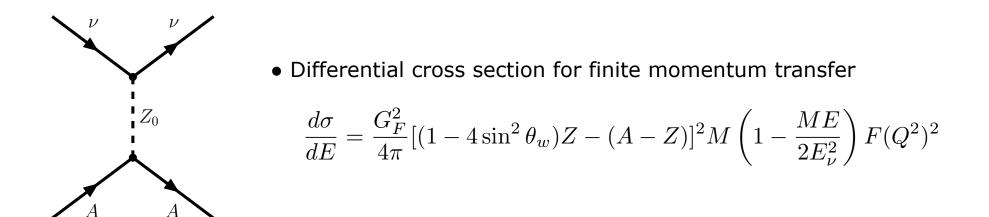
Coherent-NCvAS first predicted by Freedman in 1974 at Fermilab



• For most of the target nucleus, the coherence condition is fulfilled by neutrino energy of

$$E_{\nu} < \frac{1}{R_N} \simeq 50 \text{ MeV}$$
 $E_{max} \simeq \frac{2E_{\nu}^2}{M} \simeq \mathcal{O}(100) \text{ keV}$

Requires a ton-scale detector with ~10 keV energy threshold

$$R \simeq \mathcal{O}(10^3) \left(\frac{\sigma}{10^{-39} cm^2}\right) \times \left(\frac{\Phi}{10^{13} \nu / y ear/cm^2}\right) \times \left(\frac{M}{ton}\right) events/y ear$$

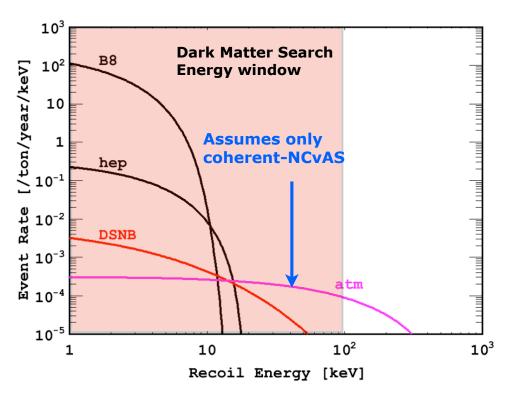
Recent innovation of Dark Matter detector technology makes it possible to access coherent-NCvAS

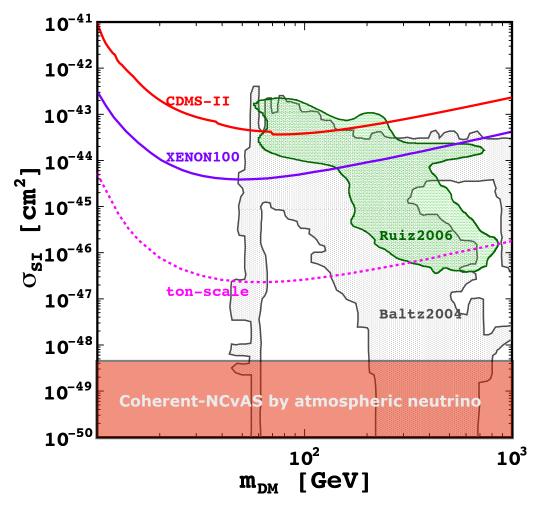
Neutrino Working Group Meeting (2011/10/24) Yoo

 Coherent scattering of atmospheric neutrino is an irreducible background in the future O(10 ton) scale Dark Matter experiments

$$\int_{A}^{\chi} \int_{A}^{\chi} \sigma_{\chi N} \simeq \frac{4}{\pi} \mu^2 [Zf_p + (A - Z)f_n]^2$$

$$\frac{dR}{dE} = \frac{\sigma_0}{m_{\chi}} \frac{A^2}{2\mu_n^2} F_A^2(E) \times \rho_0 \int_{v_m} \frac{f(v)}{v} dv$$





 Sensitivity of Dark Matter detectors will be saturated out by irreducible atmospheric neutrino backgrounds

From Booster Beam MC (S. Brice) (b) BNB neutrinos at $\cos\theta < 0.7$ - ν₋ 10² Unit · ν_u \overline{v}_{μ} v_{μ} from K [Arbitrary decay at rest 10 μ^{-} capture 1 50 100 150 200 250 0 300 \mathbf{E}_{v} [MeV] 10^{8} **BNB** neutrinos angular distribution 10 [v/cm²/sec] 20m from the target 10⁶ Flux 10⁵ ͺϻſ^ſᠺ᠂᠕᠕᠕᠕᠕᠕᠕᠕᠕᠕᠕᠕᠕᠕᠕᠕ 10⁴ -1 -0.8 -0.6 -0.4 -0.20 0.2 0.4 0.6 0.8

cosθ

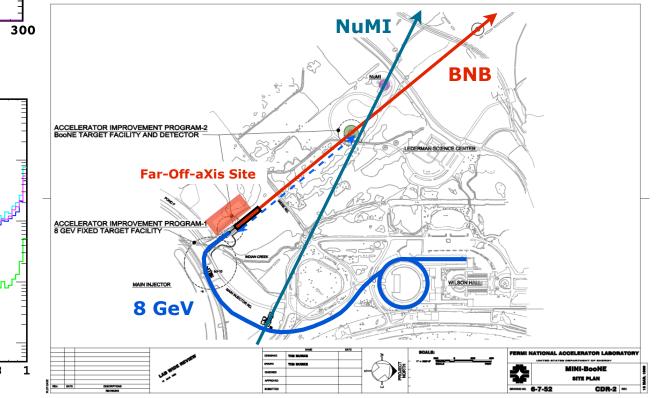
Beam MC Configuration

- Use standard Booster Beam MC
 - release stopping pion cuts in the original MC
- 8 GeV, 5Hz 5x10¹² Protons on Be target
 - 32 kW max power (w/ NUMI beam ~ 8 kW)

