
Dark Matter Session

Nikita Blinov



Becoming a Theoretical Physicist

I'm a post-doctoral researcher in the Theory and Astrophysics groups

I study the connections between particle physics and cosmology

Theory tools: pen and paper, computer programming

**Typical career path
for a theorist**

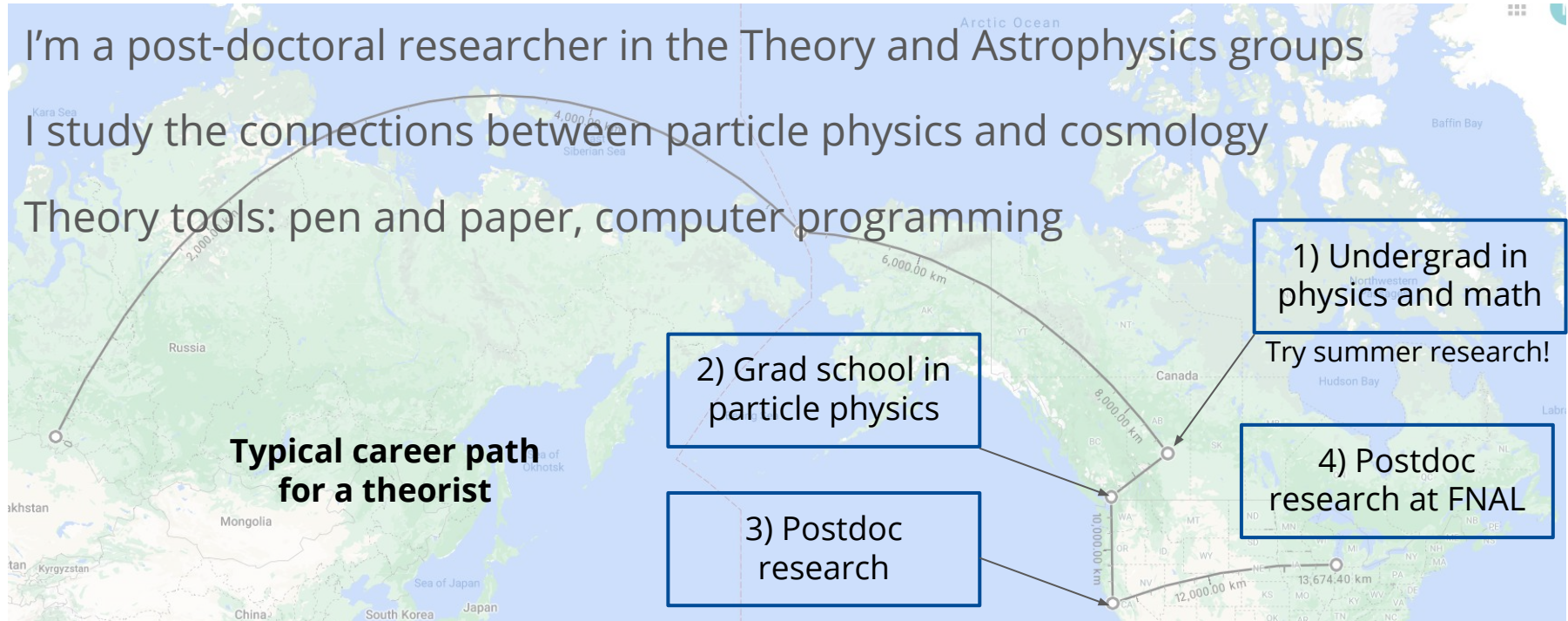
2) Grad school in
particle physics

3) Postdoc
research

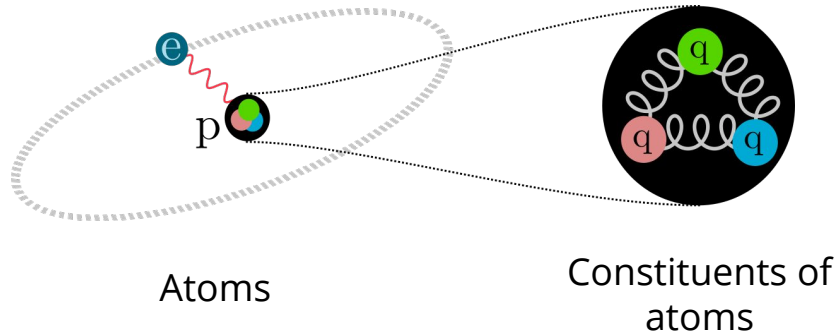
1) Undergrad in
physics and math

Try summer research!

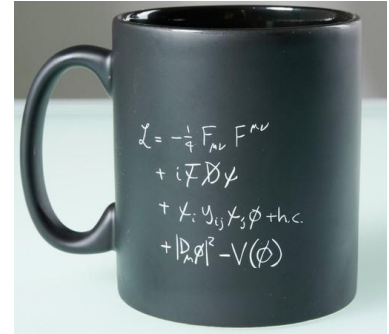
4) Postdoc
research at FNAL



Why Particle Physics and Cosmology?



