

Thoughts on the optimization of the VLENF

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DISCLAIMER:

Most of the following is based on the high energy Neutrino Factory

However: many of the basic conclusions should be transferable ...

Contents

- Neutrino factory flux
- Treatment of near detectors
- Treatment of sterile neutrinos
- Systematics/energy resolution issues
- Conclusions

Neutrino factory flux

$$\frac{d^2\Gamma}{dE_{\nu_\mu} d\cos\theta} = \frac{G_F^2 m_\mu}{24\pi^3} \gamma (1 - \beta \cos\theta) E_{\nu_\mu}^2 [3m_\mu - 4\gamma E_{\nu_\mu} (1 - \beta \cos\theta)] ,$$

$$\frac{d^2\Gamma}{dE_{\nu_e} d\cos\theta} = \frac{G_F^2 m_\mu}{4\pi^3} \gamma (1 - \beta \cos\theta) E_{\nu_e}^2 [m_\mu - 2\gamma E_{\nu_e} (1 - \beta \cos\theta)] .$$

$$\gamma = E_\mu/m_\mu = 1/\sqrt{1 - \beta^2}$$

- Sometimes useful to integrate over energy:

$$\frac{d\Gamma}{d\cos\theta} = \frac{G_F^2 m_\mu^5}{384\pi^3} \frac{1}{[\gamma(1 - \beta \cos\theta)]^2}$$

[neglect beam collimation for the moment]

