

# An external muon tagger for ND-LAr

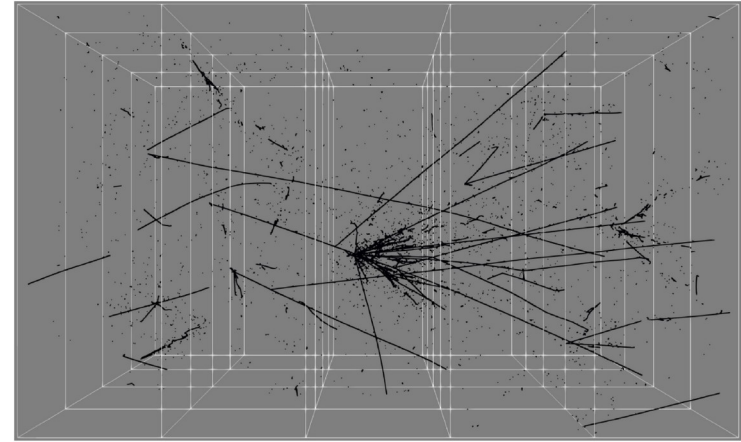
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22 March, 2021



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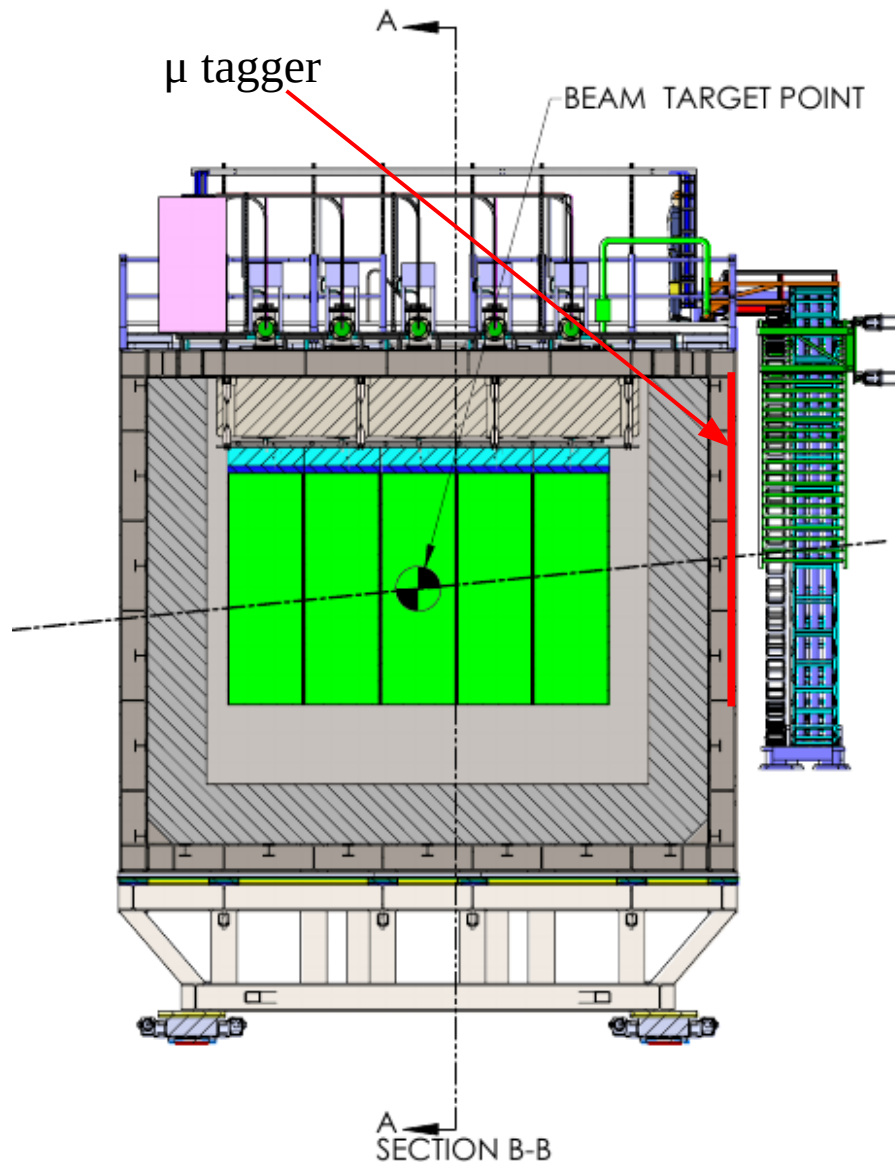


# Why



- Pile-up mitigation
  - ~30 rock muons per spill at 2.4 MW
  - Will produce light signals in ND-LAr that can be easily confused with fiducial muon tracks
  - External tagger can identify rock muons so they can be excluded from analysis
- Externally-tagged calibration sample
  - Select through-going tracks for calibrating relative energy response, argon purity, alignment, etc.
  - Also measure tracking efficiency

# What



- Two 1cm scintillator planes, with 3.85cm-wide strips in 90° stereo for 3D position
- Cover  $\sim 9 \times 4.5$ m area on upstream face of ND-LAr
- Use TMS scintillator design, fiber extrusion, readout electronics
- Piggy-back on existing TMS design and prototyping effort to reduce new costs

# How

- Propose to integrate external muon tagger into the ND-LAr design, and coordinate the work through this consortium
- Close collaboration with TMS is critical
- Next steps:
  - Continue design optimization work
  - Develop a plan for integration
  - Understand DAQ requirements and plan
  - ...

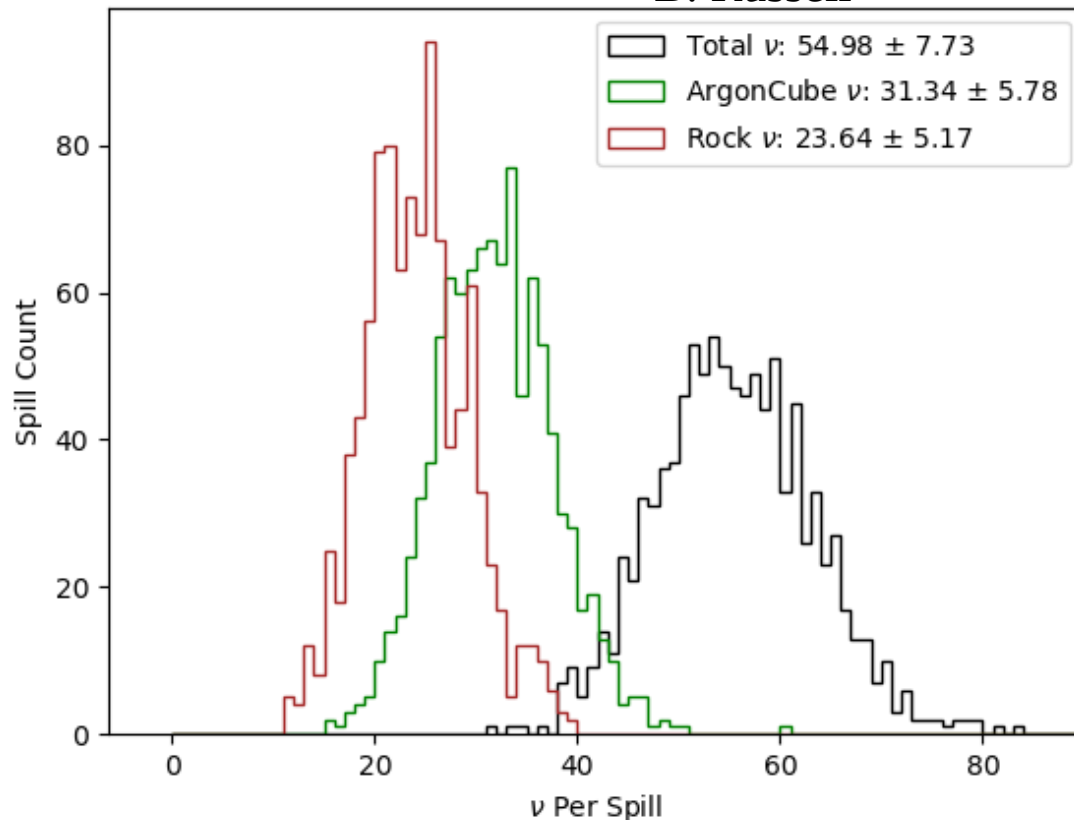
# My talk from Thursday's ArgonCube meeting follows

# Outline

- Why we need a muon tagger
  - Charge-light ambiguity
  - Calibration sample tagging
- Can't the photon detector double as the muon tagger?
- External tagger preliminary physics requirements
- Design concept
- To do