100 years of science!



CERN Gift Shop

We get to think about BIG questions like

- What are the fundamental ingredients of our universe?
- What was the early universe like?
- How do we use laboratory and astro observations to learn about these?

Missing Mass in the Universe?

What are the fundamental ingredients of our universe?

Universe Facts

Standard Model

Photons	0.005%
Neutrinos	0.004%
Atoms	5%

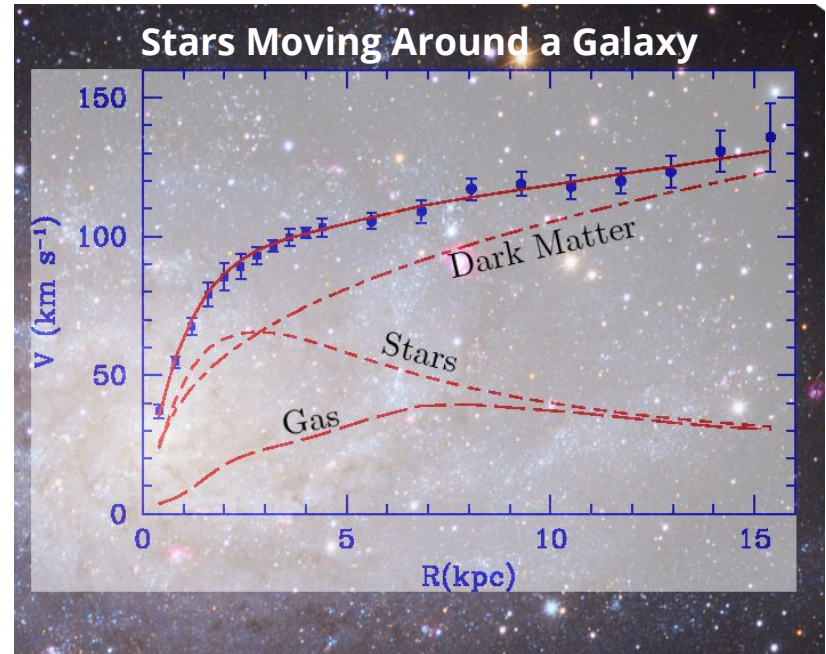
} Familiar matter

Non-Standard Model

Dark Energy	68.5%
Dark Matter	26.5%

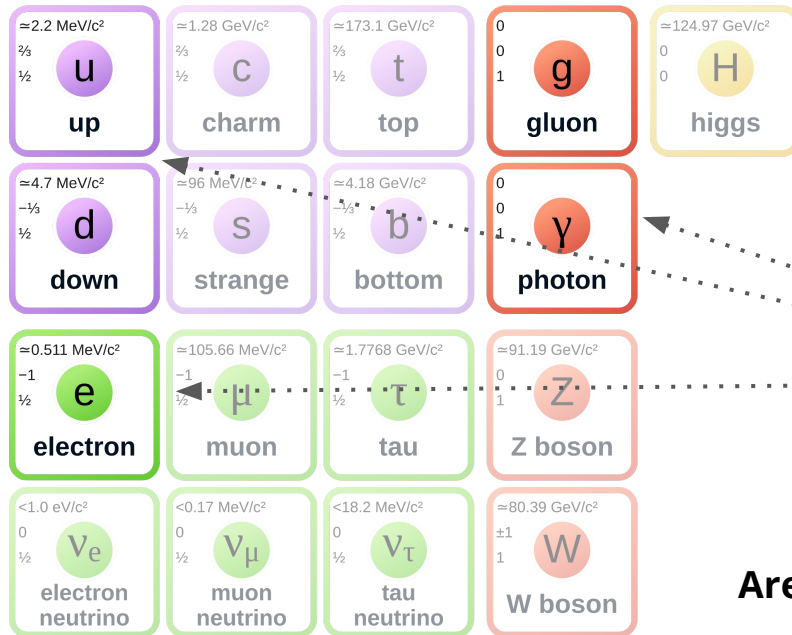
} ???

*Abundances are based on the Standard Cosmological model. They may be different in other models.



Particle Physics vs Dark Matter

Standard Model of Elementary Particles



No satisfactory particle in the Standard Model:
Need something new!

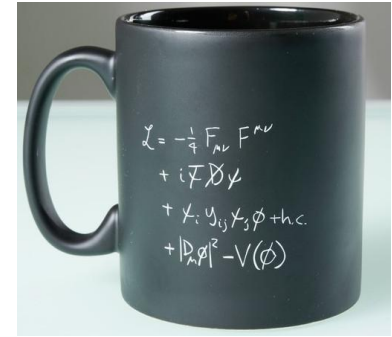
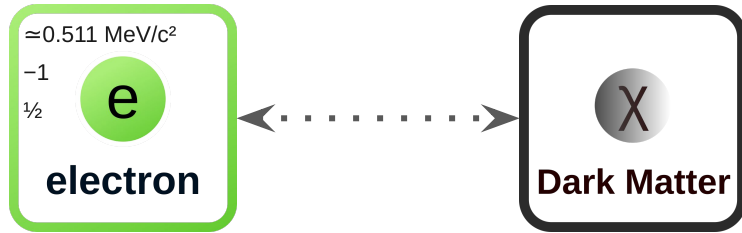


Are there any interactions with Standard Model Particles?

