

Prospects of neutrino cross-section measurements and v_e selection with the NuMI off-axis beam at ICARUS

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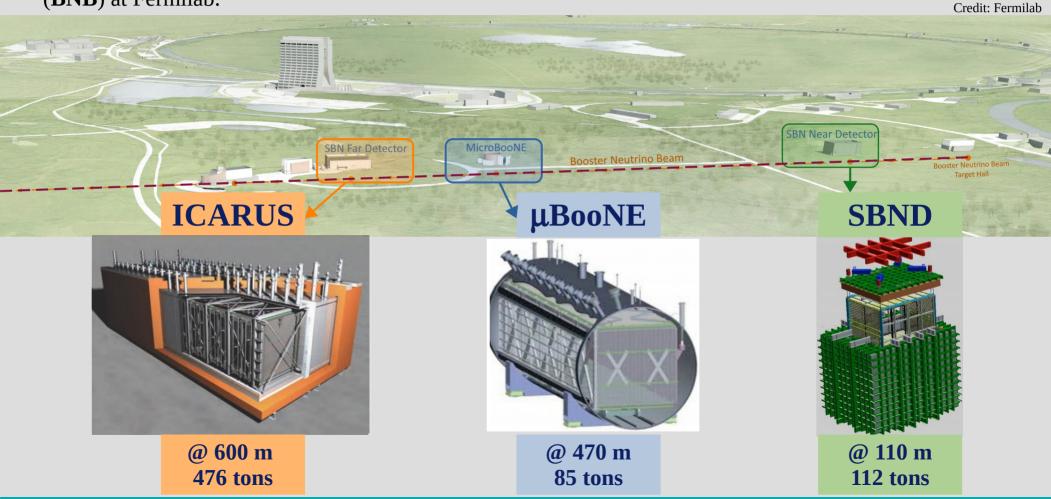
On behalf of the ICARUS collaboration



New Perspectives 2021, Fermilab August 18, 2021

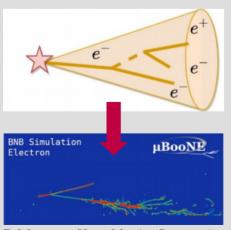
SBN Program at Fermilab

Three Liquid **Ar**gon **T**ime **P**rojection **C**hamber (**LArTPC**) detectors in the **B**ooster **N**eutrino **B**eamline (**BNB**) at Fermilab.

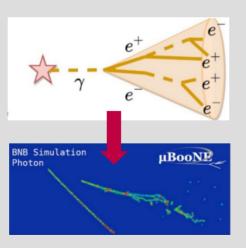


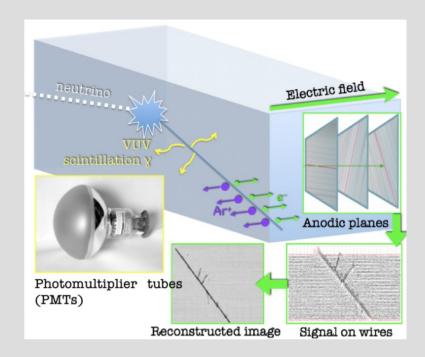
LAr TPCs

- LAr TPC detectors, provide **full 3D imaging**, **precise calorimetric energy reconstruction**, and efficient **particle identification**.
- The detailed images of particle trajectories provide significant information about final states.



E. Worcester, Users Meeting Presentation





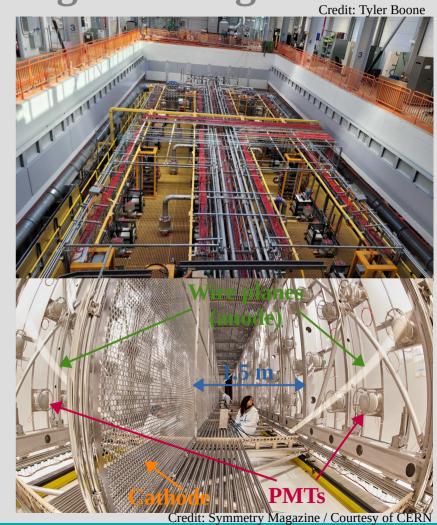
• Has the capability to distinguish electrons from photons. This by the gap between vertex and shower start for photons is $x_0 = 14$ cm and the dE/dx at shower start (e⁻ vs e⁺e⁻)