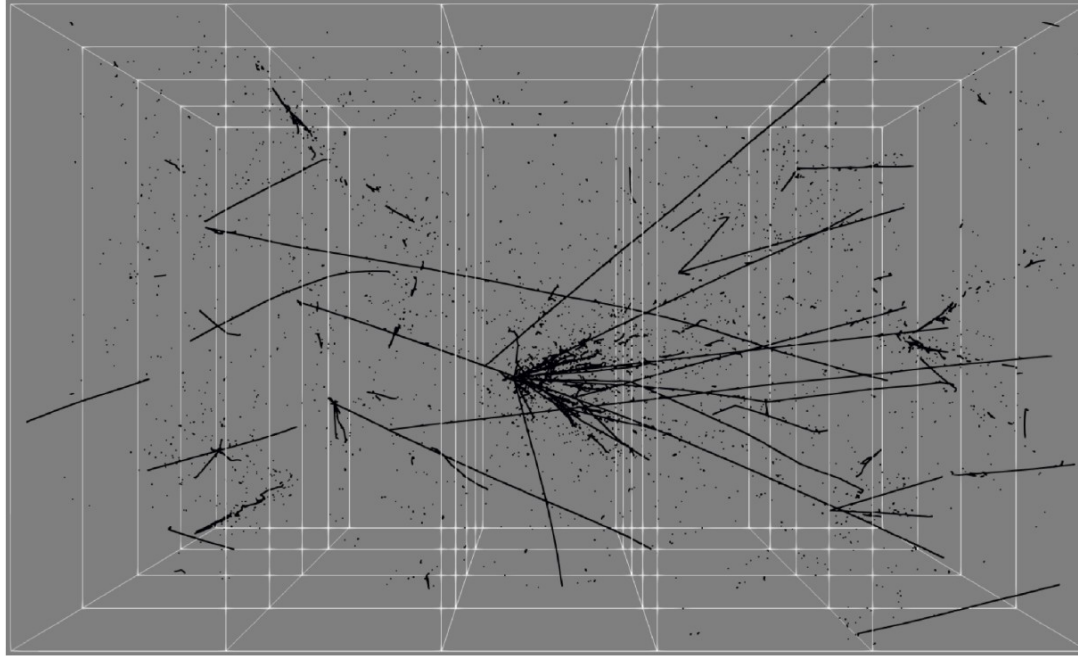
# Pile-up is going to be challenging

B. Russell



- At 1.2 MW,  $\sim 55$  neutrino interactions per spill deposit energy in ND-LAr
- $\sim 45\%$  originate in the rock – entirely external to ND-LAr
- $\sim 80\%$  of rock-originating interactions that reach ND-LAr include a muon, the rest are mostly neutrons
- You should multiply this by 2 in your head – we need to be able to handle the 2.4 MW intensity – over 100 interactions per spill!

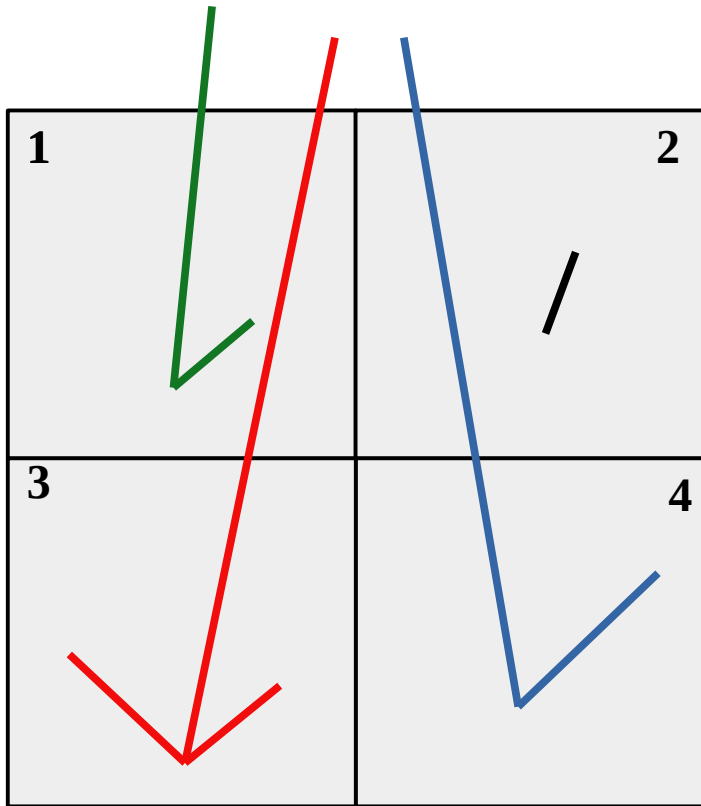
# Detached energy deposits require timing information



- Due to 3D charge readout, fully-connected interactions can *probably* be reconstructed without scintillation/timing information, though this has not been demonstrated
- Associating detached energy to the correct neutrino interaction is critical, and nearly impossible without timing in some cases
- This is unique to the ND, so getting it wrong will result in a ND/FD difference that will have to be corrected with models and lead directly to increased systematics in the OA



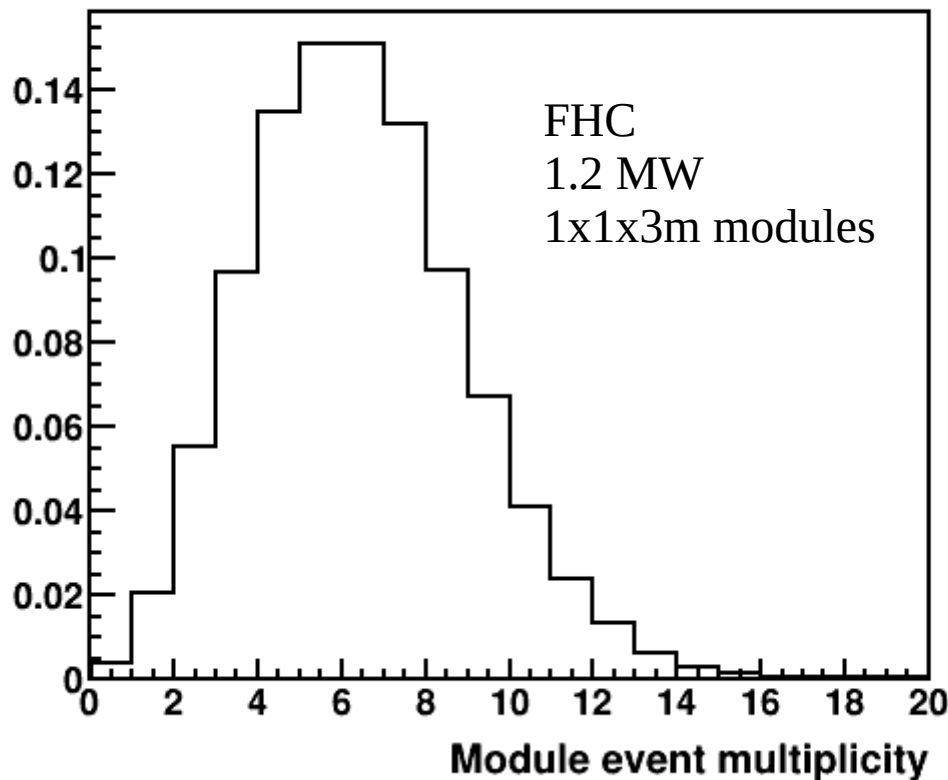
# Simple cartoon example



- Thanks to 3D readout, all the tracks and vertices can be found
- But which neutrino produced the black detached energy in module 2?
- If it's red or green, then there will be two distinct scintillation signals in module 2, and you can use timing to associate the black charge signal to the right event