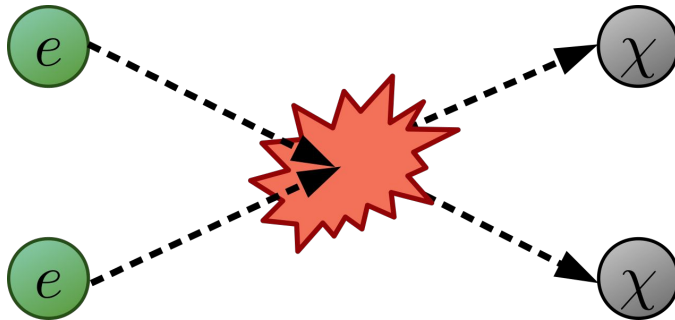
Wikipedia

How Was Dark Matter Produced?

Need to make a hypothesis, for example



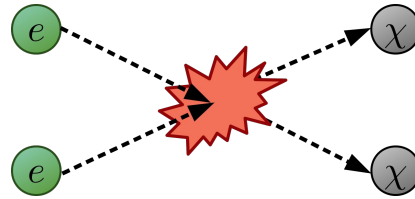
Lots of energetic electrons in the hot, dense soup of the early universe



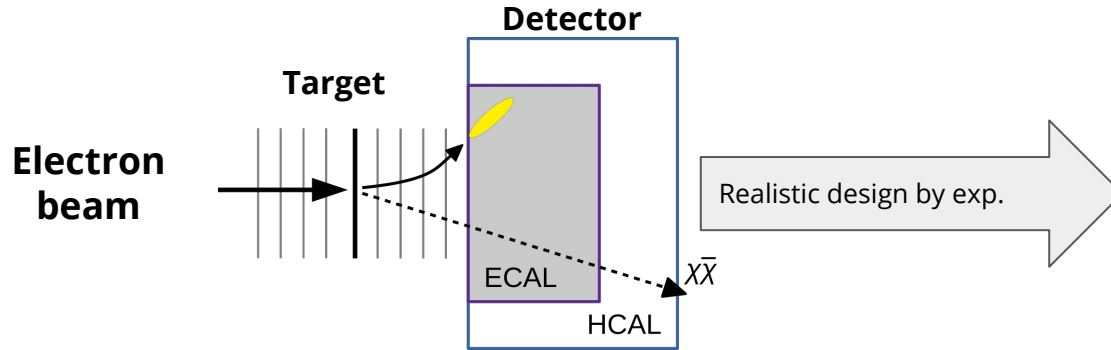
Dark Matter can be produced from collisions of familiar particles!

Searching for Dark Matter at Accelerators

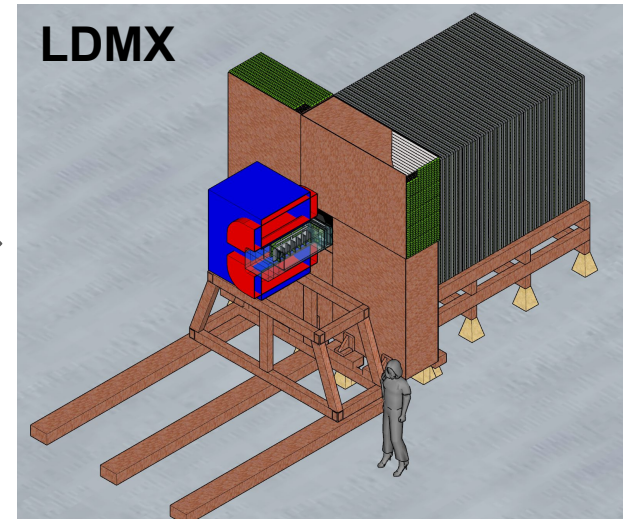
Calculations of processes in the early universe inform us what experiments we need to test a given model



Same interactions as in early universe produce Dark Matter in beam-target collisions!



Theorist cartoon of an experiment



Ana Martina Botti

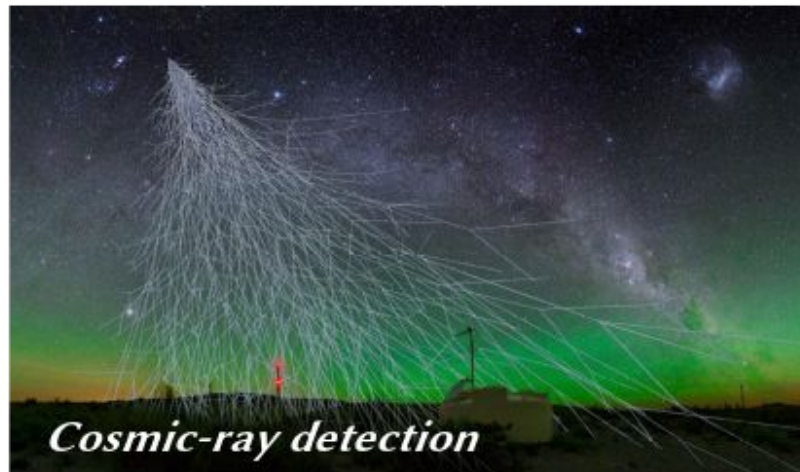


My background

1

- Physics diploma from Buenos Aires.
- Double doctoral degree in Astrophysics (Argentina/Germany).
- Post-doctoral researcher (transitioning from Buenos Aires to Fermilab).
- CPC Fellow soon.