Dominant neutrino production process at the Far-Off-aXis is pion decay at rest

$$\pi^+
ightarrow \mu^+ +
u_\mu$$
 E(v_µ) =29.9 MeV
 $\mu^+
ightarrow e^+ + ar{
u}_\mu +
u_e$ delay = 2.2 µs

 $\Rightarrow \phi \sim 5 \times 10^5 \text{ v/cm}^2/\text{s} (@20m, \cos\theta < 0.5)$



• Neutrino Source at Oak Ridge SNS

1GeV proton on Hg, 24uC/pulse, 60Hz → 1.4MW $\phi(@20m) = 10^7 \text{ v/cm}^2/\text{s}$

but practical location is 46m away from the target

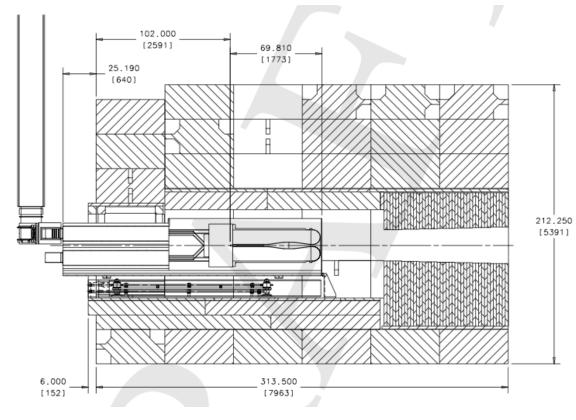
 $\phi(@46m) = 2x10^6 v/cm^2/s$

• Fermilab Far-Off-aXis BNB

8GeV proton on Be, 6.4uC/pulse, 5Hz \rightarrow 32kW $\varphi(@20m) = 5x10^5 \text{ v/cm}^2/\text{s}$

 $\phi(@10m) = 2x10^6 v/cm^2/s$

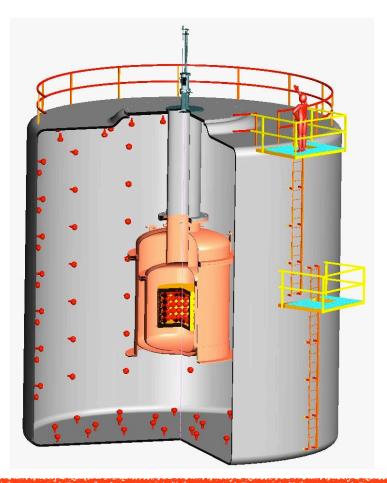
- \rightarrow beam induced neutron BG is the key issue
- \rightarrow requires a detailed MC study





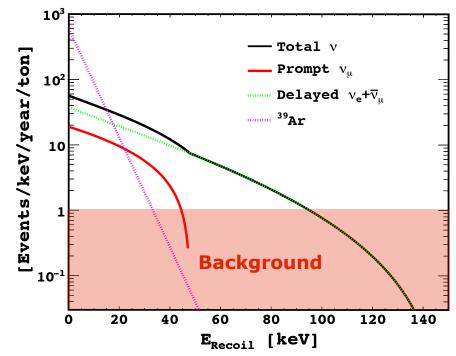
- Very impressive shielding at BNB target total ~2.5m-thick steel additional ~2.5m-thick concrete
- Ballpark estimation of beam induced neutron flux at 10m seem to be OK
 - \rightarrow need detailed MC study
- Cosmogenics would then important issue

A ton-scale LAr detector may perform the first ever observation of the coherent-NCvAS at Fermilab



Current Collaboration

Fermilab: S.Brice, F. DeJongh, S.Pordes, E.Ramberg, R. Tesarek, J.Yoo Duke: Kate Scholberg Indiana: Rex Tayloe



- Steady state background rejection factor $\sim 10^{-5}$
- Use pulse shape discrimination of nuclear recoil (fast) and electron recoil (slow) signal in LAr
- Well known detector technology (DEAP/CLEAN)
- Expected Event Rate in a single-phase LAr detector: ~200evt/ton/yr (Eth>30 keV, @20m, 32kW and w/ NUMI: ~50 evt/ton/yr)
- Project-X 8GeV, 0.3MW pulsed beam will provide intensive high energy neutrinos. The same beam will provide very precious byproduct low energy neutrinos with 4π coverage.

• We can make the first ever observation of Coherent-NCvAS at Fermilab