI shielding
 - Vibration reduction Ο











Cryogenics are Everywhere!

Dark Matter

Neutring

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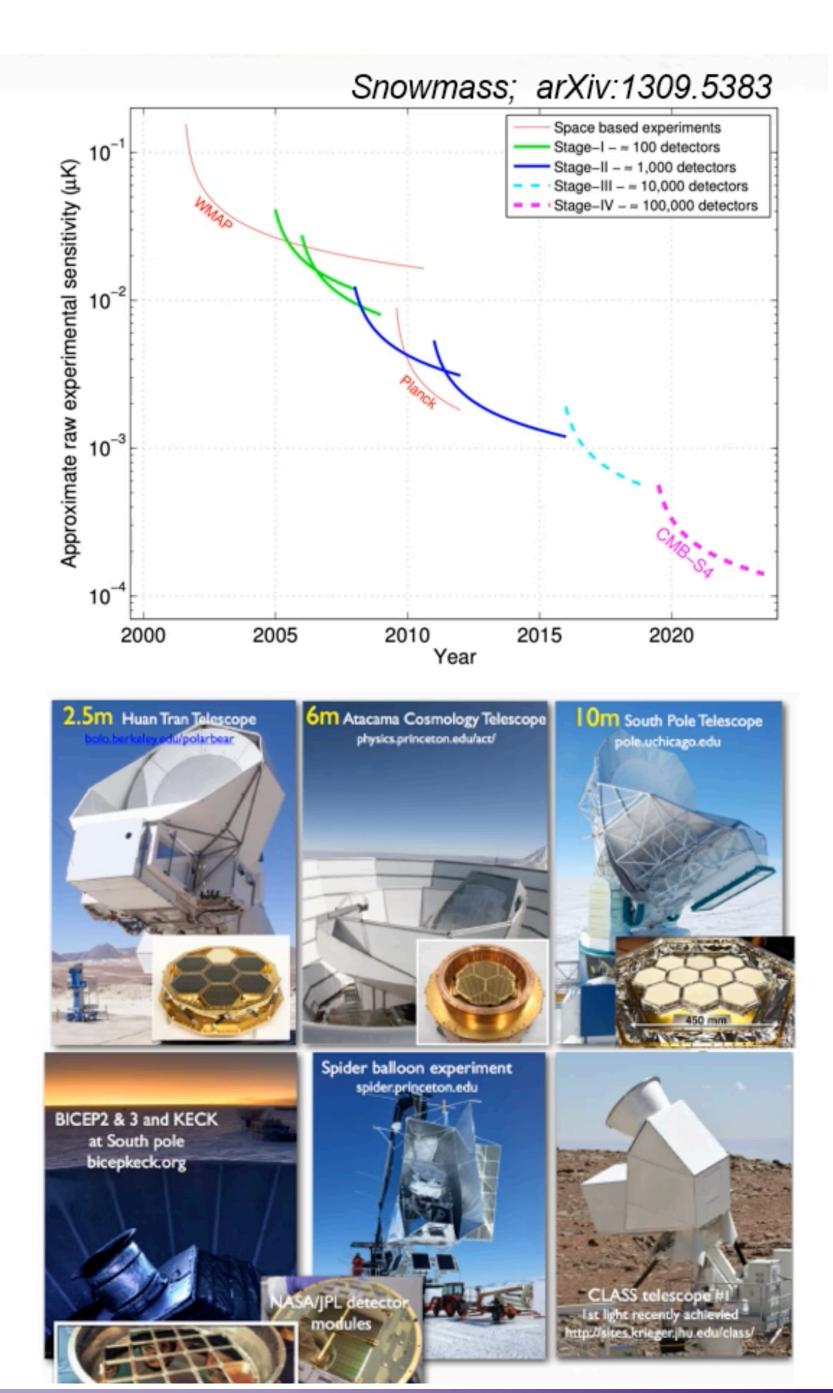




CMB-S4 scales up **proven** existing technology to unprecedented levels.

- Multiple cameras and telescopes distributed across ٠ two sites
 - each site has unique advantages and logistical ٠ challenges
- High-throughput telescope optical designs & elements
 - includes cryogenic optics
- 500,000 superconducting detectors operated at a • temperature of 100 mK distributed across all the receivers
 - multiplexed readout ٠
- Large cosmological simulations for improved ٠ theoretical modeling
- Computational methods for analysis of the resulting • data



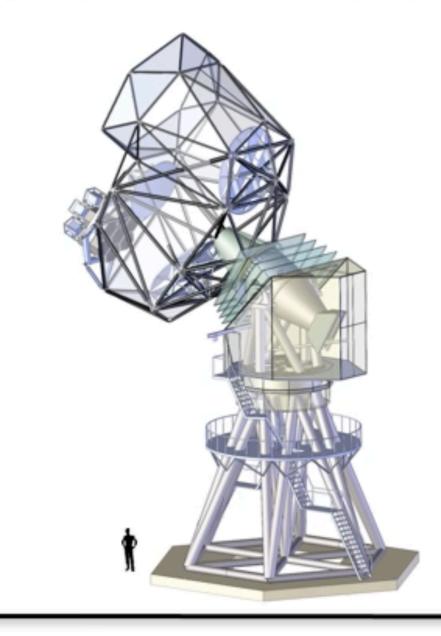




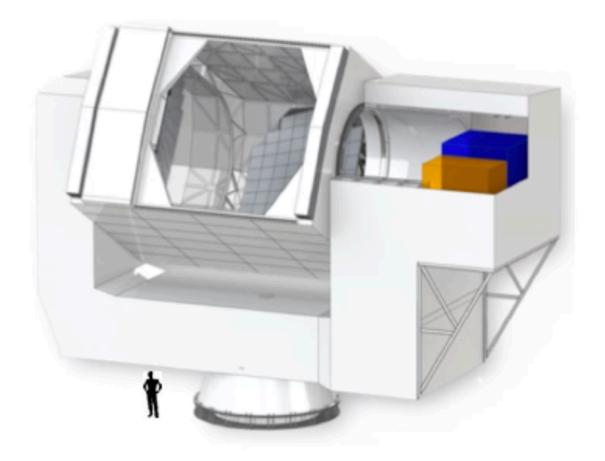
Large Aperture Telescope (LAT)

CMB-S4

Ultra-deep survey: observe ~3% of the sky with 150,000 detectors in SATs & a de-lensing LAT with 120,000 detectors.



