## Searching for New Particles Beyond the Standard Model by Proton Bremsstrahlung

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• Many proposals for new physics beyond the Standard Model predict novel, weakly interacting, light scalar or vector particles (e.g. Majorons, axions, Kaluza-Klein modes).

• Novel light particles could be responsible, among other things, for solving the strong CP problem in QCD, giving neutrinos their mass, or even explaining the origin of Dark Energy.

• Searches for such light particles will be extremely difficult at the high energy colliders designed to search for new, high mass particles.

• However, these light particles can be readily produced by proton bremsstrahlung! (Complementary to collider & dark matter searches)

# Why are neutrino experiments useful for new particle searches?

- Have a large number of protons on target (>10<sup>21</sup>!!)
- Usually located a fair distance from the target so that backgrounds from neutrino interactions are low (~1/r<sup>2</sup>), while the exotic particle flux is collimated and remains constant with r.
  - A spatially large detector contains these collimated events and makes identification easier.
  - Good beam and reconstruction timing (~nsec) allows for searches of heavy particles (> MeV).
- The "brem-ed" particles will be tightly collimated around the incoming proton beam direction (<~0.5 mrad) and will either decay or scatter in the center of the detector (assuming on-axis beam!)

## Beam-strahlung beam



# One example: "Paraphoton"<sup>1</sup>

- Gauge boson for a nearly unbroken  $U(1)_{B-L}$  symmetry
- Mass ~10 KeV (to evade bounds from 5<sup>th</sup> force measurements it needs to be short range)
- Astrophysics does not obviously exclude the light paraphoton because the mass is density dependent (in a supernova the particle becomes much more massive)
- Couples weakly (g<sup>2</sup>/α~10<sup>-9</sup>) to a "B-L" charge (for neutral matter this is proportional to number of nucleons)
- Generates an MSW-like potential in matter which affects neutrino oscillations in short baseline expts.
- Antineutrino and neutrino MSW-like effects are very different when caused by paraphotons
  <sup>1</sup>A. Nelson and J. Walsh (arXiv:0711.13663)

#### **Paraphoton Production via Bremsstrahlung**



#### **Paraphoton Production Forward Peaked**

 "beam-strahlung" cross section is forward peaked along beam direction and peaked at low energy



# Conclusions

- Neutrino Experiments can search with very high sensitivity for new scalar & vector boson production by proton bremsstrahlung
- These new particles will be produced in the forward directions and can either scatter or decay in the center of the detector
- Project X would provide 10x the current intensity (>2E21 POT per year)!
  - Coupled with a dedicated beam dump to reduce neutrino backgrounds, can make extremely sensitive searches.

# Backup

## Estimated Rates from the Paraphoton Model of Nelson & Walsh

 Described in arXiv:0711.1363, "Short Baseline Neutrino Oscillations and a New Light Gauge Boson", Ann Nelson & Jonathan Walsh

• The new light gauge vector boson ("paraphoton") has a mass of ~10 keV, a lifetime of ~2.5 ns, and a coupling strength of  $g^2/e^2 \sim 10^{-9}$ 

• BR(P -> νν) ~ 100%, BR(P->γγγ) ~ 10<sup>-7</sup>

• The paraphotons can be produced by hadronic bremsstrahlung (1/ $E_\gamma$ ) of the incident proton beam (~1%) in the forward direction (<5 mrad). Approximately all paraphotons will pass through the neutrino detector with little attenuation.

## **Paraphoton Decay In MB**

- Assume PID & Fiducial Volume efficiency ~ 50%
- Probability of decay in MB ~  $5m/\gamma c\tau = 5m/25km = 2E-4$
- Angle of decay products  $\sim 1/\gamma \sim 0.0001$
- #events = (2E20 POT)(1%)(1E-9)(2E-4)(1E-7)(50%) = 0.01 events in the forward direction

## **Paraphoton EM Shower In MB**

- Assume PID & Fiducial Volume efficiency ~ 50%
- Assume 50 cm radiation length
- Angular resolution ~3 degrees
- Number of radiation lengths in MB ~ 5m/50cm = 10
- #events = (2E20 POT)(1%)(1E-9)(1E-9)(10)(50%) = **10** events in the forward direction!