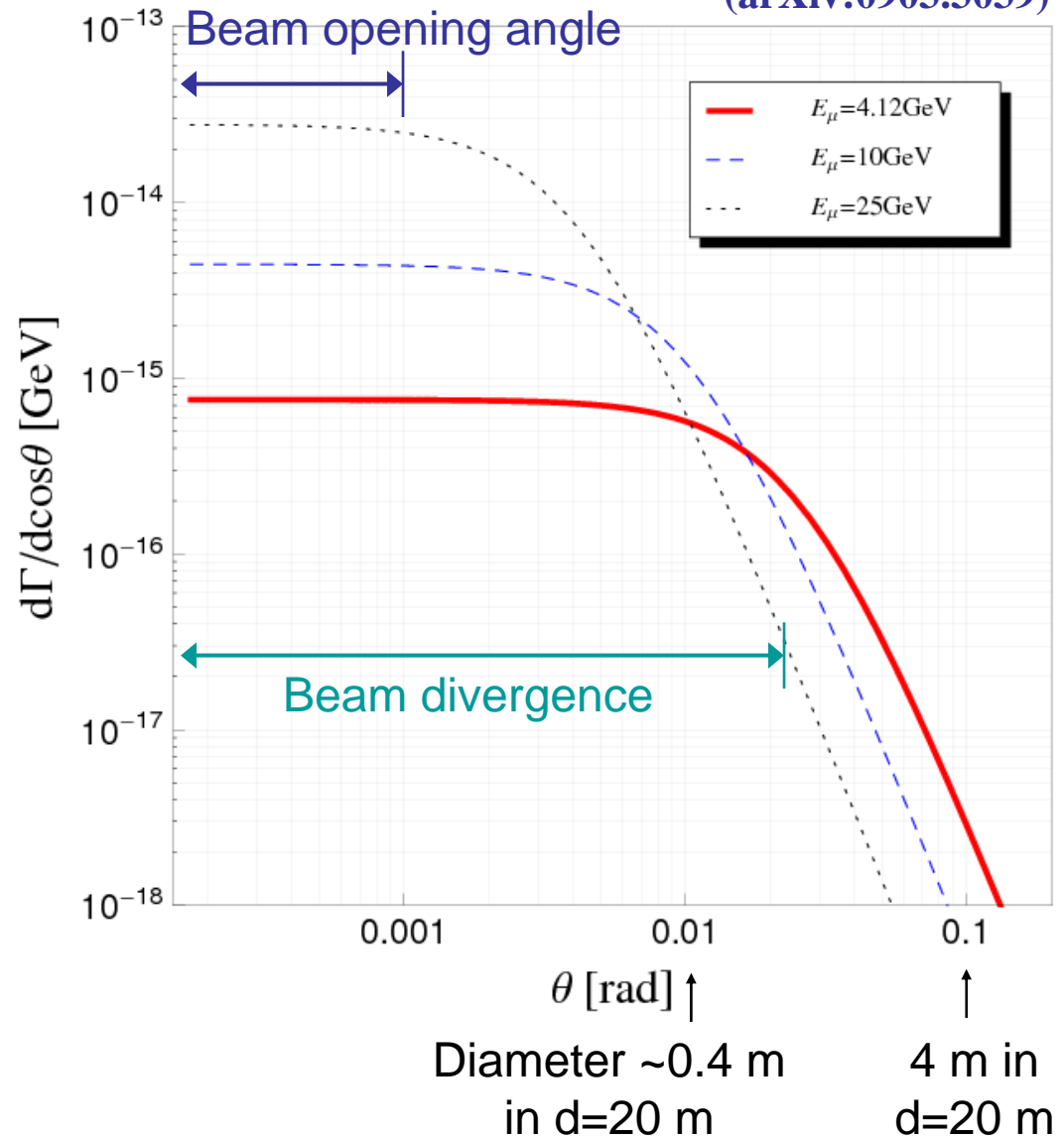
Geometry of the beam

(arXiv:0903.3039)

- Beam diameter $\sim 2 \times L \times \theta$
- We use two beam angles:
 - Beam opening angle:

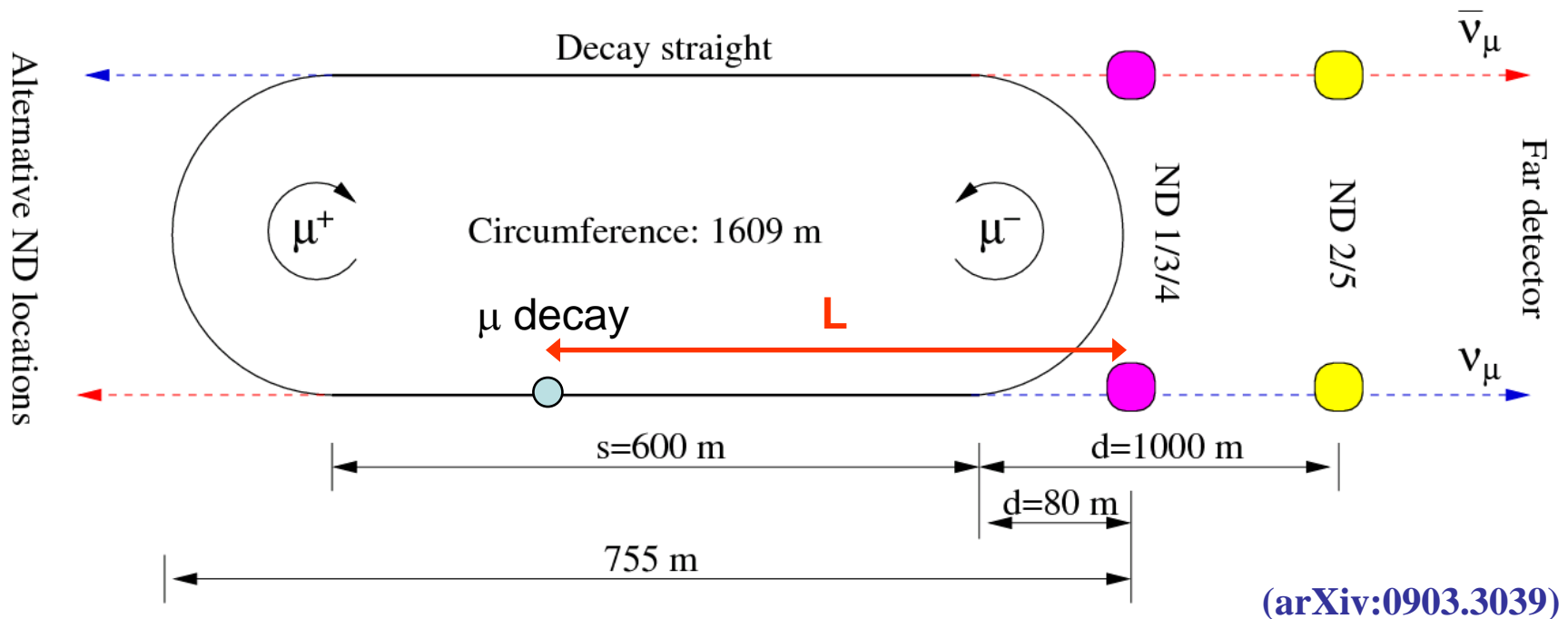
$$\left. \frac{\Gamma}{d \cos \theta} \right|_{\theta} = 0.9 \left. \frac{\Gamma}{d \cos \theta} \right|_{\theta=0}$$