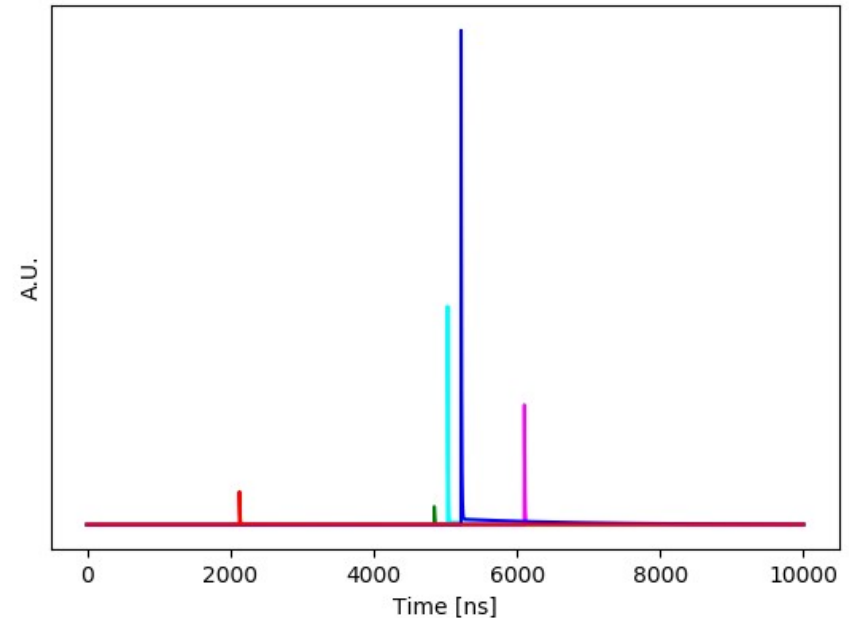
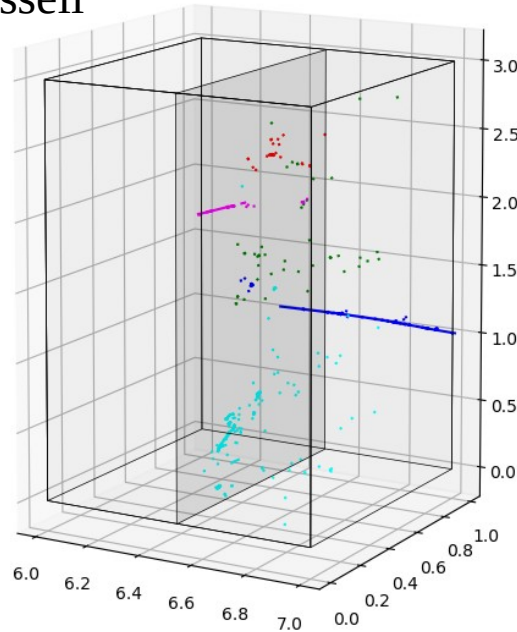
# Optical modularity is critical, but it's still nontrivial

- ~55 events per spill in entire detector
- ~6 events per spill per module, with a tail out to 15
- A typical module will record ~6 charge signals and ~6 light signals in each spill



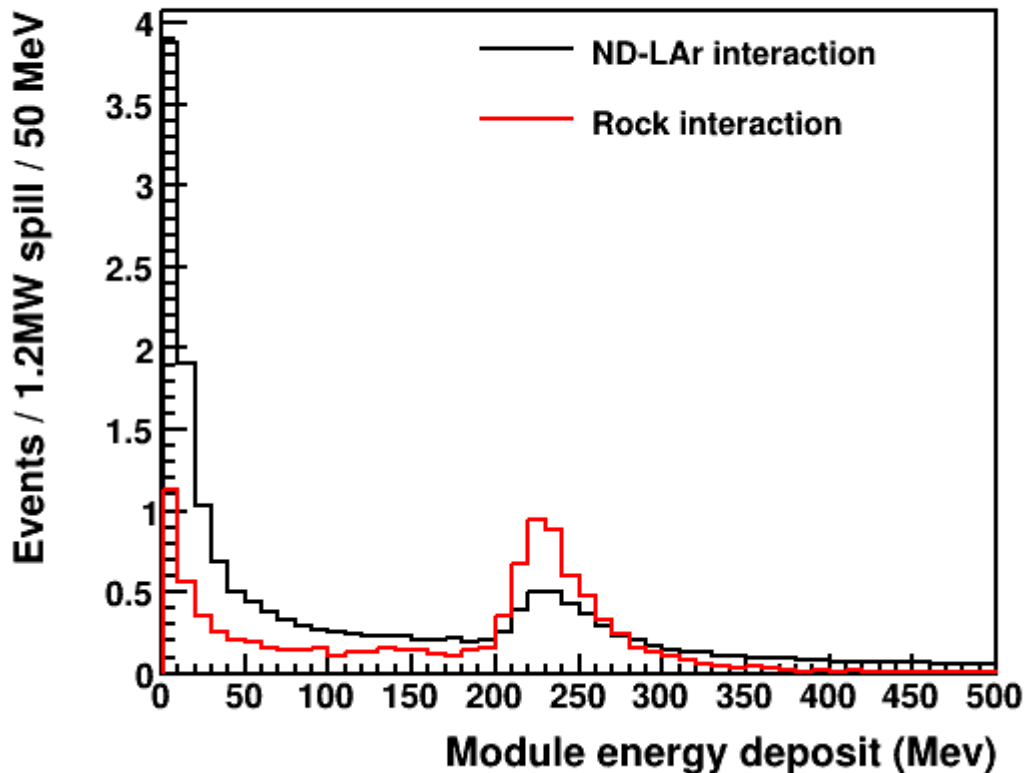
# Charge-light matching creates potential ambiguities

B. Russell



- Which light signal goes with which charge signal? “Flash matching” problem
- Modularity helps *a lot* – without it flash matching would be a 50-to-50 problem
- Three handles to help with matching: (1) position resolution, (2) pulse height resolution, (3) correlations between modules

# Through-going MIP tracks are extremely common

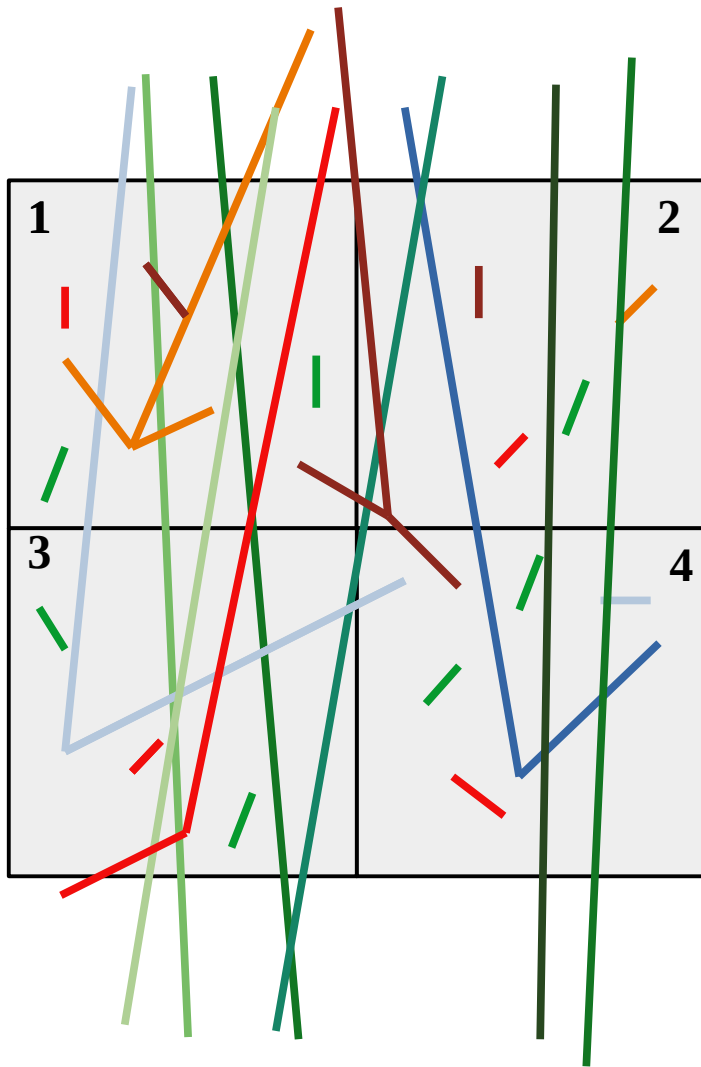


- True energy deposit per 1x1x3 module for neutrino interactions in rock, and ND-LAr (including all passive elements, cryostat, etc.)
- On average, a typical module will have ~2-3 low-energy events, which are often detached
- On average there are ~2 through-going MIP tracks per module per spill, ~1 of which is a rock muon
- Poisson fluctuations to several rock muons in one module will be common