A. Chantelauze, S. Staffi, L. Bret



But also...

~ 7 years experience in industry (software development and IT security).
I use this experience **A LOT** in my everyday work!

Why experimental physics?

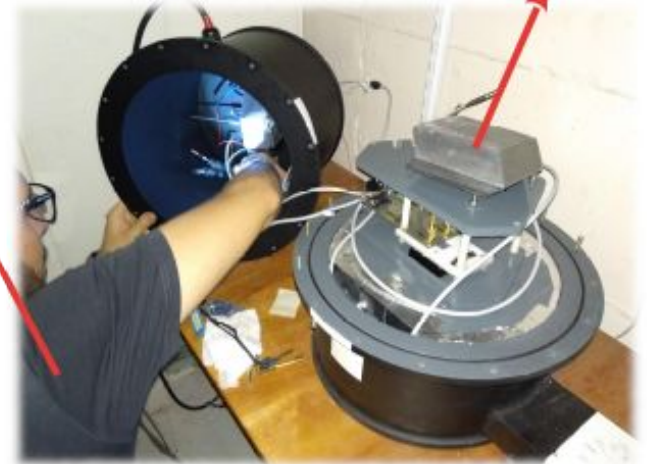


- Think, design and build experiments.
- Understand how “things” work (and why they don't!) and the Physics behind the technology.

- Contribute to answer fundamental questions about nature.
- We work a lot with engineers and technicians (and we learn a lot from them too!).

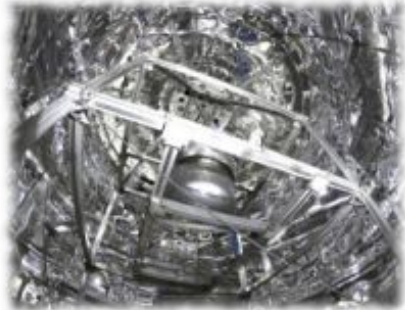
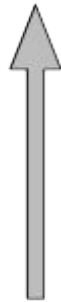
Engineer

Muon detector

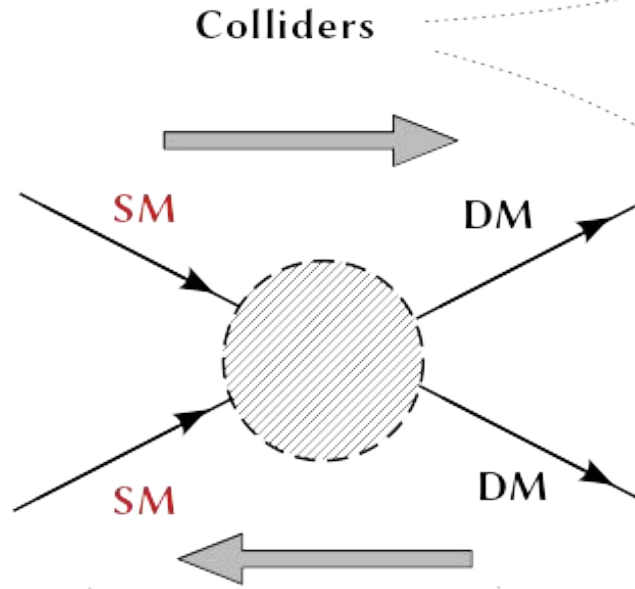


Dark Matter detection

Direct
detection
(this talk)



Xenon 1T



Colliders

SM

DM

SM

DM



Large Hadron Collider (LHC)
Credit: CERN

FermiLAT



Indirect detection
(Gammas, neutrinos and cosmic rays)

How do we build a detector for dark matter?

ParticleZoo.net

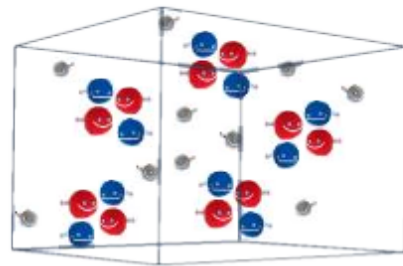


Some idea/s of what DM might be

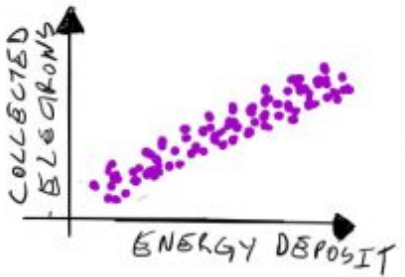
ParticleZoo.net



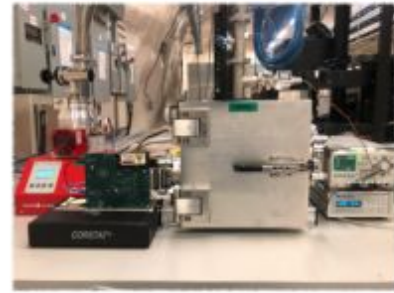
Some idea/s on how it interacts with known physics



