

# Comagnetometers and Accelerometers

William Terrano, March 29 2021



# Hugely-coherent sensors Testing new physics with >10<sup>20</sup> coherent particles

- Evade Standard Quantum Limit with very large particle numbers:

   suppress quantum fluctuations
   increases the signal size linearly if the interaction is coherent
- Comagnetometers and torsion pendula take advantage of this:
   set the best limits on: modifications to gravity,
   5th forces and ultra-low-mass dark matter;



macroscopic interactions

# **Torsion Pendulums Physics reach of scalars**

- Fundamental Incomapatibility between GR and Standard Model
- Extremely feeble new scalar particles are widely predicted
- Long-range non-gravitational interactions of Dark Matter
- Ultra-low-mass scalar dark matter direct detection
- Short-distance modifications to gravity:
  - Hierarchy Problem
  - Dark Energy 85 µm
  - Chameleon Fields non-renormalizable below ~20 μm



$$V(r) = -G\frac{m_1 m_2}{r} \left(1 + \tilde{\alpha} \left[\frac{\tilde{q}}{\tilde{g}\mu}\right]_1 \left[\frac{\tilde{q}}{\tilde{g}\mu}\right]_2\right]$$



# **Torsion Pendulums** Long-Range





- Best limits on EP violation of dark matter
  - Direct interaction
  - Long-range interaction
- Best limits on EP violation at distances < 100 km
- Within ~10 of recent satellite-missions, but much greater detection potential





# **Torsion Pendulums** Future of Long-Range

- Lower noise fibers (what is the limit)?
- Systematics improvements:
  - Improve thermal stability and modeling — systematic limit for MICROSCOPE as well
  - Improve gravitational stability and modeling



# **Torsion Pendulums Short-range gravity**







pendulum-attractor separation (µm)

- Tests gravity at 40 microns
- Goal 25 microns

Needs: Reduced vibrations, perfectly flat surfaces, lower patch-effect surfaces





# Comagnetometry **Basic Principle**

- Measure energy difference between spin-up and spin-down nuclei
- Noise limit:  $\sigma_{\rm f}^2 \ge \frac{12}{(2\pi)^2 (A/\rho)^2 T^3} C$
- High spin-densities possible O(50%) at atmospheric pressure [SEOP]
- magnetometer read-out is very low noise
- coherence times of many hours -> days



## **Basic Spin-Exchange Apparatus**



# **Comagnetometry** Best sensitivity in energy units

 Magnetic field drifts orders-of-magnitude larger than signal of interest: compare two spin-species

 $\omega_{\rm inv} = \omega_{\rm Xe} - \omega_{\rm He} (\gamma_{\rm Xe} / \gamma_{\rm He})$ 

 Many implementations: spin-species, read-out, geometry





## DM Signal: Energy-splitting of Spin along relative velocity direction

# $a = a_0 \cos \omega_{\rm C} t$

 $\mathcal{L} = (\partial_{\mu}a)\bar{\psi}\gamma^{\mu}\gamma_{5}\psi$ 



a at t = 0

### Magnitude ~ axion velocity in lab frame

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## Doubly-modulated $H_{\rm ax} \sim \sqrt{2\rho_{\rm DM}} \vec{v} \cdot \vec{\sigma}_{\psi} \cos m_{\rm a} t.$





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# DM Signal: Oscillating Torque on Spin $a = a_0 \cos \omega_{\rm C} t$ $\mathcal{L} = (\partial_{\mu} a) \bar{\psi} \gamma^{\mu} \gamma_5 \psi$ $H_{\rm ax} \sim \sqrt{2\rho_{\rm DM}} \vec{v} \cdot \vec{\sigma}_{\psi} \cos m_{\rm a} t.$



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 $H_{\rm ax} \sim 10^{-25} \, {\rm eV} \, \left( \frac{g_{a\bar{\psi}\psi}}{10^{-10} \, {\rm GeV}^{-1}} \right) \left( \frac{v}{10^{-3}} \right) \left( \sqrt{\frac{\rho_{\rm DM}}{(0.04 \, {\rm eV})^4}} \right) \cos m_{\rm a} t$ 

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# **Comagnetometry** DM limits



# **Rb-interaction free system** SQUID readout (more expensive, more complicated)



- Must control nuclear self-interactions (collisional and geometric couplings)
  - Controlled cell geometry
  - Decouple self-interaction Hamiltonian
  - Precise quantum state initialization

 Maintain or monitor magnetic field direction (Earths rotation effect)



## **Commercial Gear**

Density: 1 amagat Polarization: 50% *Cell size: 3 x 3.285 cm* SQUID noise: 1 fT SQUID distance: 2.5 cm





## 2nd Gen

T2 5hrs: Xenon-dimer

50% polarization, SQUID to 200 aT (5cm cell) **Beyond** 

Decay time T2 limited by Xenon-dimers

- Run hotter (T2 x10?) or
- Higher pressure (x20) or
- decoupling procedures

SQUID noise limited by pickup-loop *inductance:* 

- Custom SQUID coil (x7) &
- Better coupling (x5)

Big hammer:

- 50 cm cell (x30), gradients at nEDM levels, low magnetic field noise needed



# **Further Impact of High-Performance Comagnetometer** With pHz or better energy resolution

- Xenon EDM measurement in the  $10^{-32}e cm$  range? Systematics dependent
- Along with compact spin-sources: search for Goldstone bosons (pseudoscalars) beyond stellar cooling limits
- in E&M

• Using Neon (quadrupolar nucleus): Best measurements of Lorentz symmetry



# **Role of DOE and HEP in improving measurements Technical and Material Advances**

- Some technical improvements more suitable to national lab/staff scientist roles than graduate student or post-doc:
  - long time scales and incremental progress
  - need to preserve knowledge
  - need to share technical advances with multiple groups
  - 1. super High-Q silica: helps any mechanical system 2. low noise surfaces: helps all short-range gravity test 3. low magnetic noise materials: helps room-temperature magnetometer readout 4. Quantum-limited DC SQUID: comagnetometry and low-frequency axion-





# **Role of DOE and HEP in improving measurements Ultra-Stable (underground) National Facility**

• High-stability, low noise national facility O(200m deep): - along the lines of the low background facilities like Sanford - many experiments are limited by vibrations (ABRACADRABRA type, selfcompensating comagnetometers, tests of gravity) - thermal systematics are pernicious (10x lower underground - w/o stabilization) - gravitational gradient noise difficult to handle - low-frequency magnetic field variations (mHz-microHz)? Temp changes 10x lower **Gravity Gradients** (w/o stabilization!) vibrations 100x better Suppressed LISM Temperature Celsius LISM Humidity amioka /Hz<sup>1/2</sup>] Blastings Mitaka (Tokyo) Mitaka (Tokyo IGO Hanford Observatory **10**<sup>-11</sup> Displacemen 18H **10<sup>-12</sup> 1** Tempe **10<sup>-13</sup>** Quiet period at Kamioka **— TAMA Temperature** ----- Sensor noise Blasting at Kamioka **10**<sup>-14</sup> TAMA Humidity 25 9/7/01 9/14/01

0

20

100

depth [m]

120

140

160

180



Time [JST]