# Flash matching: summary

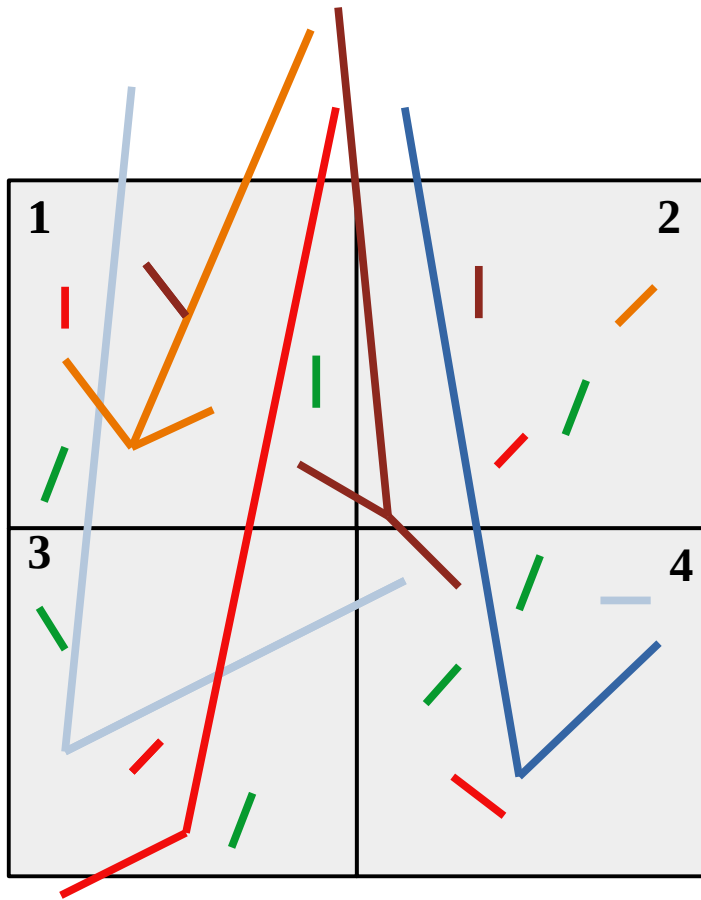
- An average 1x1x3m module in an average 2.4 MW spill will have:
  - 2-3 through-going rock muons
  - 2-3 through-going muons from ND-LAr interactions
  - 1 neutrino interaction vertex
  - 4 detached lower-energy blips/showers
- 10 time-separated prompt scintillation signals, of which 5 are essentially the exact same pulse height

# Example: realistic 2.4 MW spill



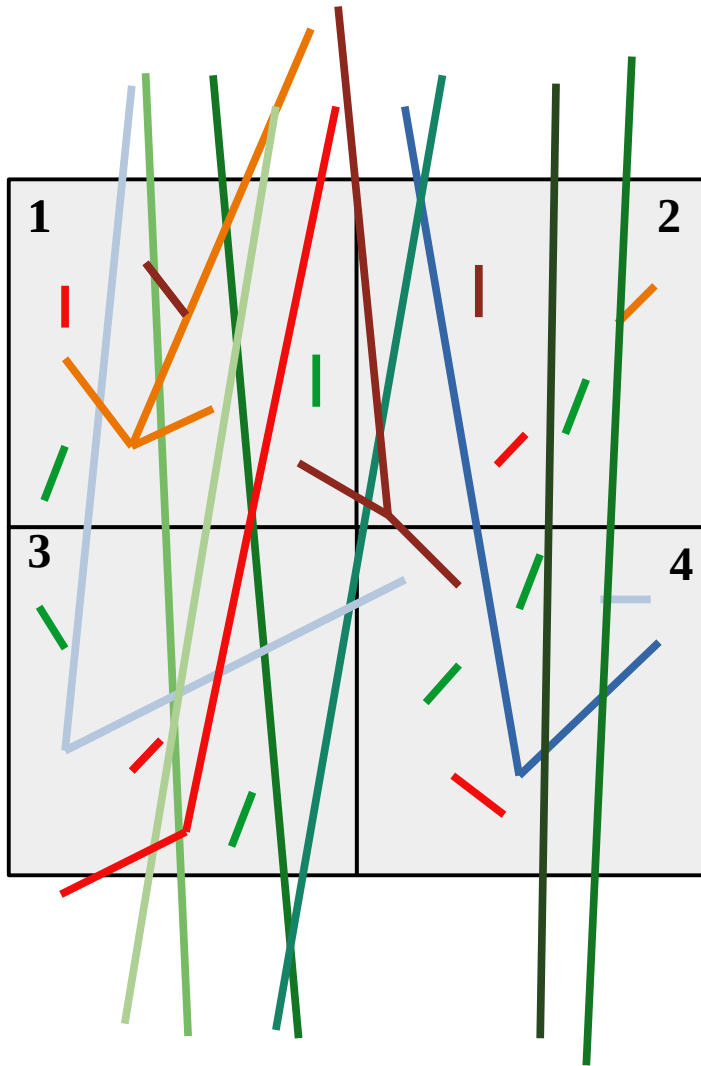
- Tracks don't actually overlap and can be reconstructed, but most detached energy requires timing
- Associating detached energy requires measuring the time of the detached deposits and the vertices
- Muon signals will look the same, position resolution is the only way to distinguish them

# Example: realistic 2.4 MW spill



- External muon tag with  $\sim$ few cm position resolution and single-bucket timing would provide external time stamps for rock muons
- The photon detector signals at these times are then known to be due to rock muons
- External tracks can just be removed from the charge-to-light matching problem
- Reduces the number of tracks by  $\sim$ factor of 2

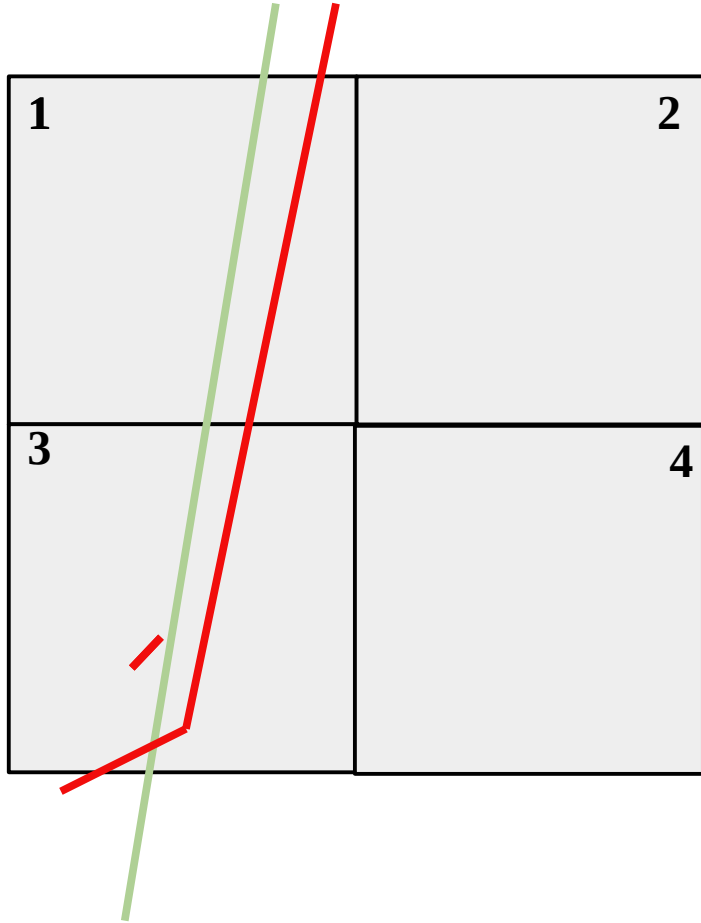
# Can't the photon detector just do this?



- If photon detector can distinguish through-going track from in-module light signal, and tiling is  $\gg 6$  per module (so that  $\sim 6$  tracks  $\sim$  never overlap)
- Then you still have a  $\sim 2x$  harder combinatorics problem at reconstruction, but it is theoretically possible



# Can't the photon detector just do this?

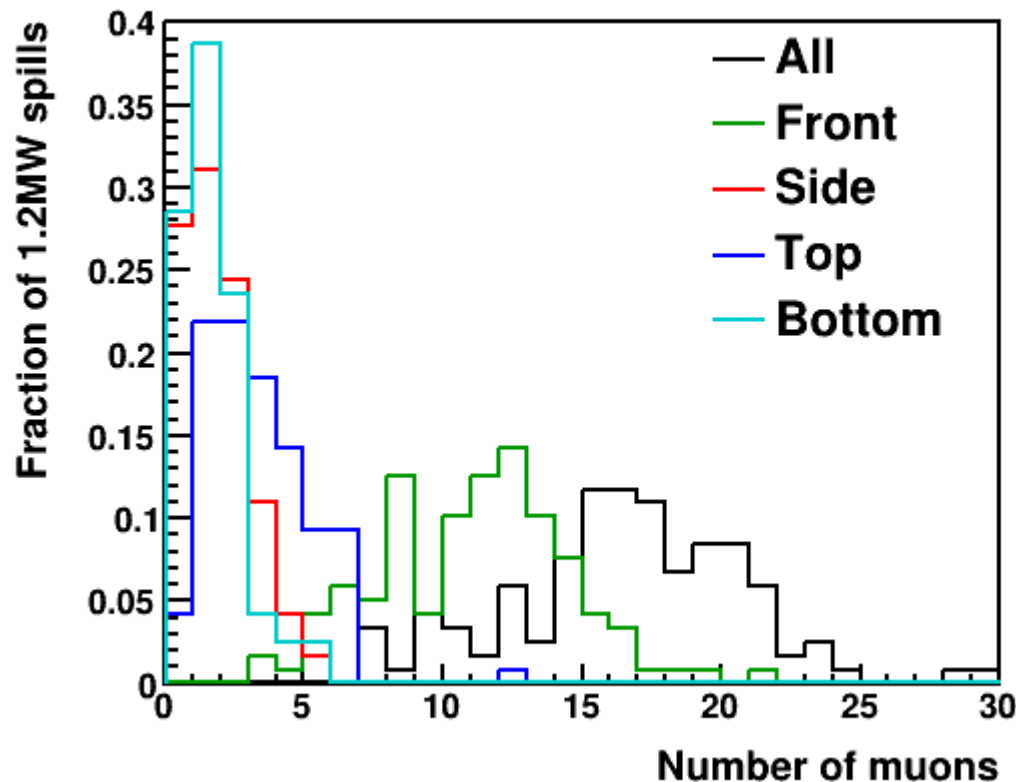


- But back-scatters could intersect light detector
- Two events like this could literally look exactly the same to the photon detector
- An external muon tag would reduce this type of ambiguity by  $>2\times$