Implement detection principle extending the available technology

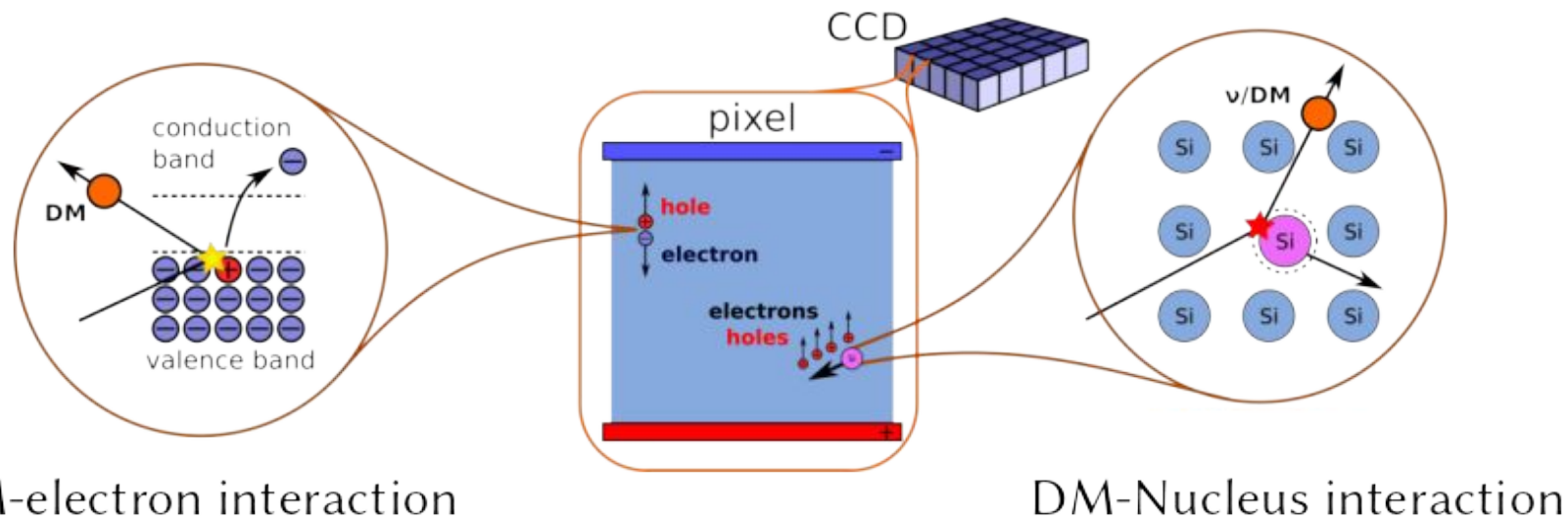


Understand the detector and go from signals to a plot



Mechanics + electronics + software

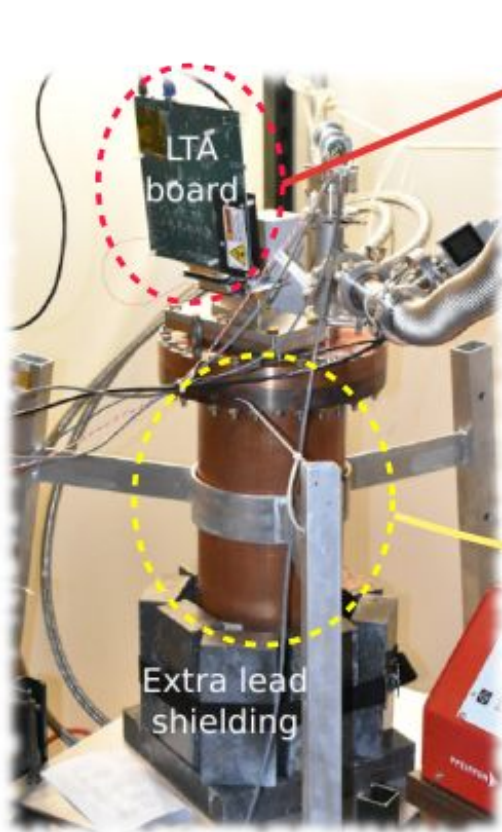
Dark Matter detection with silicon



We use science-grade CCDs (a very expensive digital camera!)



Dark Matter detection with silicon



We need vacuum, cool the CCD down (~ -215 F) and deploy it underground (~ 6800 feet or more!).



