

FNAL Neutrino
Joint Theory-Experiment
Working Group

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Radiative corrections
and neutrino scattering



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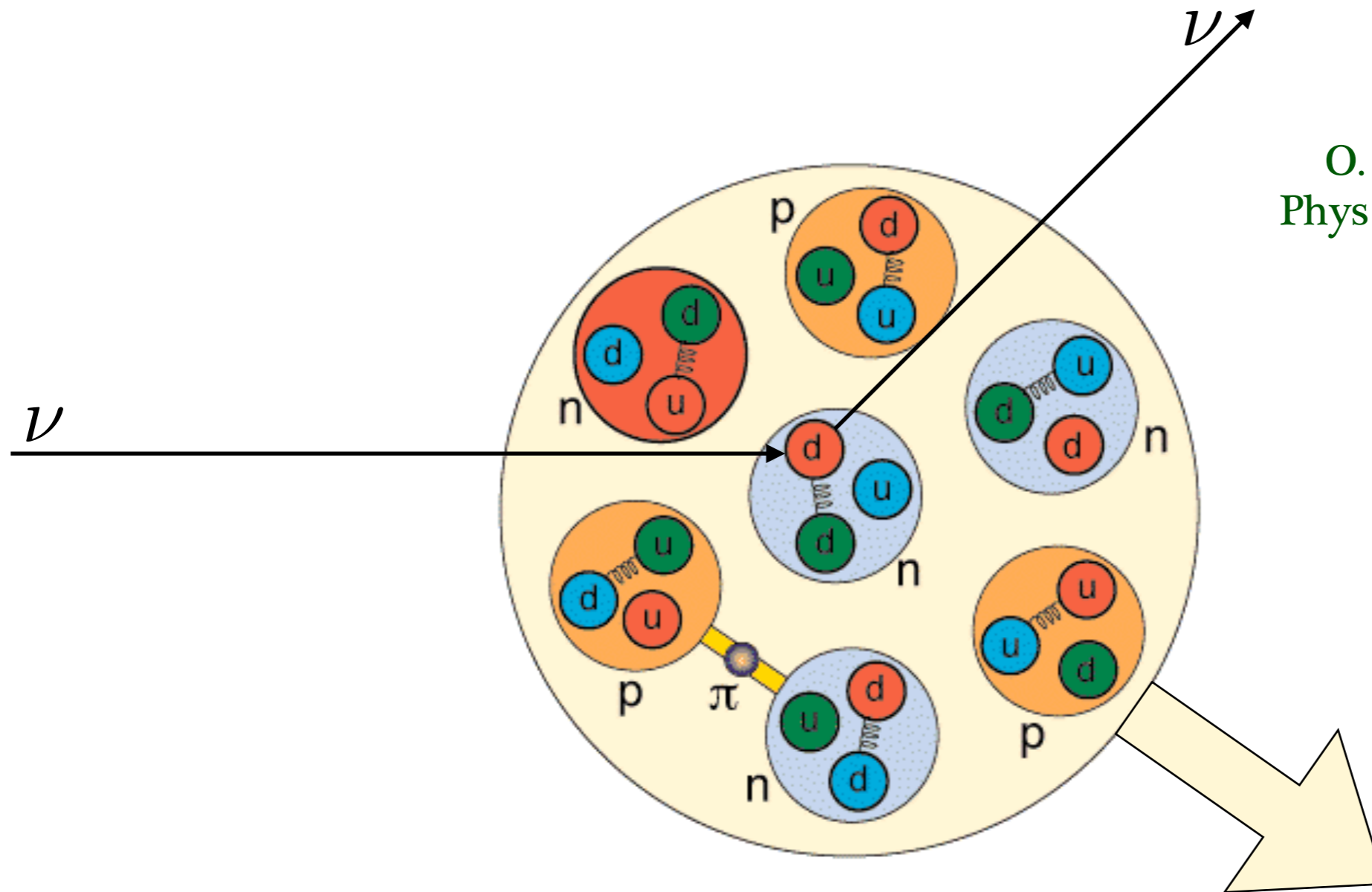


Radiative corrections in coherent elastic neutrino-nucleus scattering

O.T., P. Machado, V. Pandey and R. Plestid, JHEP 2102, 097 (2021)