Deep-wide survey: Two LATs observing ~60% of the sky with 240,000 detectors.

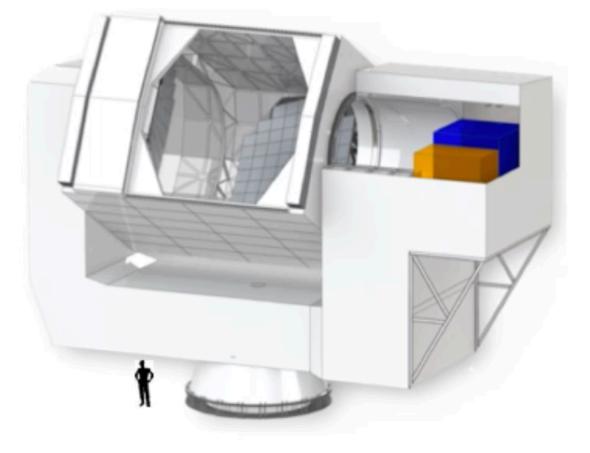


6m large telescopes, LAT e.g., like Simons Obs.

Small Aperture Telescopes (SATs)



18 x 0.55m small telescopes (3 per cryostat), e.g., like BICEP Array



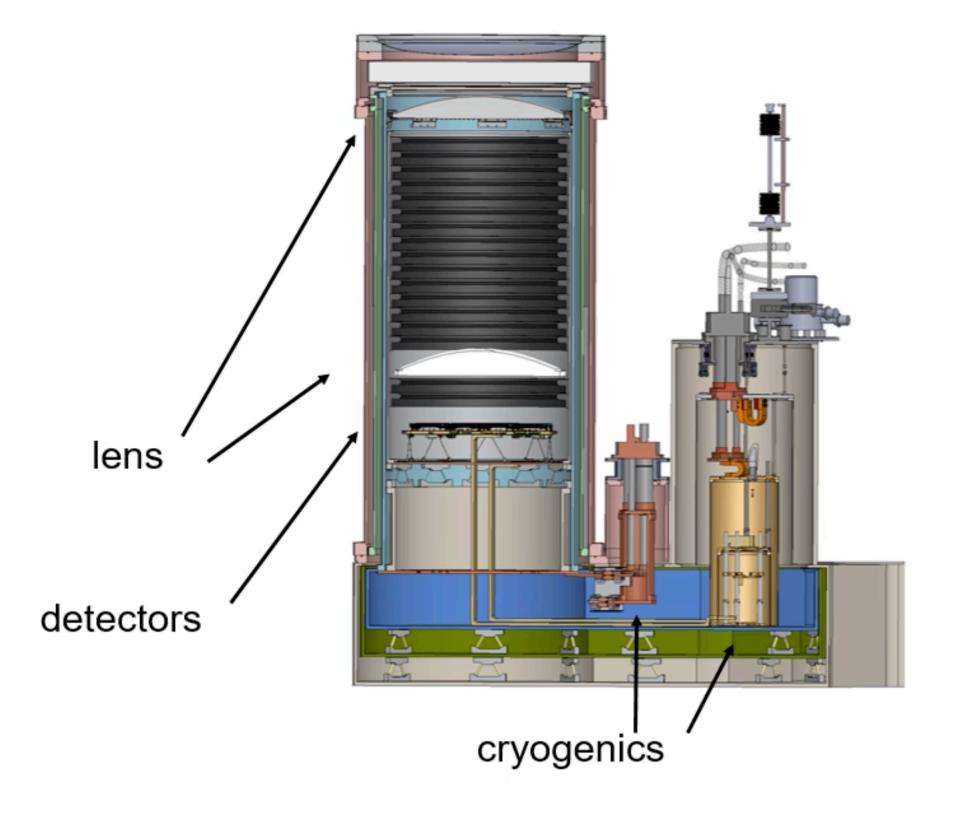
LAT

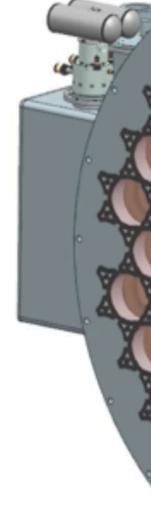






Small Area Telescope





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Large Area Telescope Detector

85 optics tubes each with a single detector wafer

2.5m



Conclusions

- each experiment there is no one-size-fits-all approach.
- future as DM/neutrino/QIS all-purpose facilities, but capability and might be useful.
- The large number of experiments using cryogenics, superconducting by the national labs to provide expertise in these areas.

Cryogenics Facilities necessarily carry over the design requirements of

 CUTE and NEXUS are currently open for business and are planning their throughput will be limited. Other facilities (like a large-magnet axion lab)

magnets, multiplexed readouts, and RF circuits provides an opportunity