Foreseen detectors: SENSEI, DAMIC, OSCURA

Noah Kurinsky

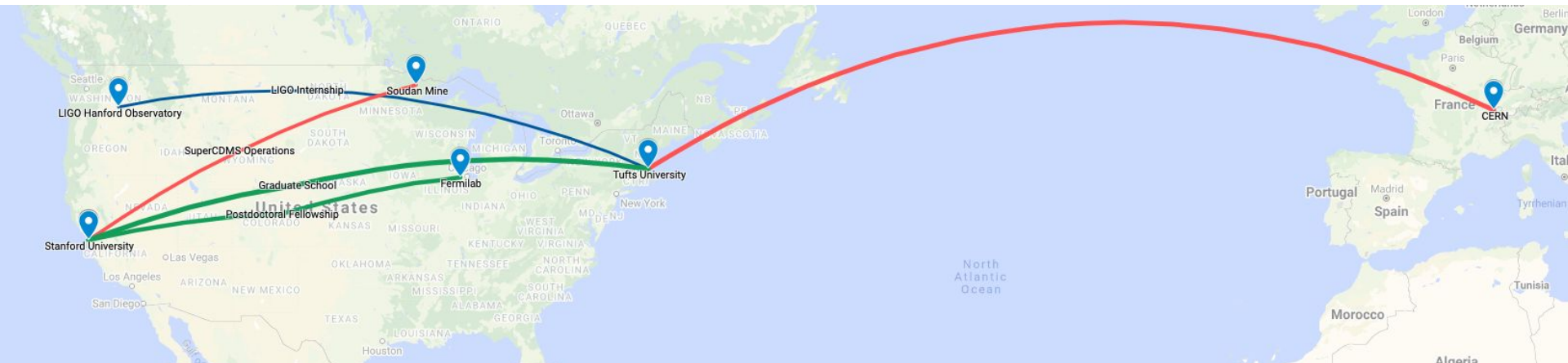


Becoming a Detector Physicist

I'm a postdoctoral fellow in the Particle Physics Division, and part of the Cosmic Physics Center

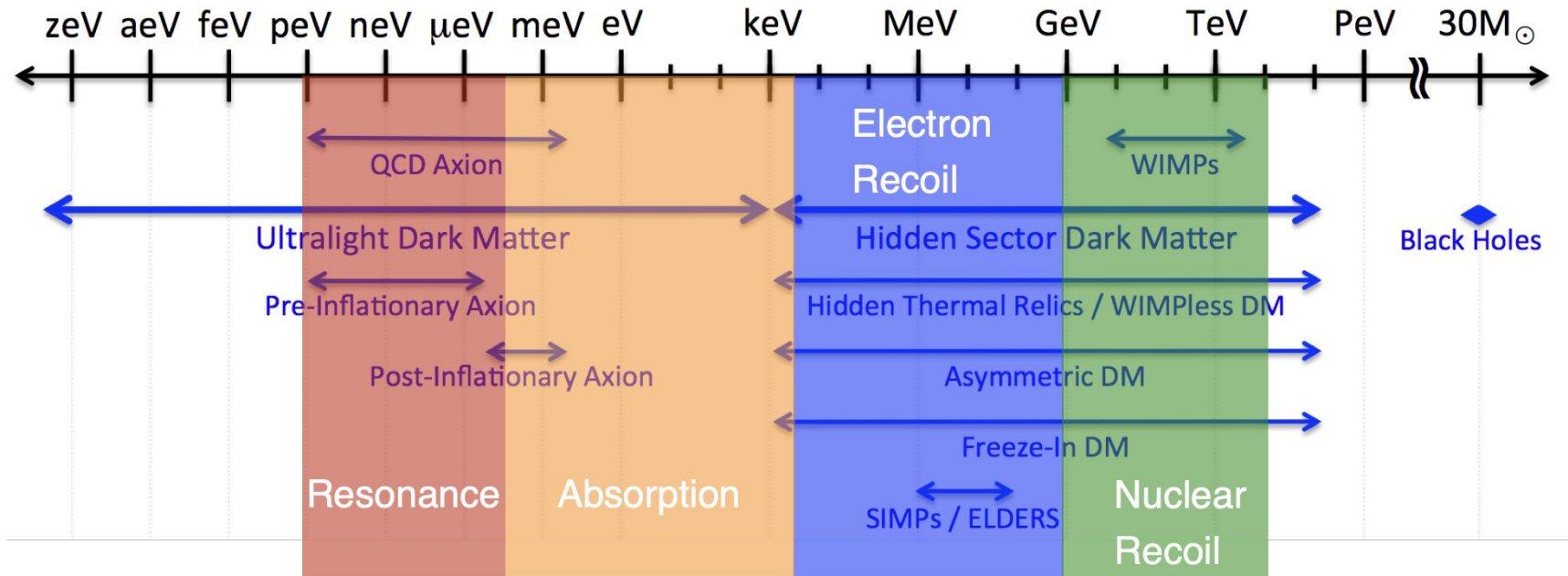
I have undergrad degrees in Astrophysics & Engineering Physics