neutrino energy < 100 MeV

From quarks to nuclei



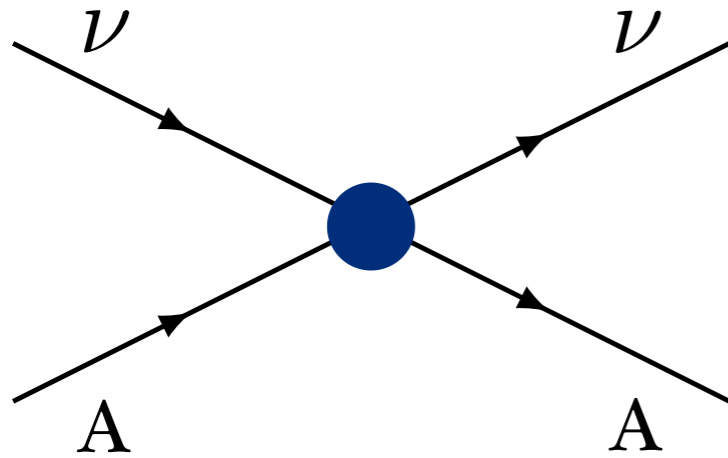
quark level

O. T. and Richard J Hill,
Phys Lett B 805 (2020) 135466

fafnir.phyast.pitt.edu

- scattering on quarks in nucleons in nucleus

CEvNS cross section on spin-0 nuclei

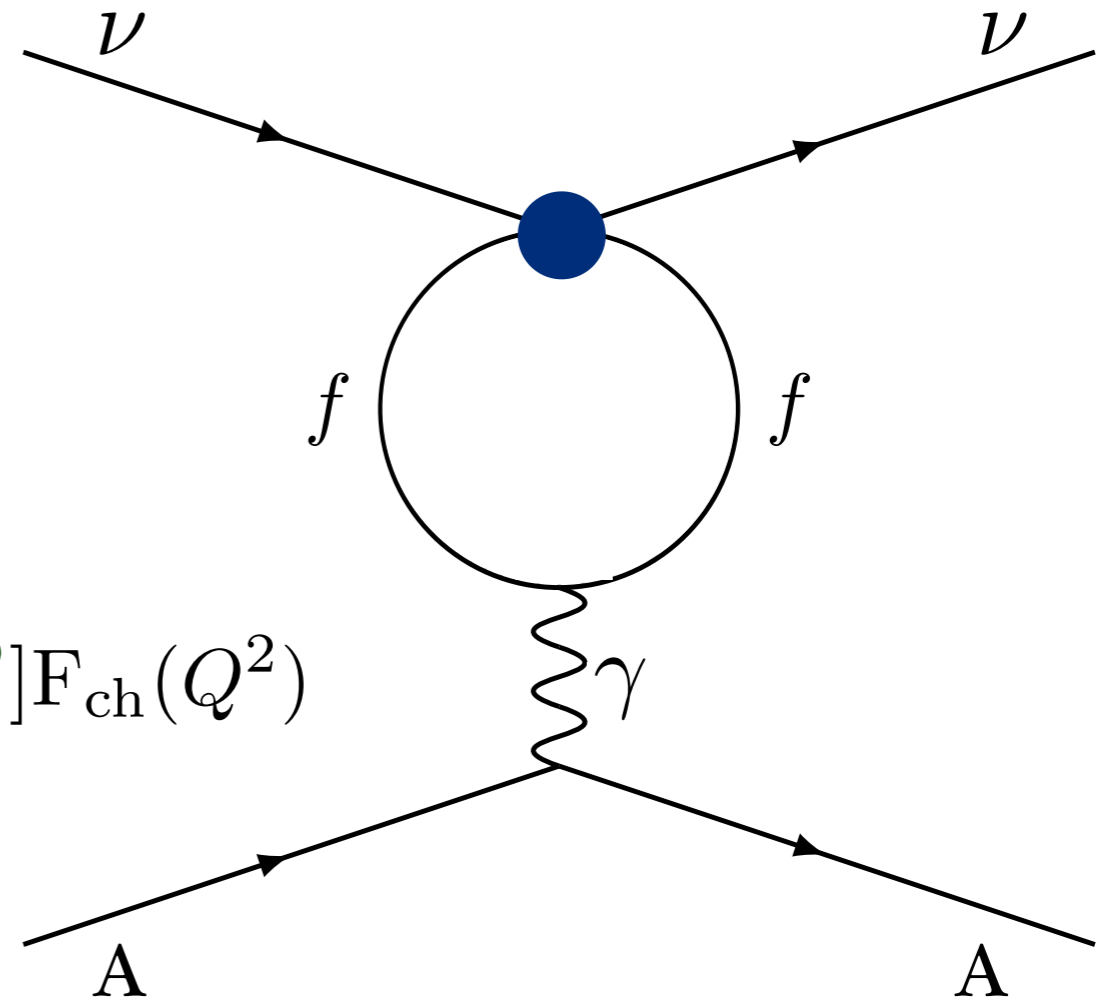


- tree-level cross section

$$\frac{d\sigma}{dT} = \frac{G_F^2 M_A}{4\pi} \left(1 - \frac{T}{E_\nu} - \frac{M_A T}{2E_\nu^2} \right) F_W^2(Q^2)$$

- effect of radiative corrections

$$F_W(Q^2) \rightarrow F_W(Q^2) + \frac{\alpha}{\pi} [\delta^{\nu e} + \delta^{\text{QCD}}] F_{\text{ch}}(Q^2)$$