- Beam divergence: contains 90% of total flux



Geometry of decay ring

- Example: high energy version



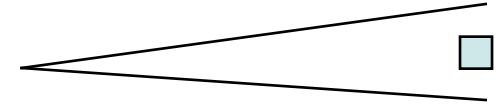
- d = distance from end of straight
- s = length of straight
- L = baseline (from decay point to detector)

Approximations

- **Far distance approximation:**

Flux in whole detector looks like on-axis flux

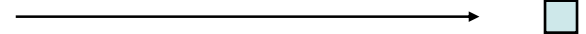
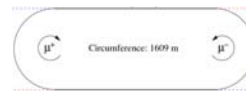
$$\frac{d^2\Gamma}{dE_{\nu_\alpha} d\cos\theta}(\cos\theta) \simeq \frac{d^2\Gamma}{dE_{\nu_\alpha} d\cos\theta} \Big|_{\theta=0}$$



- **Point source approximation:**

Extension of source can be neglected

$$d \sim L \gg s$$



Extreme cases

- **Far detector limit:**

Far distance approximation for *any* point of the decay straight, i.e., the detector diameter $D < 2 \times L \times \theta$, where θ is the **beam opening angle**

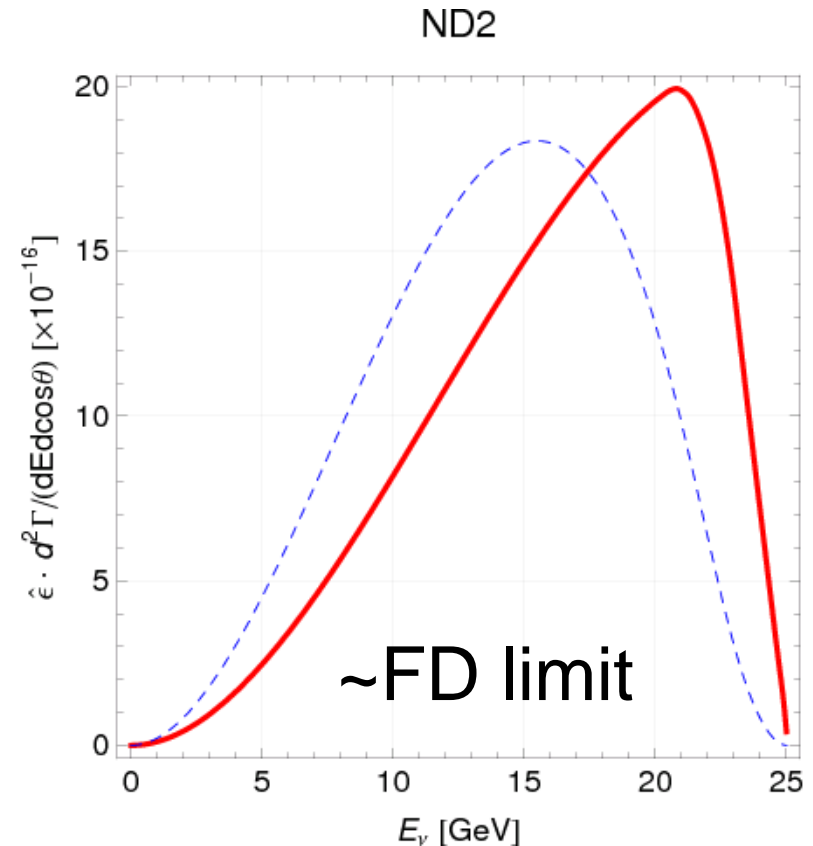
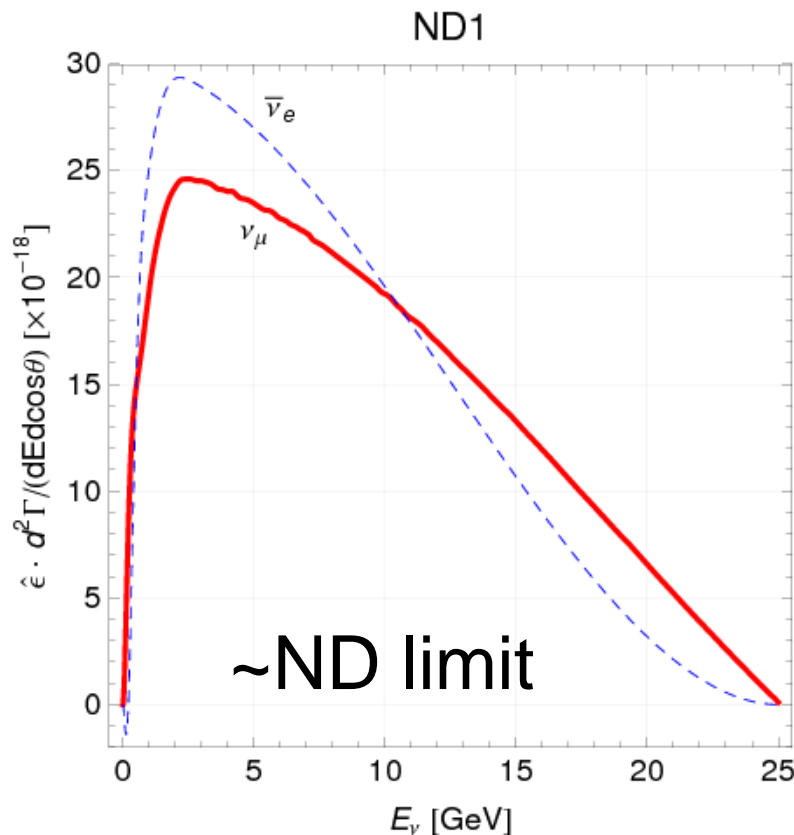
- **Near detector limit:**