ICARUS

Imaging Cosmic And Rare Underground Signals

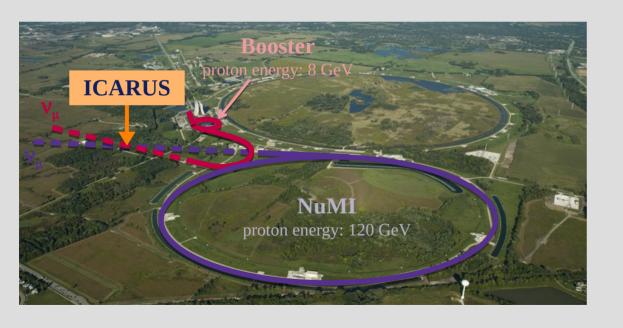
• Far detector in the SBN program

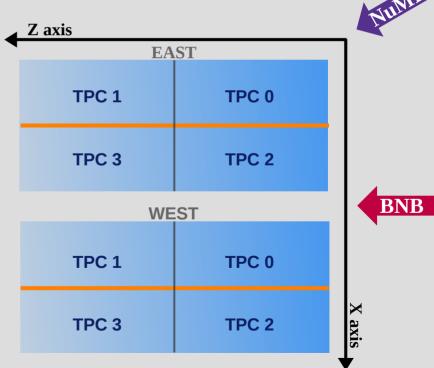
- It was operational in LNGS from 2010-2013,
- > Then refurbished at CERN.
- Two identical modules
 - \rightarrow 19.6 x 3.9 x 3.6 m³ each,
 - > **476 tons of LAr** active mass,
 - > **at 600 m** from the neutrino source.
- Two TPCs per module with central cathode
 - > 1.5 m drift lengths.
- **Three readout wire planes** (2 induction + 1 collection) per TPC
 - > ~ 54000 wires
 - \rightarrow at 0° and \pm 60°
 - with 3 mm pitch and plane spacing.
- **360 PMTs** for trigger and timing.
- Cosmic Ray Tagging System (CRT) at the bottom, side, and top.



ICARUS Two Neutrino Beams

ICARUS is located **on-axis with the Booster** beamline and **103 mrad off-axis from the NuMI** beamline, this will allow it to get a lot of data sets of ν -Ar interactions.





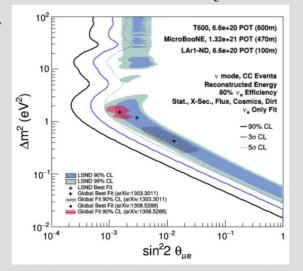
ICARUS Main Motivations Physics Program

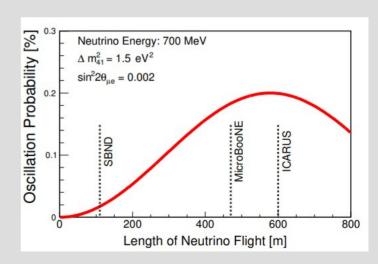
The SBN program has been designed specifically to address the sterile neutrino interpretation of the experimental at short-baseline anomalies.

- ICARUS as the SBN far detector will primarily be dedicated to the search for sterile neutrinos via $\nu_{\rm p}$ appearance.
- The combined analysis, the near detector (SBND) and ICARUS, with 3 years of data will be able to cover the 99% C.L. allowed region of the LSND signal with a sensitivity of 5 σ . Also, it will allow reducing both statistical and systematic uncertainties.

• Recently Neutrino-4* result, points to reactor anti- v_{ρ} disappearance with large Δm^2 (~ 7 eV²) and

mixing angle ($\sin^2 2\theta \sim 0.26$).