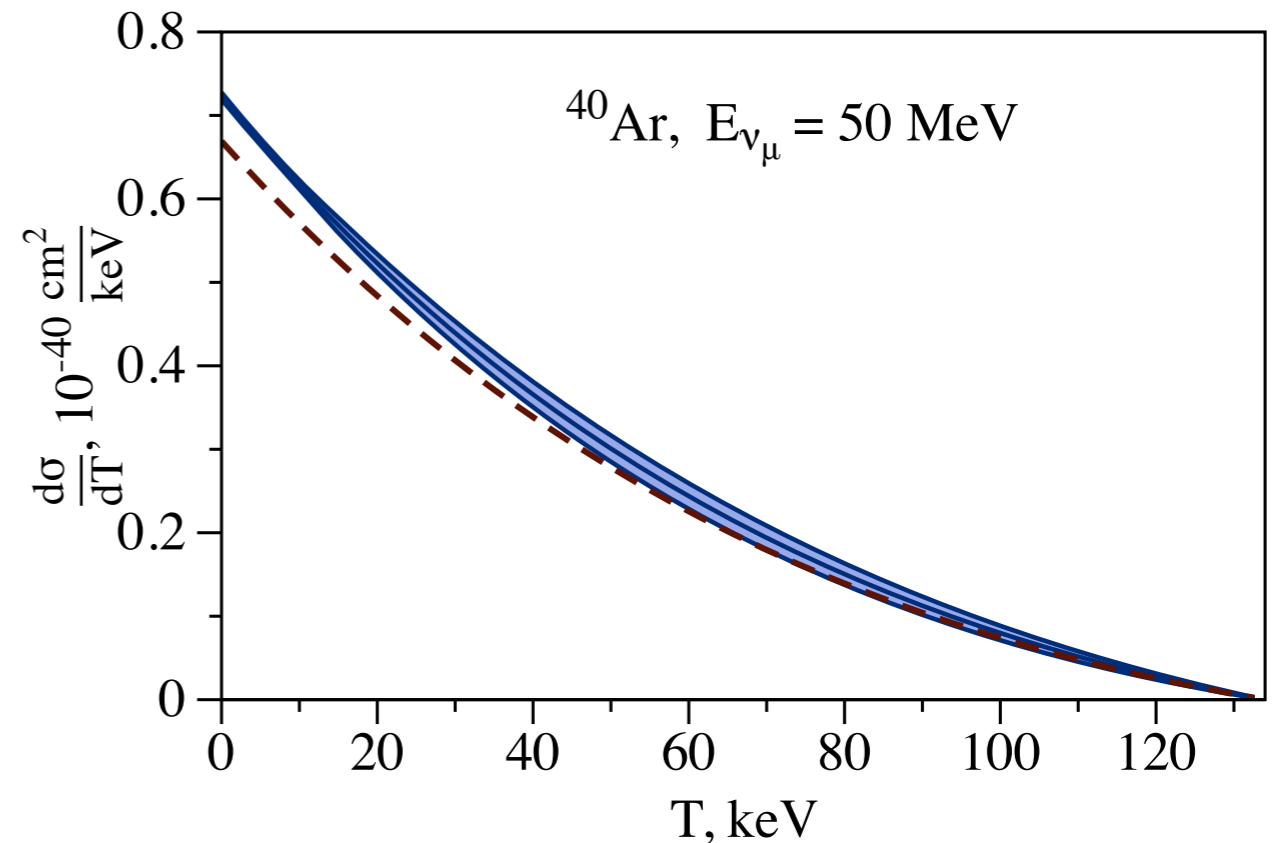
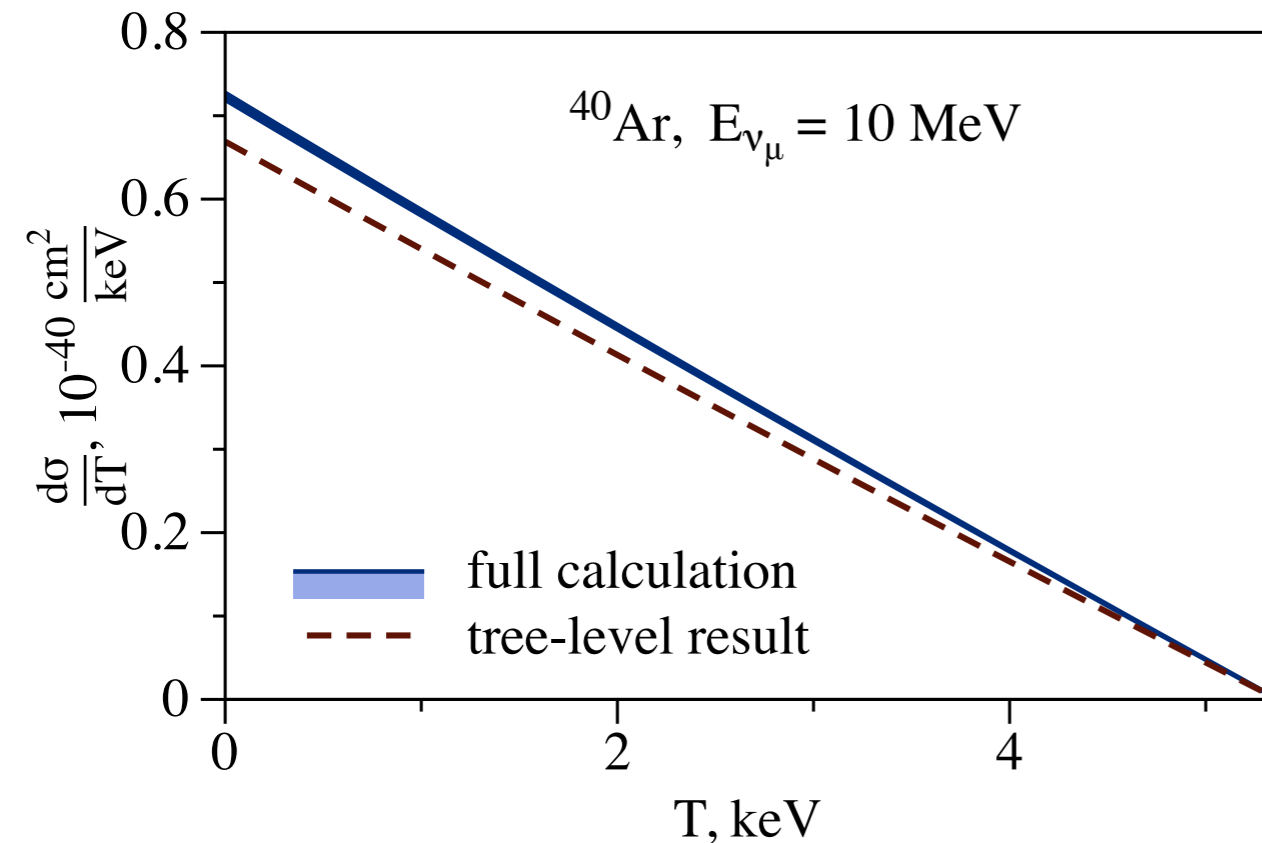