The detector catches almost the whole flux for *any* point of the decay straight, i.e., the detector diameter $D > 2 \times L \times \theta$, where θ is the **beam divergence**

Extreme cases: Spectra

- Some examples (HENF):

Parameter	ND1	ND2
Diameter D	17 m	4 m
Distance d	80 m	1000 m
Mass	450 t	25 t



Some technicalities

- How to treat arbitrary detectors in GLoBES?
(which uses the point source and far distance approximations)

1) Take into account extension of detector

$$A_{\text{eff}} = \frac{2\pi L^2}{\left. \frac{d^2\Gamma}{dE d\cos\theta} \right|_{\theta=0}} \int_0^{\frac{D}{2L}} \frac{d^2\Gamma}{dE d\cos\theta} \sin\theta d\theta \quad \text{and} \quad \varepsilon(E, L) = \frac{A_{\text{eff}}}{A_{\text{Det}}}$$

2) Take into account extension of straight

$$\frac{dN_{\text{avg}}}{dE} = \frac{dN_{\text{PS}}(L_{\text{eff}}, E)}{dE} \frac{L_{\text{eff}}^2}{s} \int_d^{d+s} \frac{\varepsilon(L, E)}{L^2} dL = \frac{dN_{\text{PS}}(L_{\text{eff}}, E)}{dE} \hat{\varepsilon}(E)$$

$$\hat{\varepsilon}(E) \equiv \frac{L_{\text{eff}}^2}{s} \int_d^{d+s} \frac{\varepsilon(L, E)}{L^2} dL$$