# Another benefit: externally-tagged calibration sample

- Through-going muons will populate the entire detector and are an extremely useful calibration sample:
  - Relative energy response calibration
  - Argon purity measurement
  - Module-to-module alignment
  - Tracking efficiency studies
  - Timing calibration of photon detector
  - All can be done vs. time:  $O(1\text{M tracks per day})$
- Matching muon tagger signal to TMS/ND-GAr reconstructed tracks gives a clean sample of through-going muons of known position and known timing that is entirely independent of any ND-LAr detector subsystem

# Rock muons per 1.2 MW spill by entry location in LAr active volume



- ~10-25 total rock muons per spill at 1.2 MW
- ~5-20 enter upstream face of ND-LAr
- Top is more important than sides or bottom due to beam angle
- Double this for 2.4 MW – need to cleanly separate ~50 events in space/time

# High-level physics requirements

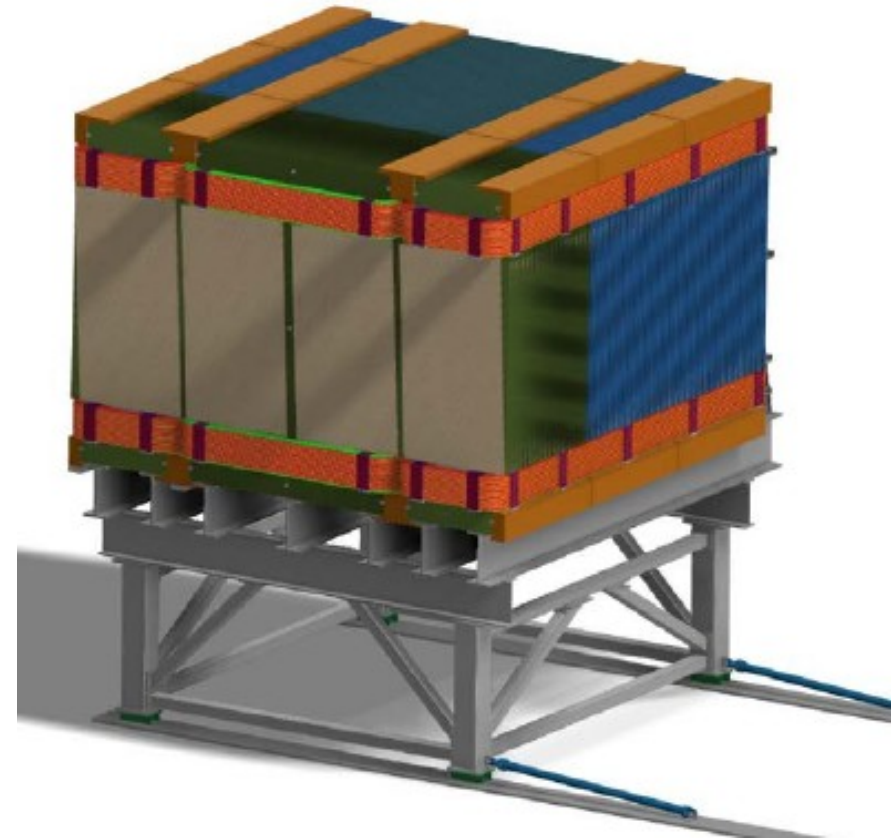
- Very high efficiency to detect MIP signals, but not necessary to measure energy precisely
- Separate  $\sim 50$  tracks with sufficient position resolution to match to ND-LAr tracks, and sufficient timing resolution to match to photon detector signals
- Cover sufficient area to detect most external muon tracks

# Lower-level physics requirements

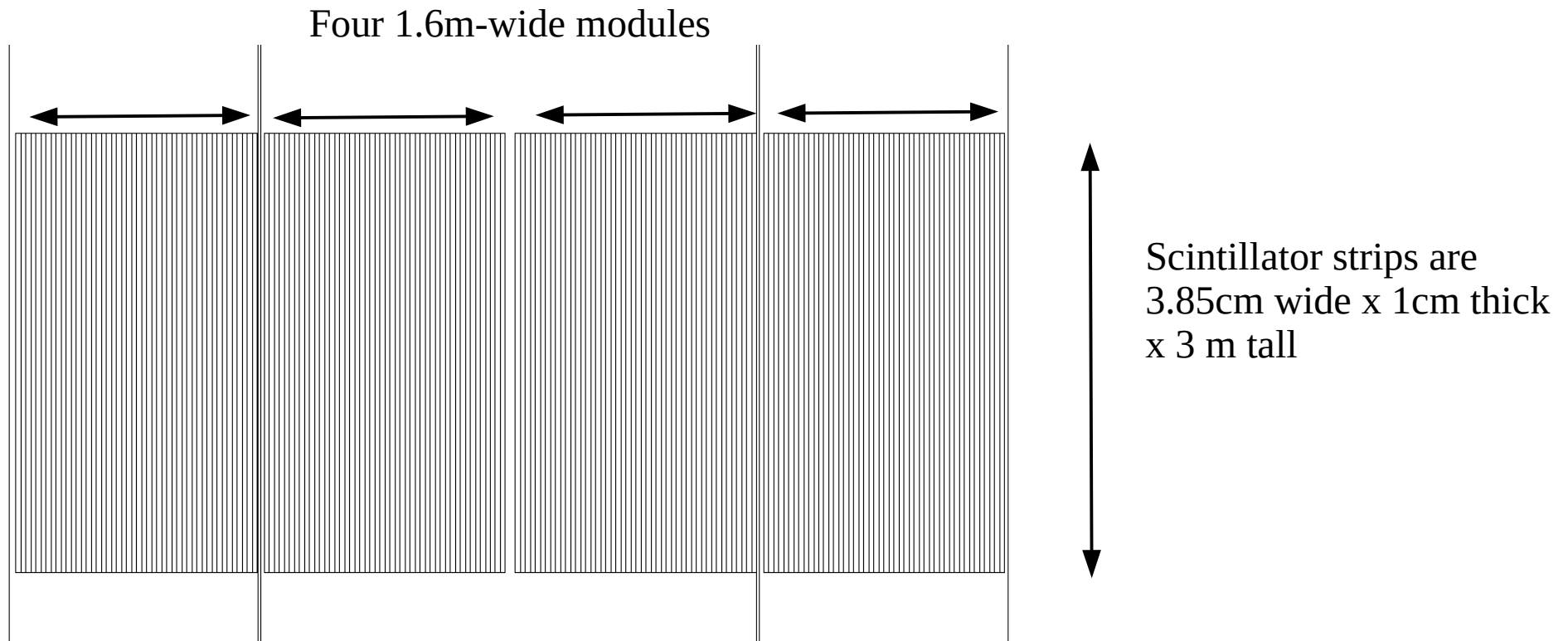
- Some further study is required, but:
  - >95% efficiency for MIP signal
  - ~5 cm position resolution in each transverse direction
  - <19 ns timing resolution
  - Cover entire front face of ND-LAr, with additional coverage above active volume
  - Move with the detector

# This all sounds familiar: TMS

- Scintillator of TMS has virtually identical requirements:
  - Be roughly the size of ND-LAr
  - Measure muons efficiently
  - Resolve single bucket timing