- radiative corrections enter with the nucleus charge form factor

Total and differential cross section

- recoil nucleus energy spectrum: one-loop vs tree level

nuclear models for point-nucleon form factors:
Yang et al. (2019), Payne et al. (2019), Hoferichter et al. (2020), Van Dessel et al. (2020)



- % effect of radiative corrections on cross sections



Radiative corrections in CCQE on free nucleons

Qing Chen, Richard J. Hill, Kevin S. McFarland and O. T. (arXiv: to appear)

soft-collinear-hard factorization framework

neutrino energy \sim GeV

Radiative corrections in CCQE

- large kinematical logarithms enhance radiative corrections

$$\frac{\alpha}{\pi} \sim 0.2 \% \quad \text{multiplied by} \quad \ln \frac{E_\nu}{m_\ell} \sim 6 - 10$$
$$E_\nu \gg m_e$$

- CCQE with electron flavor is subject to large corrections

- phase-space restrictions enhance radiative corrections

$$\frac{\alpha}{\pi} \sim 0.2 \% \quad \text{multiplied by} \quad \ln^2 \frac{E_\nu}{m_\ell} \sim 36 - 100$$

$$E_\gamma < \Delta E \quad \text{soft photons} \quad 2 \ln \frac{E_\nu}{m_\ell} \ln \frac{\Delta E}{m_\ell} \sim 35 - 60$$

smaller collinear logarithms

- crucial dependence on detector details

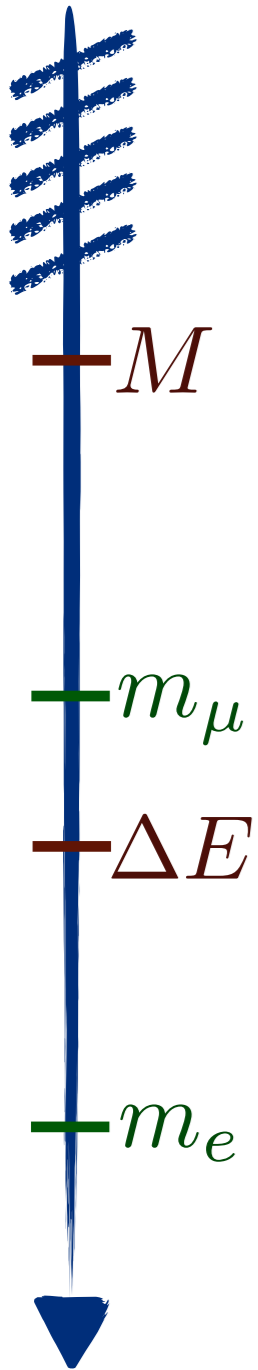
- radiative corrections crucial for %-level oscillation program