GLoBES
built-in with

$$L_{\text{eff}} = \sqrt{d(d+s)}$$

Examples for near detectors

Near detector limit

Far detector limit

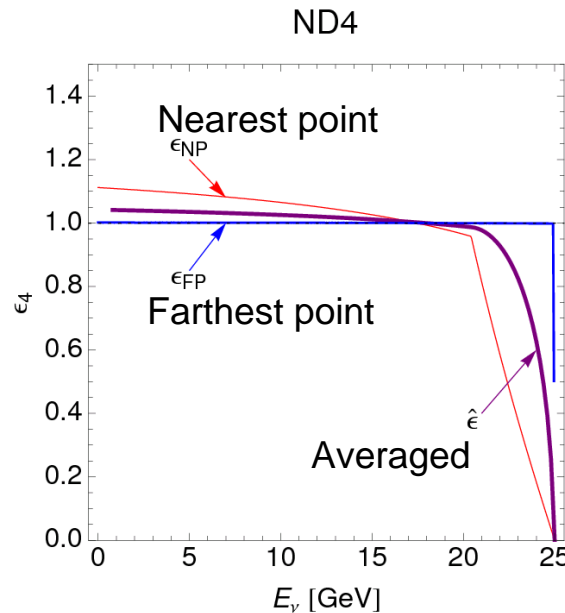
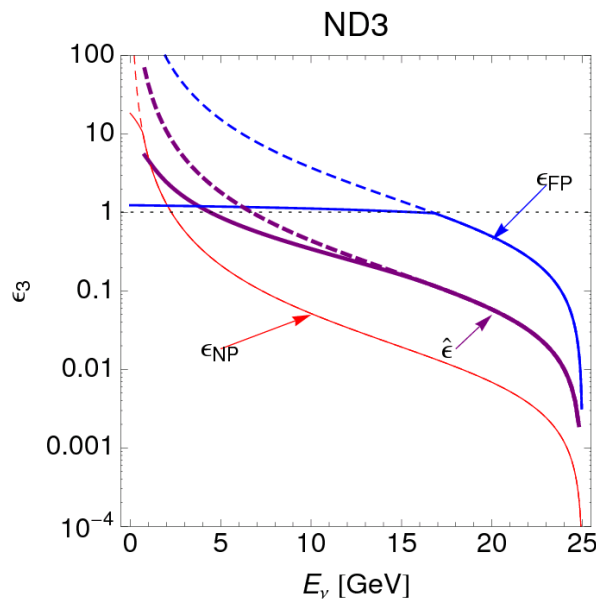
Parameter	ND1	ND2	ND3	ND4	ND5
Diameter D	17 m	4 m	4 m	0.32 m	6.8 m
Distance d	80 m	1000 m	80 m	80 m	1000 m
Mass	450 t	25 t	25 t	0.2 t	2000 t

Hypothetical

SciBar-size

Silicon-
vertex
size?

OPERA-
size



$\varepsilon=1$: FD limit
Dashed: ND limit

(Tang, Winter,
arXiv:0903.3039)

- Leads to excess of low-E events
 - near detector has to be large enough to have sufficient rates in high energy bins!
- VLENF example: 200t TASD @ 20m, 2-3 m radius: ~ qualitatively similar to ND 3
- VLENF example: 800t @ >> 600m, 6-7 m radius: ~ qualitatively similar to ND 4

Sterile neutrinos: thoughts

- First approximation (use L_{eff}):
$$\frac{\Delta m_{41}^2 L_{\text{eff}}}{4E} \sim \frac{\pi}{2}$$
- Examples (VLENF): $E \sim 1 \text{ GeV}$
 $s=100 \text{ m}, d=20 \text{ m}: L_{\text{eff}}=49 \text{ m} \Rightarrow \Delta m^2 \sim 60 \text{ eV}^2$
 $s=100 \text{ m}, d=600 \text{ m}: L_{\text{eff}} = 648 \text{ m} \Rightarrow \Delta m^2 \sim 5 \text{ eV}^2$
- The problem: are there effects from averaging over the straight?
- Oscillations depend on $x=L/E$, where

$$dx/x \sim |dL/L| + |dE/E| \sim s/L_{\text{eff}} + 0.05 \text{ (TASD)}$$

$s/L_{\text{eff}} \sim 15\%$ in far detector ($d=600\text{m}$)

\Rightarrow **Constrained by extension of straight, not energy resolution of detector!? Why need 5%?**

Treatment of steriles

- Requires generalization of ND scheme:

$$\frac{dN_{\text{avg}}}{dE} = \frac{dN_{\text{PS}}(L_{\text{eff}}, E)}{dE} \frac{L_{\text{eff}}^2}{s} \int_d^{d+s} \frac{\varepsilon(L, E)}{L^2} P_{ee}(L, E) dL = \frac{dN_{\text{PS}}(L_{\text{eff}}, E)}{dE} \hat{P}(E)$$

$$\hat{P}(E) \equiv \frac{L_{\text{eff}}^2}{s} \int_d^{d+s} \frac{\varepsilon(L, E)}{L^2} P_{ee}(L, E) dL$$

Assumption:
muon decays
per dL ~ const

Effect
of beam
geometry

Effect
of osc.
prob.