I got my PhD from Stanford designing detectors for Dark Matter searches

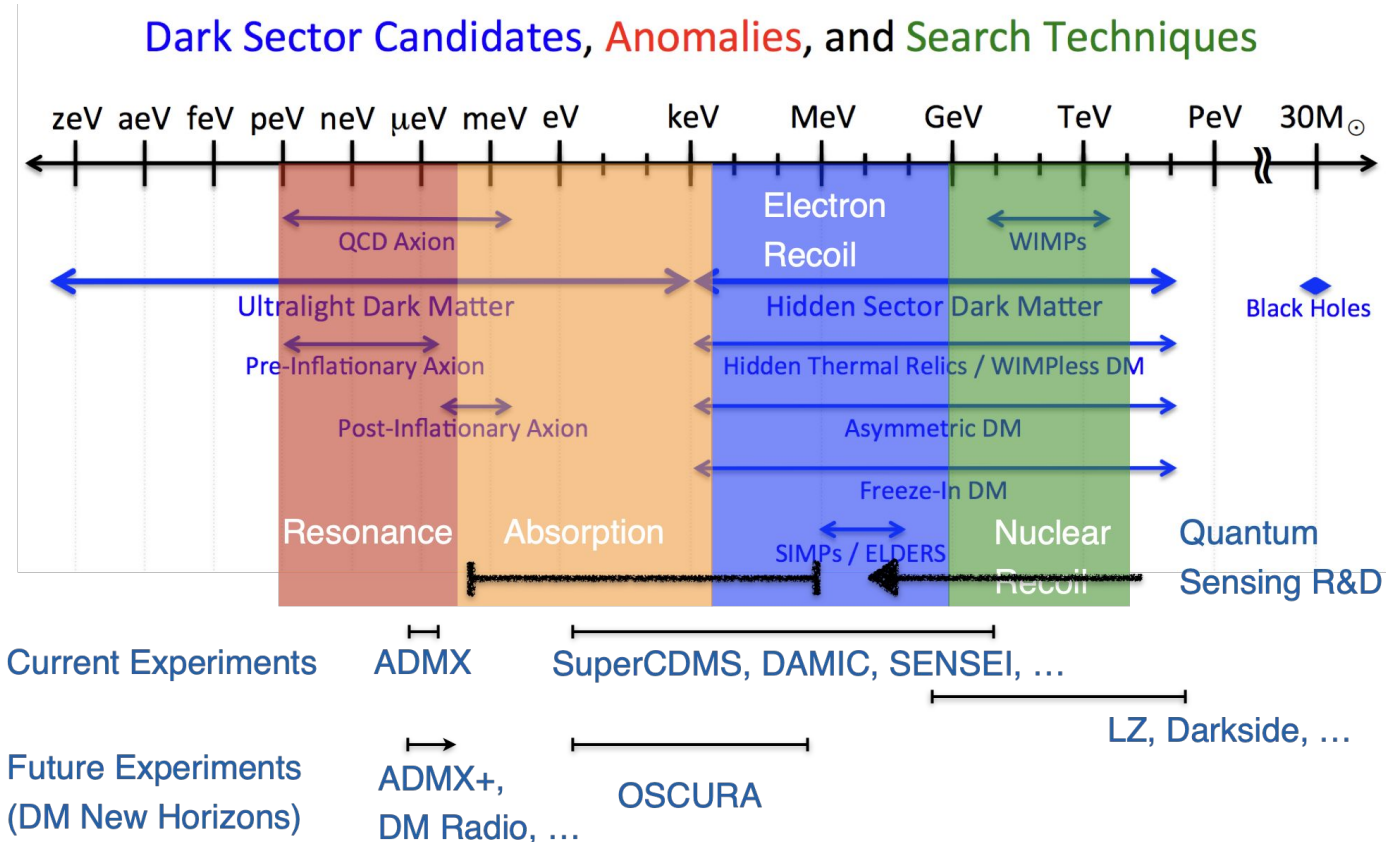


Types of DM Search Techniques

Dark Sector Candidates, Anomalies, and Search Techniques

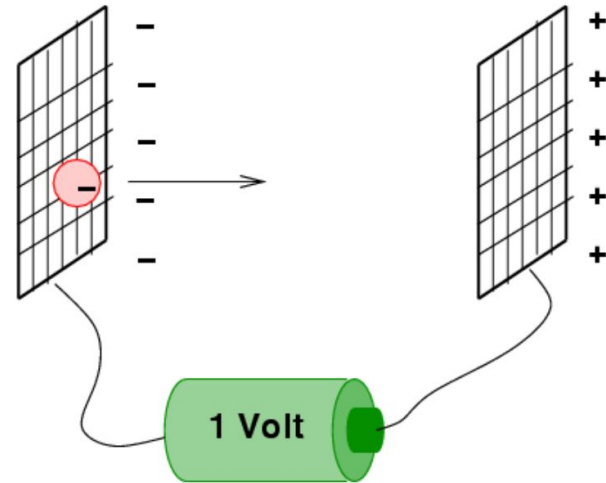


There are Many Experiments, Even Just at Fermilab!