Cross-Section

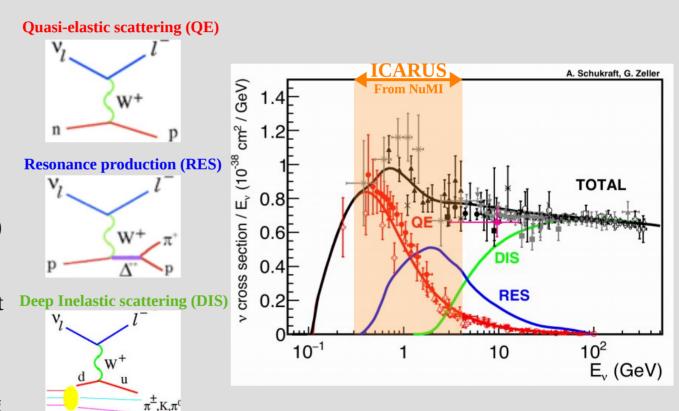
Neutrino Interactions

The ν oscillation experiments, require precise understanding of ν –Ar interaction cross section for a correct interpretation of the experimental result.

The ν cross section depends on:

- **v** interaction type (CC or NC) For CC
 - QE: nucleon changes, but NOT breaks up
 - RES: nucleon excites to resonance state
 - DIS: nucleon breaks up
- **v** target (e, nucleus, nucleon, q)
- v energy (MeV, GeV)

ICARUS will provide a large data set of v—Ar interactions from BNB and off-axis NuMI. Is particularly expected to have high statistics for v_e cross section measurement using the NuMI off axis.



In few GeV energy range, historically very few data

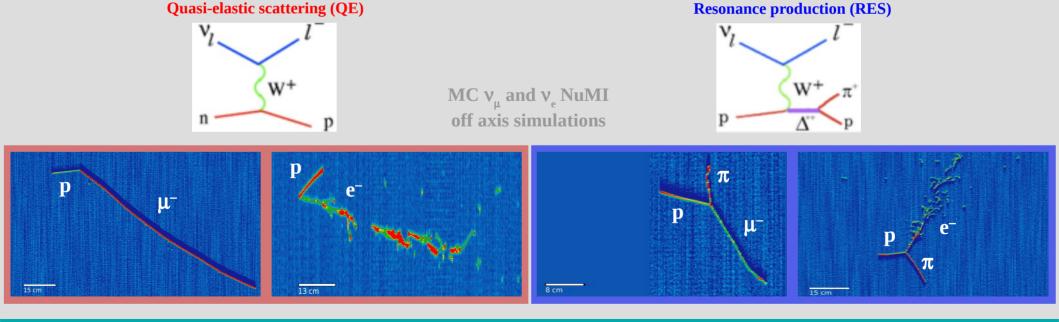
Cross-Section

Neutrino Interactions from NuMI Off Axis

The number of events from NuMI off axis (6×10^{20} POT) expected per year, are:

- v_{μ} CC = 433056 & v_{μ} NC = 191232
- v_a CC = 20608 & v_a NC = 7312

Mainly Quasi-Elastic and RESonance interactions are expected

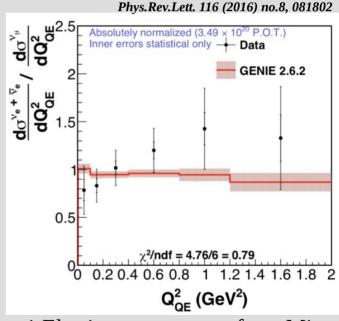


Cross-Section

Neutrino Interactions from NuMI Off Axis

Some examples of measurements that could be performed with the NuMI off axis in ICARUS include:

- to explore measurements of the electron neutrino cross section and muon to electron neutrino ratio cross section on Ar (since the statistic of these measurements is low),
- light dark matter searches
- CC non-standard interactions
- and so on.



Quasi-Elastic measurement from Minerva

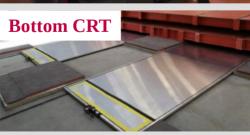
But, before measuring neutrino interactions, one needs to select the neutrinos...