Factorization approach

- cross section is given by factorization formula

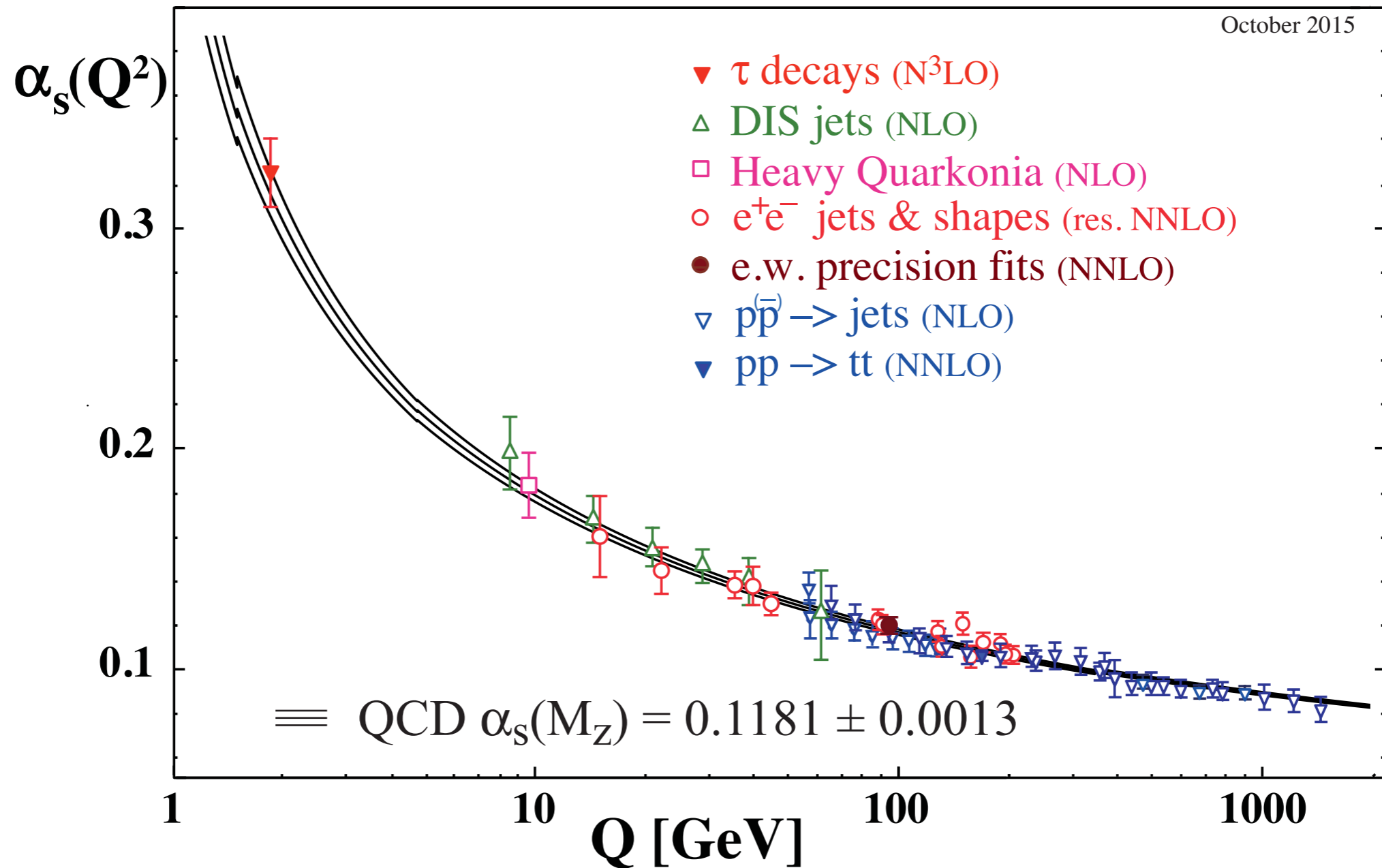
$$d\sigma \sim S \left(\frac{\Delta E}{\mu} \right) J \left(\frac{m_\ell}{\mu} \right) H \left(\frac{M}{\mu} \right) + \frac{\alpha}{\pi} O \left(\frac{\Delta E}{E_\nu}, \frac{m_\ell^2}{M^2}, (\Delta\theta)^2 \right)$$

- determine hard function at hard scale by matching experiment or model to the theory with heavy nucleon