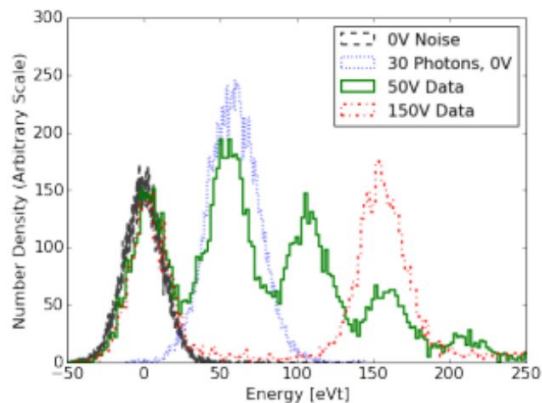
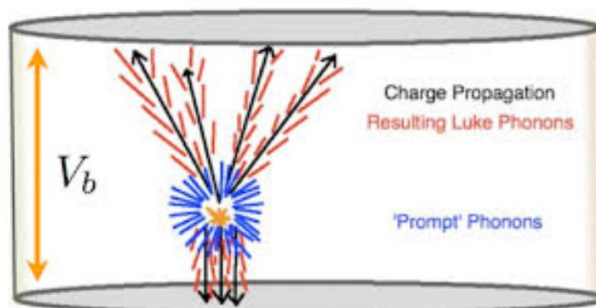
Detecting Small Energy Deposits

- The electron volt is a useful unit of energy for particles with a mass similar to the electron. It is defined as the energy gained by the electron when it passes across one volt of potential, roughly equal to the energy it would gain when going from the negative to positive terminals of a AA battery.
- To understand just how small this is, we can compare to units you may encounter in everyday life:
 - A conventional LED lightbulb uses around 5W of energy, which is $3e19$ eV (30 million trillion)
 - A grain of sand which falls from a height of 3 feet gains an energy of $1e9$ eV (one billion)
 - The grain of sand falling a distance equal to the width of a hair is 100 eV
 - For the same grain of sand to gain 1 eV during the fall, it would only need to move 1nm; roughly the size of 10 hydrogen atoms lined up!

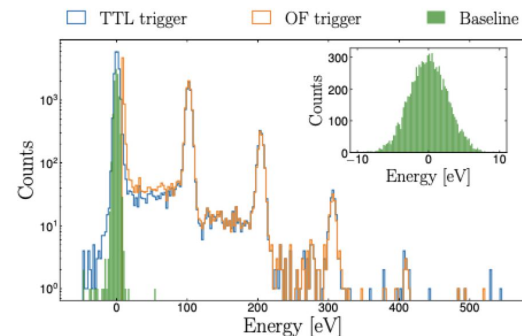


Heat Detectors

- In any recoil event, all energy is eventually converted to heat
- Heat is also produced when charges are drifted in an electric field; makes sense by energy conservation alone
- Total heat is initial recoil energy energy produced by drifting charge, as shown at right.
- Heat can be collected in sensitive 'thermometers' made with superconducting sensors!

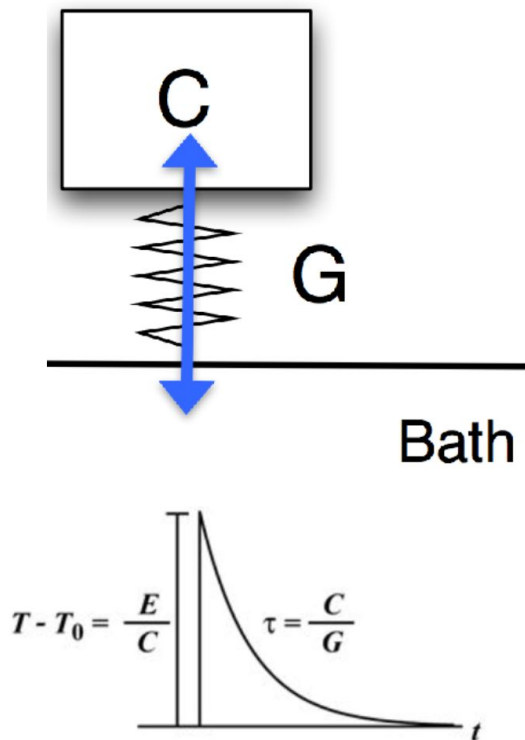


State of the Art:



Cryogenic Sensors

- Make your sensor very cold, so that a small energy deposit produces a large change in temperature
 - Most work at 10 mK, or -373C!
- Make your sensor very small, so that the heat capacity is small
 - Temperature change is E/C
- Provide a very cold heat sink to allow the detector to cool after a radiation event
 - Also helps understand the thermal response of the sensor
- A block of Tungsten at 50 mK has a heat capacity of 33 eV/K per cubic micron. How well you can measure temperature, and the size of the sensor, tells you how sensitive you are to injected energy.



NEXUS: Cryogenic Detector Test Facility

- 100m underground (right next to SENSEI)
- Base temperature of 10 mK
- Internal and external lead shielding
- Capable of running a wide variety of superconducting sensor technologies, including qubits