- So far only tested for $\varepsilon \sim 1$ (far detector limit); however, not in principle impossible if ε integrated in osc. engine (GLOBES)

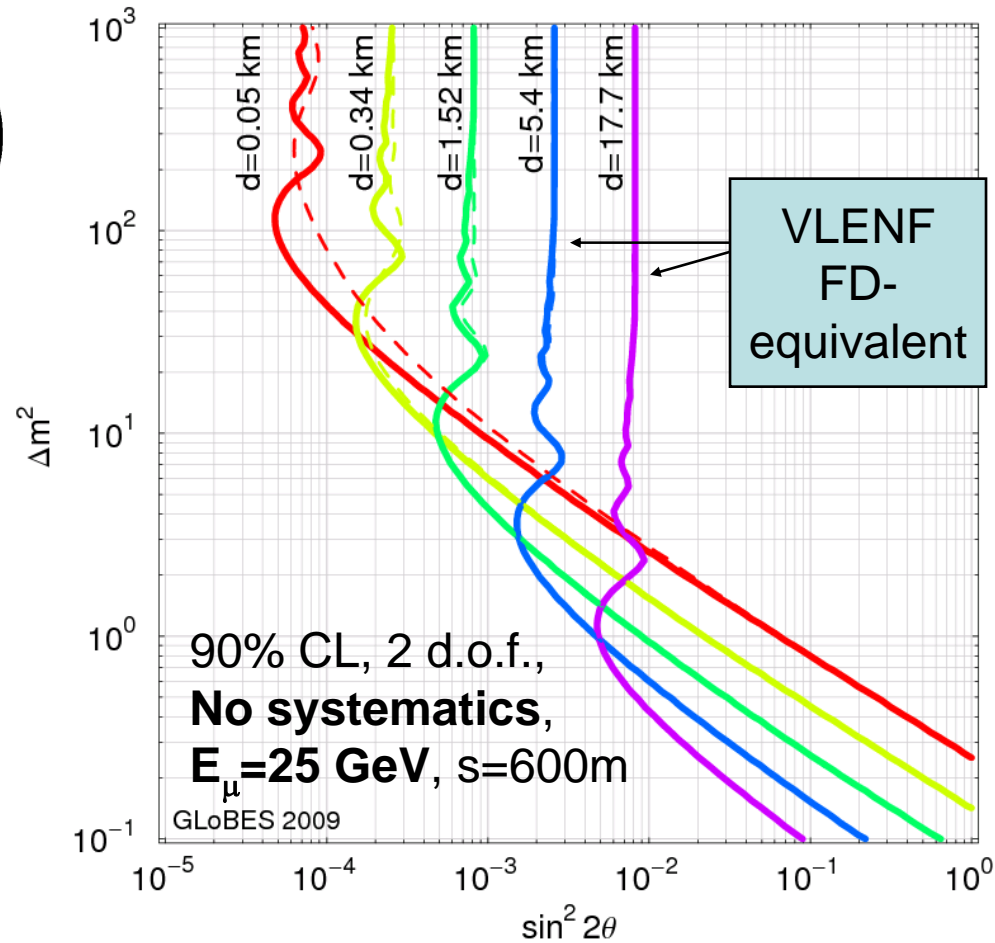
Example: ν_e disappearance

- ν_e disappearance:

$$P_{ee} = 1 - \sin^2(2\theta_\nu) \sin^2\left(\frac{\Delta m_\nu^2 L}{4E}\right)$$

- Averaging over straight important (dashed versus solid curves)
- VLENF: Expect significant averaging effects if $d < \sim s$, i.e., in near detector