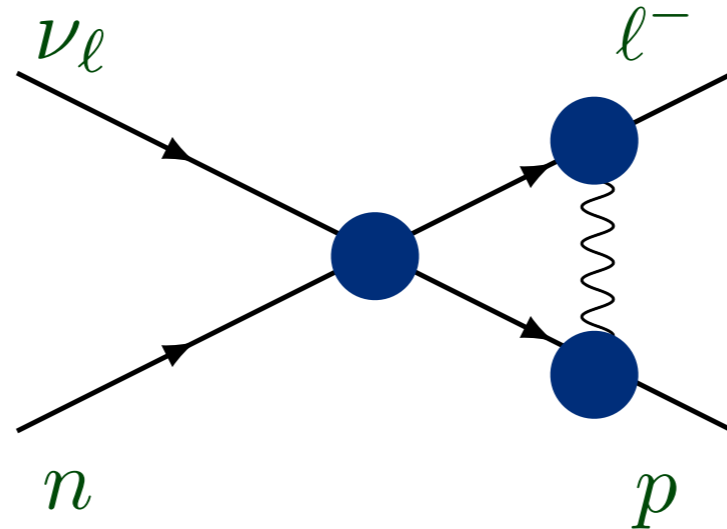
Interaction with nucleons

- QCD running coupling



- hadrons are correct degrees of freedom at GeV energy

Hadronic model at GeV scale



- exchange of photon between the charged lepton and nucleon
commonly used for 2γ , discussed for CCQE: Krzystof M. Graczyk (2013)
- assume **onshell form** for each interaction vertex

- best determination of hard function

Factorization approach

- cross section is given by factorization formula

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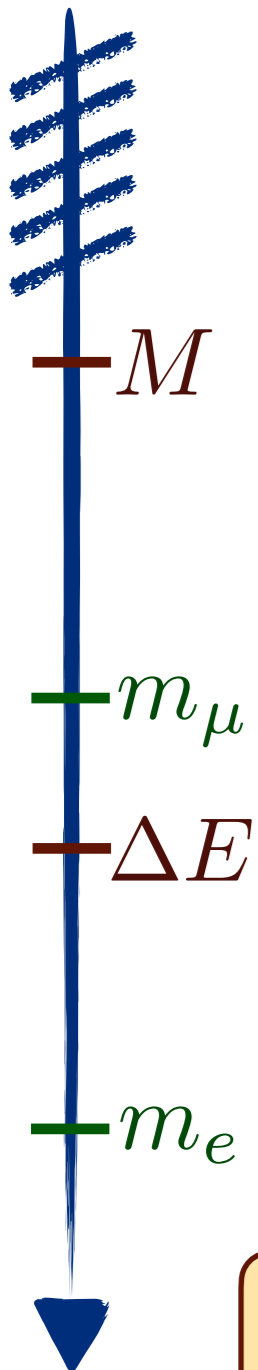
- RGE evolution of the hard function to scales $\Delta E, m_\ell$