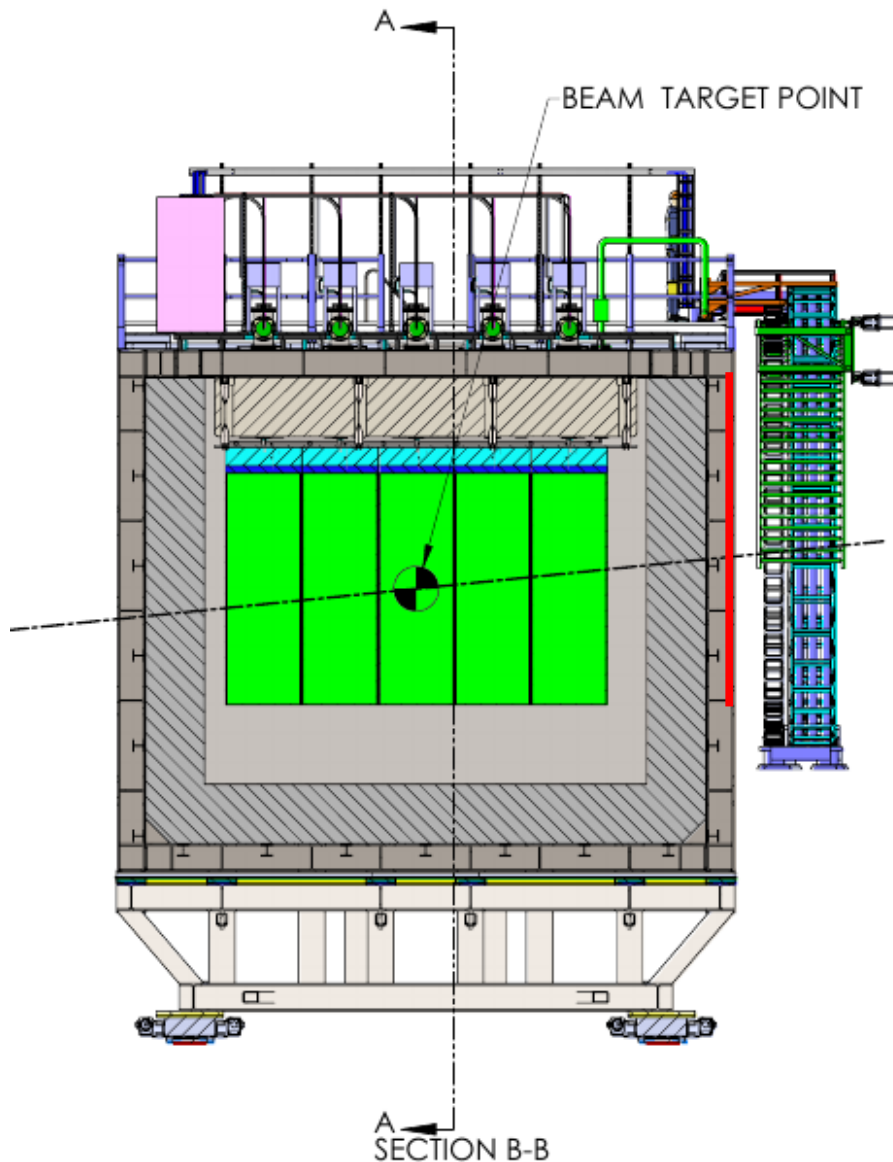


# TMS scintillator geometry



- Front/neutrino view of one plane of TMS scintillator
- Strips are very similar to MINOS

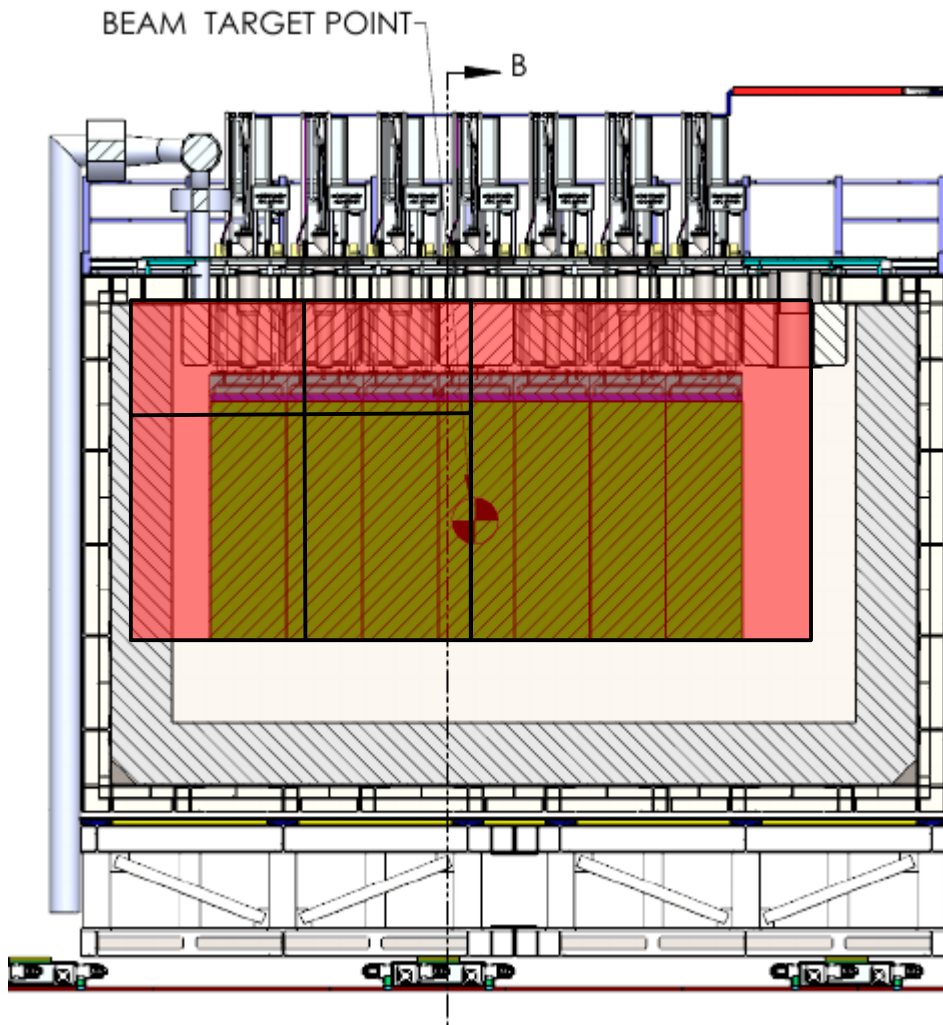
# Very preliminary design concept



- Two 1cm-thick scintillator planes, using TMS strips, fibers, readout, but 90° stereo
- Mounted to the outside of the warm steel structure on the upstream face of ND-LAr
- Extends from bottom of active volume to ~1.5m above active volume – this will allow “top-entering” muons to be tagged



# Very preliminary design concept



- Total coverage area shown here would be 9m x 4.5m, would give high acceptance for near-side entering tracks, and also top-entering
- Two 4.5x4.5m modules, so that maximum light propagation time is  $\sim 20$  ns which is also bucket spacing
- Readout on top and sides, with mirror in the middle to avoid having 9m strips