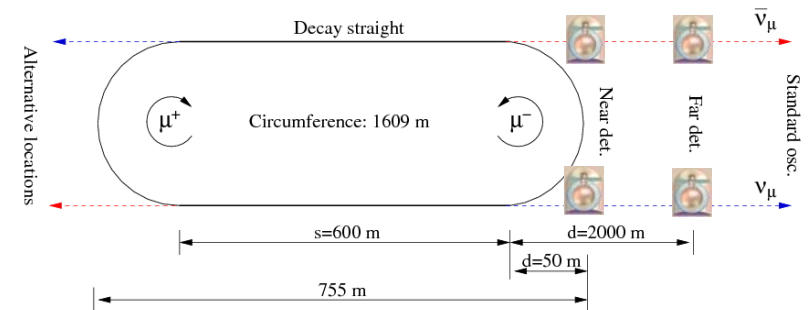
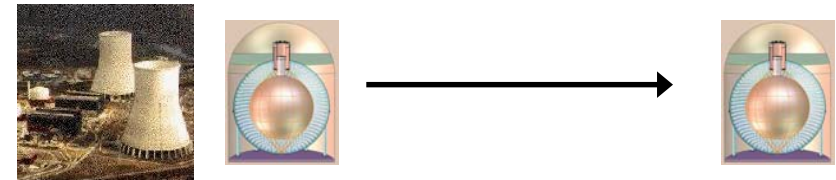
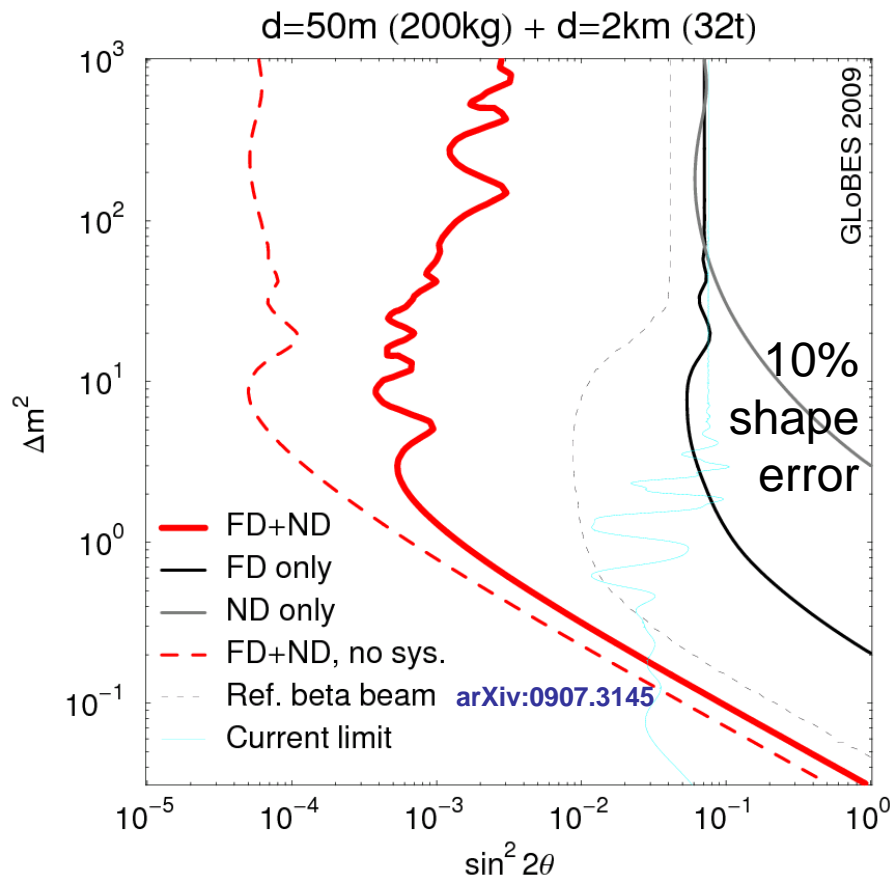
[and limitation of $x=L/E$ -resolution everywhere (see before)]



(arXiv:0907.5487)

Disappearance systematics

- Systematics similar to reactor experiments:
Use two detectors to cancel X-Sec errors



([arXiv:0907.5487](#))

On the VLENF optimization

(general conclusions)

- From the above: E_μ can be rescaled if the baselines are adjusted accordingly (e.g. if required by X-sec measurements)
- Advantage: Higher $E_\mu \Rightarrow$ longer $d \Rightarrow$ less rel. effect of averaging of the straight
- In principle: $d > \sim 2000\text{m}$ necessary if 5% energy resolution needs to be useful in FD $\Rightarrow E_\mu$ higher by a factor of three possible
- However: not so clear to me where energy resolution important ... [maybe not at first osc. maximum]

On the VLENF optimization

(technical conclusions)

- Numerical studies challenging, since optimization depends on detector geometry and straight averaging (needs some coding), but recipe clear
- But: sensitivity to sterile neutrinos will mostly depend on far detector, which can be typically approximated by far detector limit
- One has to ensure that the near detector has a sufficient event rate at all energies; it may limit the energy resolution of the system because of the decay straight averaging
- Some systematics difference between appearance and disappearance searches!

- Dedicated pheno studies should include:
 - Full N flavor framework
 - Near+far+very far detectors
 - Full Δm^2 range
- So far: only effective near detector system →