- soft and collinear functions are evaluated perturbatively

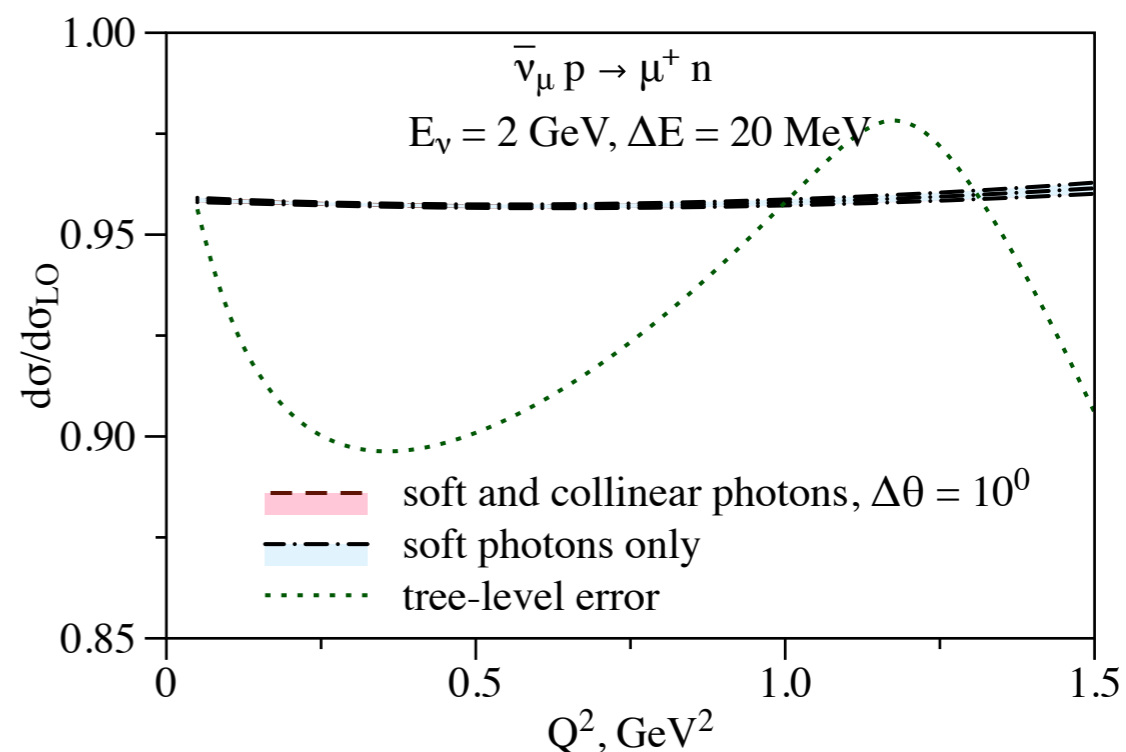
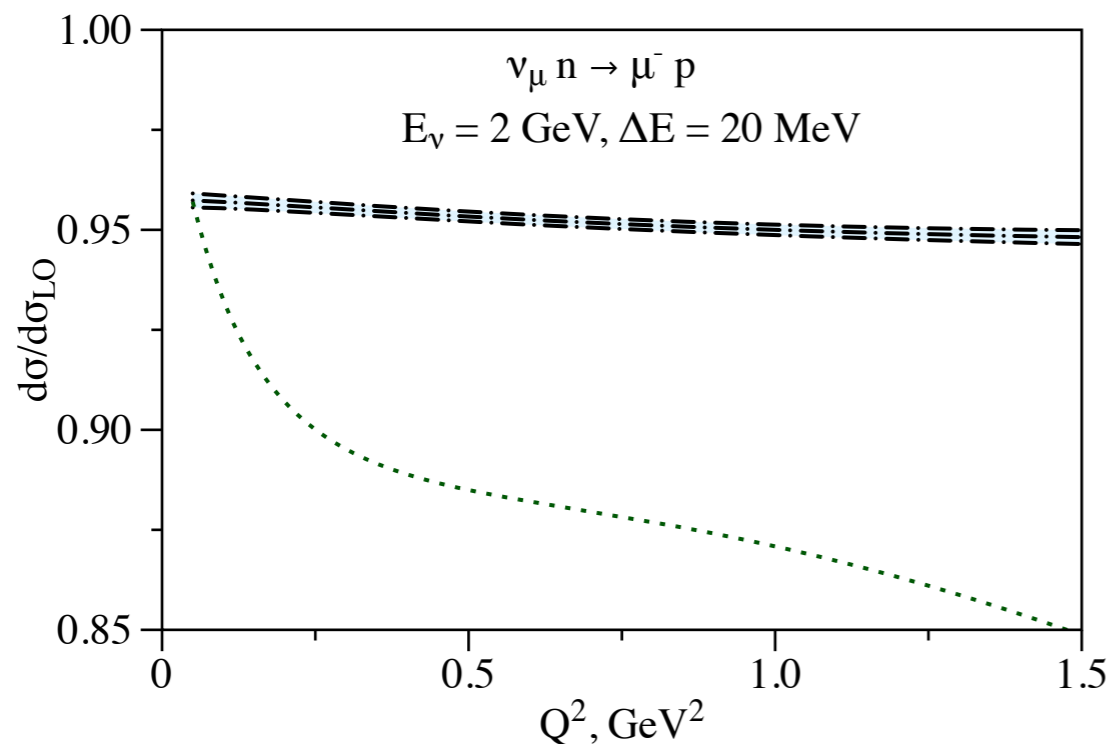
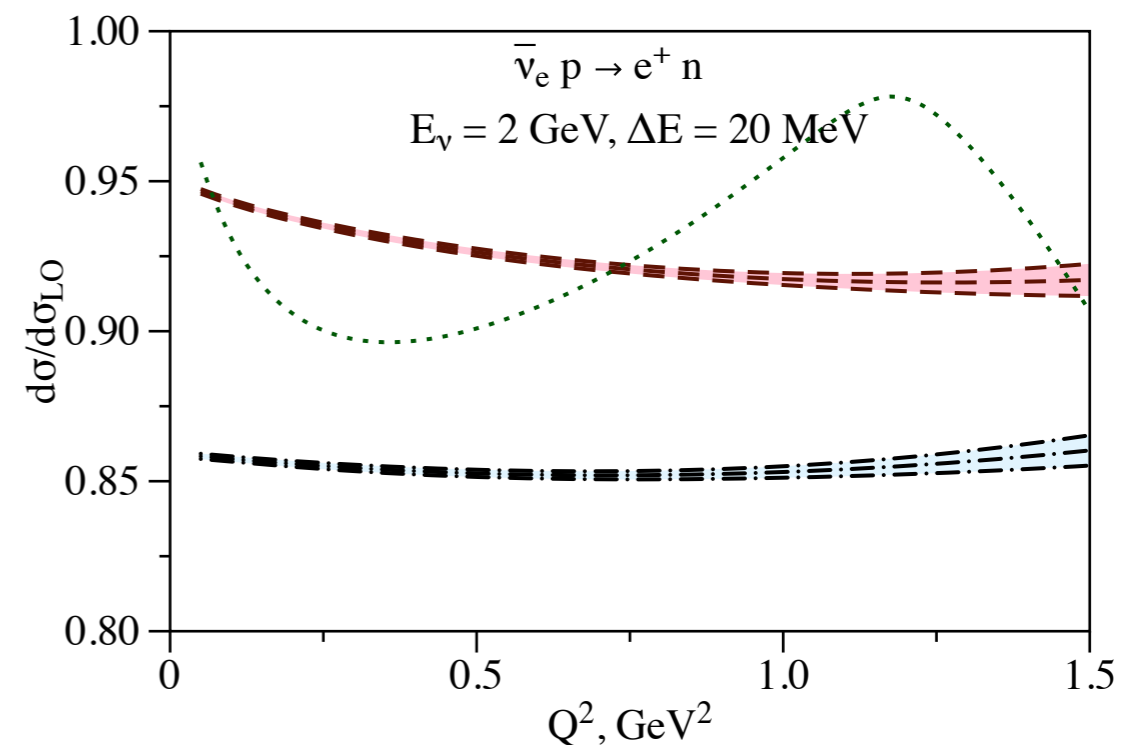
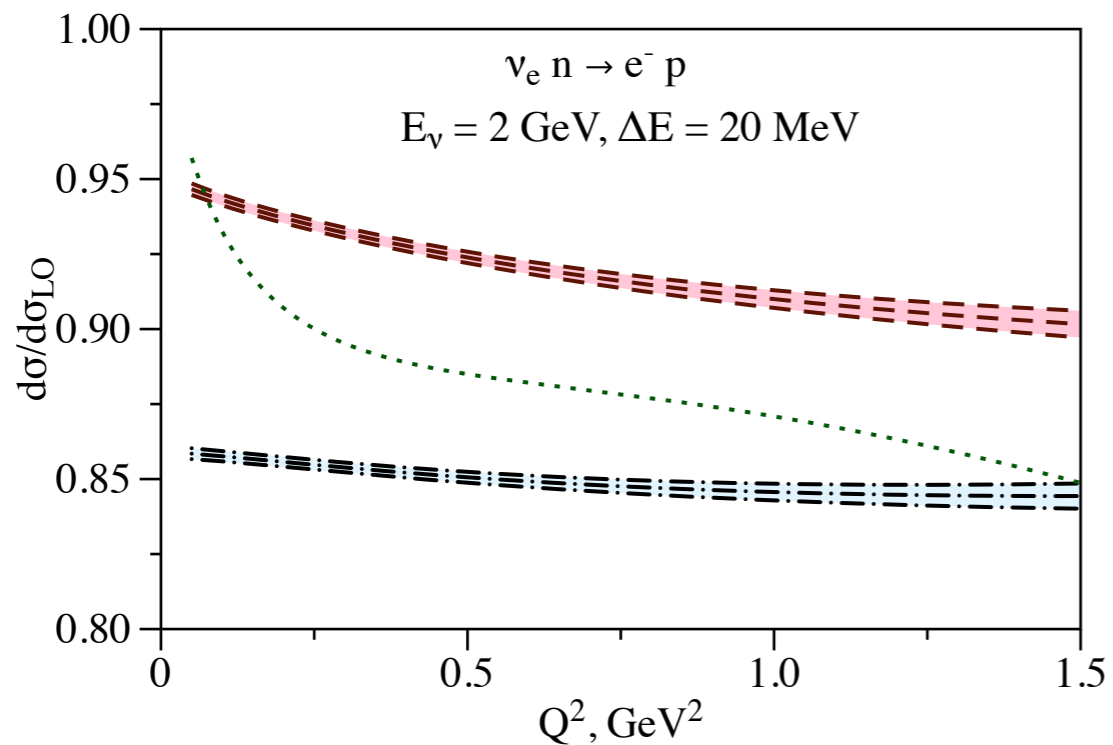
- calculate cross section at low energies accounting for **all large logs**

ep scattering with soft radiation only: R.J. Hill (2016)

- soft and collinear functions obtained **analytically**
- **hard** function describes physics at GeV energies



Cross sections



- corrections with e larger; cancellation virtual vs real

Radiative corrections

- more details at NDNN 2021 Workshop next week

Outlook

- phenomenological applications
- importance of resonances and resonance production

Ongoing work

Coulomb corrections in CCQE

Ryan Plestid, Richard J. Hill and O. T. (ongoing)

high-energy expansion

enhancement by nucleus charge Z

Stay tuned

Thanks for your attention!