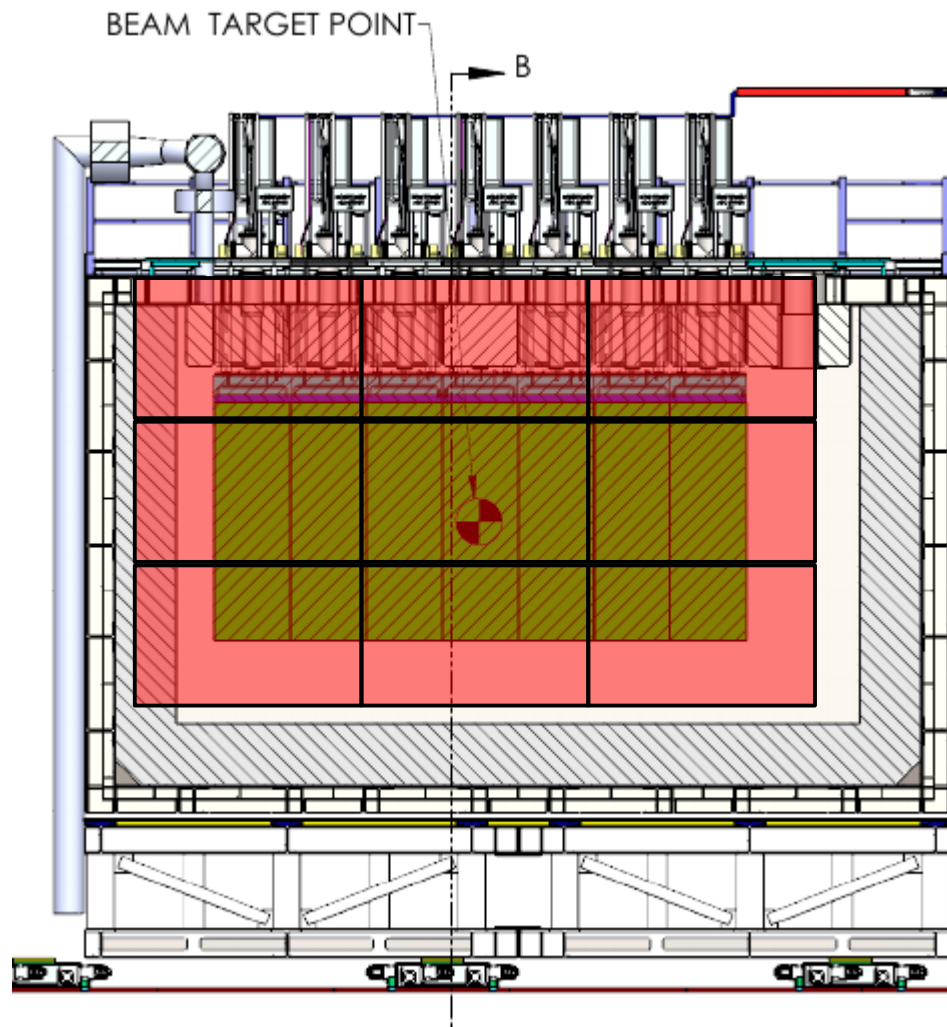
# Some numbers

- This design would have 232 rectangular scintillator strips per plane 4.5m long, mirrored on one end and read out on the other
- Strip dimension 3.85cm x 1cm x 450 cm
- Read out by 464 SiPMs coupled directly to fibers
  - Compare to 19,200 strips in TMS
  - TMS is studying groove vs. hole, we would just do the same thing
- Strips could be subdivided into smaller modules for assembly
  - i.e. 4 modules of 29 strips = one 4.5m half-plane
  - This would mean 16 modules of 112 x 450 cm

# TMS readout can be used for muon tagger

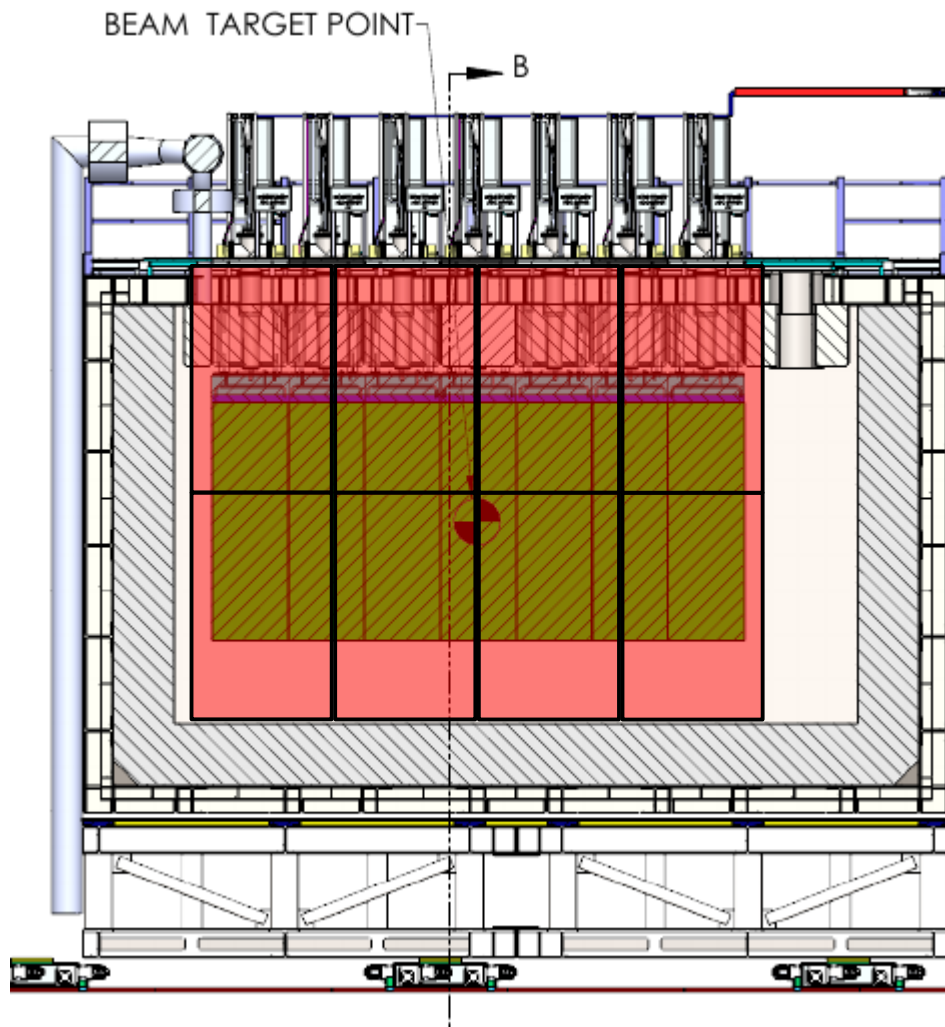
- Original TMS plan was commercial 16-channel SiPM (Hamamatsu S13361-2050AE-04), which is no longer available
- New strategy is single-channel SiPM per strip, with several options currently being considered
- Digitized by on-panel board, tentative ADC chip is 8-channel Texas Instrument AFE5807, 80 MSPS, outputs LVDS

# Could we use actual TMS modules?



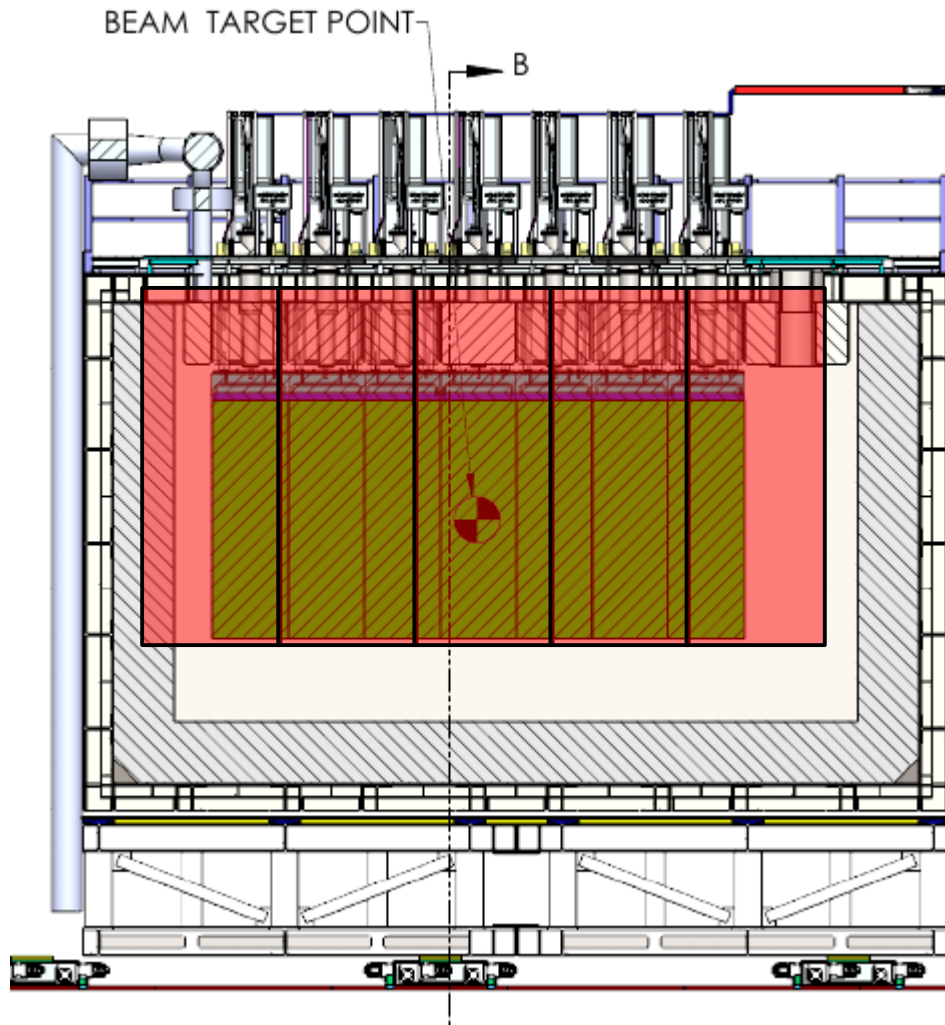
- 1 TMS module = 48 strips  
= 184 x 300 cm
- Pictured: 9 sideways modules (each 3m wide), 552cm instead of 450cm vertical coverage

# Could we use actual TMS modules?



- Other view would be 2x4: 600cm tall instead of 552cm, and 736cm wide instead of 900cm
- Or 2x5: 920cm wide
- Would require readout in the middle of the plane, which would need to be worked out but is doable