V_e Selected Events Cosmic Rejection

ICARUS, as a surface detector, **faces an additional challenge to be constantly bombarded by cosmics**. In order to decrease its incidence in the detector, we use the CRT system and the PMTs:

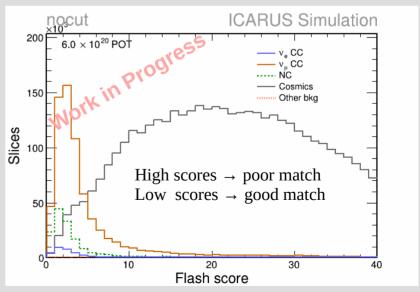
• **The CRT veto**, surrounds the cryostat with two layers of plastic scintillators, provides spatial and timing coordinates of the track entry point, match reconstructed CRT hits to activity in the TPC







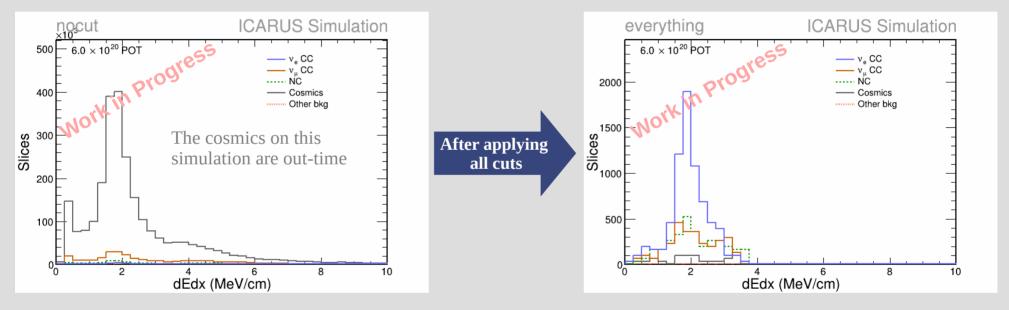
For PMT, the charge flash matching*
 associates ionized electrons (slow to read-out)
 with scintillation photons (fast to read-out). The
 main goals are to provide T₀ for each activity,
 identify a neutrino interaction from cosmics.



^{*} This simple flash match was developed by I. de Icaza and M. Stancari for SBND

V_e Selected Events from NuMI Off Axis

In addition to the CRT veto and the cut on the Flash Score, other cuts are applied in the v_e event selection, some of them in the Reconstruction, and in the fiducial volume.



- After applying all the selection cuts*, there are 56% of v_e CC and 44% of background** (of which v_u contributes with 14.6%, NC 28.1 % and cosmics are only 1.2%) in the sample.
- A similar analysis is ongoing for the BNB.

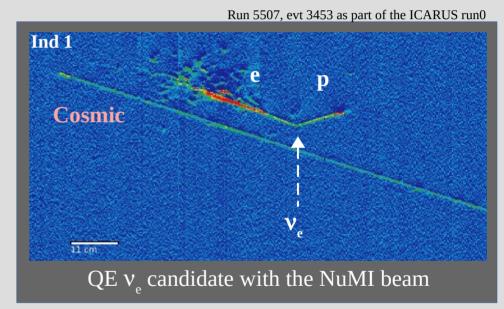
^{*} This is a preliminary event selection, criteria must be improved.

^{**} In this case, Background is all that is NOT $v_{_{\rho}}$ CC (i.e. $v_{_{II}}$, NC, cosmics).

V_e Selected Events from NuMI Off Axis

Currently, we are in the process of improving the cuts and testing another cosmic simulations (intime cosmics)

The next step will be to take this v_e event selection to test in **real data!!**

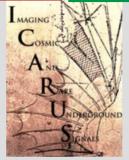


(this candidate was got by visual scanning)

Summary

- ICARUS as the Far Detector of the SBN program, has the goal to search for sterile neutrinos via v_{ρ} appearance.
- The understanding and characterization of neutrino interactions in Ar will be of great help in future experiments such as DUNE to investigate new and exciting physics.
- The v_e selection cut applied removed a great amount of cosmics (which is our principal background). However, this is still a preliminary event selection, an optimization and tunning will be done in the near future.
- The next round of ICARUS taking neutrino data is expected for this fall 2021, a lot of neutrino data awaits us!!





THANKS