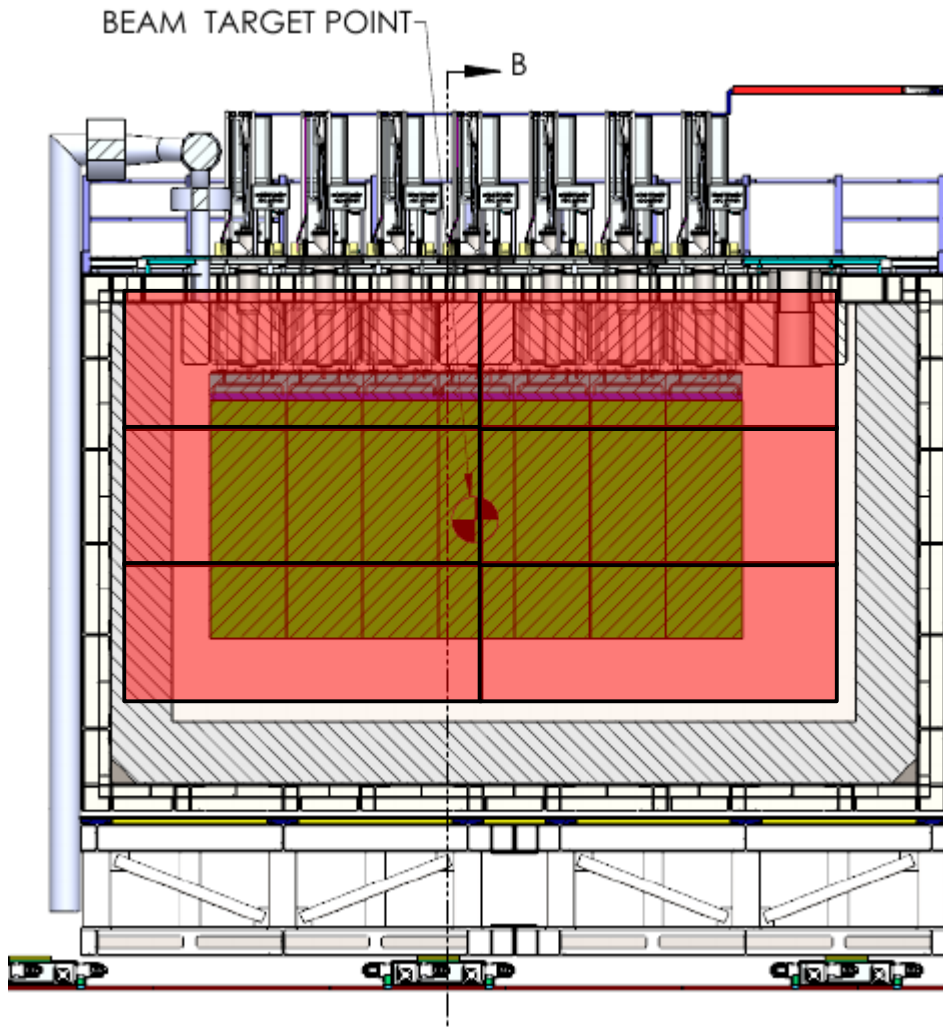
# Could we re-use existing system(s), like SBN CRTs?



- Top CRT modules are 4.5 x 1.8m with 16 11.2-cm strips
- 5 modules would cover the required area
- Position resolution would be worse – surface CRTs see ~10 muons per 2 ms readout window over similar area, we will see ~25 muons in 10 $\mu$ s



# Could we re-use existing system(s), like SBN CRTs?



- Other view would be covered by 6 top CRT modules
- What SBN CRTs will be available on the timescale required for ND-LAr?

# Do we need an actual CRT?

- Cosmic rate with our overburden is only  $\sim 2 \text{ Hz/m}^2$
- $35 \text{ m}^2$  for  $100 \mu\text{s}$  drift  $\rightarrow$  1 in-readout-window cosmic ray every few minutes  $\rightarrow$  not a problem
- We may want the ability to trigger on out-of-spill cosmic rays, but it's not clear that a top/bottom tagger is needed



# Do we need downstream muon tagger?

- I think this is worth studying
- Could simplify LAr → spectrometer matching and external calibration tag
- Marginal cost may be relatively small

# To do list

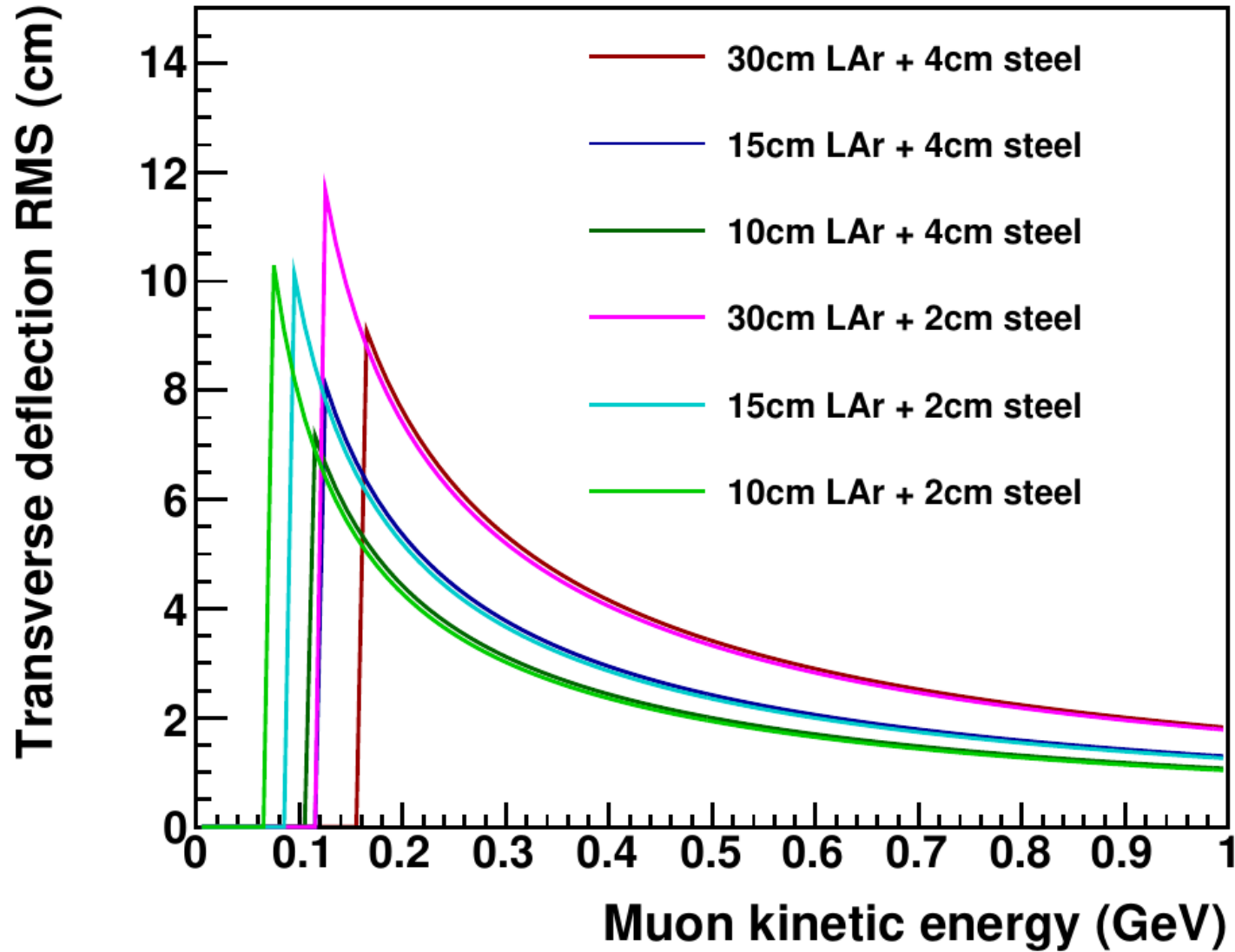
- Physics studies for design optimization
  - Update cryostat geometry, and simulate muon tagger upstream of cryostat
  - Coverage area optimization – is the current plan sufficient?
  - Impact of hadrons from near-hall rock interactions
  - View-matching
  - Estimate data rates, define DAQ requirements
  - Would SBN CRTs be sufficient?
  - Would downstream tag be useful?
- Write this up in a tech note – in progress
- Figure out what is needed to produce engineering design, cost estimates, etc.

# Conclusions

- An external muon tagger would provide a crucial additional handle on pile-up, and greatly simplify the problem of charge-light matching in ND-LAr
- It is feasible to exploit synergy with muon spectrometer requirements and re-use existing TMS design elements to keep costs low
- Opinion: this is a no-brainer for ND-LAr, and we should have put this in the design a long time ago

# Backups

# Multiple scattering is ~ few cm in cryostat → position resolution



- Transverse deflection due to scattering in cryostat for various configurations
- All basically plateau around a few cm for the energies that are of greatest interest for the external tagger system
- Position resolution better than a few cm is pointless